#### **Calculations of various muon decays in orbit**

Contents

- 1. Introduction (2 slides)
- 2. Calculating lepton wave functions (7 slides)
- 3. Decay in orbit (10 slides)
- 4.  $\mu$ -e conversion (8 slides)
- 5. Other lepton flavor violating processes

5-1.  $\mu^- \rightarrow e^- X$  (8 slides) 5-2.  $\mu^- \rightarrow e^- \gamma$  (7 slides)

5-3.  $\mu^- e^- \rightarrow e^- e^-$  (8 slides)

6. Summary

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## **1. Introduction**



### Various decays of muonic atoms

Standard processes

•  $\mu^- \to e^- \nu_\mu \overline{\nu}_e$  (Decay in orbit) •  $\mu^- \to e^- \nu_\mu \overline{\nu}_e \gamma$ 

The spectrum near the endpoint can be BG of  $\mu^- \rightarrow e^-$  conv.

- $\mu^{-}(Z, A) \rightarrow \nu_{\mu}(Z 1, A)$  (Nuclear muon capture)
- $\mu^{-}(Z, A) \rightarrow \nu_{\mu}(Z 1, A)\gamma$  (Radiative muon capture)

LFV processes

•  $\mu^- \rightarrow e^-$  •  $\mu^-(Z, A) \rightarrow e^+(Z-2, A)$ 

"Lepton Number Violation" discussed in Joe Sato's talk yesterday

•  $\mu^- \to e^- X$  •  $\mu^- \to e^- \gamma$ X: invisible boson

$$\mu^{-}e^{-} \rightarrow e^{-}e^{-}$$
  
 $e^{-}$  in atomic orbits

2/52

# 2. Calculating lepton wave functions

## Bound muon & scattering electron

For calculating the  $\mu^- \rightarrow e^-$  transition, we need wave functions of the bound  $\mu^-$  & the scattering  $e^$ to obtain the overlap integral.

(e.g.,  $\int dr \rho(r) \overline{\psi}_e(r) O \psi_\mu(r)$  for coherent  $\mu$ -e conv.)

3/52

✓ Here, the emitted  $e^-$  w.f. is not a plane wave:

Its wave function is **distorted** by the nuclear Coulomb potential. The formula for the decay rate gets complicated because the electron w.f. should be expanded by its angular momenta.

 $\checkmark$  In heavy nuclei, a bound muon is located near the nucleus.

Bohr radius:  $\sim (m_{\mu}Z\alpha)^{-1} \sim 2 \times \frac{137}{Z}$  fm Nuclear radius:  $\sim 1.2 \times A^{1/3}$  fm  $\checkmark$  comparable !!

rightarrow It is important to consider the finite nuclear size.

# Dirac equation with Coulomb potential 4/52

$$[i\partial_{\mu}\gamma^{\mu} - m + eA_{\mu}\gamma^{\mu}]\psi(\mathbf{r}) = 0$$

 $A_0$  is replaced with the nuclear Coulomb potential.

$$\Box \hspace{-1.5cm} \triangleright \hspace{-1.5cm} E\psi(\boldsymbol{r}) = [-i\boldsymbol{\alpha}\cdot\nabla + m\beta - V(\boldsymbol{r})]\psi(\boldsymbol{r})$$

Assuming V(r) is <u>spherical</u>, we obtain the radial Dirac equation,

$$\begin{cases} \frac{dg_{\kappa}(r)}{dr} + \frac{1+\kappa}{r}g_{\kappa}(r) - \left(E+m-V(r)\right)f_{\kappa}(r) = 0 \qquad \psi_{\kappa}(r) = \begin{pmatrix} g_{\kappa}(r)\chi_{\kappa} \\ if_{\kappa}(r)\chi_{-\kappa} \end{pmatrix} \\ \frac{df_{\kappa}(r)}{dr} + \frac{1-\kappa}{r}f_{\kappa}(r) + \left(E-m-V(r)\right)g_{\kappa}(r) = 0 \\ \\ \kappa \text{ is the index of angular momenta: } \kappa = \begin{cases} -(l+1) & (j=l+1/2) \\ l & (j=l-1/2) \\ l & (j=l-1/2) \end{cases} \\ l: \text{ orbital ang. mom.} \quad j: \text{ total ang. mom.} \end{cases}$$

e.g., s-wave corresponds to  $\kappa = -1$ .

