



Attractors in hydrodynamic simulations of heavy ion collisions

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Attractors and thermalization in nuclear collisions and cold quantum gases



What you will see today

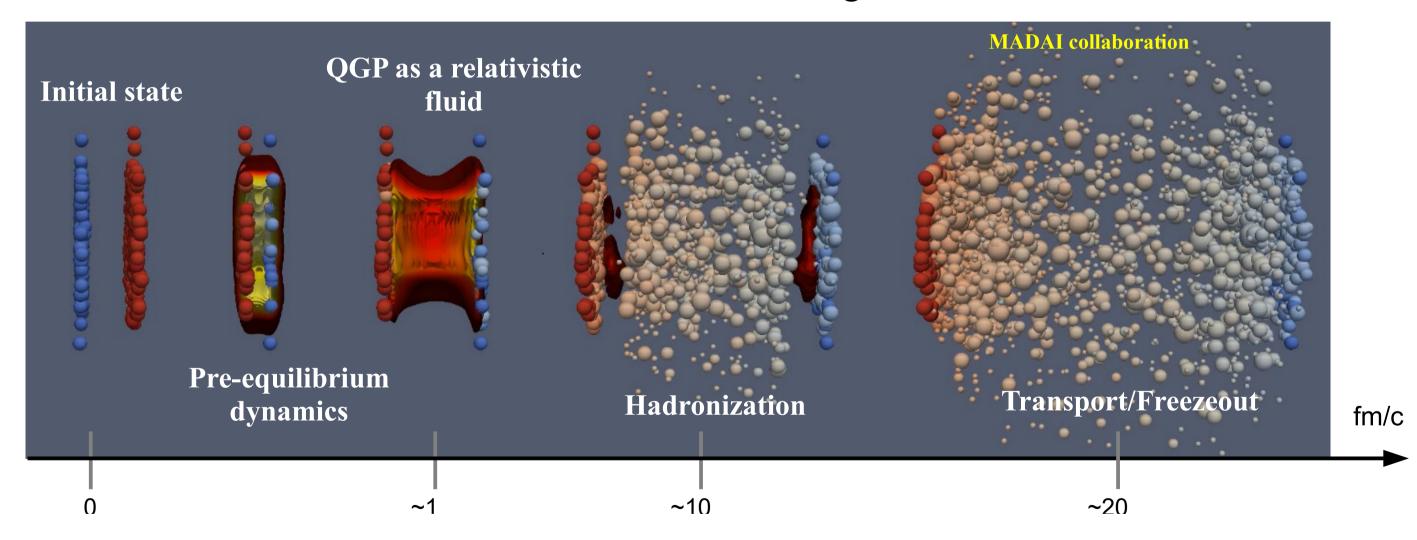
We look for hints of hydrodynamic attractors in simulations of heavy ion collisions

Outline:

- Motivation
- Attractors in (0+1)D Bjorken flow
- (2+1)D Hydrodynamic simulations and attractors
- Discussion and conclusions

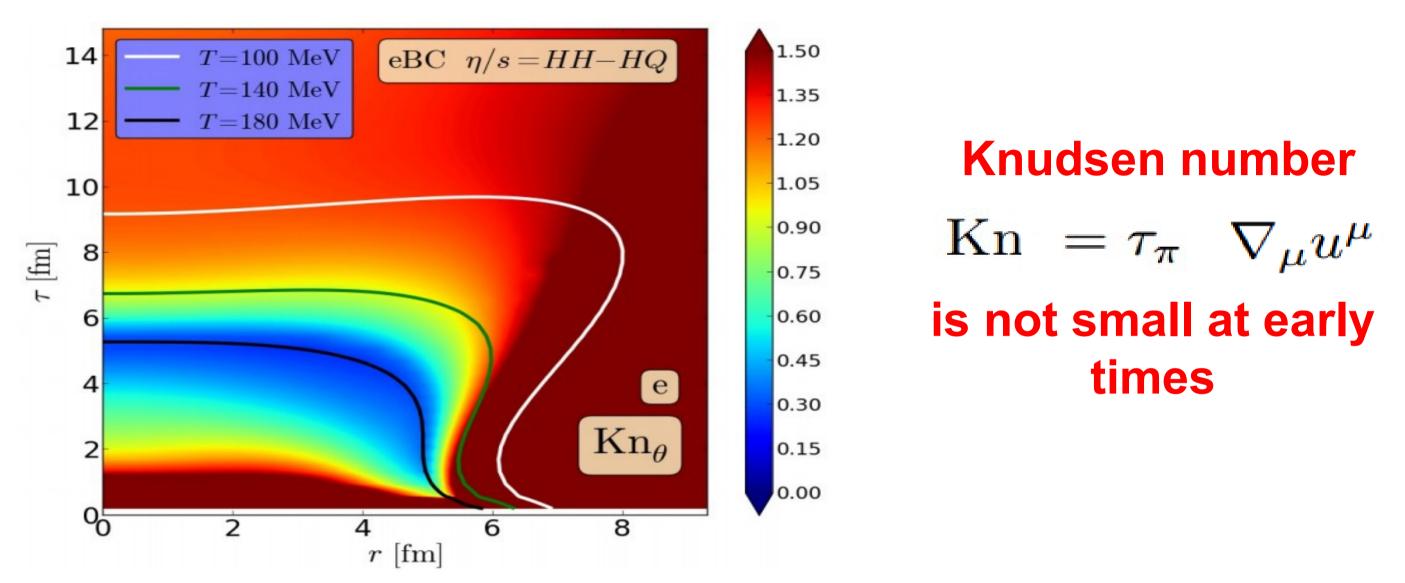
Current picture of a heavy ion collision

Empirical: Fluid-dynamical models of heavy ion collisions work well at RHIC and LHC energies



Main assumption: fluid dynamics can be applied at early times ~ 0.1--1 fm

Extreme Conditions



Can this system be close to equilibrium?

Or domain of applicability of hydrodynamic theories better than expected?

