





# Structure and Reactions of Light Nuclei within χEFT

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June 17, 2019

Neutron stars

#### **Neutrinoless Double Beta Decay**

In the hypothesis that the  $0\nu$ DBD is mediated by the exchange of a light neutrino:

$$[T_{1/2}^{0\nu}]^{-1} = G^{0\nu}(Q, Z) [M^{0\nu}]^2 m_{\beta\beta}^2$$

Javier Menendez arXiv:1703.08921 (2017)

#### Lepton space-phase integral

- Depends on the Q-value of the decay and the charge of the final state of the nucleus
- Can be calculated precisely: for most of the emitters of interest

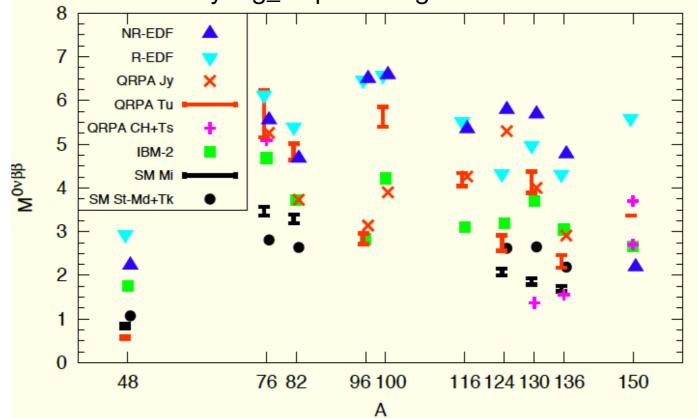
 $10^{-15} - 10^{-16} \mathrm{yr}^{-1}$ 

#### Nuclear matrix element (NME)

- Open issues for theorists
- Spread of about a factor 2-3 in the predicted values for NME for a given isotope
- Theoretical predictions for these models compared with single beta decays: g\_A quenching

#### Effective Majorana mass

- Depends on combination of neutrino masses and oscillation parameters
- Uncertainties in the parameters extracted by oscillation experiments and cosmology



#### The *basic model* of nuclear theory

The *basic model* of nuclear theory: description of the static and dynamic properties of nuclear systems.

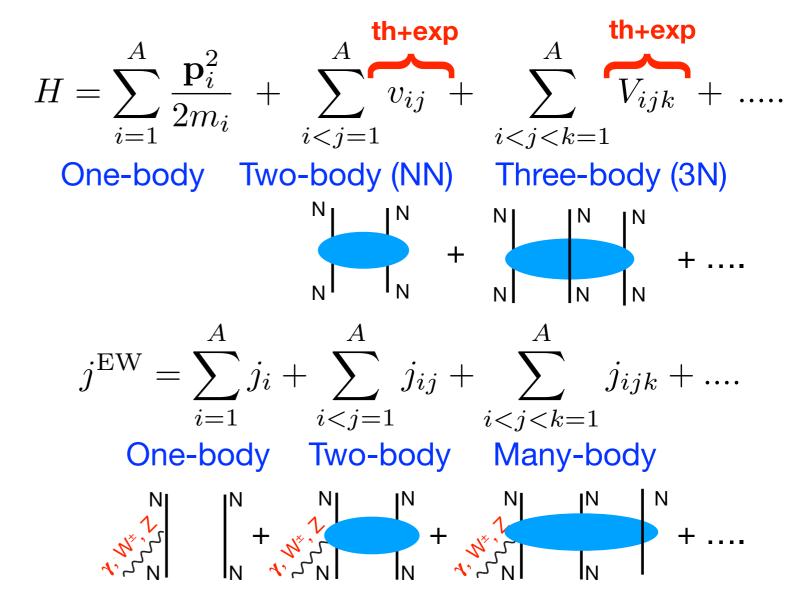
Nucleon-nucleon (NN) scattering data: "thousands" of experimental data available

The spectra, properties, and transition of nuclei: BE, radii, magnetic moments, beta decays rates, weak/radiative captures, electroweak form factors, etc. The nucleonic matter equation of state: for ex. EOS neutron matter

Inputs for the basic model:

Many-body interactions between the constituents

Electroweak current operators:



#### Chiral EFT: from QCD to nuclear systems

S. Weinberg, Phys. Lett. **B251**, 288 (1990); Nucl. Phys. **B363**, 3 (1991); Phys. Lett **B295**, 114 (1992)

QCD



Symmetries in particular the approximate chiral symmetry between hadronic d.o.f ( $\pi$ , N,  $\Delta$ )

Approximate chiral symmetry requires the pion to couple to other pions and to baryons by powers of its momentum

Effective chiral Lagrangian  $\mathcal{L}_{eff}(\pi, N, \Delta)$ 

Calculate amplitudes+prescription to obtain potentials + regularization (of high momentum components)

$$\mathcal{L}_{eff} = \mathcal{L}^{(0)} + \mathcal{L}^{(1)} + \mathcal{L}^{(2)} + \dots$$

Given a power counting scheme

 $\mathcal{L}^{(n)} \sim \left(\frac{Q}{\Lambda_{\chi}}\right)^n \sim 100 \text{ MeV soft scale} \\ \sim 1 \quad \text{GeV hard scale}$ 