### **Calculation for bound states**

5/52

What we want: the (binding) energy & wave functions

$$G(r) = rg(r) \quad \text{Boundary conditions:}$$

$$F(r) = rf(r) \quad G(0) = F(0) = 0$$

$$G(\infty) = F(\infty) = 0 \quad \frac{dG_{\kappa}(r)}{dr} + \frac{\kappa}{r}G_{\kappa}(r) - (E + m - V(r))F_{\kappa}(r) = 0$$

$$\frac{dF_{\kappa}(r)}{dr} - \frac{\kappa}{r}F_{\kappa}(r) + (E - m - V(r))G_{\kappa}(r) = 0$$

For a given *E*, we can solve the differential equations from both r = 0 and  $r = \infty$ .



If *E* is appropriate, it is expected that the two curves will be smoothly connected.

### **Calculation for bound states**

A way to find the appropriate energy to get a smooth w.f.:

Wronskian at the matching point  $r_m$ ,

 $W(E) = g_{0 \to m}(r_m)g'_{\infty \to m}(r_m) - g'_{0 \to m}(r_m)g_{\infty \to m}(r_m)$ 

(If *E* gives the appropriate binding energy, W(E) = 0.)

Step 1: Look for the zero-point of W(E):



Step 2: Normalize the wave functions to satisfy the normalization condition:

$$4\pi \int_0^\infty dr \, r^2 \{g(r)^2 + f(r)^2\} = 1$$

6/52

#### 7/52 **Bound wave functions** for ${}^{208}$ Pb (Z = 82, A = 208)



g(r) & f(r) of bound  $\mu^-$  (1s)

### **Calculation for scattering states**

Since the Coulomb potential violates the translation symmetry,

the momentum is no longer a good quantum number.

Multipole expansion: 
$$\psi_p(\mathbf{r}) = 4\pi \sum_{\kappa} i^{l_{\kappa}} (l_{\kappa}, m, 1/2, s | j_{\kappa}, v) Y_{l_{\kappa}}^{m*}(\hat{p}) e^{-i\delta_{\kappa}} \psi_{\kappa}(r)$$
  
"p" indicates the boundary condition at infinity.

> What we want: wave functions & phase shift  $\delta$  for a given <u>E</u>



The normalization is determined by the boundary condition:

8/52

 $\lim_{r \to \infty} g(r) = \cos \delta g^{reg}(r) - \sin \delta g^{irr}(r)$ 

 $g^{reg}(r)$  &  $g^{irr}(r)$  : solutions regular & irregular at r = 0of Dirac eq. without the potential

### Scattering wave functions for $^{208}$ Pb (7 - 92, 4 - 209)

for  ${}^{208}$ Pb (Z = 82, A = 208)



Distortion effects can be interpreted as follows:

① The value near r = 0 gets larger, to <u>enhance</u> the overlap with the bound muon.

(2) The effective momentum (wave number) gets larger,

to suppress the overlap with the bound muon.

# 3. Decay in orbit

J. Heeck, R. Szafron, & YU, PRD105, 053006; arXiv:2110.14667.



### **Decay in orbit (DIO)**

10/52



- the "standard" decay of a muon,  $\mu^- \rightarrow e^- \nu_\mu \overline{\nu}_e$
- electron energy :  $E_e < m_\mu E_b E_N$

cf. For free muons,  $E_e \leq m_{\mu}/2$ 

due to the energy-momentum conservation of a two-body decay.

For muonic atoms, the nuclear recoil makes the tail of the spectrum.

### Total width of DIO

The "standard" decay width of muonic atoms is given by

the sum of the DIO width and the nuclear capture width:

11/52

 $\Gamma_{\text{total}} = \Gamma_{\text{DIO}} + \Gamma_{\text{NC}}$ 

The width of DIO slightly depends on the nuclei:



(The Huff factor is important to determine  $\Gamma_{NC}$  from experiments.)

#### 12/52

### **Spectrum of the DIO electron**

E (MeV)

- $\mu^- \rightarrow e^- \nu_\mu \overline{\nu}_e$  process in the nuclear Coulomb field
- $\mu^-$  : bound wave
- Solving Dirac eq. numerically \_\_\_\_^
- $e^-$  : distorted wave

finite size of the nucleus

V

 $\rightarrow$  notential V(r)density o(r)charge

rge density 
$$\rho(r) \rightarrow \text{potential } V(r)$$
  
 $(r) = -e \int_{0}^{\infty} dr' r'^{2} \left[ \frac{\theta(r-r')}{r} + \frac{\theta(r'-r)}{r'} \right] \rho(r)$   
 $ightarrow \text{ bind a constraint of the sector of the$ 

### **DIO spectrum**

13/52



- $\checkmark$  Considering the finite size of the nucleus is important (unless Z is small) because the bound muon exists close to the nuclei.
- ✓ The  $e^-$  distortion enhances the low energy region

& suppresses high energy region.