Relativistic fluid dynamics

Effective theory describing the dynamics of a system over long-times and long-distances

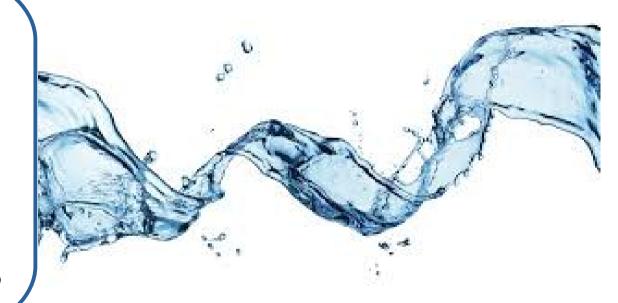
Conservation laws

+

Equation of state

+

simple constitutive relations



separation of scales:
$$K_N \sim \frac{\ell}{L} \ll 1$$

macroscopic: L

microscopic: ℓ

Relativistic Fluid dynamics

Conservation laws

$$\partial_{\mu}T^{\mu\nu} = 0$$

EoS:
$$P_0 = P_0(\varepsilon)$$

Tensor decomposition

$$T^{\mu\nu} = \varepsilon \, u^{\mu} u^{\nu} - \Delta^{\mu\nu} \, (P_0 + \Pi) + \pi^{\mu\nu}$$
Projection operator: Bulk viscous

$$\Delta^{\mu\nu} = g^{\mu\nu} - u^{\mu}u^{\nu}$$

 $\pi^{\mu\nu} = 2\eta \nabla^{\langle\mu} u^{\nu\rangle}$

Bulk viscous pressure

Shear stress tensor

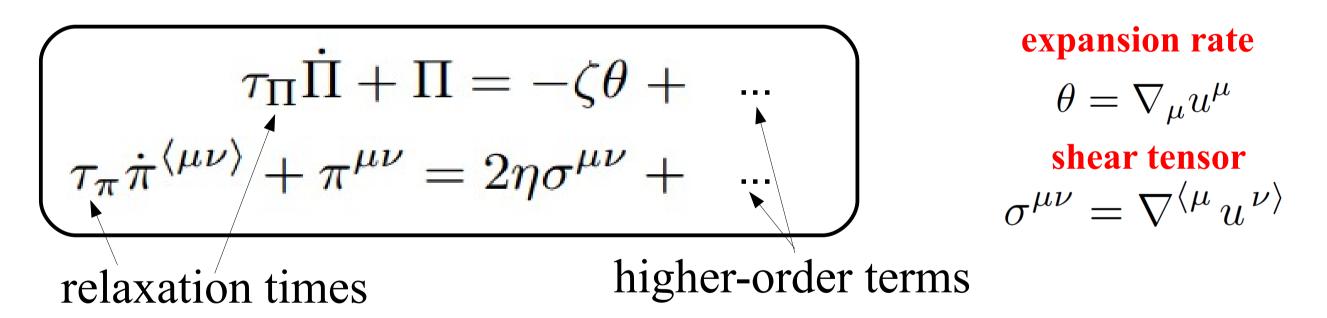
$$\Pi = -\zeta \nabla_{\mu} u^{\mu}$$



What we solve is not "traditional" fluid dynamics

Causality: constitutive relations for the dissipative currents *cannot* be imposed

Dynamical equations, e.g. Israel-Stewart theory Annals Phys. 118 (1979) 341-372



non-perturbative theory in gradients! Heller&Spalinski, PRL 115 (2015) 7, 072501

Bjorken flow (toy model of heavy ion collisions) J. D. Bjorken, Phys. Rev. D27, 140 (1983)

Simple model for *boost invariant* longitudinal expansion



Homogeneous fluid in hyperbolic coordinates (τ, x, y, ς) $\tau = \sqrt{t^2 - z^2} \quad \varsigma = \tanh^{-1}(z/t)$

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Static velocity

$$u^{\mu} = (1, 0, 0, 0)$$

Gradients $\sim 1/\tau$

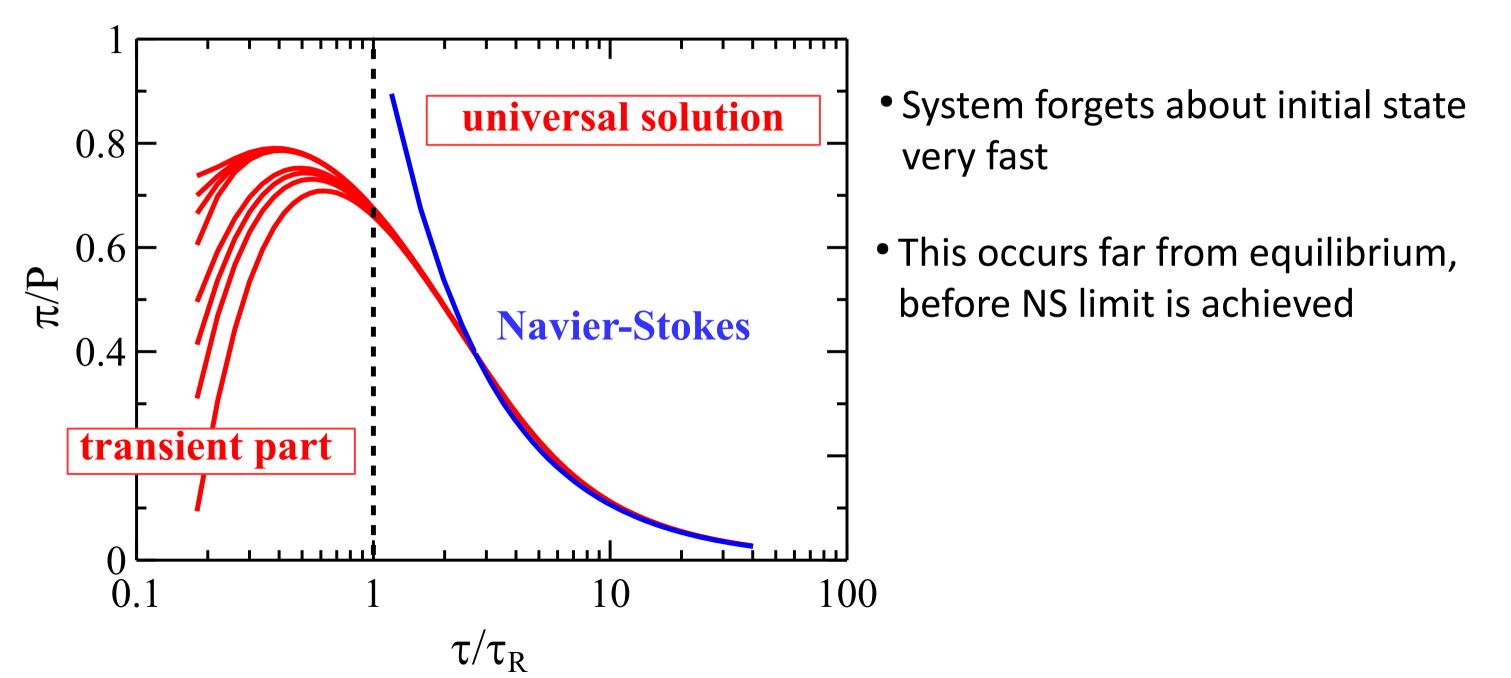
$$\sigma_{\nu}^{\mu} = \operatorname{diag}\left(0, -\frac{1}{3\tau}, -\frac{1}{3\tau}, \frac{2}{3\tau}\right)$$

