

A covariant calculation of deuteron GPDs

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April 18, 2018

Relativistic structure

Non-relativistically...

- Point particles are delta functions in configuration space.
Point particle described by plane wave in momentum space.
(These are related by Fourier transform.)
- Charge distribution is a function $\rho(\mathbf{r})$ of \mathbf{r} .
Momentum space description given by form factor $F(Q^2)$.
(These are related by Fourier transform.)

Relativistic structure

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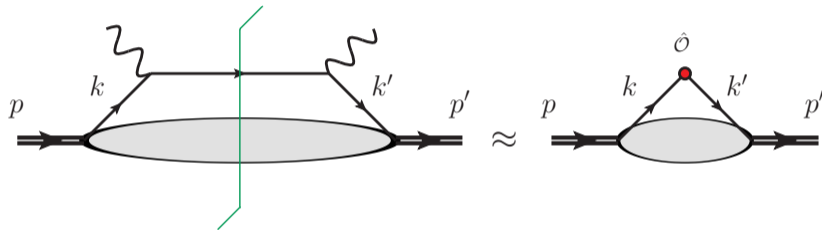
- Point particles are **not** delta functions in configuration space.
Point particle described by plane wave in momentum space.
(These are **not** related by Fourier transform.)
- Charge distribution is **not** a function $\rho(\mathbf{r})$ of \mathbf{r} .
Momentum space description given by form factor $F(Q^2)$.
(These are **not** related by Fourier transform.)

(*cf.* any comment by Jerry.)

- Relativity mucks up configuration space descriptions of many-body systems.
- We opt to take relativity seriously.
- Any description of 3D partonic structure should be a **Lorentz-invariant** function of **Lorentz-invariant** variables.
- **Generalized parton distributions** (GPDs) are exactly the kind of structure we need.

Generalized parton distributions

Look at **leading twist** GPDs. The **partonic correlator** is defined diagrammatically (leading twist approximation):



The operator \hat{O} depends on helicity-independent, axial, *etc.*:

$$\hat{O}_V = \not{n} \delta(n \cdot [xP - K])$$

$$\hat{O}_A = \not{n} \gamma_5 \delta(n \cdot [xP - K])$$

The GPDs are Lorentz-invariant functions that fall out of these matrix elements.

GPDs of the nucleon

An example: leading-twist GPDs of the nucleon.

$$\langle p', \lambda' | \not{n} \delta(n \cdot [xP - K]) | p, \lambda \rangle = \bar{u}(p', \lambda') \left[\not{n} H_N(x, \xi, t) + \frac{i\sigma^{n\Delta}}{2m_N} E_N(x, \xi, t) \right] u(p, \lambda)$$

- Similar to form factor decomposition.
- H_N and E_N are the nucleon GPDs.
- Defined wrt a lightlike four-vector n .
- Depend on **four Lorentz scalars**, the last being a factorization scale μ_F (dependence not notated here).
- Lorentz-covariant 3D structure!

$$\begin{aligned} P &= \frac{1}{2}(p + p') & x &= \frac{K \cdot n}{P \cdot n} \\ K &= \frac{1}{2}(k + k') & \xi &= -\frac{\Delta \cdot n}{2P \cdot n} \\ \Delta &= P' - P & t &= \Delta^2 \end{aligned}$$

Caveat: twist decomposition depends on choice of n .
cf. [V. Braun *et al.*, PRD89 (2014), 074022]

Choice of n should be informed by process.

Reductions of nucleon GPDs

Nucleon GPDs correspond to familiar quantities through limits or sum rules.

- Sum rules for Dirac and Pauli form factors:

$$\int_{-1}^1 H_N^q(x, \xi, t) dx = F_{1N}^q(t) \qquad \int_{-1}^1 E_N^q(x, \xi, t) dx = F_{2N}^q(t)$$

Note that ξ dependence vanishes in integral. (Special case of **polynomiality**)

- Forward limit for PDFs:

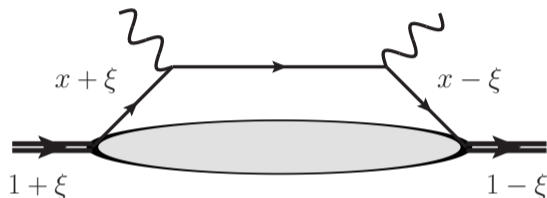
$$H_N^q(x, 0, 0) = q(x)$$

We use $x \in [-1, 1]$, with the meaning:

$$q(-x) = -\bar{q}(x)$$

The variables in the GPDs

Let's take a quick pictorial look at the GPD variables.



$$P = \frac{1}{2}(p + p')$$

$$K = \frac{1}{2}(k + k')$$

$$\Delta = P' - P$$

$$x = \frac{K \cdot n}{P \cdot n}$$

$$\xi = -\frac{\Delta \cdot n}{2P \cdot n}$$

$$t = \Delta^2$$

- Light cone fractions defined wrt $(P \cdot n)$.
- x is average momentum fraction of parton at both vertices.
- $x > 0$ quarks/gluons; $x < 0$ antiquarks/gluons.
- The **skewness** (2ξ) is the fraction lost by the target.