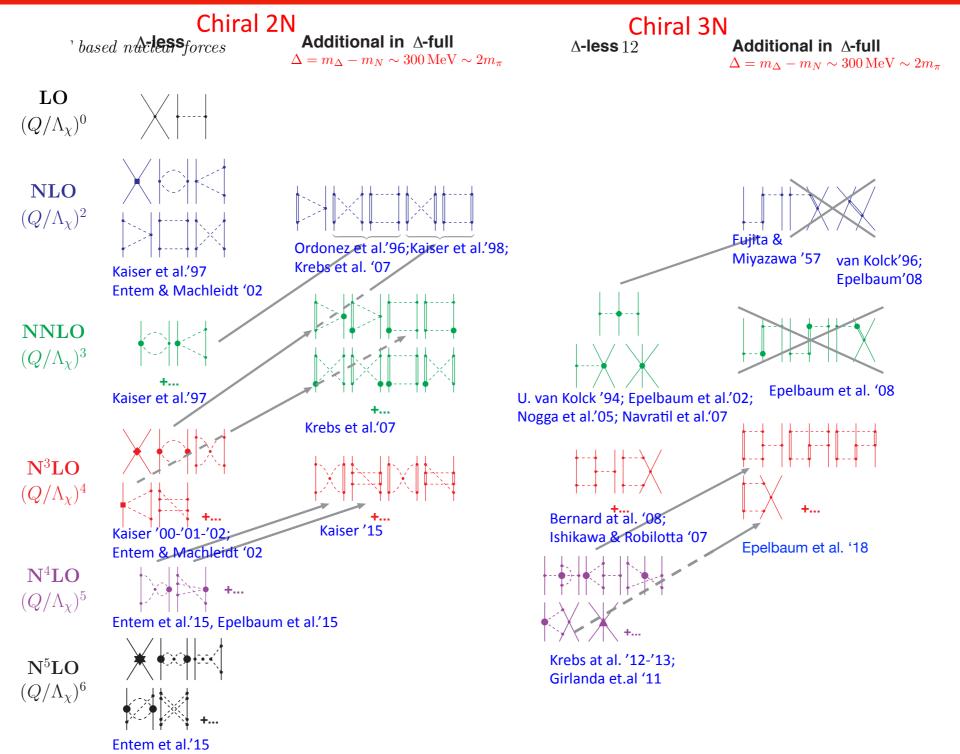
Nuclear forces and currents

Few- and many-body methods: QMC, NCSM, CC, etc



Nuclear structure and dynamics

#### Nuclear Hamiltonian: Chiral EFT formulation of the basic model



<u>Note:</u> Many of the available versions of chiral potentials are formulated in momentumspace and are strongly nonlocal:  $\Rightarrow \mathbf{p} \rightarrow -i \nabla$  hard to use in QMC methods

Gezerlis et al. PRL 111, 032501 2013; PRC 90, 054323 2014; Lynn et al. PRL 113, 192501 2014 Piarulli et al. PRC 91, 024003 2015; PRC 94, 054007 2016

 $\mathbf{p}'$ 

 $\mathbf{p}$ 

#### Local chiral NN potential with $\Delta$ 's

Piarulli et al. PRC **91**, 024003 2015; PRC **94**, 054007 2016  $v_{12} = v_{12}^{\text{EM}} + v_{12}^{\text{L}} + v_{12}^{\text{S}}$ 

 $v_{12}^{\text{EM}}$ : EM component including corrections up to  $\alpha^2$ LO:  $Q^0 \left\| \cdot \cdot \cdot \right\|^{-p'}$ 

 $v_{12}^{ extsf{L}}$  : chiral OPE and TPE component with  $extsf{\Delta}$ 's

dependence only on the momentum transfer k=p'-p

 $v_{12}^{
m S}$  : short-range contact component up to order N3LO (Q4) parametrized by (2+7+11) CI and (2+4) IB LECs

NLO:  $Q^2$   $\overrightarrow{\mathcal{A}}^{*}$   $\overrightarrow{\mathcal{A}}^{*}$   $\overrightarrow{\mathcal{A}}^{*}$   $\overrightarrow{\mathcal{A}}^{*}$ 

N2LO:  $Q^3$ 

• the functional form taken as  $C_{R_S}(r) \propto e^{-(r/R_S)^2}$  with  $R_S = 0.8~(0.7)~{
m fm}$  a (b) models

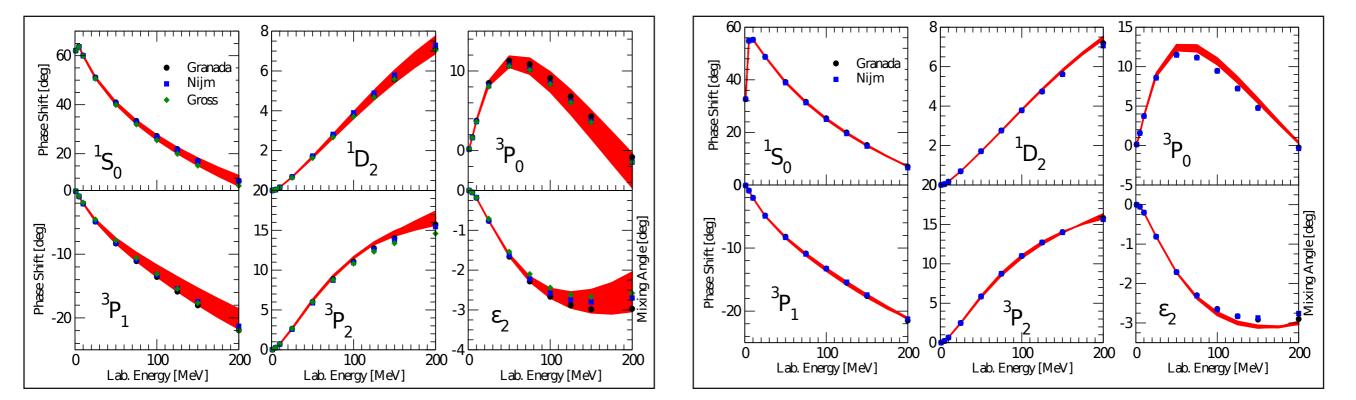
In coordinate-space it reads as:

$$\begin{aligned}
 & V_{12} = \sum_{l=1}^{16} v^l(r) O_{12}^l \\
 & O_{12}^{l=1,...,6} = [\mathbf{1}, \, \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2, \, S_{12}] \otimes [\mathbf{1}, \, \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2] \\
 & O_{12}^{l=7,...,11} = \mathbf{L} \cdot \mathbf{S}, \, \mathbf{L} \cdot \mathbf{S} \, \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2, \, (\mathbf{L} \cdot \mathbf{S})^2, \, \mathbf{L}^2, \, \mathbf{L}^2 \, \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 \\
 & O_{12}^{l=12,...,16} = T_{12}, \, (\boldsymbol{\tau}_1^z + \boldsymbol{\tau}_2^z), \, \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 \, T_{12}, \, S_{12} \, T_{12}, \, \mathbf{L} \cdot \mathbf{S} \, T_{12}
 \end{aligned}$$

