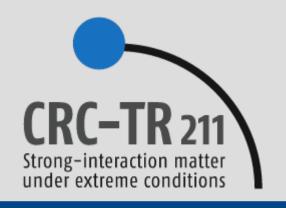
ECT* workshop: Complex structure of strong interactions in Euclidean and Minkowski space

Goldstone bosons at finite temperature

(Based on: PL, O. Philipsen, 2501.17120, 2506.XXXX)

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Outline

- 1. Goldstone's theorem in vacuum
- 2. Generalisation to finite temperature
- 3. A thermal particle? \rightarrow "Thermoparticle"
- 4. Thermal Goldstone bosons
- 5. Goldstone signatures from the lattice

1. Goldstone's theorem in vacuum

- <u>In words</u>: Goldstone's theorem states that the spontaneous breaking of a continuous symmetry implies the existence of massless (Goldstone) bosons
- QFT language: if j^{μ} is the conserved current associated with the symmetry, and A is some local field whose transformation under the symmetry has a non-trivial vev: $\langle \delta A \rangle = \lim_{R \to \infty} \langle [Q_R, A] \rangle \neq 0$, then:
 - \rightarrow The Fourier transform of $<[j^0(x), A(y)]>$ contains a $\delta(p^2)$ singularity
- In fact... current conservation and field locality means that $<\!\delta A>\neq 0$ implies the Fourier transform of $<[j^0(x), A(y)]>$ contains a $\delta(\omega)$ component as $\rho \to 0$. This is *independent* of the properties of the background state*
- The Goldstone "quasi-particle" $\delta(\omega)$ becomes a stable massless particle state $\delta(p^2)$ for relativistic QFTs

2. Generalisation to finite temperature

- For $T = 1/\beta > 0$, one defines: $\langle \Phi(x_1) ... \Phi(x_n) \rangle_{\beta} = Z^{-1} \operatorname{Tr} e^{-\beta H} \Phi(x_1) ... \Phi(x_n)$
- There are some immediate implications:
 - Lorentz invariance
 → but can retain rotational invariance
 - * **Spectral condition** $(H > 0) \times \rightarrow$ replaced by KMS condition
 - Field locality (causality) ✓ → this is important!
- Since current conservation and field locality are unaffected by T, the Fourier transform of $<[j^0(x), A(y)]>_{\beta}$ still contains a $\delta(\omega)$ as $p \to 0$
- Can we learn anything else about the properties of thermal Goldstone modes, e.g. what happens for p > 0?

Yes! The key is to determine how T modifies spectral functions $\rho_{AB}(\omega, \mathbf{p})$, the Fourier transform of the thermal expectation values $<[\Phi_A(x), \Phi_B(y)]>_{\beta}$ [Bros, Buchholz, PRD 58 (1998)]

2. Generalisation to finite temperature

• For (complex) scalar fields, the constraints imposed for T>0 imply that the spectral function has the representation*

$$\rho(\omega, \vec{p}) = \int_0^\infty ds \int \frac{d^3 \vec{u}}{(2\pi)^2} \ \epsilon(\omega) \, \delta(\omega^2 - (\vec{p} - \vec{u})^2 - s) \, \widetilde{D}_{\beta}(\vec{u}, s)$$

This is the T > 0 generalisation of the textbook *Källén-Lehmann* representation!

$$ho(\omega, ec{p}) = 2\pi\epsilon(\omega) \int_0^\infty \! ds \, \delta\!\left(p^2 - s
ight) arrho(s)$$

"Thermal spectral density"

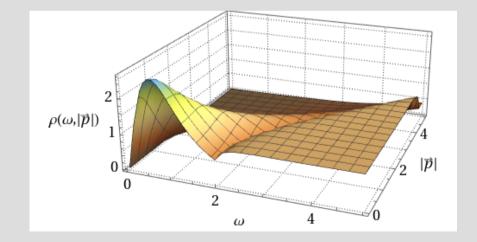
- T>0 effects amount to understanding: $\rho(s) \to \widetilde{D}_{\beta}(\boldsymbol{u},s)$, which tells us about the possible excitations that can exist in a thermal medium
- The non-trivial u dependence of $\widetilde{D}_{\beta}(u,s)$ controls the extent to which the spectral function can be off the mass-shell $p^2 = s$
- The s dependence determines whether the spectral function has energy ω thresholds, much like in the T=0 case