### **Behavior near the endpoint**

14/52

important to estimate BG of  $\mu^- \rightarrow e^-$  conv.

• nuclear recoil O. Shanker, PRD25, 1847 (1981). A. Czarnecki et al., PRD84, 013006 (2011).



15/52

### **Radiative corrections**



$$E_{\text{end}}' = E_{\text{end}} + \frac{\alpha m_{\mu} (Z\alpha)^2}{\pi} \left( \frac{11}{9} - \frac{2}{3} \log \left[ \frac{2m_{\mu} Z\alpha}{m_e} \right] \right) \qquad \delta = \frac{2\alpha}{\pi} \left( \log \left[ \frac{2m_{\mu}}{m_e} \right] - 1 \right) \simeq 0.023$$

#### 16/52

### **Nuclear charge distribution**

#### measured by electron scattering

W. Boeglin *et al.*, Nucl. Phys. A **477** (1988) 399. J. Wesseling *et al.*, PRC **55** (1997) 2773.

Different "fitting functions" are used:

#### **Examples of fitting funtions**

1. three(two)-parameter Fermi (3pF, 2pF)

$$\rho(r) = \frac{\rho_0}{1 + \exp\frac{r-c}{z}} \left(1 + \omega \frac{r^2}{c^2}\right) \qquad \omega = 0$$

2. there-parameter Gaussian (3pG)

$$\rho(r) = \frac{\rho_0}{1 + \exp\frac{r^2 - c^2}{z^2}} \left(1 + \omega \frac{r^2}{c^2}\right)$$

H. De Vries *et al.*, Atom. Data Nucl. Data Tabl. **36** (1987) 495.
G. Fricke *et al.*, Atom. Data Nucl. Data Tabl. **60** (1995) 177.
A. A. Kabir, PhD thesis, Kent State Univ., US (2015).

3. modified-harmonic oscillator (MHO)

$$\rho(r) = \rho_0 \left( 1 + \omega \frac{r^2}{a^2} \right) \exp\left( -\frac{r^2}{a^2} \right)$$

- 4. Fourier-Bessel (FB)
- 5. Sum of Gaussians (SOG)
- There are nuclides which has only data of "radius". I. Angeli & K. P. Marinova, Atom. Data Nucl. Data Tabl. **99** (2013) 69.  $\rho(r) = \frac{\rho_0}{1 + \exp \frac{r - c}{0.52} \int_{r_0}^{r_0} \frac{1}{1 + \exp \frac{r - c}{0.52} \int_$

### **Numerical results**

Table of  $E'_{end}$  & B for various nuclides and models

	$E'_{\rm end}/{\rm MeV}$	$B_{1\mathrm{pF}}/\mathrm{MeV^{-6}}$	$B_{\rm FB}/{ m MeV^{-6}}$	$B_{\rm SOG}/{ m MeV^{-6}}$	$B_{\rm MHO}/{ m MeV^{-6}}$	$B_{\rm 3pF}/{ m MeV^{-6}}$	$B_{3pG}/MeV^{-6}$
$^{4}_{2}\mathrm{He}$	104.150			$2.53\times 10^{-20}$			
${}_{3}^{6}\mathrm{Li}$	104.637	$1.20\times10^{-19}$	$1.29\times 10^{-19}$				
$_{3}^{7}$ Li	104.779	$1.28\times 10^{-19}$			$1.31\times 10^{-19}$		
$^9_4\mathrm{Be}$	104.949	$4.97\times10^{-19}$			$4.96\times10^{-19}$		
${}^{10}_5\mathrm{B}$	104.99	$1.52\times 10^{-18}$	$1.56\times 10^{-18}$		$1.50\times10^{-18}$		
${}^{11}_{5}{ m B}$	105.044	$1.53\times 10^{-18}$			$1.52\times 10^{-18}$		
${}^{12}_6\mathrm{C}$	105.059	$3.55\times10^{-18}$	$3.53\times 10^{-18}$	$3.55\times10^{-18}$			
${}^{13}_{6}{ m C}$	105.097	$3.56\times10^{-18}$			$3.63\times10^{-18}$		
$^{14}_{7}N$	105.094	$7.04\times10^{-18}$				$7.13\times10^{-18}$	
$^{16}_{8}O$	105.106	$1.22\times 10^{-17}$	$1.19\times 10^{-17}$	$1.20\times 10^{-17}$	$1.19\times 10^{-17}$		
${}^{19}_{9}{ m F}$	105.118	$1.85\times 10^{-17}$				$1.87\times 10^{-17}$	
$^{20}_{10}$ Ne	105.081	$2.79\times10^{-17}$				$2.87\times10^{-17}$	
$^{22}_{10}$ Ne	105.108	$2.89\times10^{-17}$				$2.89\times10^{-17}$	
$^{23}_{11}$ Na	105.063	$4.35\times10^{-17}$					
$^{24}_{12}\mathrm{Mg}$	105.011	$6.17\times10^{-17}$		$6.29\times10^{-17}$		$6.10\times10^{-17}$	
				•			

