

Nambu-Covariant Many-Body Theory

A Green's functions framework for superfluid systems

Dr Arnau Rios Huguet

Institute of Cosmos Sciences
Universitat de Barcelona
&
Department of Physics
University of Surrey

ECT* Workshop
Nuclear Physics Meets Condensed Matter

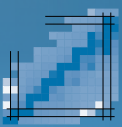
19 July 2021
ECT*/Online



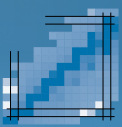
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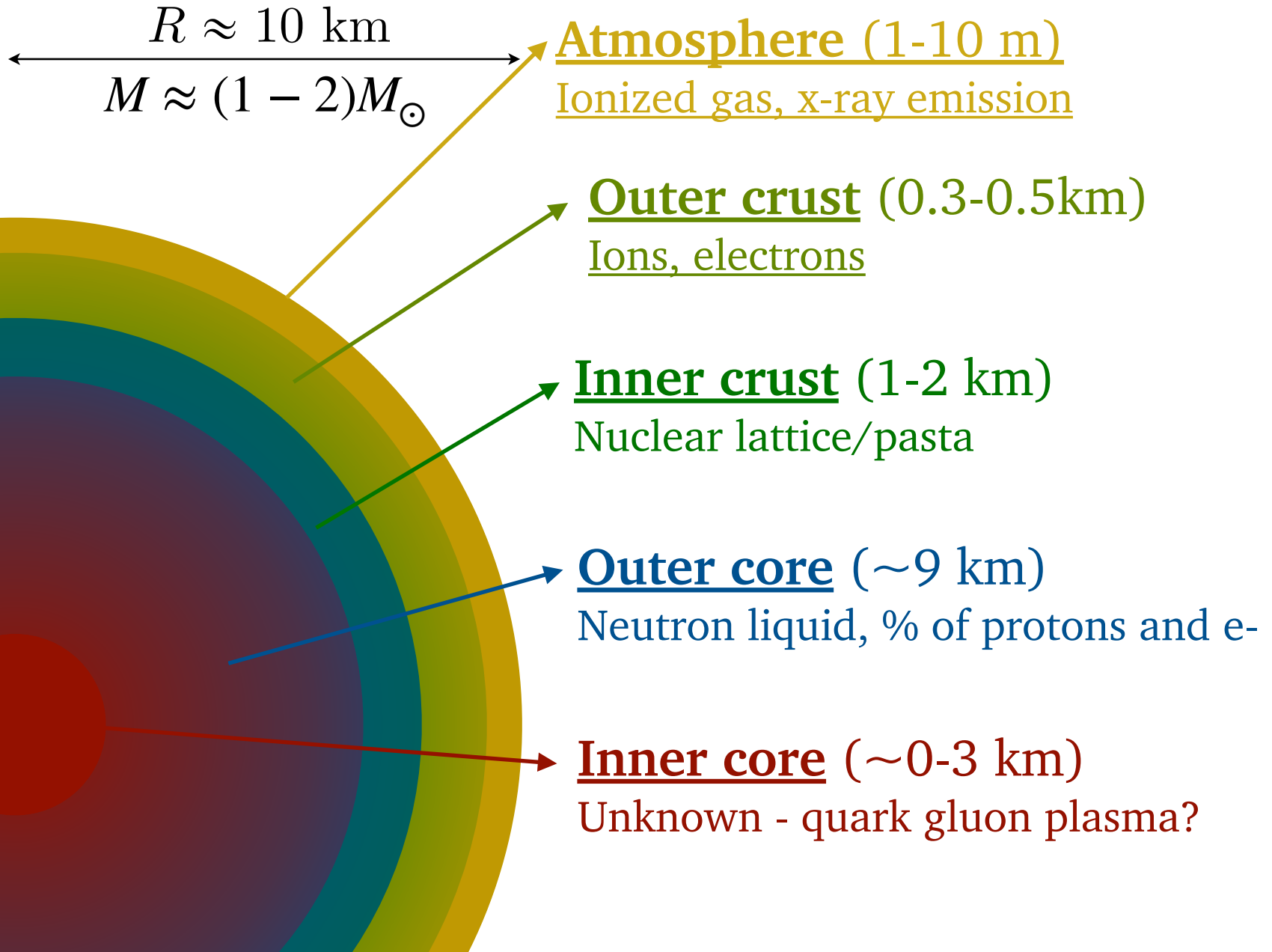
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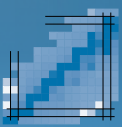


1. **Motivation**: why superfluidity in neutron stars?
2. Some **previous** results
3. **Nambu-Covariant Perturbation Theory**
4. **Nambu-Covariant Self-Consistent Green's Functions**



Stellar corpse CSI





Neutron star structure

Tolman-Oppenheimer-Volkov equations

$$\frac{dp}{dr} = -\frac{G}{c^2} \frac{(m + 4\pi pr^3)(\epsilon + p)}{r(r - 2Gm/c^2)}$$

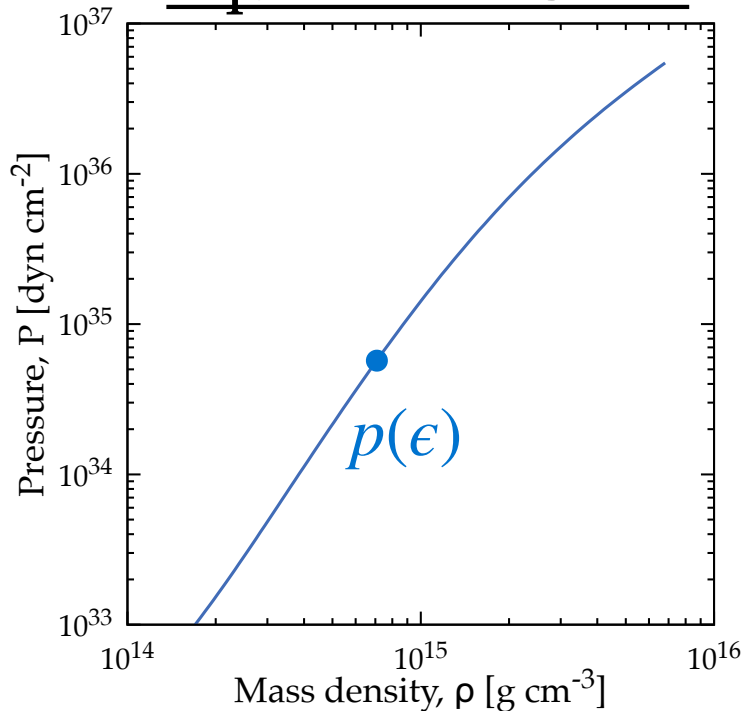
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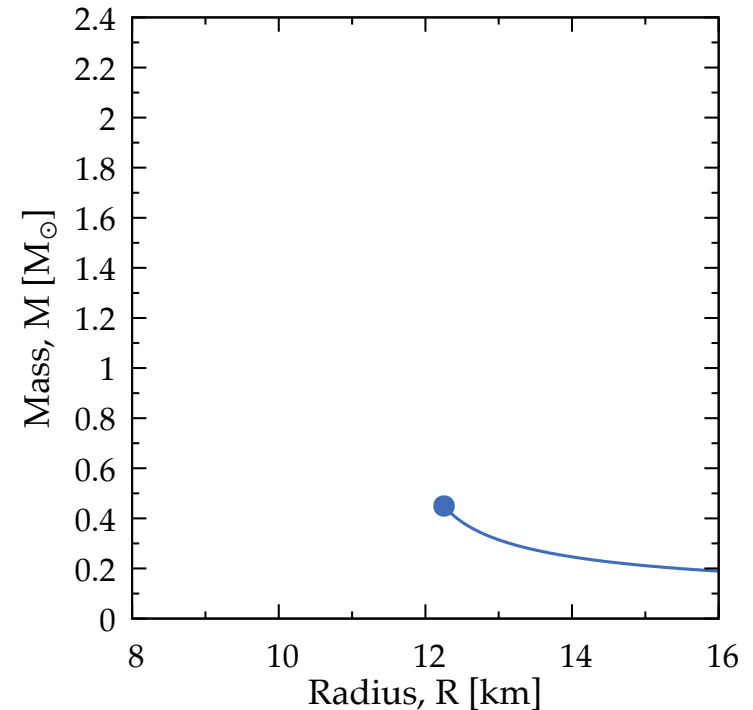
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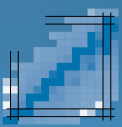
$$p \equiv p(\epsilon)$$

Equation of State



Mass-Radius





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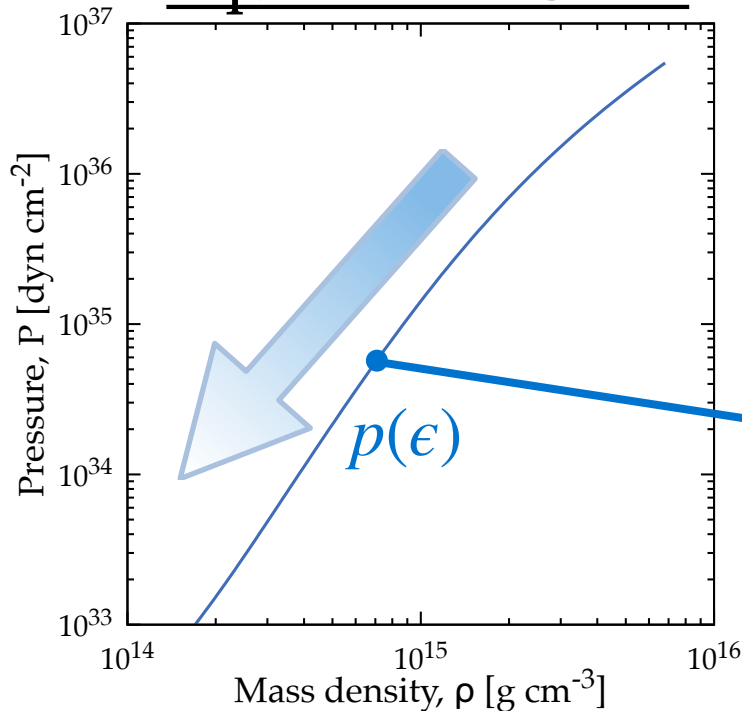
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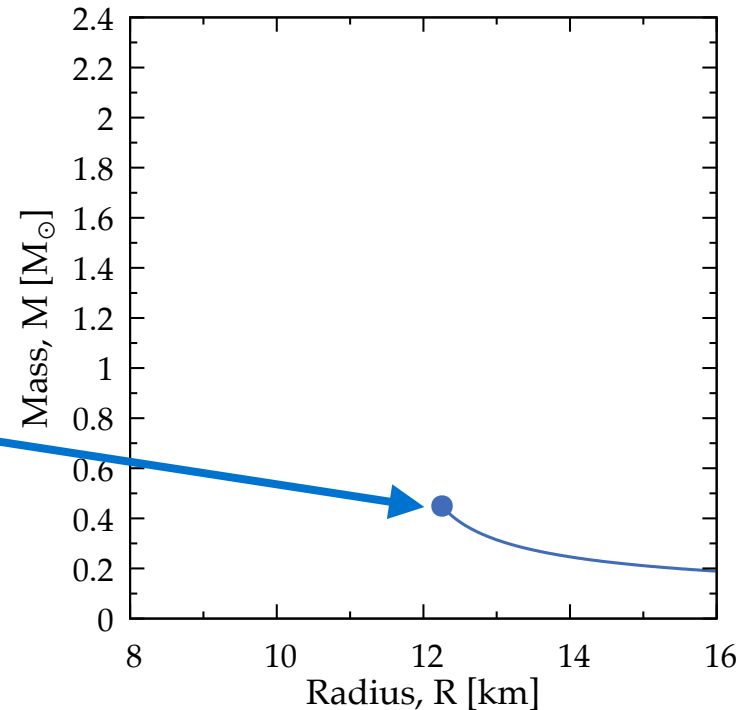
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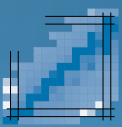
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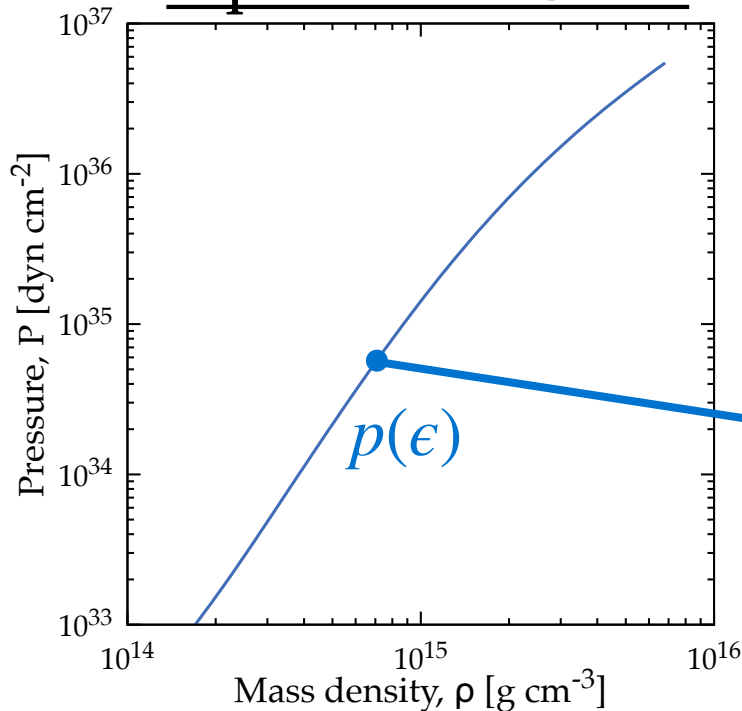
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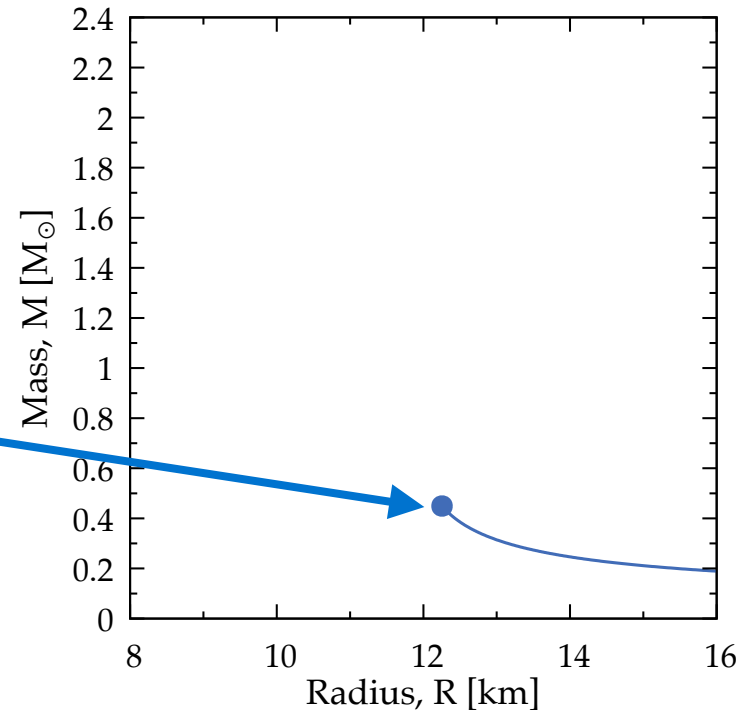
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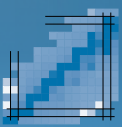
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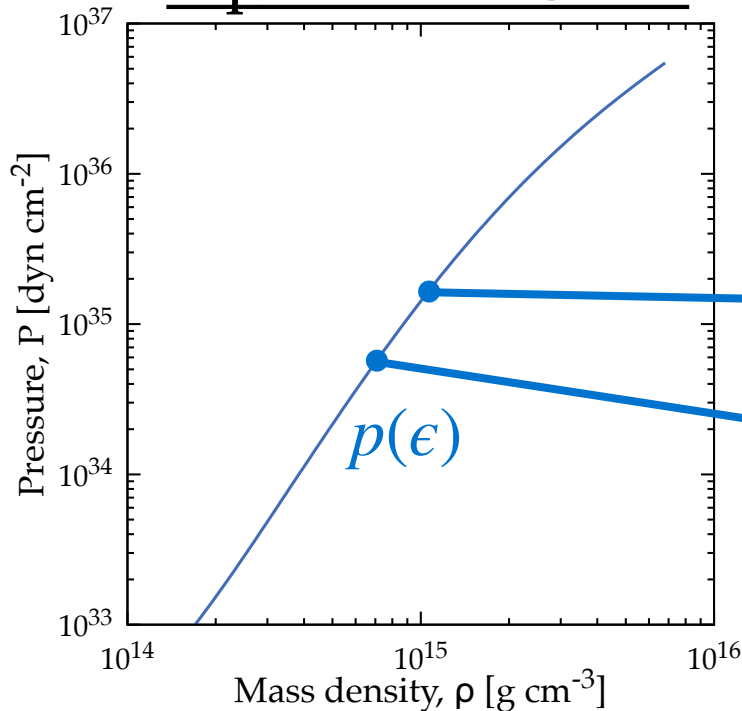
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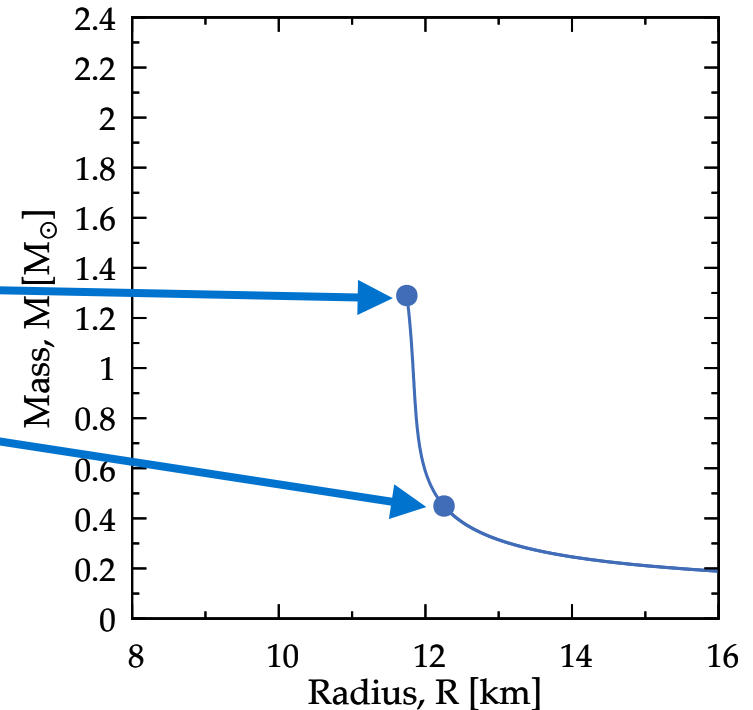
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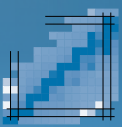
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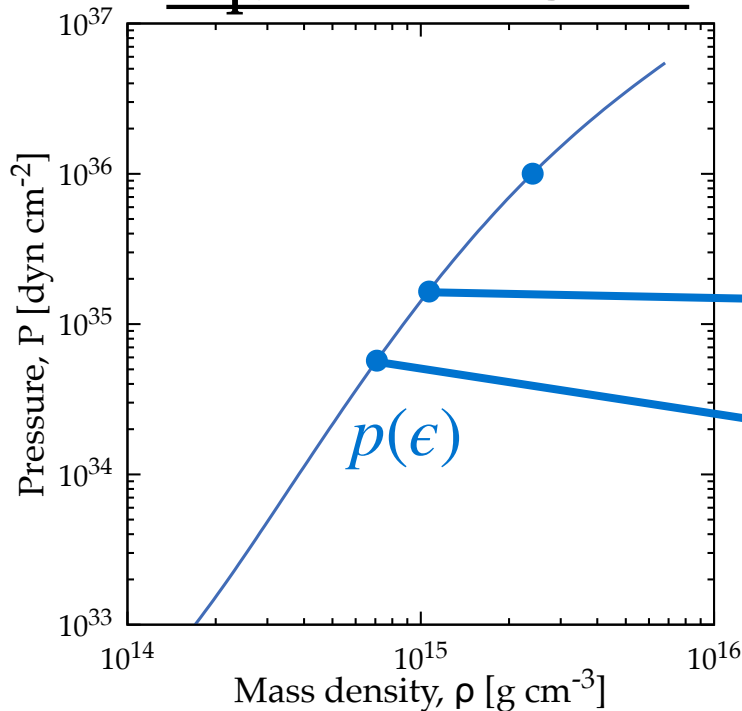
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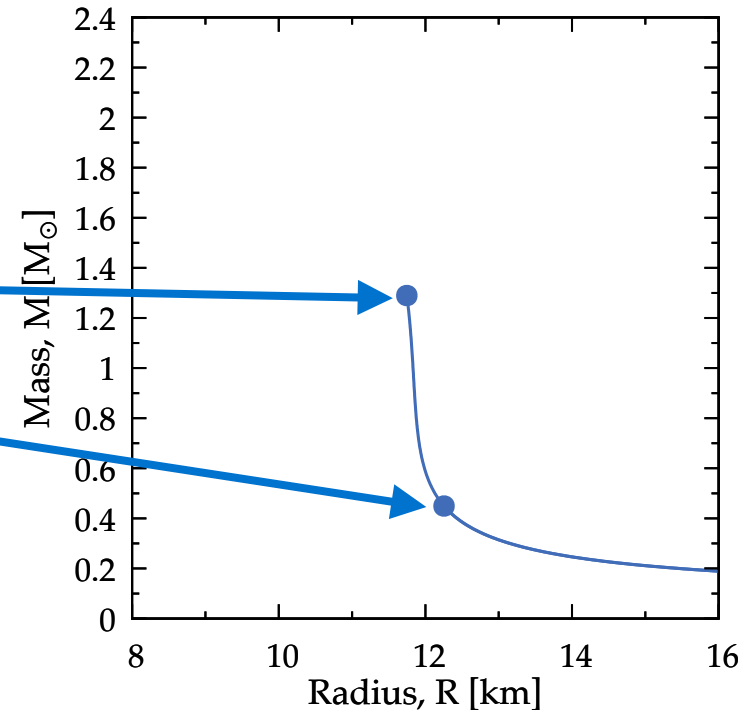
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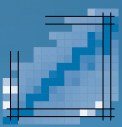
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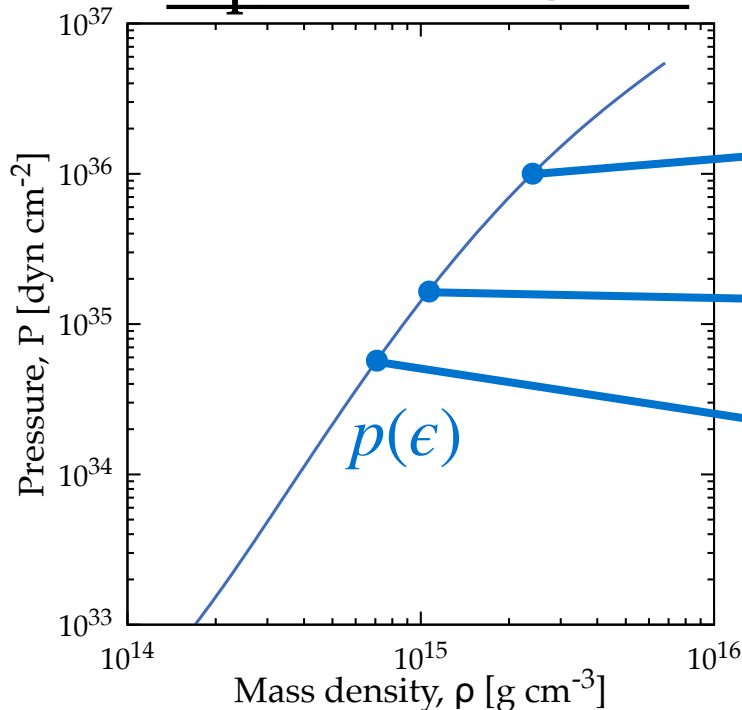
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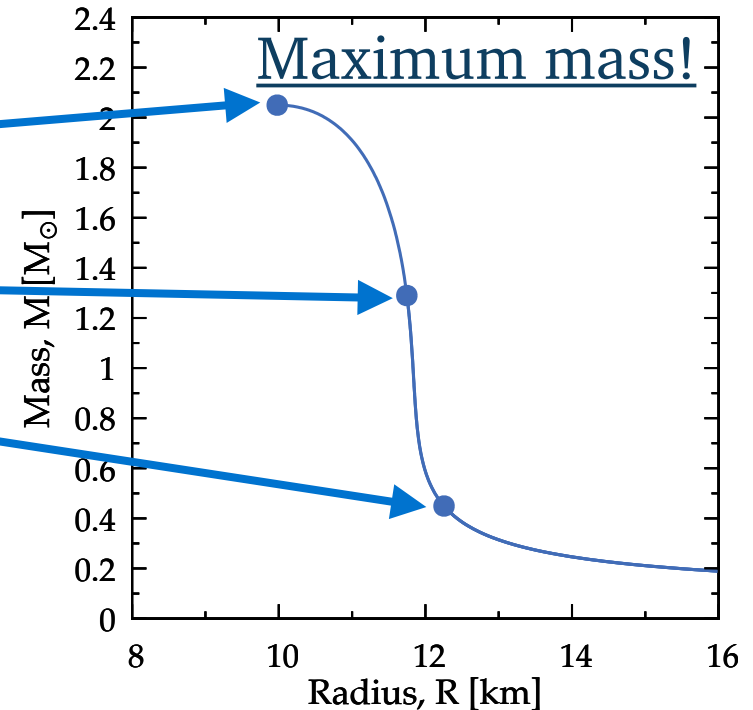
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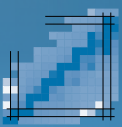
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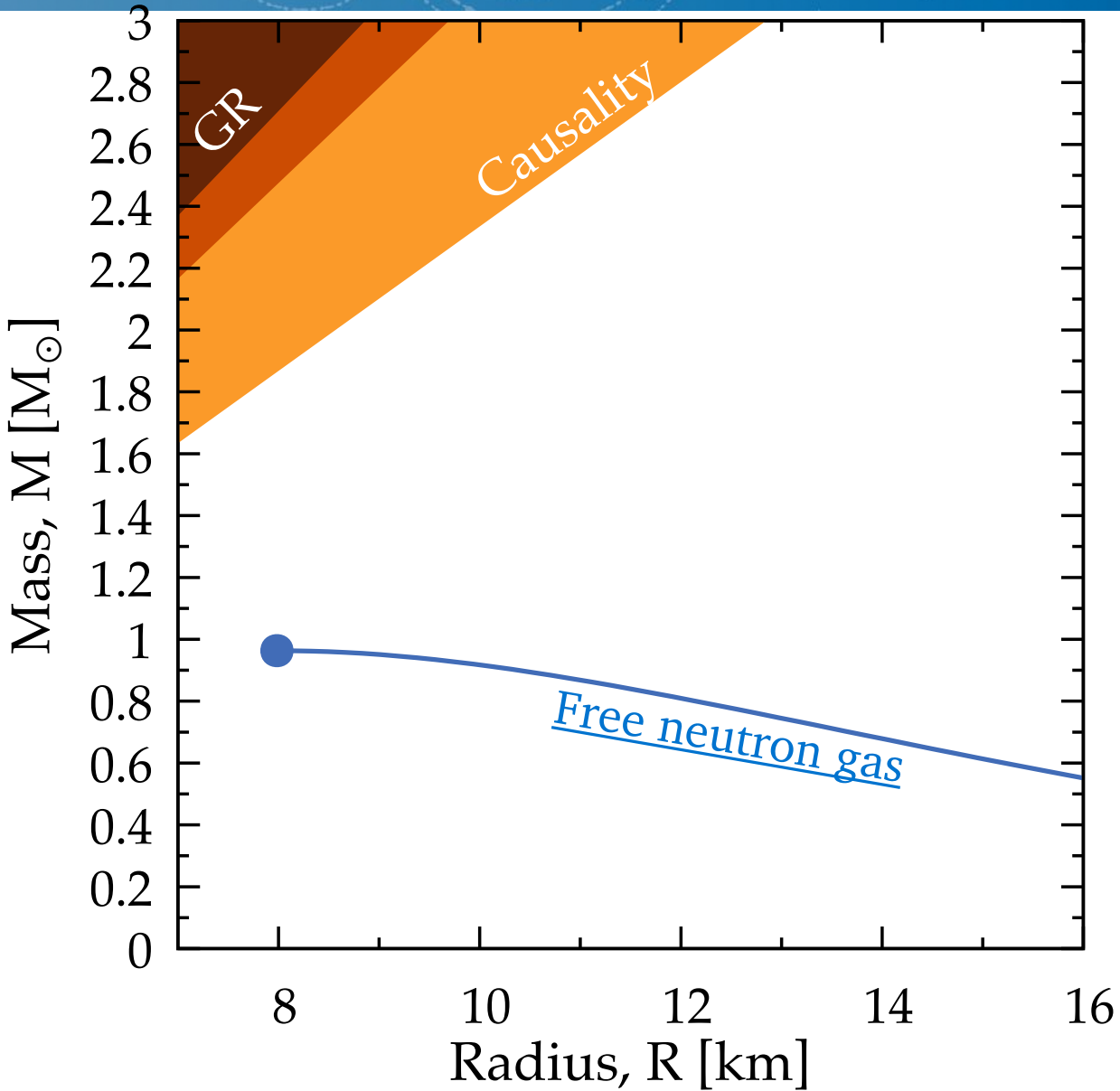