#### Fitting procedure: NN PWA and database

The 26 LECs are fixed by fitting the pp and np Granada database up to two ranges of  $E_{lab} = 125$  MeV and 200 MeV, the deuteron BE and the nn scattering length

To minimizing the  $\chi^2$  we have used the Practical Optimization Using No Derivatives (for Squares), POUNDers (M. Kortelainen, PRC 82, 024313 2010)

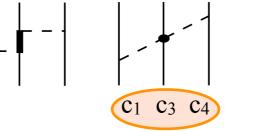


model	order	$E_{\rm Lab}~({\rm MeV})$	$N_{pp+np}$	$\chi^2/{ m datum}$
Ia	N3LO	$0\!\!-\!\!125$	2668	1.05
Ib	N3LO	$0\!\!-\!\!125$	2665	1.07
IIa	N3LO	0–200	3698	1.37
IIb	N3LO	0–200	3695	1.37

Models a (b) cutoff~500 MeV (600 MeV) in momentum-space

#### Local chiral 3N potential with $\Delta$ 's

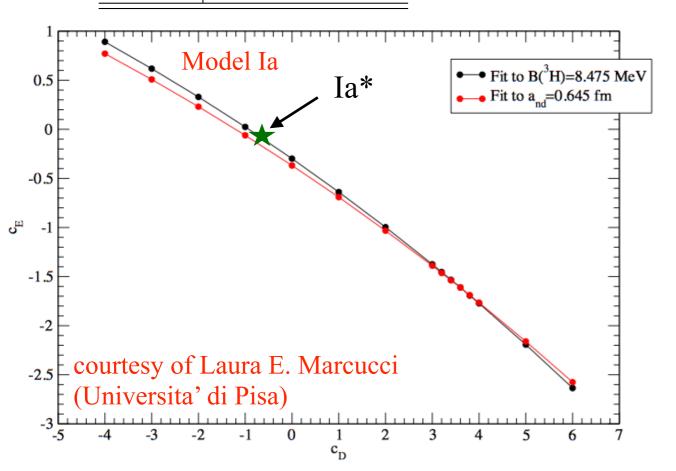
Inclusion of 3N forces at N2LO:



#### 1) Fit to:

- $\blacktriangleright E_0(^{3}\text{H}) = -8.482 \text{ MeV}$
- $a_{nd} = (0.645 \pm 0.010) \text{ fm}$

Model	$c_D$	$c_E$
Ia	3.666	-1.638
Ib	-2.061	-0.982
IIa	1.278	-1.029
IIb	-4.480	-0.412



2) Fit to:

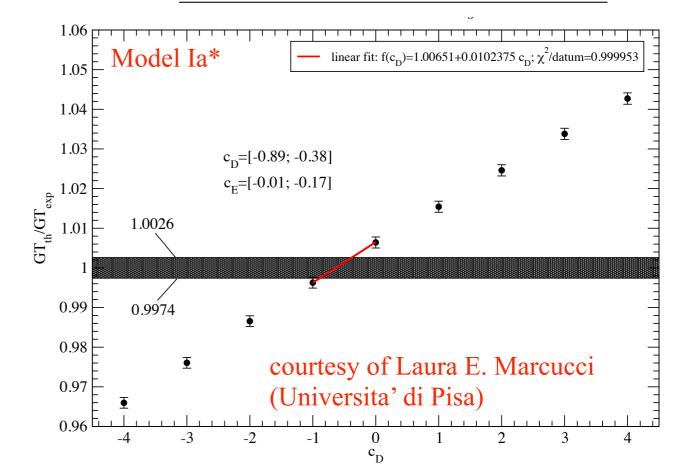
CD

 $\blacktriangleright E_0(^{3}\text{H}) = -8.482 \text{ MeV}$ 

► GT m.e. in <sup>3</sup>H  $\beta$ -decay

 $C_{E} \sim \tau_{i} \cdot \tau_{j}$ 

Model	$c_D$	$c_E$
Ia*	-0.635(255)	-0.09(8)
$Ib^*$	-4.705(285)	0.550(150)
IIa*	-0.610(280)	-0.350(100)
IIb*	-5.250(310)	0.05(180)

