

# *BSM Physics: Nuclei as Laboratories*

M.J. Ramsey-Musolf

*U Mass Amherst*



<http://www.physics.umass.edu/acfi/>

*My pronouns: he/him/his*

Nuclei as BSM Laboratories  
ECT\* April 2019

## *Key Theme for This Talk*

- *Fundamental questions motivate the search for physics beyond the Standard Model*
- *Tests of fundamental symmetries at low-energy are poised to*
  - *discover the BSM physics that answers several of these questions*
  - *determine its character*
- *Robust hadronic & nuclear computations plus high sensitivity experiments are essential*

## ***Goals For This Talk***

- *Put nuclear tests of fundamental symmetries & neutrino property studies in broader BSM context*
- *Highlight a few areas of inter-frontier implications*
- *Emphasize new directions for theoretical work*

# Outline

*I. The BSM Context*

*Origin of Matter*

*II. Lepton Number*

*III. CP (Flavor Conserving)*

*IV. Precision Tests:  $\beta$ -Decay Highlight*

*V. Outlook*

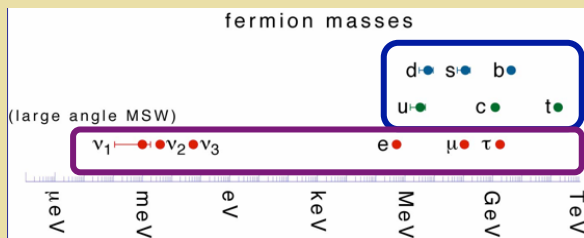
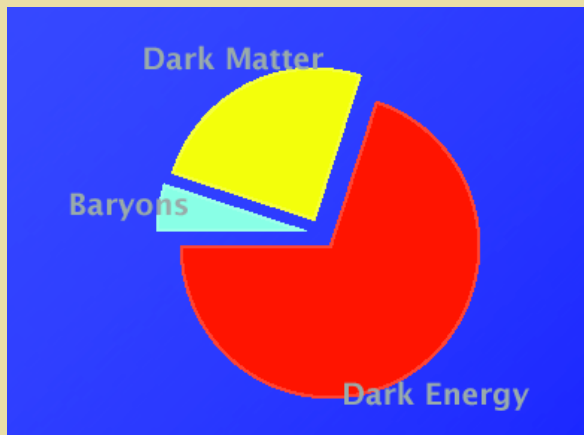
*BSM Mass Scale*



# ***I. The BSM Context***

# Fundamental Questions

**MUST** answer



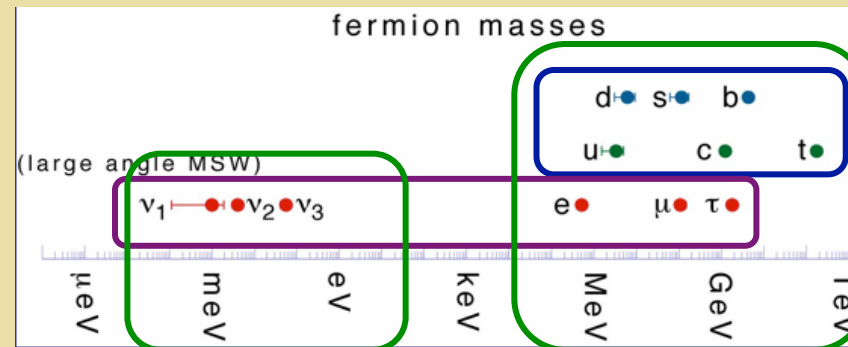
Origin of  $m_\nu$

**SHOULD** answer

A Feynman diagram showing a loop of Higgs bosons ( $H^0$ ) with a new scalar field ( $\phi_{NEW}$ ) interacting with them. The diagram is labeled  $\Delta m^2 \sim \lambda \Lambda^2$ .

$\Lambda_{Cosmological}$

# Fermion Masses & Baryon Asymmetry



*Partners*

*Partners*

*Something else ?*

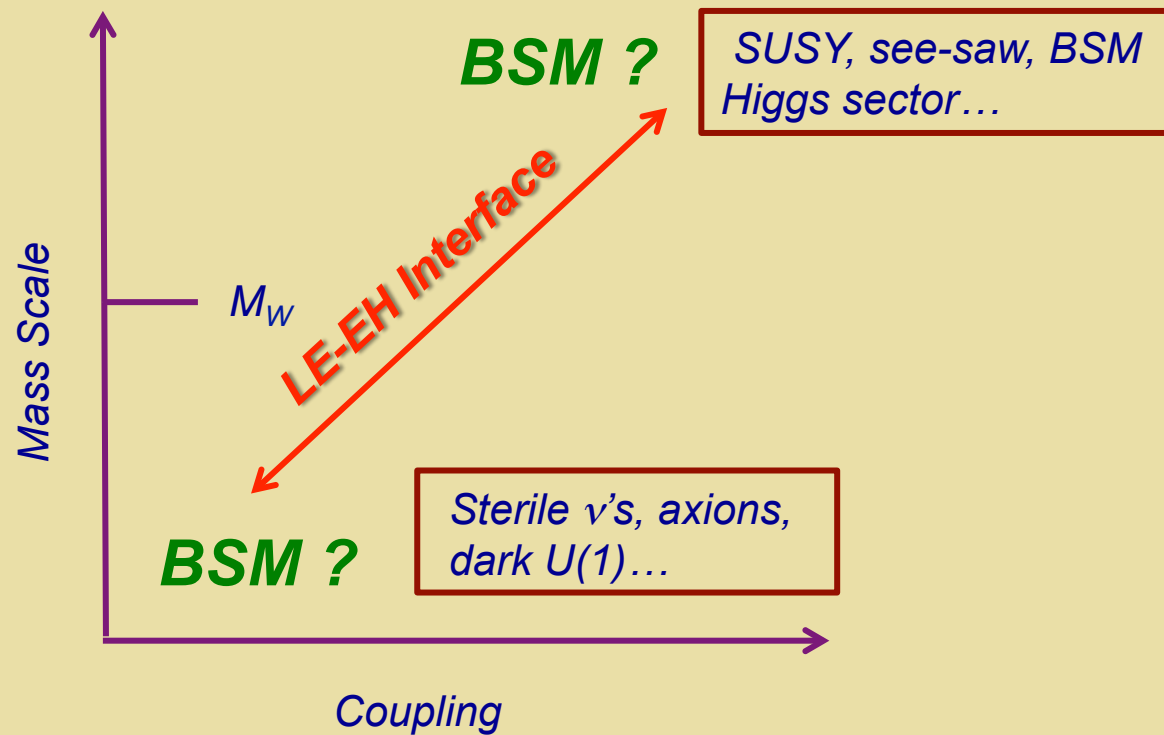
*Higgs Mechanism*

**Leptogenesis: Baryon asymmetry &  $m_\nu$  from lepton number violation**

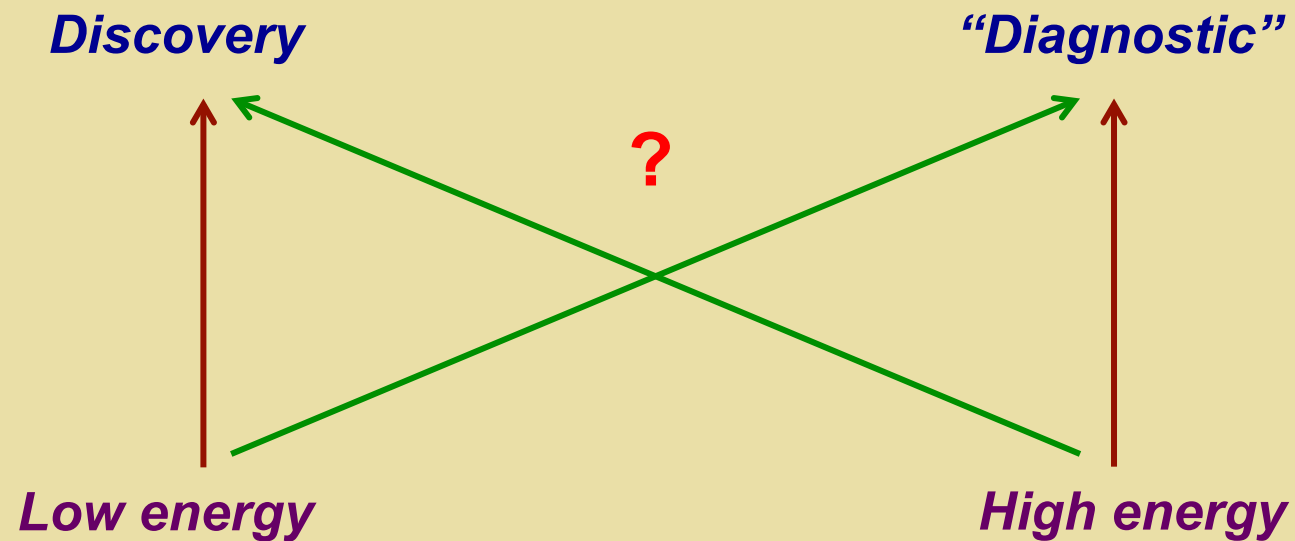
**Electroweak baryogenesis: Baryon asymmetry &  $m_f$  from EW symmetry breaking**

# ***BSM Physics: Where Does it Live ?***

# BSM Physics: Where Does it Live ?



# *Low-Energy / High-Energy Interplay*



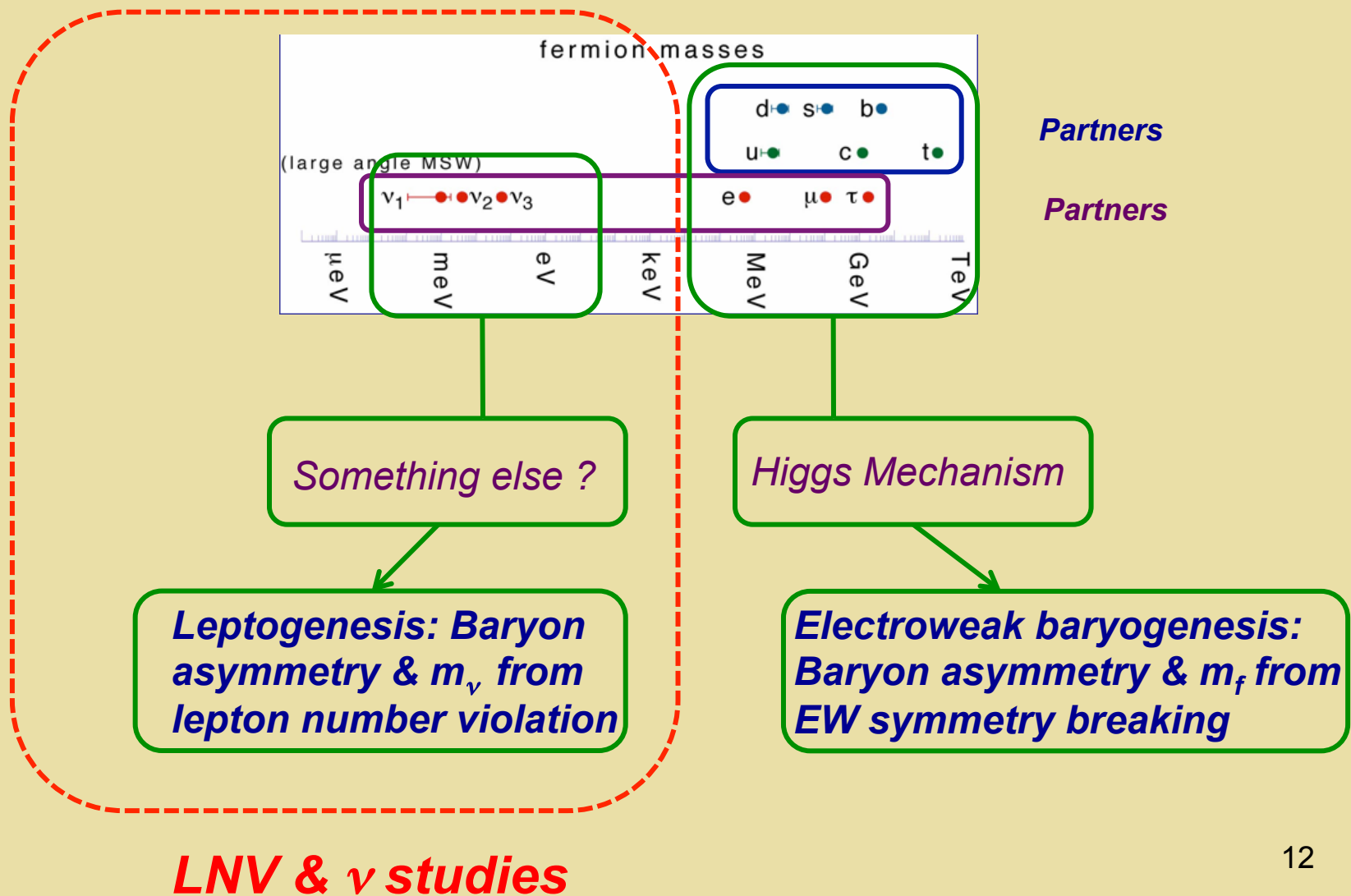
## ***II. Lepton Number***

*Highlight: TeV LNV*

*Collaborators: T. Peng, P. Winslow; J. Harz, S. Quirroga, S. Shen*

*PRD 93 (2016) 093002*

# Fermion Masses & Baryon Asymmetry





# *$0\nu\beta\beta$ -Decay: LNV? Mass Term?*

$$\mathcal{L}_{\text{mass}} = y \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

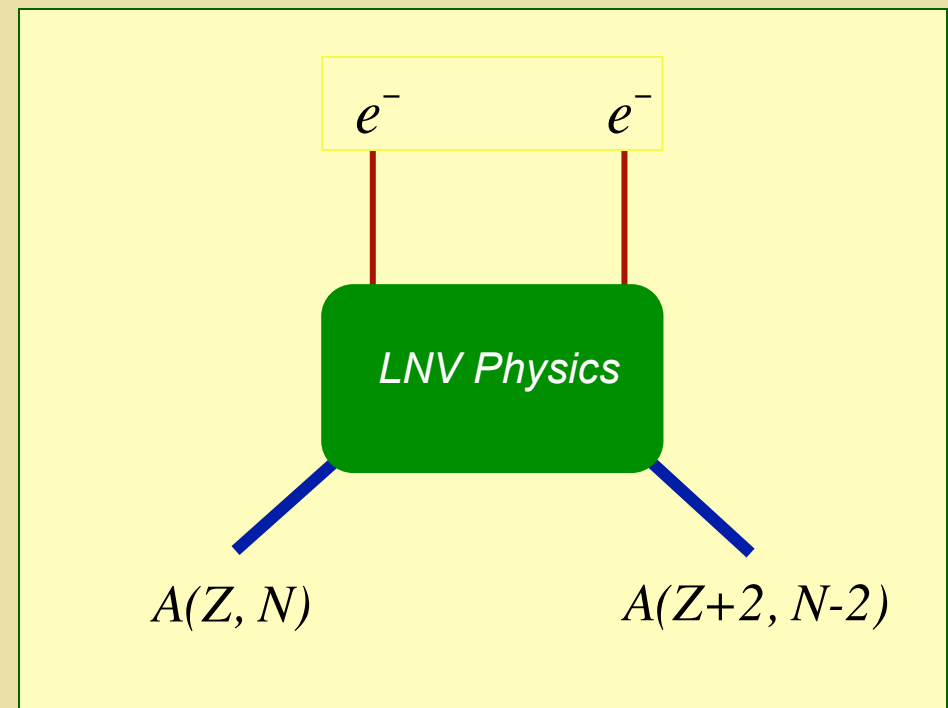
*Dirac*

$$\mathcal{L}_{\text{mass}} = \frac{y}{\Lambda} \bar{L}^c H H^T L + \text{h.c.}$$

*Majorana*

## *Impact of observation*

- Total lepton number not conserved at classical level*
- New mass scale in nature,  $\Lambda$*
- Key ingredient for standard baryogenesis via leptogenesis*



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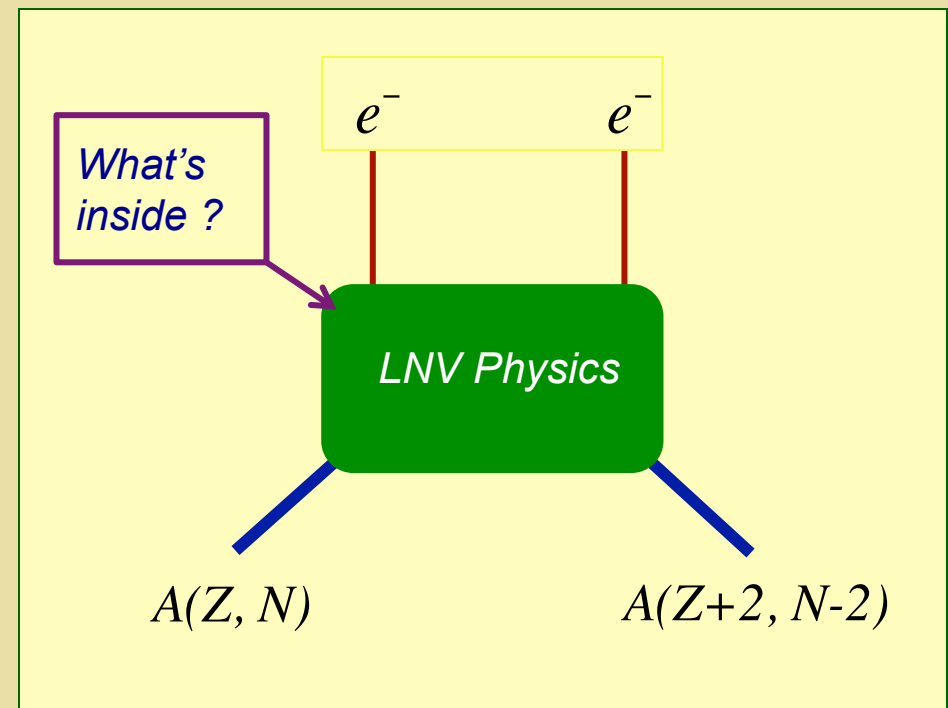
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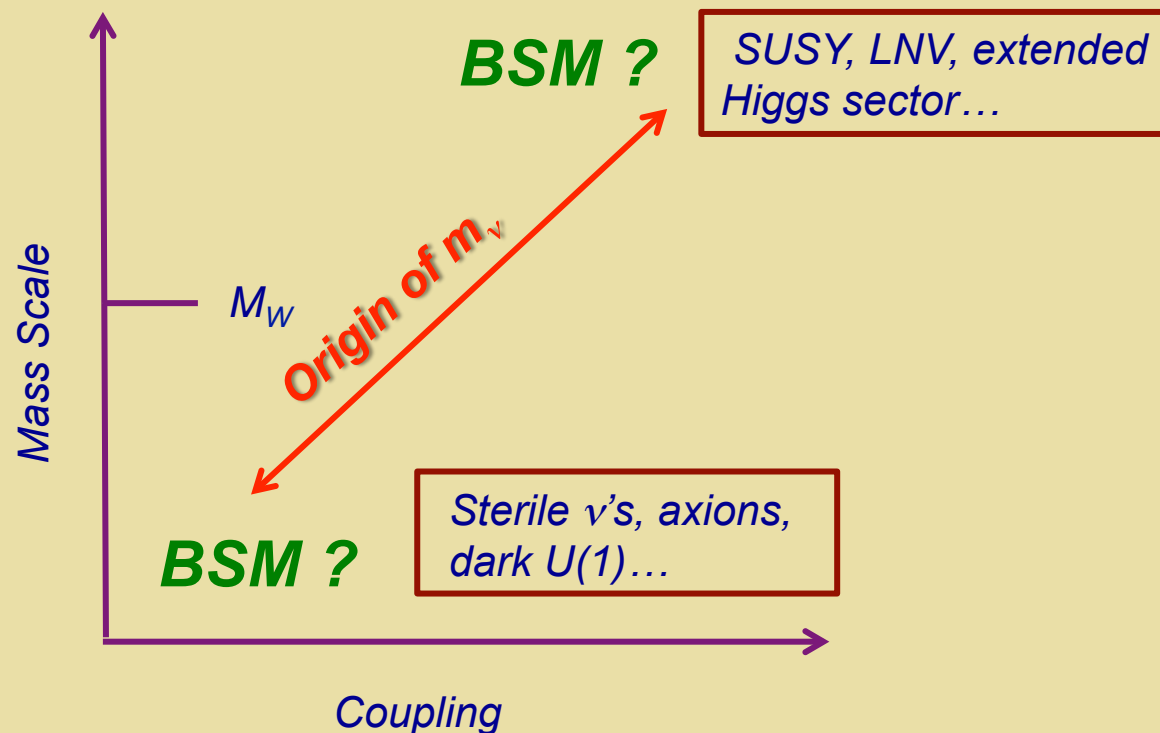
*Majorana*