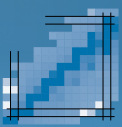
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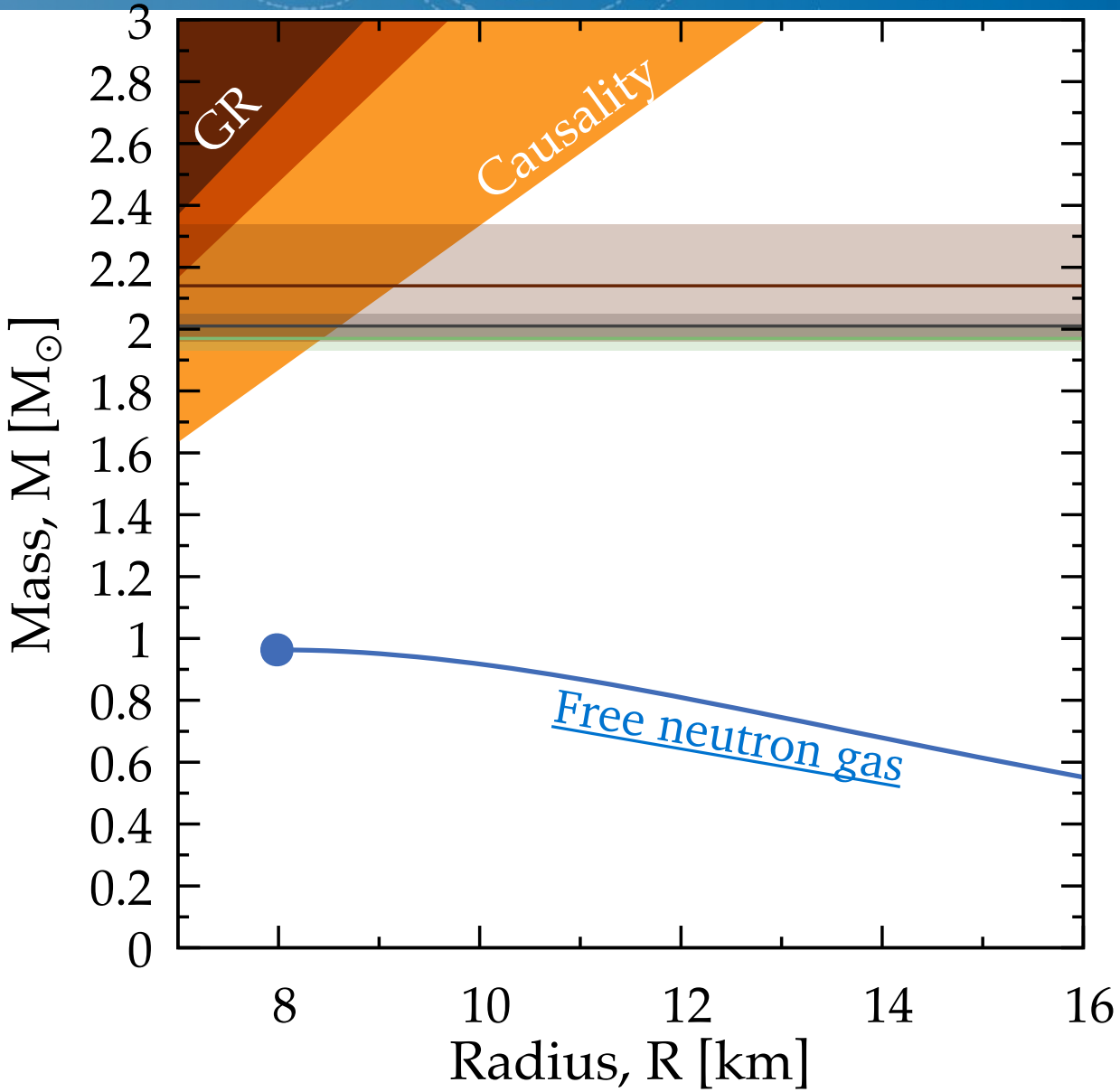


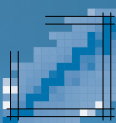
How important is our job?



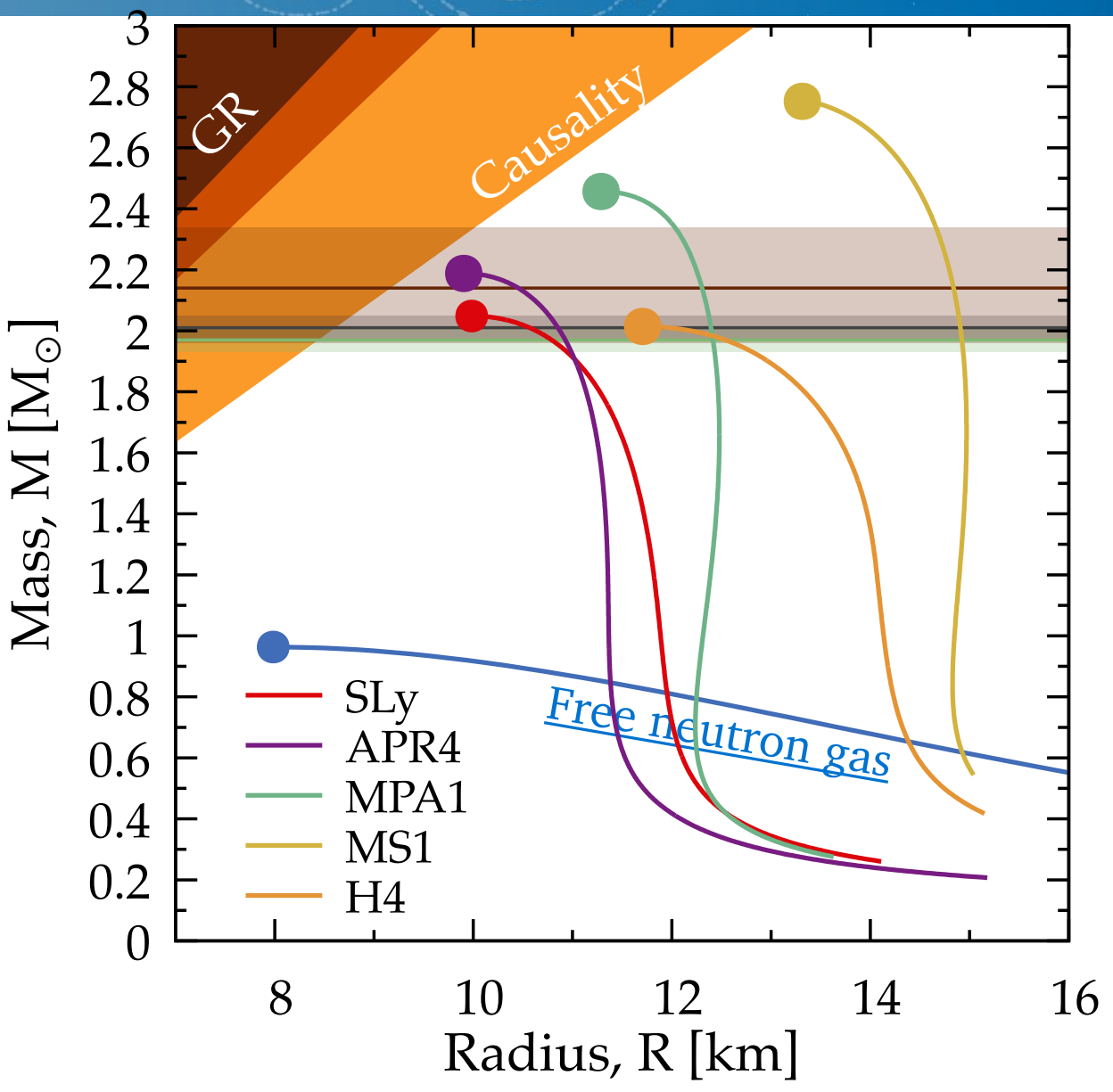


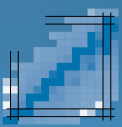
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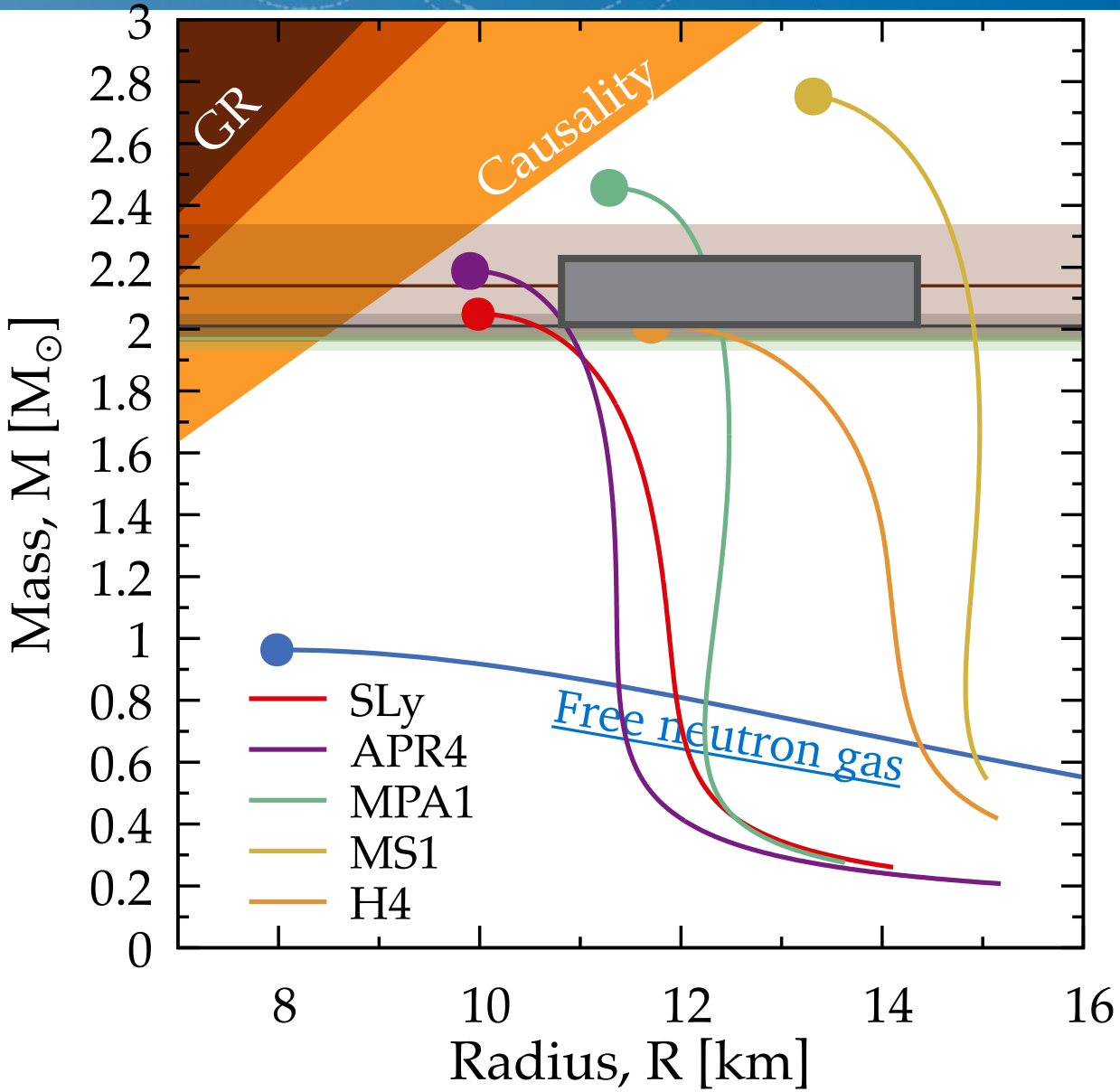


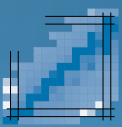
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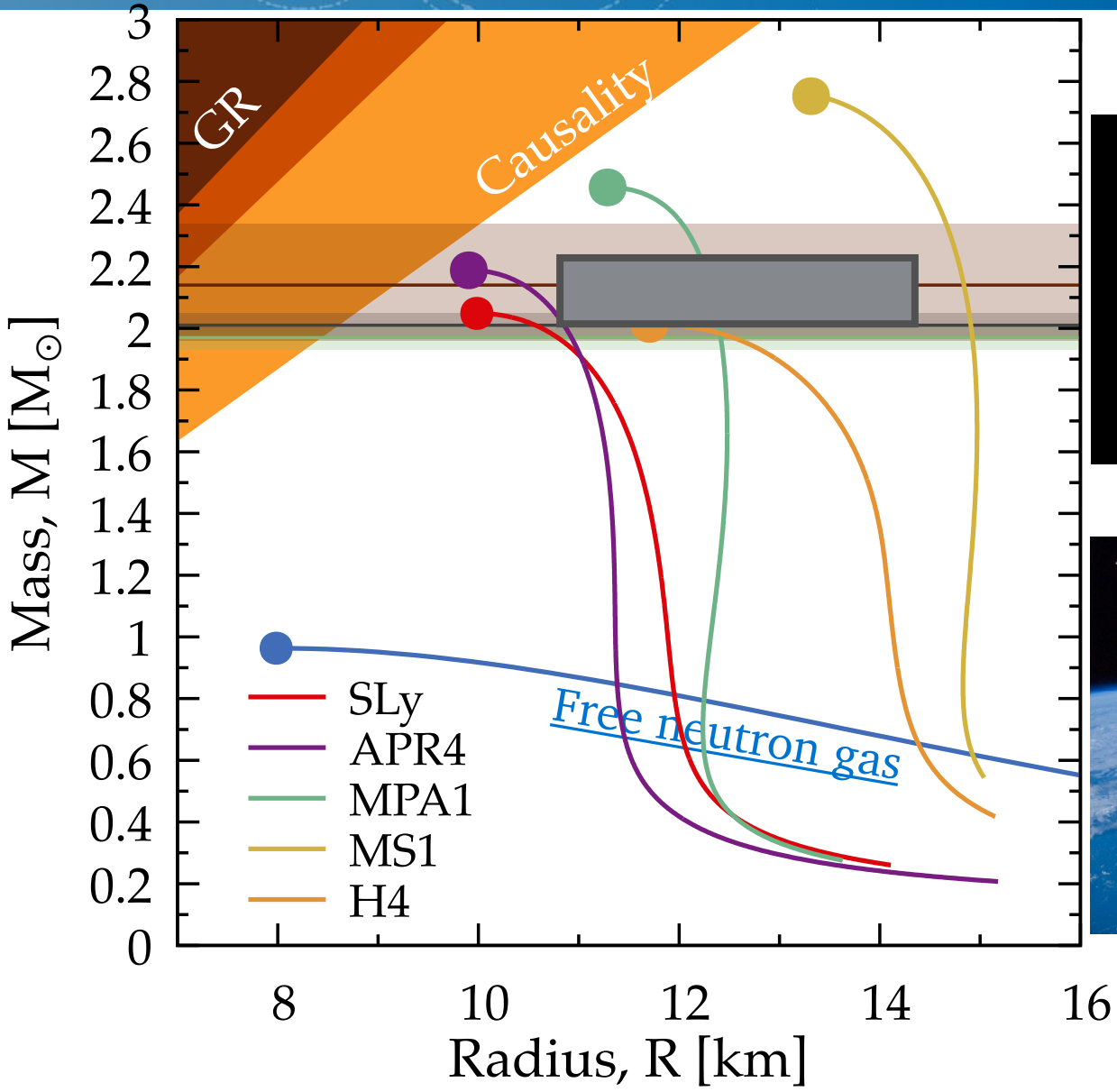


How important is our job?

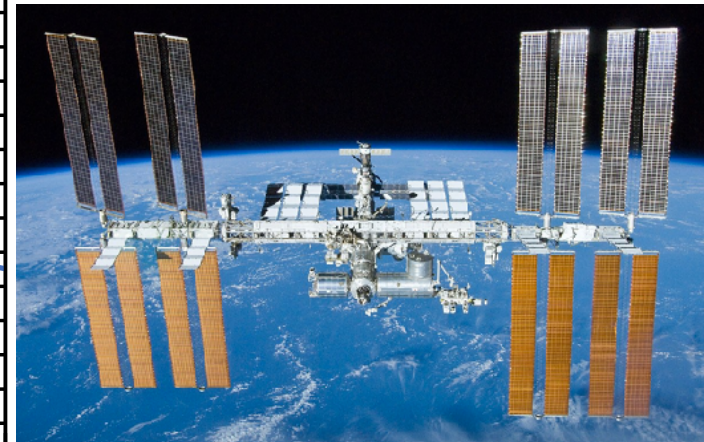
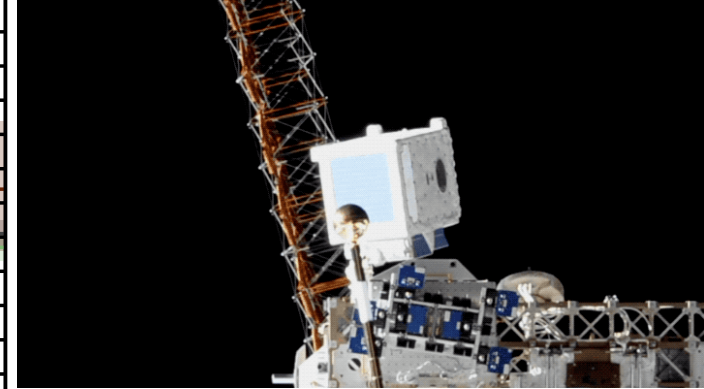




How important is our job?