energy-momentum tensor

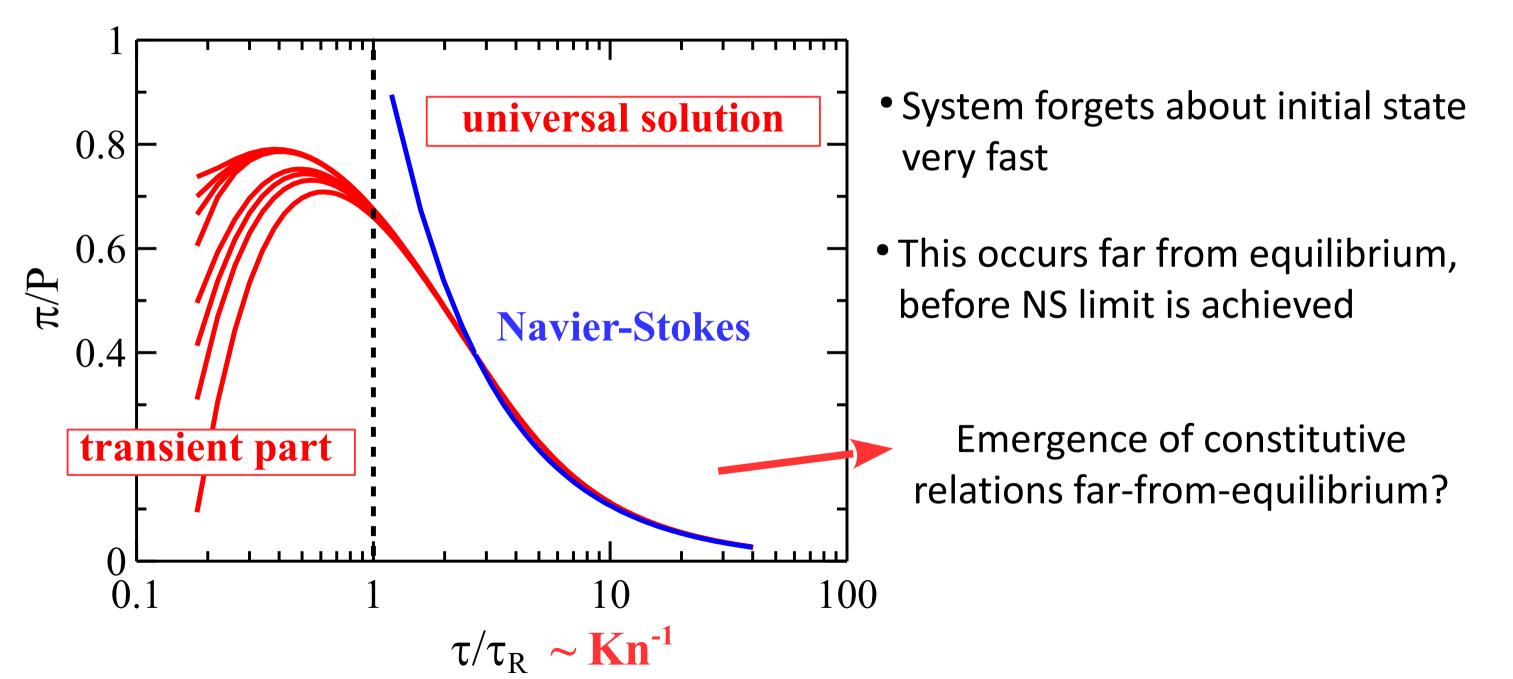
$$u^{\mu} = (1,0,0,0) \qquad \sigma^{\mu}_{\nu} = \operatorname{diag}\left(0, -\frac{1}{3\tau}, -\frac{1}{3\tau}, \frac{2}{3\tau}\right) \qquad T^{\mu}_{\nu} = \operatorname{diag}\left(\varepsilon, P - \pi/2, P - \pi/2, P + \pi\right)$$
dissipative correction