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J. Heeck, R. Szafron, & YU, PRD105, 053006; arXiv:2110.14667.

17/52

 $E_b \& E'_{end}$ 

18/52



### **B** coefficient



- > The uncertainty from fitting functions is about 10%.
- The isotope differences are often of the same order

or even larger than the uncertainty.

19/52

➤ How large is the quadrupole deformation effects?

# 4. $\mu$ -e conversion

J. Heeck, R. Szafron, & YU, NPB980, 115833; arXiv:2203.00702.



# (Coherent) $\mu^- \rightarrow e^-$ conversion

20/52



- charged lepton flavor violating (CLFV) process
- the energy of emitted  $e^-$ :  $E_e = m_\mu E_b E_{\text{recoil}} \sim 105 \text{ MeV}$  $\begin{bmatrix} E_b : \text{binding energy of muon} \\ E_{\text{recoil}} : \text{nuclear recoil energy} \end{bmatrix} \sim 0.5 \text{ MeV}$  for Al

# Effective Lagrangian beyond the SM <sup>21/52</sup>



### Branching ratio

R. Kitano *et al.*, PRD**84**, 013006 (2011).

22/52

$$BR_{\rm SI} = \frac{32G_F^2}{\Gamma_{\rm capture}} \left[ \left| C_{D,L} \frac{D}{4} + \sum_{N=p,n} \left( C_{S,L}^{NN} S^{(N)} + C_{V,R}^{NN} V^{(N)} \right) \right|^2 + \{L \Leftrightarrow R\} \right]$$

• overlap integrals

$$D = \frac{4m_{\mu}}{\sqrt{2}} \int_0^\infty dr \, r^2 \left[-E(r)\right] \left(g_e^- f_{\mu}^- + f_e^- g_{\mu}^-\right)$$
  
electric field

$$S^{(p)} = \frac{1}{2\sqrt{2}} \int_{0}^{\infty} dr \, r^{2} Z \rho^{(p)} \left(g_{e}^{-} g_{\mu}^{-} - f_{e}^{-} f_{\mu}^{-}\right) \qquad \rho^{(p)} : \text{proton density}$$

$$S^{(n)} = \frac{1}{2\sqrt{2}} \int_{0}^{\infty} dr \, r^{2} N \rho^{(n)} \left(g_{e}^{-} g_{\mu}^{-} - f_{e}^{-} f_{\mu}^{-}\right) \qquad \rho^{(n)} : \text{neutron density}$$

$$V^{(p)} = \frac{1}{2\sqrt{2}} \int_{0}^{\infty} dr \, r^{2} Z \rho^{(p)} \left(g_{e}^{-} g_{\mu}^{-} + f_{e}^{-} f_{\mu}^{-}\right)$$

$$V^{(n)} = \frac{1}{2\sqrt{2}} \int_{0}^{\infty} dr \, r^{2} N \rho^{(p)} \left(g_{e}^{-} g_{\mu}^{-} + f_{e}^{-} f_{\mu}^{-}\right)$$

## **Overlap integrals (dipole integral D)**



> For Z < 30, the uncertainty due to  $\rho_c$  is less than 2%,

but grows to 8% for large Z.

23/52

> Isotope dependence exceeds the uncertainty at medium Z.



For  $S^{(p)}$  and  $V^{(p)}$ , the uncertainty is estimated to range from 5% at low Z to 10% at high Z.

> For  $S^{(n)}$  and  $V^{(n)}$ , the uncertainty is also 5% at low Z,

but more than 10% at high Z.