3. A thermal particle? \rightarrow "Thermoparticle"

<u>Proposition</u>: the medium contains "Thermoparticles": particle-like excitations which differ from collective quasi-particle modes, and show up as **discrete** contributions to $\widetilde{D}_{\beta}(\boldsymbol{u},s)$ [Bros, Buchholz, NPB 627 (2002)]

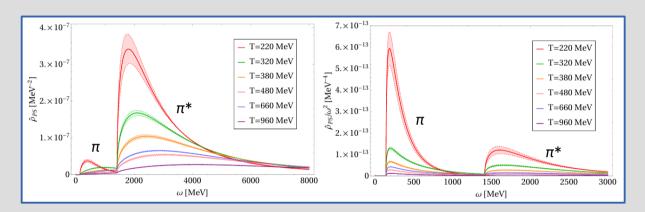
$$\widetilde{D}_{\beta}(\vec{u},s) = \widetilde{D}_{m,\beta}(\vec{u})\,\delta(s-m^2) + \widetilde{D}_{c,\beta}(\vec{u},s)$$

- o Thermoparticle components $\widetilde{D}_{\beta}(\textbf{\textit{u}})\delta(s-m^2)$ reduce to those of a vacuum particle with mass m in the limit $T \to 0$
- \rightarrow Non-trivial "Damping factor" $\widetilde{D}_{\beta}(\boldsymbol{u})$ results in thermally-broadened peaks in the spectral function: this parametrises the effects of collisional broadening
- ightharpoonup Component $\widetilde{D}_{c,\beta}(\boldsymbol{u},s)$ contains all other types of excitations, including those that are *continuous* in s



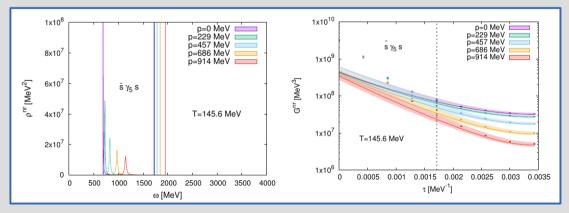
3. A thermal particle? \rightarrow "Thermoparticle"

- There is mounting evidence for low-energy thermoparticle excitations, e.g. spatial correlator $C_{PS}(z)$ of the pseudo-scalar meson operator $O_{PS}^a = \overline{\psi} \gamma_5 \frac{\tau^a}{2} \psi$ in lattice QCD
- Studies extracting pseudo-scalar spectral function in various channels:



Light-light pseudo-scalar meson (pion) channel [P.L., O. Philipsen, *JHEP* 10, 161 (2022)]

Light-strange (kaon) and strange-strange (eta) pseudo-scalar meson channels [D. Bala, O. Kaczmarek, P. L., O. Philipsen, and T. Ueding, *JHEP* 05, 332 (2024)]



Data in *all* channels consistent with a thermoparticle-type ground state: suggests light pseudo-scalar mesons (pions, kaons,..) still have a bound-state-like structure, even at high T

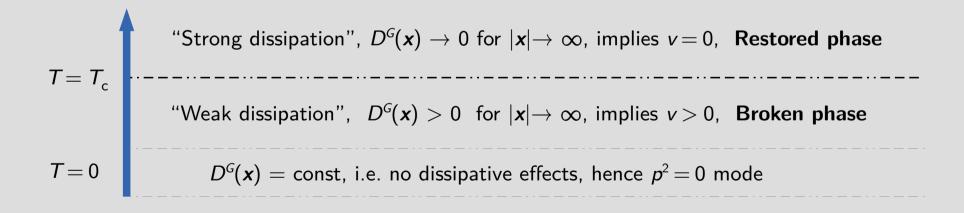
• Using the extra information given by the T>0 spectral representation, in [Bros, Buchholz, 1998] the authors were able to prove that the SSB condition $<\delta A>\neq 0$ implies that $<[j^0(x),\ A(0)]>_{\beta}$ contains a massless thermoparticle component, which in position space has the form:

$$D_{\beta}^{G}(\vec{x},s) = D_{\beta}^{G}(\vec{x})\delta(s)$$

- This a **thermal Goldstone boson**: in the $T \to 0$ limit $D^G(x) \to \text{const}$, hence the current-field spectral function contains a vacuum Goldstone component $\delta(p^2)$, as expected!
- When the damping factor $D^{G}(x)$ is non-trivial, this causes the stable massless Goldstone peak at $p^{2} = 0$ to become broadened