## *Impact of observation*

- *Total lepton number not conserved at classical level*
- *New mass scale in nature,  $\Lambda$*
- *Key ingredient for standard baryogenesis via leptogenesis*



# BSM Physics: Where Does it Live ?



*Is the mass scale associated with  $m_\nu$  far above  $M_W$  ? Near  $M_W$  ? Well below  $M_W$  ?*

# ***LVN Mass Scale & $0\nu\beta\beta$ -Decay***

$$A(Z,N) \rightarrow \boxed{\text{Underlying Physics}} \rightarrow A(Z+2, N-2) + e^- e^-$$

- *3 light neutrinos only: source of neutrino mass at the very high see-saw scale*
- *3 light neutrinos with TeV scale source of neutrino mass*
- *> 3 light neutrinos*

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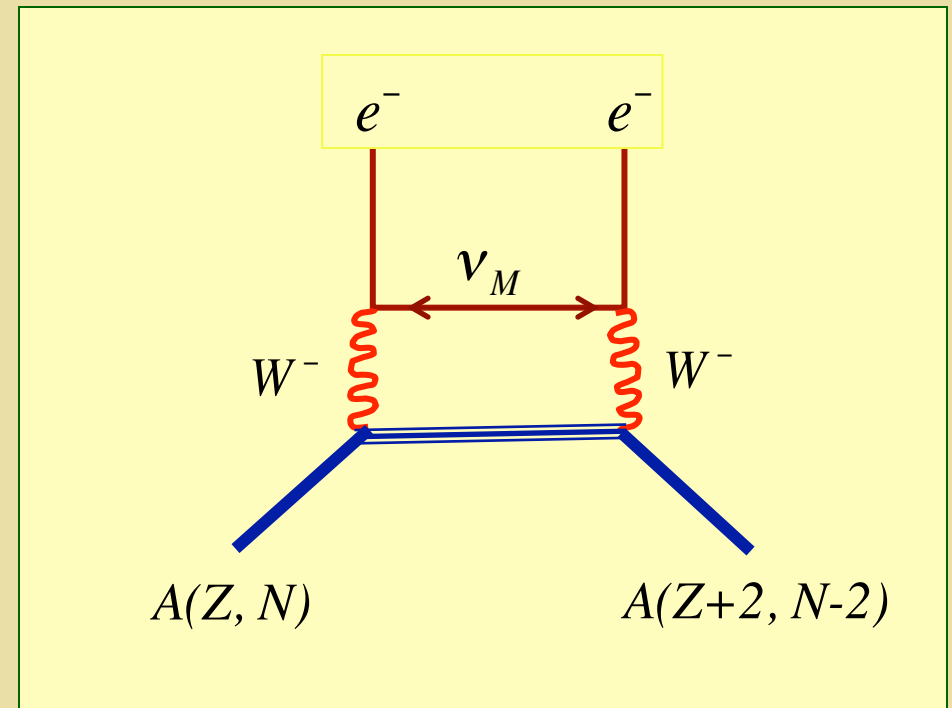
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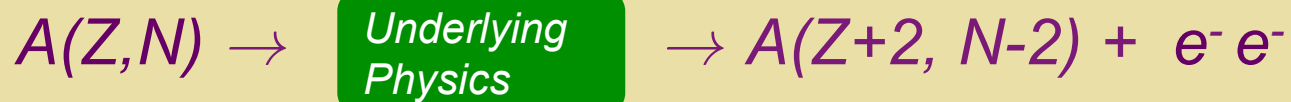
*Majorana*

## ***“Standard” Mechanism***

- Light Majorana mass generated at the conventional see-saw scale:  $\Lambda \sim 10^{12} - 10^{15} \text{ GeV}$*
- 3 light Majorana neutrinos mediate decay process*



# ***LVN Mass Scale & $0\nu\beta\beta$ -Decay***



- *3 light neutrinos only: source of neutrino mass at the very high see-saw scale*
- *3 light neutrinos with TeV scale source of neutrino mass*
- *> 3 light neutrinos*

*Two parameters: **Effective coupling** & **effective heavy particle mass***

# *$0\nu\beta\beta$ -Decay: LNV? Mass Term?*

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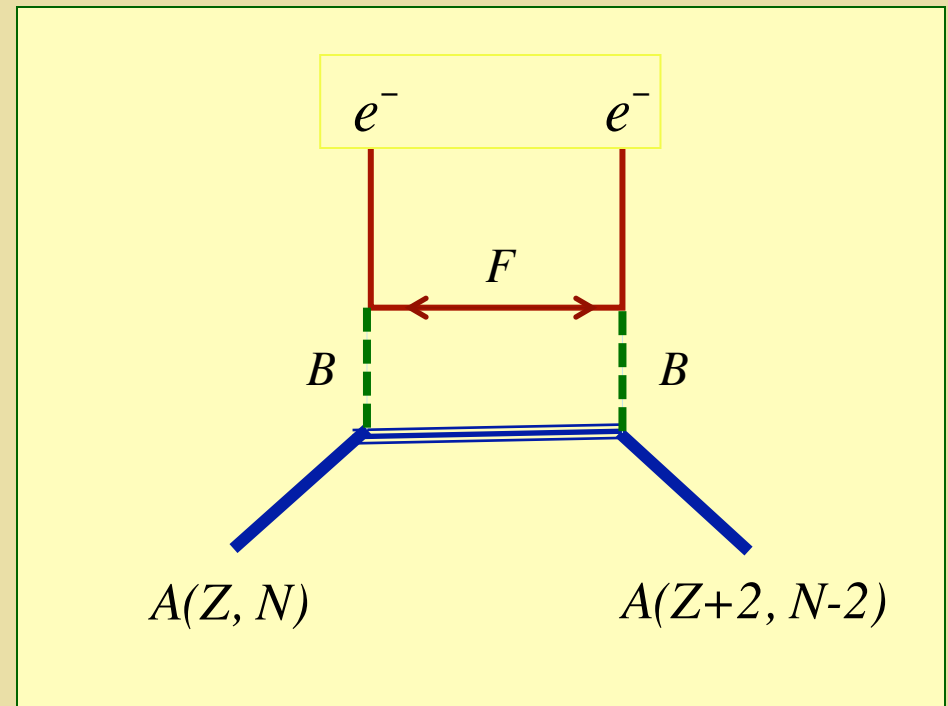
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*Majorana*

## *TeV LNV Mechanism*

- *Majorana mass generated at the TeV scale*
- *Low-scale see-saw*
- *Radiative  $m_\nu$*
- *$m_{\text{MIN}} \ll 0.01 \text{ eV}$  but  $0\nu\beta\beta$ -signal accessible with tonne-scale exp'ts due to heavy Majorana particle exchange*





# *$0\nu\beta\beta$ -Decay: LNV? Mass Term?*

$$\mathcal{L}_{\text{mass}} = y \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

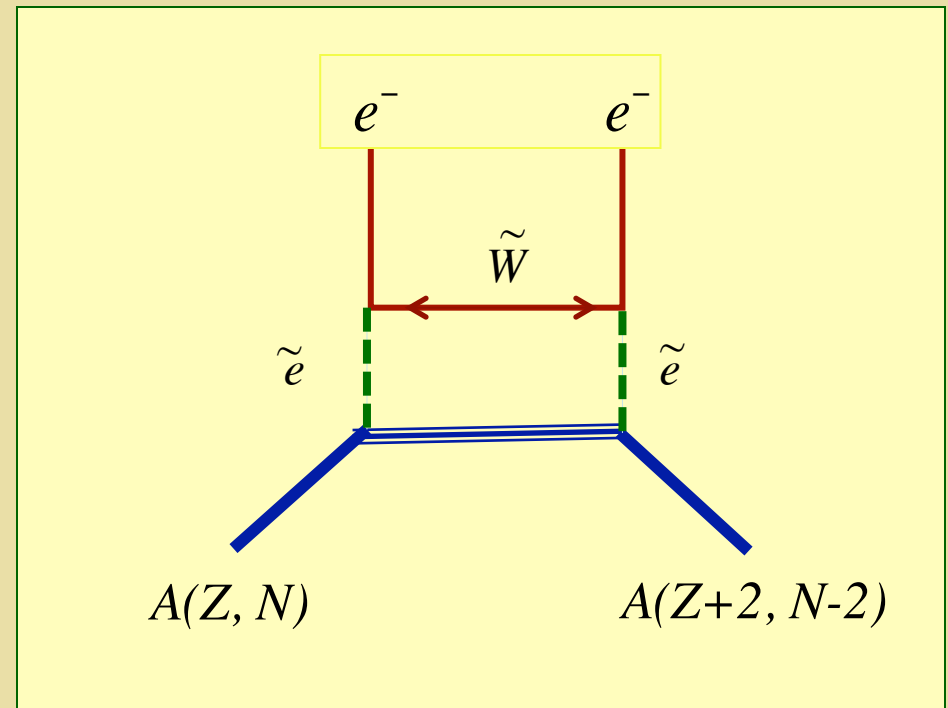
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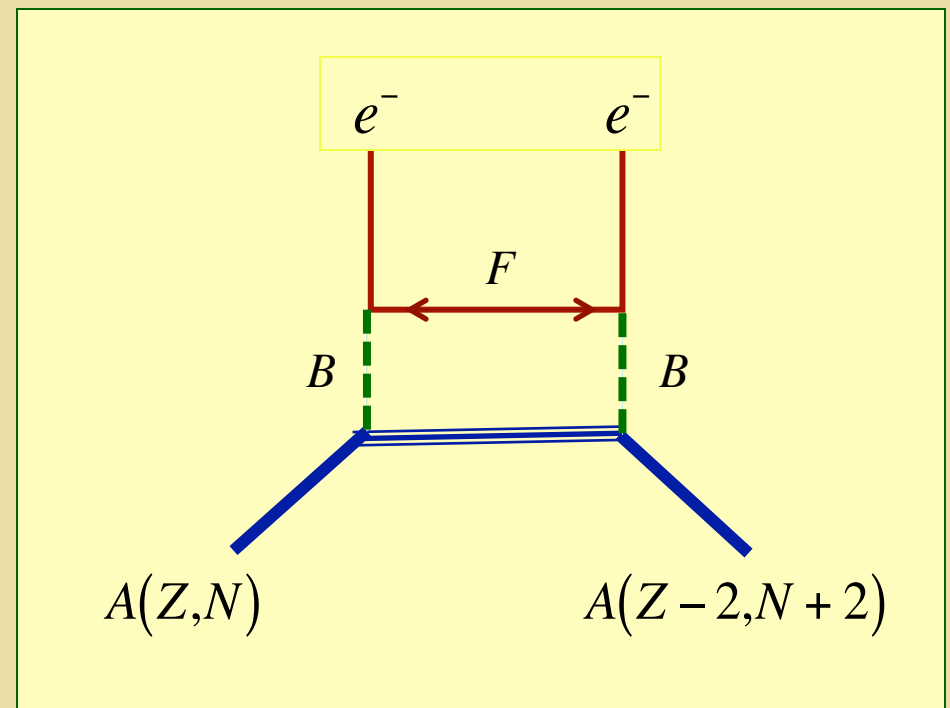
*Majorana*

## *TeV LNV Mechanism*

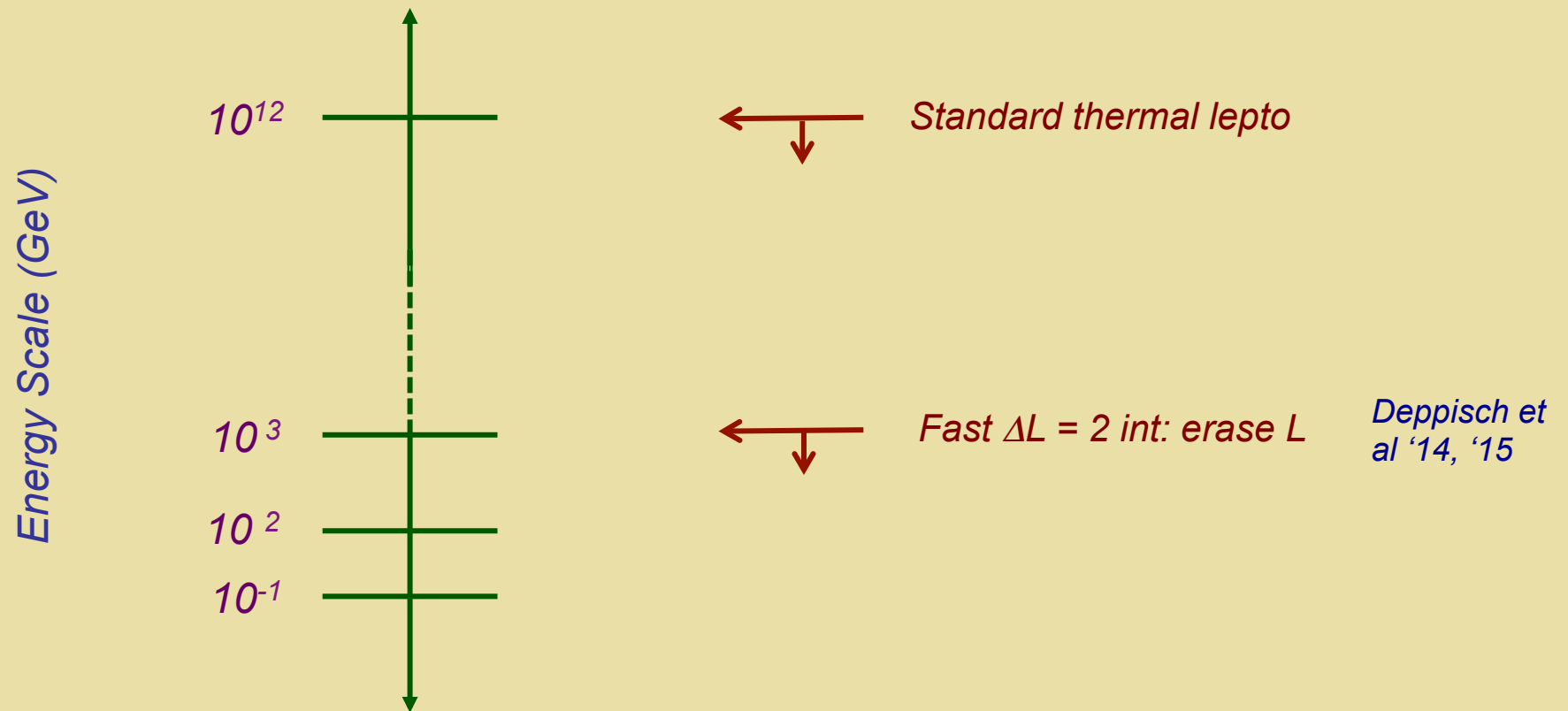
$$\frac{A_H}{A_L} \sim \frac{M_W^4 \bar{k}^2}{\Lambda^5 m_{\beta\beta}}$$

*$O(1)$  for  $\Lambda \sim 1 \text{ TeV}$*

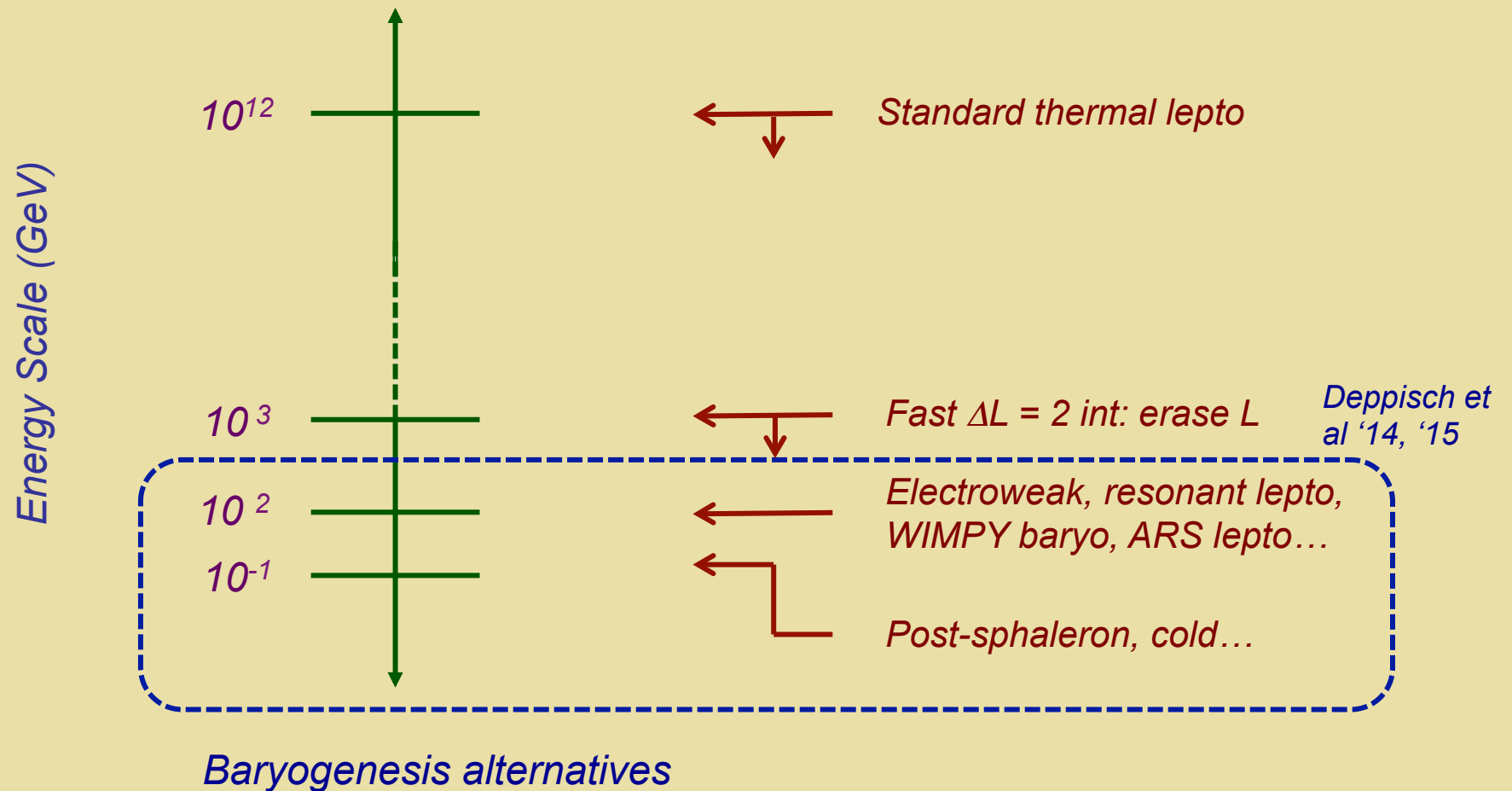
*Implications*



# TeV LNV & Leptogenesis



# TeV LNV & Leptogenesis



# $0\nu\beta\beta$ -Decay: TeV Scale LNV

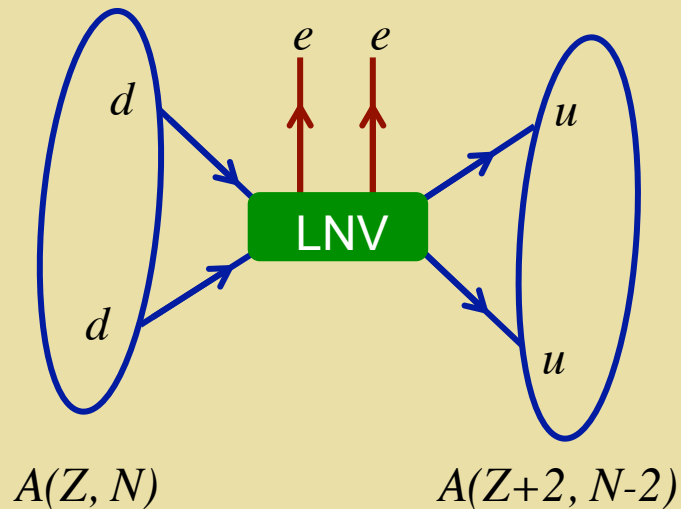
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Dirac

