





Analogue Kerr black hole and Penrose process

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Plan

- Analogue physics with quantum vortices
- 2D acoustic Kerr black hole
- Penrose process
- Conclusion

Analogue physics with vortices

- Maxwell 1861:
 - Magnetic field by analogy with fluid dynamics
- Popov 1973:
 - 2D quantum vortices mapped to 3D magnetostatic problem
 - 2+1 relativistic electrodynamics:
 - vortices = charges
 - phonons = photons
- Other works:
 - 2+1 QED with vortices (Arovas, 1997), Maxwell's equations
 - Cosmic strings (Volovik, 1995)
 - Half-vortices as magnetic monopoles (Solnyshkov, 2012)

Quantum vortex

- Quantum fluid described by a wave function
- Phase gradient determines velocity $\nabla \times \mathbf{v} = 0$
- Velocity curl zero everywhere
- Except zero-density points quantum vortices
- Quantized circulation (phase winding) $\oint \nabla \varphi \cdot \mathbf{dl} = 2\pi n$



Wave function of a single vortex:

$$\psi \approx \frac{r / \xi}{\sqrt{2 + (r / \xi)^2}} e^{i\phi}$$

V.N. Popov, JETP 37, 341 (1973)

Quantum vortex as a relativistic charge

- Action for a vortex = action for a relativistic charge $S = -m_0 c \int ds + \int \frac{e}{c} \frac{dx_{\alpha}}{ds} A^{\alpha} ds$
- Relativistic particle in an electromagnetic field F

$$\frac{d^2 x^{\mu}}{ds^2} = \frac{q}{m_V} F^{\mu\beta} \frac{dx^{\alpha}}{ds} \eta_{\alpha\beta}$$

• In curved spacetime:

$$\frac{d^{2}x^{\mu}}{ds^{2}} = -\Gamma^{\mu}_{\alpha\beta} \frac{dx^{\alpha}}{ds} \frac{dx^{\beta}}{ds} + \frac{q}{m_{V}} F^{\mu\beta} \frac{dx^{\alpha}}{ds} g_{\alpha\beta}$$
Time-like geodesics

The motion of the vortex is described by the acoustic metric!

Christoffel symbols: $\Gamma^{\mu}_{\alpha\beta} = 0.5 g^{\mu\nu} (\partial g_{\nu\alpha} / \partial x^{\beta} + \partial g_{\nu\beta} / \partial x^{\alpha} - \partial g_{\alpha\beta} / \partial x^{\nu})$

Kerr black hole features

- Rotating black hole
- Maximal angular momentum *a*/*M*=1
- Strong frame dragging
- Two particular surfaces
 - Horizon (light cannot get out)
 - Static limit (light cannot go against the rotation)
- Penrose process (extraction of the rotation energy)
- Ergosphere

PRD 70, 124006 (2004).

Kerr metric for a condensate in 2D

Equatorial plane only!



Condensate wavefunction



Asymptotic series expansion at large r



Parameters typical for exciton-polariton condensates

Analogue Kerr BH parameters

• Comparing the metric element g_{rr} we obtain

 $M_{cond} \sim r_H \sim \zeta \xi$ Mass is controlled by the drain

• Maximal number of vortices in an analogue BH

$$v_{\max} \sim \frac{r_H}{\xi}, a_{\max} \sim r_H$$

• Maximal angular momentum $\frac{a_{\max}}{M} \sim 1$

Vortices inside the BH are distributed along the horizon

Penrose process

- A particle p falls into the ergosphere
- It splits into two (p' and p'')
- p" falls into the BH
 - p" had negative energy!
- p' escapes
- Negative energy does not mean an antiparticle
- One needs more than mc² to get the particle away from BH

E.F. Taylor, Exploring Black Holes



Penrose process





- A vortex-antivortex pair is formed from a density dip
- The anti-vortex falls into BH
- The vortex escapes to infinity
- BH rotation decreases (energy loss)

-50

-50

Penrose process: snapshots

50

b)

AV





- V/AV interaction slows the AV
- AV rotates *slower* than the condensate
- E_{AV}<0
- V gains energy from AV and escapes

Vortex trajectories and time-like geodesics of the Kerr metric



Both metrics dominated by the divergent term $g_{rr} \sim (r - r_H)^{-1}$

Phase of the condensate



Outlook

- Dynamical metric
 - The angular momentum is not fixed externally
- Natural presence of quantum fluctuations
 - Towards quantum gravity
 - Comparable scales of quantum and gravitational effects
- Control of quantum fluctuations
 - Interactions
 - Particle mass

$$\frac{n^{(1)}(0) - n^{(1)}(\infty)}{n^{(1)}(0)} \sim \frac{\alpha m}{\hbar^2}$$

• Thermal fluctuations negligible

$$n^{(1)}(s) \sim \left(\frac{s_T}{s}\right)^{\nu}, \nu = \frac{k_B T m}{2\pi \hbar^2 n_s}$$

Conclusions

- Quantum vortices relativistic massive test particles
- Analogue electrodynamics in curved spacetimes

More info: PRB 99, 214511 (2019)