→ Describes the dissipative effects of the Goldstone moving through the thermal medium

- This analysis reveals some very important characteristics:
 - \rightarrow The thermal Goldstone mode is never on-shell for T > 0
 - \rightarrow This component can persist at *any* temperature, even if the symmetry is restored when $T > T_{\rm c}$
 - \rightarrow The order parameter v depends **entirely** on the functional form of the damping factor $D^G(x)$



 This captures the physics! Sufficiently strong dissipative effects destroy the long-range order and lead to symmetry restoration

- The phase of the theory is determined *entirely* by the dissipative effects experienced by the thermal Goldstone mode
- Only $D^G(x) \to 0$ for $|x| \to \infty$ is required to ensure the symmetry is restored, but this can happen at *any* rate
- If the damping factor has the functional form $D^{G}(x) \sim |x|^{-\varepsilon}$ with some $\varepsilon > 0$ at large |x|, the two-point function decays as a pure power-law

Interesting possibility: a symmetry could be restored at high temperatures without there being a finite correlation length!

- Analysis of *massive* thermoparticle states [Bros, Buchholz, (2002)] demonstrates that damping factors are fixed by the (asymptotic) dynamics of the theory
- If this is also the case for thermal Goldstone modes, this suggests that universality arguments alone may not be sufficient to fully characterise finite-temperature phase transitions

- If thermal Goldstone modes are present one can look for their signatures in (Euclidean) correlation functions
- For simplicity, consider the QFT of a single complex scalar field at finite temperature, with two-point function $C(\tau, \vec{x}) = \langle \phi(\tau, \vec{x}) \phi^{\dagger}(0) \rangle_{\beta}$
- If a thermal Goldstone mode is present, it follows from the thermoparticle structure, and the spectral function representation, that:

$$C^{G}(0, \vec{x}) = \frac{\coth\left(\frac{\pi|\vec{x}|}{\beta}\right)}{4\pi\beta|\vec{x}|} D_{\beta}^{G}(\vec{x})$$

• The mode dissipation is determined by the damping factor $D^{G}(x)$

 \rightarrow For $T \rightarrow 0$ the vacuum behaviour is recovered:

$$C^G(0, \vec{x}) \xrightarrow{T \to 0} \frac{\alpha_0}{4\pi^2 |\vec{x}|^2}$$

See: [PL,O. Philipsen, 2022]

- Now we know what the signatures of thermal Goldstone modes are, one can look for them in lattice data
- Consider a simple model with SSB: U(1) complex scalar field theory

$$\mathcal{L} = \frac{1}{2} \partial_{\mu} \phi^{\dagger} \partial^{\mu} \phi - \frac{1}{2} m^2 \phi^{\dagger} \phi - \frac{\lambda}{4!} (\phi^{\dagger} \phi)^2$$

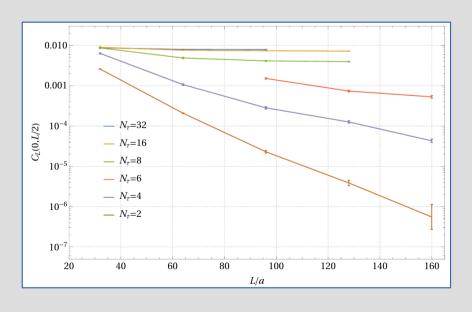
- In the broken phase at T=0 the model contains a massless Goldstone boson and a resonance-like mode
 - ightharpoonup Model expected* to undergo a second-order phase transition: for $T > T_{\rm c}$ the U(1) symmetry is restored, and $|v|^2 = <\phi><\phi^\dagger>=0$
- Investigate theory on a $L_{\tau} \times L^3$ lattice $(L_{\tau} = aN_{\tau}, L = aN_{s})$ with action

$$S = a^4 \sum_{x \in \Lambda_a} \left[\sum_{\mu} \left(\frac{1}{2} \Delta_{\mu}^f \phi^*(x) \Delta_{\mu}^f \phi(x) \right) + \frac{m_0^2}{2} \phi^*(x) \phi(x) + \frac{g_0}{4!} \left(\phi^*(x) \phi(x) \right)^2 \right]$$

 \rightarrow Avoid potential triviality by keeping lattice spacing fixed throughout, hence $T = (aN_{\tau})^{-1}$ is varied in discrete steps