### **Quadrupole deformation**

L. Borrel, D. G. Hitlin, S. Middleton, arXiv:2401.15025

25/52



(also shown yesterday by A. Czarnecki)

#### 26/52

### **Complementarity of targets**

cf. S. Davidson, Y. Kuno, & M. Yamanaka, PLB790 (2019) 380.

$$BR_{\rm SI} = \frac{32G_F^2}{\Gamma_{\rm capture}} \left[ |\boldsymbol{v} \cdot \boldsymbol{C}_L|^2 + |\boldsymbol{v} \cdot \boldsymbol{C}_R|^2 \right]$$
$$\left[ \boldsymbol{C}_L = \left( C_{D,R}, C_{V,L}^{pp}, C_{S,R}^{nn}, C_{V,L}^{nn}, C_{S,R}^{nn} \right) \right]$$

coefficient-space vector :

$$\boldsymbol{v} = \left(\frac{D}{4}, V^{(p)}, S^{(p)}, V^{(n)}, S^{(n)}\right)$$

If their directions of v are different, these targets are "complimentary".

misalignment angle : 
$$\theta_{Al} = \arccos\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}_{Al}}{|\boldsymbol{v}||\boldsymbol{v}_{Al}|}\right)$$

Large  $\theta_{Al}$  means high complimentarity to Al.

### **Complementarity of targets**





High-Z targets have large complementary with AI.

- Low-Z targets
  - Li-7, Ti-50 have large  $\theta_{Al}$ .

27/52

- cf. A/Z = 2.33 for Li-7
  - A/Z = 2.27 for Ti-50
  - (A/Z = 2.08 for Al-27)
- Ti would be a suitable next target for Al-based experiment ?
- cf. Ti-50, Ti-49, Cr-54 : natural abundance low Li-7, V-51 : (practically preferable?) natural abundance >90%

## **5. Other LFV processes**

# Other LFV processes in muonic atoms

μ<sup>-</sup> → e<sup>-</sup>X
 X. G. i Tormo, D. Bryman, A. Czarnecki, & M. Dowling, PRD84, 113010; arXiv:1110.2874.
 YU, PRD102, 095007; arXiv:2005.07894.

•  $\mu^- \rightarrow e^- \gamma$ 

YU, M. Yamanaka, & Y. Kuno, PRD111, 035017; arXiv:2411.10304.

•  $\mu^-e^- \rightarrow e^-e^-$ 

M. Koike, Y. Kuno, J. Sato, & M. Yamanaka, PRL105, 121601; arXiv:1003.1578.
YU, Y. Kuno, J. Sato, T. Sato, & M. Yamanaka, PRD93, 076006; arXiv:1603.01522.
YU, Y. Kuno, J. Sato, T. Sato, & M. Yamanaka, PRD97, 015017; arXiv:1711.08979.
Y. Kuno, J. Sato, T. Sato, YU, & M. Yamanaka, PRD100, 075012; arXiv:1908.11653.

## **5-1.** $\mu^- \to e^- X$

YU, PRD102, 095007; arXiv:2005.07894.

### $\mu^+ \rightarrow e^+ X$ searches

29/52

➤ A. Jodidio *et al.* PRD **34**, 1967 (1986).

- +  $1.8 \times 10^7 \ \mu^+$  that was highly polarized
- search for  $e^+$  emitted in opposite direction for  $\mu^+$  polarization
- ${\rm Br}(\mu^+ \to e^+ X) < 2.6 \times 10^{-6}$  for  $m_X = 0$



Mu3e Collab. A. Schöning, Talk at Flavour and Dark Matter Workshop, Heidelberg, September 28 (2017).

• Br <  $10^{-8}$  (for 25MeV <  $m_X$  < 95MeV )

## $\mu^- \rightarrow e^- X$ in a muonic atom

originally proposed by X. G. i Tormo et al., PRD 84, 113010 (2011).

Advantages over free muon decay

1. less background

---: 
$$\mu^+ \rightarrow e^+ X$$
 (free)  
---:  $\mu^+ \rightarrow e^+ \nu_e \overline{\nu}_\mu$  (free)  
:  $\mu^- \rightarrow e^- X$  ( $\mu$ -gold)  
:  $\mu^- \rightarrow e^- \overline{\nu}_e \nu_\mu$  ( $\mu$ -gold

different peak positions of signal & BG



30/52

- 2. more information : "spectrum", "dependence on nucleus", ...
- 3. huge # of muonic atoms in coming experiments (COMET, Mu2e, DeeMe)

Disadvantages

✓ non-monochromatic signal

✓ shorter lifetime of muonic atom

### **Effective models**

#### A. Scalar X

(e.g. majoron induced by R-parity violation, ...)

also analyzed by X. G. i Tormo et al., PRD 84, 113010 (2011).