Knudsen number: $K_N \sim \hat{\tau}^{-1} \equiv \tau_R/\tau$

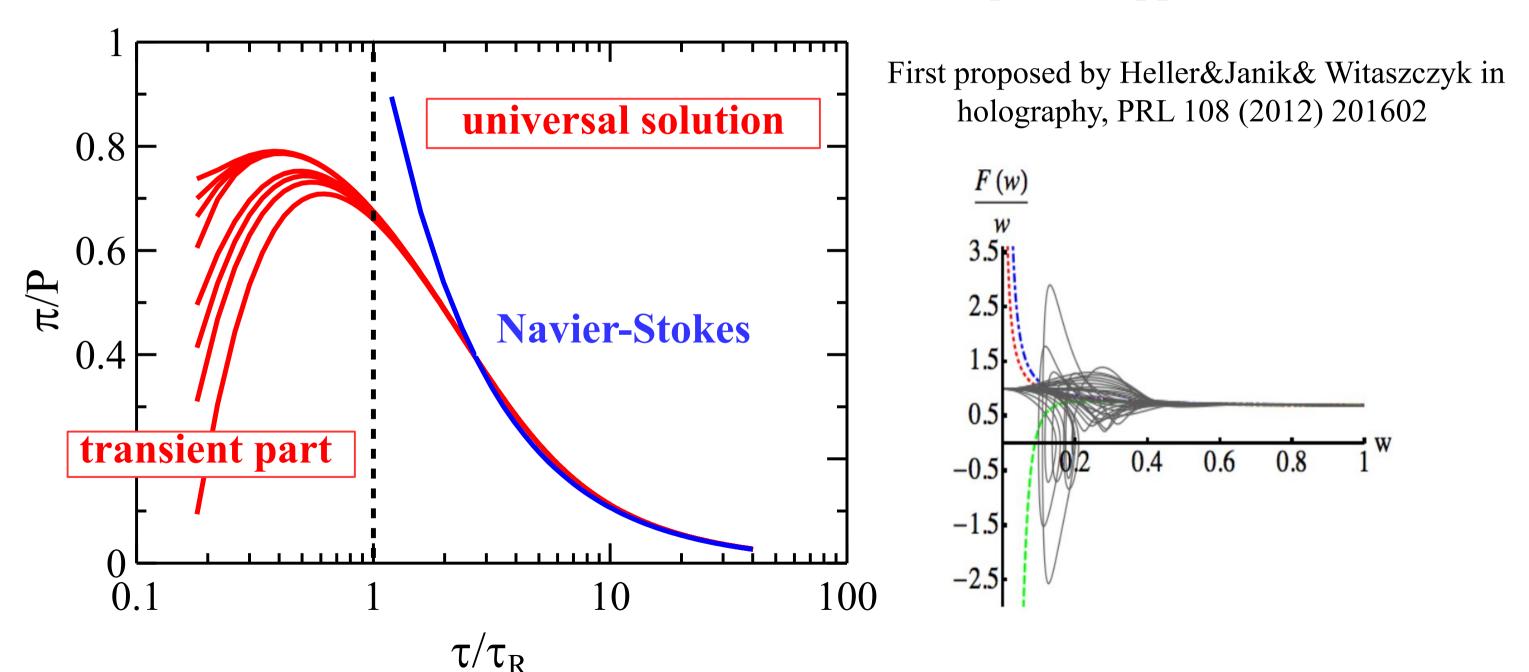
Late-time solution of the conformal (RTA) Boltzmann equation appears to be universal



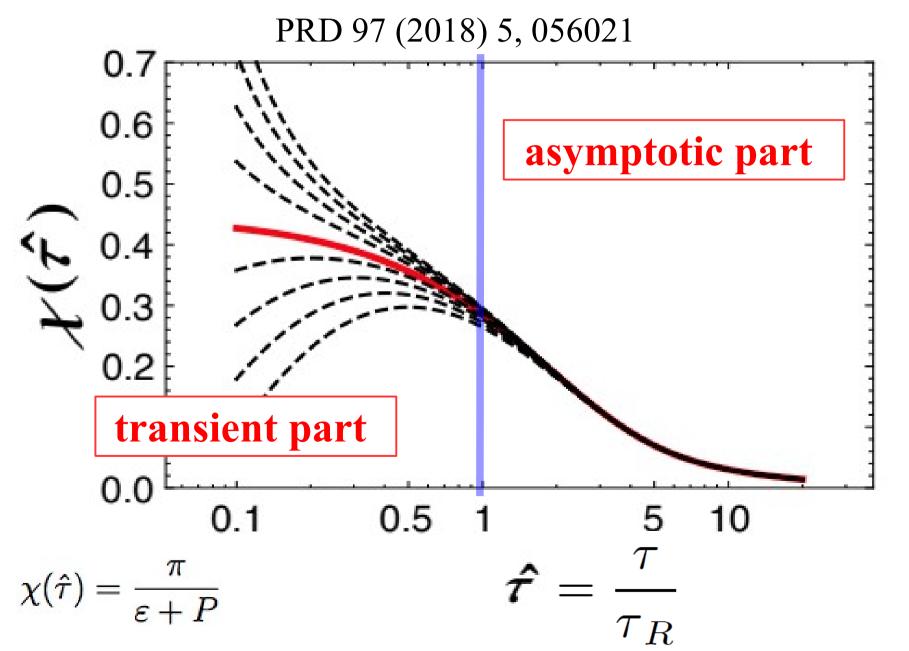
Late-time solution of the conformal (RTA) Boltzmann equation appears to be universal



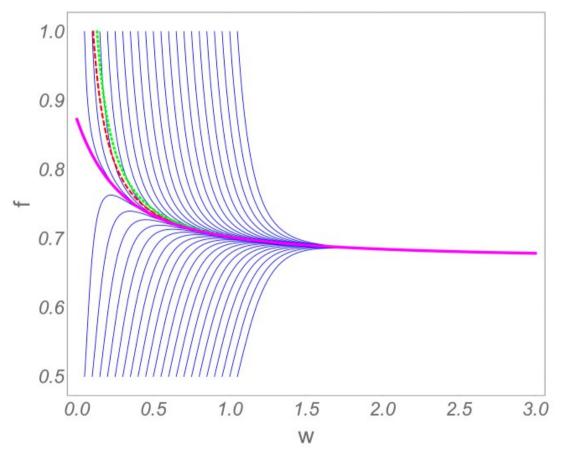
Late-time solution of the conformal (RTA) Boltzmann equation appears to be universal