NICER



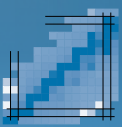
NICER Collaboration

Miller et al. arXiv:2105.06979

Riley et al. arXiv:2105.06981

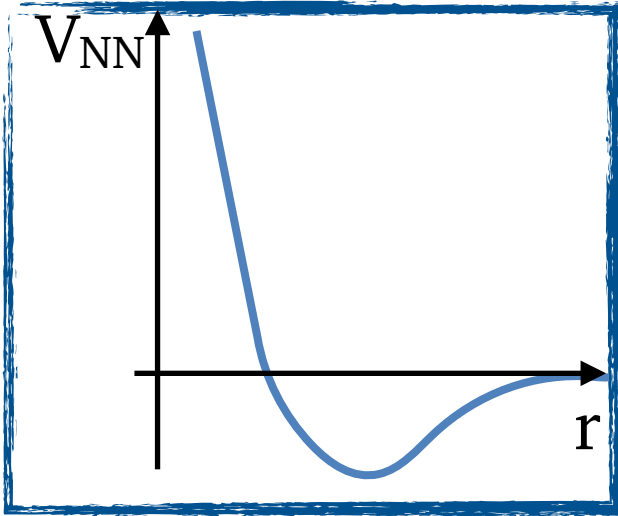
Raaijmakers et al. arXiv:2105.06981

<https://heasarc.gsfc.nasa.gov/docs/nicer/>

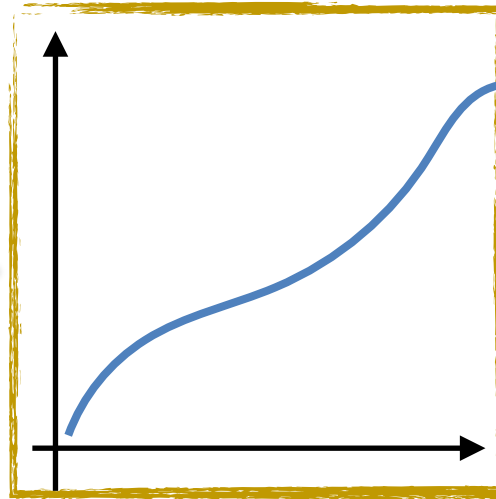


Nuclear predictions 19xx style

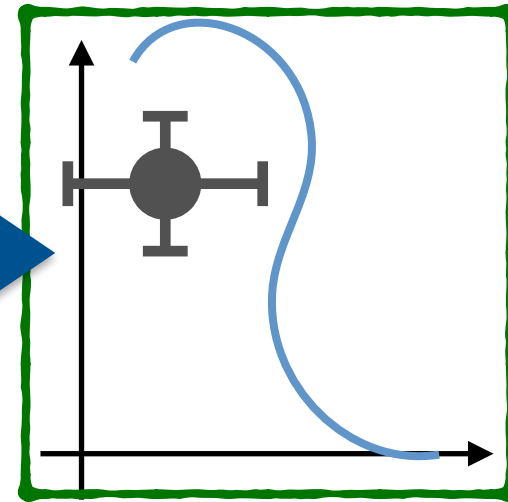
Hamiltonian



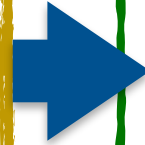
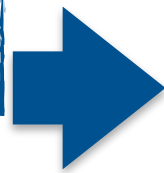
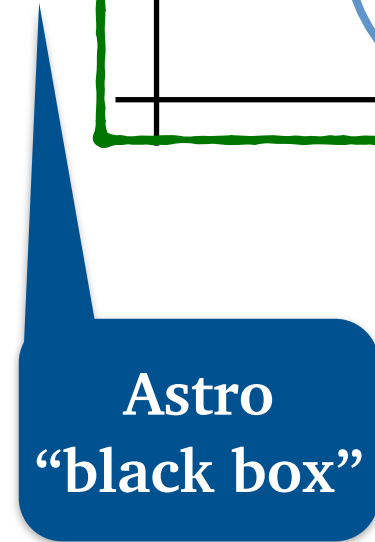
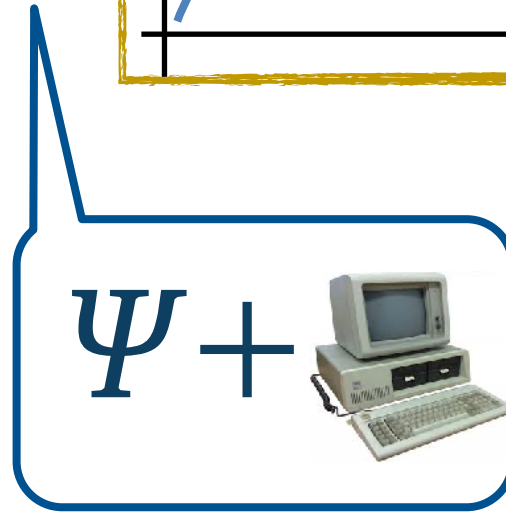
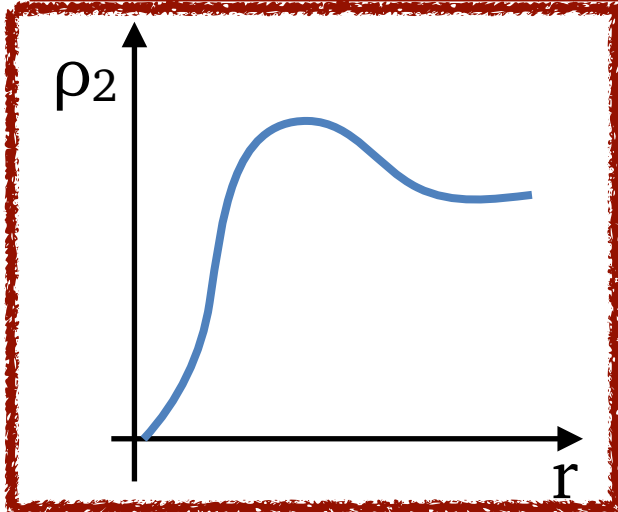
Astronuclear property

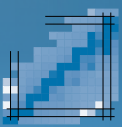


Neutron star observations



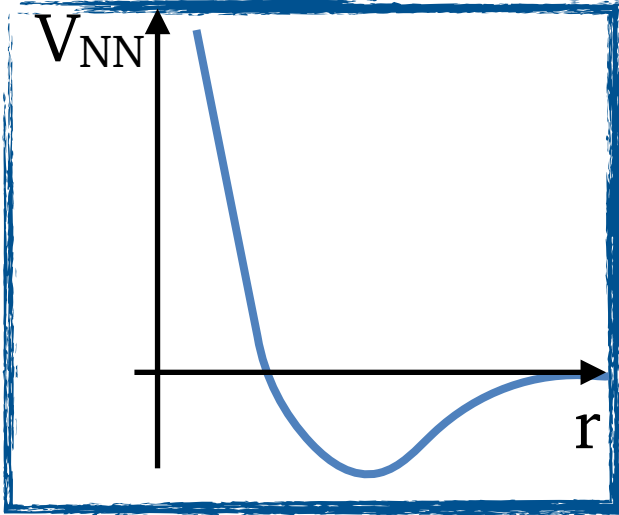
Many-body method



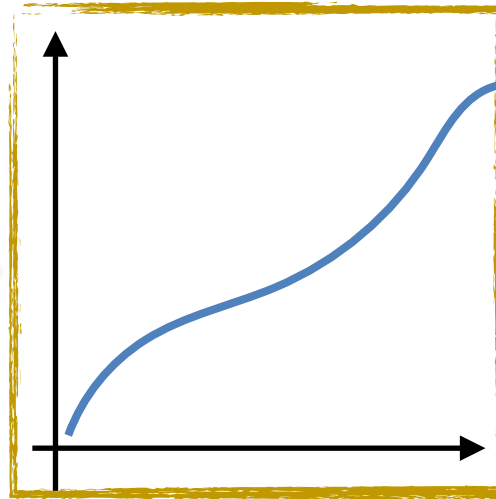


Nuclear predictions 19xx style

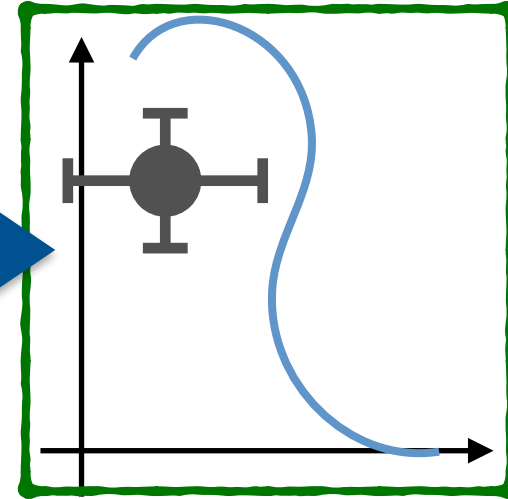
Hamiltonian



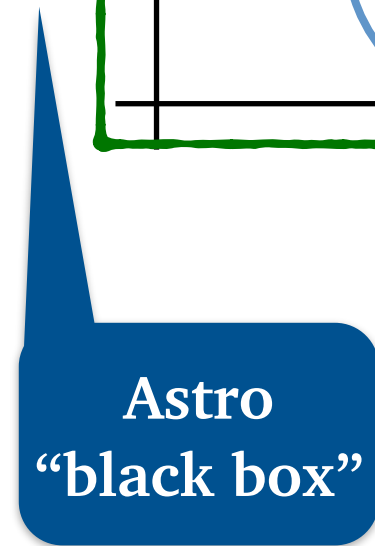
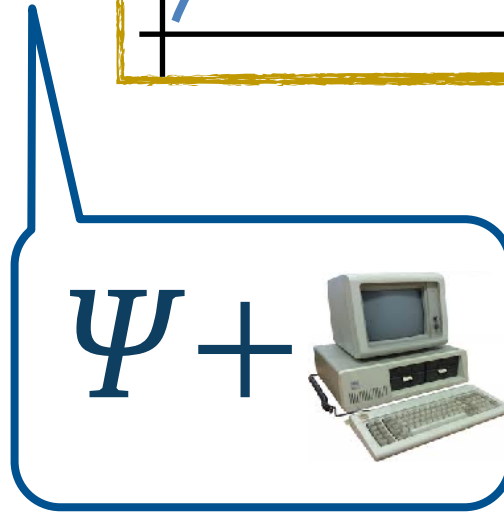
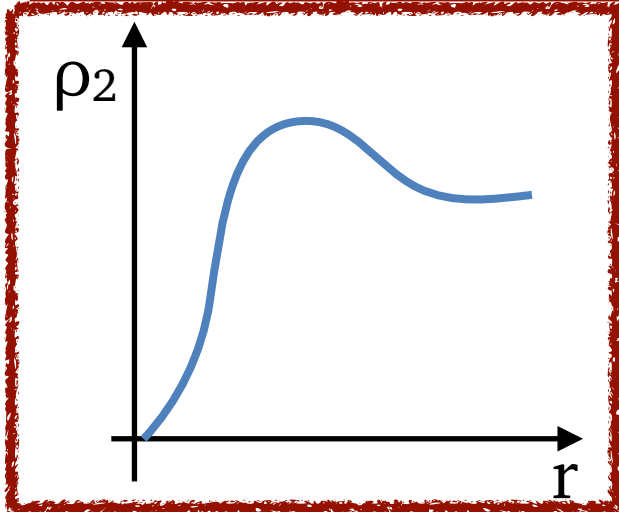
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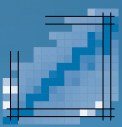


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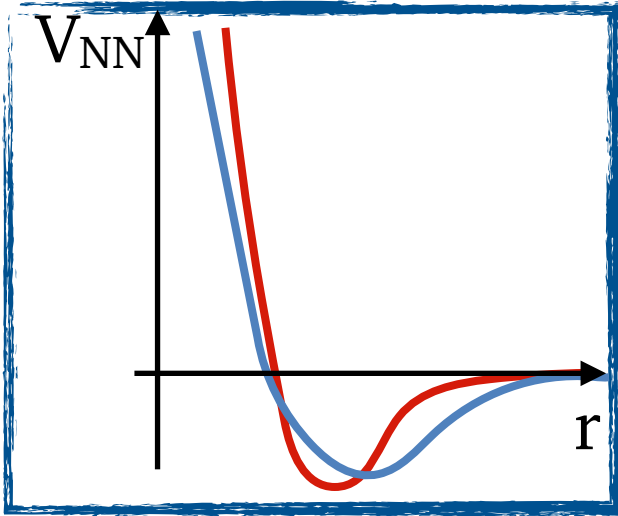
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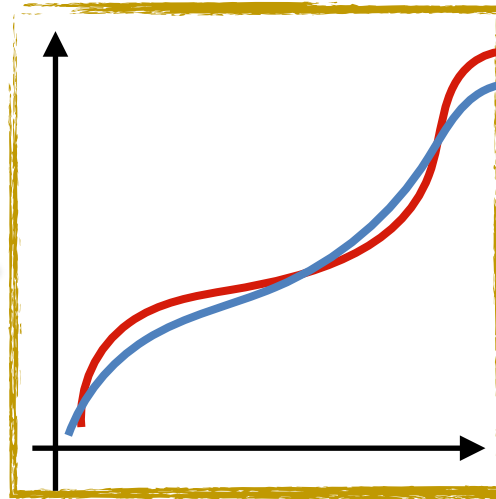


Nuclear predictions 19xx style

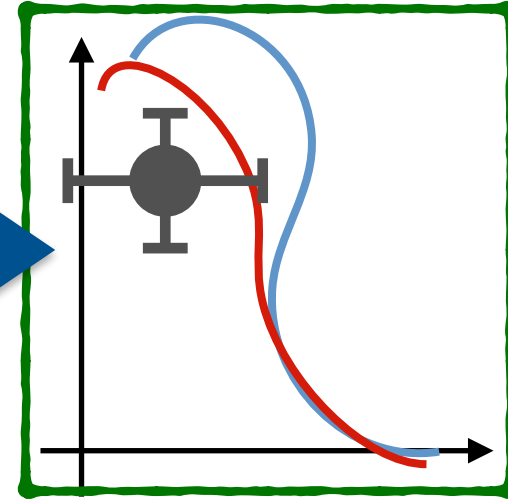
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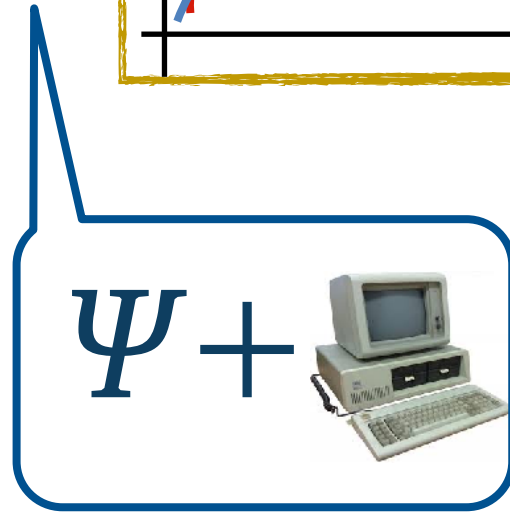
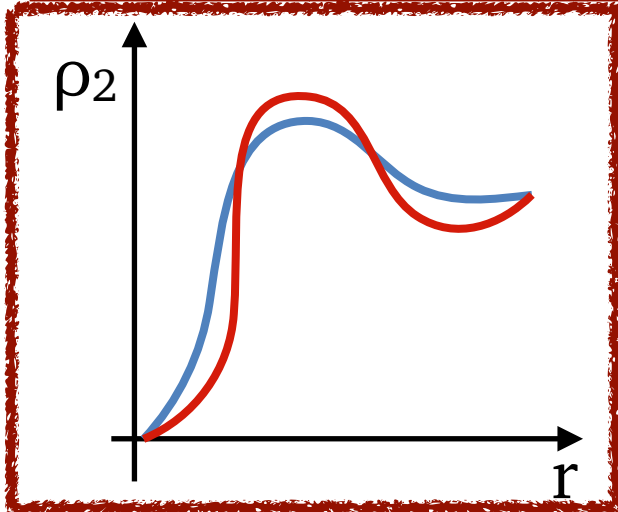
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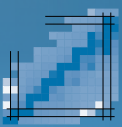
Neutron star observations



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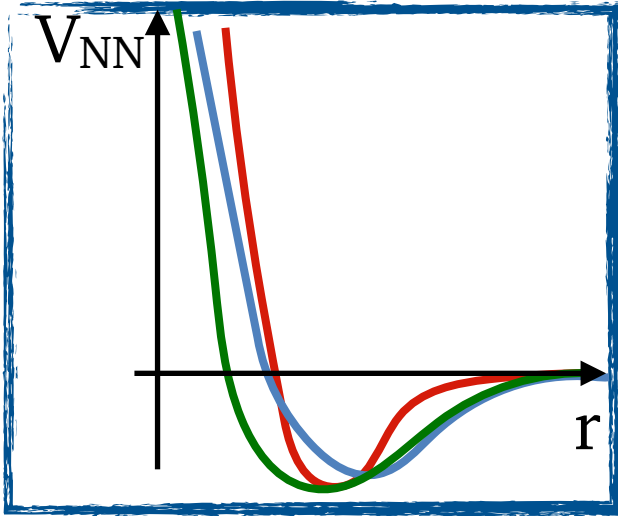


Astro
"black box"

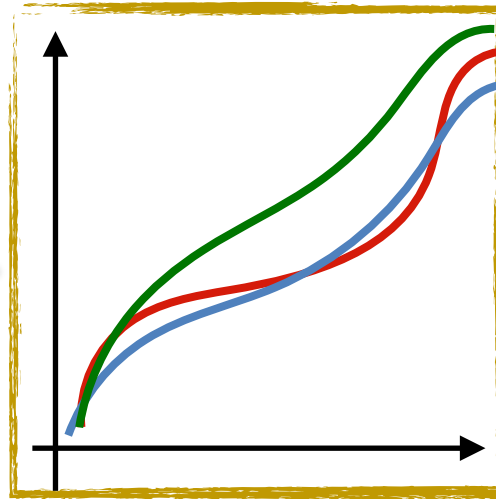


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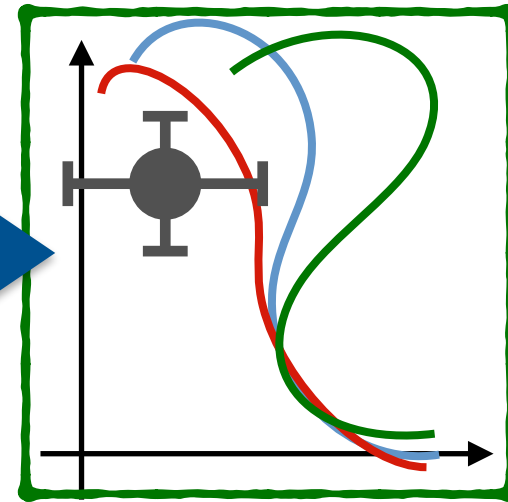
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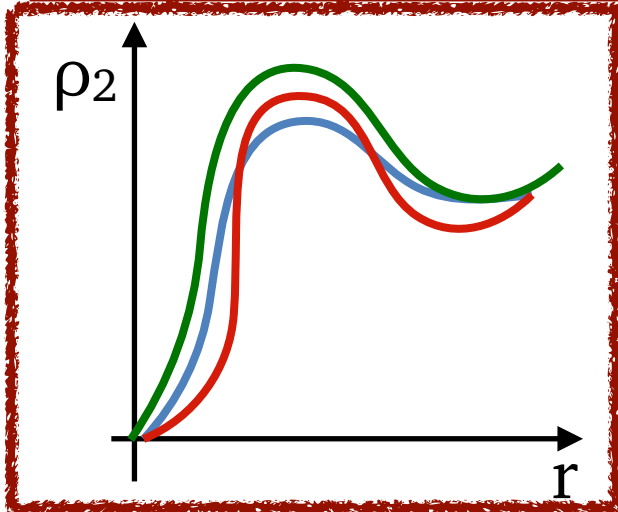
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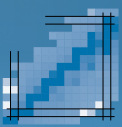


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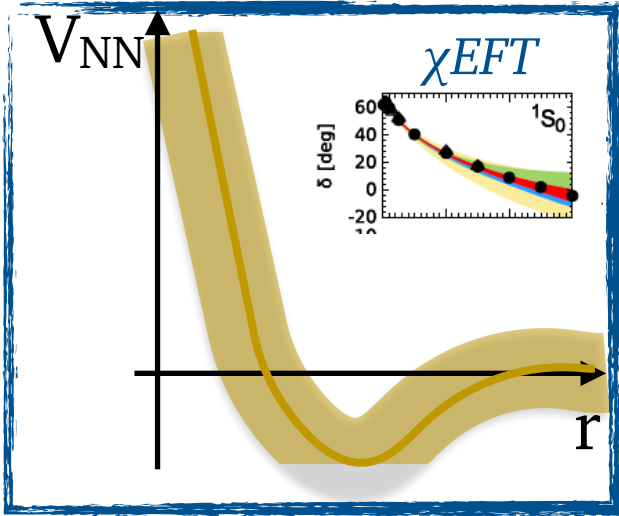
Ψ + 

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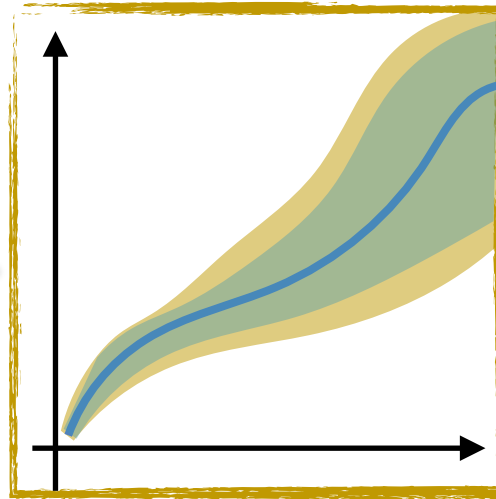


Nuclear error quantification

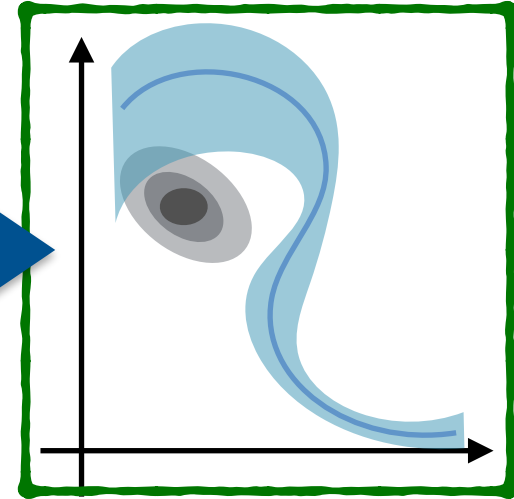
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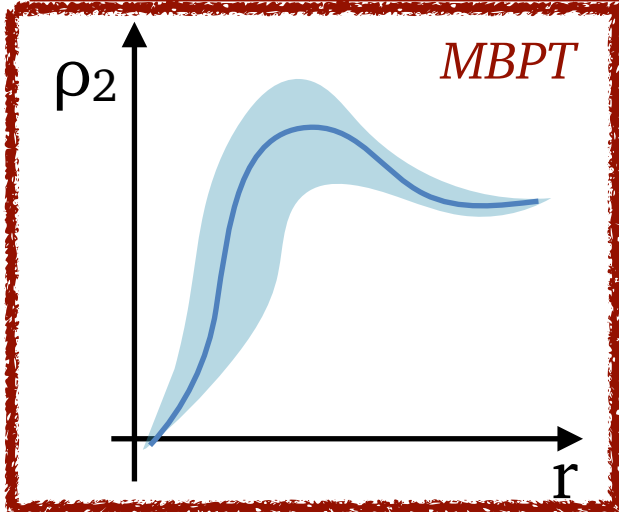
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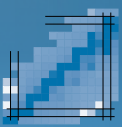


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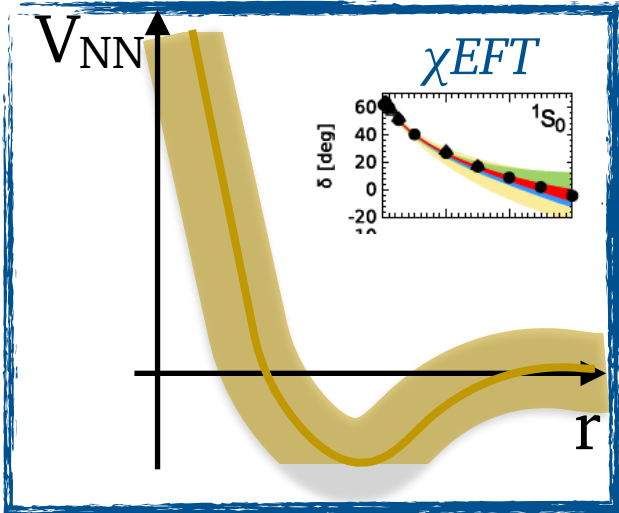
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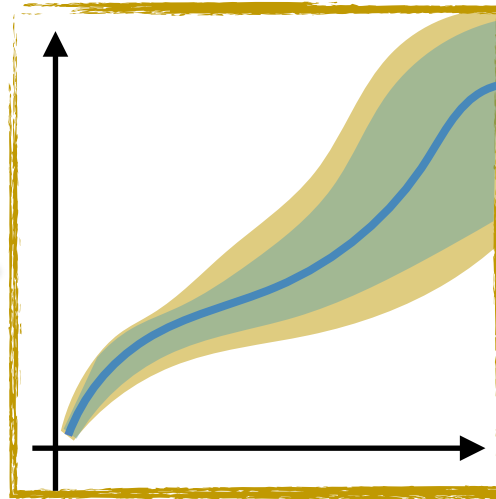


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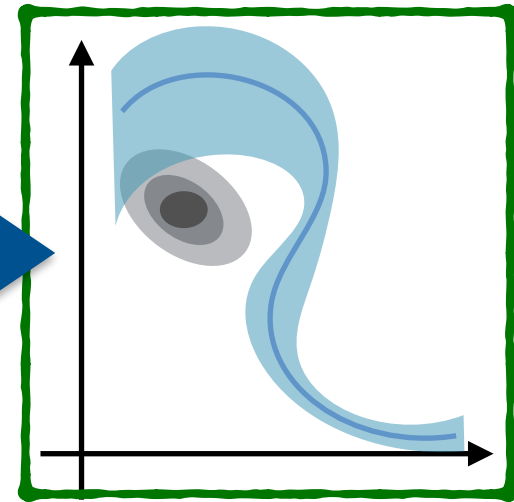
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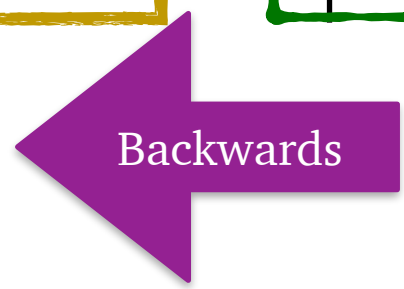
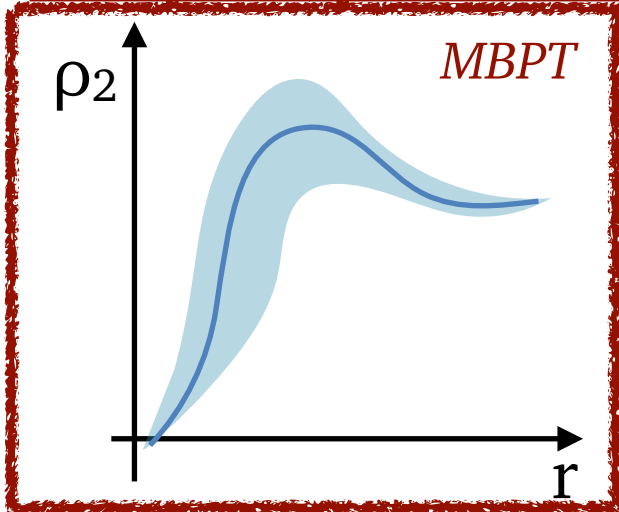
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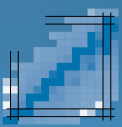
Neutron star observations