31/52

$$\mathcal{L}_{S0} = g_{S0}(\overline{e}\mu)X + [H.c.]$$

derivative coupling

yukawa coupling

(e.g. majoron, familon, axion, ...)

$$\mathcal{L}_{S1} = \frac{g_{S1}}{\Lambda_{S1}} (\overline{e} \gamma_{\alpha} \mu) \partial^{\alpha} X + [H.c.]$$

B. Vector X

$$\mathcal{L}_{V1} = \frac{g_{V1}}{\Lambda_{V1}} \left( \overline{e} \sigma_{\alpha\beta} \mu \right) X^{\alpha\beta} + [H.c.] \\ X^{\alpha\beta} = \partial^{\alpha} X^{\beta} - \partial^{\beta} X^{\alpha}$$

 $e^-$  spectrum ( $m_X = 0$ )

32/52



Spectrum does not strongly depend on properties of X.

The sharper peak is obtained for the lighter nucleus because the width reflects the shape of the bound muon w.f.

#### e<sup>-</sup> spectrum near endpoint

33/52



We can see spectra depending on operators!

# Characteristic behavior of spectrum <sup>34/52</sup>



$$\frac{d\Gamma}{dE_e} = \frac{g_Y^2}{4\pi^2} p_e p_X \sum_{\kappa} (2j_{\kappa} + 1) |I_{\kappa}|^2$$
$$I_{\kappa} = \int_0^\infty dr r^2 j_{l_{\kappa}}(p_X r) \{g_{p_e}^{\kappa}(r) g_{\mu}^{1s}(r) - f_{p_e}^{\kappa}(r) f_{\mu}^{1s}(r)\}$$

# **Characteristic behavior of spectrum**

✓ Main contribution comes from s-wave of emitted electron.

s-wave 
$$(\kappa = -1)$$
 amplitude  

$$I_{-1} = m_{\mu} \int_{0}^{\infty} dr r^{2} j_{0}(p_{X}r) \{g_{p_{e}}^{-1}(r)g_{\mu}^{1s}(r) - f_{p_{e}}^{-1}(r)f_{\mu}^{1s}(r)\}$$

$$j_{0}(p_{X}r) \simeq const. \text{ near the endpoint}$$



# Characteristic behavior of spectrum <sup>36/52</sup>



Q. The reduced mass treatment is insufficient to include the nuclear mass?

## **5-2.** $\mu^- \rightarrow e^- \gamma$

YU, M. Yamanaka, & Y. Kuno, PRD111, 035017; arXiv:2411.10304.

## $\mu^- \rightarrow e^- \gamma$ in a muonic atom

• Rare decay of  $\mu^-$  in orbit (not free  $\mu^+$ )

- signal: a pair of  $e^- \& \gamma$  with  $E_e + E_{\gamma} = m_{\mu} E_b$ ( $E_b$ : binding energy)
- Typically, the signal  $e^- \& \gamma$  are emitted back-to-back with  $E_e \simeq E_{\gamma} \simeq 50$  MeV
  - $\checkmark~$  But it's not a strict two-body decay,

the spectrum is smeared.

- As well as dipole operator,
   <u>diphoton operator</u>(FF & FF̃) can be studied.
  - $\checkmark$  The diphoton ope. can be directly restricted

as  $\mu^+ \rightarrow e^+ \gamma \gamma$ .

Disadvantages:

- Muonic atoms have shorter lifetime than free muons.
- Invariant mass  $m_{e\gamma} \neq m_{\mu}$ ; Although the energy is (approximately) conserved, the 3-momentum is not in  $\mu^- \rightarrow e^-\gamma$ .