Late-time solution of Israel-Stewart theory displays the same feature



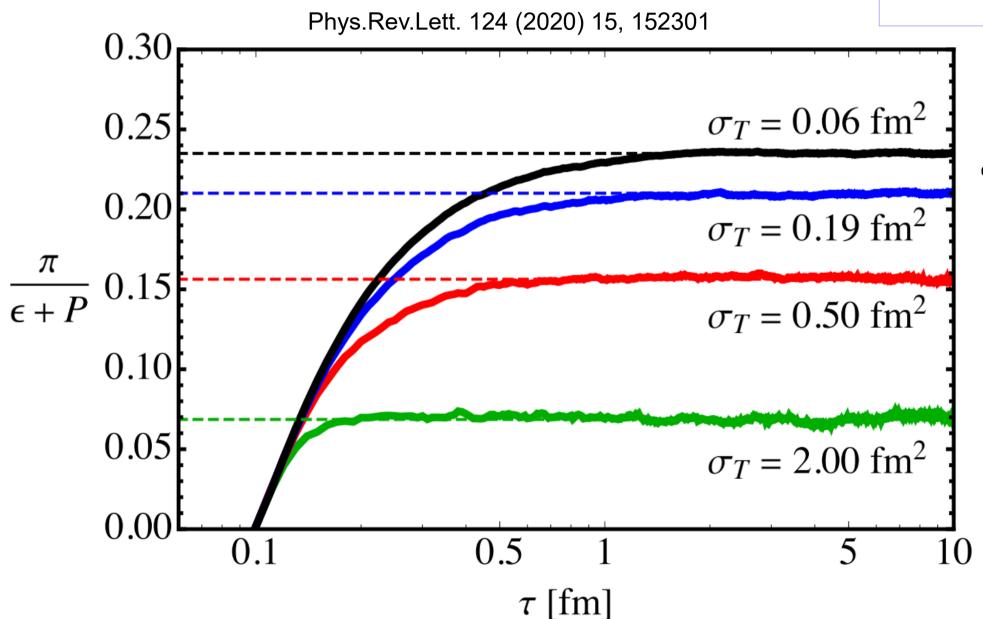
First proposed by Heller&Spalinski, PRL 115 (2015) 7, 072501



Relativistic gas of hard spheres

System with a contant Knudsen number

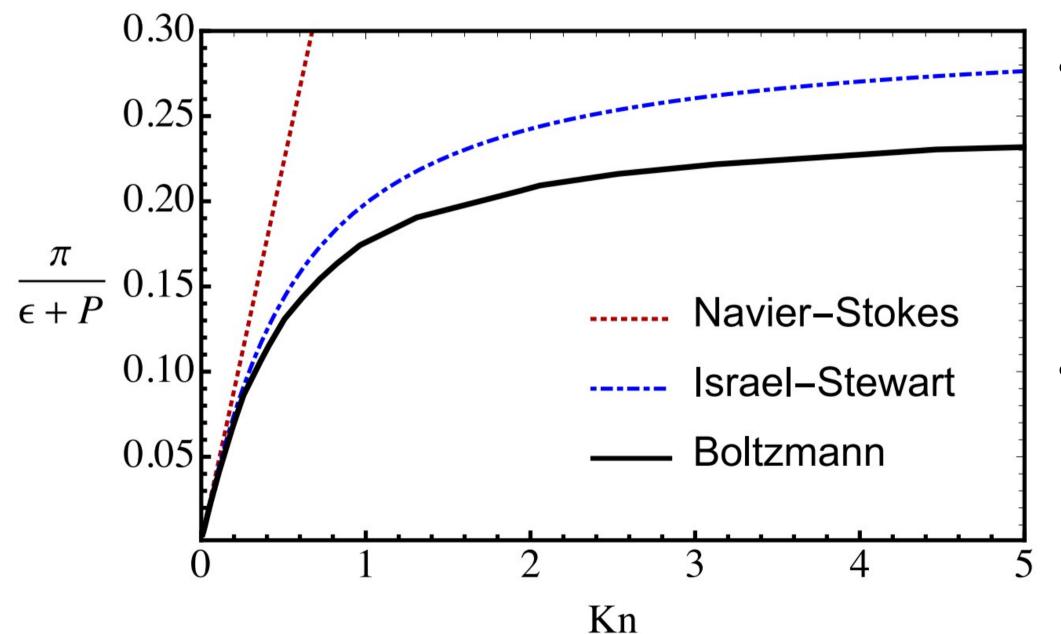
$$Kn = \frac{\ell_{\rm mfp}}{\tau} = \frac{1}{n_0 \tau_0 \sigma_T}$$



 Solution for the normalized shear-stress tensor relaxes to a constant (attractor!)

Relativistic gas of hard spheres

Attractor solution of the Boltzmann equation and DNMR theory can be calculated



- DNMR attractor is qualitatively and even quantitatively similar to the attractor of the Boltzmann equation
- Navier-Stokes only works for small
 Knudsen number

Can we study attractors in more general flow configurations?

In Bjorken flow (boost invariant), every fluid element is identical → same attractor for all.

In more realistic flow profiles, different fluid elements (in a Lagrangian view) may each approach their own local attractor ????

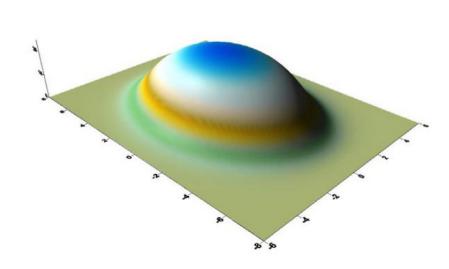
Here we investigate this using Lagrangian solvers: attractor behavior is easier to identify for each moving fluid element

Simulation

- ✓ We solve the fluid-dynamical equations using a relativistic version of the SPH algorithm
 Phys.Rev.C 80 (2009) 064901
- ✓ Ideal equation of state: ε = 3P
- \checkmark Transport coefficients: $\eta/s = cte$ $\tau_{\pi} = 5\frac{\eta}{\varepsilon + P}$
- Initial condition: Optical Glauber, central collision

$$\varepsilon_0 \sim 40 \text{ GeV/fm}^3$$

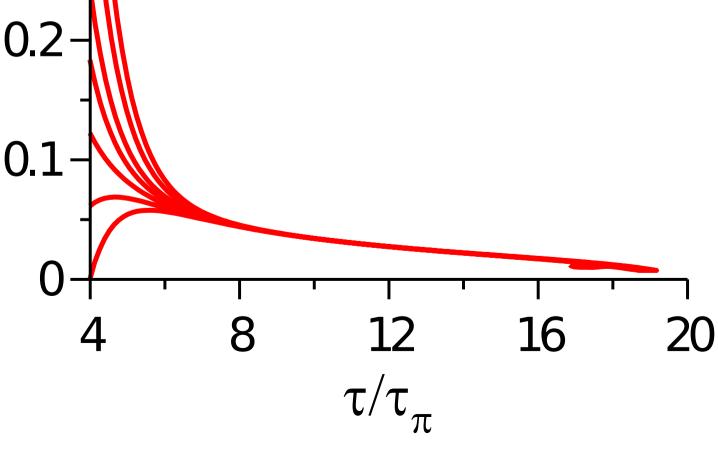
initial shear stress $\sim \epsilon + P$