- SSB does not occur in a finite spatial volume $V = L^3$, i.e. there is no notion of a "vev" on the lattice
 - \rightarrow One needs to perform an $L\rightarrow\infty$ extrapolation of lattice results!
- For the finite-volume correlator $C_L(\tau, \mathbf{x})$: $\lim_{L \to \infty} C_L(\tau, \vec{x}) \xrightarrow{|\vec{x}| \to \infty} |v|^2$
- Based on this property there are different approaches* for extracting $|v|^2$
- Given periodic spatial boundary conditions one can use:

$$|v|^2 = \lim_{L \to \infty} C_L(0, |\vec{x}| = L/2)$$



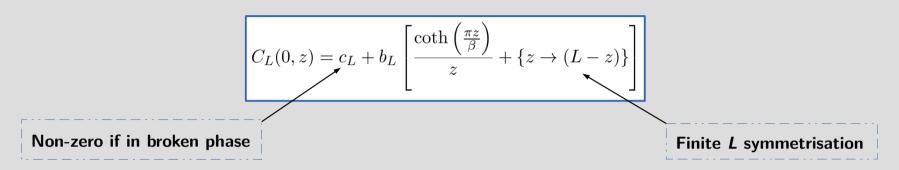
- $N_{\tau} = 8, 16, 32$ non-zero for large $L \rightarrow U(1)$ broken
- $N_{\tau} = 2, 4, 6$ vanishing for small $L \rightarrow U(1)$ restored
- Infinite-volume extrapolation requires a parametrisation for $C_I(\tau=0, |x|=z)$
- Use finite-volume version of $C^G(0, \vec{x}) = \frac{\coth\left(\frac{\pi |\vec{x}|}{\beta}\right)}{A 21|\vec{x}|} D^G_{\beta}(\vec{x})$

$$C^G(0, \vec{x}) = \frac{\coth\left(\frac{\pi |\vec{x}|}{\beta}\right)}{4\pi\beta |\vec{x}|} D^G_{\beta}(\vec{x})$$

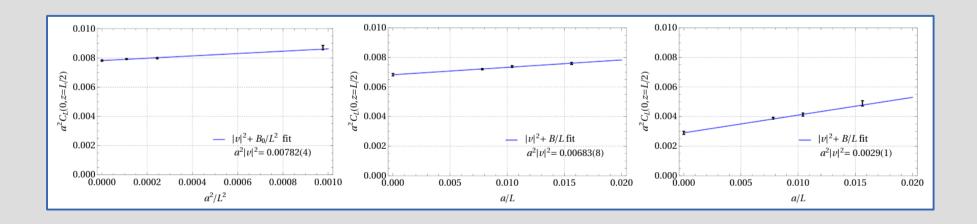
^{*} See eg. [H. Neuberger, PRL 60,(1988).]

Broken phase ($N_{\tau} = 8, 16, 32$):

• In this case we fit the ansatz (assuming $D^{G}(x) \sim \text{const}$):



• Functional form provides very good description of the data for each volume considered (L/a=32,64,96,128). Use this to do $L\rightarrow\infty$ extrapolation:

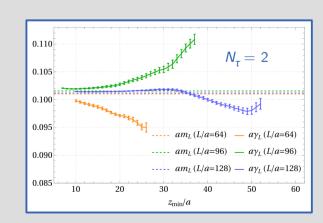


Restored phase $(N_{\tau} = 2, 4, 6)$:

- As outlined previously, the Goldstone mode can still persist in this regime, and would have the properties of a massless thermoparticle
- What is the structure of the damping factor $D^{G}(x)$? Use spatial correlator:

$$C_L(z) = d_L \left[e^{-m_L z} + \{ z \to (L - z) \} \right] \longrightarrow D_\beta^G(\vec{x}) = \alpha e^{-\gamma |\vec{x}|} \longrightarrow C^G(0, \vec{x}) = \frac{\coth\left(\frac{\pi |\vec{x}|}{\beta}\right)}{4\pi\beta |\vec{x}|} \alpha e^{-\gamma |\vec{x}|}$$

- Spatial correlator fits provide excellent description of data over full range [0,L/2] for each of the (large) volumes considered (L/a=64,96,128,160)
- Test damping by fitting: $C_L(0,z) = b_L \left[\frac{\coth(\frac{\pi z}{\beta})}{z} e^{-\gamma_L z} + \{z \to (L-z)\} \right]$
- Consistency of the thermoparticle hypothesis requires that the fit parameters γ_L and m_L approach one another in the infinite-volume $L \rightarrow \infty$ limit