37/52

### **Decay rate**

38/52

> Nucleus is treated as a static Coulomb potential.

$$d\Gamma = \frac{d^3 p_e}{(2\pi)^3 2E_e} \frac{d^3 p_{\gamma}}{(2\pi)^3 2E_{\gamma}} (2\pi) \delta \left(E_e + E_{\gamma} - E_{\mu}\right) \frac{1}{2} \sum_{spins} |\mathcal{M}|^2$$

transition amplitude

$$\mathcal{M} = -\frac{2im_{\mu}}{v^{2}} \int d^{3}r \overline{\psi}_{e}\left(\boldsymbol{r}\right) \sigma_{\alpha\beta} \left(D_{L}P_{L} + D_{R}P_{R}\right) \psi_{\mu}^{1s}\left(\boldsymbol{r}\right) p_{\gamma}^{\alpha} \epsilon^{s_{\gamma}*\beta} \exp\left(-i\boldsymbol{p}_{\gamma}\cdot\boldsymbol{r}\right)$$
$$-\frac{4i}{v^{3}} \int d^{3}r \overline{\psi}_{e}\left(\boldsymbol{r}\right) \left(C_{L}P_{L} + C_{R}P_{R}\right) \psi_{\mu}^{1s}\left(\boldsymbol{r}\right) p_{\gamma}^{\alpha} \epsilon^{s_{\gamma}*\beta} \exp\left(-i\boldsymbol{p}_{\gamma}\cdot\boldsymbol{r}\right) \left\langle N \mid F_{\alpha\beta} \mid N \right\rangle$$
$$-\frac{4i}{v^{3}} \int d^{3}r \overline{\psi}_{e}\left(\boldsymbol{r}\right) i\gamma_{5} \left(\tilde{C}_{L}P_{L} + \tilde{C}_{R}P_{R}\right) \psi_{\mu}^{1s}\left(\boldsymbol{r}\right) p_{\gamma}^{\alpha} \epsilon^{s_{\gamma}*\beta} \exp\left(-i\boldsymbol{p}_{\gamma}\cdot\boldsymbol{r}\right) \left\langle N \mid \tilde{F}_{\alpha\beta} \mid N \right\rangle$$

 $\rightarrow$  The field strength F is replaced with nuclear electric field E.

$$\langle N | F_{\alpha\beta} | N \rangle = \begin{cases} -E_i & (\alpha = i, \beta = 0) \\ E_j & (\alpha = 0, \beta = j) \\ 0 & (\alpha = i, \beta = j) \end{cases} \qquad \left\langle N \left| \tilde{F}_{\alpha\beta} \right| N \right\rangle = \begin{cases} -\epsilon_{ijk} E_k & (\alpha = i, \beta = j) \\ 0 & (\alpha = 0 \text{ or } \beta = 0) \end{cases}$$

 $\therefore$  In this calculation, we restrict ourselves to the case that the proton number is small.  $\rightarrow$  We assume "The nucleus is point-charge." & "The electron is plane wave."

### Z dep. of the decay rate ( $Z \le 30$ )

39/52



Upper limits of BR from the past experiments





(green dashed: case that dipole & diphoton are equally interfered)

✓ diphoton: momentum distribution spreads more than dipole
 (dipole & diphoton are interfered, but 分布の形状が異なる=完全に打ち消されることはない)

### Backgrounds

41/52

✓ Electrons & photons are produced in the ordinary muon decays:



▶ BG1: 2つのミューオンから出た電子と光子のエネルギー和が

▶ BG2:  $\mu^- \rightarrow e^- \nu_\mu \overline{\nu}_e \gamma$  におけるニュートリノが低エネルギーである事象 (physics BG) (ただしこれは無視できるほど小さい)

### BG $e^-$ & $\gamma$ spectra

42/52



### **Effective branching ratio**

43/52

Benchmark :	$r_{e\gamma} = 78 \text{ ps} \qquad \Delta_{x+y} =$	ergy resolution	Effective BR $B_{acc} = R_{\mu}\Delta t_{e\gamma} \frac{\Delta\Omega_{e\gamma}}{4\pi} f_{acc}$					
Signal ( <mark>diphoton</mark> )	Al ( $Z = 13$ ) $4.6 \times 10^{-12}$	Zn ( $Z = 30$ ) 2.2 × 10 <sup>-11</sup>	$R_{\mu}$ : produced $\mu^{-}$ / time $\Delta t_{e\gamma}$ : time resolution					
Signal (dipole)	$1.2 \times 10^{-13}$	$2.0 \times 10^{-14}$	$\Delta\Omega_{e\gamma}$ : angular resolution					
accidental BG	$1.2 \times 10^{-10}$	$1.2 \times 10^{-11}$	$f_{acc}$ : the rate that BG $e^- \& \gamma$ satisfy the energy condition					
BG $(\mu^- \rightarrow e^- \nu \overline{\nu} \gamma)$	$1.5 \times 10^{-15}$	$5.7 \times 10^{-16}$	large BGs					
$\int_{e_{\gamma}} constraint \theta_{e_{\gamma}} to optimize diphoton signal  \theta_{e_{\gamma}}^{peak} - 0.01 < \theta_{e_{\gamma}} < \theta_{e_{\gamma}}^{peak} + 0.01 = = = = = = = = = = = = = = = = = = =$								
	<b>AI (</b> <i>Z</i> = 13 <b>)</b>	<b>Zn (</b> <i>Z</i> = 30 <b>)</b>						
Signal (diphoton)	$7.1 \times 10^{-14}$	$3.3 \times 10^{-13}$	Signal is twice larger					
Signal (dipole)	$1.2 \times 10^{-17}$	$1.6 \times 10^{-17}$	than BG !					
accidental BG	$8.5 \times 10^{-13}$	$1.1 \times 10^{-13}$	↑ There may be rooms					
BG $(\mu^- \rightarrow e^- \nu \overline{\nu} \gamma)$	$2.6 \times 10^{-19}$	$6.1 \times 10^{-19}$						