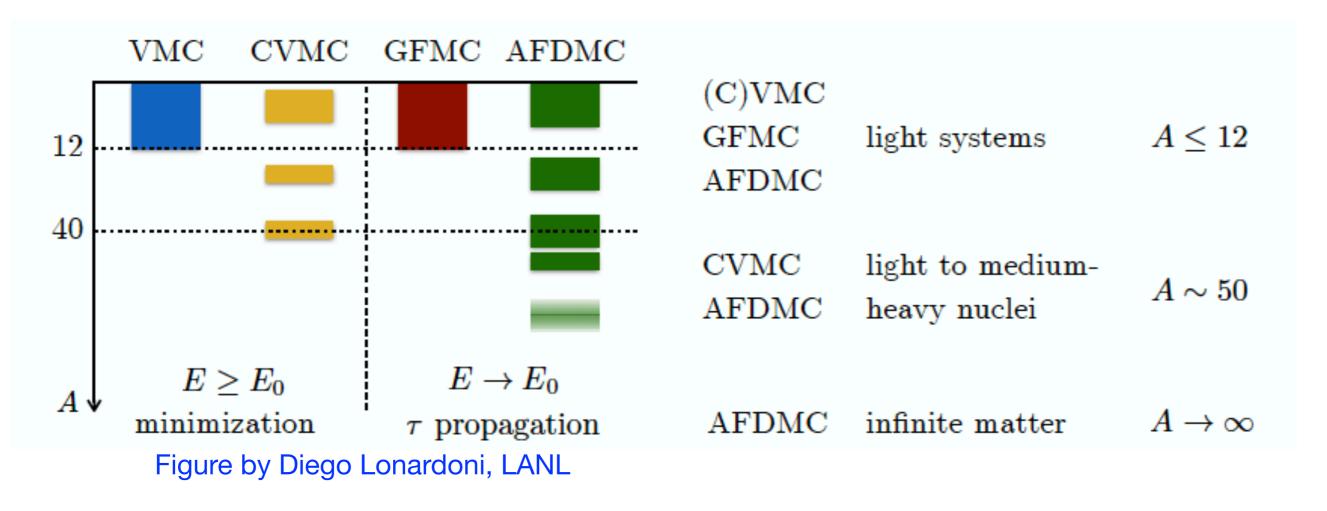


#### Ab initio Methods: HH and QMC

Hyperspherical Harmonics (HH) expansion for A=3 and 4 bound and continuum states



Quantum Monte Carlo (QMC) methods encompass a large family of computational methods whose common aim is the study of complex quantum systems



#### QMC: Variational Monte Carlo (VMC)

R.B. Wiringa, PRC 43, 1585 (1991)

Minimize the expectation value of *H*:

Trial wave function (involves variational parameters):

$$E_T = \frac{\langle \Psi_T | H | \Psi_T \rangle}{\langle \Psi_T | \Psi_T \rangle} \ge E_0$$
$$\Psi_T \rangle = \left[ 1 + \sum_{i < j < k} U_{ijk} \right] \left[ S \prod_{i < j} (1 + U_{ij}) \right]$$

 $|\Psi_J
angle$ 

 $|\Psi_J\rangle = \left[\prod_{i < j} f_c(r_{ij})\right] |\Phi(JMTT_z)\rangle$  (s-shell nuclei): Jastrow wave function, fully antisymmetric  $S \prod_{i < j}$ : represents a symmetrized product

$$U_{ij} = \sum_{p=2,6} u_p(r_{ij}) O_{ij}^p$$
: pair correlation operators

$$U_{ijk} = \sum_{x} \epsilon_x V_{ijk}^x \text{ : three-body correlation operators}$$
$$|\Psi_T\rangle \text{ are spin-isospin vectors in 3A dimension with } 2^A \begin{pmatrix} A \\ Z \end{pmatrix}$$

The search in the parameter space is made using COBYLA (Constrained Optimization BY Linear Approximations) algorithm available in NLopt library

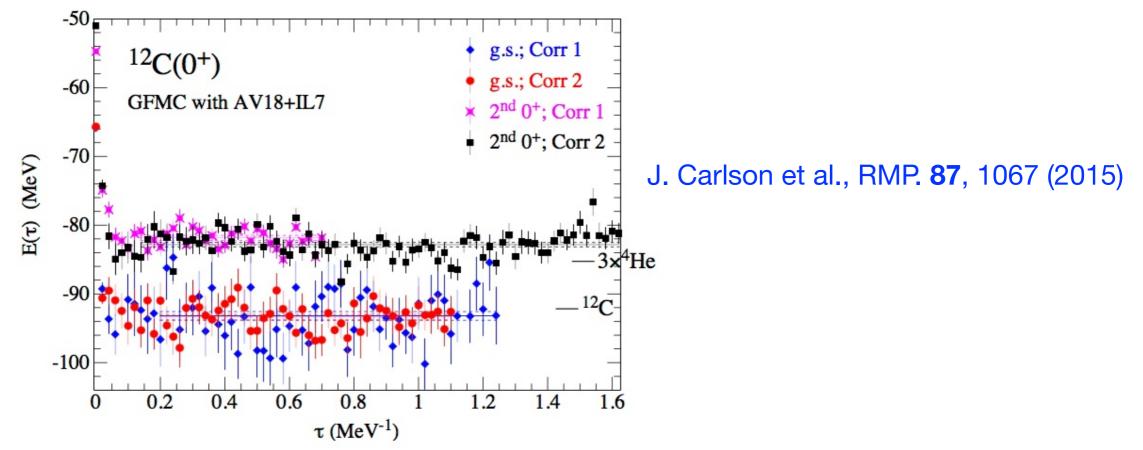
The diffusion Monte Carlo (DMC) method (ex. GFMC or AFDMC) overcomes the limitations of VMC by using a projection technique to determine the true ground-state

The method relies on the observation that  $\Psi_T$  can be expanded in the complete set of eigenstates of the Hamiltonian according to