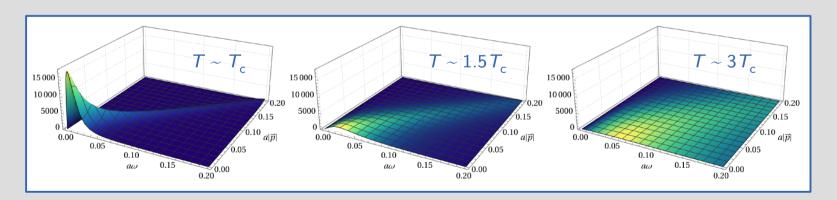


Restored phase $(N_{\tau} = 2, 4, 6)$:

- Data at each value of N_{τ} consistent with the correlator being dominated by an exponentially-damped massless thermoparticle \rightarrow a thermal Goldstone!
- One can use the extracted damping factor $D^G(\mathbf{x}) = \alpha \, e^{-\gamma |\mathbf{x}|} \, \text{to compute the spectral function}$ $\rho_G(\omega, \mathbf{p}) \, \text{of the Goldstone mode}$

$$\rho_G(\omega, \vec{p}) = \frac{4\alpha \,\omega \gamma}{(\omega^2 - |\vec{p}|^2 - \gamma^2)^2 + 4\omega^2 \gamma^2}$$

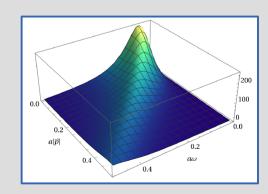
• Spectral properties are very different to the vacuum case $\rho_G(\omega, \mathbf{p}) \sim \delta(p^2)$



- Broadened peak structure around $p^2 = 0$ increases with temperature
 - → Represents the increasingly strong dissipative effects of the medium!

Summary & outlook

- Goldstone's theorem in vacuum has well-known consequences, but at finite temperature there still remain open questions
- One can use the non-perturbative constraints imposed by *causality* to gain new insights $\rightarrow \rho(\omega, \mathbf{p})$ have spectral representations. This narrows down the potential excitations that exist \rightarrow "Thermoparticles"
- SSB for T>0 implies existence of Goldstone modes which have the structure of massless thermoparticles → these modes can persist at any temperature, even if the symmetry is restored
- Phase structure is determined by the damping of the Goldstone mode:
 (i) Weak dissipation → SSB, (ii) Strong dissipation → symmetry restoration
- We find evidence for the existence of such modes in the U(1) complex scalar field theory on the lattice below and above $T_c \rightarrow$ General feature of T > 0 QFTs



Backup: Rigorous definition of SSB

• In order to define SSB rigorously one needs to define a regularised charged operator Q_R , and this only converges for $R \rightarrow \infty$ within commutator $[Q_R, A]$

$$Q_{\delta R} = \int d^4x \,\alpha_{\delta}(t) g_R(\vec{x}) j_0(t, \vec{x})$$

- (i) $g_R(x)=1$ for $|x| \le R$, $g_R(x)=0$ for $|x| \ge R(1+\varepsilon)$
- (ii) $a_{\delta}(t)$ has compact support, and $a_{\delta}(t){
 ightarrow}\delta(t)$
- The condition: $\lim_{R\to\infty} <[Q_R, A]>=q$ is then **always** well-defined, and q=0 is a necessary and sufficient condition for the existence of a charge operator Q, defined by: $< u, Qv> := \lim_{R\to\infty} < u, Q_Rv>$
 - \rightarrow If this is non-vanishing for any A then no such charge exists, i.e. SSB!
- SSB condition: $\lim_{\delta \to 0} \lim_{R \to \infty} \int_0^\infty ds \int \frac{d^3 \vec{u}}{(2\pi)^3} \frac{d^3 \vec{p}}{(2\pi)^3} \; \widetilde{D}_{\beta}^{(+)}(\vec{u}, s) \, \widetilde{g}(\vec{p}) \, \widetilde{\alpha} \left(\delta R \sqrt{((\vec{p}/R) \vec{u})^2 + s} \right) = iq\widetilde{\alpha}(0)$
- In the vacuum case this implies: $\widetilde{D}^{(+)}(\pmb{u},s) o iq (2\pi)^3 \delta(\pmb{u})\delta(s)$
- Value of q is **determined** by Goldstone damping factor $D^{(+)}(x)$ for $|x| \to \infty$

