



Hierarchical Bayesian Unfolding in the Oslo Method

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ISNET-11 — Information and Statistics in Nuclear Experiment and
Theory

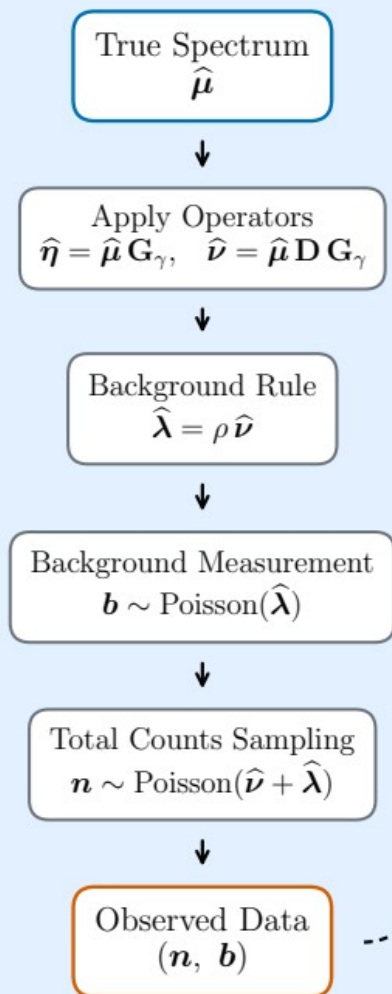
ECT*, Trento · Nov 17–21, 2025

Overview

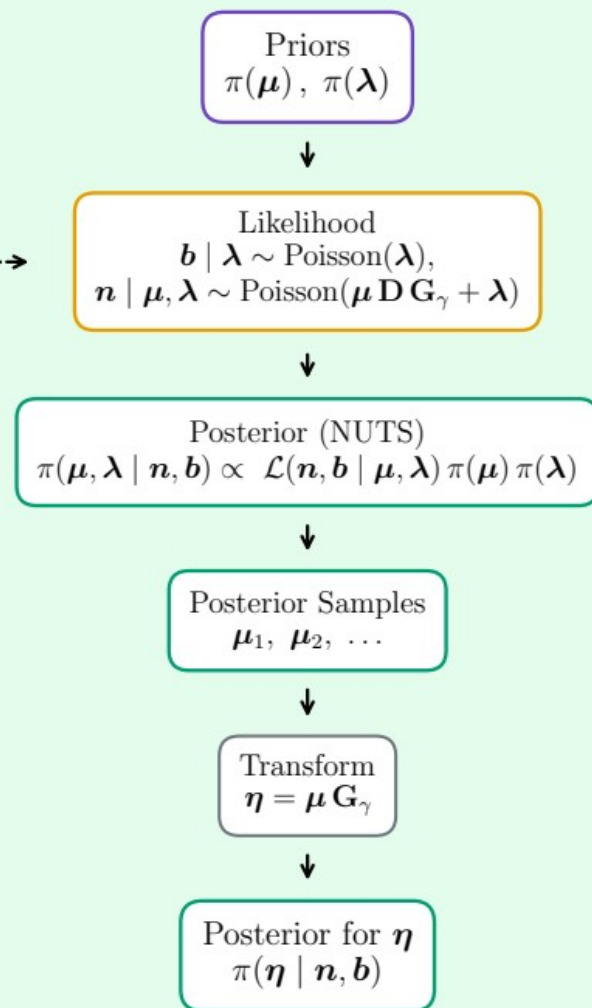
- The unfolding problem
- Hierarchical Bayesian Model
- Unfolding results and validity

Bayesian Unfolding Framework (Signal + Background)

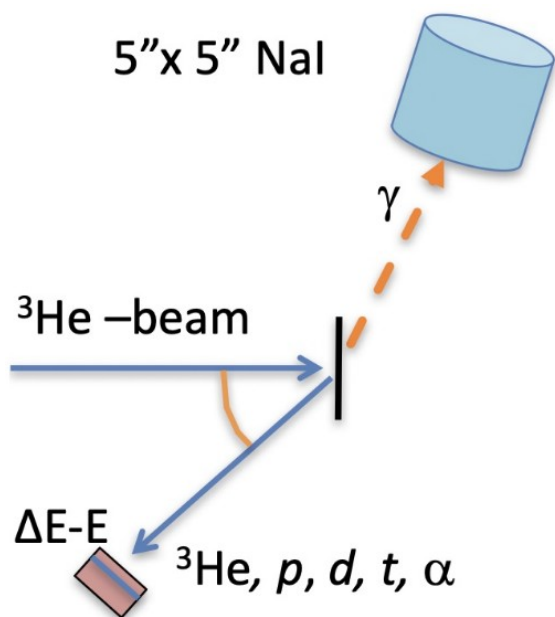
Forward Model (Synthetic Data Generation)



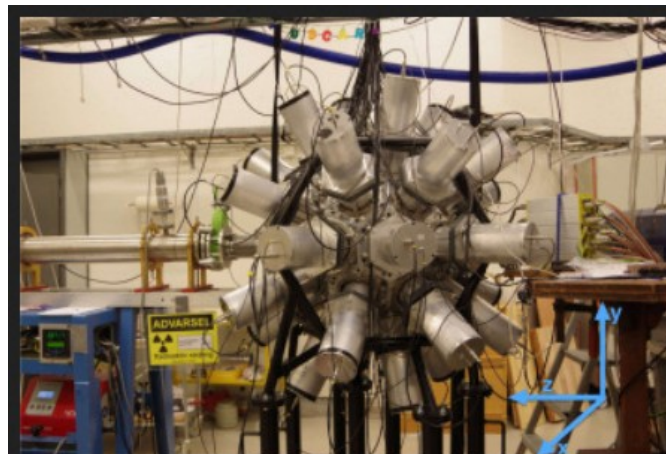
Statistical Inference (Bayesian)