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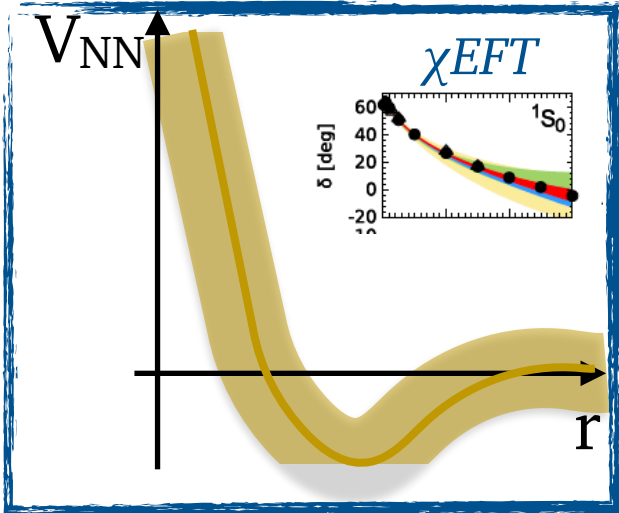


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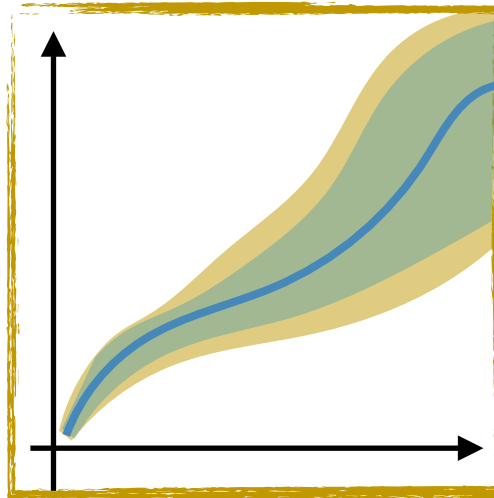


Nuclear error quantification

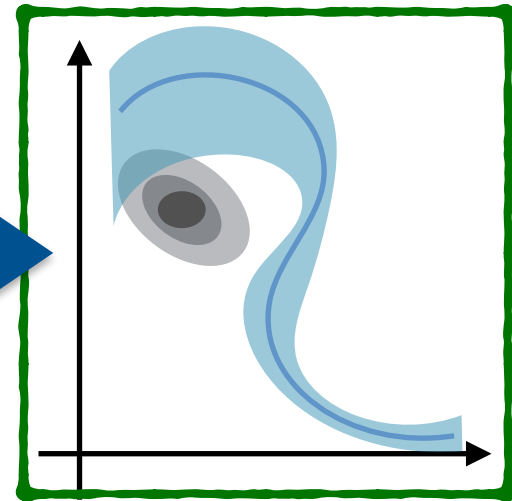
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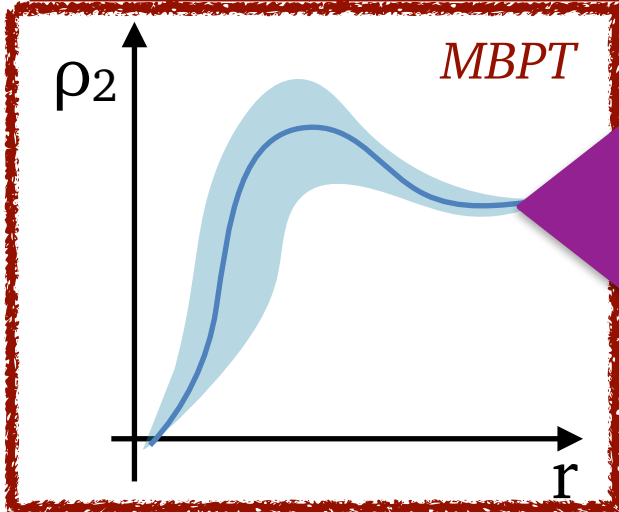
Astronuclear property



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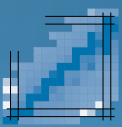


Many-body method



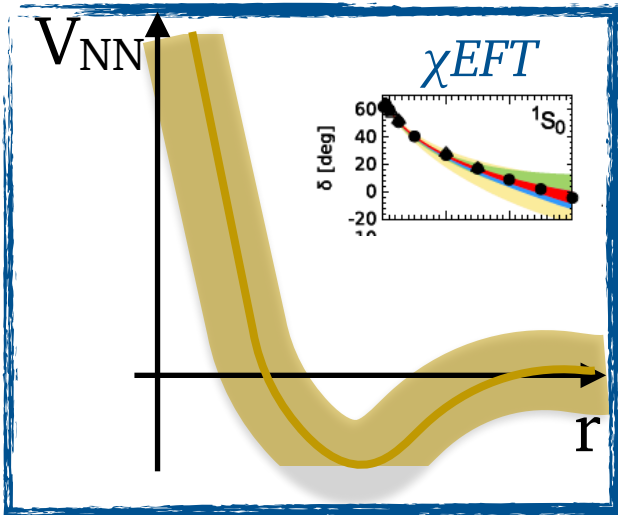
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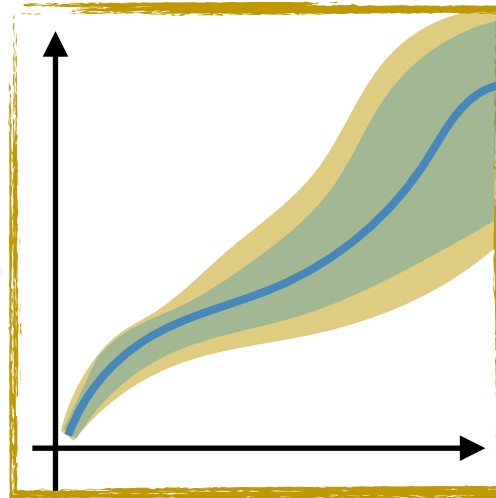


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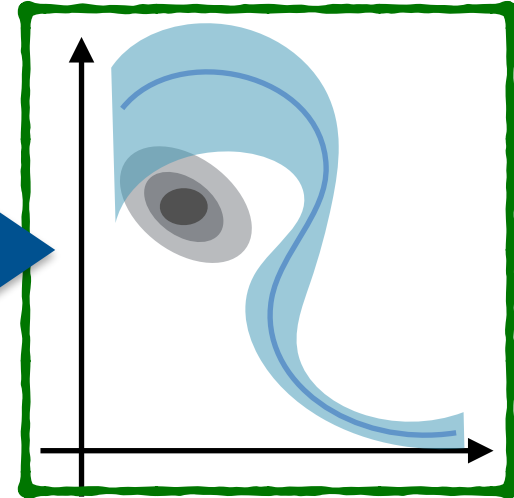
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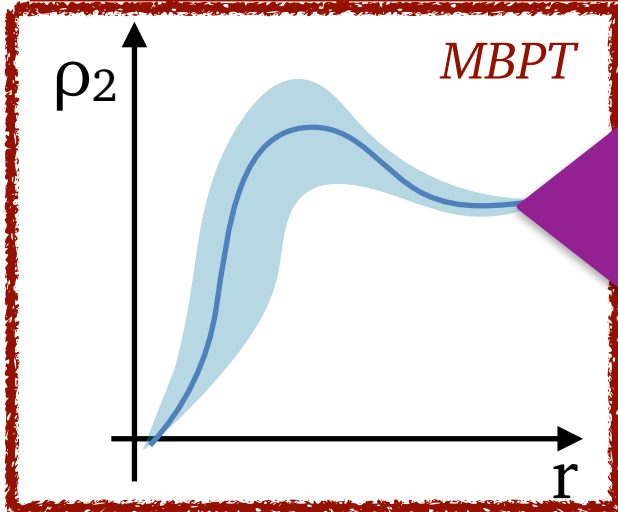
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Neutron star observations



Many-body method

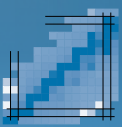


Backwards

Backwards

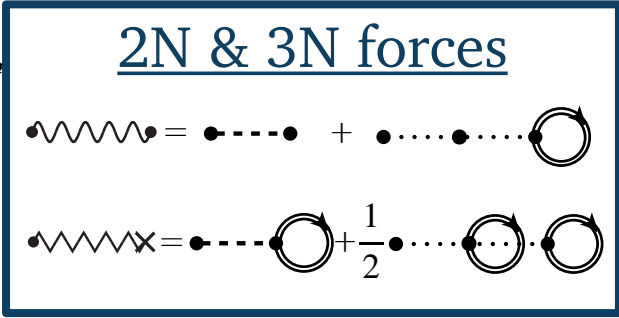
Normal phase (EoS) ✓

Superfluid phase (gaps) ✗

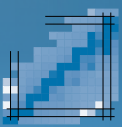


Self-Consistent Green's Functions

(ρ, T)



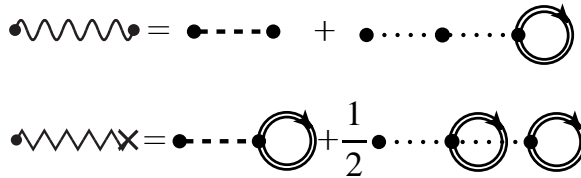
Carbone, Rios & Polls PRC 88 044302 (2013);
 PRC 90, 054322 (2014);
 Carbone PhD Thesis



Self-Consistent Green's Functions

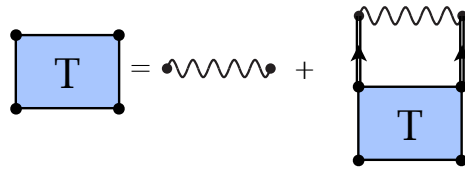
(ρ, T)

2N & 3N forces

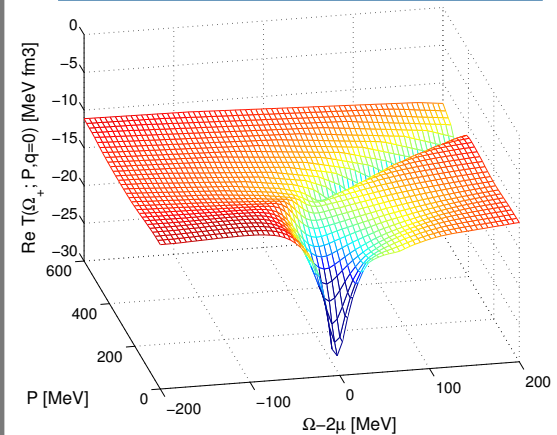


Carbone, Rios & Polls PRC **88** 044302 (2013);
 PRC **90**, 054322 (2014);
 Carbone PhD Thesis

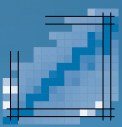
In-medium interaction



T-matrix at T=5 MeV



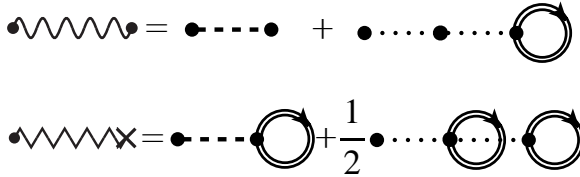
Ramos, Polls & Dickhoff, NPA **503** 1 (1989)
 Alm *et al.*, PRC **53** 2181 (1996)
 Dewulf *et al.*, PRL **90** 152501 (2003)
 Frick & Muther, PRC **68** 034310 (2003)
 Rios, PhD Thesis, U. Barcelona (2007)
 Soma & Bozek, PRC **78** 054003 (2008)
 Rios & Soma PRL **108** 012501 (2012)



Self-Consistent Green's Functions

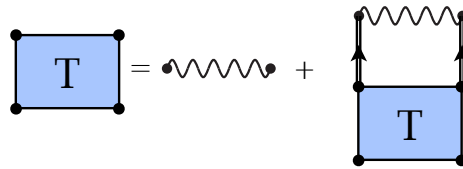
(ρ, T)

2N & 3N forces

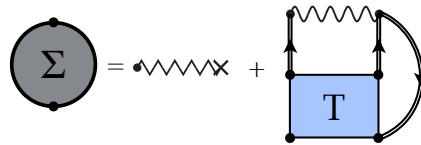


Carbone, Rios & Polls PRC **88** 044302 (2013);
 PRC **90**, 054322 (2014);
 Carbone PhD Thesis

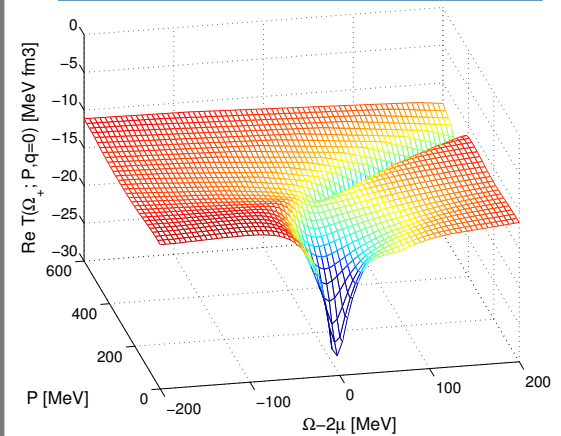
In-medium interaction



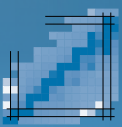
Self-energy



T-matrix at T=5 MeV



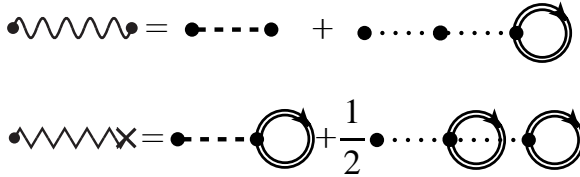
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 Rios & Soma PRL **108** 012501 (2012)



Self-Consistent Green's Functions

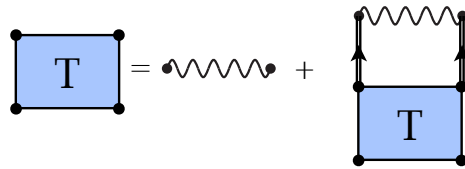
(ρ, T)

2N & 3N forces

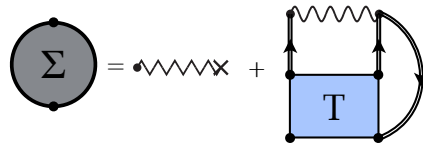


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 Carbone PhD Thesis

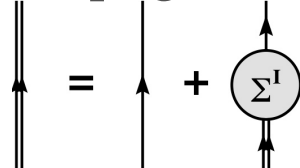
In-medium interaction



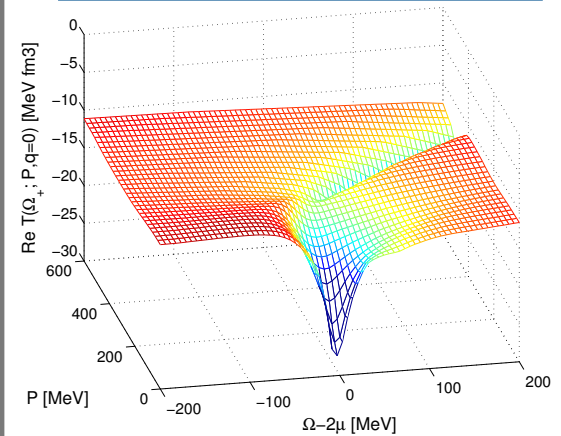
Self-energy



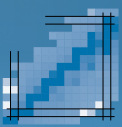
Propagator



T-matrix at T=5 MeV



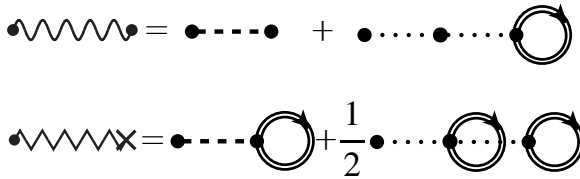
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Self-Consistent Green's Functions

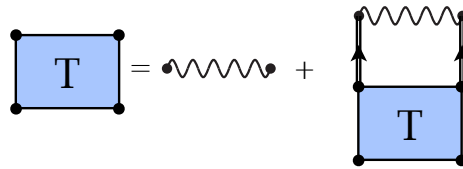
(ρ, T)

2N & 3N forces

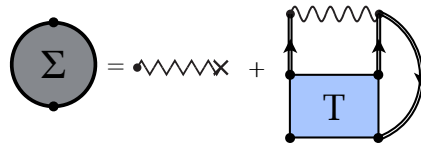


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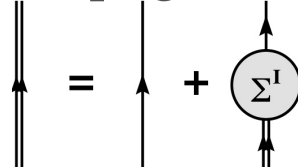
In-medium interaction



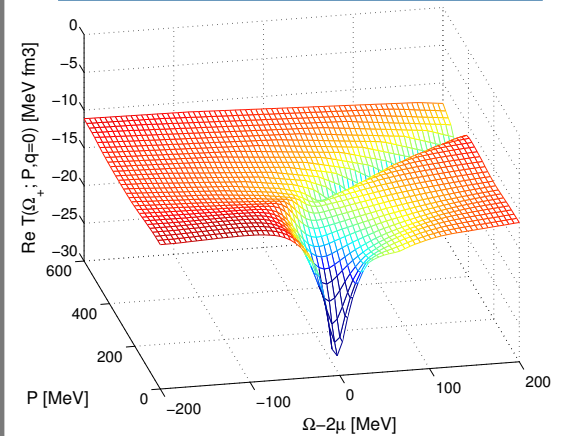
Self-energy



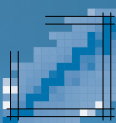
Propagator



T-matrix at T=5 MeV



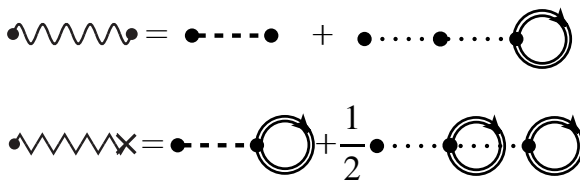
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 Rios & Soma PRL **108** 012501 (2012)



Self-Consistent Green's Functions

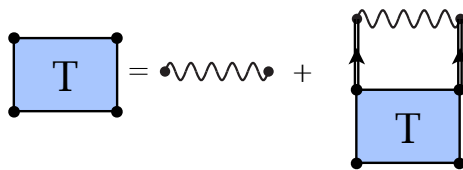
(ρ, T)

2N & 3N forces

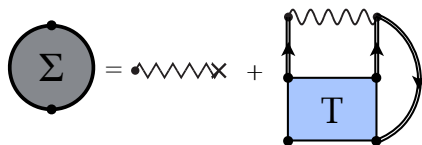


Carbone, Rios & Polls PRC **88** 044302 (2013);
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 Carbone PhD Thesis

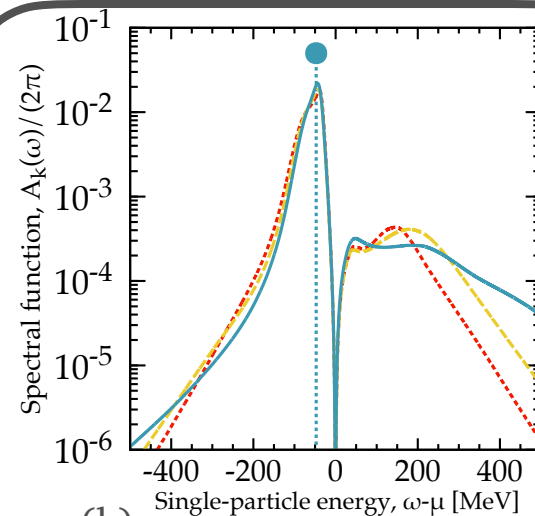
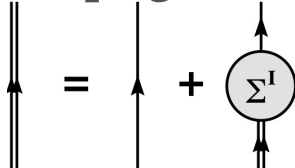
In-medium interaction



Self-energy



Propagator



$n(k)$

Thermodynamics & EoS

Transport

Ramos, Polls & Dickhoff, NPA **503** 1 (1989)

Alm *et al.*, PRC **53** 2181 (1996)

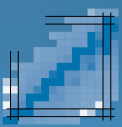
Dewulf *et al.*, PRL **90** 152501 (2003)

Frick & Muther, PRC **68** 034310 (2003)

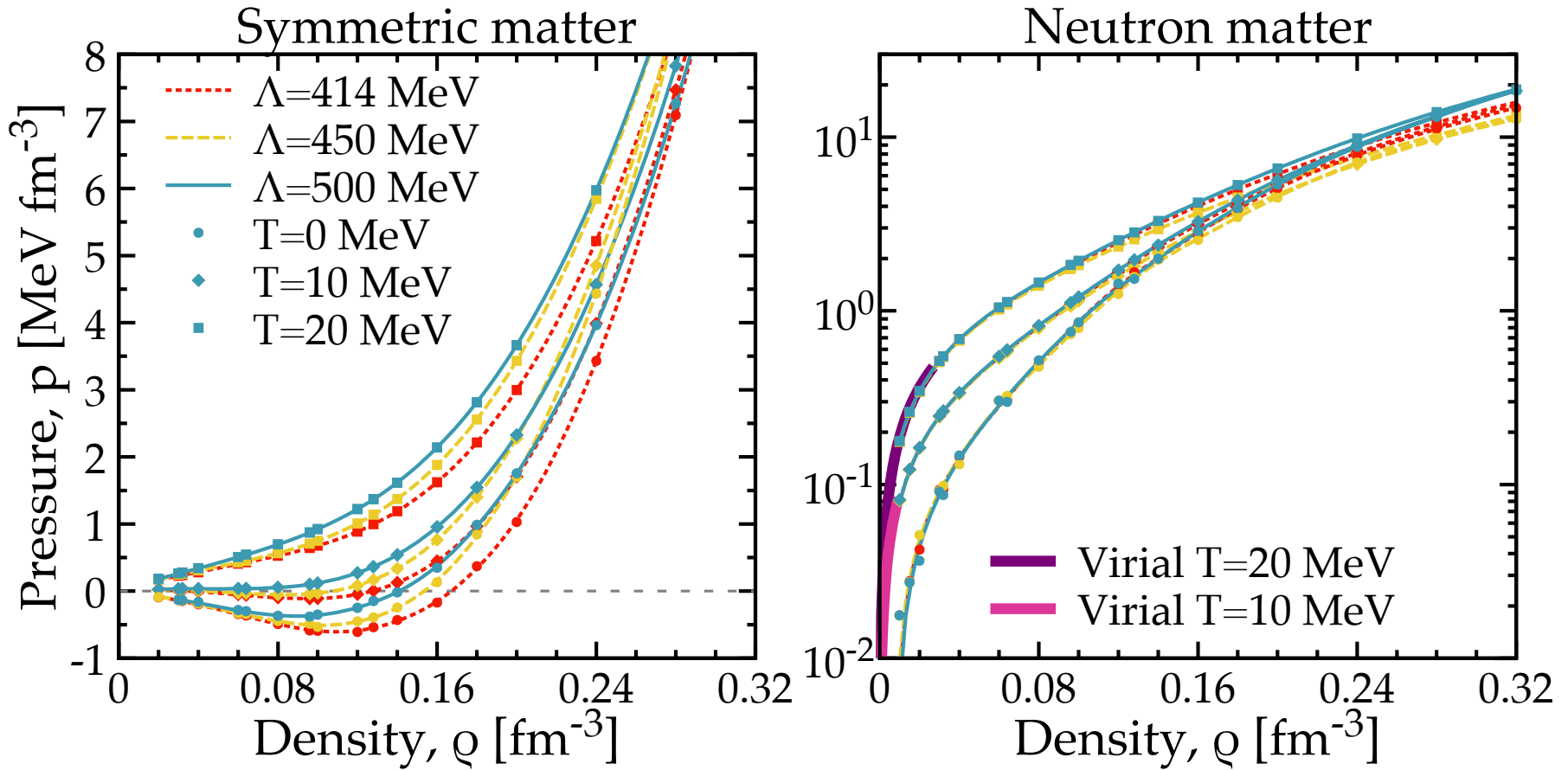
Rios, PhD Thesis, U. Barcelona (2007)