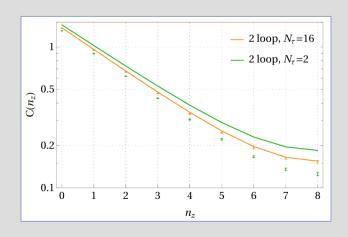
Backup: NRT theorem

• The NRT theorem implies that states with purely real dispersion relations $\omega = E(\mathbf{p})$ cannot exist in interacting theories when T > 0

QFT reason: thermal states satisfy the KMS condition, and this gives rise to very different spectral constraints than in the vacuum case

Physics reason: Dissipative effects of the thermal medium are everywherepresent \rightarrow always need to take these into account (i.e. always a width!)

- This has significant implications for perturbation theory: *neither free field,* nor quasi-particle propagators with real poles can form the basis of finite-temperature perturbative expansions [Landsman, Ann. Phys. 186, 141 (1988)]
- There is both analytic [Weldon, PRD 65 (2002)] and now numerical [PL, O. Philipsen, JHEP 08, (2024)] evidence for this perturbative breakdown



Backup: Perturbation theory

- How might one resolve inconsistencies of T > 0 perturbation theory?
- In the vacuum theory we know that the basis of perturbation theory is the Gell-Mann/Low formula [Landsman, Ann. Phys. 186, 141 (1988)]:

$$iG(x_1\cdots x_n) = \sum_{i_1\cdots i_n} Z_{i_1}^{1/2}\cdots Z_{i_n}^{1/2} \frac{\langle \Omega \mid T[\Phi_{i_1}(x_1)\cdots \Phi_{i_n}(x_n) U(\infty, -\infty)] \mid \Omega \rangle}{\langle \Omega \mid U(\infty, -\infty) \mid \Omega \rangle}.$$

$$U(t_1, t_2) = e^{iH_0[\Phi]t_1} e^{-iH[\Phi](t_1 - t_2)} e^{-iH_0[\Phi]t_2}$$

$$= T \exp\left(-i \int_{t_2}^{t_1} dt \ H_I(t)\right),$$

- Correlation functions of *interacting* fields can be computed from correlators of free fields \rightarrow This is derived from the fact that at large times the interactions between fields diminish: the asymptotic fields/states are free!
- The standard perturbative series is defined by expanding the exponential in the evolution operator as a series in the coupling parameter \rightarrow each term in the expansion is determined by the propagators of the asymptotic fields

Key point: free field propagators form the basis of perturbation theory at T=0 because the large-time states experience no interactions

Backup: Thermoparticles in QCD data

Goal: Extract information about the finite T spectral function $\rho_{\Gamma}(\omega, \mathbf{p})$ from data of *Euclidean* correlator $C_{\Gamma}(\tau, \vec{x}) = \langle O_{\Gamma}(\tau, \vec{x}) O_{\Gamma}(0, \vec{0}) \rangle_T$ $O_{\Gamma} = \text{scalar operator}$

Standard approach: extract $\rho_{\Gamma}(\omega, \mathbf{p})$ from temporal correlator $C_{\Gamma}(\tau, \mathbf{p})$

$$\widetilde{C}_{\Gamma}(\tau,\vec{p}) = \int_{0}^{\infty} \frac{d\omega}{2\pi} \frac{\cosh\left[\left(\frac{\beta}{2} - |\tau|\right)\omega\right]}{\sinh\left(\frac{\beta}{2}\omega\right)} \, \rho_{\Gamma}(\omega,\vec{p}) \qquad \qquad \rightarrow \text{ Problem is ill-conditioned, }$$
 need more information!

- Instead, one can use the **spatial** correlator, where one integrates $C_{\Gamma}(\tau, \mathbf{x})$ over $\{\tau, x, y\}$ and fixes a spatial direction z

$$C_{\Gamma}(z) = \int_{-\infty}^{\infty} \frac{dp_z}{2\pi} e^{ip_z z} \int_{0}^{\infty} \frac{d\omega}{\pi\omega} \, \rho_{\Gamma}(\omega, p_x = p_y = 0, p_z)$$

• It turns out that *if* thermoparticles exist, then they will give a distinct contribution to C(z)

$$C(z) \approx \frac{1}{2} \int_{|z|}^{\infty} dR \ e^{-mR} D_{m,\beta}(R)$$

[P.L., PRD 106 (2022); P.L., O. Philipsen, JHEP 10, 161 (2022)]

This component can be extracted *directly* from data

Backup: Thermoparticle characteristics

Given a specific QFT, what form should the damping factors take?