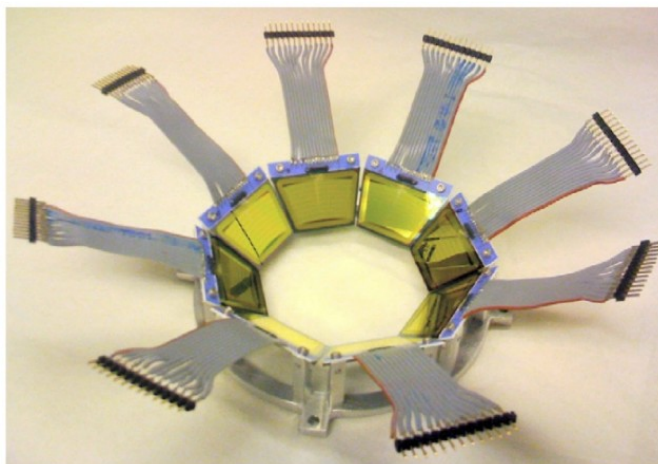
Experimental setup



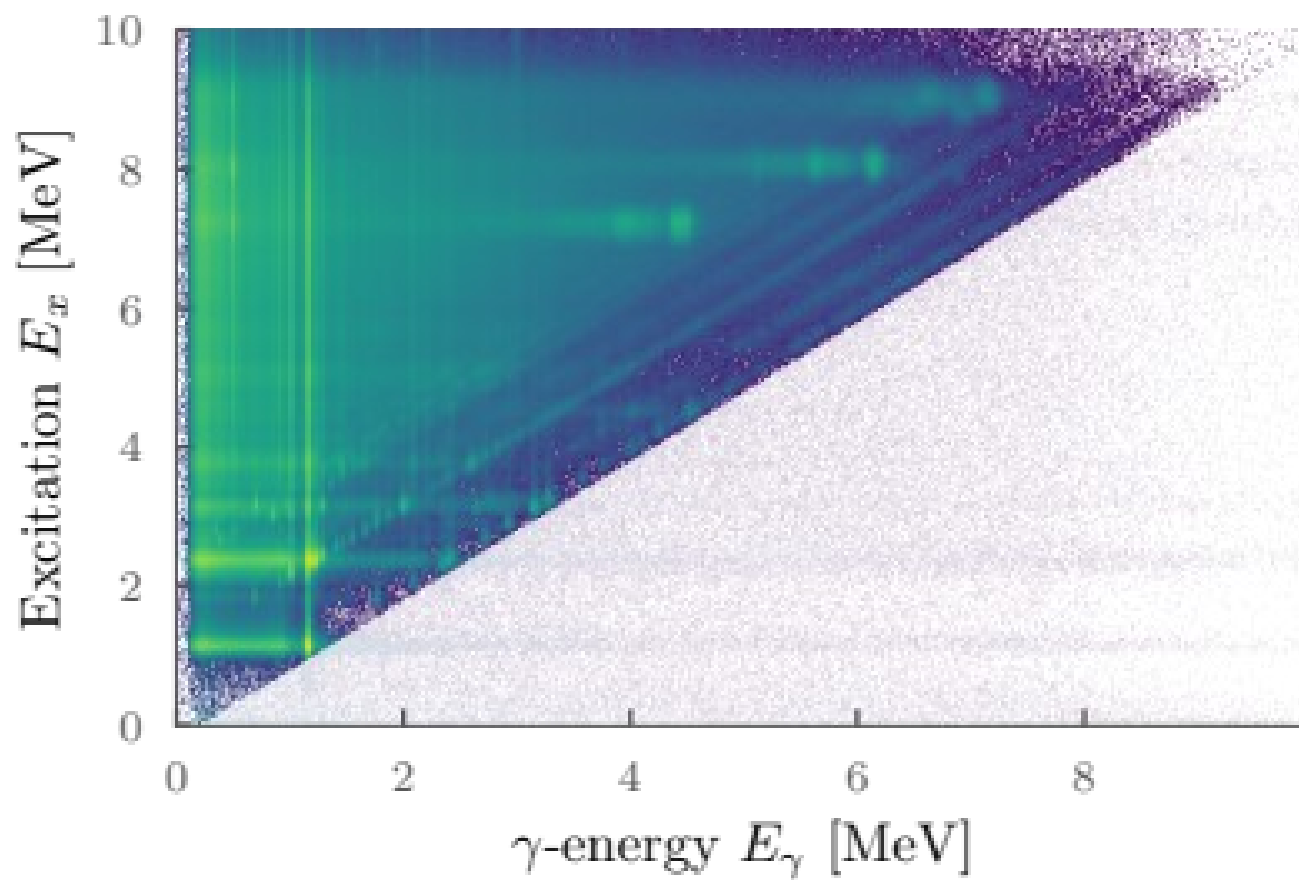
OSCAR



SiRi



Observed data



The Oslo Method

- Objective: *Simultaneous extraction of the nuclear level density (NLD) and the γ -ray strength function (γ SF) of a nucleus.*
- The main steps:
 - 1) **Unfolding of γ -ray spectra for each excitation energy to compensate for the detector response.**
 - 2) Determination of first-generation γ -ray spectra.
 - 3) Extraction of the NLD and the γ SF.
 - 4) Normalizing the NLD and the γ SF to eliminate degrees of freedom.
- **A rigorous uncertainty propagation is needed for each step.**

The old Unfolding Method

- 1) Start with a trial function $\mathbf{u} = \mathbf{n}$ at iteration $i = 0$.
 - 2) Calculate the folded spectrum $\mathbf{f}_i = \mathbf{R}\mathbf{u}_i$.
 - 3) $\mathbf{u}_{i+1} = \mathbf{u}_i + (\mathbf{n} - \mathbf{f}_i)$.
 - 4) Iterate until $\mathbf{f}_i \approx \mathbf{u}_i$
- *New frequentist method of unfolding by Erlend Lima, University of Oslo*

Why is unfolding hard: ill-conditioning, non-uniqueness

- $\mathbf{R} := \mathbf{D} \mathbf{G}_y$
- The order matters: $\mathbf{D} \mathbf{G}_y \neq \mathbf{G}_y \mathbf{D}$
- The naive approach: $\boldsymbol{\mu} = \mathbf{R}^{-1} \mathbf{n} \rightarrow$ very large fluctuations
- $\mathbf{R}$ is ill-conditioned $\rightarrow$ regularization is necessary
- The null space of $\mathbf{R}$ is non-empty in general:

$$\mathbf{v} = \mathbf{R}(\boldsymbol{\mu} + \mathbf{x}) \quad \forall \mathbf{x} : \mathbf{R}\mathbf{x} = 0$$

- The prior is a powerful tool to avoid unphysical solutions

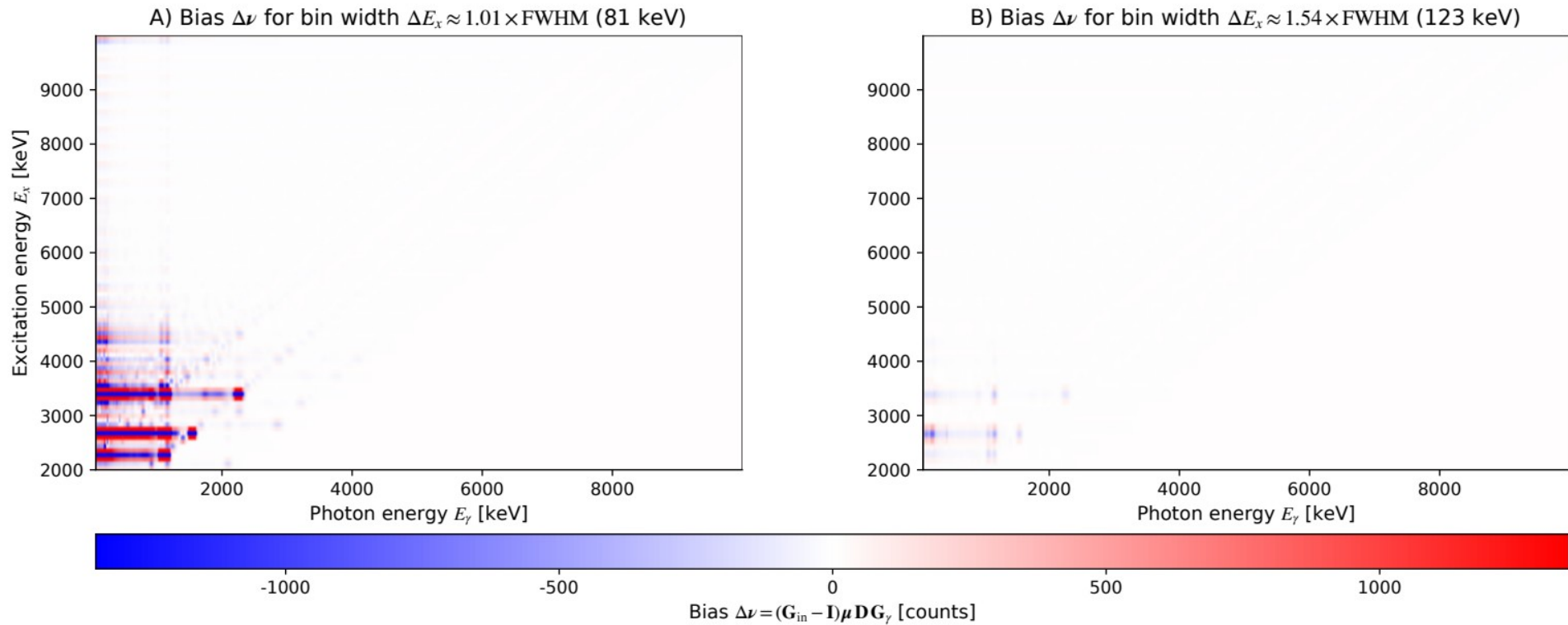
Assumptions

- Negligible “leakage” of counts between E_x bins.
- Perfect knowledge about the response .
- No systematic errors in the background between OFF and ON runs.