Differing conventions

Personally, I get confused because there are **many differing notational conventions**. I'll explicitly define the conventions I'm using here:

$$\begin{aligned} P &= \frac{1}{2}(p + p') & K &= \frac{1}{2}(k + k') \\ x &= \frac{K \cdot n}{P \cdot n} \in [-1, 1] & x_A &= Ax \in [-A, A] \\ \xi &= -\frac{\Delta \cdot n}{2P \cdot n} \in [-1, 1] & \xi_A &= A\xi \in [-A, A] \end{aligned}$$

Note that if I write x rather than x_A , it is **not scaled by A** .

Yes, there's a good reason to use x (rather than x_A) even for nuclear systems: higher Mellin moments in x_A pick up extra powers of A that make things messy and confusing!

I'll use x_A at the end, but it will be explicitly notated.

Covariance and Polynomiality

Polynomiality rules for the nucleon

Nucleon GPDs are known to obey **polynomiality sum rules** [X. Ji, J.Phys. G24 (1998) 1181]:

$$\int_{-1}^1 x^s H_N(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s A_{s+1,l}(t) (2\xi)^l + \text{mod}(s, 2) C_N(t) (2\xi)^{s+1}$$

$$\int_{-1}^1 x^s E_N(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s B_{s+1,l}(t) (2\xi)^l - \text{mod}(s, 2) C_N(t) (2\xi)^{s+1}$$

- A , B , and C are called **generalized form factors**.
- These rules are a **result of Lorentz covariance**.
- They are **violated for models that break covariance** (e.g., models with Fock space truncations or which use non-relativistic nuclear wave functions).

Spin-1 systems will have polynomiality rules too (due to Lorentz symmetry).

Polynomiality sum rules for the deuteron

I have derived the following sum rules for spin-1 systems (with $x \in [-1, 1]$ convention):

$$\int_{-1}^1 x^s H_1(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{A}_{s+1,l}(t) (2\xi)^l + \text{mod}(s, 2) \mathcal{F}_{s+1}(t) (2\xi)^{s+1}$$

$$\int_{-1}^1 x^s H_2(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{B}_{s+1,l}(t) (2\xi)^l$$

$$\int_{-1}^1 x^s H_3(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{C}_{s+1,l}(t) (2\xi)^l + \text{mod}(s, 2) \mathcal{G}_{s+1}(t) (2\xi)^{s+1}$$

$$\int_{-1}^1 x^s H_4(x, \xi, t) dx = \sum_{\substack{l=1 \\ 2 \nmid l}}^s \mathcal{D}_{s+1,l}(t) (2\xi)^l$$

$$\int_{-1}^1 x^s H_5(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^{s-1} \mathcal{E}_{s+1,l+1}(t) (2\xi)^l.$$

Only H_1 and H_3 (related to **electric charge distribution**, but not magnetic) have the $(2\xi)^{s+1}$ term.

Special cases of generalized form factors

The first Mellin moments ($s = 0$) give **electromagnetic form factors**:

$$\begin{aligned}\int_{-1}^1 H_1(x, \xi, t) dx &= F_1(t) & \int_{-1}^1 H_2(x, \xi, t) dx &= F_2(t) \\ \int_{-1}^1 H_3(x, \xi, t) dx &= F_3(t) & \int_{-1}^1 H_4(x, \xi, t) dx &= \int_{-1}^1 H_5(x, \xi, t) dx = 0.\end{aligned}$$

Forward limits relate the GPDs to **parton distribution functions**:

$$H_1(x, 0, 0) = f(x) \qquad H_5(x, 0, 0) = b_1(x)$$

(No forward limits defined for the other GPDs.)

Combining these entails valence version of Kumano-Close sum rule:

$$\int_{-1}^1 b_1(x) dx = 0$$

Violation of the usual Kumano-Close sum rule is possible, and would indicate tensor polarization in the sea.

Special cases of generalized form factors

The second Mellin moments ($s = 1$) give **gravitational form factors**:

$$\begin{aligned}\int_{-1}^1 xH_1(x, \xi, t)dx &= \mathcal{G}_1(t) + (2\xi)^2\mathcal{G}_3(t) & \int_{-1}^1 xH_2(x, \xi, t)dx &= \mathcal{G}_5(t) \\ \int_{-1}^1 xH_3(x, \xi, t)dx &= \mathcal{G}_2(t) + (2\xi)^2\mathcal{G}_4(t) & \int_{-1}^1 xH_4(x, \xi, t)dx &= (2\xi)\mathcal{G}_6(t) \\ \int_{-1}^1 xH_5(x, \xi, t)dx &= \mathcal{G}_7(t) & \sum_{\text{partons}} \mathcal{G}_7(t) &= -\frac{t}{2M_D^2} \sum_{\text{partons}} \mathcal{G}_6(t)\end{aligned}$$

(Notation from Simonetta's papers, modulo small differences like factors of 2.)

The last relation is a consequence of energy-momentum conservation.

It also entails:

$$\sum_{\text{partons}} \int_{-1}^1 xb_1(x)dx = 0$$

Violation of this sum rule over quarks alone would indicate momentum-sharing with gluons.