<u>Idea</u>: thermal scattering states are defined by imposing an asymptotic field condition [Bros, Buchholz, (2002)]:

Asymptotic fields Φ_0 are assumed to satisfy dynamical equations, but only at large x_0

In Φ^4 theory

$$(\partial^2 + m^2)\phi_0(x) + \frac{\lambda}{3!}\phi_0^3(x) \xrightarrow{|x_0| \to \infty} 0$$

- Since thermoparticles dominate the large-time behaviour of correlators, they are natural candidates for describing such states. It turns out that their damping factors $\widetilde{D}_{m.\beta}(u)$ are uniquely fixed by the asymptotic condition
- In Φ^4 theory one finds (where κ is a thermal width):

For
$$\mathbf{g} < \mathbf{0}$$
: $D_{m,\beta}^{(-)}(\vec{u}) = \frac{2\kappa}{\kappa^2} \delta(|\vec{u}| - \kappa),$
$$\widetilde{G}_{\beta}^{(-)}(k_0, \vec{p}) = \frac{1}{4|\vec{p}|\kappa} \ln \left[\frac{-k_0^2 + m^2 + (|\vec{p}| + \kappa)^2}{-k_0^2 + m^2 + (|\vec{p}| - \kappa)^2} \right]$$

For
$$\mathbf{g} < \mathbf{0}$$
:
$$\widetilde{D}_{m,\beta}^{(-)}(\vec{u}) = \frac{2\pi^2}{\kappa^2} \delta(|\vec{u}| - \kappa), \qquad \mathbf{For} \ \mathbf{g} > \mathbf{0}$$
:
$$\widetilde{D}_{m,\beta}^{(+)}(\vec{u}) = \frac{4\pi}{\kappa_0 \left(|\vec{u}|^2 + \kappa^2\right)},$$
$$\widetilde{G}_{\beta}^{(-)}(k_0, \vec{p}) = \frac{1}{4|\vec{p}|\kappa} \ln \left[\frac{-k_0^2 + m^2 + (|\vec{p}| + \kappa)^2}{-k_0^2 + m^2 + (|\vec{p}| - \kappa)^2} \right]$$
$$\widetilde{G}_{\beta}^{(+)}(k_0, \vec{p}) = \frac{i}{2|\vec{p}|\kappa_0} \ln \left[\frac{\sqrt{-k_0^2 + m^2 - i|\vec{p}| + \kappa}}{\sqrt{-k_0^2 + m^2 + i|\vec{p}| + \kappa}} \right]$$

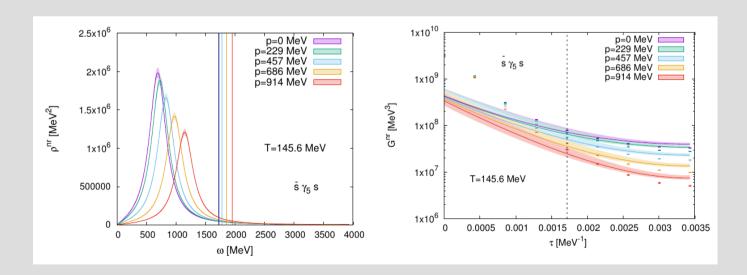
Backup: Robustness of spectral approach

• The robustness of the thermoparticle hypothesis can also be tested by comparing with different causal models, e.g. a Breit Wigner

$$\rho_{\rm BW}(\omega, \vec{p}) = \frac{4\omega\Gamma}{(\omega^2 - |\vec{p}|^2 - m^2 - \Gamma^2)^2 + 4\omega^2\Gamma^2},$$

$$C_{\rm BW}(z) = rac{e^{-\sqrt{m^2 + \Gamma^2}|z|}}{2\sqrt{m^2 + \Gamma^2}}.$$

• Same procedure as with the thermoparticle case: (i) extract the width parameter Γ and coefficient from the spatial lattice data (ii) use this to predict the corresponding temporal correlator



→ Data is *not* consistent with a Breit-Wigner-type ground state!