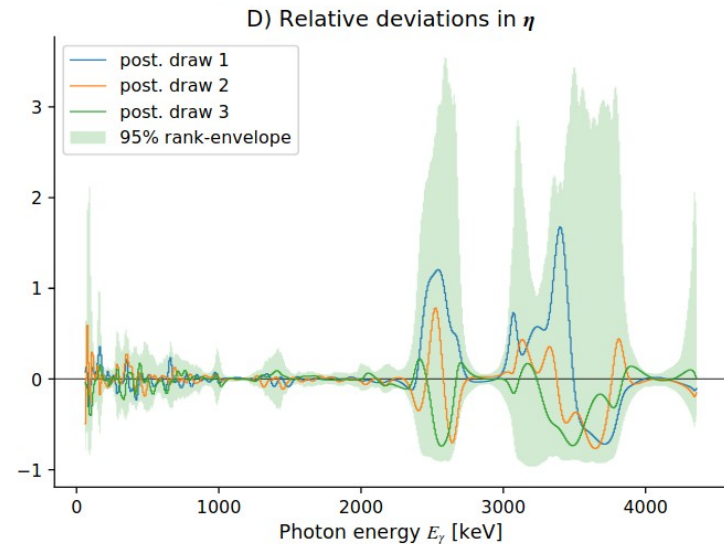
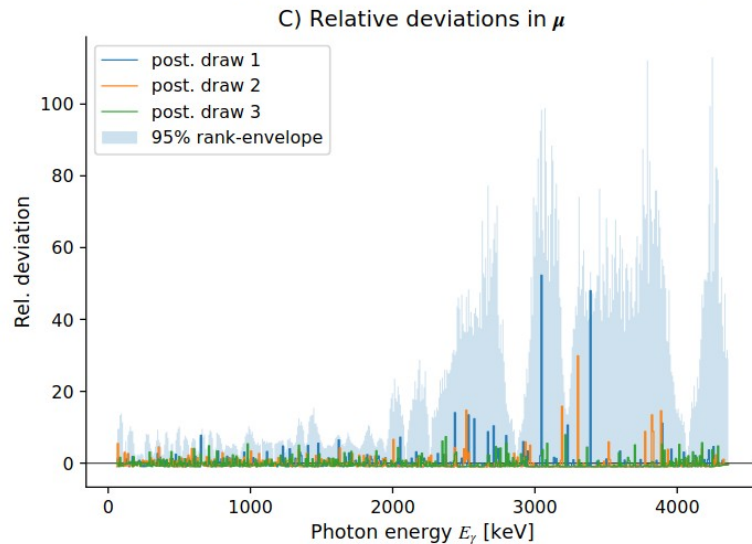
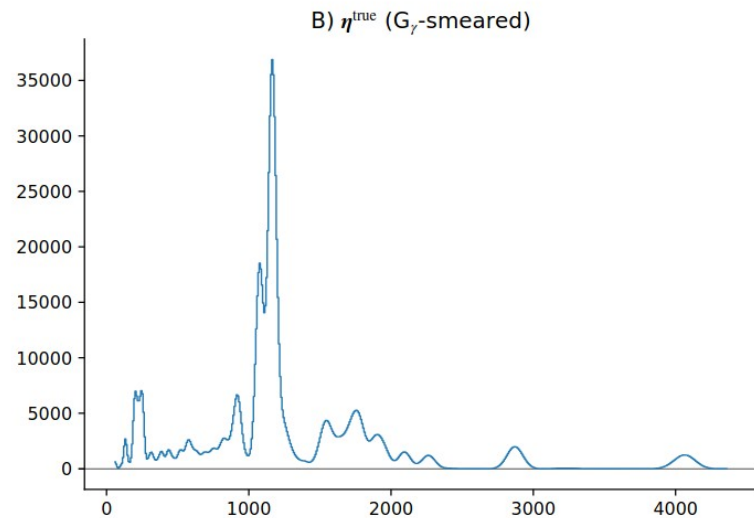
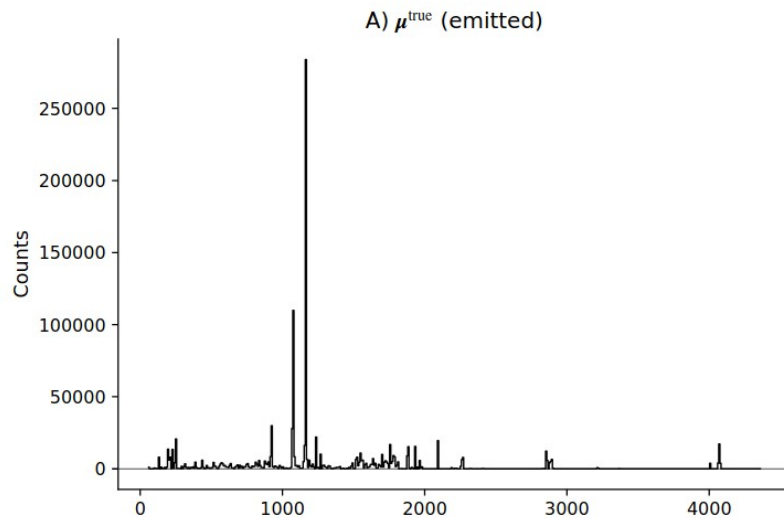
Beyond the scope of this work:

- Uncertainty in the response itself

The effect of the E_x -smearing



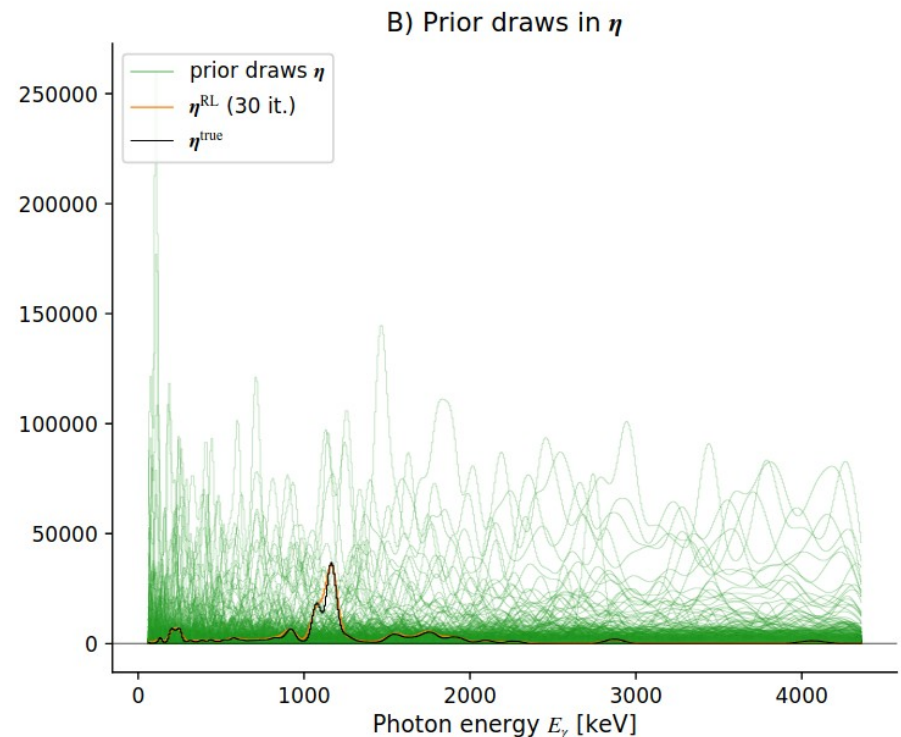
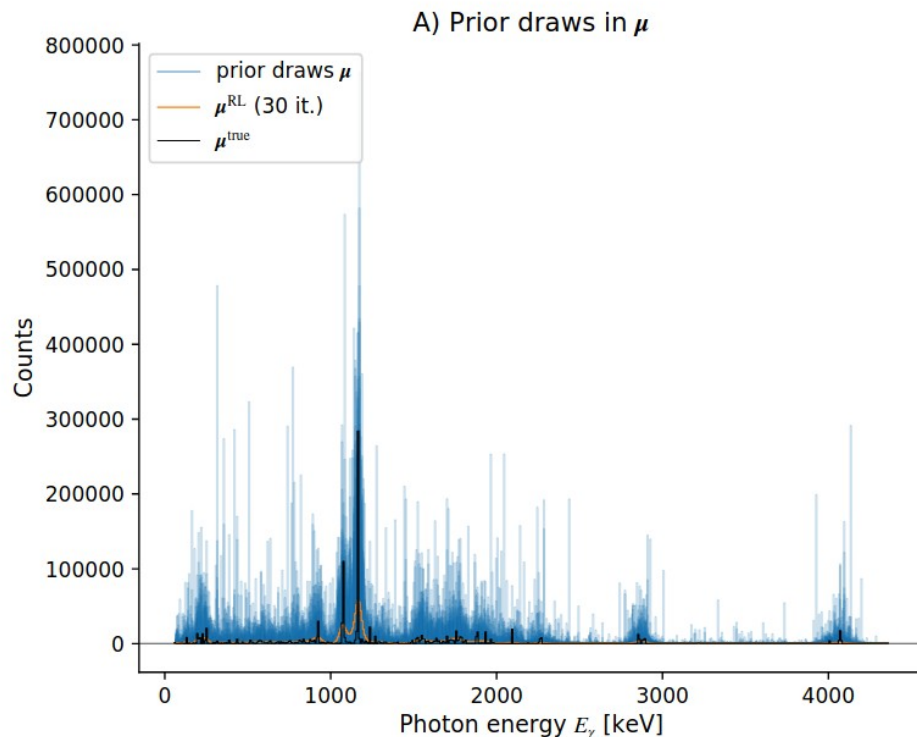
Identifiability μ vs η



Richardson-Lucy (RL) estimate

$$\boldsymbol{\mu}^{(t+1)} = \boldsymbol{\mu}^{(t)} \circ \left[\left(\frac{\mathbf{n} - \mathbf{b}}{\boldsymbol{\mu}^{(t)} \mathbf{R}} \right) \mathbf{R}_\gamma^\top \right], \quad t = 0, 1, \dots$$

Citation: W. H. Richardson, Journal of the optical society of America, 1972 and L. B. Lucy - Astronomical Journal, 1974



Priors

$$\mu_j \mid M_j \sim \text{Gamma}\left(\alpha, \beta_j = \frac{\alpha}{M_j}\right),$$

$$\log M_j \sim \mathcal{N}\left(\log m_j - \frac{1}{2}\sigma_j^2, \sigma_j^2\right),$$

where m_j is set to the RL-estimate μ_j^{RL}

$$\sigma_j = \sigma_{\min} + \frac{\sigma_{\max} - \sigma_{\min}}{1 + \frac{\mu_j^{\text{RL}}}{\mathbb{E}[\boldsymbol{\mu}^{\text{RL}}]}}.$$

$$\lambda_j \sim \text{Gamma}(a_0, b_0)$$

typically with $a_0 = 1.0$

and $b_0 = a_0 / \text{avg}\{\mathbf{b}^{\text{obs}}\}$.