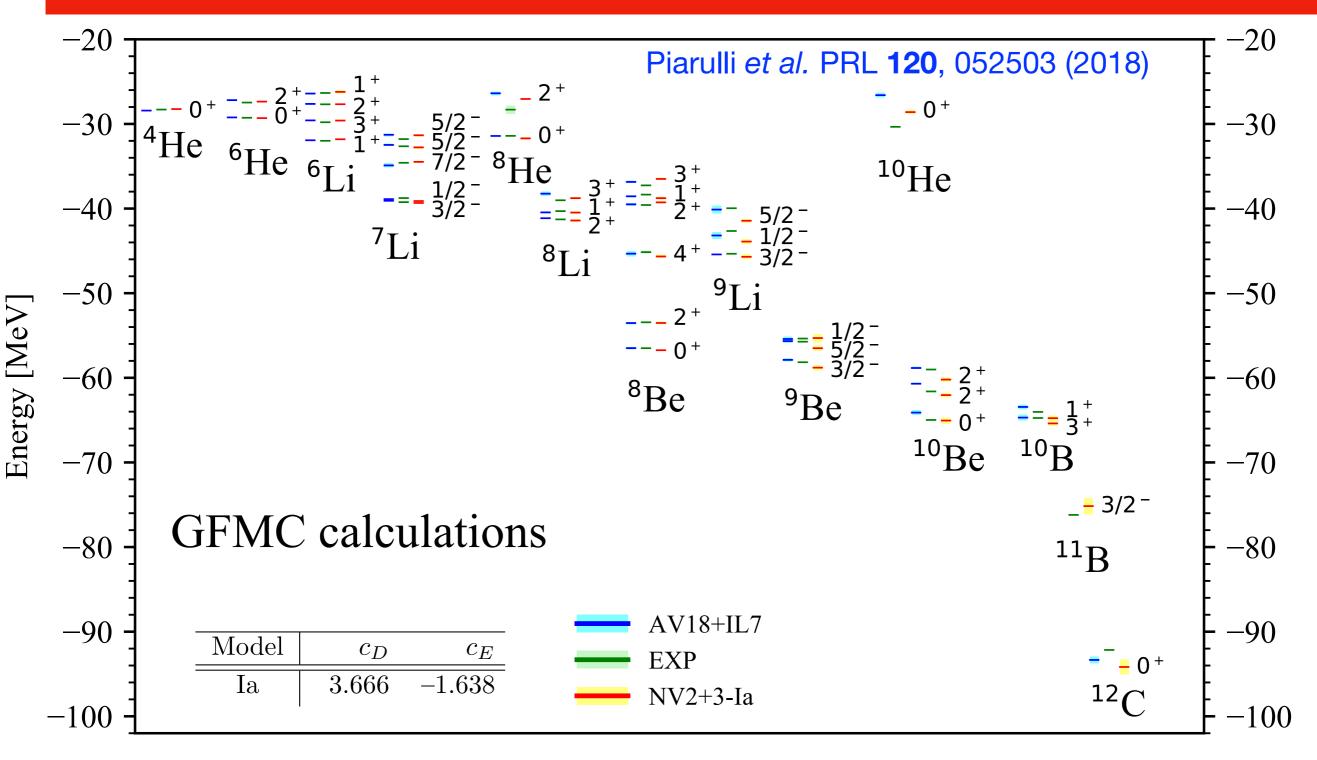
$$\begin{aligned} |\Psi_T\rangle &= \sum_n c_n |\Psi_n\rangle & H|\Psi_n\rangle = E_n |\Psi_n\rangle \\ \lim_{\tau \to \infty} |\Psi(\tau)\rangle &= \lim_{\tau \to \infty} e^{-(H - E_0)\tau} |\Psi_T\rangle = c_0 |\Psi_0\rangle & |\Psi(\tau = 0)\rangle = |\Psi_T\rangle \end{aligned}$$

where  $\boldsymbol{\tau}$  is the imaginary time

The evaluation of  $\Psi(\tau)$  is done stochastically in small time steps  $\Delta \tau$  ( $\tau = n \Delta \tau$ ) using a Green's function formulation