Soma & Bozek, PRC **78** 054003 (2008)

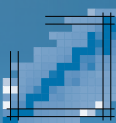
Rios & Soma PRL **108** 012501 (2012)



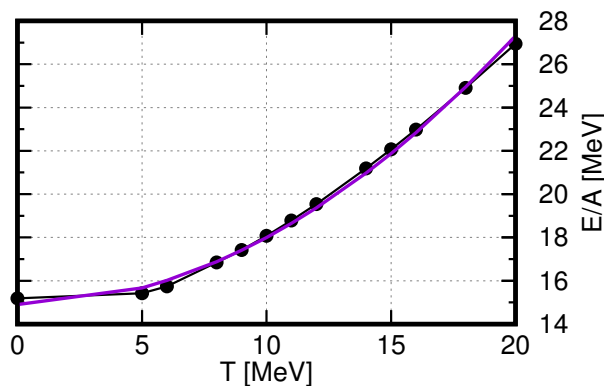
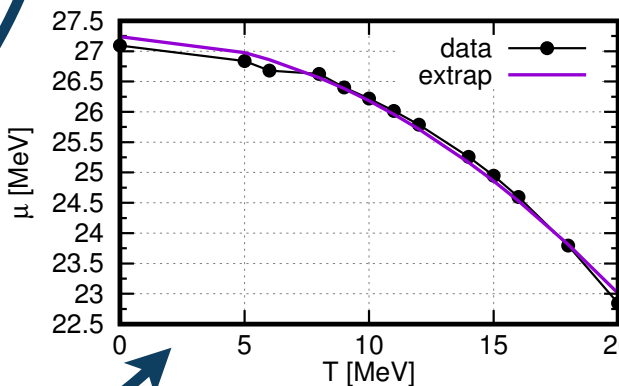
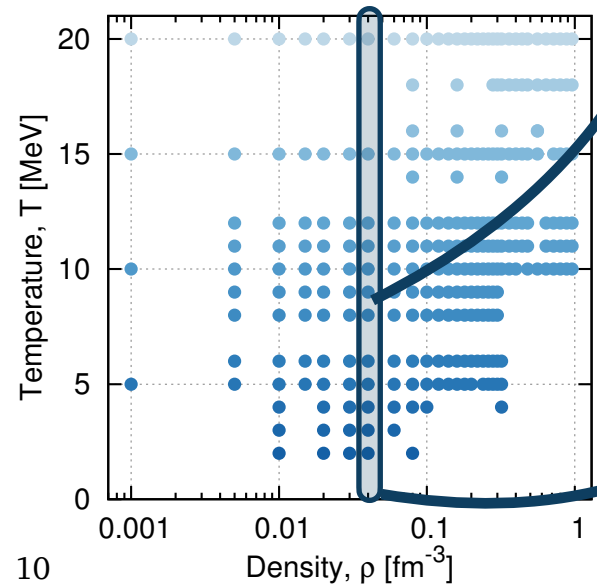
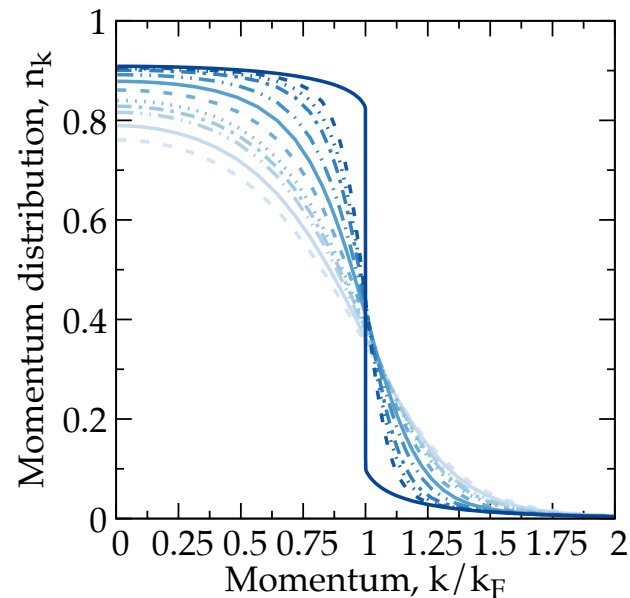
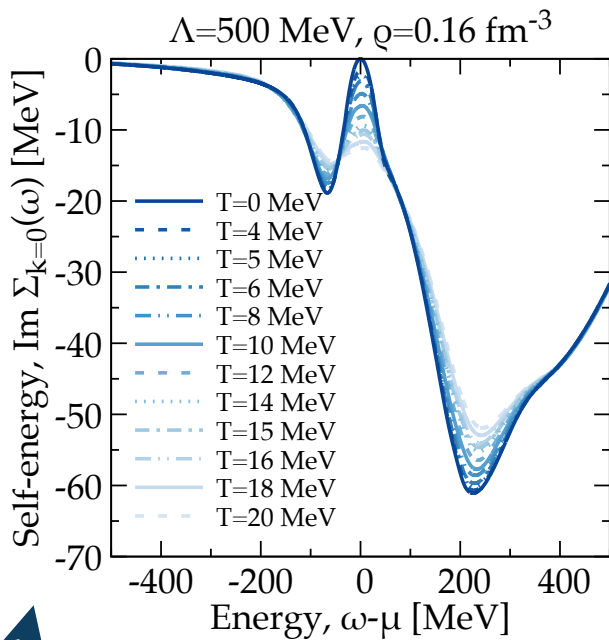
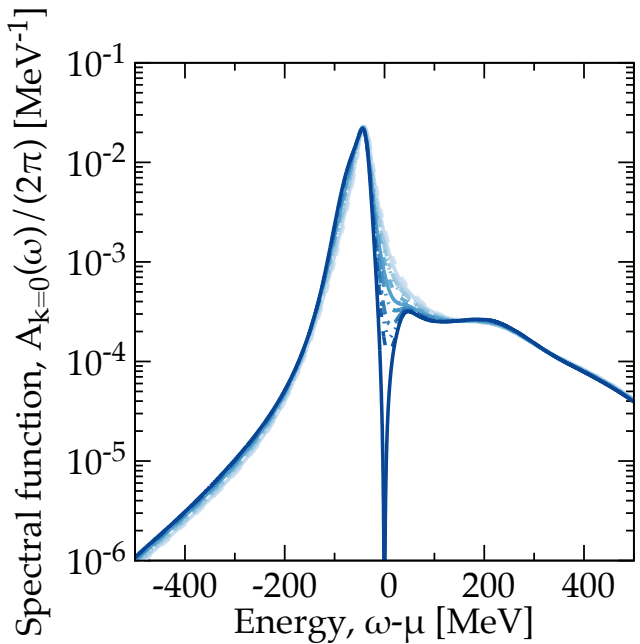
EoS at finite temperature

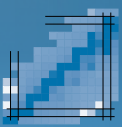


- Relevant & **necessary** for astro simulations

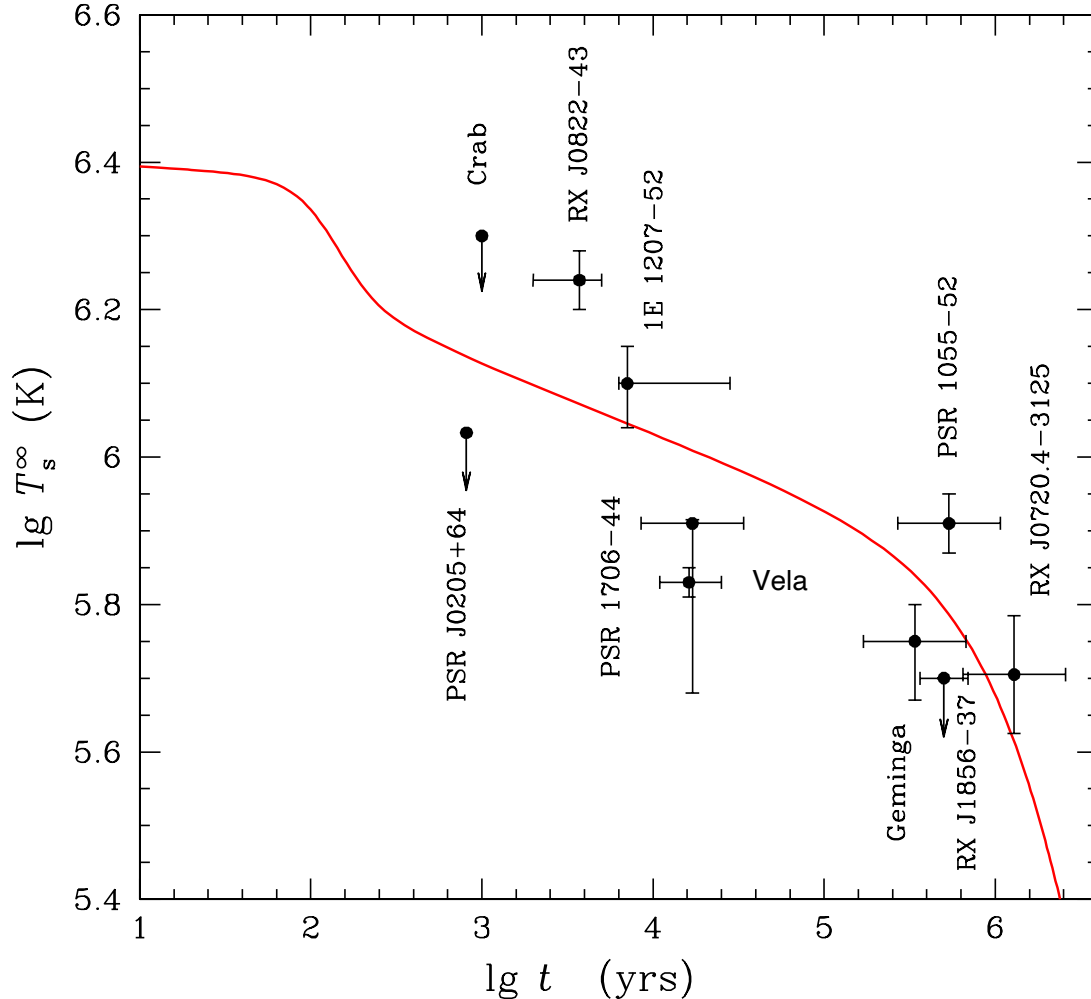


Zero temperature extrapolation

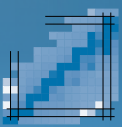




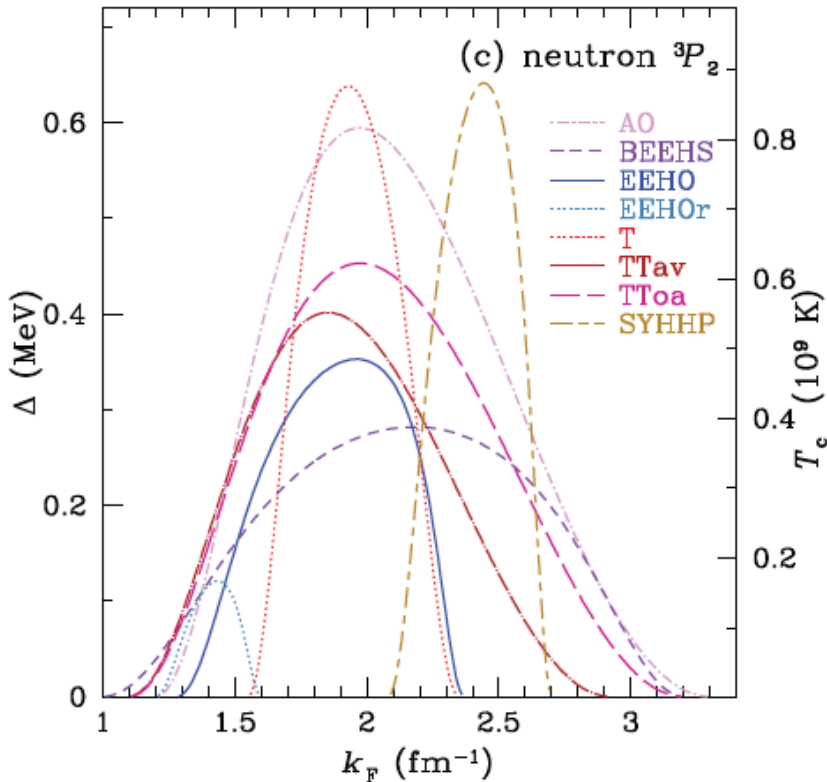
Cooling curve of neutron stars



- Observational data available for a handful of NS
- Sensitive to **interior** physics (mostly **pairing**)

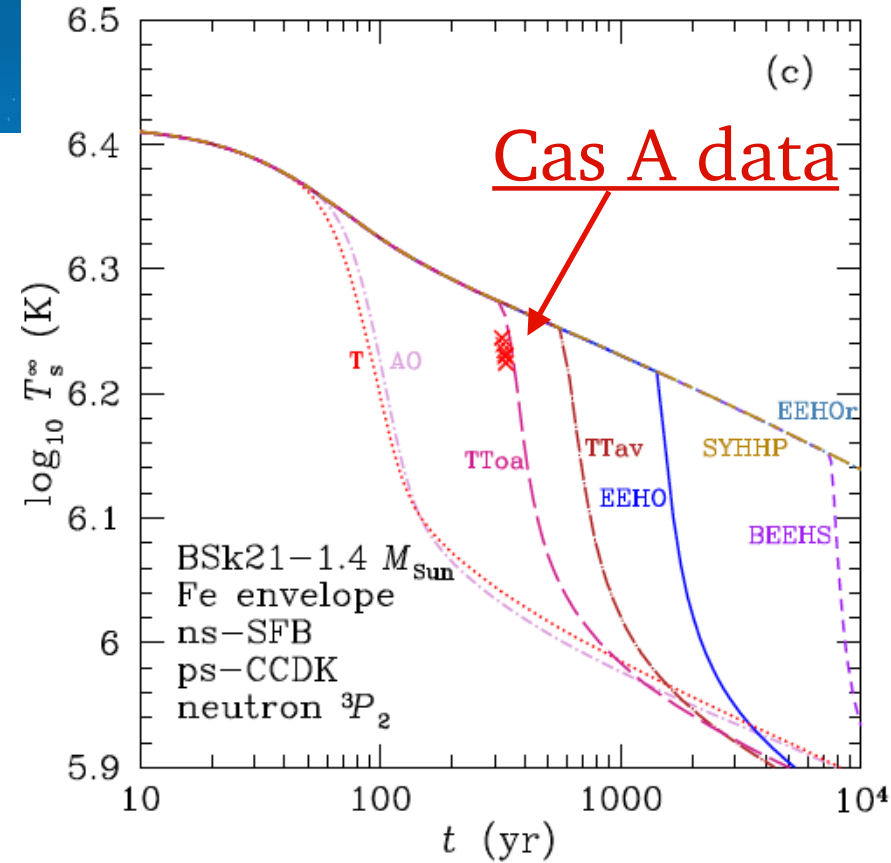


Cooling of CasA



[Ho, et al., PRC 91 015806 \(2015\)](#)

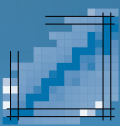
[Page, et al., PRL 106 081101 \(2011\)](#)



Ingredients

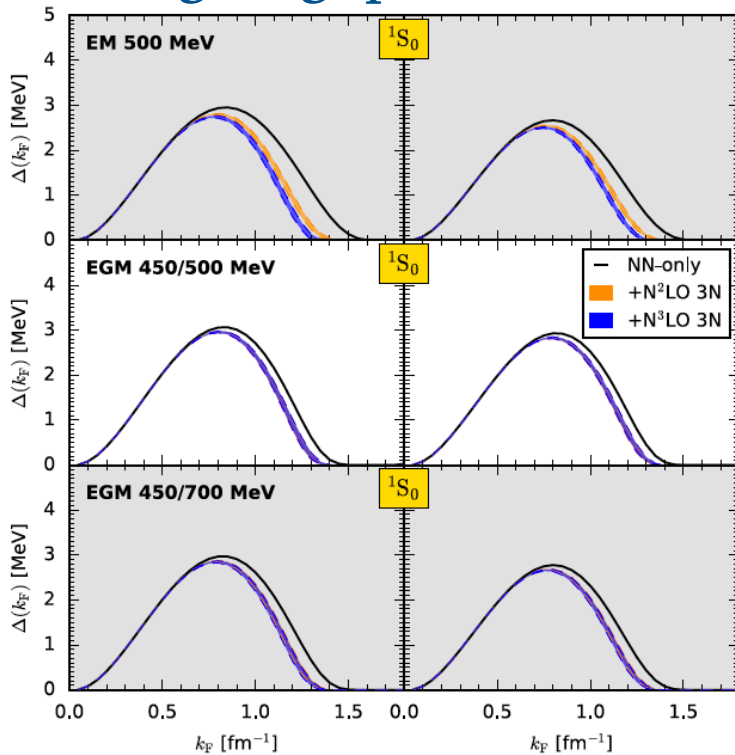
- (a) Mass of pulsar
- (b) EoS (determines radius)
- (c) Internal composition
- (d) **Pairing gaps (1S_0 & 3PF_2 channels)**
- (e) Atmosphere composition

Name	Process	Emissivity ($\text{erg cm}^{-3} \text{s}^{-1}$)
Modified Urca (neutron branch)	$n+n \rightarrow n+p+e^- + \bar{\nu}_e$	$\sim 2 \times 10^{21} RT_9^8$
	$n+p+e^- \rightarrow n+n+\nu_e$	
Modified Urca (proton branch)	$p+n \rightarrow p+p+e^- + \bar{\nu}_e$	$\sim 10^{21} RT_9^8$
	$p+p+e^- \rightarrow p+n+\nu_e$	
Bremsstrahlungs	$n+n \rightarrow n+n+\nu+\bar{\nu}$	$\sim 10^{19} RT_9^8$
	$n+p \rightarrow n+p+\nu+\bar{\nu}$	
	$p+p \rightarrow p+p+\nu+\bar{\nu}$	
Cooper pair	$n+n \rightarrow [nm] + \nu + \bar{\nu}$	$\sim 5 \times 10^{21} RT_9^7$
	$p+p \rightarrow [pp] + \nu + \bar{\nu}$	$\sim 5 \times 10^{19} RT_9^7$
Direct Urca (nucleons)	$n \rightarrow p+e^- + \bar{\nu}_e$	$\sim 10^{27} RT_9^6$
	$p+e^- \rightarrow n+\nu_e$	

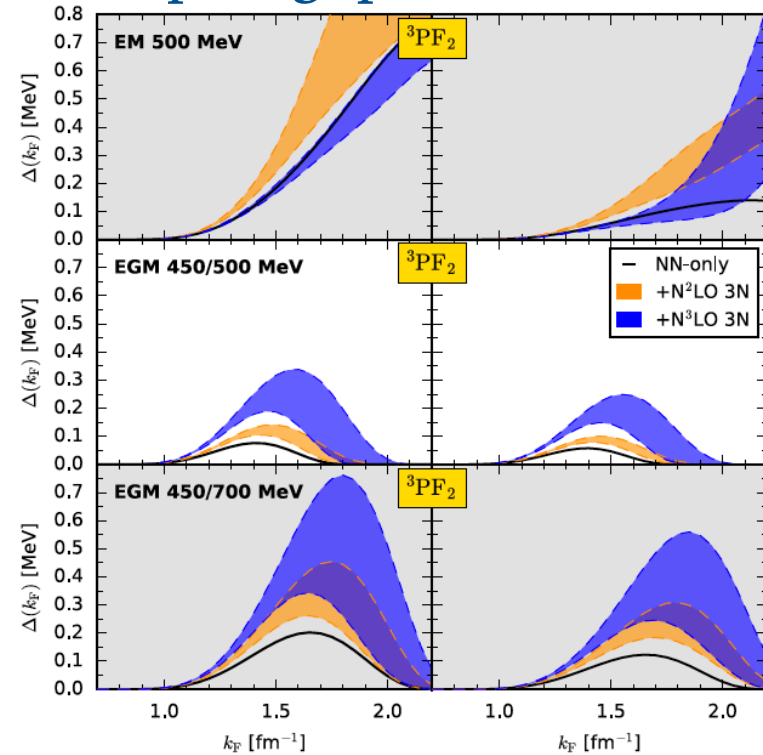


BCS+HF gaps in neutron matter

Singlet gaps with 3NF



Triplet gaps with 3NF

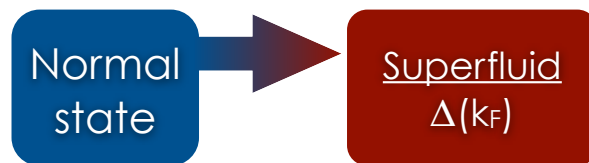
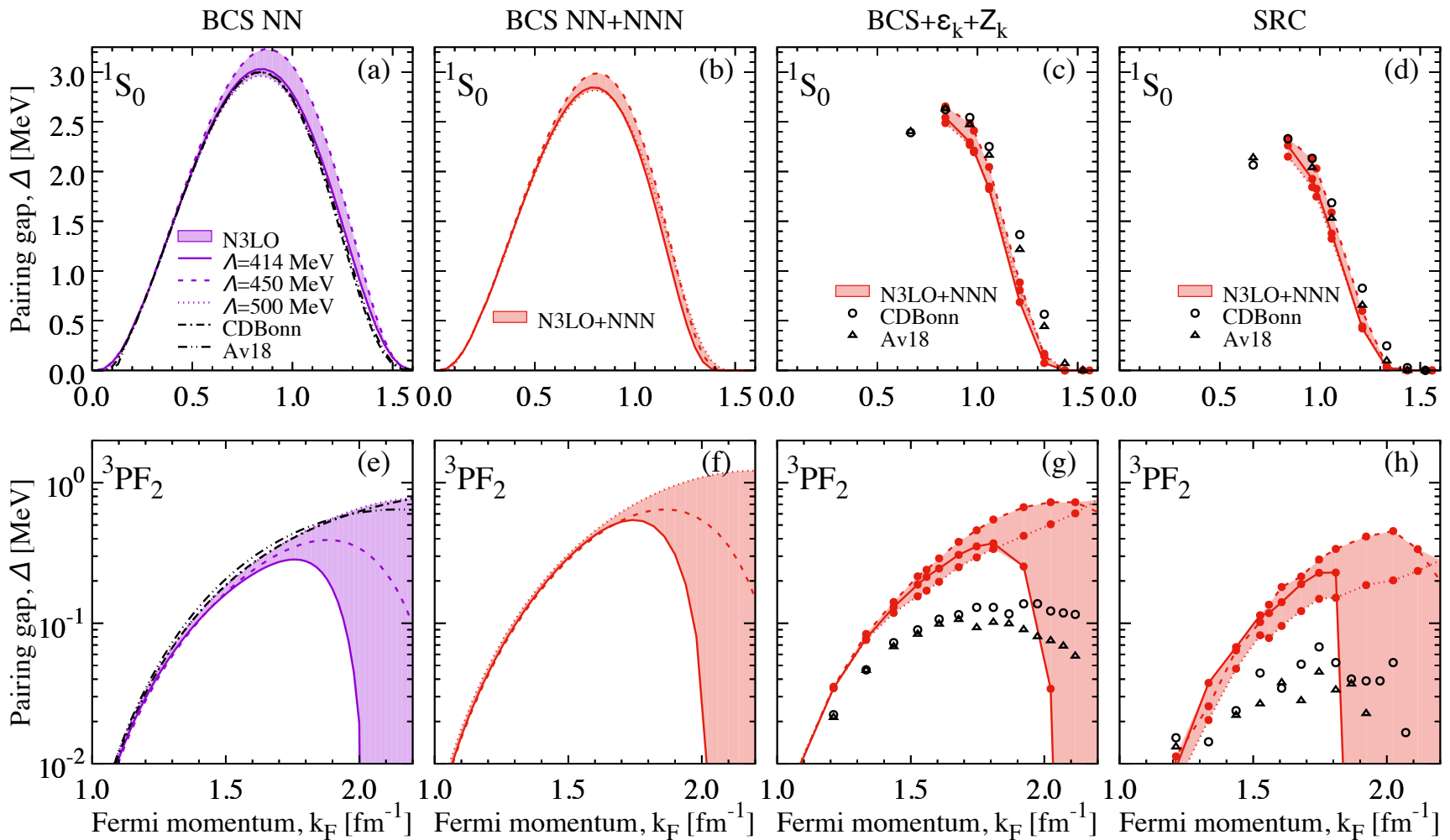