Likelihood

$$\mathcal{L}(n, \mathbf{b} \mid \boldsymbol{\mu}, \boldsymbol{\lambda}) = \text{Poisson}(\mathbf{b} \mid \boldsymbol{\lambda}) \text{Poisson}(n \mid \boldsymbol{\mu} \mathbf{D} \mathbf{G}_\gamma + \boldsymbol{\lambda}).$$

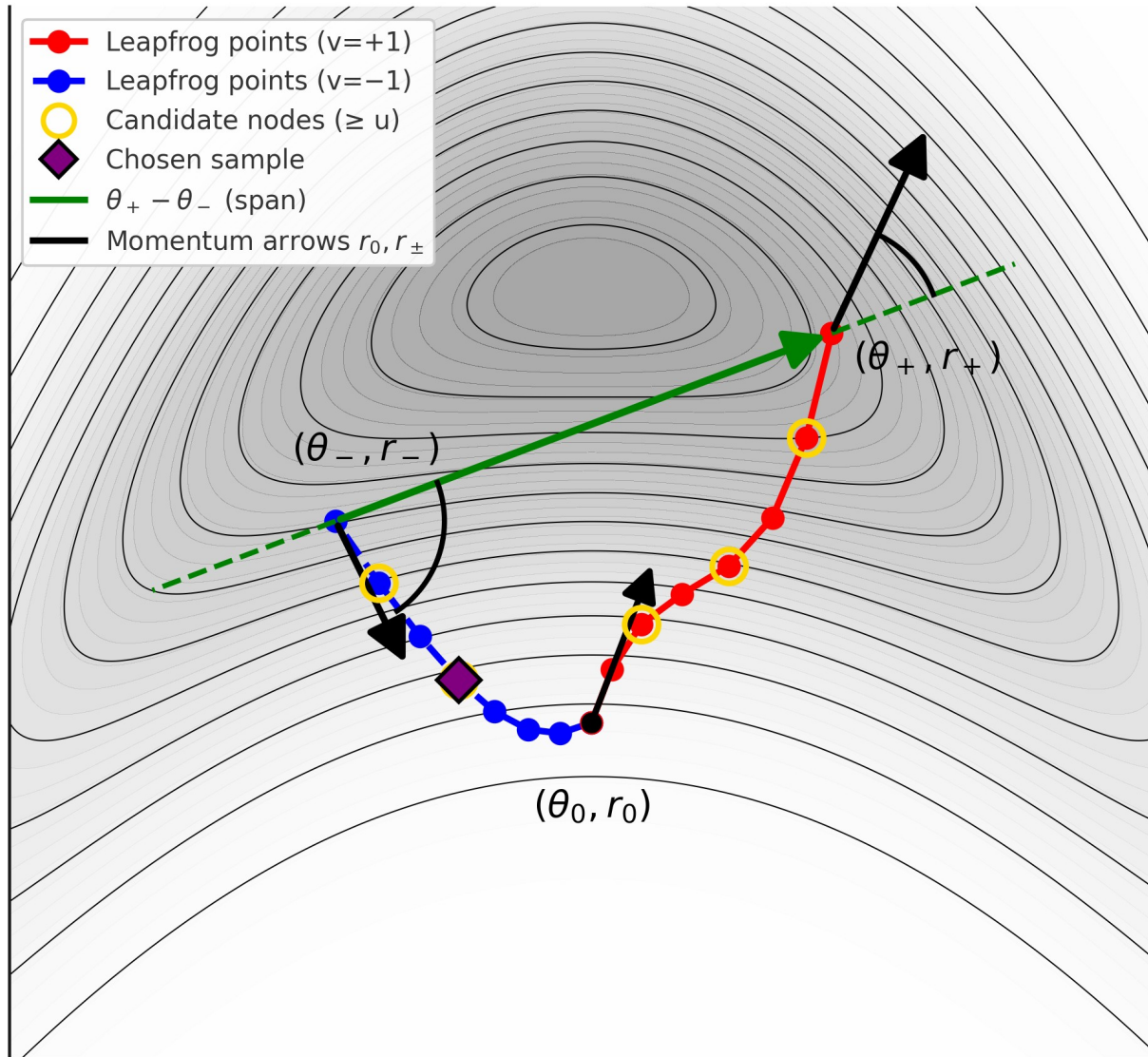
Hessian:

$$\mathbf{H}_{\log \boldsymbol{\mu}}^{(-\ell)} = -\text{diag}(\boldsymbol{\mu}) \mathbf{H}_\mu \text{diag}(\boldsymbol{\mu}) - \text{diag}(\boldsymbol{\mu} \odot \mathbf{g}_\mu),$$

$$\mathbf{g}_\mu = (\mathbf{D} \mathbf{G}_\gamma)^\top \left(\frac{\mathbf{n}}{\boldsymbol{\nu}} - \mathbf{1} \right), \quad \boldsymbol{\nu} = \boldsymbol{\mu} \mathbf{D} \mathbf{G}_\gamma + \boldsymbol{\lambda},$$

$$\mathbf{H}_\mu = -(\mathbf{D} \mathbf{G}_\gamma)^\top \text{diag}\left(\frac{n}{\boldsymbol{\nu}^{\odot 2}}\right) (\mathbf{D} \mathbf{G}_\gamma).$$

No-U-Turn Sampler (NUTS)

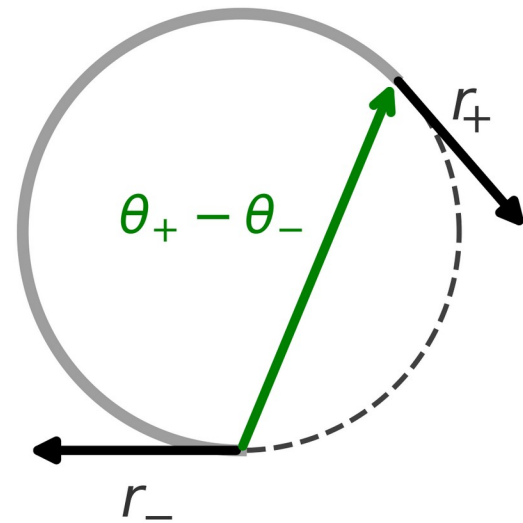
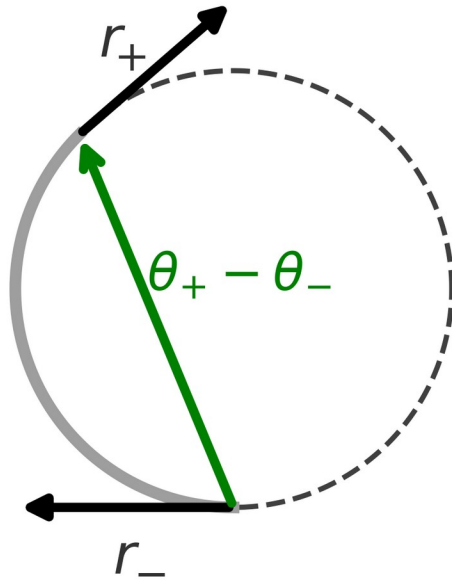