Fluid element near the center (r=0.25 fm)

near
$$0.5$$
 0.4 0.4 0.4 0.4 0.3 0.2 0.1

$$\eta/s=1/8\pi$$



0.2 Fluid element near the center (r=0.25 fm) $\eta/s = 1/8\pi$ **Navier-Stokes** 0.080.04-Knudsen number starts

to increase again

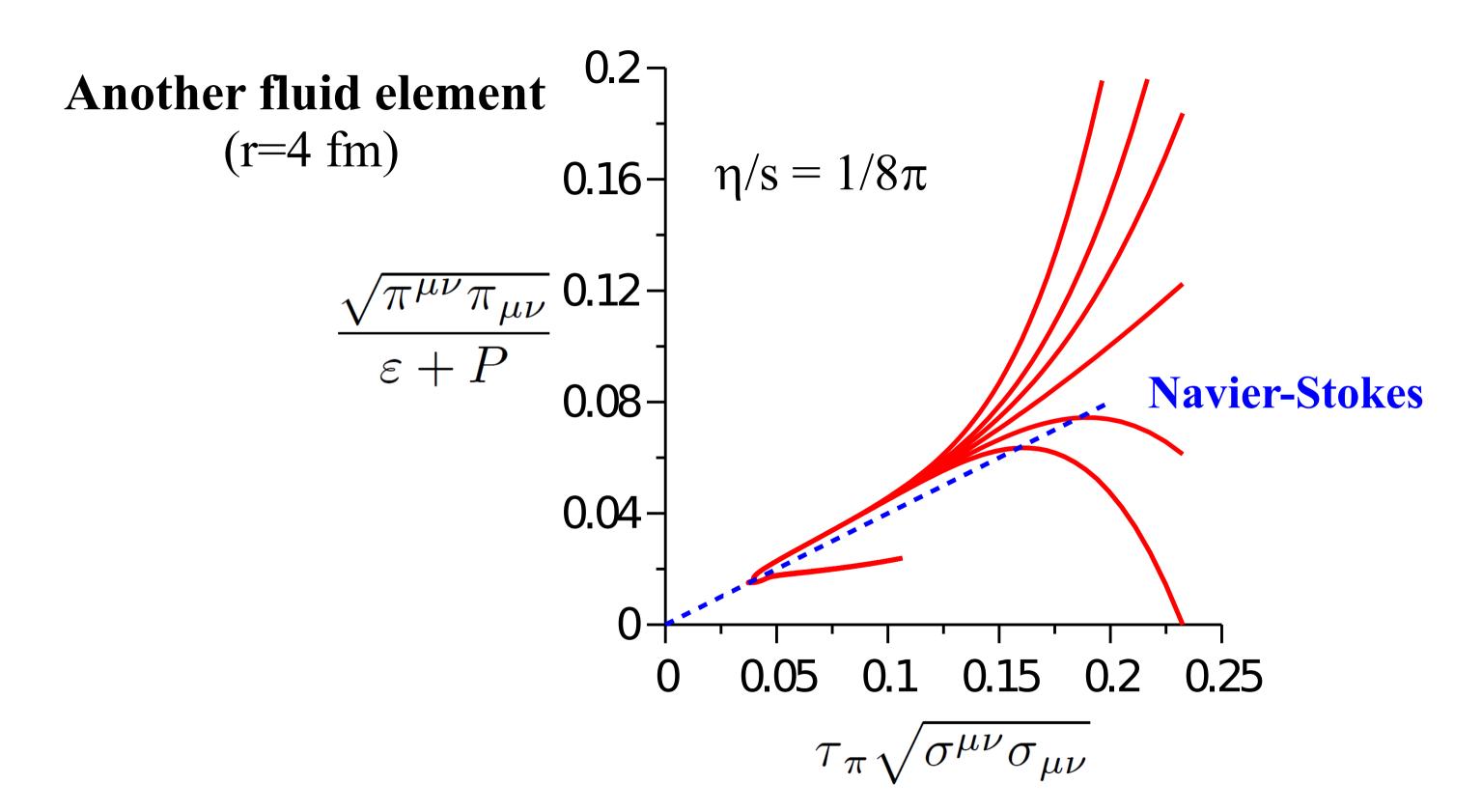
Fluid element starts to

Fluid element starts to deviate from equilibrium

$$au_{\pi} \sqrt{\sigma^{\mu
u} \sigma_{\mu
u}}$$

0.15

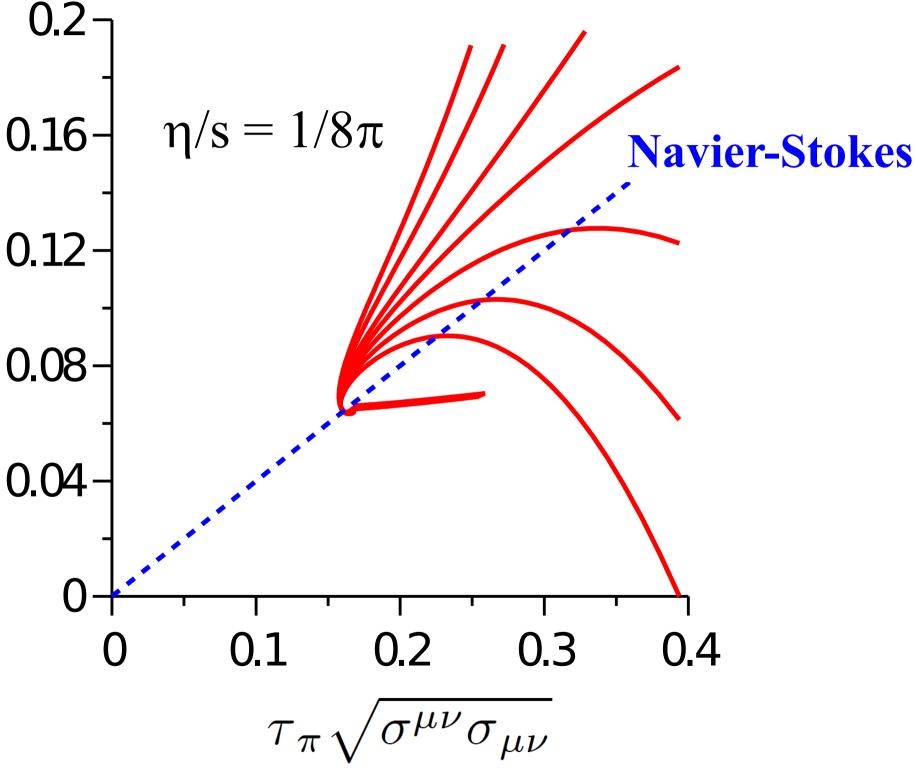
Another fluid element (r=2 fm) $\eta/s = 1/8\pi$ 0.16 **Navier-Stokes** 0.08-0.04-0.15 0.2 0.25