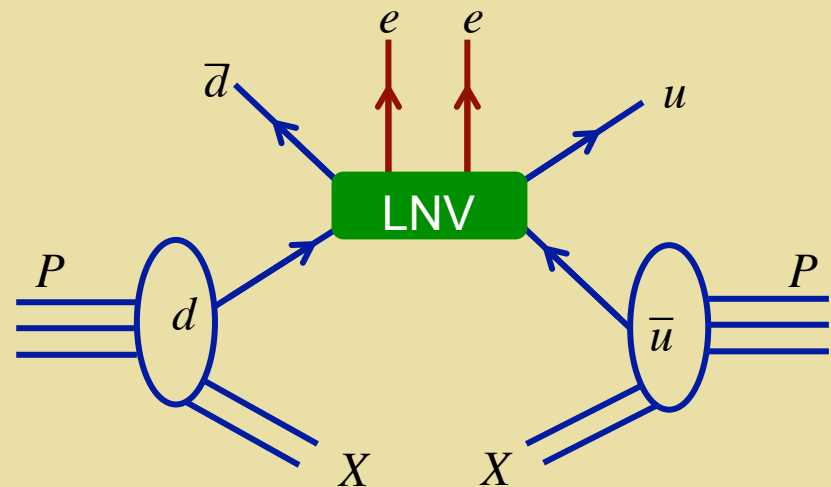
$$\mathcal{L}_{\text{mass}} = \frac{y}{\Lambda} \bar{L} H H^T L + \text{h.c.}$$

Majorana

$0\nu\beta\beta$ -Decay



$pp$  Collisions



# $0\nu\beta\beta$ -Decay: TeV Scale LNV

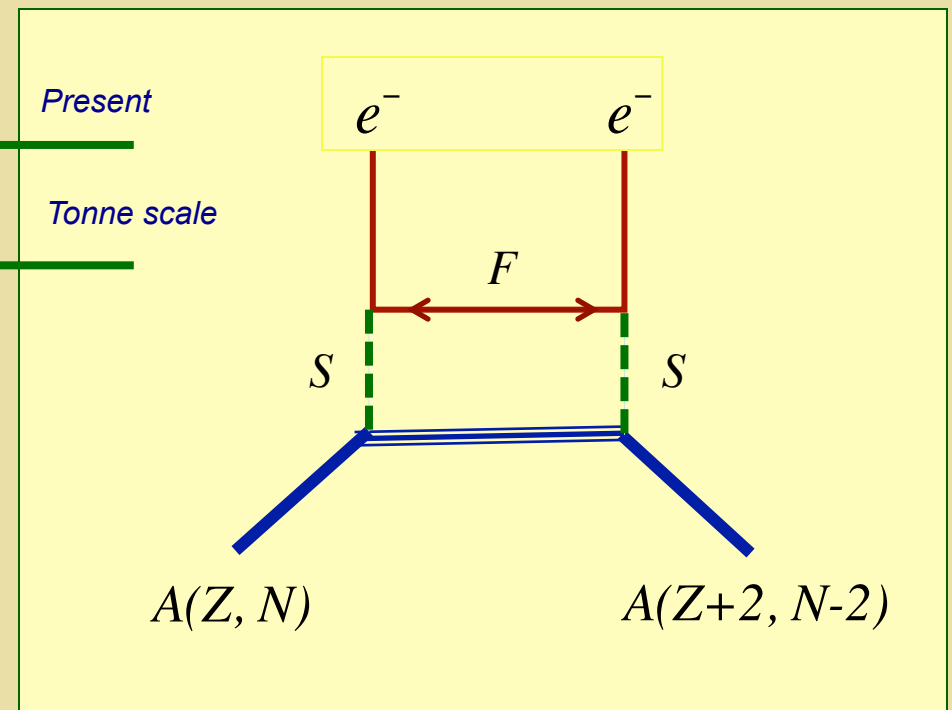
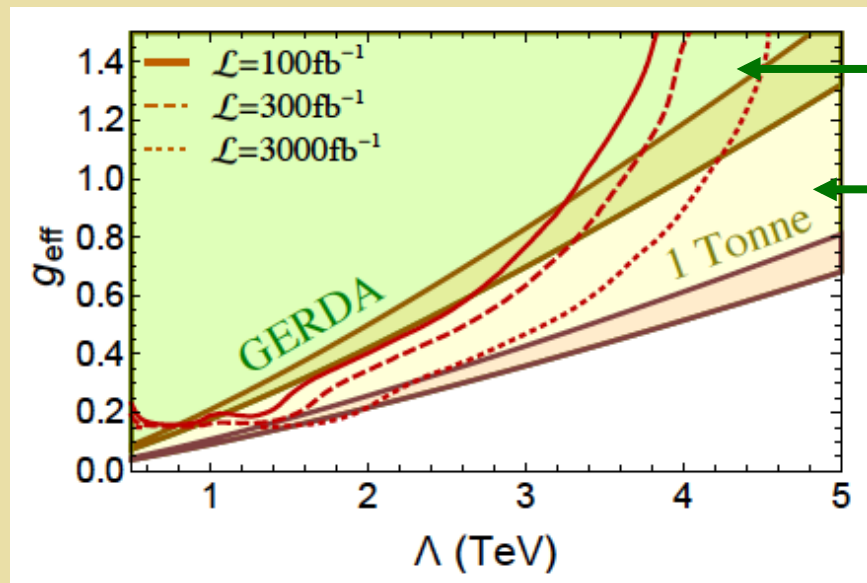
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*Majorana*

## Benchmark Sensitivity: TeV LNV



T. Peng, MRM, P. Winslow 1508.04444

# $0\nu\beta\beta$ -Decay: TeV Scale LNV

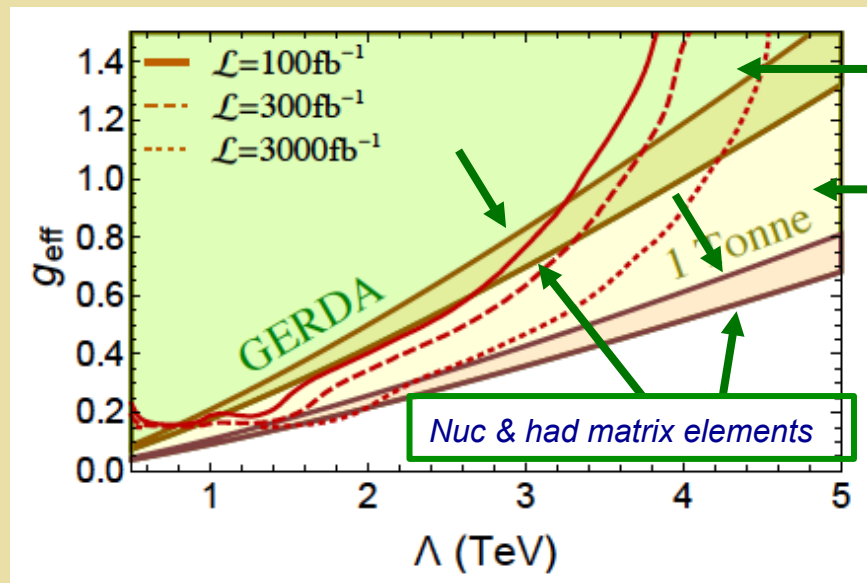
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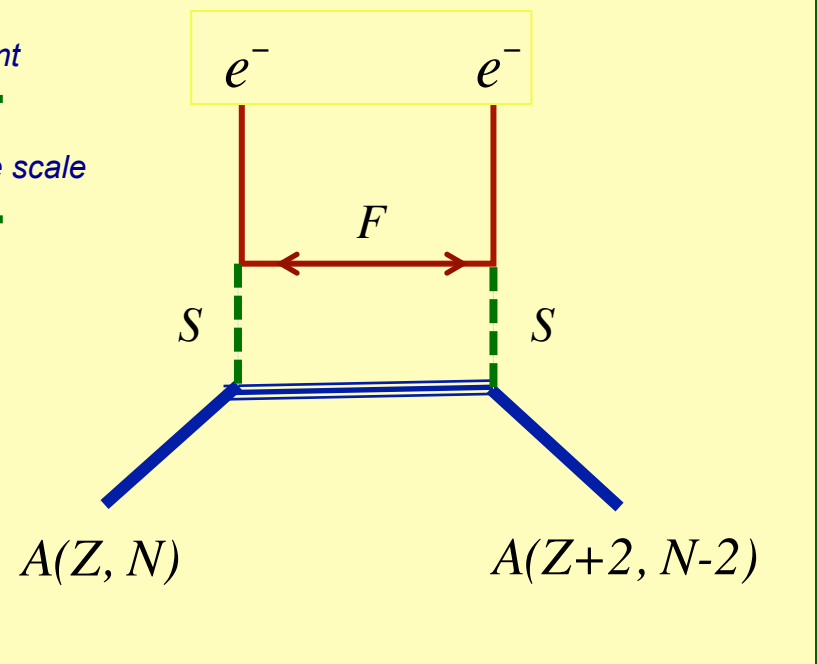
*Majorana*

## Benchmark Sensitivity: TeV LNV



*Present*

*Tonne scale*



# $0\nu\beta\beta$ -Decay: TeV Scale LNV

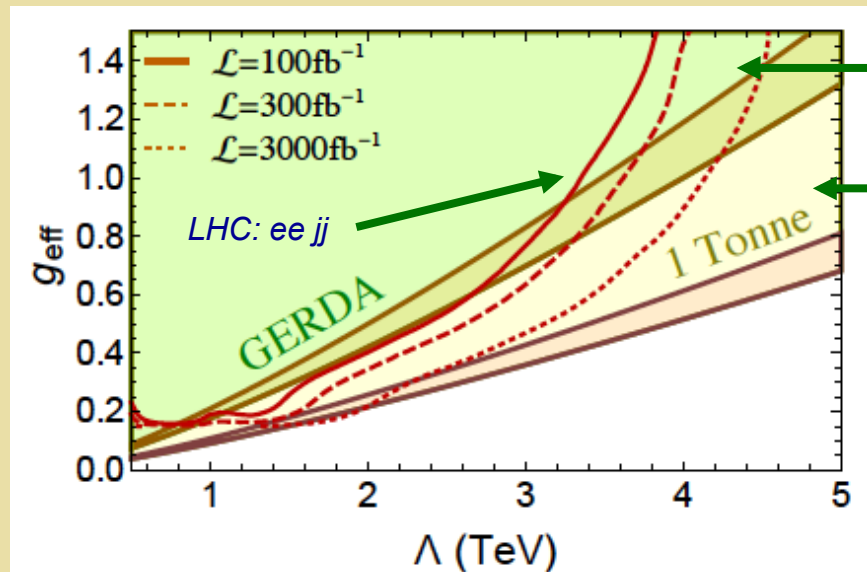
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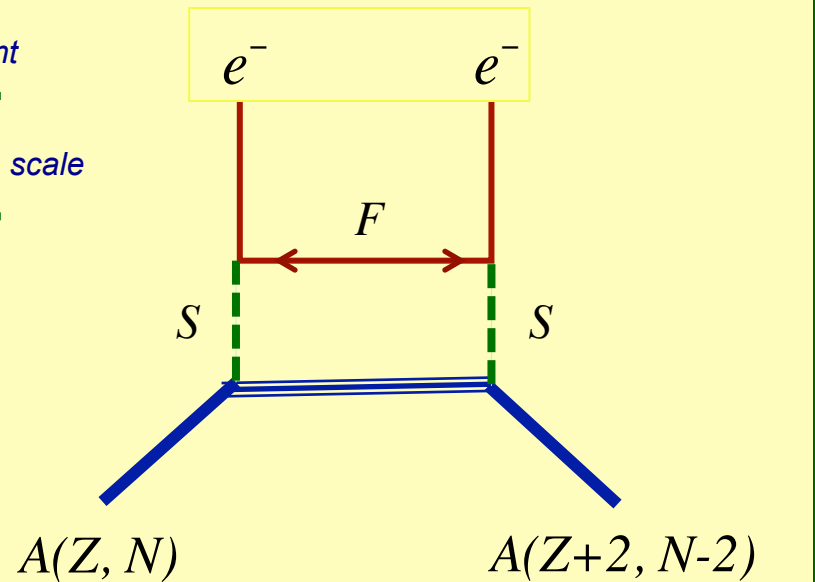
*Majorana*

## Benchmark Sensitivity: TeV LNV



*Present*

*Tonne scale*



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# $0\nu\beta\beta$ -Decay: TeV Scale LNV & $m_\nu$

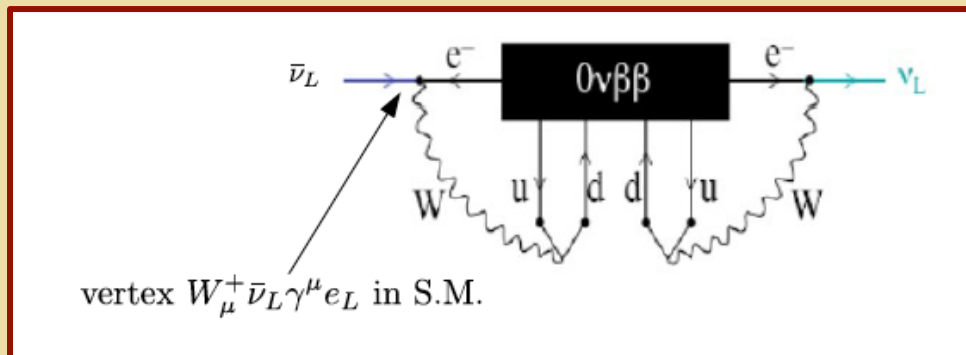
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*Dirac*

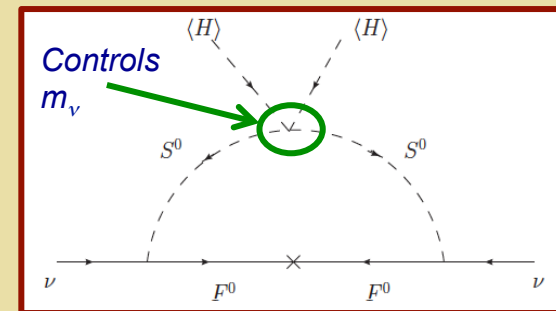
$$\mathcal{L}_{\text{mass}} = \frac{y}{\Lambda} \bar{L}^c H H^T L + \text{h.c.}$$

*Majorana*

*Implications for  $m_\nu$ :*



*Schechter-Valle: non-vanishing Majorana mass at (multi) loop level*



*Simplified model: possible (larger) one loop Majorana mass*

# $0\nu\beta\beta$ -Decay: TeV Scale LNV & $m_\nu$

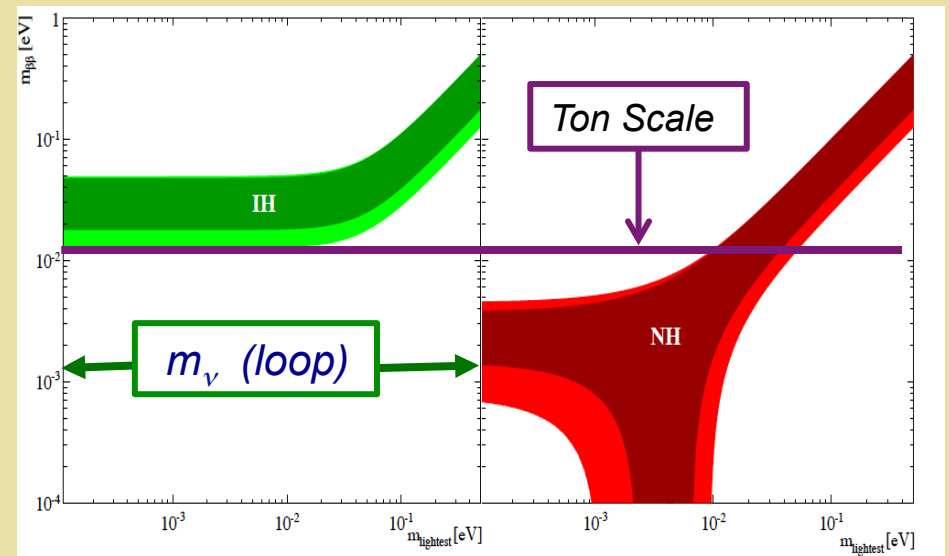
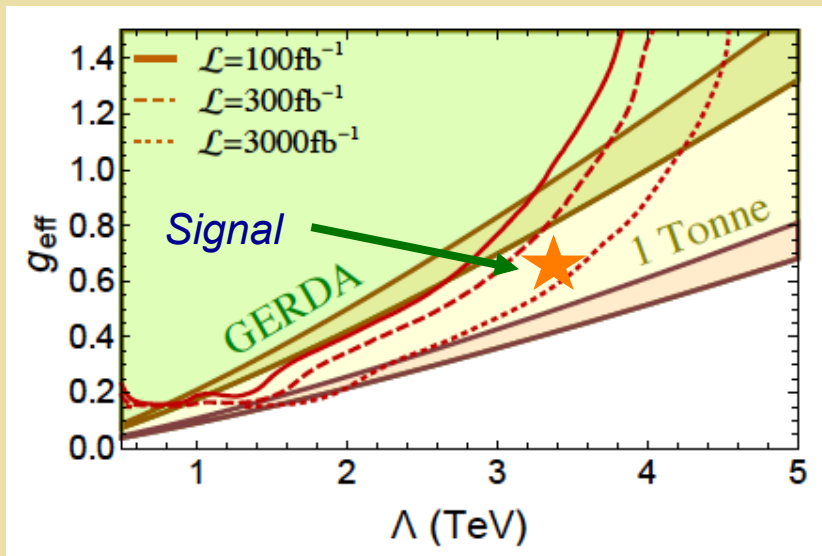
$$\mathcal{L}_{\text{mass}} = y \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