- $v \in \{-1, 1\}$
- 2^j leapfrog steps,
 $j=\{0,1,\dots,N\}$

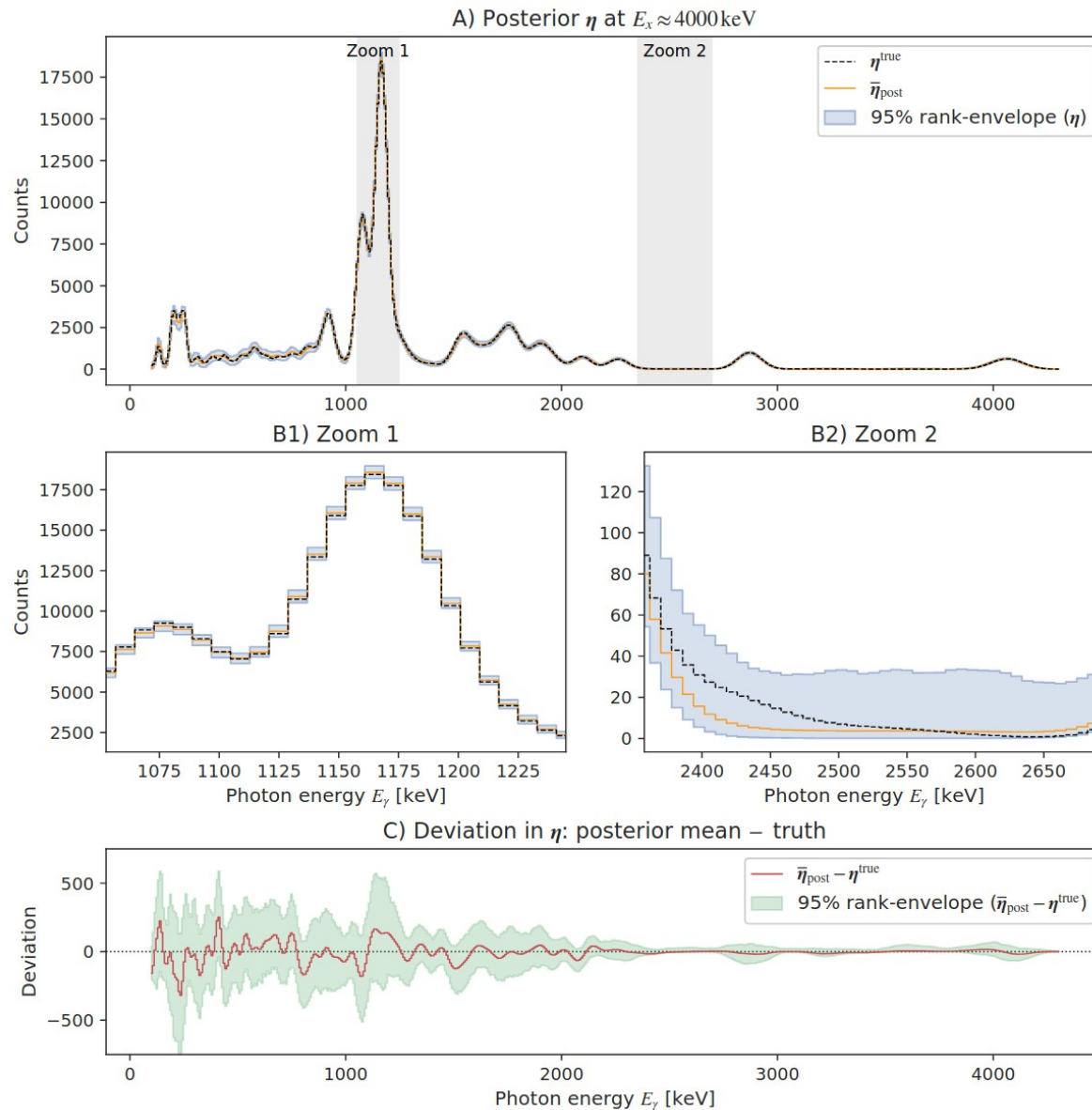
$\text{Uniform}(u; [0, \exp\{\mathcal{L}(\theta) - \frac{1}{2}r \cdot r\}])$

No-U-Turn condition:

$$(\theta^+ - \theta^-) \cdot r^- < 0 \quad \text{or} \quad (\theta^+ - \theta^-) \cdot r^+ < 0.$$

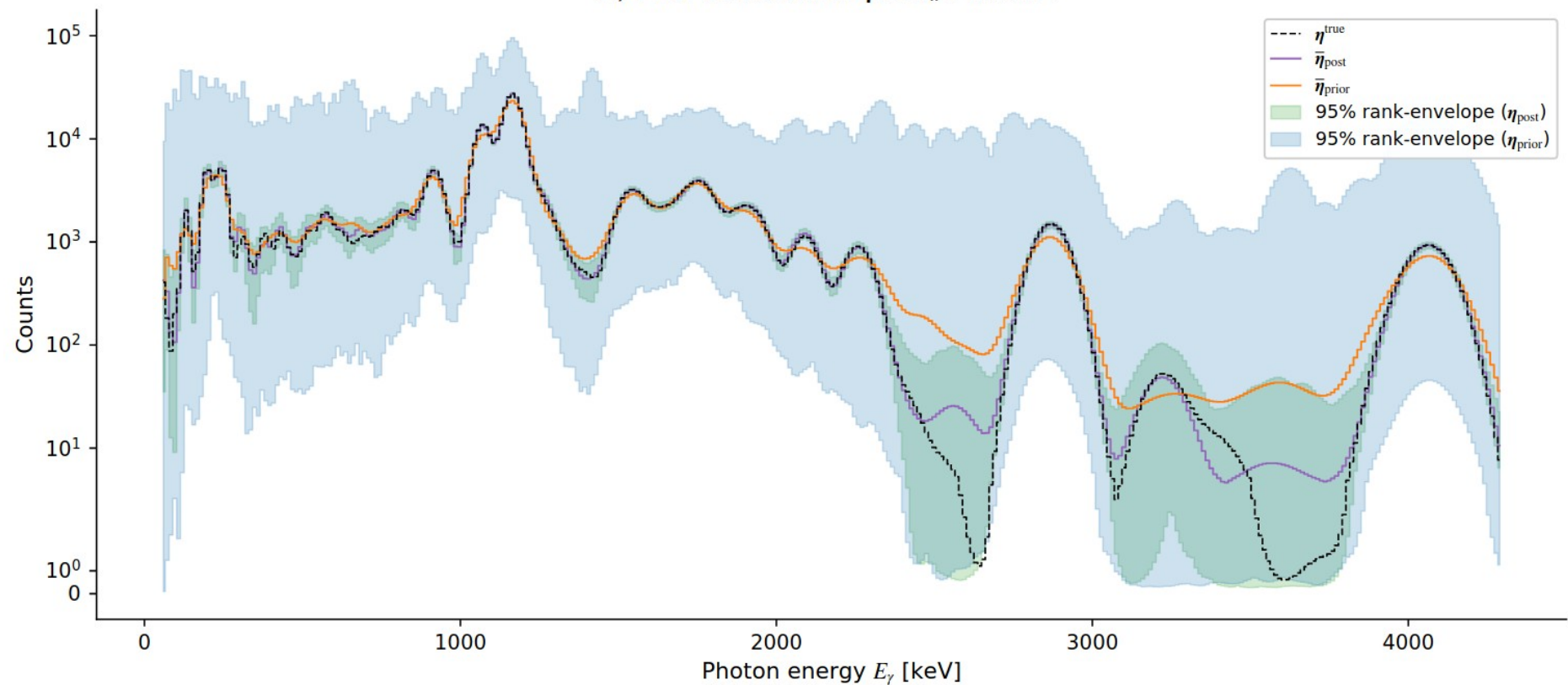


Posterior on unfolded solution

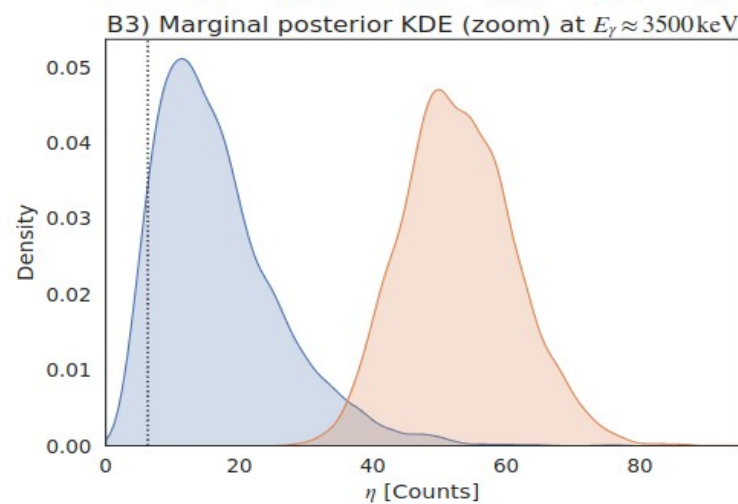
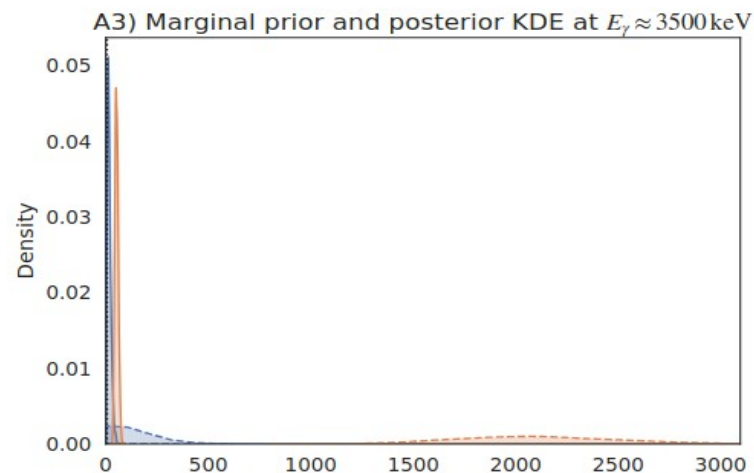
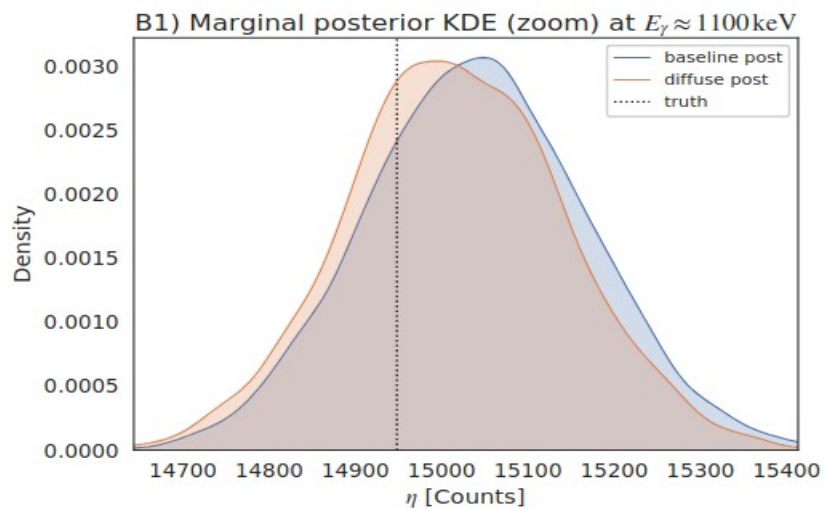
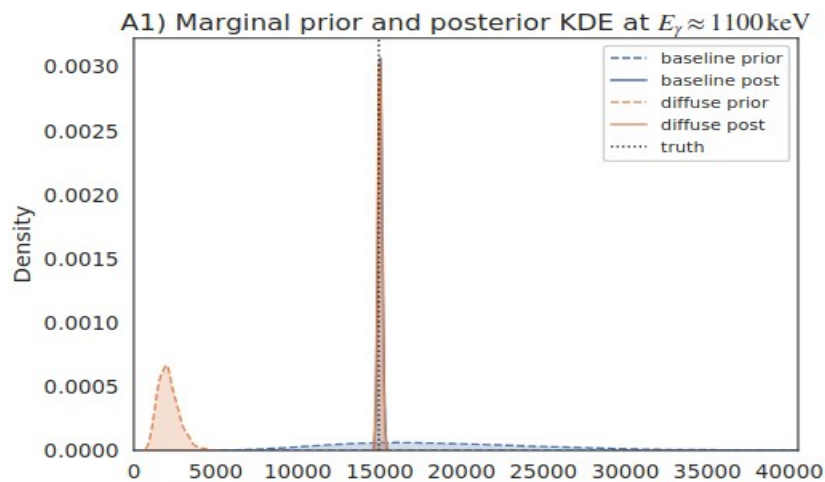