BCS equation

$$\Delta_k^L = - \sum_{L'} \int_{k'} \frac{\langle k | V_{nn}^{LL'} | k' \rangle}{2\sqrt{\chi_{k'}^2 + |\Delta_{k'}^{L'}|^2}} \Delta_{k'}^{L'} + \begin{aligned} \chi_k &= \varepsilon_k - \mu \\ \varepsilon_k &= \frac{k^2}{2m} + U(k) - \mu \end{aligned}$$

- Error estimates from nuclear force (chiral expansion) ✓
- Many-body uncertainty? ✗

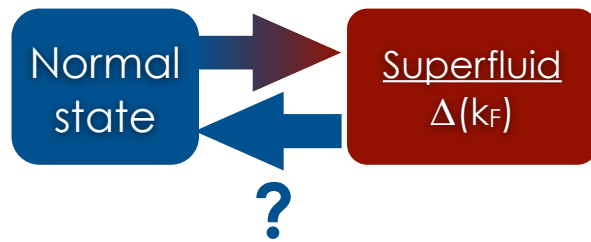
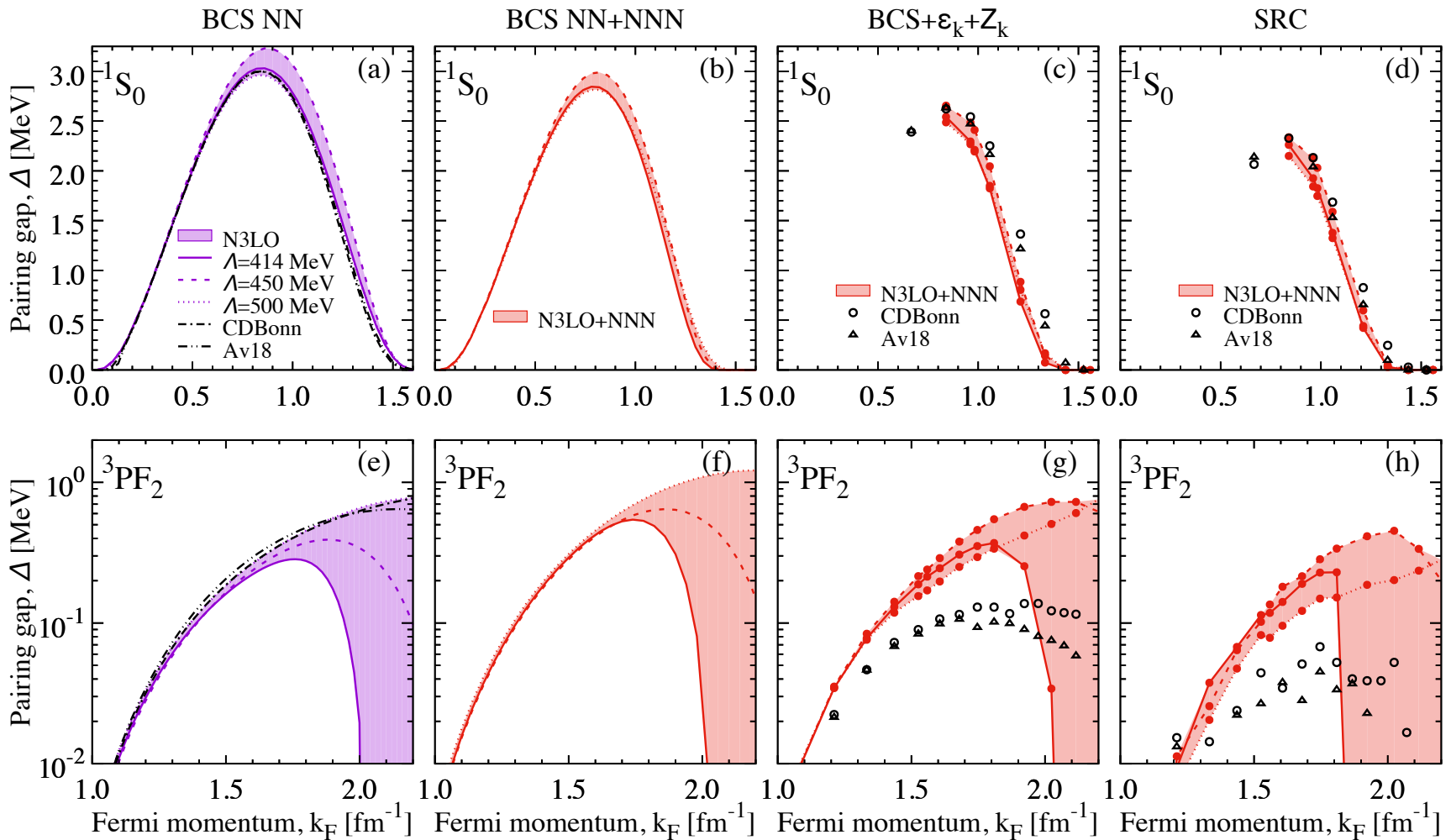
Beyond-BCS in neutron matter: SRC



Rios, Dickhoff, Polls, JLTJ **189**, 234 (2017)
[arXiv:1707.04140]

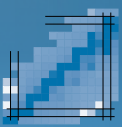
Rios, Dickhoff, Polls, et al PRC **94**, 025802 (2016)
[arxiv:1601.01600]

Beyond-BCS in neutron matter: SRC

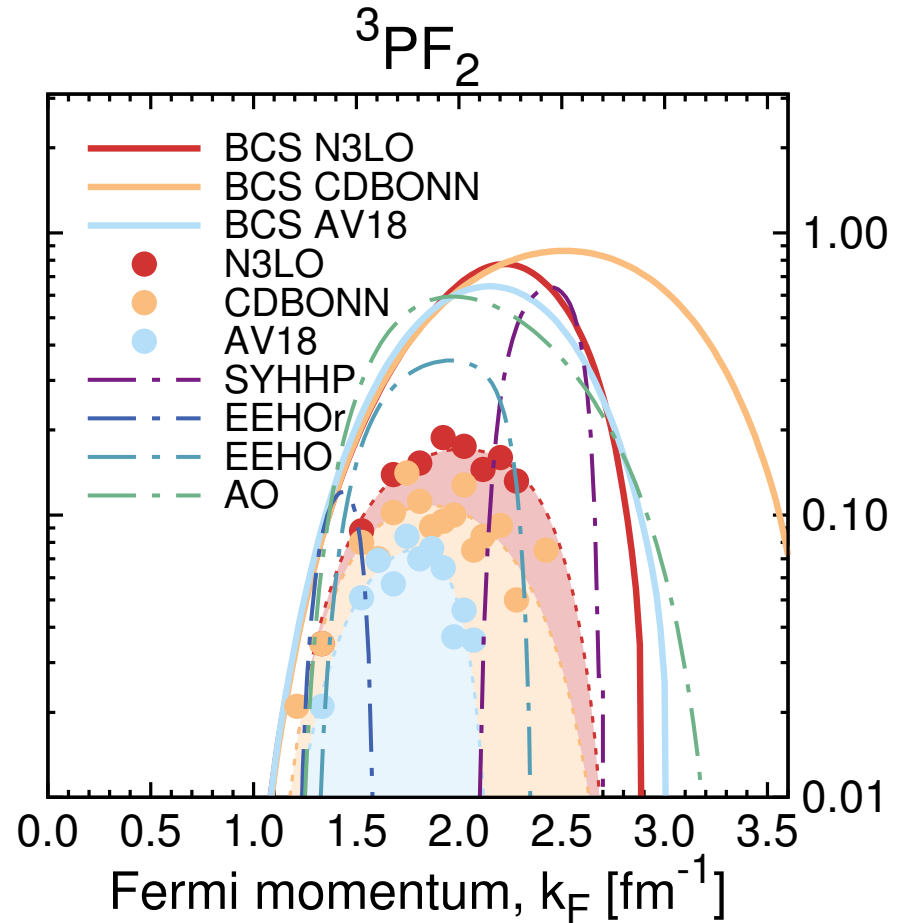
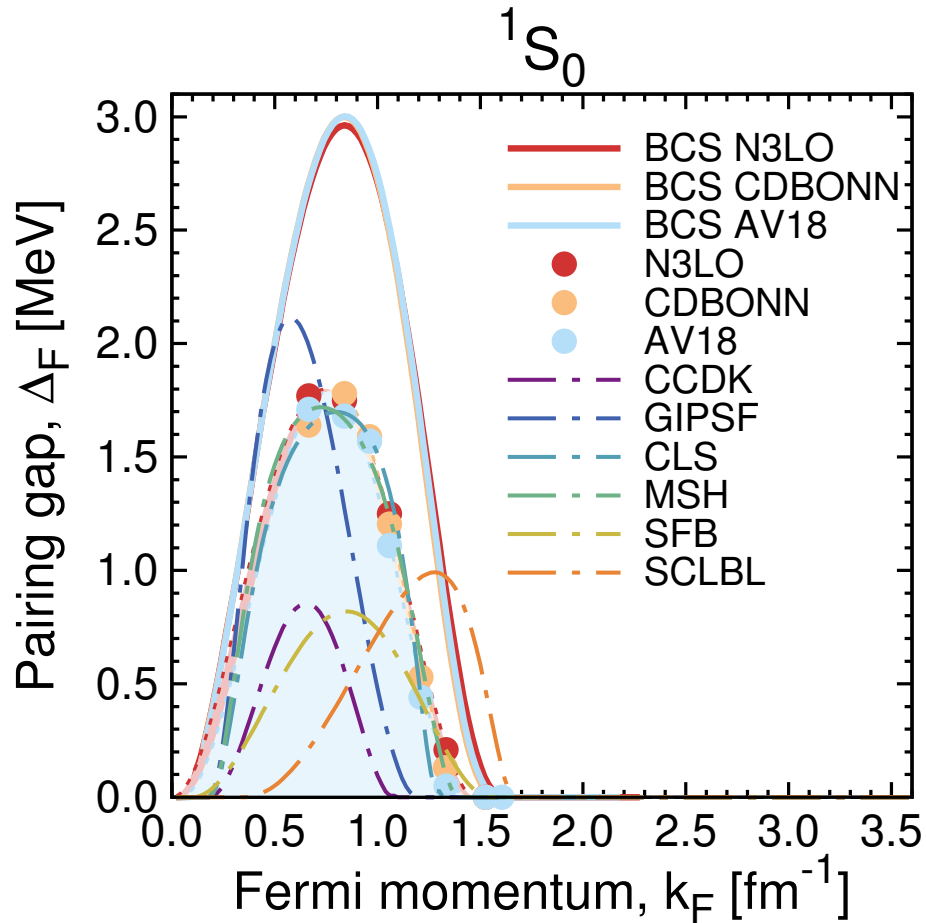


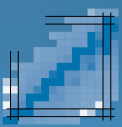
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[arxiv:1601.01600]

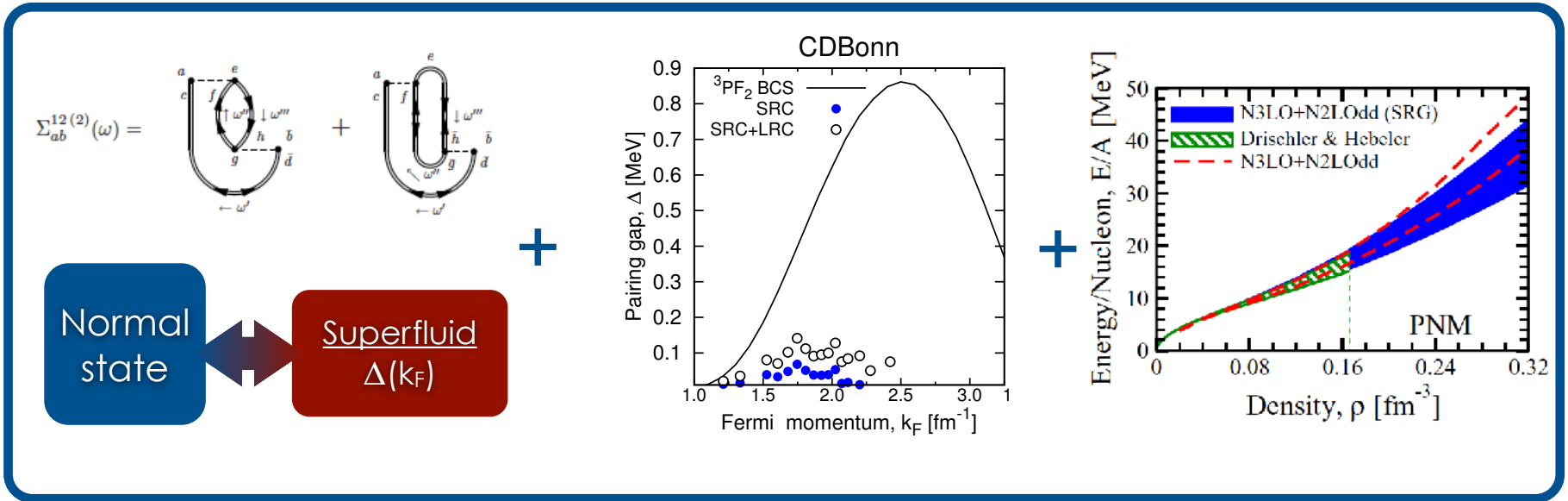


Beyond-BCS pairing: overview





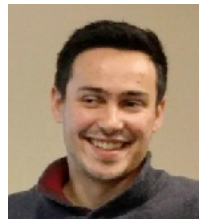
How to go beyond BCS?



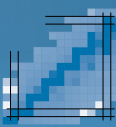
Existing frameworks difficult to generalise

Nambu-covariant **SCGF** technique

- Symmetry breaking ✓
- Finite temperature ✓
- Systematic expansion w diagrams ✓
- 3 nucleon forces ✓



M. Drissi



Nambu-Covariant Perturbation Theory

Nambu fields

[Anderson, 1958] [Nambu, 1960]

- \mathcal{B} and $\bar{\mathcal{B}} \equiv$ orthonormal bases $|b\rangle \rightarrow |\bar{b}\rangle$
- Let $\bar{\cdot}$ be the involution ($\bar{1} = 2, \bar{2} = 1$)
- Define $\mu \equiv (b, g)$ and $\bar{\mu} \equiv (\bar{b}, \bar{g})$ where $g \in \{1, 2\}$

- Then Nambu fields are defined as

$$A^\mu \equiv A^{(b,g)} \equiv \begin{pmatrix} a_b \\ a_{\bar{b}}^\dagger \end{pmatrix}_g$$

$$A_\mu^\dagger \equiv A_{(b,g)}^\dagger \equiv \begin{pmatrix} a_b^\dagger & a_{\bar{b}} \end{pmatrix}_g$$

- Canonical anticommutation relation

$$\{A^\mu, A^\nu\} = \delta_{\mu\bar{\nu}} \quad , \quad \{A_\mu^\dagger, A_\nu^\dagger\} = \delta_{\mu\bar{\nu}} \quad , \quad \{A^\mu, A_\nu^\dagger\} = \delta_{\mu\nu}$$

Tensor definitions

- Let \mathcal{W} a unitary Bogoliubov transformation

$$B^\mu = \sum_\nu (\mathcal{W}^\dagger)^\mu_\nu A^\nu$$

$$B_\mu^\dagger = \sum_\nu \mathcal{W}^\nu_\mu A_\nu^\dagger$$

- Definition: (p, q) -tensor is multi-dim array s.t.

$$t^{\mu_1 \dots \mu_p}_{\nu_1 \dots \nu_q} \equiv \sum_{\kappa_1 \dots \kappa_p} \sum_{\lambda_1 \dots \lambda_q} (\mathcal{W}^\dagger)^{\mu_1}_{\kappa_1} \dots (\mathcal{W}^\dagger)^{\mu_p}_{\kappa_p} t^{\kappa_1 \dots \kappa_p}_{\lambda_1 \dots \lambda_q} (\mathcal{W})^{\lambda_1}_{\nu_1} \dots (\mathcal{W})^{\lambda_q}_{\nu_q}$$

- p contravariant & q covariant indices

Operators

- Operators as polynomial of Nambu fields

$$O \equiv \sum_{\mu_1 \dots \mu_{2k}} o^{\mu_1 \dots \mu_k}_{\mu_{k+1} \dots \mu_{2k}} A_{\mu_1}^\dagger \dots A_{\mu_k}^\dagger A^{\mu_{k+1}} \dots A^{\mu_{2k}}$$

$$O \equiv \sum_{\mu_1 \dots \mu_{2k}} o_{\mu_1 \dots \mu_{2k}} A^{\mu_1} \dots A^{\mu_{2k}}$$

$$O \equiv \sum_{\mu_1 \dots \mu_{2k}} o^{\mu_1 \dots \mu_{2k}} A_{\mu_1}^\dagger \dots A_{\mu_{2k}}^\dagger$$

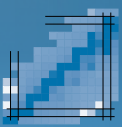
Metric tensor

- Definition: $(0, 2)$ -, $(1, 1)$ -, $(2, 0)$ -tensors

$$\left. \begin{aligned} g_{\mu\nu} &\equiv \delta_{\mu\bar{\nu}} \\ g^\mu_\nu &\equiv \delta_{\mu\nu} \\ g^{\mu\nu} &\equiv \delta_{\mu\bar{\nu}} \end{aligned} \right\} + \text{transform like a tensor}$$

- Raising/lowering indices of a tensor:

$$o_{\mu_1 \dots \mu_{2k}} = \sum_{\alpha_1 \dots \alpha_k} g_{\mu_1 \alpha_1} \dots g_{\mu_k \alpha_k} o^{\alpha_1 \dots \alpha_k}_{\mu_{k+1} \dots \mu_{2k}}$$

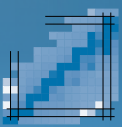


Why Bogoliubov tensor algebra?

$$\text{Tensor product: } r^{\mu_1}_{\nu_1}{}^{\mu_2\mu_3} = s^{\mu_1}_{\nu_1} t^{\mu_2\mu_3}$$

$$\text{Tensor contraction: } r^{\mu}_{\nu} = \sum_{\alpha} s^{\mu}_{\alpha} t^{\alpha}_{\nu}$$

- Co(contra-)variance under Bogoliubov transforms provide **invariant** expressions in any basis
- Potential to **optimise** the extended basis
- **Tensor-network** structure becomes **transparent**
- Leads to **diagrammatic** expansion (à la de Dominicis-Martin or Haussmann)
- Other **formalisms** through specific basis or metric



Perturbative expansion

Hamiltonian partitioning

$$\Omega = \Omega_0 + \Omega_1$$

$$\Omega_0 = \frac{1}{2} \sum_{\mu\nu} U_{\mu\nu} A^\mu A^\nu$$

$$\Omega_1 = \sum_{k=1}^n \frac{1}{(2k)!} \sum_{\mu_1 \dots \mu_{2k}} v_{\mu_1 \dots \mu_{2k}}^{(k)} A^{\mu_1} \dots A^{\mu_{2k}}$$

Covariant k-body vertices

Green's functions

• Contravariant k-body Green's function

$$(-1)^k \mathcal{G}^{\mu_1 \dots \mu_{2k}}(\tau_1, \dots, \tau_{2k}) \equiv \langle T [A^{\mu_1}(\tau_1) \dots A^{\mu_{2k}}(\tau_{2k})] \rangle$$

with $\langle . \rangle = \text{Tr} (. \rho)$ and $\rho \equiv \frac{e^{-\beta\Omega}}{\text{Tr} (e^{-\beta\Omega})}$

• Unperturbed case: $\Omega \longleftrightarrow \Omega_0$

Expansion

• Interaction picture expression

$$(-1)^k \mathcal{G}^{\mu_1 \dots \mu_{2k}}(\tau_1, \dots, \tau_{2k}) = \frac{\langle T [e^{-\int_0^\beta ds \Omega_1(s)} A^{\mu_1}(\tau_1) \dots A^{\mu_{2k}}(\tau_{2k})] \rangle_0}{\langle T e^{-\int_0^\beta ds \Omega_1(s)} \rangle_0}$$

• Perturbative expansions

$$\langle T [e^{-\int_0^\beta ds \Omega_1(s)} A^{\mu_1}(\tau_1) \dots A^{\mu_{2k}}(\tau_{2k})] \rangle_0 = \sum_{n=0}^{+\infty} \frac{(-1)^n}{n!} \int_0^\beta d\tau'_1 \dots \int_0^\beta d\tau'_n \langle T [\Omega_1(\tau'_1) \dots \Omega_1(\tau'_n) A^{\mu_1}(\tau_1) \dots A^{\mu_{2k}}(\tau_{2k})] \rangle_0$$

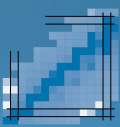
• Statistical Wick theorem + Linked-cluster theorem

\Rightarrow **Feynman diagrammatics almost as usual**

• We provide a set of **Feynman rules**

• Also rules to evaluate **Matsubara sums**

• **Simpler** expressions than in other approaches (Gorkov Green's functions or BMPT)



Feynman diagram building blocks

Formulations

Time-independent partitioning

- ▶ Time representation
 - ▶ Energy representation
- ↔ Fourier Transform

Matsubara frequencies

$$\mathcal{G}^{\mu\nu}(\omega_p) \equiv \int_0^\beta d\tau e^{i\omega_p\tau} \mathcal{G}^{\mu\nu}(\tau)$$

$$\mathcal{G}^{\mu\nu}(\tau) = \frac{1}{\beta} \sum_p e^{-i\omega_p\tau} \mathcal{G}^{\mu\nu}(\omega_p)$$

Time-dependent partitioning (not today!)

Propagators

$$-\mathcal{G}^{\mu\nu}(\omega_p) = \begin{array}{c} \mu \\ \parallel \\ \nu \end{array} \uparrow \omega_p$$

$$-(\mathcal{G}^{(0)})^{\mu\nu}(\omega_p) = \begin{array}{c} \mu \\ | \\ \nu \end{array} \uparrow \omega_p$$

Fully antisymmetric vertex

Definition

$$v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}] }^{(k)} \equiv \frac{1}{(2k)!} \sum_{\sigma \in S_{2k}} \epsilon(\sigma) v_{\mu_{\sigma(1)} \mu_{\sigma(2)} \dots \mu_{\sigma(2k-1)} \mu_{\sigma(2k)}}^{(k)}$$

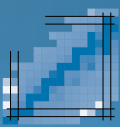
Antisymmetrisation defines a new $(0, 2k)$ -tensor

Not the case in a mixed representation

k-body vertex

$$v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}] }^{(k)} = \begin{array}{c} \mu_1 \quad \mu_2 \\ \diagdown \quad | \\ \bullet \\ \diagup \quad | \\ \mu_{2k} \quad \mu_{2k-1} \end{array}$$

(Note: The diagram shows a central black dot with four solid lines and two dashed lines extending from it. The solid lines are labeled with indices $\mu_1, \mu_2, \mu_{2k-1}, \mu_{2k}$ and the dashed lines are unlabeled.)



Feynman diagram building blocks

Formulations

Time-independent partitioning

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Covariant **k-body vertices**

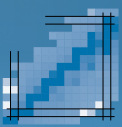
Antisymmetrisation defines a new $(0, 2k)$ -tensor

Not the case in a *mixed* representation

k-body vertex

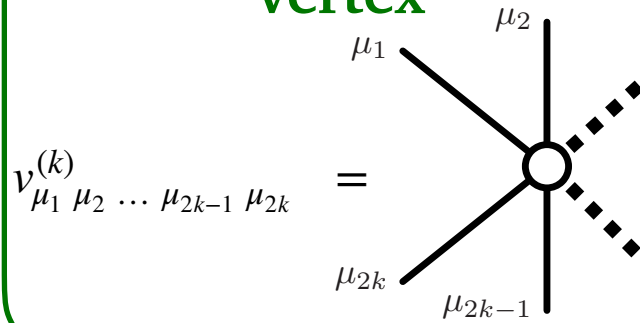
$$v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}] }^{(k)} = \begin{array}{c} \mu_1 \quad \mu_2 \\ \diagdown \quad | \\ \bullet \\ \diagup \quad | \\ \mu_{2k} \quad \mu_{2k-1} \end{array}$$

(Note: The diagram shows a central black dot with four solid lines and two dashed lines extending from it, representing a k-body vertex.)

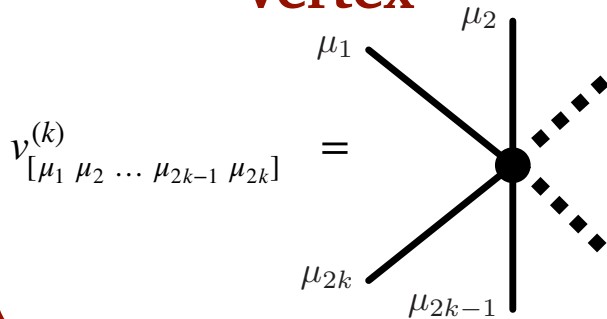


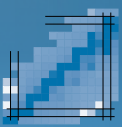
Why antisymmetric vertices?