GFFs and the matter distribution

One can construct a **monopole gravitational form factor** for a spin-one system:

$$\mathcal{G}_N(t) = \left(1 + \frac{2}{3}\tau\right) \mathcal{G}_1(t) - \frac{2}{3}\tau \mathcal{G}_5(t) + \frac{2}{3}\tau(1 + \tau) \mathcal{G}_2(t)$$

where $\tau = -t/(4M_D^2)$.

(Analogous to the Coulomb electric form factor. N is for Newton.)

Related to a **gravitational radius**:

$$\langle r_G^2 \rangle = 6 \frac{d}{dt} [\mathcal{G}_N(t)]$$

This should look very familiar to the charge radius! But:

$$r_G \neq r_E,$$

The model I'll present soon gives:

$$r_E = 2.09 \text{ fm}$$

$$r_G = 1.95 \text{ fm}$$

GFFs and much more

There's even more information hidden in GFFs.

- That's either **generalized** or **gravitational** form factor.

Taneja *et al.* (Phys.Rev. **D86** (2012) 036008) tell us that

$$J(t) = \frac{1}{2}\mathcal{G}_5(t)$$

We can learn how quarks and gluons share and distribute their angular momentum!

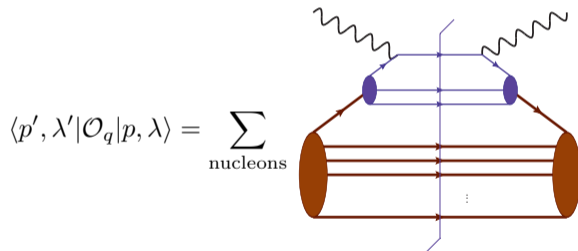
Also, the **purely spacelike** components of the stress-energy tensor give **shears and pressures**:

- $\mathcal{G}_3(t)$, $\mathcal{G}_4(t)$, and $\mathcal{G}_6(t)$ all contribute to these.
- In particular, both $\mathcal{G}_3(t)$ and $\mathcal{G}_6(t)$ contribute to Polyakov's and Schweitzer's coveted static quantity D .
- This is an ongoing topic of research with Simonetta, Whit, and Ian.

Convolution Formalism

GPD convolution

- A real nucleus has modifications (**EMC effect**—the whole point of this workshop!).
- But the “EMC effect” exists as a *discrepancy with a baseline prediction*.
- So what do nuclear GPDs look like if we assume unmodified nucleons?



Unmodified convolution formalism

To have an EMC effect for the GPDs, we need an **unmodified convolution formalism**. This is ostensibly straightforward:

- Get a model for the nucleon GPDs H_N and E_N .
- Compute the matrix element

$$\langle p', \lambda' | \left[\not{n} H_N + \frac{i\sigma^{n\Delta}}{2m_N} E_N \right] | p, \lambda \rangle$$

assuming pointlike, on-shell nucleons.

(The factors H_N and E_N fold in the non-pointlike structure.)

- An **ambiguity** arises: identities like Gordon decomposition that are true for on-shell nucleons **will lead to different results for kinematically off-shell nucleons**.
- This turns out to matter for the nucleon D-terms.

The D-term and Gordon decomposition

In models such as [Goeke *et al.*, Prog. Part. Nucl. Phys. 47 (2001)], the nucleon GPD is broken into a **double distribution** and a **D-term**:

$$H_N(x, \xi, t) = H_{DD}(x, \xi, t) + D\left(\frac{x}{\xi}, t\right) \quad E_N(x, \xi, t) = E_{DD}(x, \xi, t) - D\left(\frac{x}{\xi}, t\right)$$

- The D-term here contributes to the $(2\xi)^{s+1}$ GFF in the polynomiality sum rules.
- The same D-term enters both H_N and E_N with opposite sign.
- **This is due to Lorentz invariance.** [X. Ji, J.Phys. G24 (1998) 1181]

Using **Gordon decomposition**, we can write:

$$\bar{u}(\mathbf{p}', \sigma') \left[\not{n} H_N + \frac{i\sigma^{n\Delta}}{2m_N} E_N \right] u(\mathbf{p}, \sigma) = \bar{u}(\mathbf{p}', \sigma') \left[\not{n} H_{DD} + \frac{i\sigma^{n\Delta}}{2m_N} E_{DD} + \frac{p \cdot n}{m_N} D_N \right] u(\mathbf{p}, \sigma)$$

for *on-shell* spinors.

We must decide between the LHS and RHS for the “unmodified” deuteron GPD.
(I've chosen the RHS since it emphasizes there is *one nucleon D-term*.)

The master convolution formula

Evaluating the matrix element

$$\langle p', \lambda' | \left[\not{n} H_{DD} + \frac{i\sigma^{n\Delta}}{2m_N} E_{DD} + \frac{P_N \cdot n}{m_N} D_N \right] | p, \lambda \rangle$$

gives a **master convolution formula**:

$$H_i(x, \xi, t) = \int_{-1}^1 \frac{dy}{y} \left[h_i(y, \xi, t) H_{DD} \left(\frac{x}{y}, \frac{\xi}{y}, t \right) + e_i(y, \xi, t) E_{DD} \left(\frac{x}{y}, \frac{\xi}{y}, t \right) + y d_i \left(\frac{y}{\xi}, t \right) D_N \left(\frac{x}{\xi}, t \right) \right]$$

(*cf.* also work by Sergio, Simonetta.)