Prior vs posterior

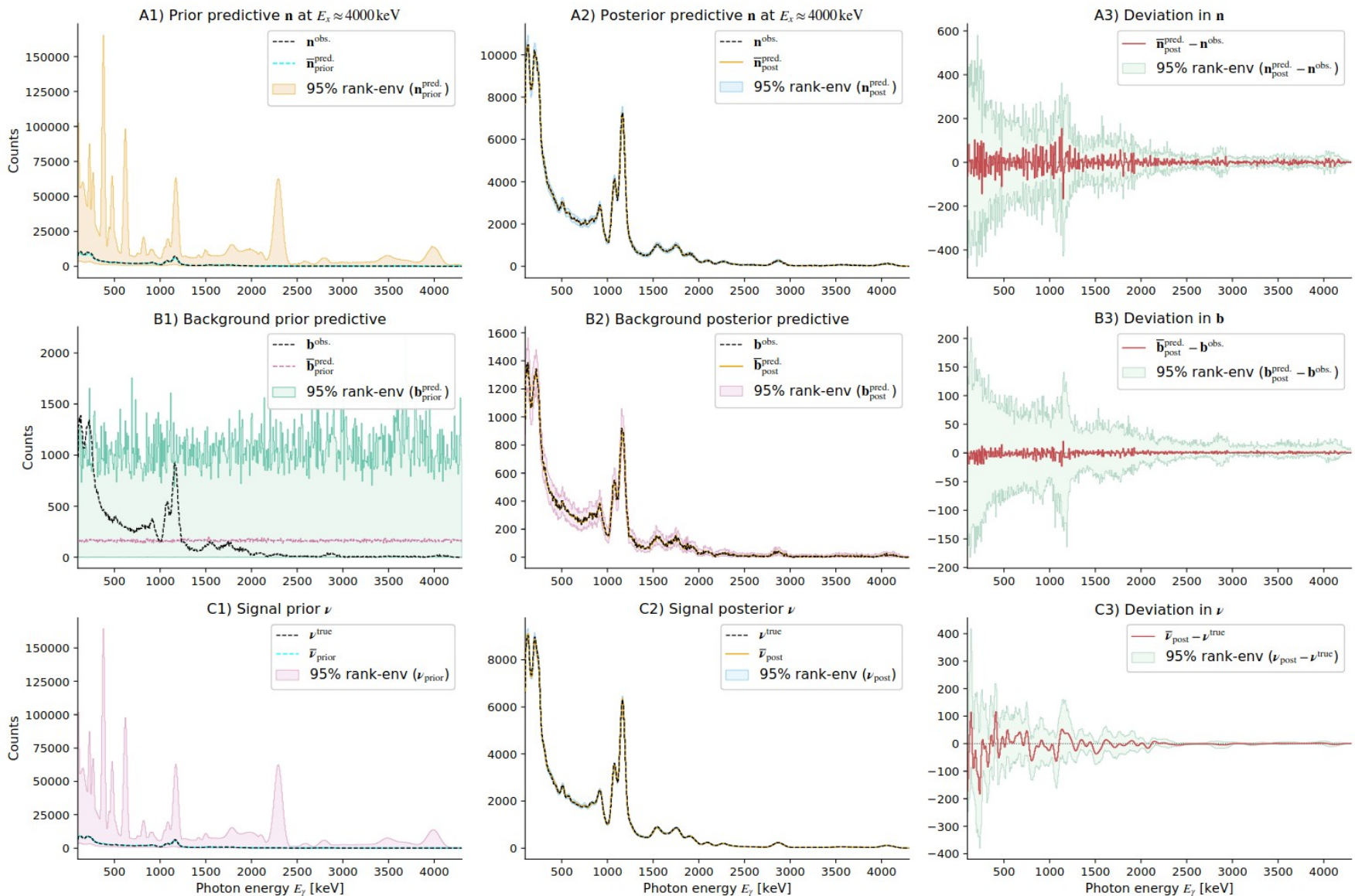
A) Prior vs Posterior η at $E_x \approx 4000$ keV