*Dirac*

$$\mathcal{L}_{\text{mass}} = \frac{y}{\Lambda} \bar{L}^c H H^T L + \text{h.c.}$$

*Majorana*

*Implications for  $m_\nu$ :*



*A hypothetical scenario*

# $0\nu\beta\beta$ / LHC Interplay: Matrix Elements

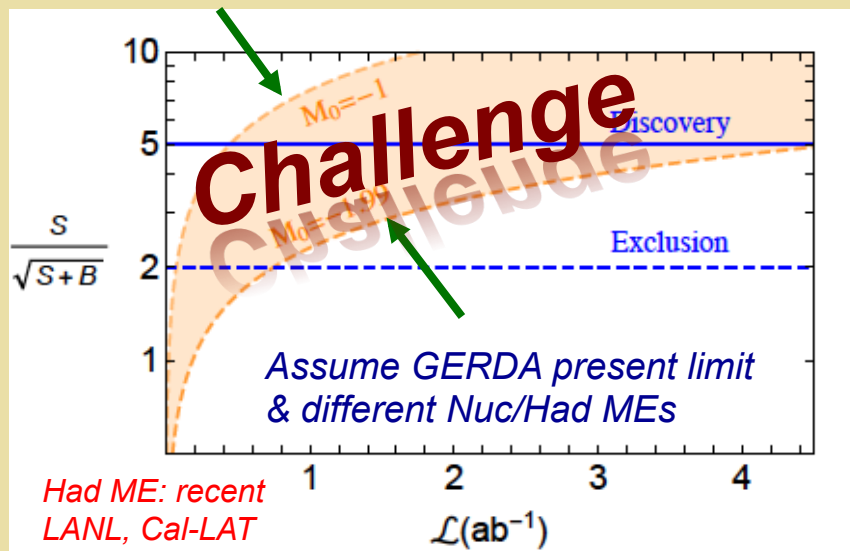
$$\mathcal{L}_{\text{mass}} = y \bar{L} \tilde{H} \nu_R + \text{h.c.}$$

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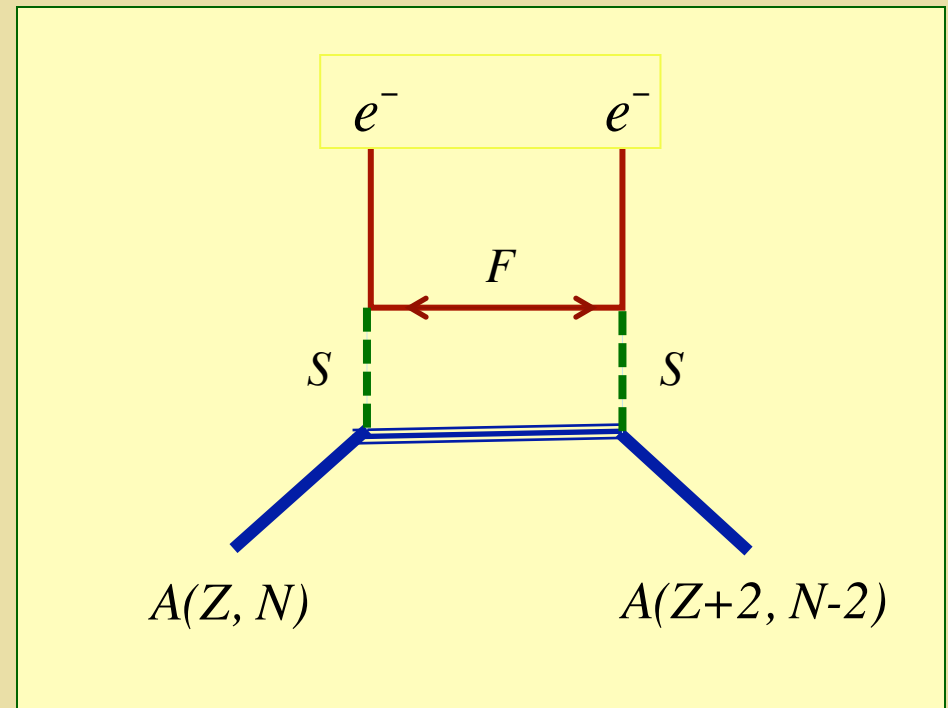
$$\mathcal{L}_{\text{mass}} = \frac{y}{\Lambda} \bar{L}^c H H^T L + \text{h.c.}$$

*Majorana*

**Benchmark Sensitivity: TeV LNV**



T. Peng, MRM, P. Winslow 1508.04444

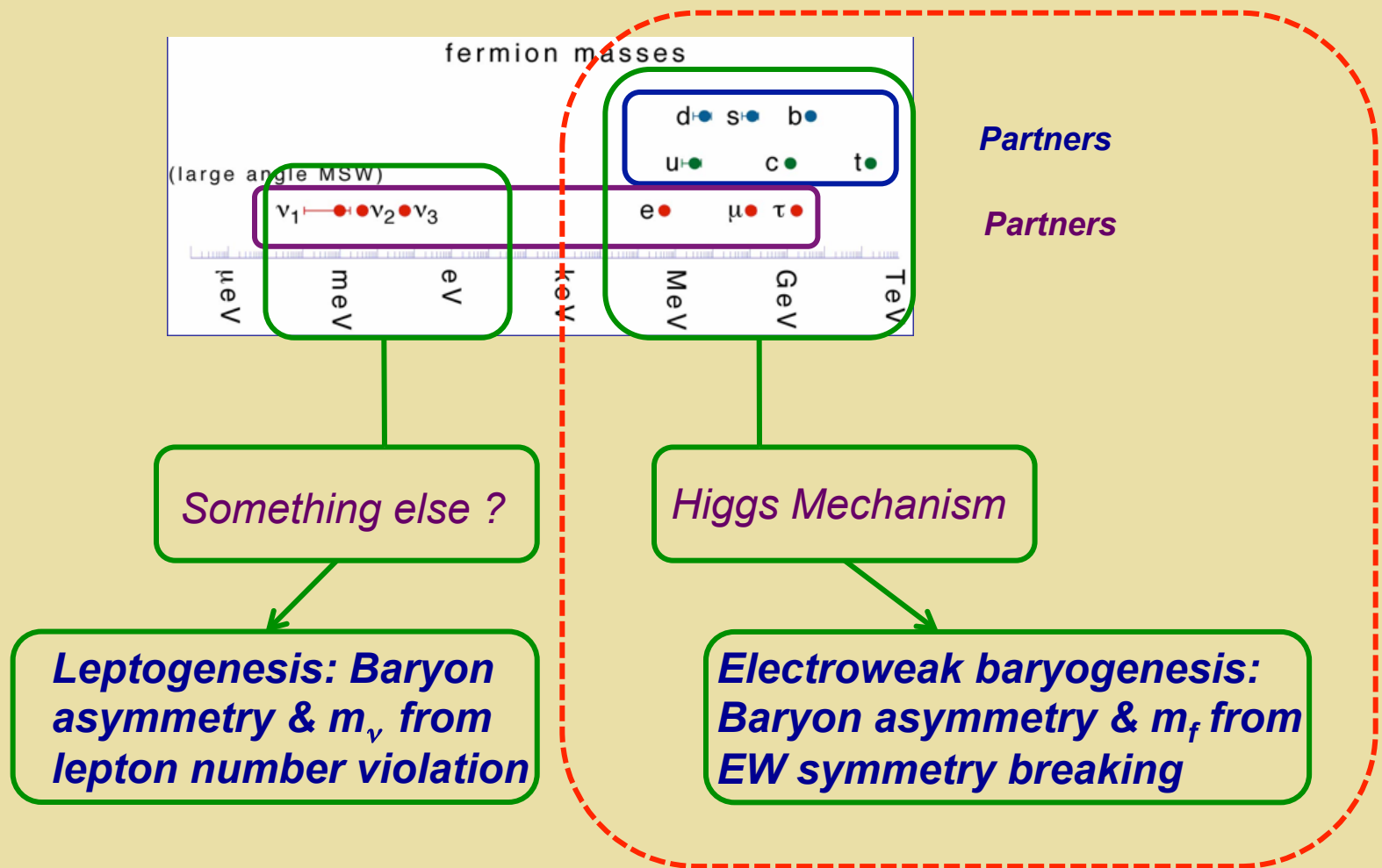


### ***III. CP (Flavor Conserving)***

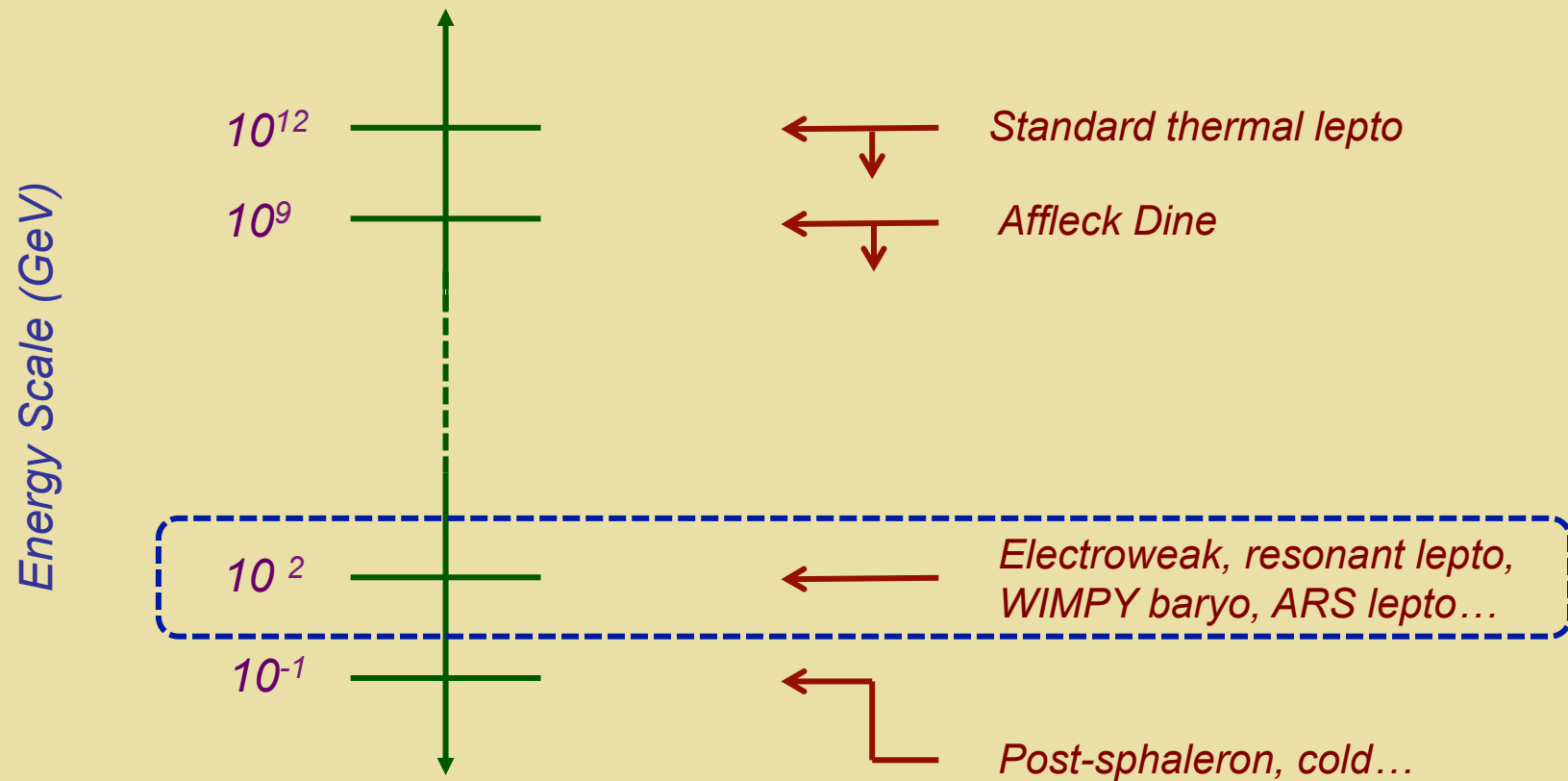
*Highlight: EW Baryogenesis, Dark Z*

*Collaborators: T. Chupp, K. Fuyuto, X.-G. He, S. Inoue, G. Li, Y. Li, G. Ovanesyan, S. Profumo, S. Shen*

# Fermion Masses & Baryon Asymmetry



# Baryogenesis Scenarios

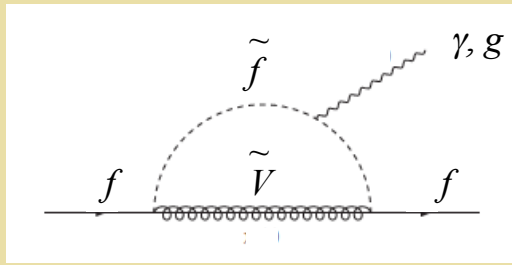


Era of EWSB:  $t_{\text{univ}} \sim 10 \text{ ps}$

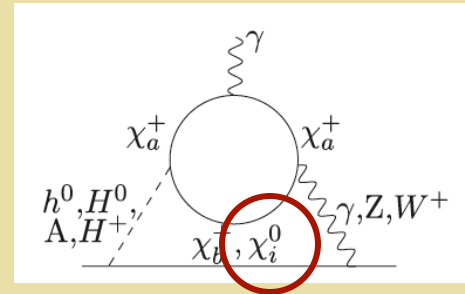
# ***Electroweak Baryogenesis***

- *Was  $Y_B$  generated in conjunction with electroweak symmetry-breaking?*
- *To what extent can EDM searches test this scenario?*

# EDMs & EWBG: MSSM & Beyond

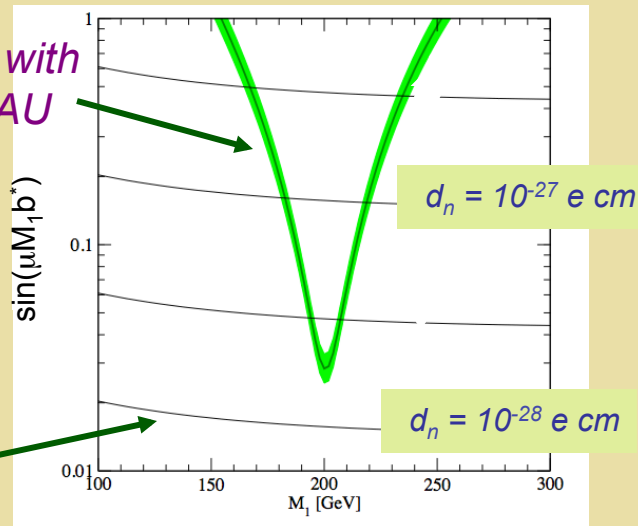


Heavy sfermions: LHC consistent & suppress 1-loop EDMs



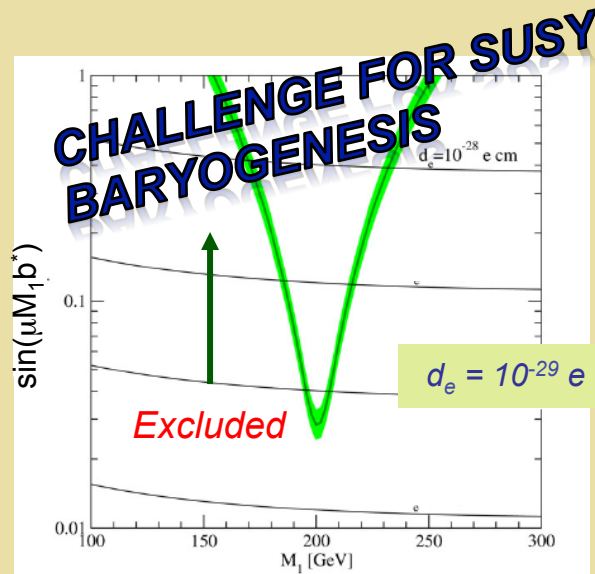
Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases

Compatible with observed BAU



Next gen  $d_n$

Li, Profumo, RM '09-'10

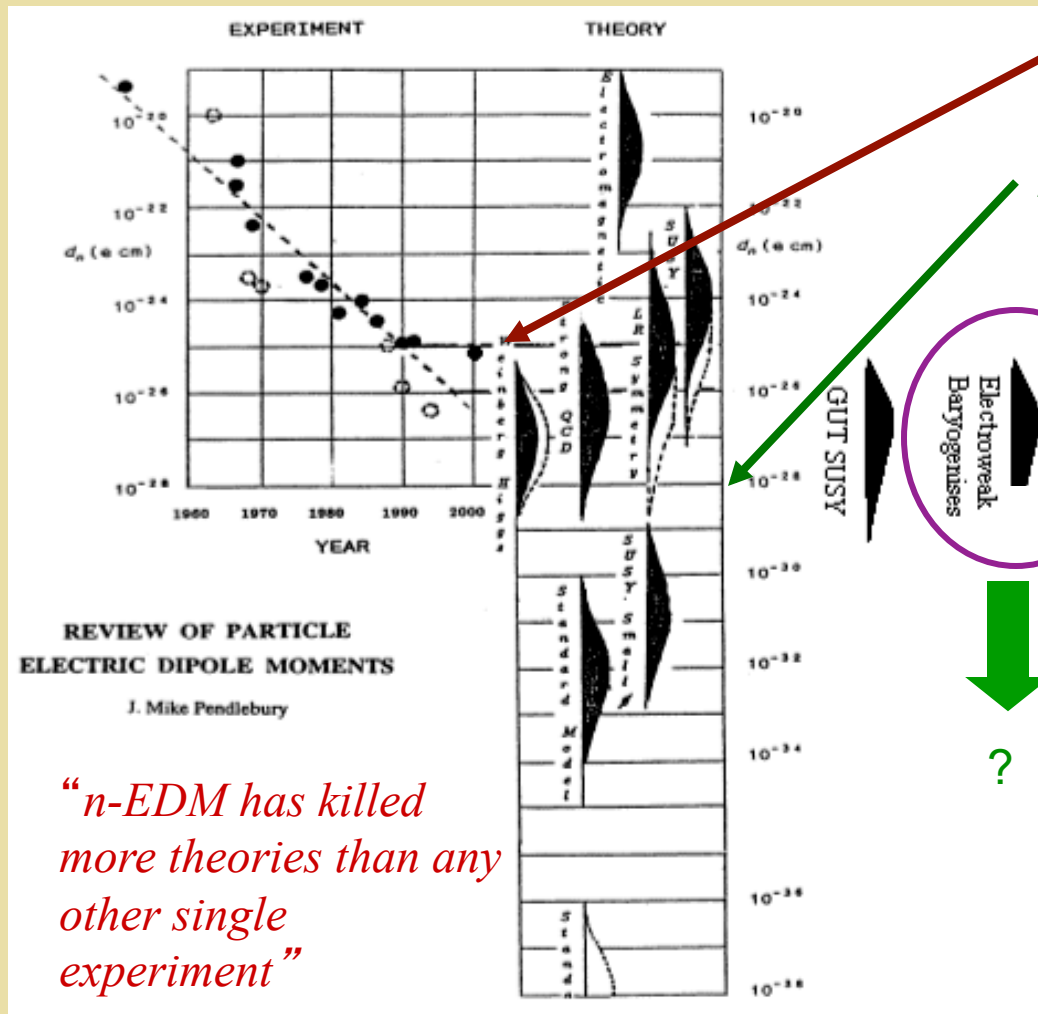


CHALLENGE FOR SUSY BARYOGENESIS

ACME: ThO



# EDMs: What We May Learn



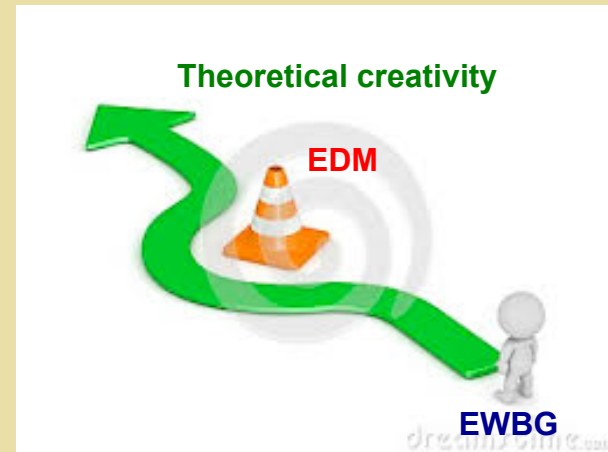
Present *n*-EDM limit

Proposed *n*-EDM limit

Matter-Antimatter  
Asymmetry in  
the Universe:  
**MSSM**

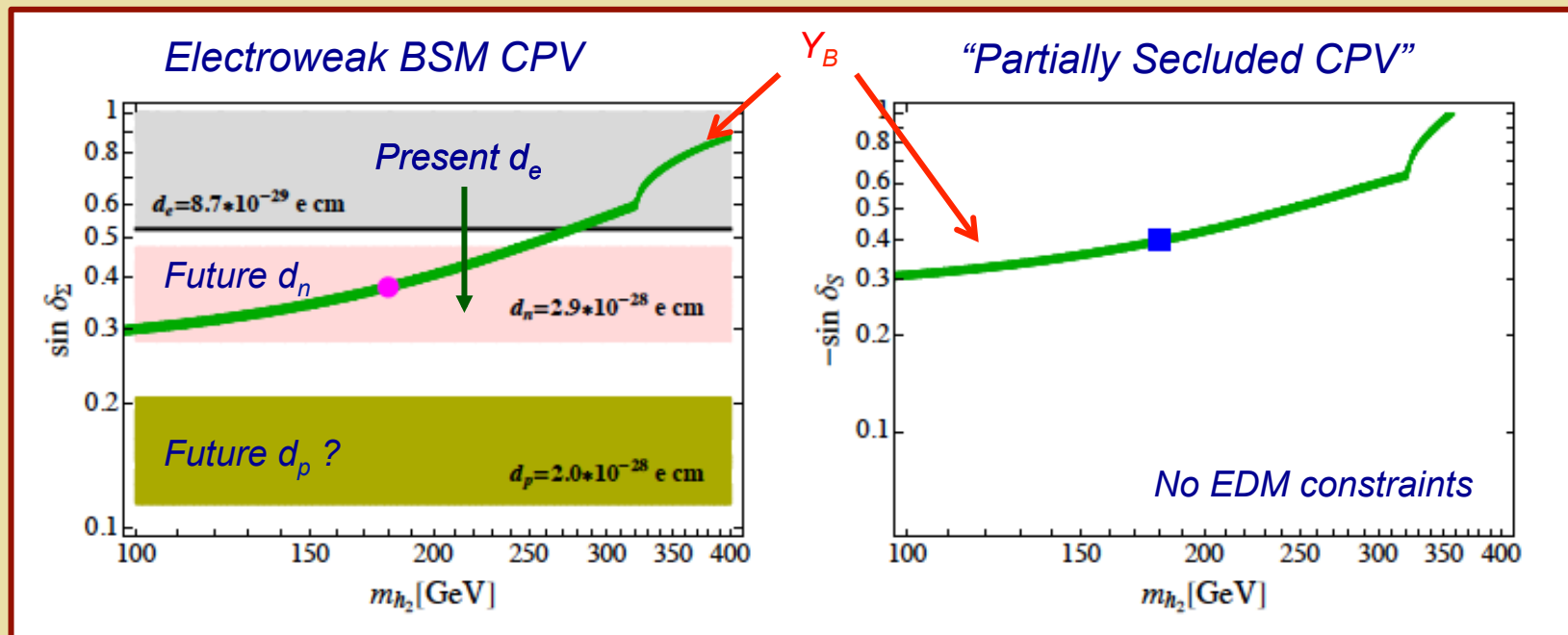
Theory: How robust?  
Can EDMs kill EW  
baryogenesis?

# ***CPV for EWBG***

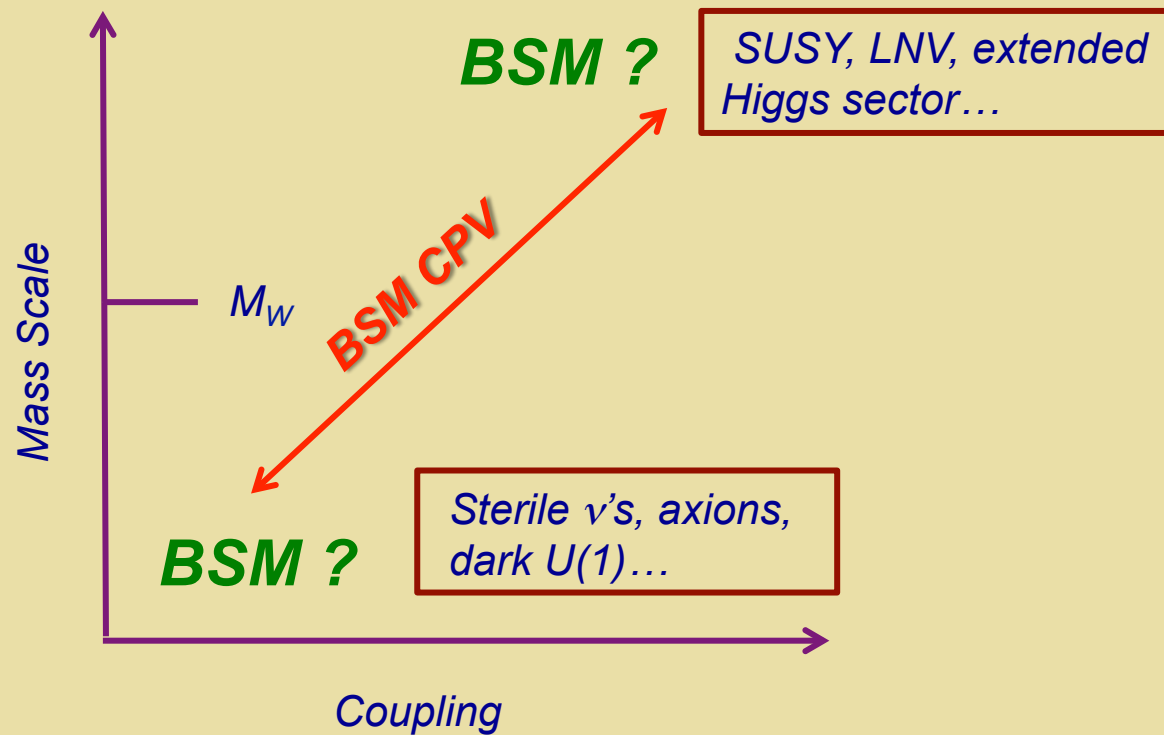


# “Two-Step EW Baryogenesis”

Two CPV sources for baryon asymmetry



# ***BSM Physics: Where Does it Live ?***



## ***BSM Physics: $T$ (CP) Invariant ?***

# *Dark Photon Portal*

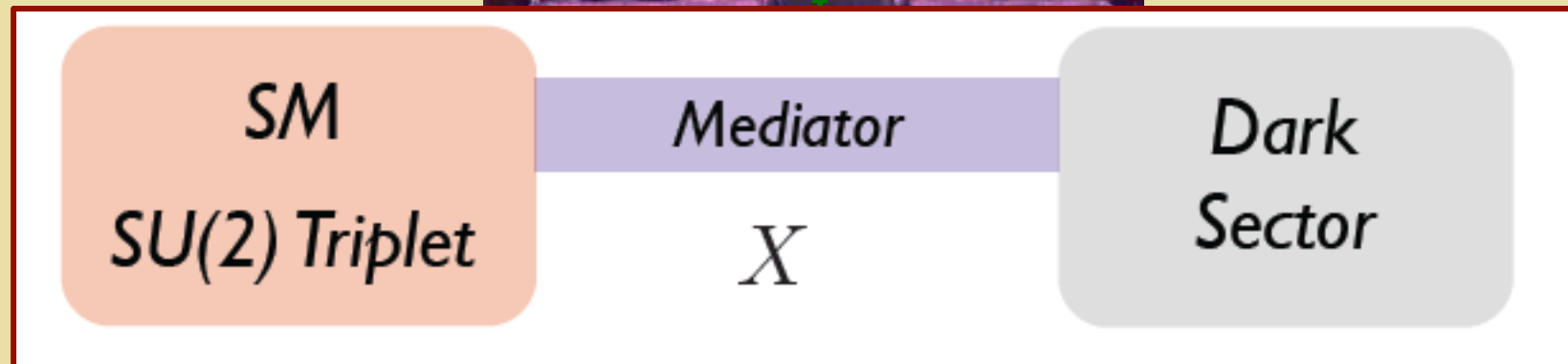


*Standard Model*

*Hidden Sector*

*New CPV ?*

# *Dark Photon Portal*



*Thanks: K. Fuyuto*

# CPV Dark Photon

$$\mathcal{L}^{(d=5)} = -\frac{\beta}{\Lambda} \text{Tr}[W_{\mu\nu} \Sigma] X^{\mu\nu} - \frac{\tilde{\beta}}{\Lambda} \text{Tr}[W_{\mu\nu} \Sigma] \tilde{X}^{\mu\nu}$$

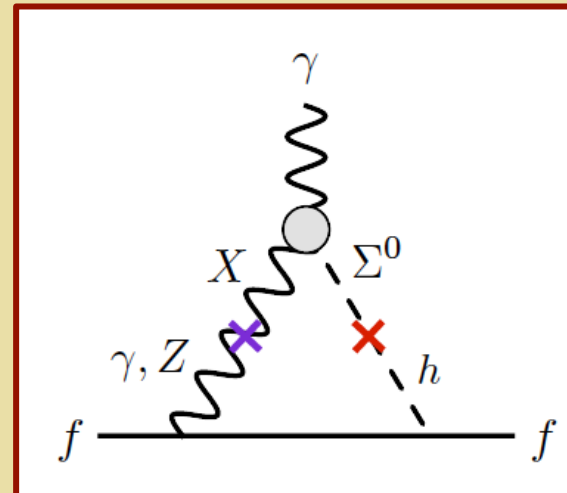
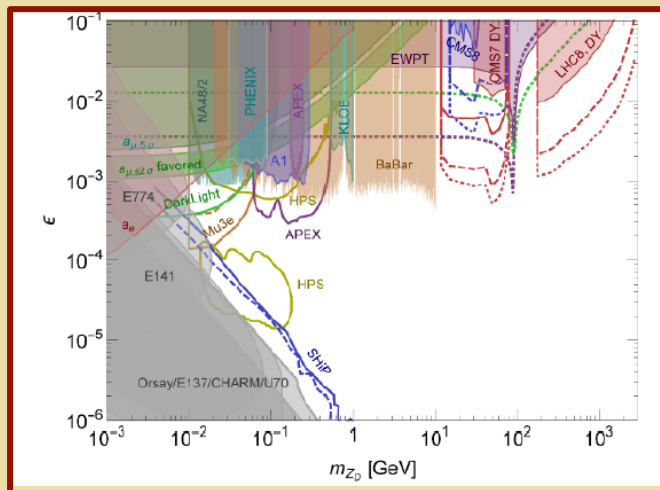
CP-conserving

CP-violating

Thanks: K. Fuyuto

$X - \gamma$  Mixing

EDM



# CPV Dark Photon

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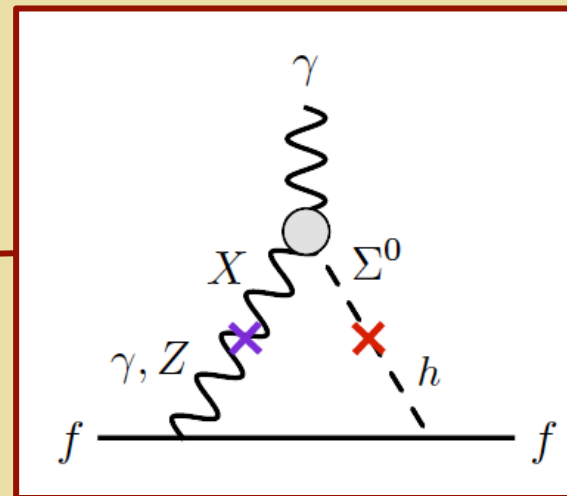
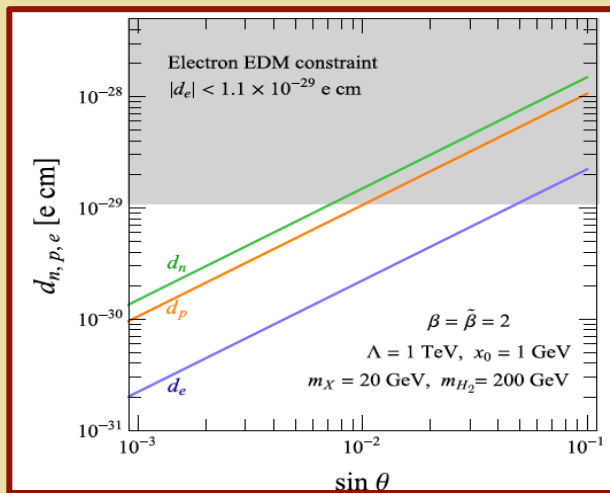
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$X - \gamma$  Mixing

EDM

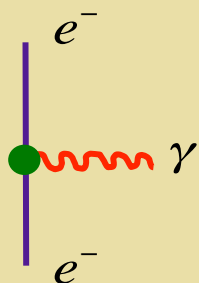




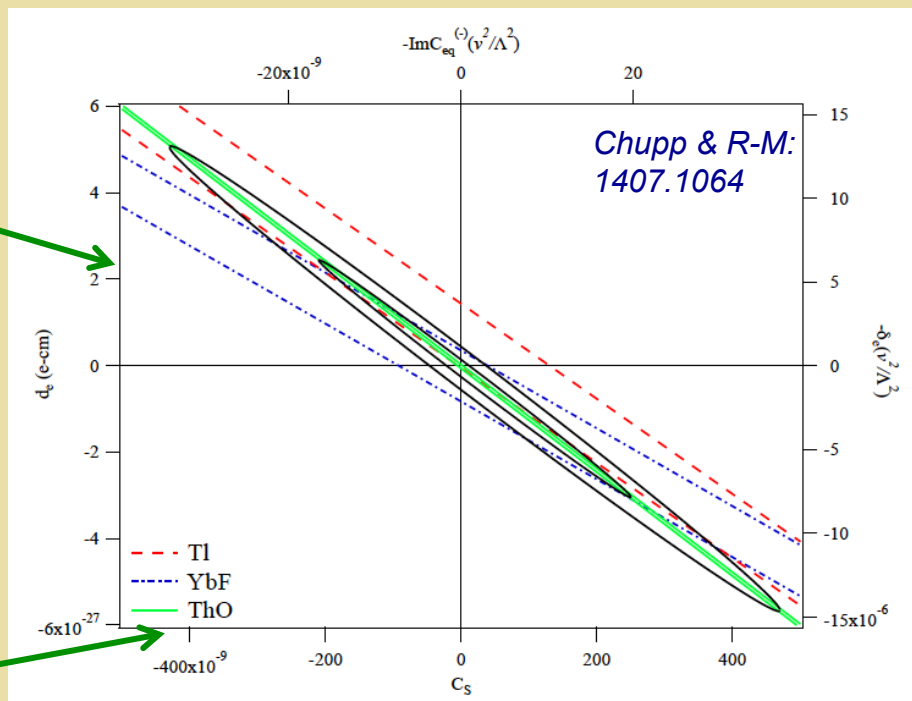
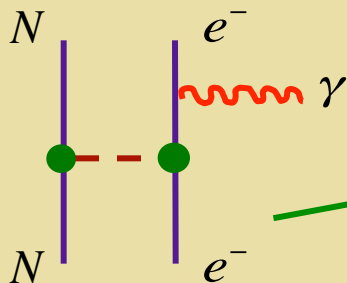
# *EDM Complementarity*

# Paramagnetic Systems: Two Sources

Electron  
EDM



(Scalar  $q$ )  
 $\times$  (PS  $e^-$ )



$$\Lambda \gtrsim (1.5 \text{ TeV}) \times \sqrt{\sin \phi_{\text{CPV}}}$$

Electron EDM (global)

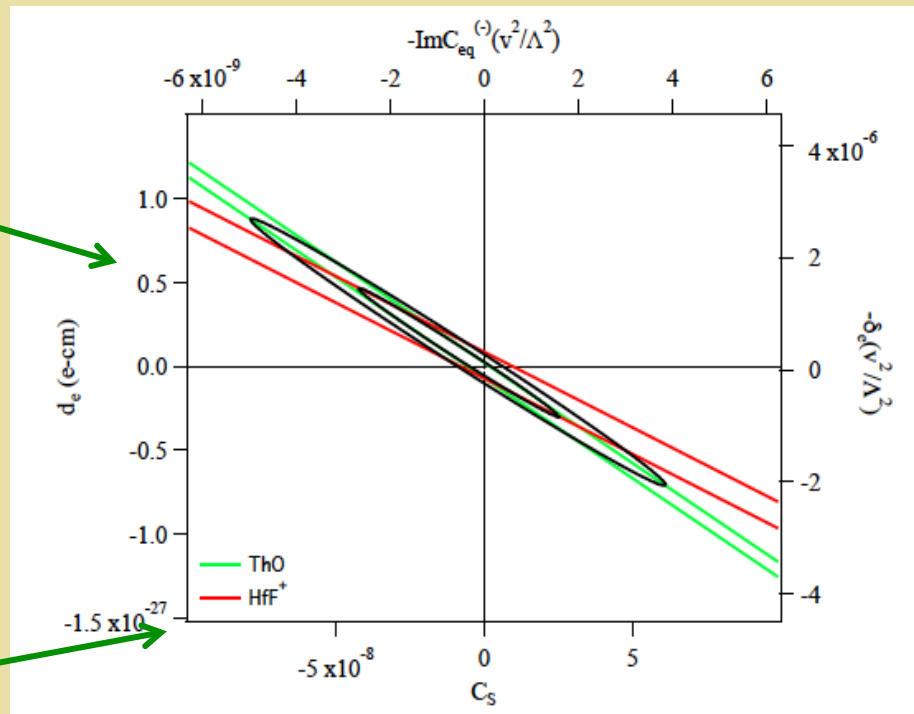
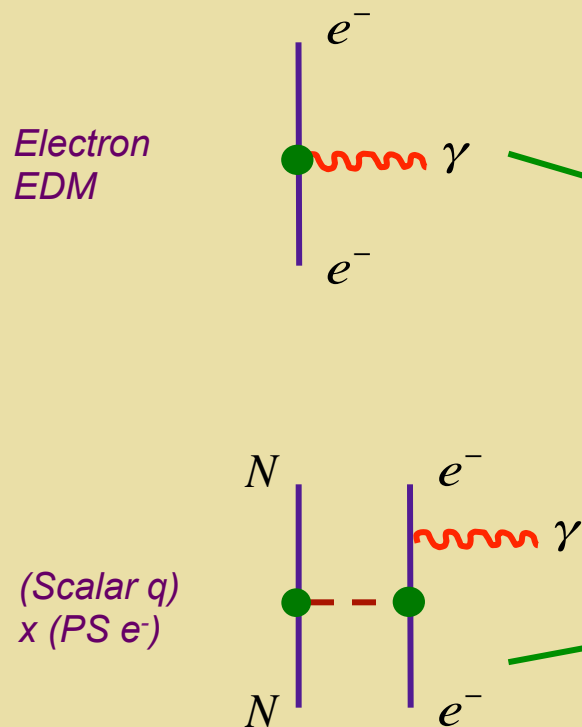
$$\Lambda \gtrsim (1300 \text{ TeV}) \times \sqrt{\sin \phi_{\text{CPV}}}$$