- h_i , e_i , and d_i describe how the nucleons are distributed in the nucleus, using GPD language. Call them **generalized nucleon distributions** (GNDs).
- By construction, H_{DD} , E_{DD} , and D_N already obey polynomiality.
- **We can prove that when the GNDs obey polynomiality sum rules, so do the deuteron GPDs.**
- **The only ingredient needed to ensure the GNDs observe polynomiality is a Lorentz-covariant model of nuclear structure.**

Discrete convolution formulas

The master convolution formula entails a collection of **discrete convolution formulas** for the generalized form factors.

- Define “generalized body form factors” from the GNDs, e.g.,

$$\int_{-1}^1 y^s h_1(y, \xi, t) dy = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{A}_{s+1,l}^H(t) (2\xi)^l + \text{mod}(s, 2) \mathcal{F}_{s+1}^H(t) (2\xi)^{s+1}$$

with similar relations for all i , for e_i , and for d_i .

- The full GFFs can be found through, e.g.:

$$\mathcal{A}_{s+1,l}(t) = \sum_{\substack{r=0 \\ 2|r}}^l [\mathcal{A}_{s-r+1,l-r}^H(t) \mathcal{A}_{s+1,r}^N(t) + \mathcal{A}_{s-r+1,l-r}^E(t) \mathcal{B}_{s+1,r}^N(t)]$$

- Similar relations for \mathcal{B} through \mathcal{E} , except \mathcal{D} involves odd r .
- \mathcal{F} and \mathcal{G} are slightly more complicated (they have an extra term).

Discrete convolution formulas

Convolution formulas for \mathcal{F} and \mathcal{G} :

$$\mathcal{F}_{s+1}(t) = \sum_{\substack{r=0 \\ 2|r}}^s [\mathcal{F}_{s-r+1}^H(t)A_{s+1,r}^N(t) + \mathcal{F}_{s-r+1}^E(t)B_{s+1,r}^N(t)] + \mathcal{A}_{1,0}^D \mathcal{C}_{s+1}^N(t)$$

$$\mathcal{G}_{s+1}(t) = \sum_{\substack{r=0 \\ 2|r}}^s [\mathcal{G}_{s-r+1}^H(t)A_{s+1,r}^N(t) + \mathcal{G}_{s-r+1}^E(t)B_{s+1,r}^N(t)] + \mathcal{C}_{1,0}^D \mathcal{C}_{s+1}^N(t)$$

A reminder that these originate from Mellin moments of H_1 and H_3 :

$$\int_{-1}^1 x^s H_1(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{A}_{s+1,l}(t) (2\xi)^l + \text{mod}(s, 2) \mathcal{F}_{s+1}(t) (2\xi)^{s+1}$$

$$\int_{-1}^1 x^s H_3(x, \xi, t) dx = \sum_{\substack{l=0 \\ 2|l}}^s \mathcal{C}_{s+1,l}(t) (2\xi)^l + \text{mod}(s, 2) \mathcal{G}_{s+1}(t) (2\xi)^{s+1}$$

Utility of discrete convolution

There is a point to these discrete convolutions.
Say we want **deuteron electric form factors**:

$$F_{iD}(t) = F_{iV}(t)[F_{1p}(t) + F_{1n}(t)] + F_{iT}(t)[F_{2p}(t) + F_{2n}(t)]$$

- F_{iV} and F_{iT} (“body form factors”) are matrix elements of \not{n} and $i\sigma^{n\Delta}/(2m_N)$, respectively.
- This is an easy calculation, bypassing a full GPD computation followed by an additional integration over x (which would be expensive in core-hours!).

Say we want **deuteron gravitational form factors**:

$$\mathcal{G}_i(t) = \mathcal{G}_{iV}(t)[A_p(t) + A_n(t)] + \mathcal{G}_{iT}(t)[B_p(t) + B_n(t)] + \mathcal{G}_{(i-2)D}(t)[C_p(t) + C_n(t)]$$

(the pink terms contribute only for $i = 3, 4$).

- \mathcal{G}_{iV} , \mathcal{G}_{iT} , and $\mathcal{G}_{(i-2)D}$ are matrix elements of $i(n \cdot \overleftrightarrow{\partial})\not{n}$, $i(n \cdot \overleftrightarrow{\partial})\sigma^{n\Delta}/(2m_N)$, and $1/m_N$, respectively.
- This again bypasses an expensive computation.

Covariant Contact Model

Motivation for a contact model

For computing the GPDs themselves, covariance is of the utmost importance.

- Can be difficult to maintain covariance while solving a bound state equation.
- Covariantly solving a four-Fermi contact interaction is tractable.
- Success of the Nambu-Jona-Lasinio (NJL) model suggests this approach has promise.
- The skeptic may ask: what about the deuteron's D-wave?

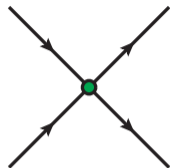
What about the deuteron's huge quadrupole moment?

- The magic of relativity will produce these things, even in a contact interaction model.