Marginal prior vs posterior



Prior and posterior predictive checks



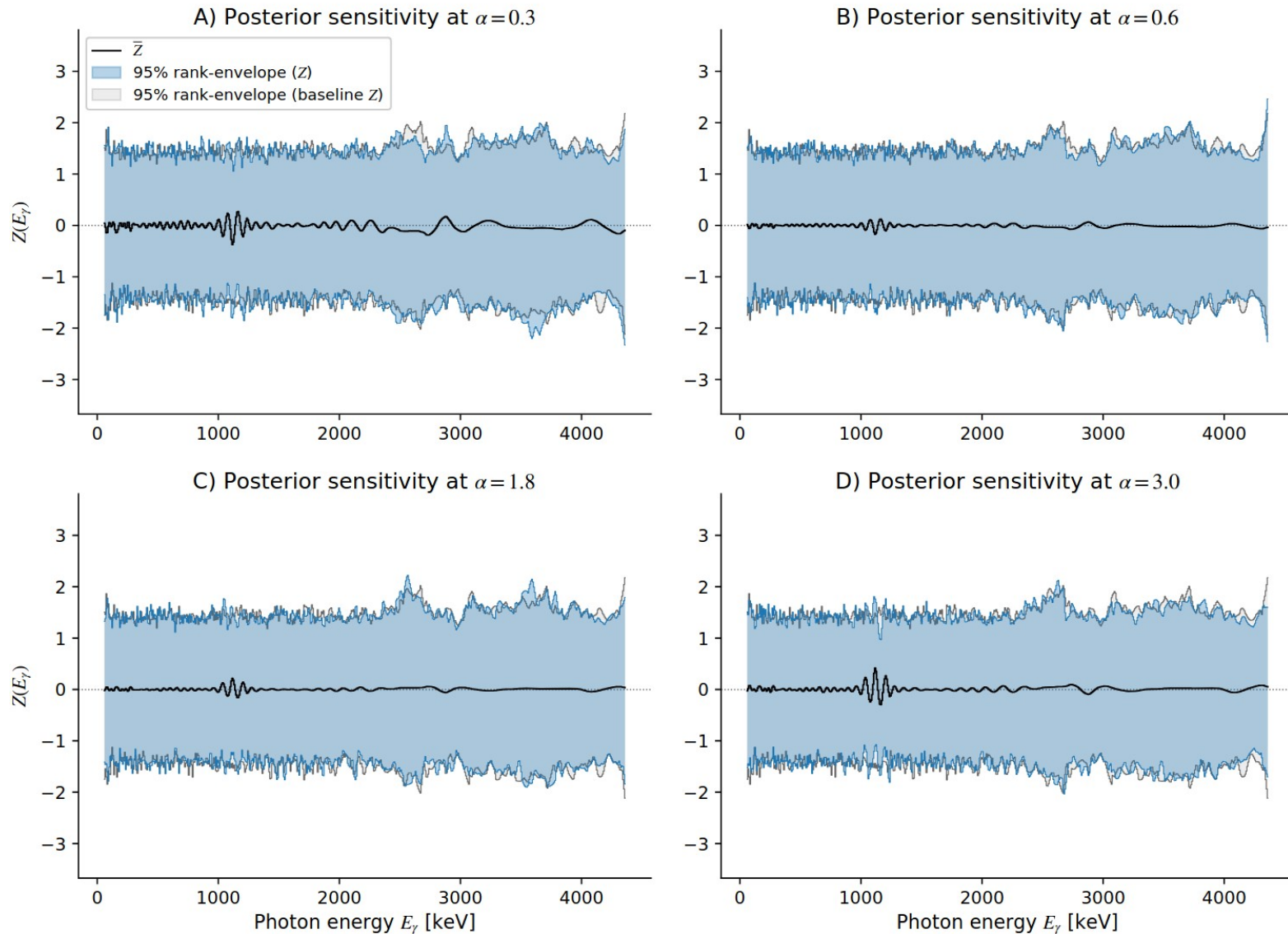
Sensitivity and robustness

- Sensitivity: how much does the posterior change if we *change a parameter of the prior/hyperprior?*
- Robustness: how much does the posterior change if we *change more fundamental properties of the prior?*
- Z-statistic (baseline prior vs changed prior):

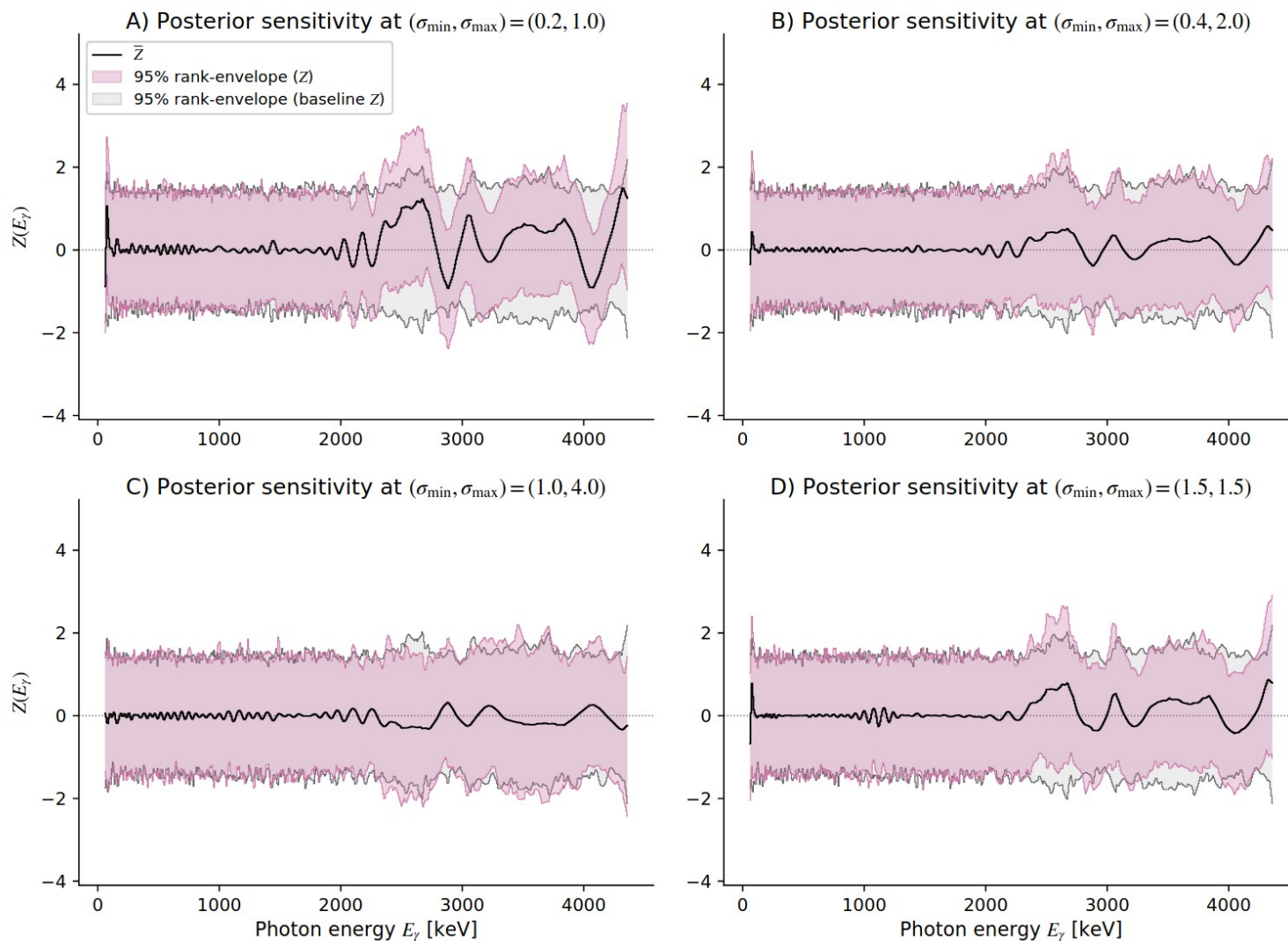
$$Z_s(j) = \frac{C_s(j) - B_s(j)}{w(j)}$$

$$Z_s^0(j) = \frac{B_s^{(2)}(j) - B_s^{(1)}(j)}{w(j)}$$

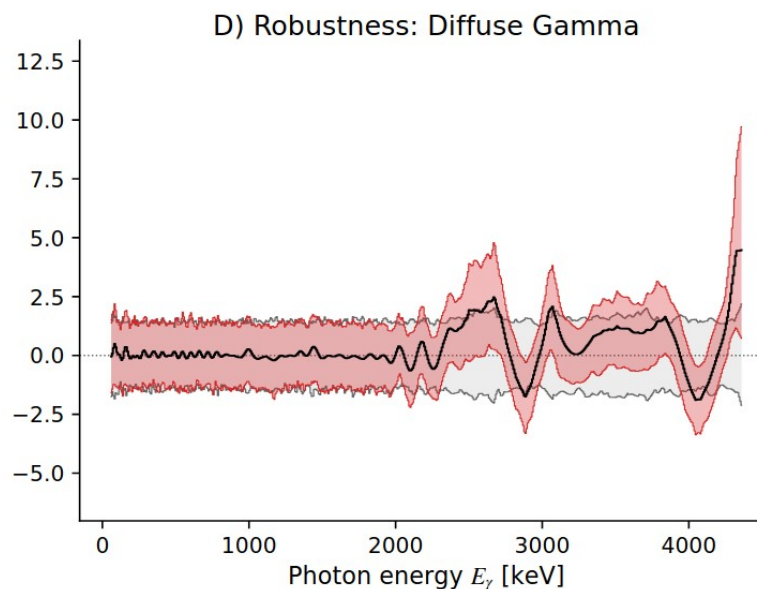
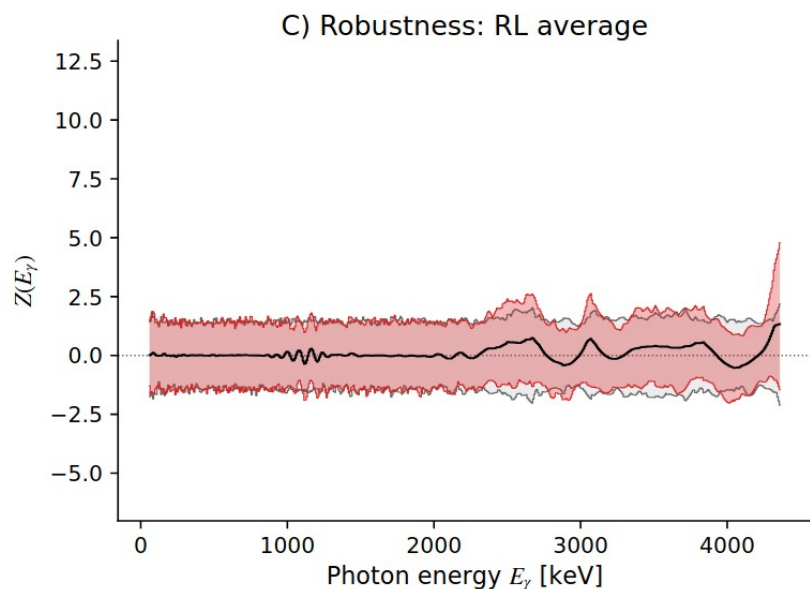
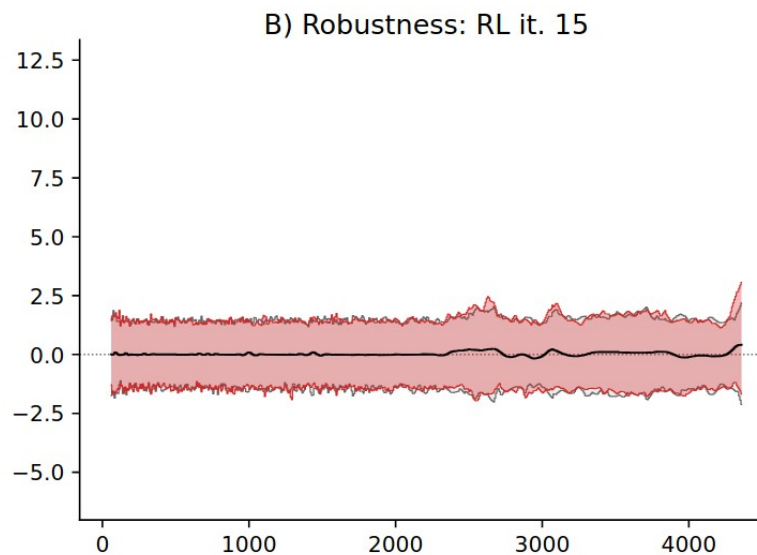
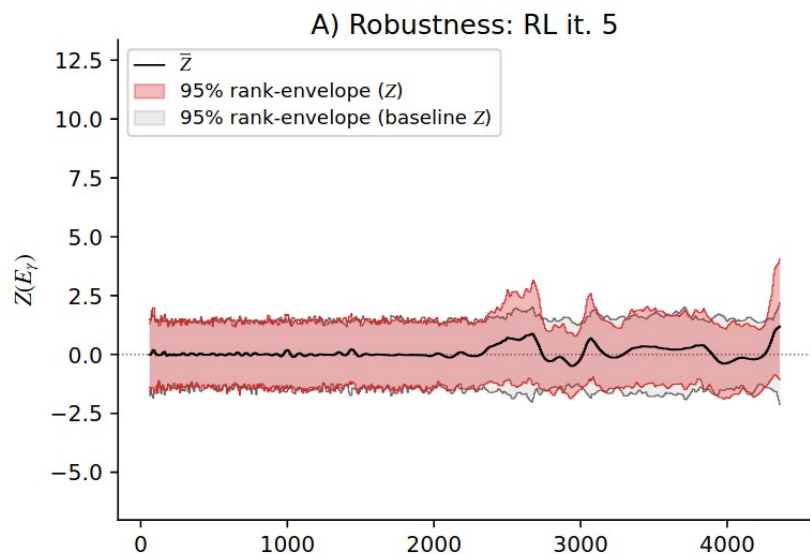
Sensitivity to prior α