## **5-3.** $\mu^- e^- \rightarrow e^- e^-$

YU, Y. Kuno, J. Sato, T. Sato, & M. Yamanaka, PRD93, 076006; arXiv:1603.01522. YU, Y. Kuno, J. Sato, T. Sato, & M. Yamanaka, PRD97, 015017; arXiv:1711.08979.

# $\mu^-e^- \rightarrow e^-e^-$ in a muonic atom 44/52

M. Koike, Y. Kuno, J. Sato & M. Yamanaka, Phys. Rev. Lett. **105**, 121601 (2010).



# Effective Lagrangian for $\mu^-e^- \rightarrow e^-e^-$





### **Branching ratio**

47/52



Phys. Rev. Lett. **105**,121601 (2010).

# 48/52 To improve calculation for decay rate

✓ previous formula of CLFV decay rate by Koike et al.

Note

$$\Gamma_{\mu^- e^- \to e^- e^-} = 2\sigma v_{\rm rel} |\psi^e_{1S}(0)|^2 \propto (Z-1)^3$$

- $\succ$  "Z dependence" comes from only  $|\psi_{1S}^e(0)|^2$  (always  $\Gamma \propto (Z-1)^3$ )
- $\succ$  emitted  $e^-$ s are expected to be back-to-back with equal energies



More quantitative estimation is needed ! (important for large Z)

### Formulation for decay rate

49/52

$$\begin{split} \Gamma &= \sum_{f} \sum_{\overline{\iota}} (2\pi) \delta(E_{f} - E_{i}) \left| \left\langle \psi_{e}^{p_{1},s_{1}} \psi_{e}^{p_{2},s_{2}} \middle| H \middle| \psi_{\mu}^{1s,s_{\mu}} \psi_{e}^{1s,s_{e}} \right\rangle \right|^{2} \\ \text{partial wave expansion to express the distortion} \\ &\psi_{e}^{p,s} = \sum_{\kappa,\mu,m} 4\pi \, i^{l_{\kappa}} (l_{\kappa},m,1/2,s|j_{\kappa},\mu) Y_{l_{\kappa},m}^{*}(\hat{p}) e^{-i\delta_{\kappa}} \psi_{p}^{\kappa,\mu} \\ &\kappa : \text{index of angular momentum} \\ \text{solving "Dirac eq. with $\phi$" numerically to get radial wave functions} \\ &\frac{dg_{\kappa}(r)}{dr} + \frac{1+\kappa}{r} g_{\kappa}(r) - (E+m+e\phi(r))f_{\kappa}(r) = 0 \\ &\frac{df_{\kappa}(r)}{dr} + \frac{1-\kappa}{r} f_{\kappa}(r) + (E-m+e\phi(r))g_{\kappa}(r) = 0 \\ \end{split}$$

### **Upper limits of BR (contact process)**

50/52



# 51/52 Upper limits of BR (photonic process)



# 6. Summary

### Summary

52/52

> For quantitively calculating muon decays in nuclei, we need to solve the Dirac equation with the nuclear Coulomb field, for  $\mu^- \& e^-$ .

Some LFV processes on nuclei are discussed:

• DIO,  $\mu^- \rightarrow e^- \nu_\mu \overline{\nu}_e$ 

•  $\mu^- \rightarrow e^-$  conversion

Other possible LFVs

- $\mu^- \rightarrow e^- X$
- $\mu^- \rightarrow e^- \gamma$
- $\mu^-e^- \rightarrow e^-e^-$