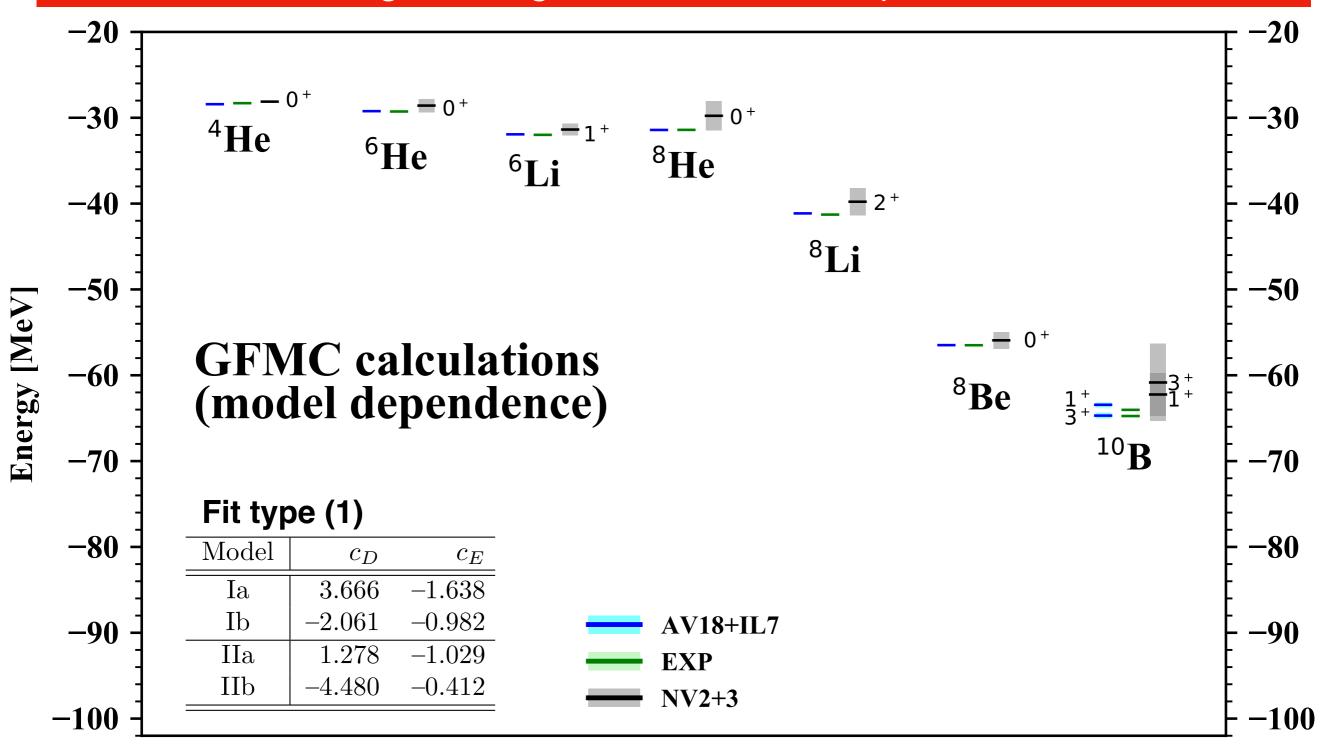


Spectra of Light Nuclei: Phenomenology vs  $\chi$ EFT



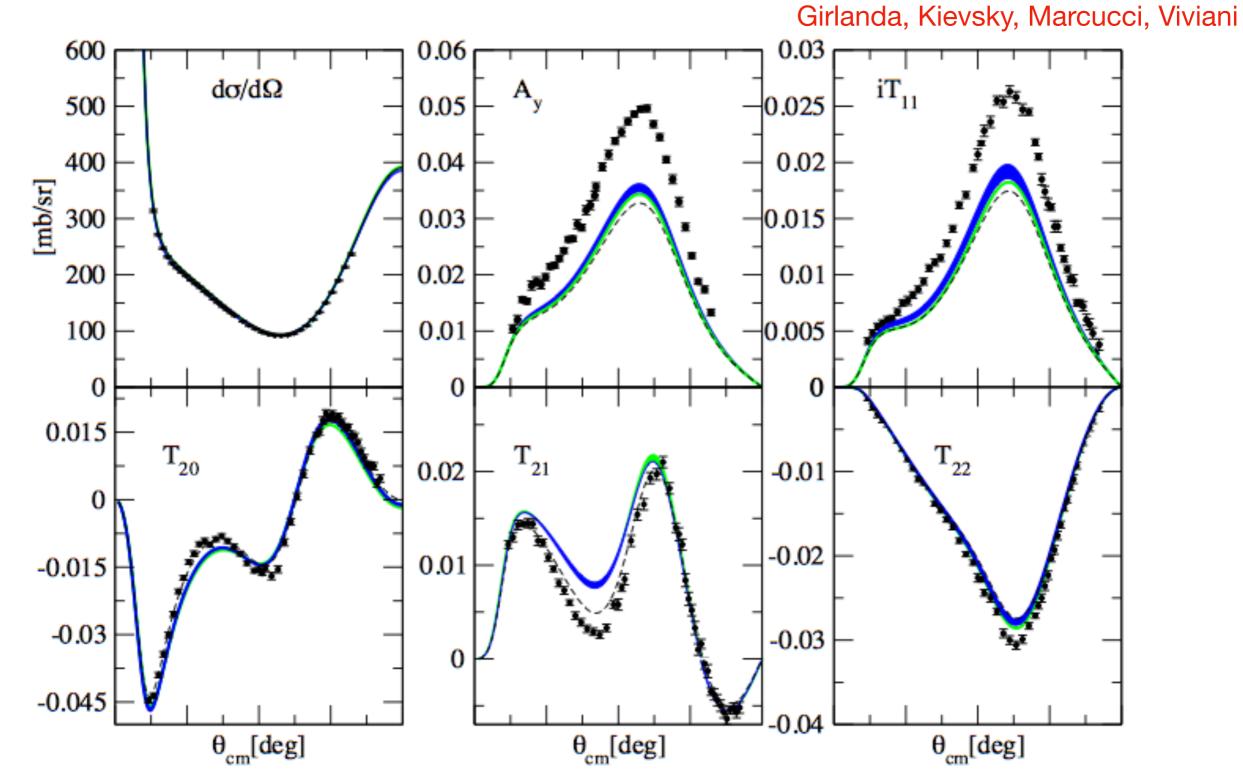
The rms from experiment is 0.72 MeV for NV2+3-Ia compared to 0.80 MeV for AV18+IL7  $c_E < (>)0$ : repulsion (attraction) in light-nuclei (the opposite effect in PNM)  $c_D < (>)0$ : repulsion (attraction) in light-nuclei (same effect in PNM but very small)

#### **Energies of Light Nuclei: Model-dependence**



Model-dependence for NV2+3 up to 7-8% of the total binding energy

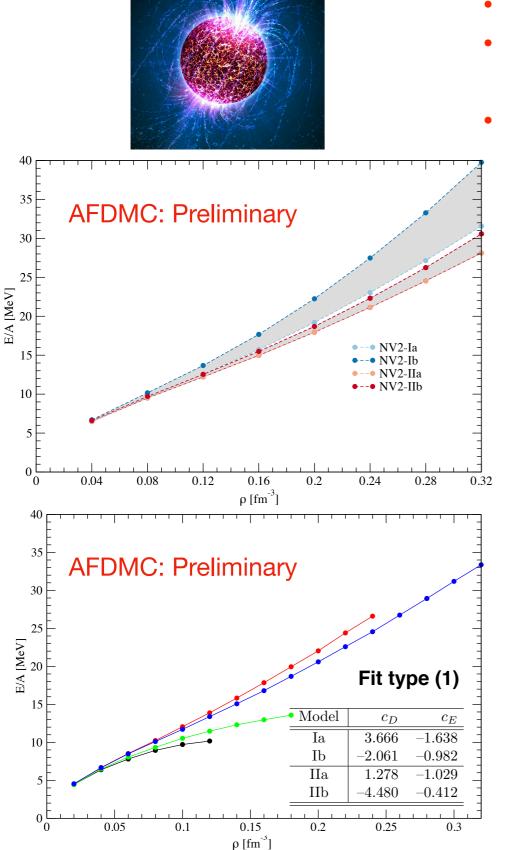
 $c_E < (>)0$ : repulsion (attraction) in light-nuclei (the opposite effect in PNM)  $c_D < (>)0$ : repulsion (attraction) in light-nuclei (same effect in PNM but very small) Polarization observables in pd elastic scattering at 3 MeV: HH calculations with the NV2+3 models Ia-Ib (IIa-IIb), are shown by the green (blue) band. The black dashed line are results obtained with only the two-body interaction NV2-Ia



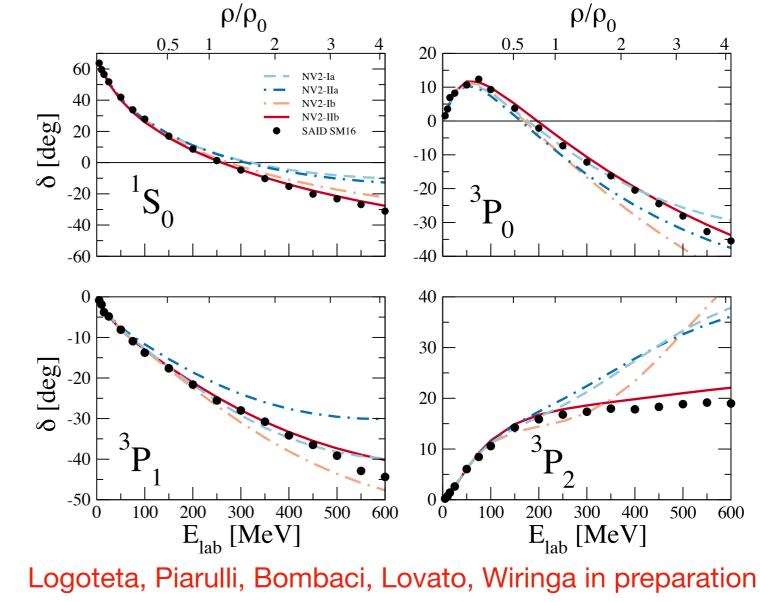
More sophisticated 3N force??? Different way to fix the 3N??? subleading contact terms in 3N interaction???