Sensitivity to hyperprior σ

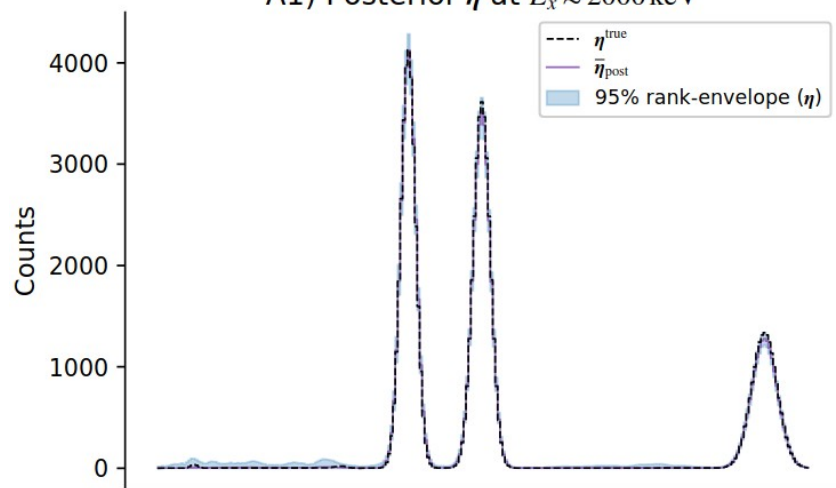


Robustness to different priors

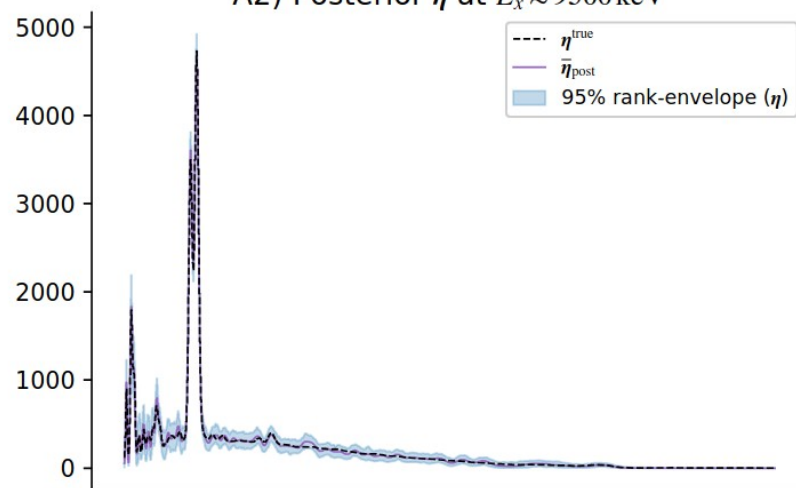


Unfolding results for different E_x

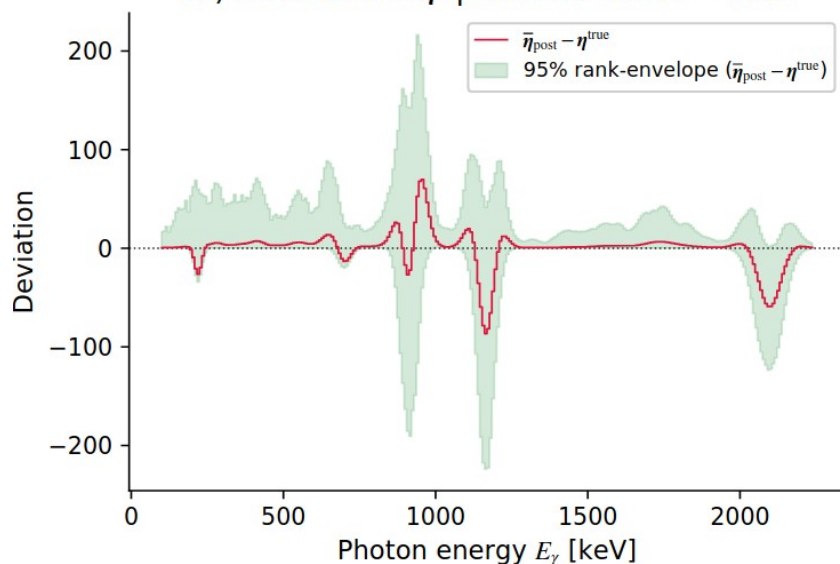
A1) Posterior η at $E_x \approx 2000$ keV



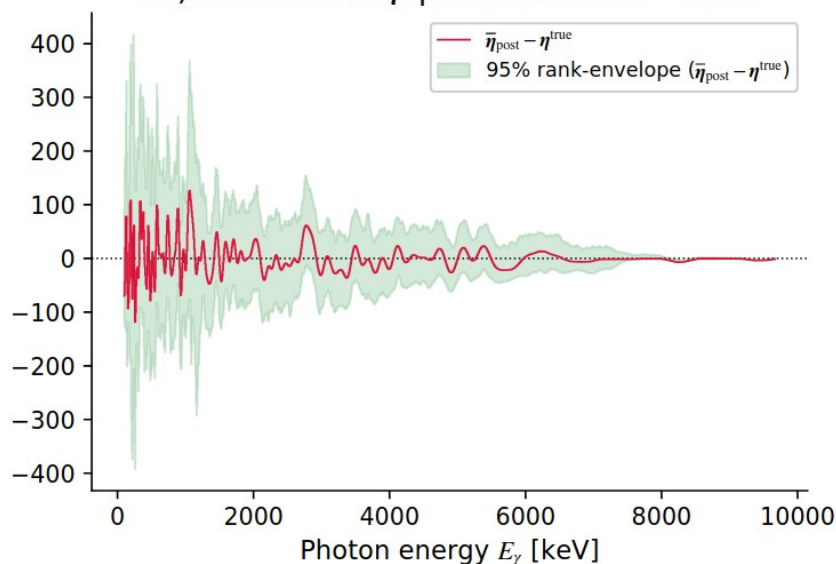
A2) Posterior η at $E_x \approx 9500$ keV



B1) Deviation in η : posterior mean – truth



B2) Deviation in η : posterior mean – truth



Summary and outlook

- The choice of prior is most important for low count energy bins, where the data is less informative.
- The prior needs to be sufficiently wide to avoid posterior bias.
- The RL estimate helps us determining the critical regions where the prior variance should be relatively larger.
- Outlook: Uncertainty propagation of posterior draws through the next steps of the Oslo Method