Lagrangian

Construct most general possible NN Lagrangian that:

- Has four-fermi contact interactions.
- Has no derivatives in interaction terms.
- Obeys $SU(2)_V \times SU(2)_A$ isospin symmetry.
- Satisfies Pauli exclusion principle (enforced by ψ being Grassmann-number-valued!).

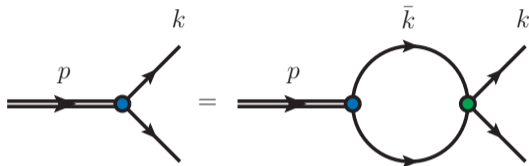


$$\begin{aligned}\mathcal{L}_{NN} = & \bar{\psi}(i\not{\partial} - m)\psi \\ & - G_S [(\bar{\psi}\tau_j C\tau_2\bar{\psi}^T) (\psi^T C^{-1}\tau_2\tau_j\psi) - (\bar{\psi}\tau_j\gamma^5 C\tau_2\bar{\psi}^T) (\psi^T C^{-1}\tau_2\gamma^5\tau_j\psi)] \\ & - G_V [(\bar{\psi}\tau_j\gamma^5\gamma^\mu C\tau_2\bar{\psi}^T) (\psi^T C^{-1}\tau_2\gamma^5\gamma_\mu\tau_j\psi) + (\bar{\psi}\gamma^\mu C\tau_2\bar{\psi}^T) (\psi^T C^{-1}\tau_2\gamma_\mu\psi)] \\ & - \frac{1}{2}G_T [(\bar{\psi}i\sigma^{\mu\nu} C\tau_2\bar{\psi}^T) (\psi^T C^{-1}\tau_2i\sigma_{\mu\nu}\psi)]\end{aligned}$$

Neglect charge-symmetry violation (assume $m_p = m_n \equiv m_N$).
Interactions decouple into separate isoscalar and isovector sectors.

Bethe-Salpeter vertex

Bethe-Salpeter equation in the covariant contact model:



Solution is the Bethe-Salpeter vertex:

$$\Gamma_D(p, \lambda) = \left[\alpha_V \not{p}(p, \lambda) + i\alpha_T \frac{\sigma^{\varepsilon p}}{2M_D} \right] C\tau_2$$

We can solve for α_V and α_T in terms of couplings G_V and $G_T \dots$ and a UV regulator Λ (from proper time regularization).

Solution and static observables

Solution has parameters: G_V , G_T , and Λ .
These must be chosen somehow.

Fit to static observables:

- Deuteron binding energy
- Deuteron electromagnetic moments
- 3S_1 - 3D_1 scattering parameters.

	Contact model	Empirical
ϵ_D (MeV)	2.18	2.22
r_E (fm)	2.09	2.14
μ_D	0.879	0.857
Q_D (fm ²)	0.285	0.286
3a_1 (fm)	5.26	5.42
3r_1 (fm)	1.78	1.76
Λ (MeV)	139	—
G_V (GeV ⁻²)	-683	—
G_T (GeV ⁻²)	-715	—

- $\Lambda = 139$ MeV is a result of a fit—is not chosen by us.
- Suggests the model “knows” it breaks down when pion exchange becomes relevant.
- Note we have a non-zero, almost correct quadrupole moment.
- We do actually have a D-wave!

Origin of the D-wave

Whence the D-wave?

Bethe-Salpeter wave function takes the form

$$\psi_D(p, k, \lambda) = S(k)\Gamma_D(p, \lambda)S^T(p - k)$$

The numerator of the top-right 2×2 corner (where both nucleons have **positive energy**):

$$\psi_D^{(++)}(p, k, \lambda) \propto m_N(M_D + m_N)(\alpha_V + \alpha_T)(\boldsymbol{\varepsilon} \cdot \boldsymbol{\sigma}) + 2(\alpha_V - \alpha_T)(\mathbf{k} \cdot \boldsymbol{\varepsilon})(\mathbf{k} \cdot \boldsymbol{\sigma})$$

D-wave comes from second part of structure.

Ensures that even non-relativistic reductions, with:

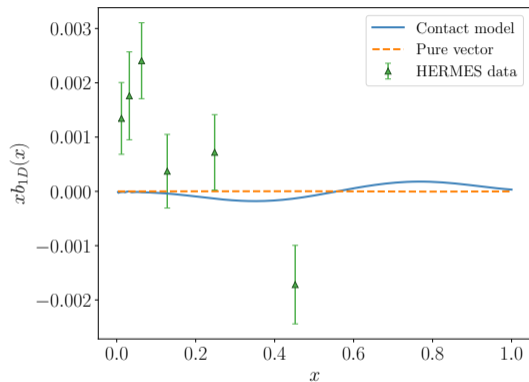
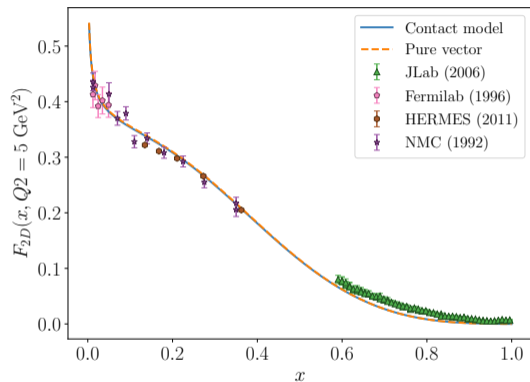
$$\psi_{\text{NR}}(p, k, \lambda) \propto \bar{u}(k, s_1)\Gamma_D(p, \lambda)\bar{u}^T(p - k, s_2)$$

have D-wave—that is, $(\mathbf{k} \cdot \boldsymbol{\varepsilon})(\mathbf{k} \cdot \boldsymbol{\sigma})$ terms—in them.