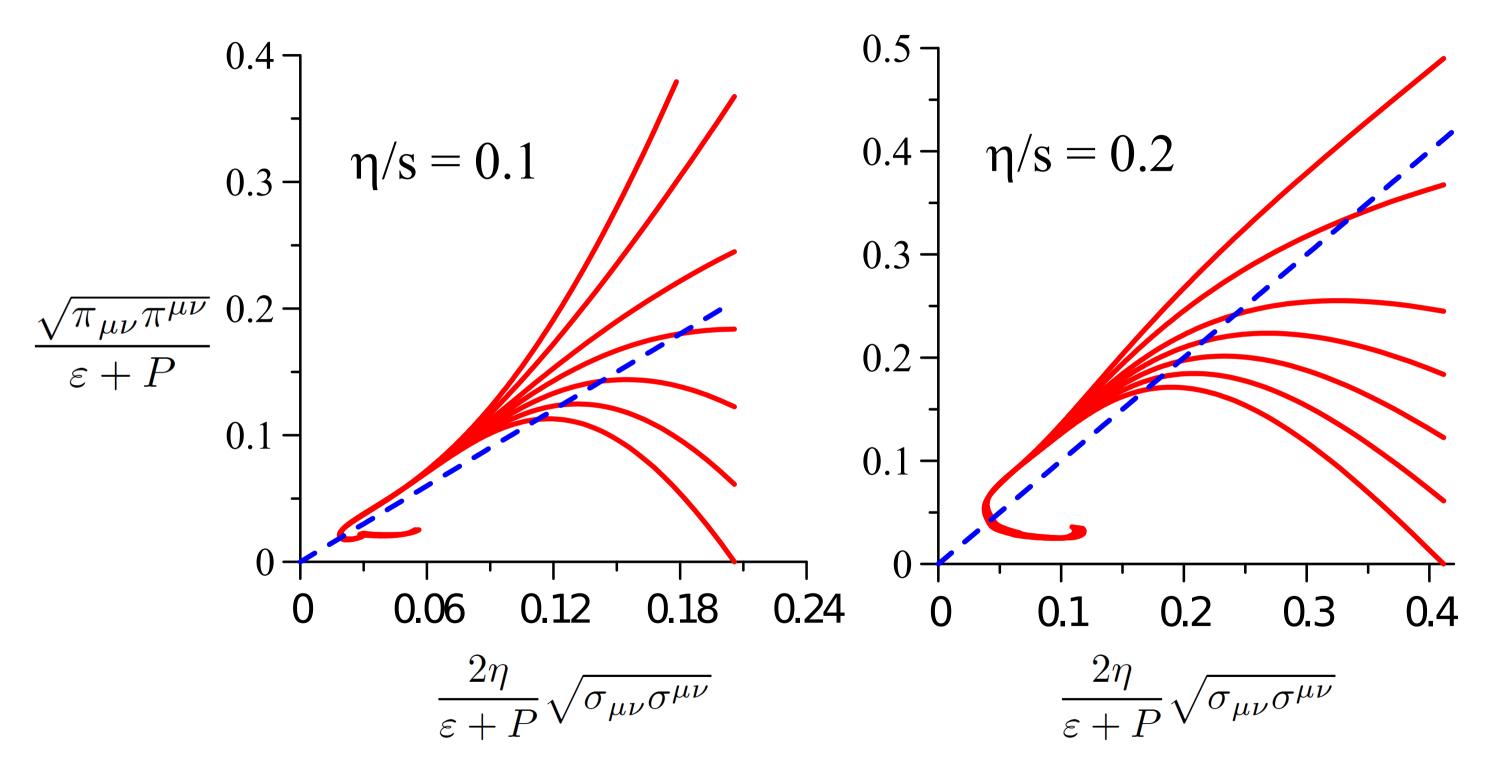
Fluid element at the edge (r=6 fm)

 $\frac{\sqrt{\pi^{\mu\nu}\pi_{\mu\nu}}}{arepsilon+P}$ 0.12

• Transient dynamics still 0.04goes away, but system does not display 0hydrodynamic behavior



Different viscosities



Far-from-equilibrium hydrodynamics for general flows

Consider the usual gradient expansion (Kn <<1)

$$\pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} + \eta_1\sigma_{\lambda}^{\langle\mu}\sigma^{\nu\rangle\lambda} + \eta_2\sigma_{\lambda}^{\langle\mu}\omega^{\nu\rangle\lambda} + \eta_3\omega_{\lambda}^{\langle\mu}\omega^{\nu\rangle\lambda} + \eta_4\theta\sigma^{\mu\nu} + \eta_5\nabla_{\perp}^{\langle\mu}P\nabla_{\perp}^{\nu\rangle}P + \eta_6\nabla_{\perp}^{\langle\mu}\nabla_{\perp}^{\nu\rangle}P + \mathcal{O}\left[K_N^3\right],$$

- When Kn ~1: Large rearrangement of the series
 - 3rd-order terms like $\sim \sigma^{\alpha\beta} \sigma_{\alpha\beta} \sigma^{\mu\nu}$ may be grouped with $2\eta \sigma^{\mu\nu}$