$C_S$  (global)

*LHC inaccessible*

*Tl, YbF, ThO...*

# Paramagnetic Systems: Two Sources



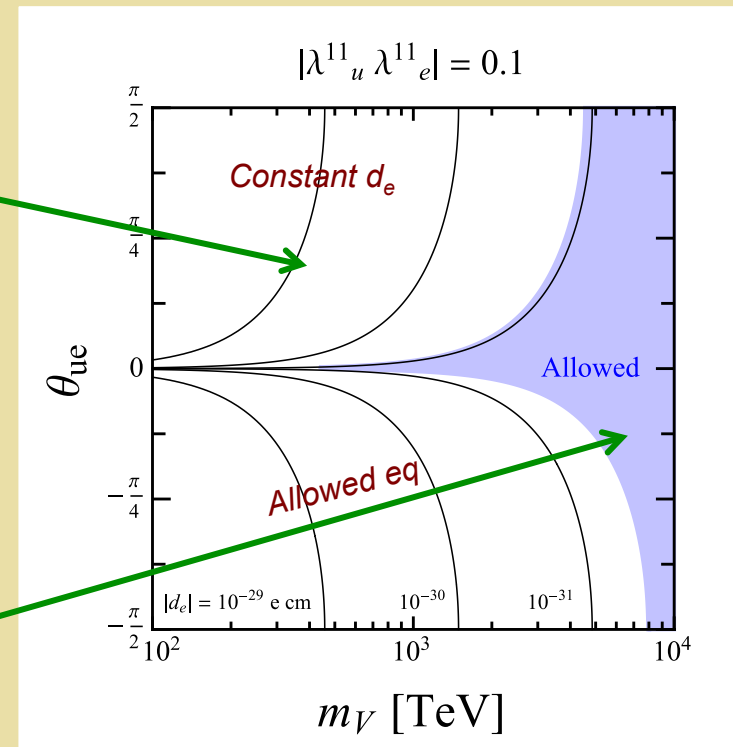
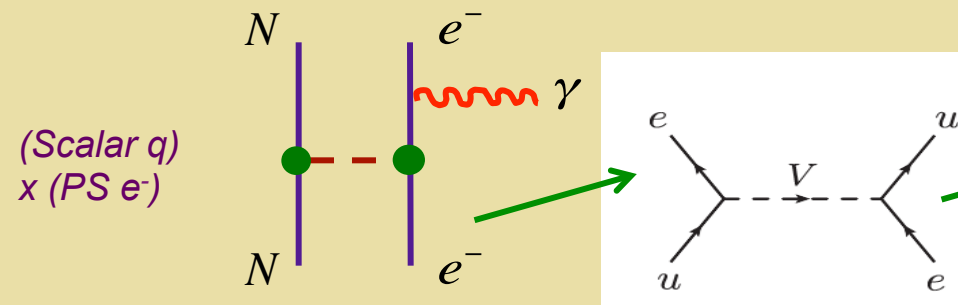
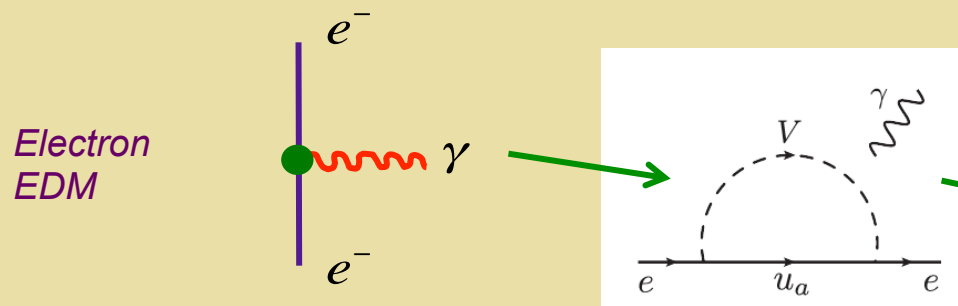
Chupp, Fierlinger, R-M, Singh 1710.02504;  
Fleig & Jung 1802.02171

Inclusion of HfF+ :  $\sim 6$  times stronger  
bounds on  $d_e$  &  $C_s \rightarrow 2.5$  higher on  $\Lambda$

New ThO  $\rightarrow$  even stronger !

Tl, YbF, ThO, HfF+

# Illustrative Example: Leptoquark Model



(3, 2, 7/6)

Fuyuto, R-M, Shen 1804.01137

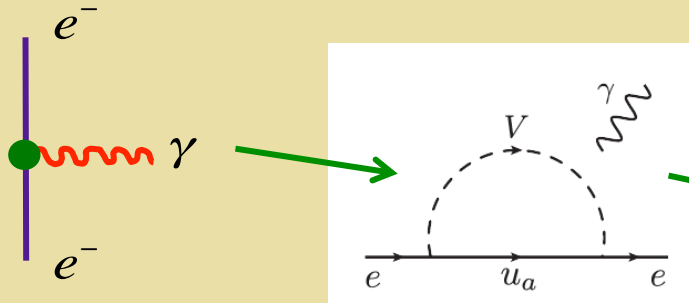
$$\mathcal{L} \ni -\lambda_u^{ab} \bar{u}_R^a X^T \epsilon L^b - \lambda_e^{ab} \bar{e}_R^a X^\dagger Q^b + \text{h.c.}$$

# Illustrative Example: Leptoquark Model

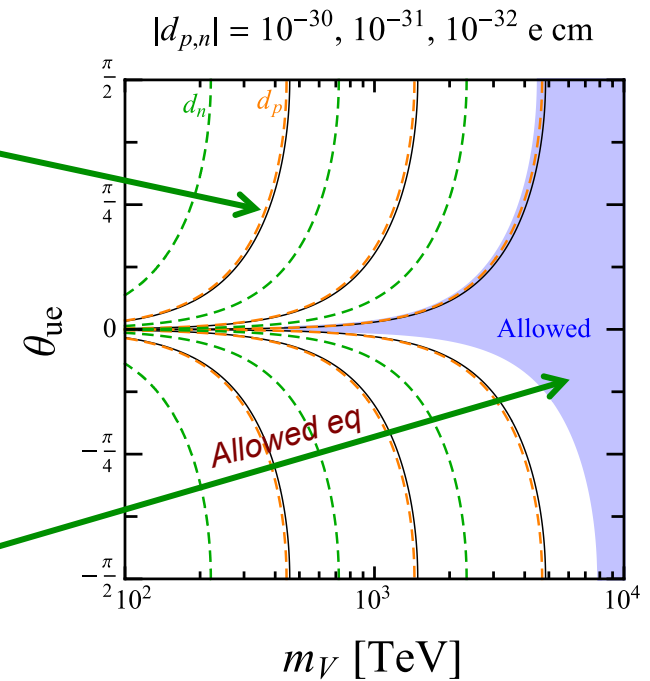
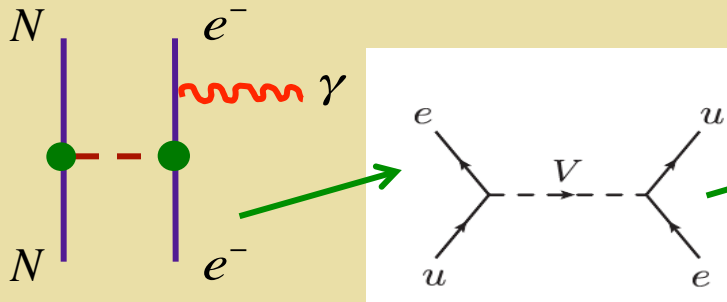
Orange:  $|d_p| = 10^{-30}, 10^{-31}, 10^{-32} \text{ e cm}$

Green:  $|d_n| = 10^{-30}, 10^{-31}, 10^{-32} \text{ e cm}$

Electron  
EDM



(Scalar  $q$ )  
 $\times$  (PS  $e^-$ )



(3, 2, 7/6)

Fuyuto, R-M, Shen 1804.01137

$$\mathcal{L} \ni -\lambda_u^{ab} \bar{u}_R^a X^T \epsilon L^b - \lambda_e^{ab} \bar{e}_R^a X^\dagger Q^b + \text{h.c.}$$

## ***IV. Precision Tests***

*Highlight:  $\beta$ -decay*

*Collaborators: M. Gorchtein, H. Patel, C. Seng*

*PRL 121 (2018) 241804 [1807.10197], 1812.03352*

## ***Precision ~ BSM Mass Scale***

$$\Delta_{SM} = (O_{EXP} - O_{SM}) / O_{SM}$$

*Precision Goal: \**

$$\delta \Delta_{SM} \sim O(10^{-4})$$

*Heavy BSM Physics:*

$$\Delta_{SM} \sim C (v/\Lambda)^2$$

$$\Lambda \sim 10 \text{ TeV (tree)}$$

$$\Lambda < 1 \text{ TeV (loop)}$$

*\* Can be  $\sim 10^{-3}$  or larger if  $O_{SM}$  suppressed*

# Weak Decays: CKM Unitarity

$$d \rightarrow u e^- \bar{\nu}_e$$

$$s \rightarrow u e^- \bar{\nu}_e$$

$$b \rightarrow u e^- \bar{\nu}_e$$

$$\begin{pmatrix} u & c & t \end{pmatrix} \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

$$\Delta_{\text{CKM}} = (|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2)_{\text{exp}} - 1$$

$$0.94906 \pm 0.00041$$

$$0.05031 \pm 0.00022$$

$$0.00002$$

$$\Delta_{\text{CKM}} = -0.0006 \pm 0.0005$$



# ***Precision ~ BSM Mass Scale***

*Precision Goal:*

$$\delta \Delta_{CKM} \sim O(10^{-4})$$

*Heavy BSM Physics:*

$$\Delta_{CKM} \sim C \left( v/\Lambda \right)^2$$

$$\Lambda \sim 10 \text{ TeV (tree)}$$

$$\Lambda < 1 \text{ TeV (loop)}$$

*Ultralight BSM Physics:*


$$\Delta_{CKM} \sim \varepsilon^2 \left( \alpha/4\pi \right)$$

$$\varepsilon < 1 \text{ (loop)}$$


## ***Error Budget***

$$\Delta_{\text{CKM}} = (|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2)_{\text{exp}} - 1$$

0.94906 ± 0.00041



0.05031 ± 0.00022



0.00002



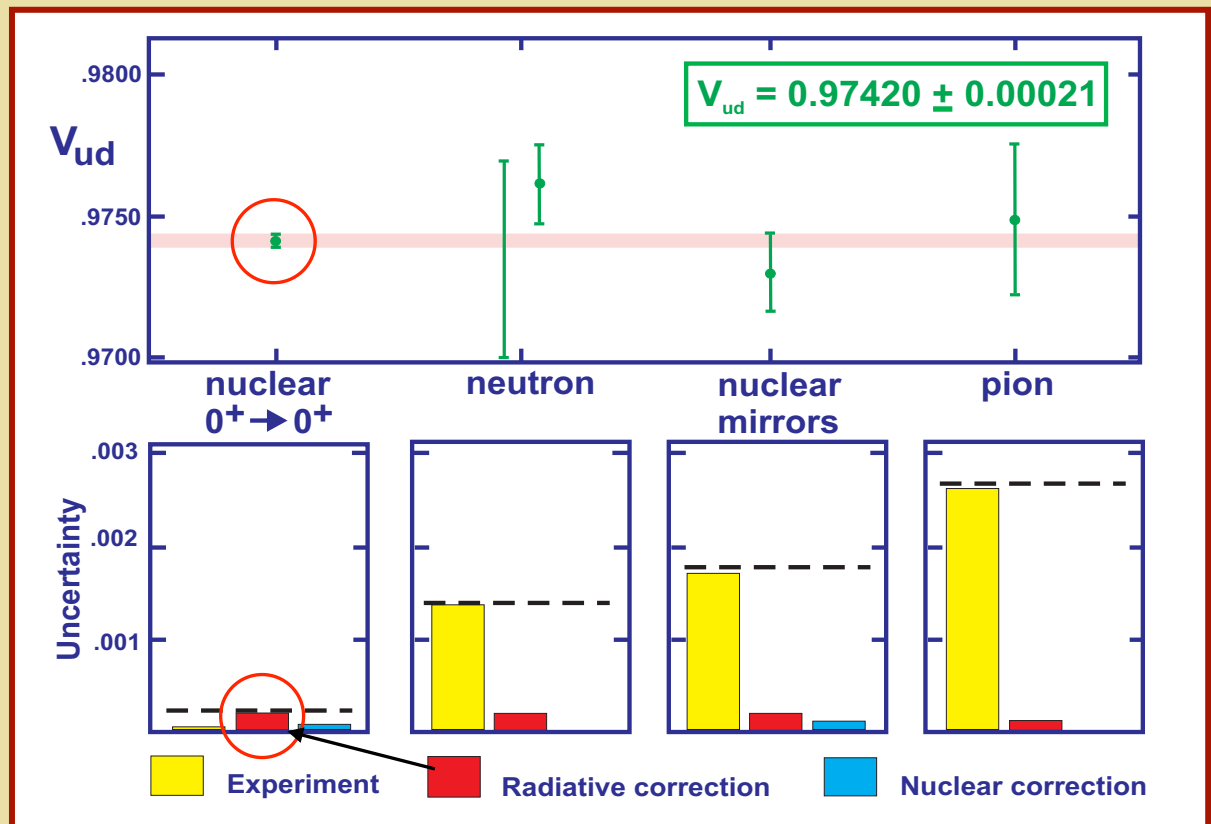
# Error Budget

$$\Delta_{\text{CKM}} = (|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2)_{\text{exp}} - 1$$

$$0.94906 \pm 0.00041$$

**Radiative Correction**

*Factor of 2 reduction  
using disp relations*



Thanks: J. Hardy

# Error Budget

$$\Delta_{\text{CKM}} = (|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2)_{\text{exp}} - 1$$

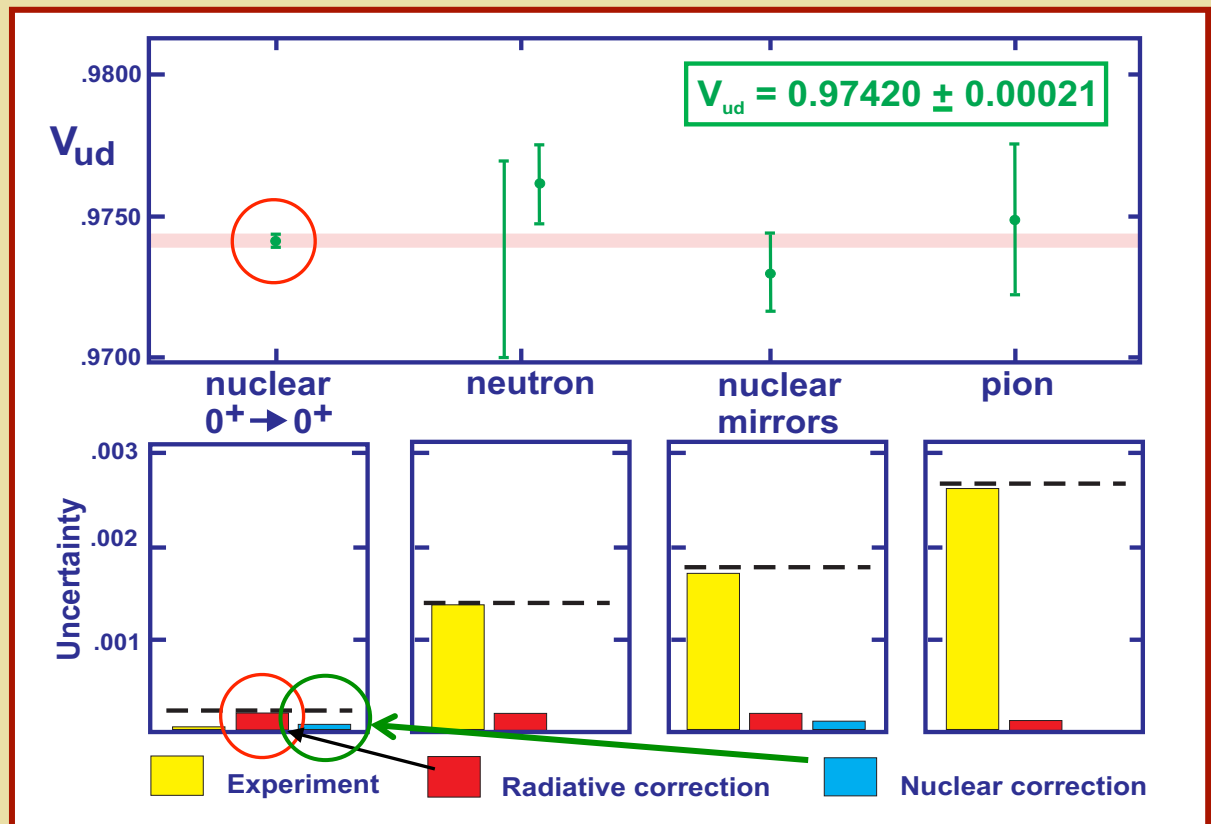
$$0.94906 \pm 0.00041$$

## Radiative Correction

*Factor of 2 reduction  
using disp relations*

## Nuclear Correction

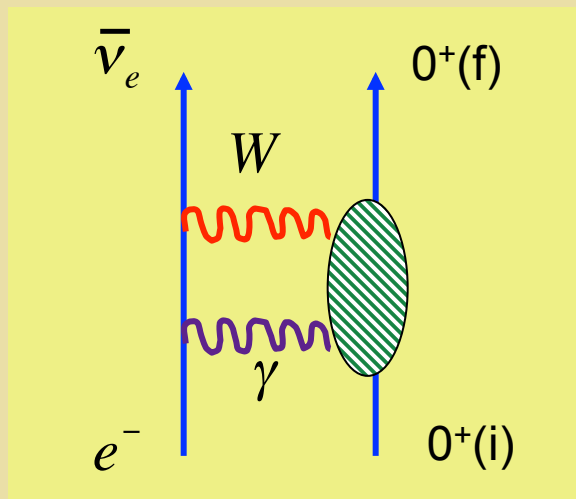
*Increase due to previously omitted contributions*



*Thanks: J. Hardy*

# Radiative Corrections

*Dominant source of uncertainty:*



$$M_{\gamma W} = \frac{G_F V_{ud}}{\sqrt{2}} \frac{\alpha}{8\pi} \left[ \ln \left( \frac{M_Z^2}{\Lambda^2} \right) + C_{\gamma W}(\Lambda) \right]$$