Answer to **whence**: the **lower components** of u ! **This is a relativistic effect.**

DIS structure functions

How well can this model describe DIS structure functions?
(Use CJ15 for nucleon PDFs.)

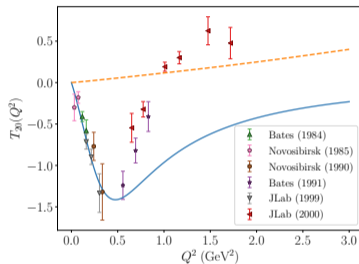
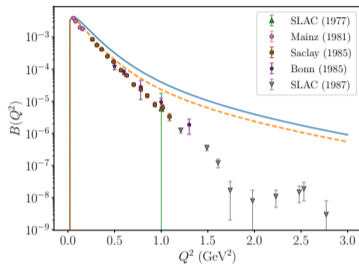
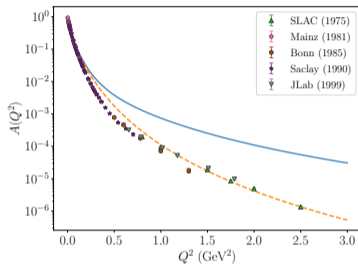


- Not bad for $F_2(x, Q^2)$ (underestimate at high x due to lack of short range correlations).
- Doesn't describe HERMES data for $b_1(x, Q^2)$, but that's expected.

Electromagnetic form factors

What about electromagnetic form factors?

(Use Kelly-Riordan nucleon form factors ... Blue is full model, orange is G_V only.)



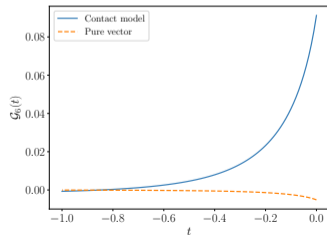
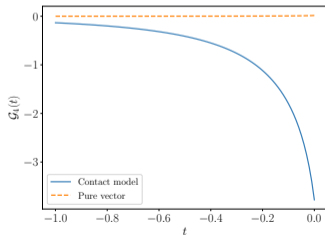
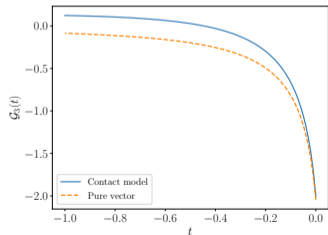
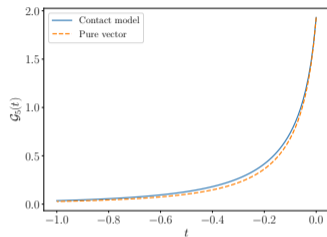
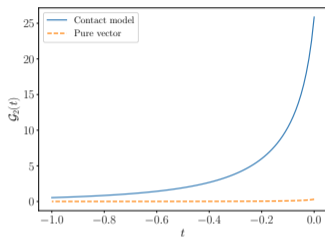
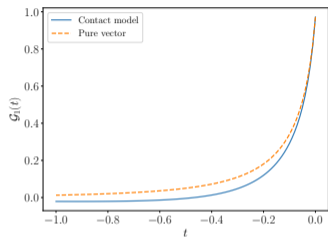
- Absolute size is too big at moderate-to-large Q^2 .
- Agreement is OK for $Q^2 \lesssim 0.5 \text{ GeV}^2$.
- Suggests our GPDs will be applicable to only low $-t$.

Gravitational form factors

Let's see what the **gravitational** form factors look like.

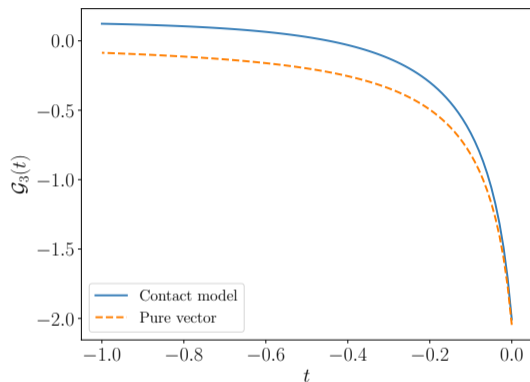
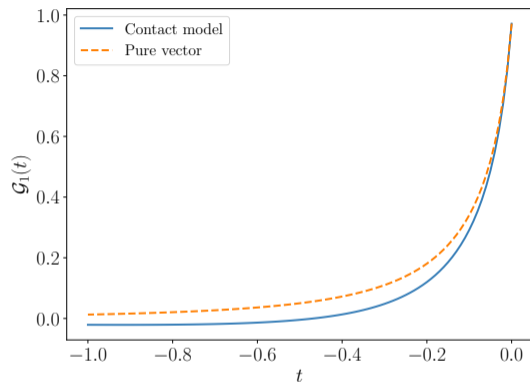
(Summed over all partons ... nucleon GPD model from [Goeke *et al.*, Prog. Part. Nucl. Phys 47 (2001)].)

