# Scattering of co-current surface waves on an analogue black hole

Based on: arXiv:1806.05539 (2018/19)



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## Analogue black holes

Unruh (1981): on large scales, waves in moving media propagate as if in curved spacetime

fixed background

 $\phi = \phi_0 + \delta \phi$ 

 $\phi_0 \qquad g_{\mu\nu}(\phi_0) dx^{\mu} dx^{\nu} = c^2 dt^2 - (d\bar{x} - \bar{v} dt)^2$ • linear perturbations  $\delta \phi$   $g^{\mu\nu}(\phi_0) \nabla_{\mu} \nabla_{\nu}(\delta \phi) = 0$ 





Picture courtesy of Piotr Pieranski

#### The surface wave analogy

(Unruh and Schützhold, 2002)



**Total wave speed** =  $v \pm c$ 

 $\delta h \leftrightarrow$  canonical momentum in the (2+1)-d spacetime metric  $g_{\mu\nu}dx^{\mu}dx^{\nu} = c^{2}\left[c^{2}dt^{2} - (dx - \nu dt)^{2} - dy^{2}\right]$  *"conformal factor"*, multiplies simpler metric generates effective potential *Veff* **allows mixing between co- and counter-propagating waves** 

## Wave scattering (subcritical flows)



## Wave scattering (transcritical flows)



## Wave scattering (transcritical flows)

![](_page_5_Figure_1.jpeg)

## Wave scattering (transcritical flows)

![](_page_6_Figure_1.jpeg)

dispersive wave x

#### Key surface wave experiments in AG

## All previous experiments (in 1 dimension) have been in purely subcritical flows

Rousseaux et al., *New J. Phys.* **10**, 053015 (2008) Weinfurtner et al., *Phys. Rev. Lett.* **106**, 021302 (2011) Euvé et al., *Phys. Rev. Lett.* **117**, 121301 (2016)

→ No horizon in the effective metric

➤ No thermality in the sense of Hawking/Unruh

#### (Nontrivial scattering allowed thanks to **dispersion**)

(Michel and Parentani, *Phys. Rev. D* **90**, 044033 (2014)) (Robertson, Michel and Parentani, *Phys. Rev. D* **93** 124060 (2016))

![](_page_8_Figure_0.jpeg)

#### The current Poitiers setup

![](_page_9_Picture_1.jpeg)

#### Scattering of incident probe

![](_page_10_Figure_1.jpeg)

#### Example data

 $A_{\rm wm} = 0.25 \,\rm mm$  $\omega = 3.46 \,\rm Hz$ 

![](_page_11_Figure_2.jpeg)

#### **Dispersion relation**

![](_page_12_Figure_1.jpeg)

#### **Dispersion relation**

![](_page_13_Figure_1.jpeg)

Full theoretical dispersion relation:

$$\omega - vk = \pm \sqrt{gk \tanh(hk)}$$

- Observations lie close to theoretical dispersion relations
- Close to the linear (non-dispersive) regime, so metric description valid
- Counter-propagating branch has flipped sign, indicating transcriticality and (indirectly) the existence of negative-energy waves

## Fitting the waveforms

FT in time at wave maker frequency, yielding real and imaginary parts of  $\delta h_{\omega}(x)$ 

Fit to sum of two plane waves, separately in sub- and supercritical regions

![](_page_14_Figure_3.jpeg)

### Dispersion relation

#### Latest results

Found by FT'ing in time then fitting to sum of two plane waves

Different colours represent different wave maker amplitudes: 0.25 mm, 0.5 mm, 1 mm

Allows fitting of *v* and *c* 

- very close to values inferred from depth in subcritical region
- small difference in supercritical region (likely due to presence of vorticity)

![](_page_15_Figure_7.jpeg)

### Scattering coefficients

(i.e., ratios of wave amplitudes to amplitude of incident wave)

#### Latest results

Found by FT'ing in time then fitting to sum of two plane waves

Different colours represent different wave maker amplitudes: 0.25 mm, 0.5 mm, 1 mm

- Reasonable agreement with theory, especially at high frequencies
- Considerable scatter and errors at low freq., especially in supercritical region

![](_page_16_Figure_7.jpeg)

### Scattering coefficients

#### (normalized)

#### Latest results

Found by FT'ing in time then fitting to sum of two plane waves

Different colours represent different wave maker amplitudes: 0.25 mm, 0.5 mm, 1 mm

- Reasonable agreement with theory, especially at high frequencies
- Considerable scatter and errors at low freq., especially in supercritical region
- Unable to verify unitarity:

$$\left|\mathscr{R}\right|^{2} - \left|\mathscr{N}\right|^{2} + \left|\mathscr{T}\right|^{2} = 1$$

![](_page_17_Figure_9.jpeg)

### Summary

**Transcritical black-hole flow** realized (a first in water wave Analogue Gravity), probed by scattering of incident co-current waves off effective potential

Have access to dispersion relation and scattering coefficients

Results reasonably close to predictions of **effective metric** description

But what next...?

#### Future prospects

- Stimulation of Hawking radiation
  - exciting **incident dispersive modes** is technologically challenging
- "Unstimulated" scenario (i.e. without wave maker)
  - incident modes provided by **noise** already present
- Stimulated Hawking radiation in a time-dependent black hole formation (à la Hawking '74-75)

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#### Thank you for your attention!