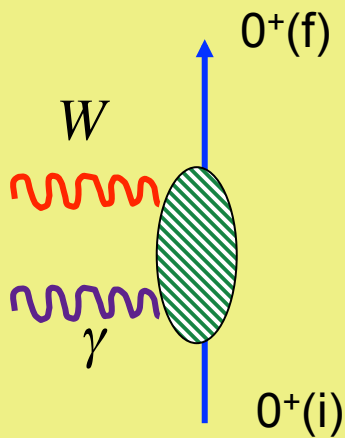
*Short distance*

*Long distance*

*Long distance*

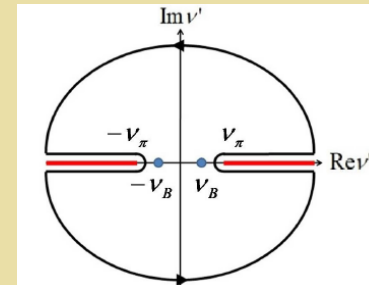
*Sensitive to hadronic & nuclear dynamics*

# Dispersion Relations



**Dispersion relation:**

Write  $T_3$  as integral over discontinuity along cut



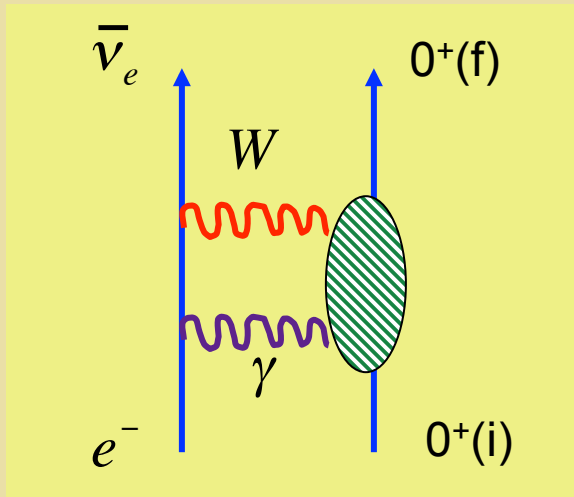
$$T_3^{(I)}(\nu, Q^2) = \frac{2}{i} \int_0^\infty d\nu' \left[ \frac{1}{\nu' - \nu} + \frac{\xi^I}{\nu' + \nu} \right] F_3^{(I)}(\nu', Q^2)$$

**Electroproduction structure functions:**

$$\begin{aligned} W_{\gamma W}^{(I)\mu\nu} &= \frac{1}{8\pi} \sum_X (2\pi)^4 \delta^4(p + q - p_X) \langle p | J_{em}^{(I)\mu} | X \rangle \langle X | J_W^\nu | n \rangle \\ &= \left[ -g^{\mu\nu} + \frac{q^\mu q^\nu}{q^2} \right] F_1^{(I)} + \frac{\hat{p}^\mu \hat{p}^\nu}{(p \cdot q)} F_2^{(I)} + \frac{i\epsilon^{\mu\nu\alpha\beta} p_\alpha q_\beta}{2(p \cdot q)} F_3^{(I)} \end{aligned}$$

$$d\sigma \propto L_{\mu\nu} W^{\mu\nu}$$

# Dispersion Relations



**Radiative Correction:**

$$\begin{aligned} \square_{\gamma W}^{VA(0)} &= \frac{\alpha}{\pi M} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} \int_0^\infty d\nu \frac{(\nu + 2q)}{\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2) \\ &= \frac{3\alpha}{2\pi} \int_0^\infty \frac{dQ^2 M_W^2}{Q^2 [M_W^2 + Q^2]} \boxed{M_3^{(0)}(1, Q^2)} \end{aligned}$$

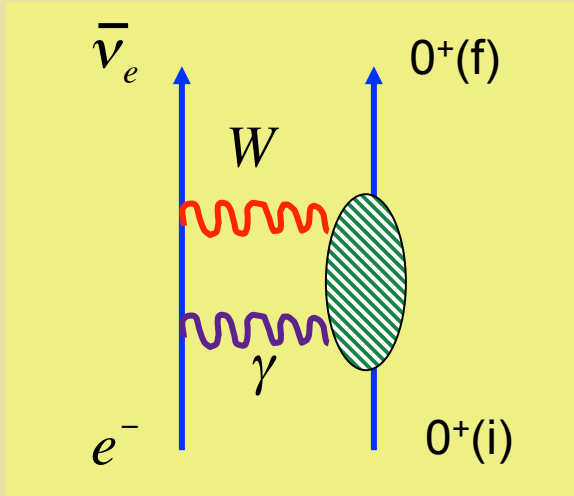
**Nachtmann Moments:**

$$\boxed{M_3^{(0)}(N, Q^2) = \frac{N+1}{N+2} \int_0^1 \frac{dx \xi^N}{x^2} \left[ 2x - \frac{N\xi}{N+1} \right] F_3^{(0)}}$$

$$\xi = 2x \left( 1 + \frac{4M^2 x^2}{Q^2} \right)^{-1}$$

$$\boxed{d\sigma \propto L_{\mu\nu} W^{\mu\nu}}$$

# Dispersion Relations



**Radiative Correction:**

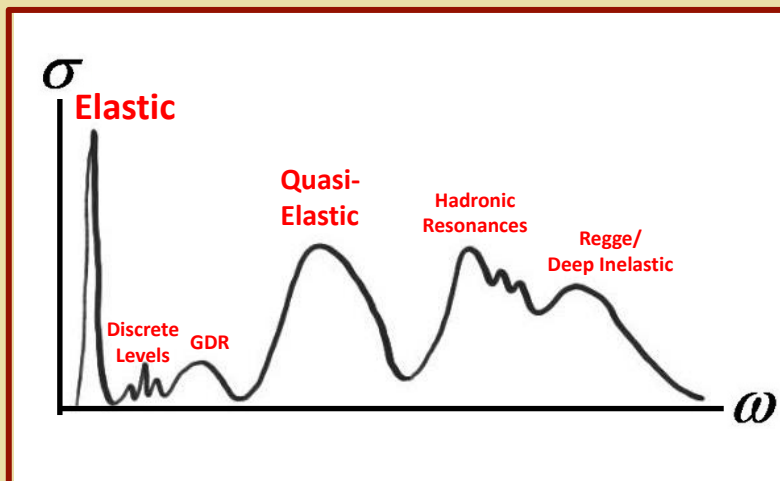
$$\begin{aligned} \square_{\gamma W}^{VA(0)} &= \frac{\alpha}{\pi M} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} \int_0^\infty d\nu \frac{(\nu + 2q)}{\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2) \\ &= \frac{3\alpha}{2\pi} \int_0^\infty \frac{dQ^2 M_W^2}{Q^2 [M_W^2 + Q^2]} M_3^{(0)}(1, Q^2) \end{aligned}$$

- **Relate  $F_3^{(0)}$  and  $M_3^{(0)}$  to data and/or**
- **Compute  $F_3^{(0)}$  and  $M_3^{(0)}$  using same methods used to describe semi-leptonic scattering processes with nucleon & nuclear targets**

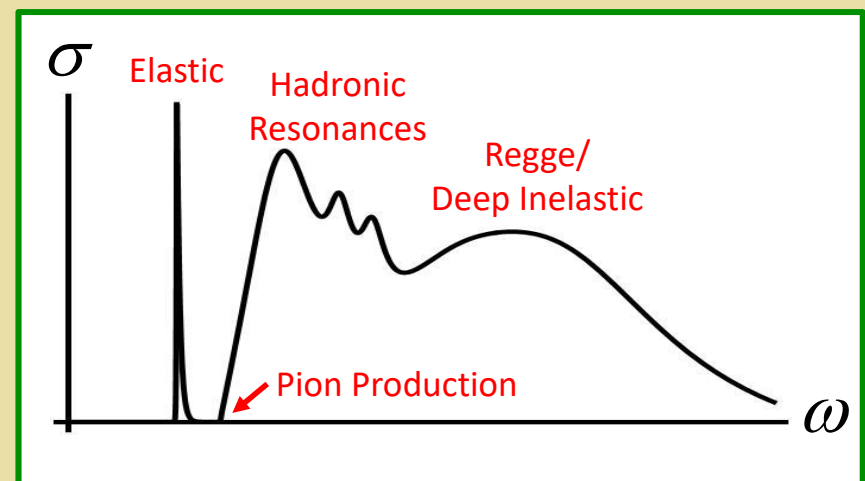


# Leptonproduction: Had & Nuc Response

## Nuclei

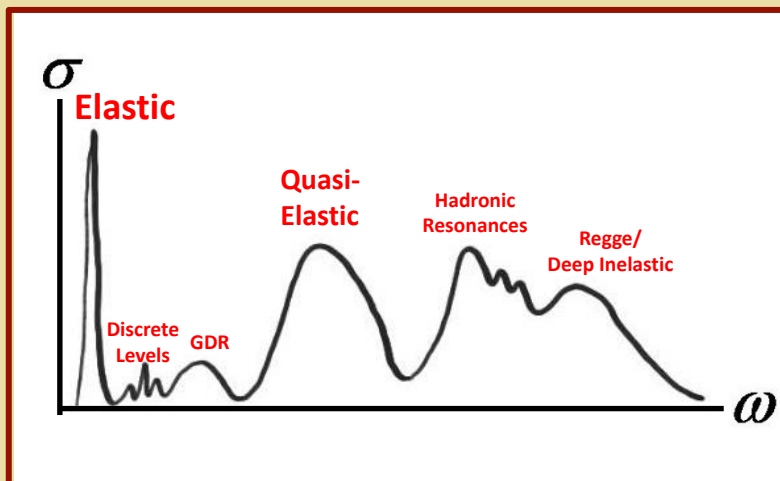


## Free nucleons

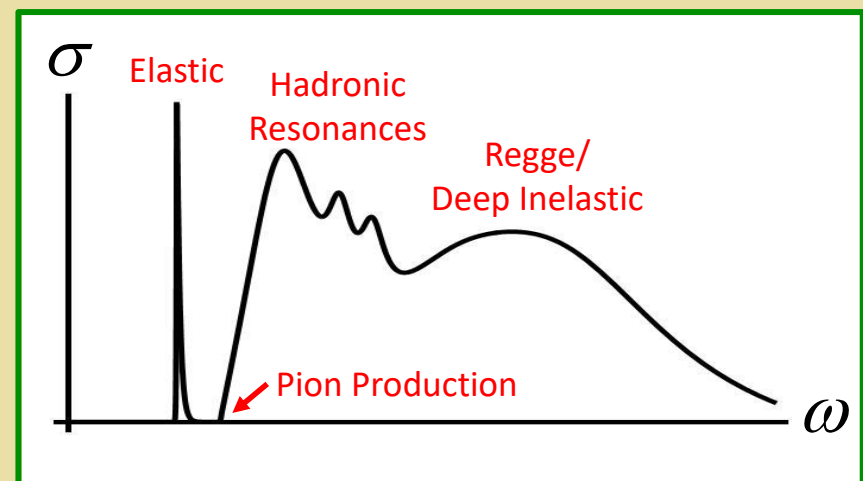


# Leptoproduction: Had & Nuc Response

## Nuclei



## Free nucleons



Single nucleon: *PRL* 121 (2008) 241804

$$\Delta_R^V = 0.02361(38) \rightarrow 0.02467(22)$$

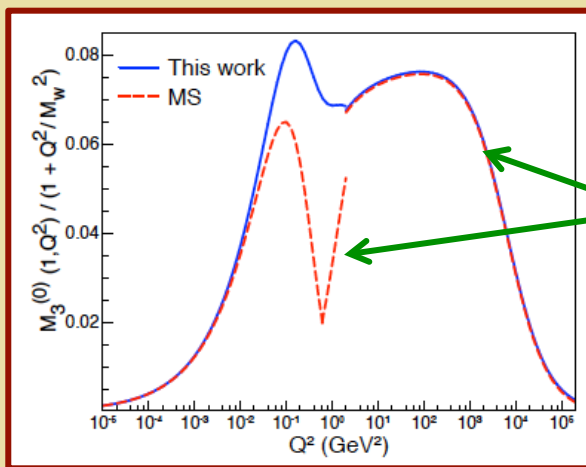
## ***$W_\gamma$ Box: Update from 2006***

$$\square_{\gamma W}^{VA} = \frac{\alpha}{2\pi} [C_{DIS} + C_B + C^{Regge} + C^{\pi N} + C^{Res}]$$

$$C_{DIS}^{MS} = 1.84 \rightarrow C_{DIS}^{new} = 1.87$$

$$C_B^{MS} = 0.829(83) \rightarrow C_B^{new} = 0.91(5)$$

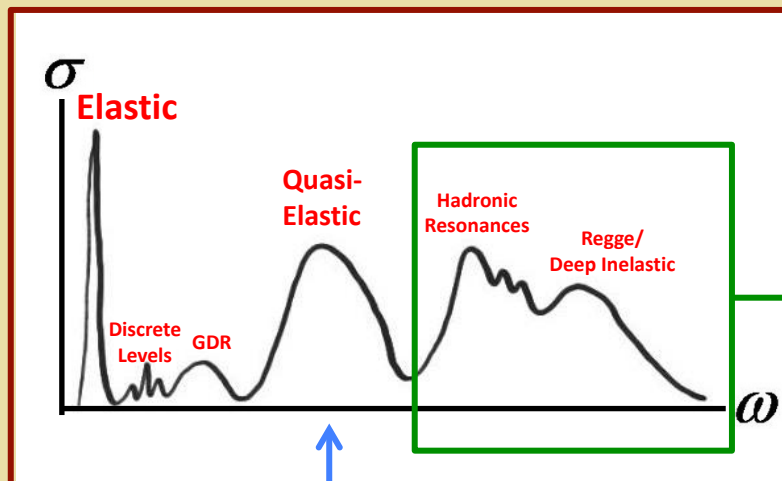
$$C_{INT}^{MS} = 0.14(14) \rightarrow C^{Regge} + C^{\pi N} + C^{Res} = 0.48(7)$$



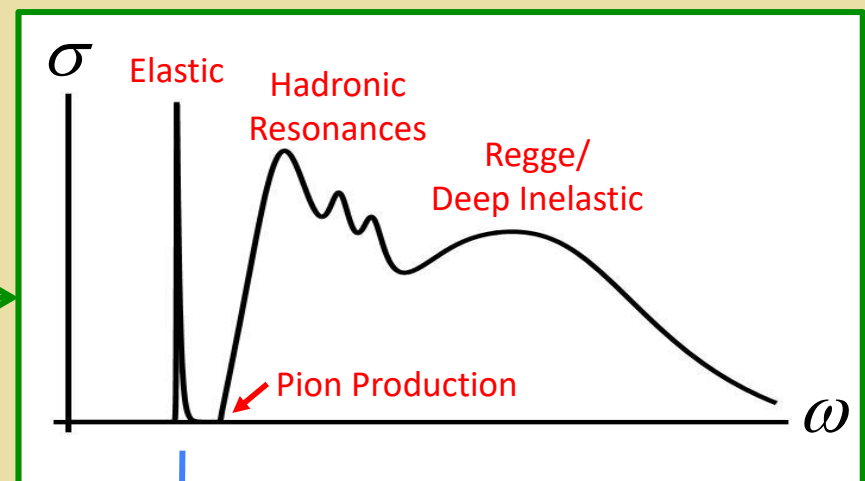
$$F_{MS}(Q^2) = \frac{12}{Q^2} M_3^{(0)}(1, Q^2)$$

# Leptoproduction: Had & Nuc Response

*Nuclei*



*Free nucleons*

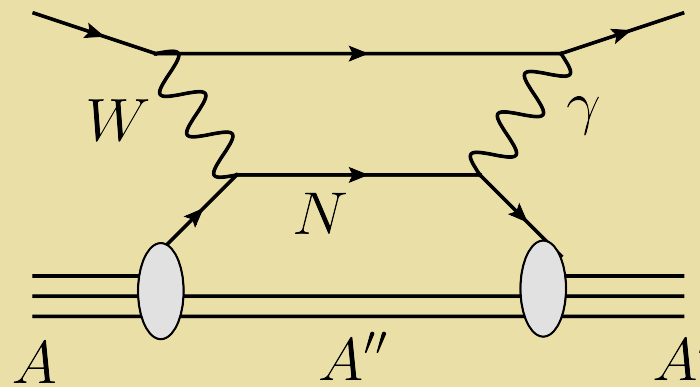


*Quasielastic response*

*Part of  $\delta_{NS}$  : “ $C_B^{Nucl}$ ”*

*New work*

## Impact on $\delta_{NS}$



$$\Delta \delta_{NS} = \frac{\alpha}{\pi} \left( C_{QE} - q_S^{(0)} q_A C_B \right) = -(4.6 \pm 0.9) \times 10^{-4}$$

**Our new work:**  
QE response

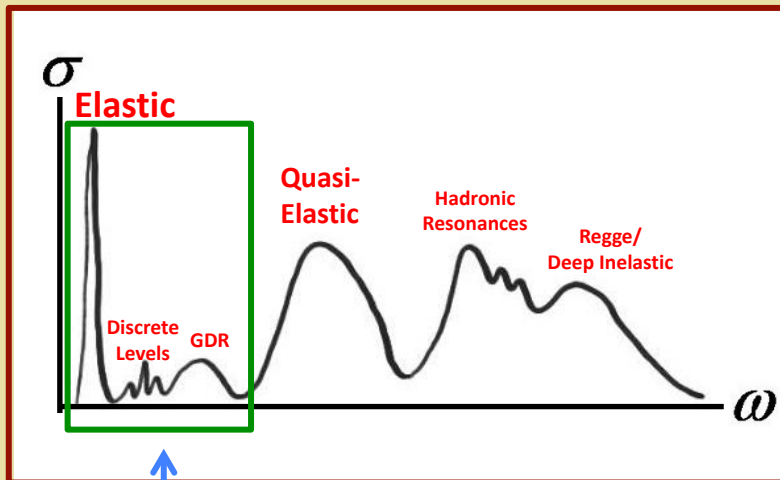
**Towner &  
Hardy**

**Features:**

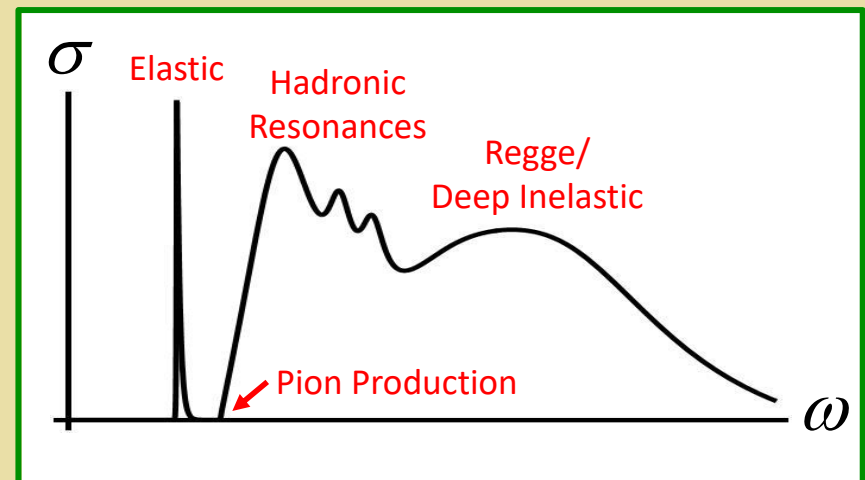
- $\text{Few} \times 10^{-4}$
- Error bar ?
- Refinements ?