No data due to lack of graviton exchange dominated scattering experiments...



Gravitational form factors

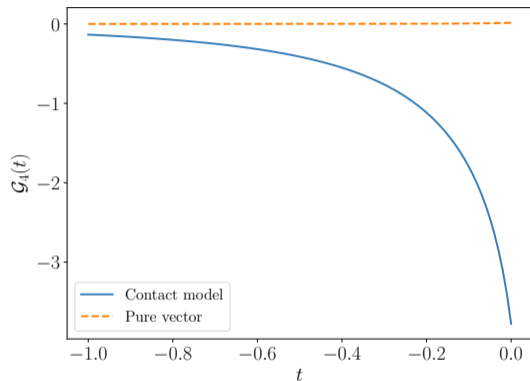
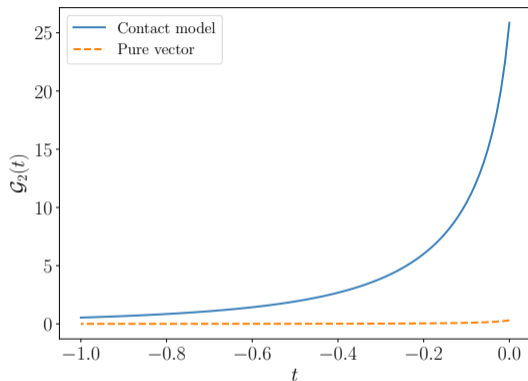
Features of the monopole-dominated GFFs.



- $\mathcal{G}_1(t)$ is the analogue of electromagnetic $F_1(t)$.
- $\mathcal{G}_1(0) = 1$ is a mass sum rule—it's a statement of **energy conservation**.
- $\mathcal{G}_3(t)$ dominates the Polyakov-Schwitzer D . It's related to internal balancing of forces.

Gravitational form factors

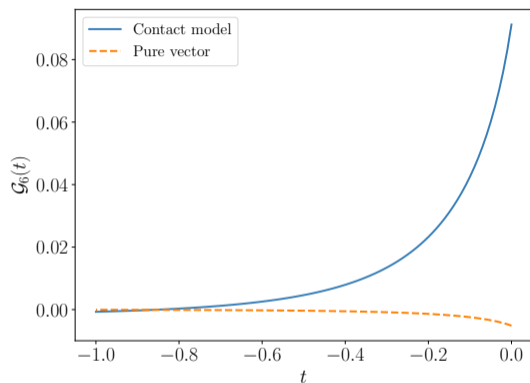
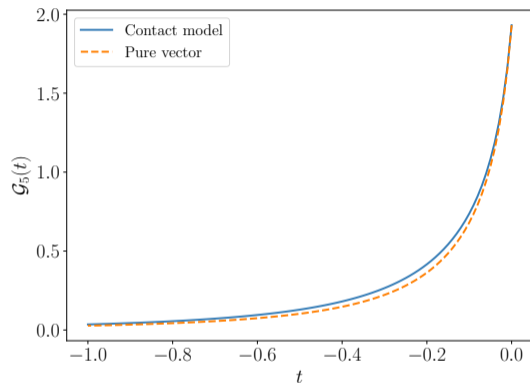
Features of the quadrupole-dominated GFFs.



- $\mathcal{G}_2(t)$ is the analogue of electromagnetic $F_3(t)$.
- Deuteron has a *large gravitational quadrupole moment*. Large tidal forces for a small nucleus.
- $\mathcal{G}_4(t)$ is a $(2\xi)^{s+1}$ term—it seems like a “second deuteron D-term”—mandated by the deuteron’s quadrupole moment.

Gravitational form factors

Features of the other GFFs.

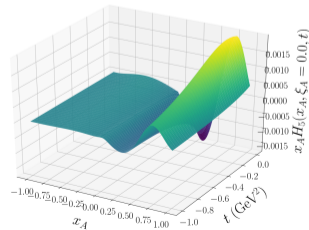
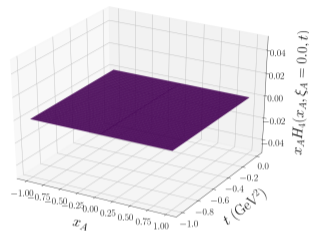
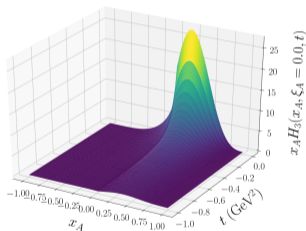
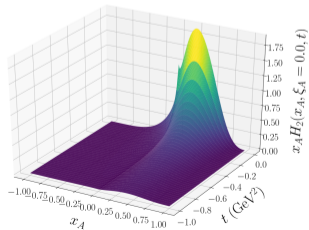
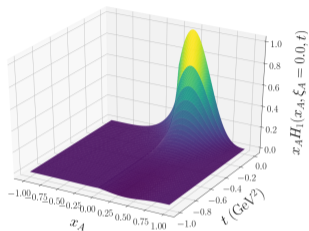


- $\mathcal{G}_5(t)$ is angular momentum distribution. (See Simonetta's talk.)
- $\mathcal{G}_6(t)$ is a tensor polarization. (It's small, like $b_1(x)$.)