Un-symmetrised vertex



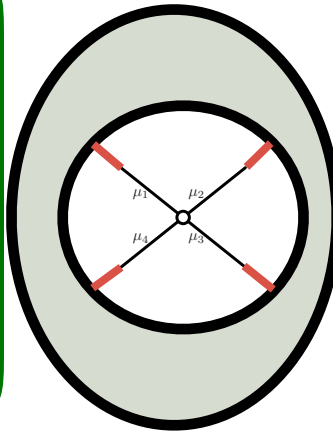
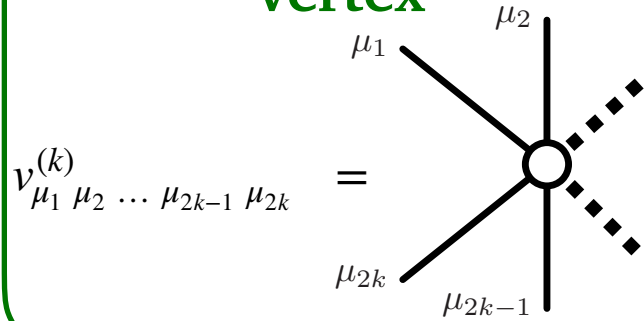
Antisymmetrized vertex



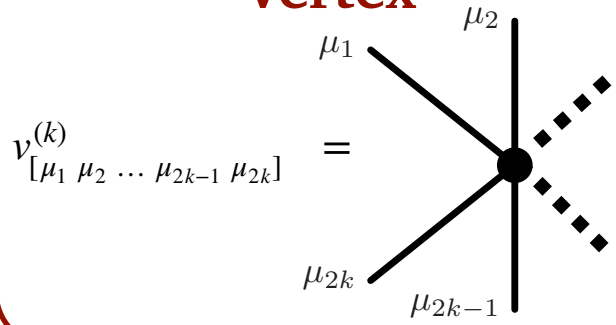


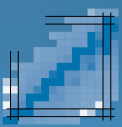
Why antisymmetric vertices?

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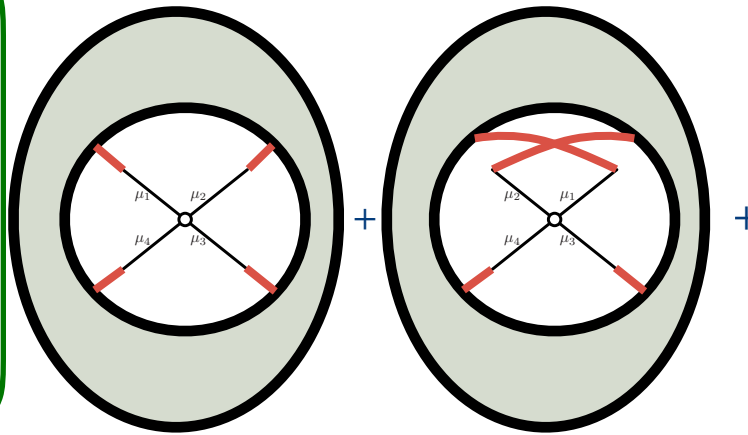
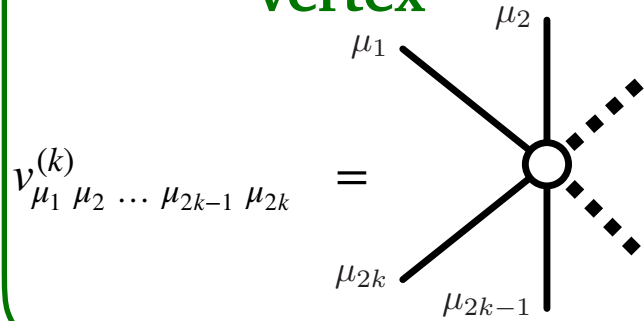
Antisymmetrized vertex



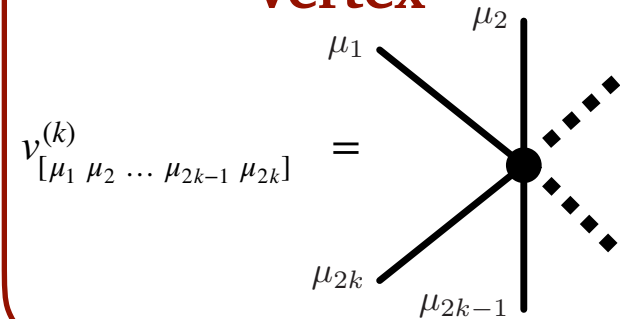


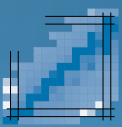
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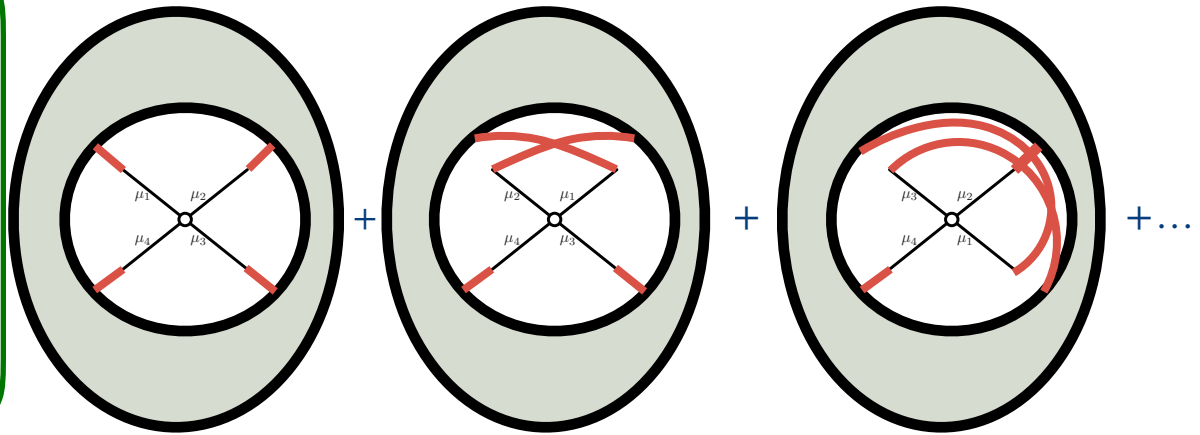
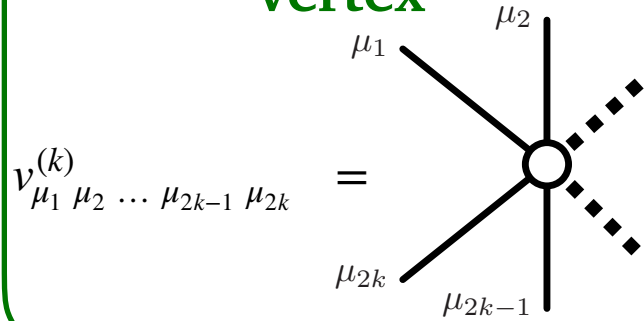
Antisymmetrized vertex



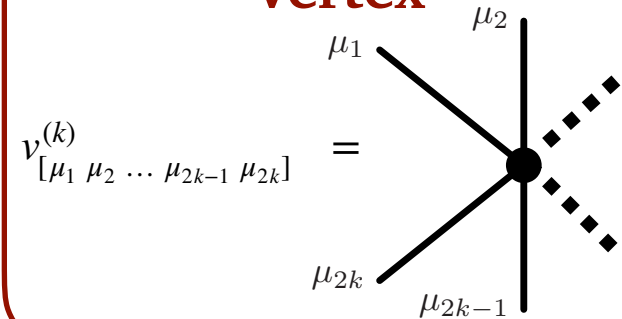


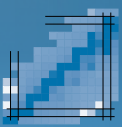
Why antisymmetric vertices?

Un-symmetrised vertex



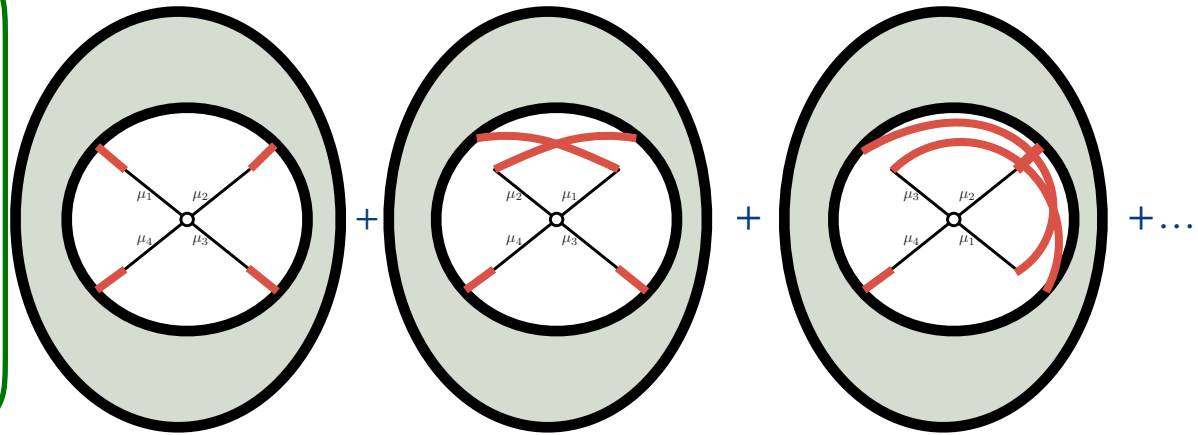
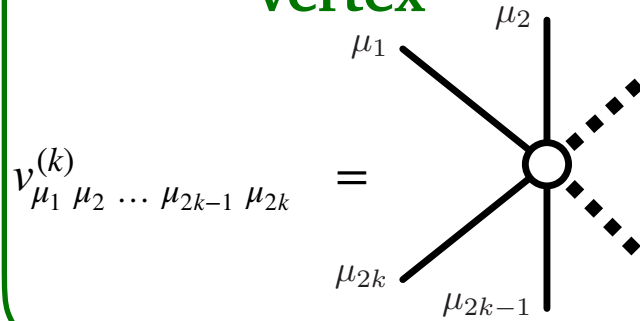
Antisymmetrized vertex





Why antisymmetric vertices?

Un-symmetrised vertex



Antisymmetrized vertex

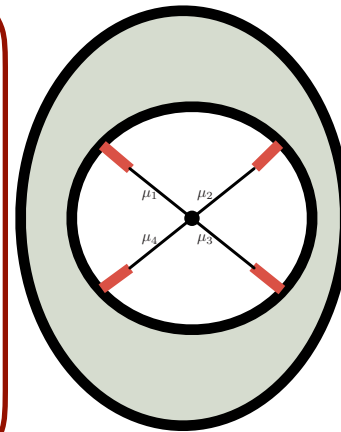
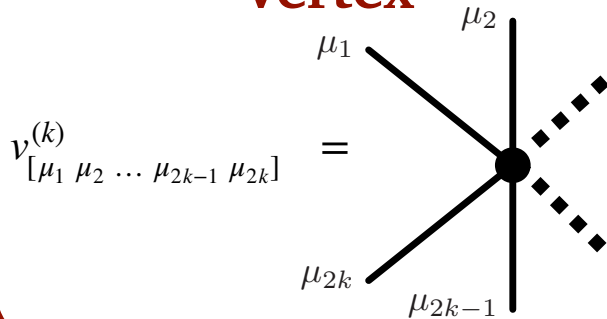
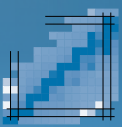


Diagram factorisation

- Derivations rely on
 - ▶ Wick theorem \Rightarrow sum over pairing
 - ▶ Sum over single-particle and Nambu indices
- \rightarrow Extends Hugenholtz antisymmetrisation
- Antisym is a one-off pre-computing cost



Perturbative expansion

Order n graphical rules

- Draw all topologically distinct **connected unlabelled** diagrams
 - with $2k$ external legs
 - with n vertices (for order n contributions)

Feynman rules

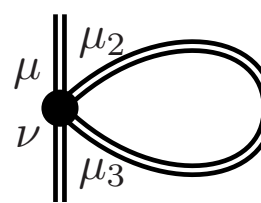
- Label vertices from 1 to n
 - S is the number of vertex labels permutations leaving the diagram invariant
- For each line multiply by $-(\mathcal{G}^{(0)})^{\mu\nu}(\omega_e)$
- For each k -body vertex multiply by $v_{[\mu_1 \mu_2 \dots \mu_{2k-1} \mu_{2k}]}$
- Sum over each internal μ index and each independent ω_e frequency

5. Multiply by
$$\frac{(-1)^{n+L}}{S \times 2^T \prod_{l=2}^{l_{\max}} (l!)^m}$$

Gaudin rules

- These simplify Matsubara sums
- Require **spanning trees**

Tadpoles are exceptional



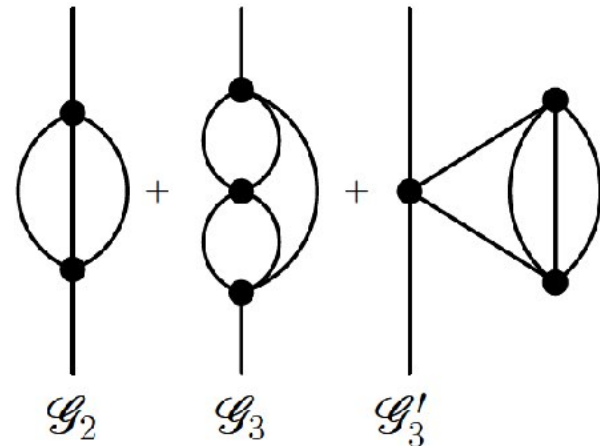
$$I_{\mu\nu} = \sum_{\mu_2 \dots \mu_{2k-1}} \frac{(-1)^k}{2^{k-1}(k-1)!} v_{[\mu \mu_2 \mu_3 \nu]}^{(k=2)} \times \frac{1}{\beta} \sum_{\omega_e} -\mathcal{G}^{\mu_2 \mu_3}(\omega_e) e^{-i\omega_e \eta_p}$$

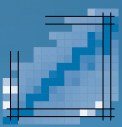
- Partially antisymmetrized vertices needed:

$$v_{[\mu_1 \dots \mu_x \dots \mu_y \dots \mu_{2k}]}^{(k)} \equiv \frac{2^p p!}{(2k)!} \sum_{\sigma \in S_{2k}/S_2^p \times S_p} \epsilon(\sigma) v_{\mu_{\sigma(1)} \dots \mu_x \dots \mu_y \dots \mu_{\sigma(2k)}}^{(k)}$$

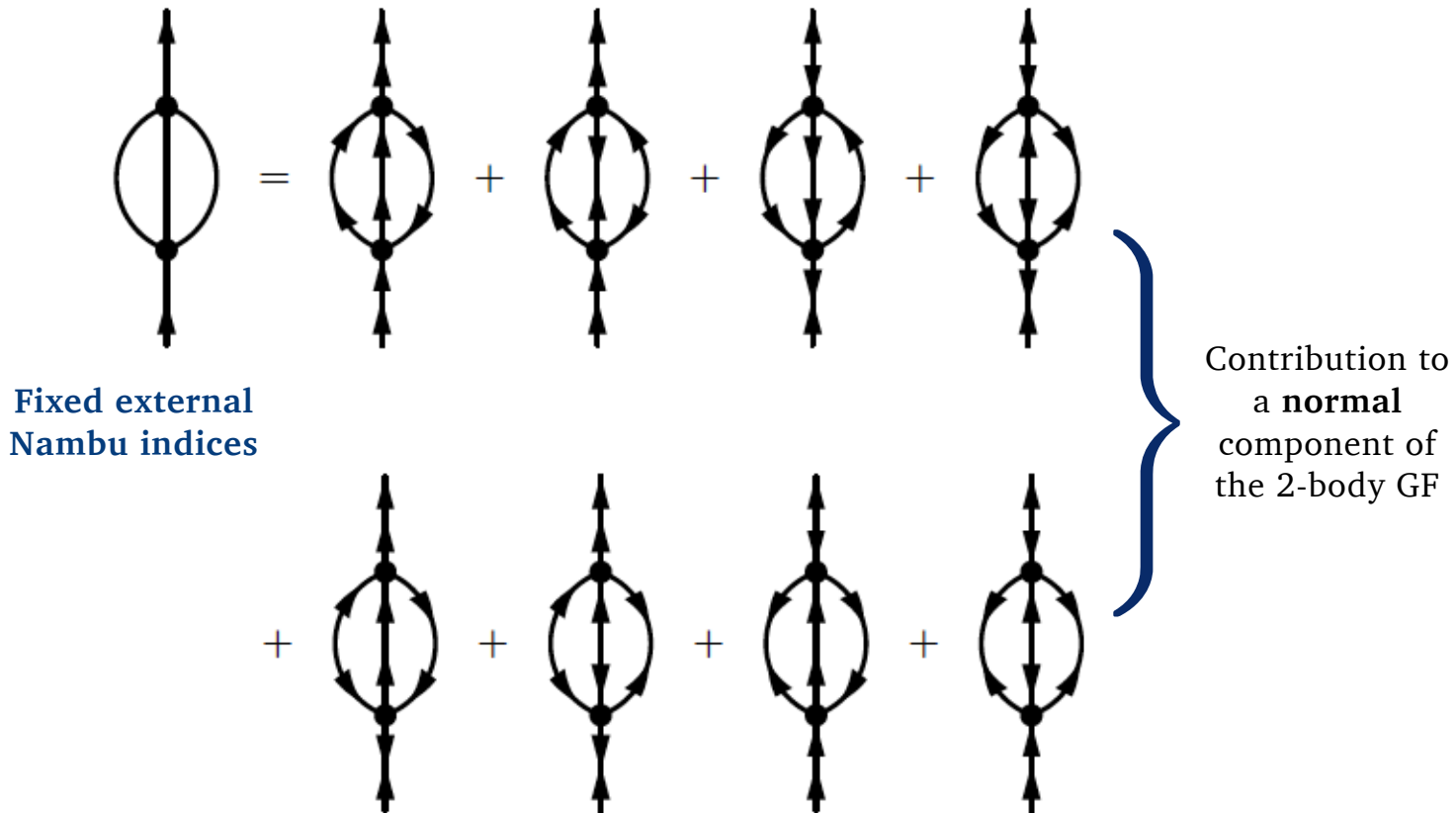
- p internal lines are **fixed**
- k -body generalisation works**

HFB partitioning 3rd order

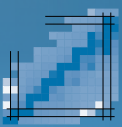




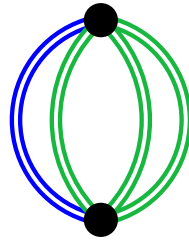
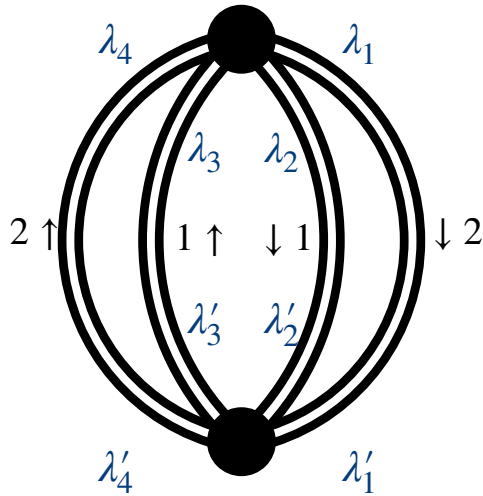
Connection with Gorkov



- Changes of **single-particle basis** bring admixtures in normal & anomalous components...
- but leave Nambu-convariant diagrams **invariant!**

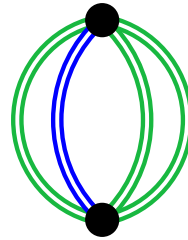


Basic example



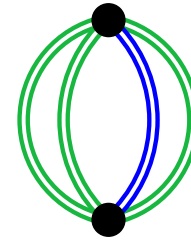
\mathcal{A}_4

$$\frac{-f(\epsilon_3) f(-\epsilon_2) f(-\epsilon_1)}{\epsilon_4 + \epsilon_3 - \epsilon_2 - \epsilon_1}$$



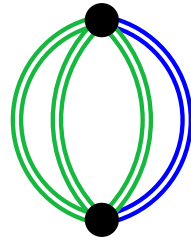
\mathcal{A}_3

$$\frac{f(-\epsilon_4) f(-\epsilon_2) f(-\epsilon_1)}{\epsilon_4 + \epsilon_3 - \epsilon_2 - \epsilon_1}$$



\mathcal{A}_2

$$\frac{f(-\epsilon_4) f(-\epsilon_3) f(-\epsilon_1)}{-\epsilon_4 - \epsilon_3 + \epsilon_2 + \epsilon_1}$$



\mathcal{A}_1

$$\frac{f(-\epsilon_4) f(-\epsilon_3) (-f(\epsilon_2))}{-\epsilon_4 - \epsilon_3 + \epsilon_2 + \epsilon_1}$$

$$I = \frac{1}{48} \sum_{\substack{\lambda_1 \lambda_2 \lambda_3 \lambda_4 \\ \lambda'_1 \lambda'_2 \lambda'_3 \lambda'_4}} v_{[\lambda_1 \lambda_2 \lambda_3 \lambda_4]}^{(2)} v_{[\lambda'_1 \lambda'_2 \lambda'_3 \lambda'_4]}^{(2)} \int_{-\infty}^{+\infty} \frac{d\epsilon_1}{2\pi} \frac{d\epsilon_2}{2\pi} \frac{d\epsilon_3}{2\pi} \frac{d\epsilon_4}{2\pi} S^{\lambda_1 \lambda_1}(\epsilon_1) S^{\lambda_2 \lambda_2}(\epsilon_2) S^{\lambda_3 \lambda_3}(\epsilon_3) S^{\lambda_4 \lambda_4}(\epsilon_4) \\ \times \frac{f(\epsilon_1) f(\epsilon_2) f(-\epsilon_3) f(-\epsilon_4) - f(-\epsilon_1) f(-\epsilon_2) f(\epsilon_3) f(\epsilon_4)}{\epsilon_1 + \epsilon_2 - \epsilon_3 - \epsilon_4}$$

Spectral function

$$\mathcal{G}^{\mu\nu}(\omega_p) = \int_{-\infty}^{+\infty} \frac{d\omega'}{2\pi} \frac{S^{\mu\nu}(\omega')}{i\omega_p - \omega'}$$

Lehmann representation

$$S^{\mu\nu}(\omega) \equiv \frac{1}{Z} \sum_{m,n} \langle \Psi_m | A^\mu | \Psi_n \rangle \langle \Psi_n | A^\nu | \Psi_m \rangle \\ \times e^{-\beta\Omega_m} (1 + e^{-\beta\omega}) (2\pi) \delta(\Omega_n - \Omega_m - \omega)$$

Properties

Symmetries

- ▶ **Hermiticity:** $(S^\mu_\nu(\omega))^* = S^\nu_\mu(\omega)$
- ▶ **Antisymmetry:** $S^\mu_\nu(\omega) = S^\nu_\mu(-\omega) \quad (\neq S^\nu_\mu(-\omega))$

Positive bounds

- ▶ $S(\omega) > 0 \iff$ each principal minor is strictly positive

Self-consistent Green's function resummation

Dyson equation

- Partitioning considered

$$\Omega = \underbrace{\frac{1}{2!} \sum_{\mu\nu} U_{\mu\nu} A^\mu A^\nu}_{\Omega_0} + \underbrace{\frac{1}{4!} \sum_{\alpha\beta\gamma\delta} v_{\alpha\beta\gamma\delta}^{(2)} A^\alpha A^\beta A^\gamma A^\delta}_{\Omega_1}$$

- Dyson equation

$$\mathcal{G}^{\mu\nu}(\omega_n) = \mathcal{G}^{(0)\mu\nu}(\omega_n) + \sum_{\lambda_1\lambda_2} \mathcal{G}^{(0)\mu\lambda_1}(\omega_n) \Sigma_{\lambda_1\lambda_2}(\omega_n) \mathcal{G}^{\lambda_2\nu}(\omega_n)$$

Diagrammatic expansion of

$$\Sigma_{\mu\nu}(\omega_n)$$

- with unperturbed propagators

$$\Sigma_{\mu\nu}(\omega_n) = \frac{\mathcal{F}_{\mu\nu}(\omega_n) - \mathcal{F}_{\nu\mu}(-\omega_n)}{2}$$

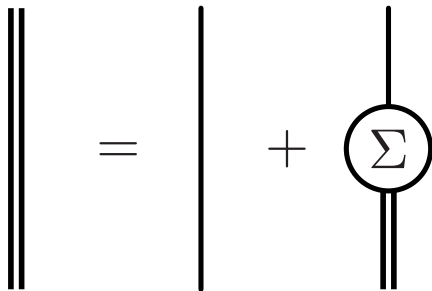
$$\mathcal{F}_{\mu\nu}(\omega_n) = \sum \text{1PI diagrams with } \mathcal{G}^{(0)}$$