#### Equation of State of Pure Neutron Matter in $\chi$ EFT

#### The EoS of pure neutron matter (PNM): neutrons stars



- Compact objects: R ~ 10km,  $M_{\rm max}^{\rm obs} \sim 2 M_{\odot}$
- Composed predominantly of neutrons between the inner crust and the outer core
- NS from gravitational collapse of a massive star after a supernova explosion



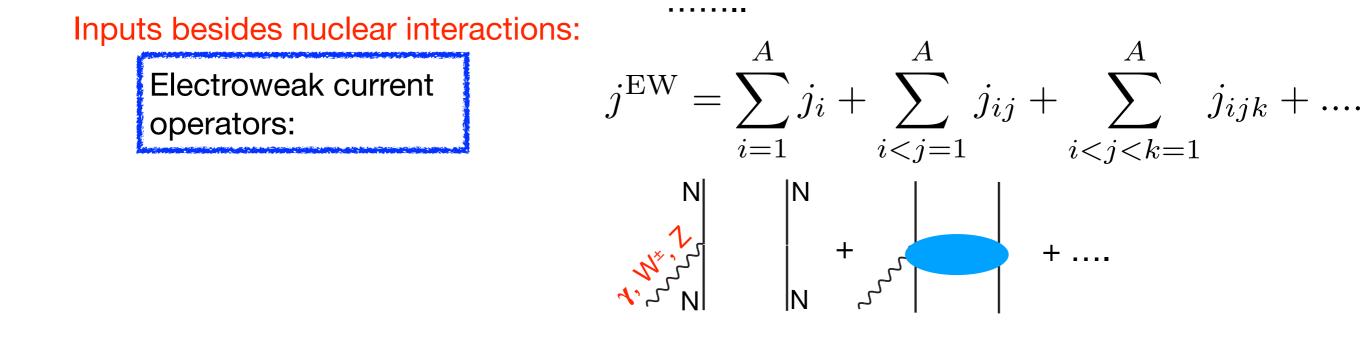
Cutoff sensitivity: modest in NV2 models; very large in NV2+3 models

#### **Beyond Energy Calculations**

Electroweak structure and reactions:

Electroweak form factors Magnetic moments and radii Electroweak Response functions Radiative/weak captures

G.T. matrix elements involved in beta decays



Current operators constructed in correspondence to the phenomenological interactions based on meson-exchange approach Marcucci *et al.* PRC **72**, 014001 (2005)

Current operators derived in  $\chi$ EFT: Pastore *et al.* PRC **78**, 064002 (2008), PRC **80**, 034004 (2009); Piarulli *et al.* PRC **87**, 014006 (2013), Baroni *et al.* PRC **93**, 015501 (2016); Kölling *et al.* PRC **86**, 047001 (2012), Krebs et al., Ann. Phys. **378**, 317 (2017)

#### Nuclear axial currents and beta-decays in light-nuclei

Matrix Element  $< \Psi_f |GT| \Psi_i > \sim g_A$  and decay rate  $\sim g_A^2$ 

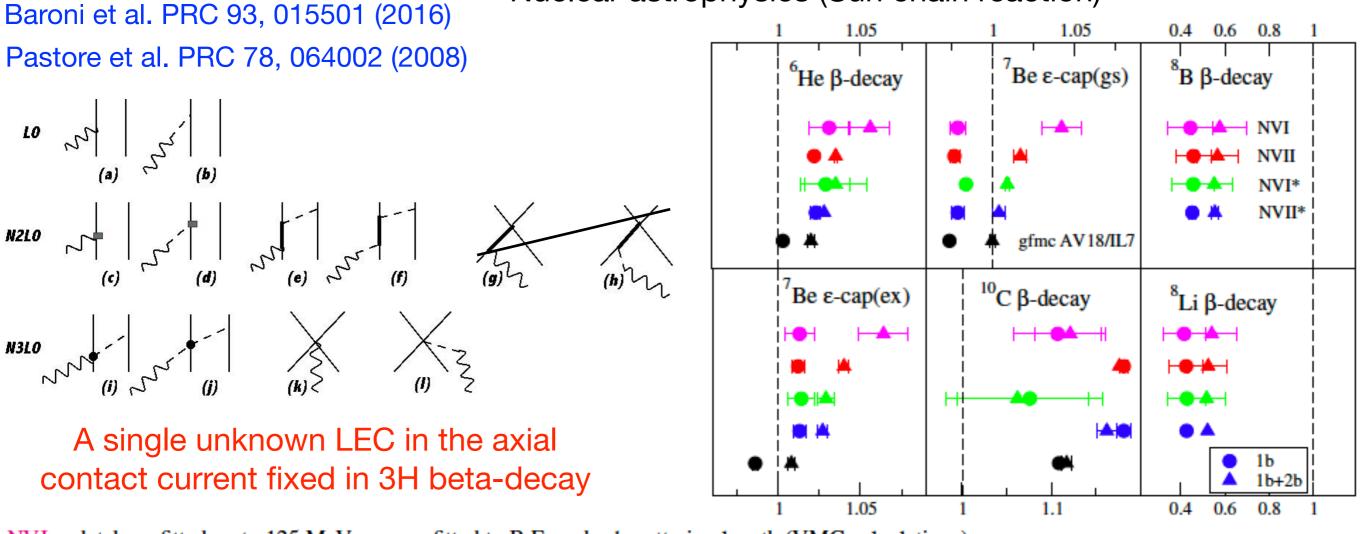
 $(Z,N) \rightarrow (Z+1,N-1) + e + \bar{v}_e$ 