Generalized parton distributions

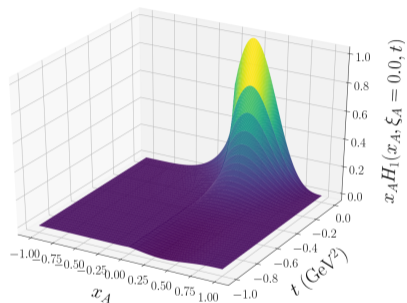
So, what do the coveted deuteron GPDs actually look like?

Look at $\xi = 0$; has clearer correspondences with well-known PDFs and form factors.

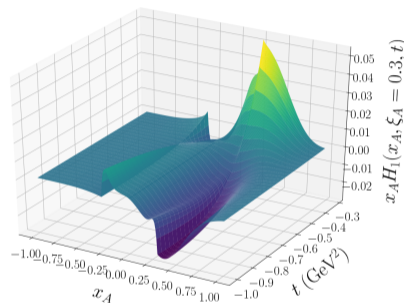


Generalized parton distributions

Let's take a closer look... at $H_1(x_A, \xi_A, t)$.



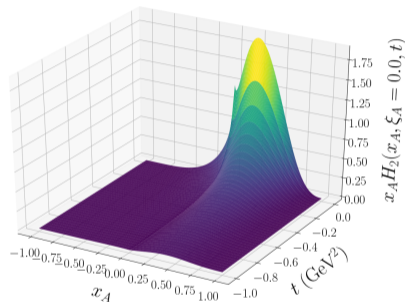
finite skewness \rightarrow



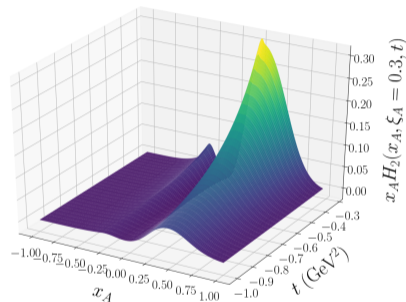
- $H_1(x_A, \xi_A, t)$ is monopole-dominated.
- Forward limit ($\xi_A = 0, t = 0$) is old-fashioned PDF.
- Region $-\xi_A < x_A < \xi_A$ (“ERBL region”) is dominated by D-term.

Generalized parton distributions

Closer look at $H_2(x_A, \xi_A, t)$.



finite skewness \rightarrow



- $H_2(x_A, \xi_A, t)$ is spin-dominated.
- Ridges at $x_A = \pm \xi_A$, despite lack of D-term.

Generalized parton distributions

Closer look at $H_3(x_A, \xi_A, t)$.



- $H_3(x_A, \xi_A, t)$ is quadrupole-dominated.
- The absolute magnitude is large because of the large quadrupole moment.
- ERBL region is dominated by D-term.

Generalized parton distributions

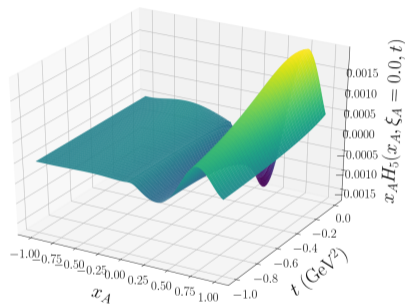
Closer look at $H_4(x_A, \xi_A, t)$.



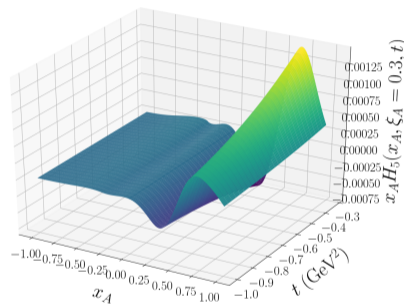
- $H_4(x_A, \xi_A, t)$ is dominated by tensor polarization.
- This GPD has neither a forward limit, nor a relation to EM form factors.
- It's zero at $\xi_A = 0$ because it is odd in ξ_A .

Generalized parton distributions

Closer look at $H_5(x_A, \xi_A, t)$.



finite skewness \rightarrow



- $H_5(x_A, \xi_A, t)$ is tensor polarization dominated.
- Forward limit ($\xi_A = 0, t = 0$) is partonic $b_1(x)$.

Conclusions and outlook

In conclusion:

- We have calculated deuteron GPDs in a manifestly covariant contact model.
- Our GPDs obey polynomiality sum rules, and allow an unambiguous extraction of generalized form factors.
- We have computed gravitational form factors within this model, too.

Future work to be done:

- The model will be extended to other light nuclei (triton and helium).
- The NJL model can be used to compute covariant *nucleon* GPDs.
- A deeper understanding of the stress-energy tensor and gravitational form factors for spin-1 systems is needed.