- with self-consistent propagators

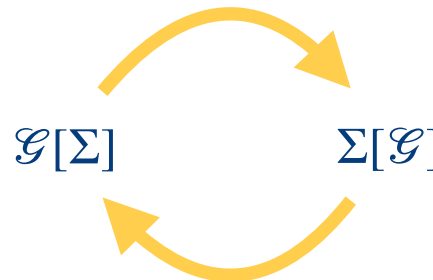
$$\Sigma_{\mu\nu}(\omega_n) = \frac{\mathcal{F}_{\mu\nu}(\omega_n) - \mathcal{F}_{\nu\mu}(-\omega_n)}{2}$$

$$\mathcal{F}_{\mu\nu}(\omega_n) = \sum \text{2PI diagrams with } \mathcal{G} \quad \left(= \mathcal{F}_{\mu\nu}(\omega_n) \right)$$

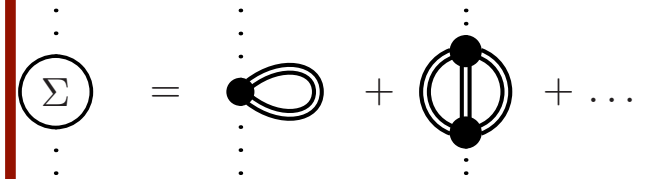
Diagrammatic representation

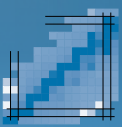


SCGF cycle



Self-energy expression





T-matrix: ladder's rung

Approximations on $\Gamma_{2\text{PFI}}^{(2)}$

- Sum of all possible rungs

$$\Gamma_{2\text{PFI}}^{(2)} = \text{cross} + \text{loop} + \dots$$

Ladder approximation

T-matrix $\equiv \Gamma^{(2)}$ in ladder approximation

$$T = \text{cross} + \text{loop} + \dots$$

$$T = \text{cross} + \text{loop} + \text{two loops} + \dots$$

Ladder approximation

- Analytic/Retarded/Advanced/Sp function \Rightarrow as usual

- T-matrix equation

$$T_{MN}(Z) = V_{MN}^{(2)} + \frac{1}{2} \sum_{LL'} V_{ML}^{(2)} \Pi^{LL'}(Z) T_{L'N}(Z)$$

where $V_{MN}^{(2)} \equiv v_{[\mu_1\mu_2\nu_1\nu_2]}^{(2)}$, $M \equiv (\mu_1, \mu_2)$ & $N \equiv (\nu_1, \nu_2)$

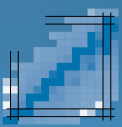
Solving the ladder

- Spectral representation

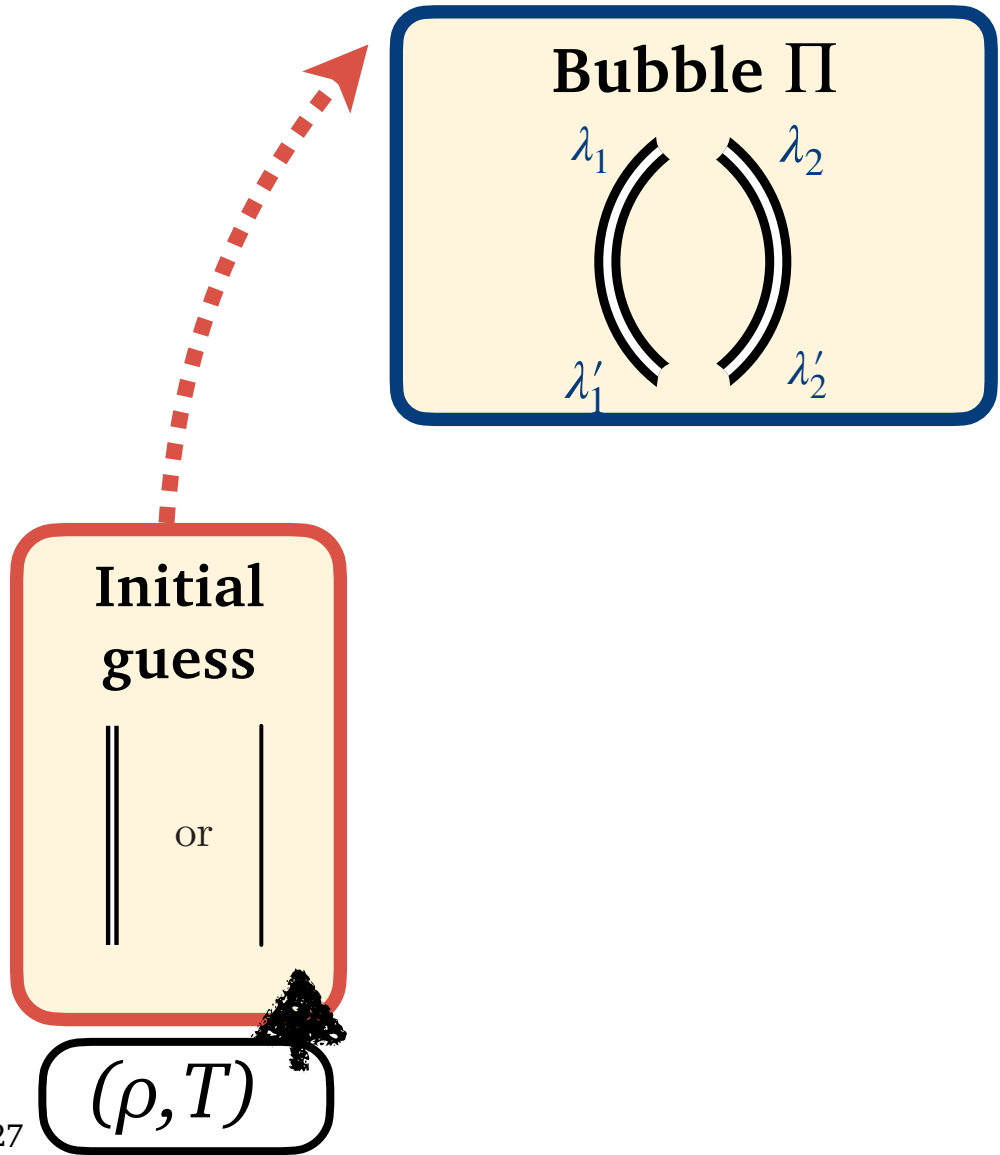
$$T_{MN}(Z) \equiv \underbrace{V_{MN}^{(2)}}_{\text{Instantaneous part}} + \int_{-\infty}^{+\infty} \frac{d\Omega}{2\pi} \frac{\mathcal{T}_{MN}(\Omega)}{Z - \Omega} \left. \vphantom{\int} \right\} T^C(Z) \equiv \text{continuous part}$$

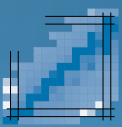
- Solution

$$\mathcal{T}(\Omega) = iV^{(2)} \left\{ \left(gg - \frac{1}{2} \Pi^R(\Omega) V^{(2)} \right)^{-1} - \left(gg - \frac{1}{2} \Pi^A(\Omega) V^{(2)} \right)^{-1} \right\}$$



Nambu-Covariant Ladders





Nambu-Covariant Ladders

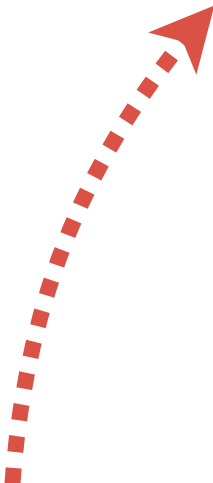
Initial guess

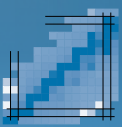
|| or |

(ρ, T)

Bubble Π

T-matrix





Nambu-Covariant Ladders

Initial guess

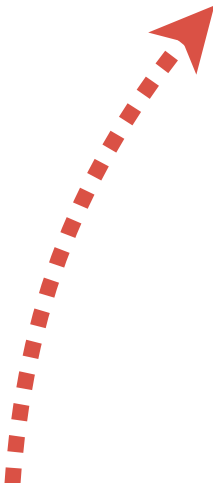
or

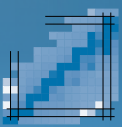
(ρ, T)

Bubble Π

T-matrix

Self-energy Σ





Nambu-Covariant Ladders

Initial guess

or

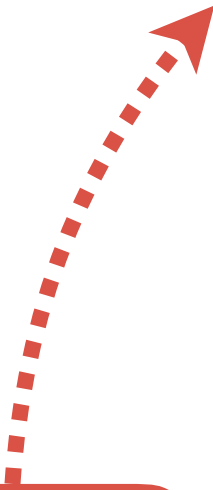
(ρ, T)

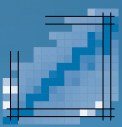
Bubble Π

T-matrix

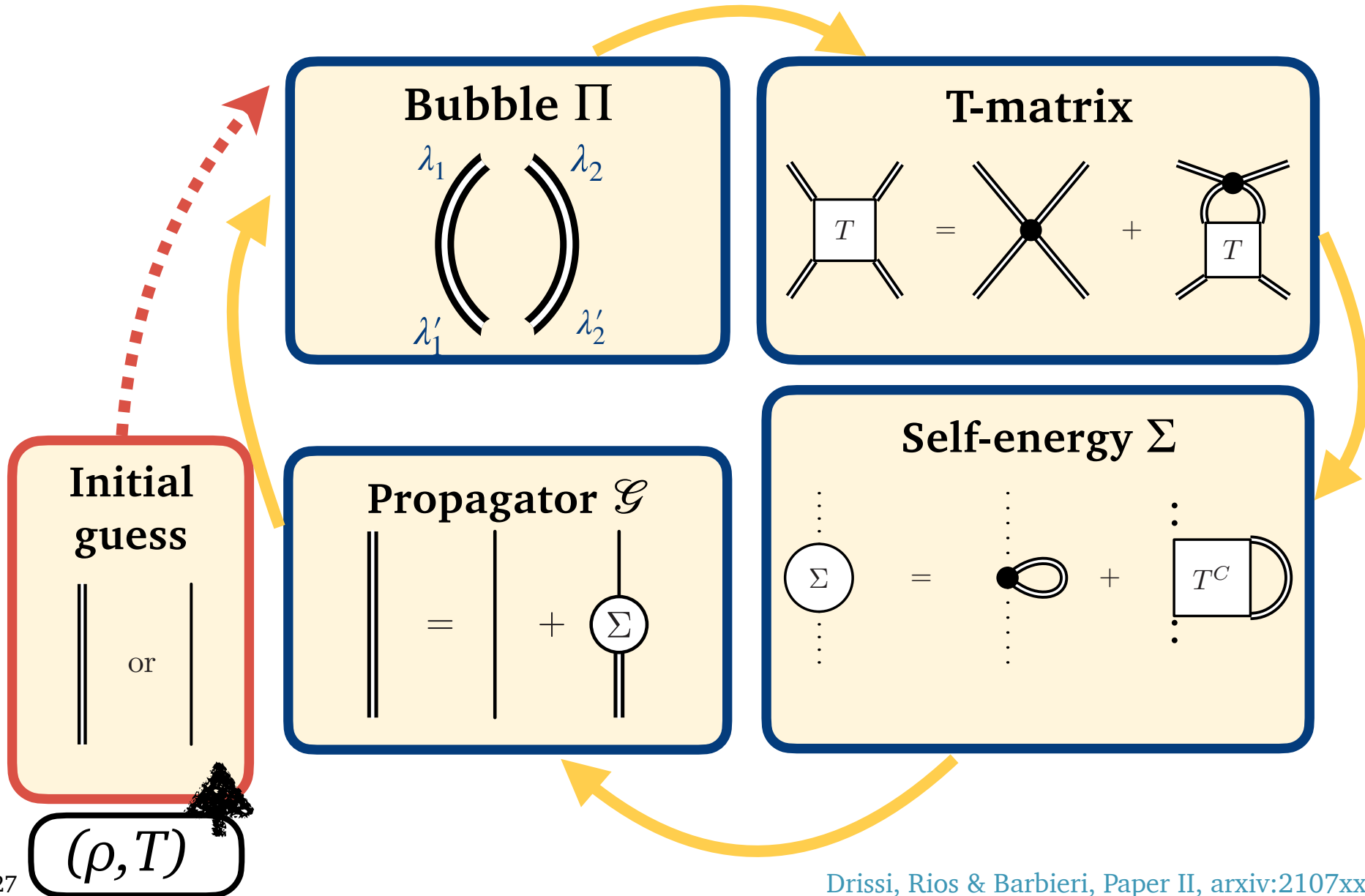
Propagator \mathcal{G}

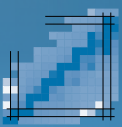
Self-energy Σ



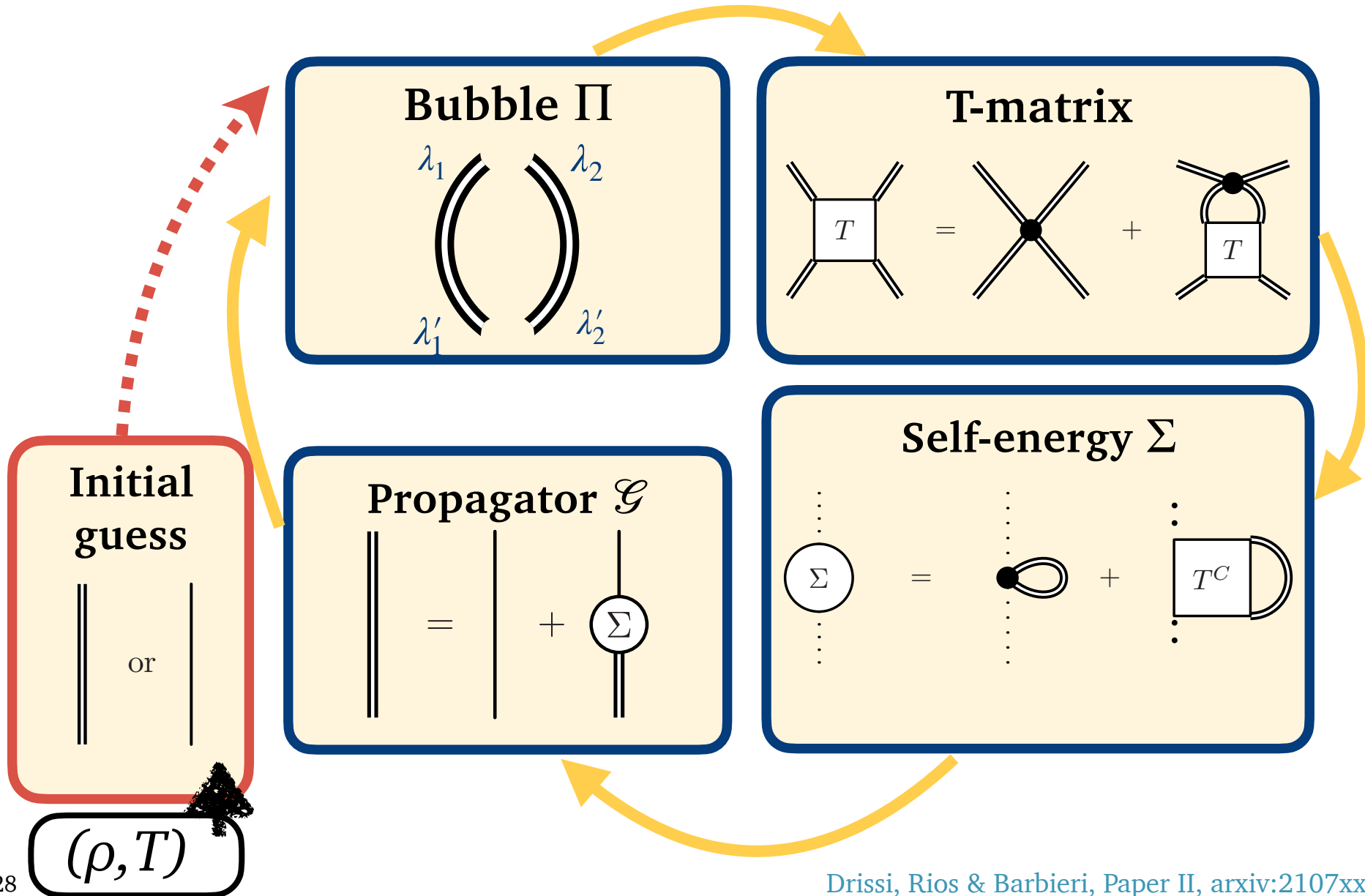


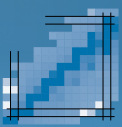
Nambu-Covariant Ladders



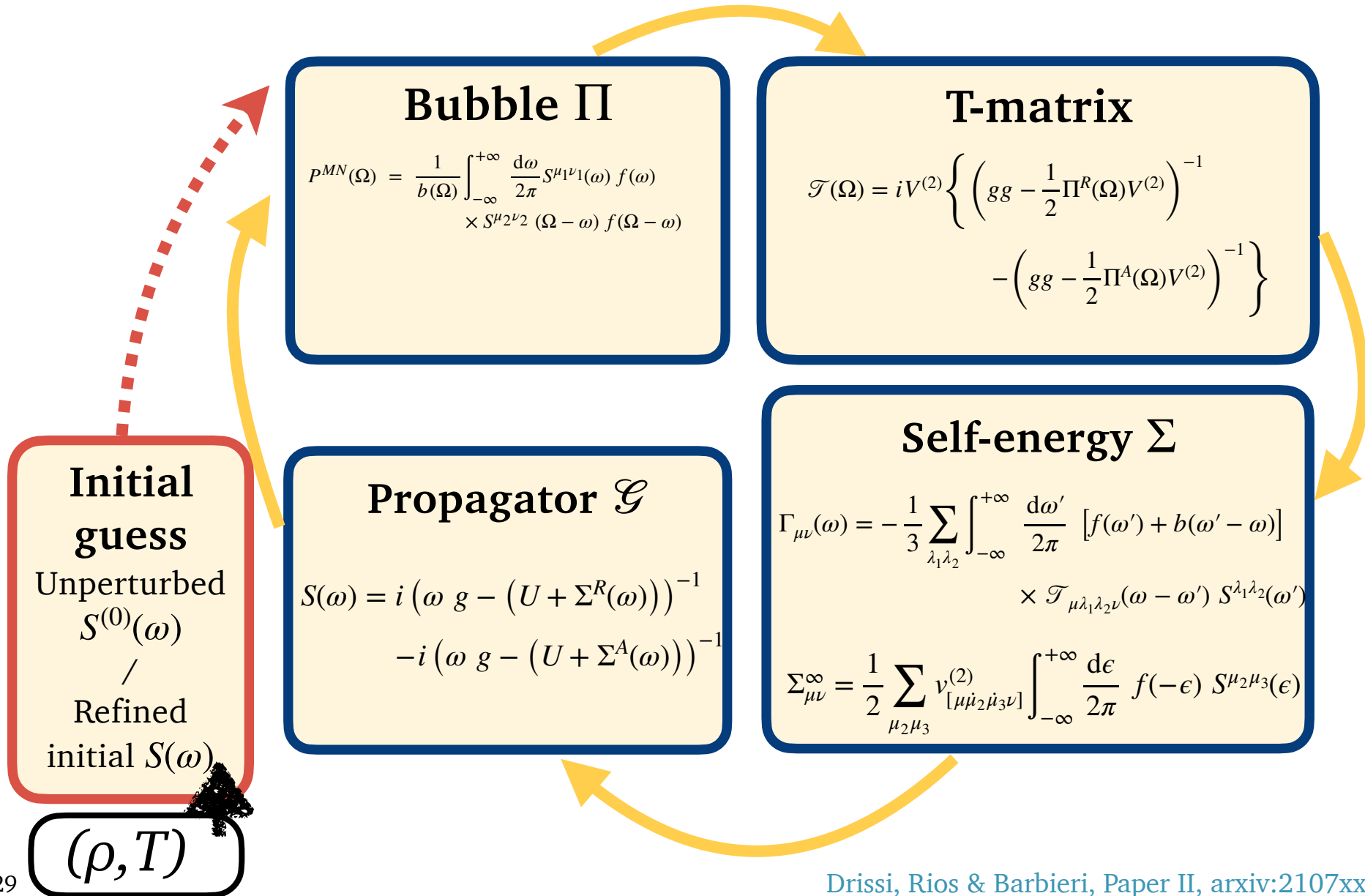


Nambu-Covariant Ladders





Nambu-Covariant Ladders



Initial guess
Unperturbed $S^{(0)}(\omega)$
/
Refined initial $S(\omega)$

(ρ, T)

Bubble Π

$$P^{MN}(\Omega) = \frac{1}{b(\Omega)} \int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} S^{\mu_1 \nu_1}(\omega) f(\omega) \times S^{\mu_2 \nu_2}(\Omega - \omega) f(\Omega - \omega)$$

T-matrix

$$\mathcal{T}(\Omega) = iV^{(2)} \left\{ \left(gg - \frac{1}{2} \Pi^R(\Omega) V^{(2)} \right)^{-1} - \left(gg - \frac{1}{2} \Pi^A(\Omega) V^{(2)} \right)^{-1} \right\}$$

Propagator \mathcal{G}

$$S(\omega) = i \left(\omega g - (U + \Sigma^R(\omega)) \right)^{-1} - i \left(\omega g - (U + \Sigma^A(\omega)) \right)^{-1}$$

Self-energy Σ

$$\Gamma_{\mu\nu}(\omega) = -\frac{1}{3} \sum_{\lambda_1 \lambda_2} \int_{-\infty}^{+\infty} \frac{d\omega'}{2\pi} [f(\omega') + b(\omega' - \omega)] \times \mathcal{T}_{\mu \lambda_1 \lambda_2 \nu}(\omega - \omega') S^{\lambda_1 \lambda_2}(\omega')$$

$$\Sigma_{\mu\nu}^{\infty} = \frac{1}{2} \sum_{\mu_2 \mu_3} v_{[\mu \mu_2 \mu_3 \nu]}^{(2)} \int_{-\infty}^{+\infty} \frac{d\epsilon}{2\pi} f(-\epsilon) S^{\mu_2 \mu_3}(\epsilon)$$

1) Normal-phase SCGF

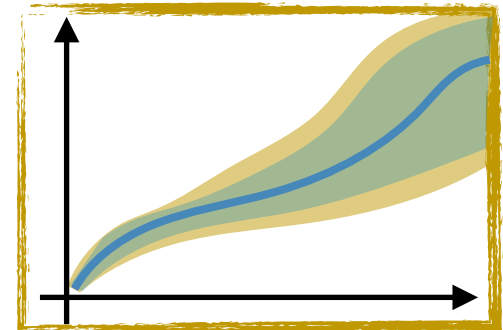
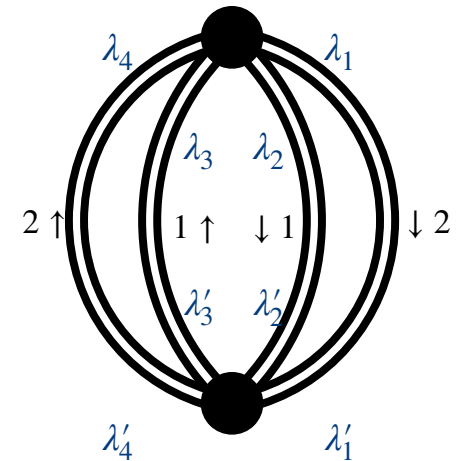
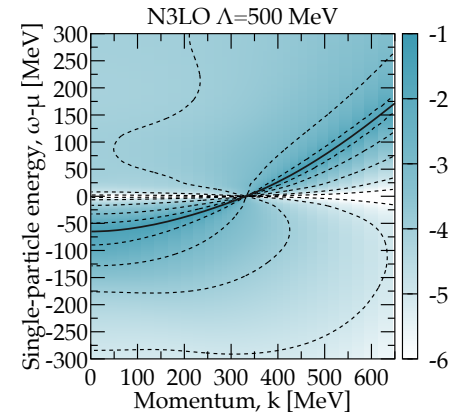
- Spectral strength is available
- EoS & thermodynamics too!

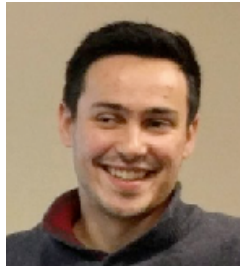
2) Nambu-covariant SCGFs

- Formally relevant
- Perturbative expansion simplified
- Allows for different approximation schemes

Next:

- Numerical implementation
- Uncertainties in predictions?





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 DI MILANO

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(@TRIUMF from 10/2021)

Thank you!

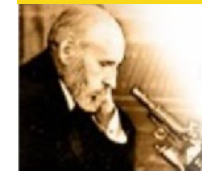
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