





EUROPEAN CENTRE FOR THEORETICAL STUDIES IN NUCLEAR PHYSICS AND RELATED AREAS

TRENTO, ITALY

Institutional Member of the European Expert Committee NUPECC



Microscopic description of fission for the r-process in neutron-star mergers ejecta J.-F. Lemaître Institut d'Astronomie et d'Astrophysique - Brussels University

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Castello di Trento ("Trint"), watercolor 19.8 x 27.7, painted by A. Dürer on his way back from Venice (1495). British Museum,

Nuclear and astrophysics aspects for the rapid neutron capture process in the era of multimessenger observations

Trento, July 1-5, 2019





Understand the structure of solar r-process abundances



- From NS to r-process elements - Fission path & fission barrier

- Fission in nucleosynthesis
- Fission process

- Fission fragments SPY model

From NS to r-process elements

- **inspiral phase** (~ millions of years) gravitational-wave emission

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- coalescence phase/dynamical ejecta (~ ms)

 $0.001 - 0.02 \text{ M}_{\odot}$ nuclear matter are ejected from NS merger $(1.35M_{\odot}x2)$ initially cold, gets shock-heated during the ejection $T > 10^{10}$ K (1MeV) (see A. Bauswein's talk)

- nucleosynthesis in ejected matter

starts when T<10¹⁰ K & ρ < 4.2 10¹¹ g/cm³ (drip density) many neutron captures

 ς neutron capture > β decay

move away from the valley of stability

 \hookrightarrow neutron capture + β decay

production of heavy neutron rich nuclei ⁸/₇







Fission in nucleosynthesis

- Fission of heavy elements plays a fundamental role
- **Recycling the matter** during the neutron irradation

$$(n,\gamma) + \beta + (\gamma,n) + \beta dn + \beta d2n + \beta d3n - \frac{2-3 \text{ cycles } (0.1 \rightarrow 1 \text{ s})}{\text{fission}}$$

$$(n,f) + \beta df \stackrel{\text{heavy neutron}}{\text{rich nuclei}}$$



Fig. 13. (Color online) time evolution of the neutron density N_n and average mass number $\langle A \rangle$ for a specific representative mass element ejected from the 1.35–1.35 M_{\odot} NS merger.



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- Neutron density falls at t ~ 1s no more neutron captures
- Fission shapes the final r-abundance distribution in 110 ≤ A ≤ 170 mass region : 1→10 s (heating dominated by βdf & sf)

- β decay for t > 10 s

- Nuclear flow & abundances are affected by fission barriers & fission fragments distribution



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Fig. 14. (Color online) time evolution of the total massaveraged radioactive heating rate per unit mass $Q_{\rm tot}$ (black solid line) and its β -decay Q_{β} (dashed line) and fission $Q_{\rm f}$ (dotted line) contributions for the matter ejected from the 1.35– $1.35M_{\odot}$ NS merger. The $Q_{\rm f}/Q_{\rm tot}$ ratio is also given (dashdotted line; right axis).

S. Goriely, EPJ A (2015) 51:22 5/33









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Fission process modelling : TALYS, PES (D1M+ATDHF), SPY (BSk27/D1M)



Fission path & fission barrier Potential Energy Surface

PES : constrainted HFB Gogny D1M (Q20,Q30) + triaxial correction + beyond mean-field correction (ATDHF)

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Fission path & fission barrier Fission barrier (D1M)

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Fission path & fission barrier Least action path

U236 ; shortest-path tree : red edged white arrows, LAP in black



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Fission path & fission barrier

Spontaneaous fission lifetimes

$$T_{1/2}^{sf} = 2.86 \ 10^{-21} (1 + e^{2S(E)/\hbar}) \ [s]$$

- Experimental half-lifes are fairly well reproduced
- Fission half-life are very sensitive to excitation energy





Fission path & fission barrier $Q_{\beta} - S_n - B_f \& sf - \beta (D1M)$



Fission path & fission barrier U235(n,f) cross section



Fission fragments - SPY model presentation

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What is the role of the nuclear structure of fission fragments during the fission process ?

Can experimental data be understood/reproduced considering <u>only the</u> <u>nuclear structure of the fission fragments</u> ?



Fission fragments - SPY model a scission point model

- **Hypo. to determine the frag. properties** : fission process (CN \rightarrow frag.) \approx scission line **Scission configuration** : defined by the proton density at scission neck between 2 frag. Fragments are at rest (no prescission kinetic energy)
- Fragments are axially symmetric

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Inputs : frag. Eind, spl & proton density from HFB calculations (Gogny or Skyrme)







- \rightarrow ONLY based on fission fragments & first-chance fission
- \rightarrow Evolution (quasi static) between saddle point to scission point is neglected
- \rightarrow Isolated fragments
- \rightarrow Well defined fragments characteristics (Z, N, β)
- \rightarrow Fragmentation probability \propto number of available states
 - **4** Fragments observables





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 - available energy for each fragmentation of the system : AE
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- 235 U(n_{th},f) : ~500 fragmentations
 - fragmentation → 57 x 57 deformations
 L Ecoul : the most time-consuming numerically computed

 \mapsto AE \approx 20 MeV \rightarrow 20 AS/fragmentation

1,6 million AE

32 million AS

Fission fragments - SPY model available energy & available states

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Fission fragments - SPY model Fission of U236, Pu240 & Cf252

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Fission fragments - SPY model Fission of « actinides »

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Fission fragments - SPY model Peak multiplicity

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Fission fragments - SPY model Doubly asymmetric fission & abundances

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Conclusions & outlooks



Fission path

- Fission barriers for 500 e-e nuclei (J.-F. Lemaître et al, PRC 98, 024623 (2018))
- Half lives in good agreement with experimental data
- * To be extended to odd nuclei
- * LEP/LAP \rightarrow Fission cross sections
- * R-process in full Gogny D1M framework

Fission fragments

- Scission point, static frag., statistical (microcanonical description)
- Definition of the scission point based on realistic proton distribution
- All ingredients are calculated coherently in the same microscopic framework (Skyrme BSk27 eff. N-N interaction ; J.-F. Lemaître et al, Phys Rev C 99, 034612 (2019))
- Applied to the r-process, doubly asymm. fission (S. Goriely et al, Phys. Rev. Lett. 111, 242502 (2013))
- * Improve the description of the kinetic energy, the neutron evaporation
- * Octupole deformations
- * New version with Gogny-D1M eff. N-N interaction \rightarrow link with PES