Resummation of transport coefficients $\eta \rightarrow \eta_R(Kn)$

$$\pi^{\mu\nu} = 2\eta^R \sigma^{\mu\nu} + \eta_1^R \sigma_{\lambda}^{\langle \mu} \sigma^{\nu \rangle \lambda} + \eta_2^R \sigma_{\lambda}^{\langle \mu} \omega^{\nu \rangle \lambda} + \eta_3^R \omega_{\lambda}^{\langle \mu} \omega^{\nu \rangle \lambda} + \eta_5^R \nabla_{\perp}^{\langle \mu} P \nabla_{\perp}^{\nu \rangle} P$$

$$+ \eta_6^R \nabla_{\perp}^{\langle \mu} \nabla_{\perp}^{\nu \rangle} P + \mathcal{O}\left[(K_N^R)^3 \right],$$

Example: slow-roll solution of Israel-Stewart theory

Simplified Israel-Stewart theory:

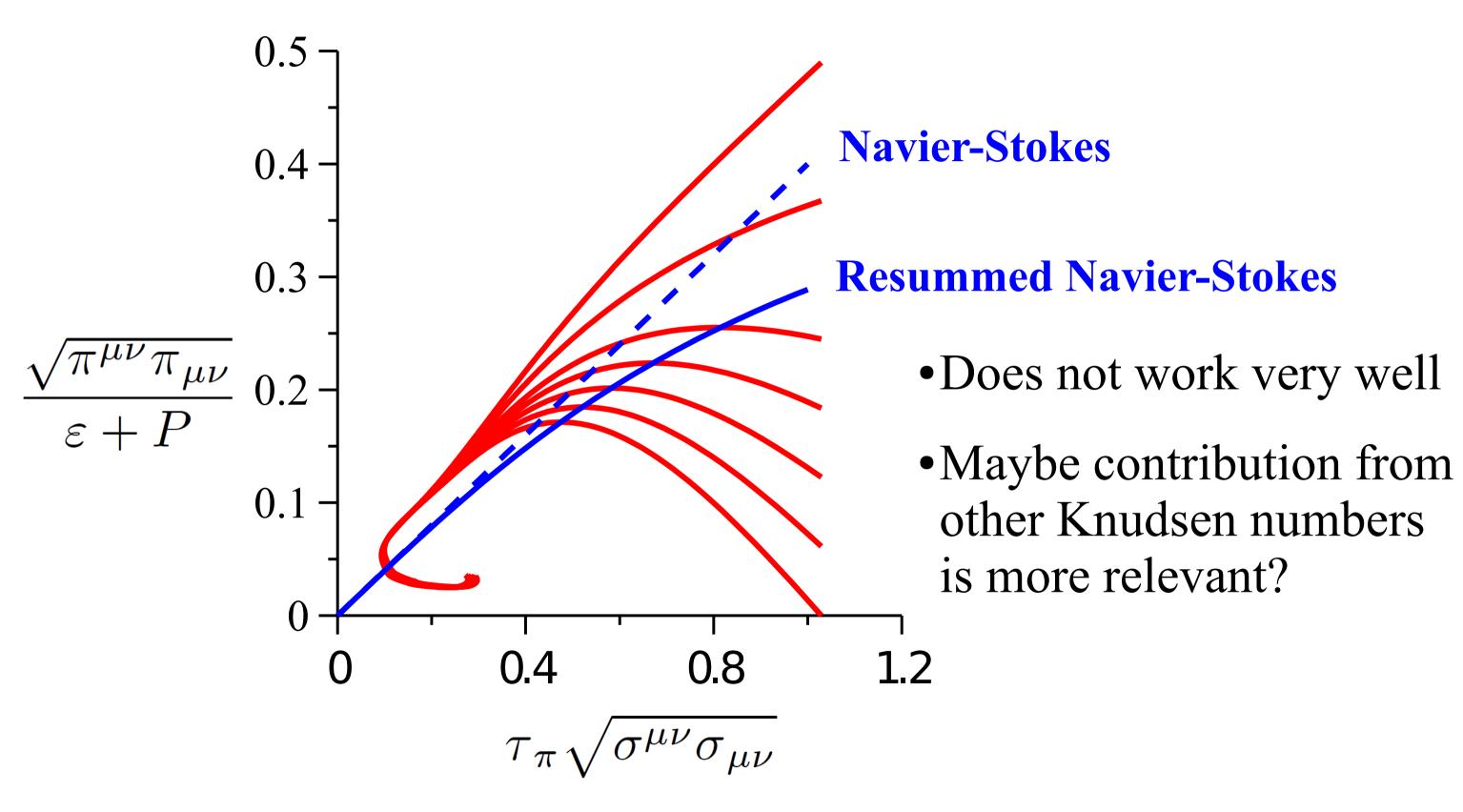
$$\chi^{\mu\nu} = \frac{\pi^{\mu\nu}}{\varepsilon + P} \qquad \tau_{\pi}D\chi^{\langle\mu\nu\rangle} + \chi^{\mu\nu} = \frac{2}{5}\tau_{\pi}\sigma^{\mu\nu} - \frac{4}{3}\tau_{\pi}\chi^{\mu\nu}\chi^{\alpha\beta}\sigma_{\alpha\beta}$$

Slow-roll expansion:
$$\epsilon \tau_{\pi} D X^{\langle \mu \nu \rangle} = -X^{\mu \nu} + \frac{2}{5} \tau_{\pi} \sigma^{\mu \nu} - \frac{4}{3} \tau_{\pi} X^{\mu \nu} X^{\alpha \beta} \sigma_{\alpha \beta}$$

 $X^{\mu \nu}(x; \epsilon) = \sum_{n=0}^{\infty} \epsilon^n X_n^{\mu \nu}(x)$

Leading solution:
$$X_0^{\mu\nu} = \frac{2}{5}\tau_\pi \sigma^{\mu\nu} S$$
 $S(K_N) = \frac{15}{16K_N^2} \left(\sqrt{1 + \frac{32}{15}K_N^2} - 1\right)$

$$K_N = \tau_\pi \sqrt{\sigma_{\mu\nu}\sigma^{\mu\nu}}$$





Conclusions

Attractor behavior, previously seen in symmetric (0+1)D Bjorken flow, also manifests locally in complex (2+1)D heavy-ion collisions

• Each fluid element can approach its own far-from-equilibrium attractor, provided it evolves long enough (i.e., not at the very edge of the fireball)

What is the correct functional form of the attractor for a general flow? Is it a function of one Knudsen number, or multiple independent invariants?