# Other Nuclear Corrections

*Nuclei*



*Free nucleons*



*Low-lying transitions*

*Part of  $\delta_{NS}$*

**Challenges & opportunities for nuclear th'y**

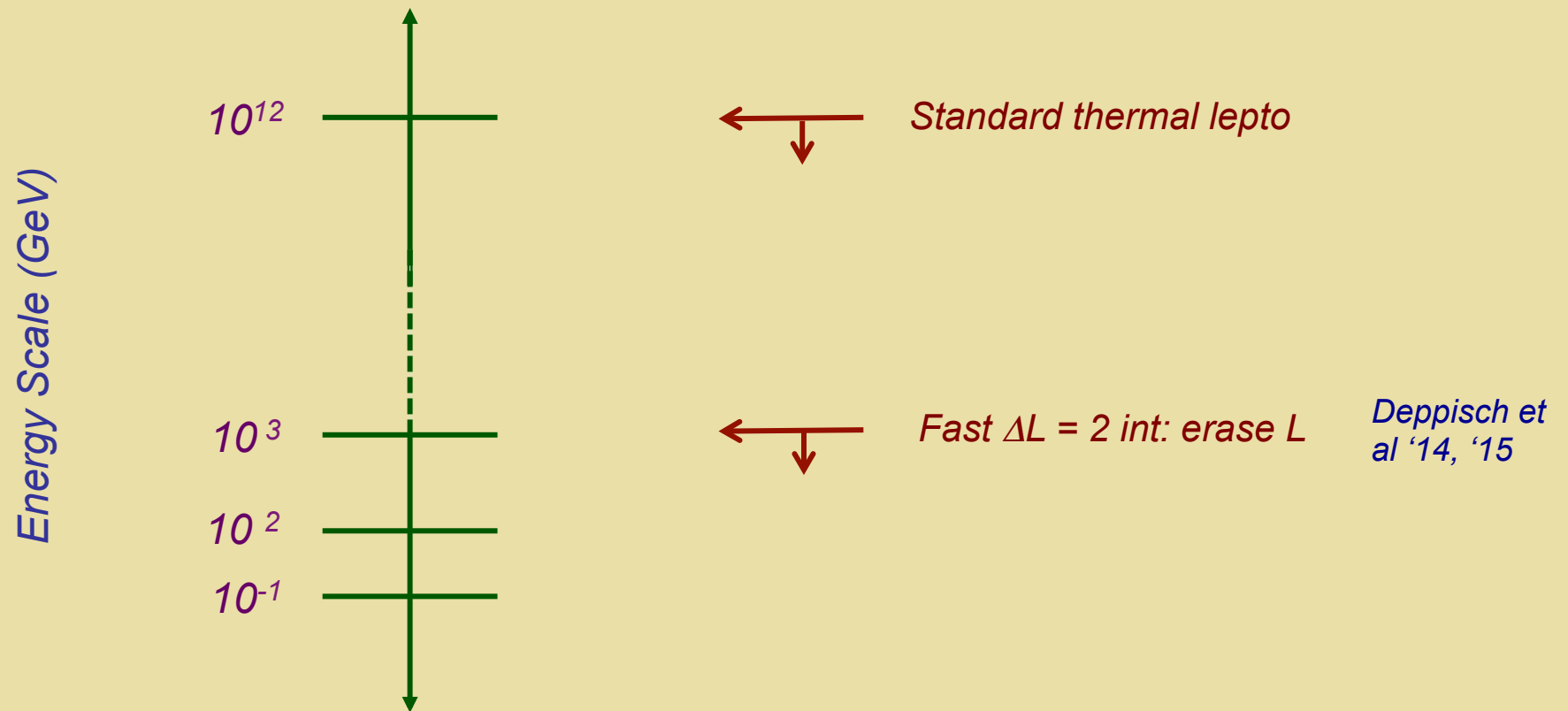
# Outlook

- *Fundamental questions motivate the search for physics beyond the Standard Model*
- *Tests of fundamental symmetries at low-energy are poised to*
  - *discover the BSM physics that answers several of these questions*
  - *determine its character*
- *Robust hadronic & nuclear computations plus high sensitivity experiments are essential*

# ***Back Up Slides***



# TeV LNV & Leptogenesis



# EDMs: New CPV?

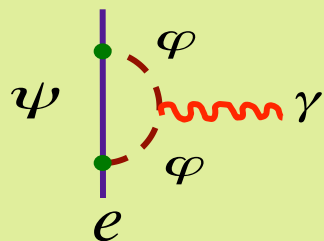
System	Limit (e cm)*	SM CKM CPV	BSM CPV
$^{199}\text{Hg}$	$7.4 \times 10^{-30}$	$10^{-35}$	$10^{-30}$
ThO	$8.7 \times 10^{-29} **$	$10^{-38}$	$10^{-29}$
n	$3.3 \times 10^{-26}$	$10^{-31}$	$10^{-26}$

\* 95% CL

\*\* e<sup>-</sup> equivalent

*New ACME:  $< 1.1 \times 10^{-29}$*

Mass Scale Sensitivity



$$\sin\phi_{\text{CP}} \sim 1 \rightarrow M > 5000 \text{ GeV}$$

$$M < 500 \text{ GeV} \rightarrow \sin\phi_{\text{CP}} < 10^{-2}$$

# EDMs: New CPV?

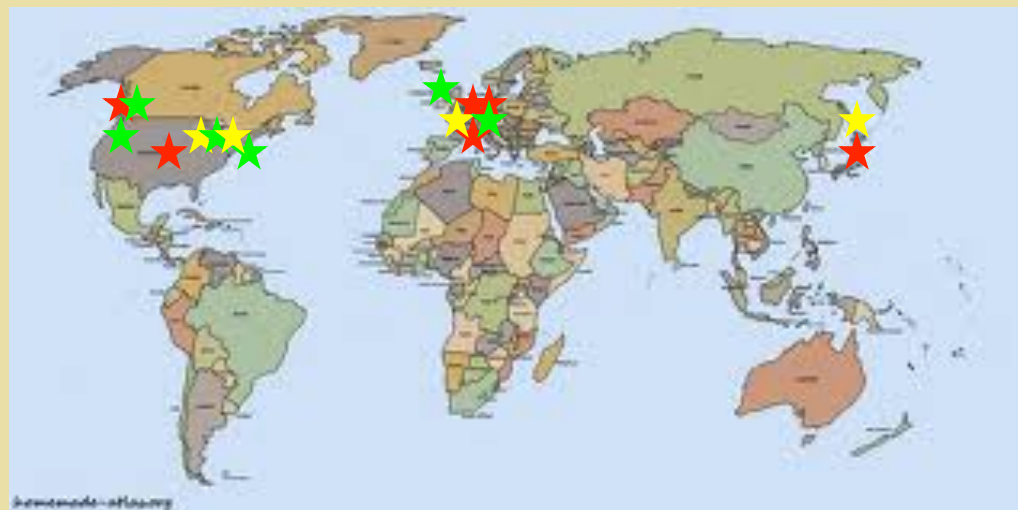
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n	$3.3 \times 10^{-26}$	$10^{-31}$	$10^{-26}$

\* 95% CL

\*\* e<sup>-</sup> equivalent

*New ACME:  $< 1.1 \times 10^{-29}$*

Not shown:  
muon



- ★ neutron
- ★ proton & nuclei
- ★ atoms

**~ 100 x better sensitivity**

# ***Future Tests***

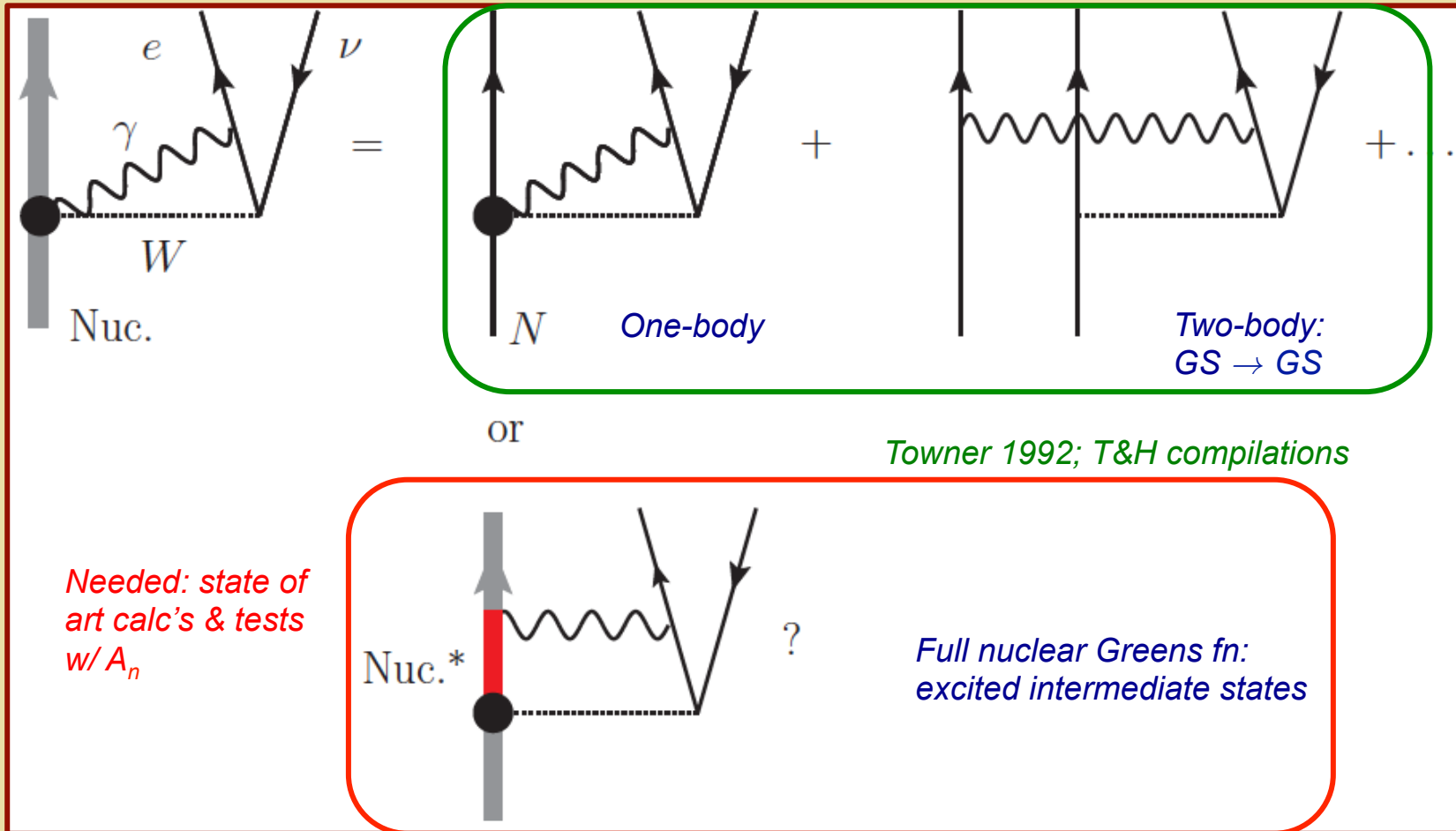
- *Lattice computation of  $M_3^{(0)}(Q^2)$*
- *PV electron scattering*

## ***Isospin relation***

$$4F_3^{(0)} = F_{3,\gamma Z}^p - F_{3,\gamma Z}^n$$

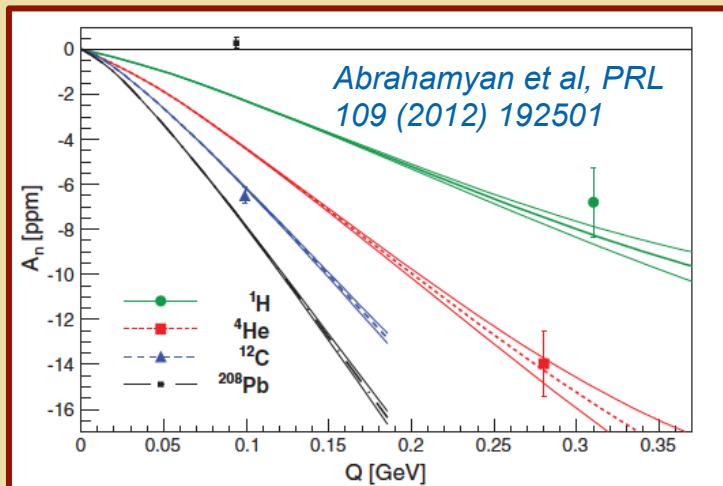
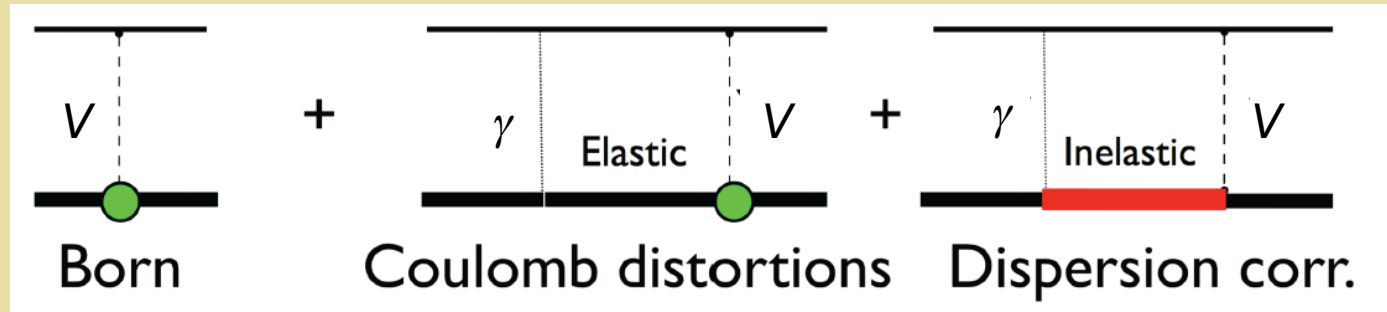
- ***SoLID ?***
- ***EIC ?***
- *More neutrino data for  $M_3^{(0)}(Q^2)$*

# $0^+ \rightarrow 0^+$ Decay: $\delta_{NS}$



# Dispersion Corrections

*Two-boson exchange in semileptonic processes: important for elastic PV eN & eA scattering ( $^{12}\text{C}$ ) & nuclear  $\beta$ -decay; beam normal asymmetry, Olympus... provide tests*

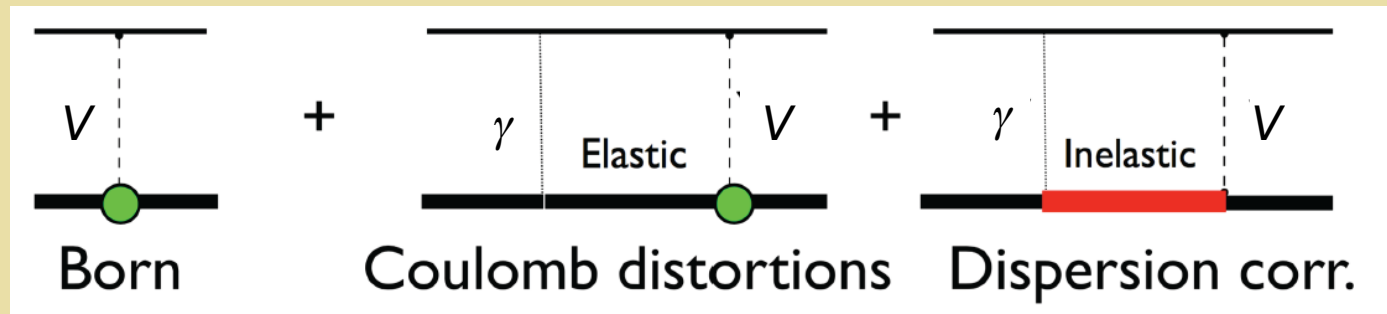


$V = \gamma$  Beam normal asymmetry

- J Lab Hall A
- Future: Mainz, J Lab

# Dispersion Corrections

*Two-boson exchange in semileptonic processes: important for elastic PV eN & eA scattering ( $^{12}\text{C}$ ) & nuclear  $\beta$ -decay; beam normal asymmetry provides, Olympus... provide tests*



$$V = Z^0, W, \gamma$$

*Important for  $O(0.1\%)$  probes of PV  $^{12}\text{C}(e, e')$  & superallowed  $\beta$ -decay*

$$V = \gamma$$

*Beam normal asymmetry*

$$V = Z^0, W$$

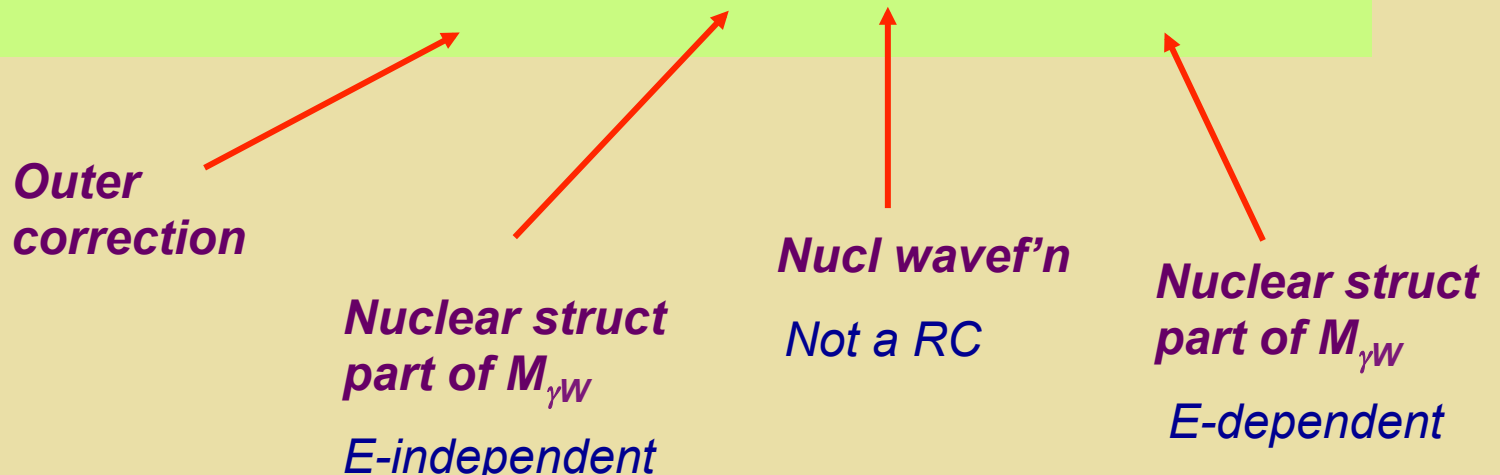
*Nucleus-dependent QED & EW corrections*

# Radiative Corrections & Ft Values

Corrected ft values:

$$\mathcal{F}t = ft (1 + \delta'_R) (1 + \delta_{NS} - \delta_C) (1 + \Delta_E^{NS})$$

Outer  
correction



Nuclear struct  
part of  $M_{\gamma W}$   
*E-independent*

Nucl wavef'n  
Not a RC

Nuclear struct  
part of  $M_{\gamma W}$   
*E-dependent*



# ***Radiative Corrections & $V_{ud}$***

## ***Superallowed***

$$|V_{ud}|^2 = \frac{2984.43s}{\mathcal{F}t(1 + \Delta_R^V)}$$

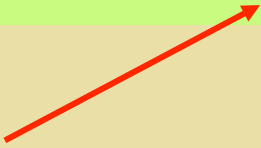
***Hadronic & short  
distance part of  $M_{\gamma W}$***



## ***Neutron***

$$|V_{ud}|^2 = \frac{5099.34s}{\tau_n(1 + 3\lambda^2)(1 + \Delta_R^V)}$$

***$(g_A/g_V)^2$***



***Contains  $\Delta_R^V$***