Schiavilla et al. PRC 99, 034005 (2019)

Understanding "quenching" of  $\sim g_A$ 

Relevant for neutrinoless double beta decay since rate  $\sim g_A^4$ 

Nuclear astrophysics (Sun chain reaction)



NVI - database fitted up to 125 MeV - c<sub>D</sub>, c<sub>E</sub> fitted to B.E. and nd-scattering length (VMC calculations)
 NVII - database fitted up to 200 MeV - c<sub>D</sub>, c<sub>E</sub> fitted to B.E. and nd-scattering length (VMC calculations)
 NVI\* - database fitted up to 125 MeV - c<sub>D</sub>, c<sub>E</sub> fitted to B.E. and GT triton (VMC calculations)
 NVII\* - database fitted up to 200 MeV - c<sub>D</sub>, c<sub>E</sub> fitted to B.E. and GT triton (VMC calculations)

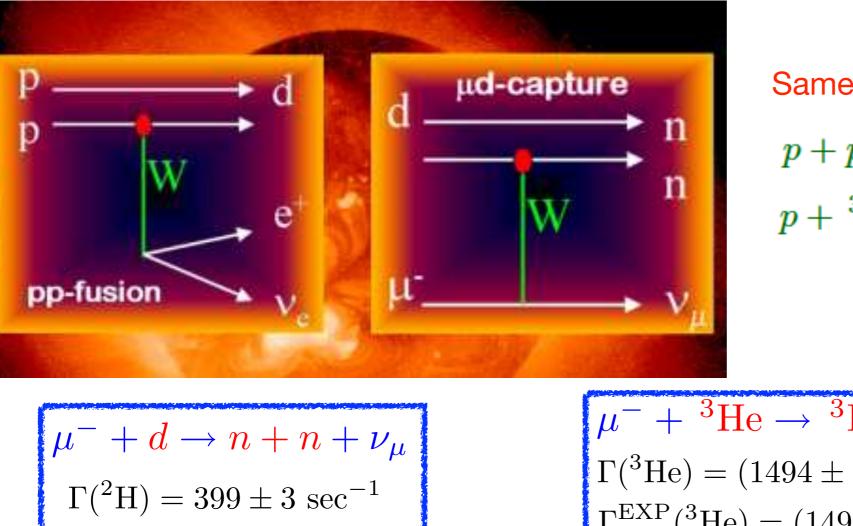
Pastore, Piarulli, Schiavilla, Wiringa, Baroni, Carlson, Gandolfi, in preparation

#### PRELIMINARY

AV18+IL7 - database fitted up to 350 MeV - c<sub>D</sub> fitted to GT triton (GFMC calculations) Pastore et al. PRC 97 022501 (2018)

#### $\mu^{-}$ Capture on <sup>2</sup>H and <sup>3</sup>He in $\chi$ EFT

#### http://www.npl.illinois.edu/exp/musun/



Same theoretical inputs of:

$$p + p \rightarrow d + e^+ + \nu_e$$
  
 $n + {}^{3}\text{He} \rightarrow {}^{4}\text{He} + e^+ + \nu_e$  (her)

$$\mu^{-} + d \rightarrow n + n + \nu_{\mu}$$
$$\Gamma(^{2}\text{H}) = 399 \pm 3 \text{ sec}^{-1}$$

$$\mu^{-} + {}^{3}\text{He} \rightarrow {}^{3}\text{H} + \nu_{\mu} (70\%)$$
  

$$\Gamma({}^{3}\text{He}) = (1494 \pm 21) \text{ sec}^{-1}$$
  

$$\Gamma^{\text{EXP}}({}^{3}\text{He}) = (1496 \pm 4) \text{ sec}^{-1}$$

Good agreement with available experimental data although the muon-capture on the deuteron have large errors

•Upcoming measurement of  $\Gamma(^{2}H)$  by the MuSun collaboration at PSI with a 1% error

Two-body currents important to achieve this agreement

Marcucci & Piarulli FBS **49**, 35-39 (2011) Marcucci et al. PRL 108, 052502 (2012)

Marcucci et al. PRC 83, 014002 (2011)

#### Conclusions

We are testing our models of NN+3N interactions with  $\Delta$ -isobar based on chiral EFT framework in both light-nuclei and infinite nuclear matter

We mainly focused our attention on studying properties of nuclei up to A=12 and EoS of infinite neutron matter

For the time being, we are interested in studying the model-dependence of the nuclear observables by exploring different cutoffs and range of energies used to fit the NN interactions as well as analyzing different strategies fo fit the TNI

It looks like that the formulation of the TNI with only  $c_D$  and  $c_E$  terms is too simplistic if we want to have a good descriptions of spectra, properties of light-nuclei, infinite nuclear matter, three-body observables with a certain degree of accuracy

We are investigating the effect of subleading 3N contact interactions in light-nuclei (we will do so also for infinite nuclear matter)

## THANK YOU

### Washington University in St. Louis







Alessandro Baroni, University of South Carolina, USA Joe Carlson, Los Alamos National Lab, USA Stefano Gandolfi, Los Alamos National Lab, USA Luca Girlanda, University of Salento, Italy Alejandro Kievsky, INFN-Pisa, Italy Alessandro Lovato, INFN-Trento, Italy Laura E. Marcucci, INFN-Pisa, University of Pisa, Italy Saori Pastore, Washington University in St. Louis, USA Steven Pieper\*\*, Argonne National Lab, USA (\*\*deceased) Rocco Schiavilla, Old Dominion University/Jefferson Lab, USA Michele Viviani, INFN-Pisa, Italy Robert Wiringa, Argonne National Lab, USA

NUCE LEI Nuclear Computational Low-Energy Initiative



