

Introduction to $\mu \rightarrow e$ Conversion

Yoshitaka KUNO RCNP, Osaka University

April 14th, 2025 Workshop on "Lepton Flavour Change in Nuclei" ECT*, Trento, Italy

Thanks for the organizers, my last visit to ECT* was 1998.

Outline



1: CLFV Physics 2: Muon CLFV 3: What is $\mu^- \rightarrow e^-$ Conversion ? 4: $\mu^- \rightarrow e^-$ Conversion Experiment 5: $\mu^- \rightarrow e^-$ Conversion Phenomenology 6: $\mu^- \rightarrow e^+$ Conversion 7: Future : PRISM/PRIME 8: Summary



CLFV Physics

Charged Lepton Flavour Violation (CLFV)

老

CLFV in the Standard Model



S.T. Petcov, Sov.J. Nucl. Phys. 25 (1977) 340 W.J. Marciano et al.. Phys. Lett. B 67 (1977) 303 B.W. Lee et al., Phys. Rev. Lett. 38 (1977) 937 B.W. Lee et al., Phys. Rev. D 16 (1977) 1444.



BSM Physics Scale Reach



from European particle Physics Strategy Update (2019)

light colour: present dark colour: future prospect

5

Various CLFV Theoretical Models



The MECO Experiment to Search for Coherent Conversion of Muons to Electrons September 27, 2002 6

Model Dependent CLFV Predictions CLFV Predictions (for $\mu \rightarrow e\gamma$ and μ -e conversion),



Muon CLFV



Present Upper Limits for CLFV Searches

process	present limit	future
$\overline{ au o \mu \eta}$	$< 6.5 \times 10^{-8}$	$10^{-9} - 10^{-10}$
$ au o \mu \gamma$	$< 6.8 \times 10^{-8}$	
$ au o \mu \mu \mu$	$< 3.2 \times 10^{-8}$	
$\tau \rightarrow eee$	$< 3.6 \times 10^{-8}$	
$\overline{K_L \to e\mu}$	$< 4.7 \times 10^{-12}$	
$K^+ \to \pi^+ e^- \mu^+$	$< 1.3 \times 10^{-11}$	
$B^0 \to e\mu$	$< 7.8 \times 10^{-8}$	
$B^+ \to K^+ e \mu$	$< 9.1 \times 10^{-8}$	
$\overline{Z^0 \to e\mu}$	$< 7.5 \times 10^{-7}$	
$Z^0 \to e \tau$	$< 1.2 \times 10^{-5}$	
$Z^0 o \mu \tau$	$< 9.8 \times 10^{-6}$	
$H^0 \to e\mu$	$< 3.5 \times 10^{-4}$	
$H^0 \to e \tau$	$< 3.7 \times 10^{-3}$	
$H^0 \to \mu \tau$	$< 2.5 \times 10^{-3}$	
$\mu^+ \to e^+ \gamma$	$< 4.2 \times 10^{-13}$	10^{-14} (MEG II)
$\mu^+ \to e^+ e^+ e^-$	$< 1.0 \times 10^{-12}$	10^{-16} (Mu3e)
$\mu^{-}\mathrm{Au} \rightarrow e^{-}\mathrm{Au}$	$< 7.0 \times 10^{-13}$	10^{-17} (COMET, Mu2e)
$\mu^{-}\mathrm{Ti} \rightarrow e^{+}\mathrm{Ca}$	$< 3.6 \times 10^{-11}$	10^{-17} (COMET, Mu2e)
$\mu^+e^- \rightarrow \mu^-e^+$	$< 8.3 \times 10^{-11}$	

<mark>x10⁻⁴</mark>



7.1 History - how it all began

In the fall of 1976 rumors spread about an experiment performed at SIN for the search of the decay $\mu \rightarrow e\gamma$. A debate was going on, whether or not the decay had been observed. The rumors traveled from SIN via email to R. Eichler at Stanford and from him to a graduate student in the lecture-class of James Bjorken. The next week, J. Bjorken in turn gave the students an exercise to compute the decay rate and also confronted his colleague Steven Weinberg with the rumor. It took a few weeks after Weinberg's talk at the APS meeting to reach the New York Times. There it read on February 8th 1977: *Experimenters in Switzerland have reportedly observed an "impossible" transmutation of atomic particles. This has thrown the world community of theoretical physicists into a frenzy of speculations, calculations and publications (S. Weinberg).* This inspired R. Hofstadter of Stanford to initiate an experiment at LAMPF for $\mu^+ \rightarrow e^+\gamma$ to try to resolve the dispute around the SIN experiment.

Permanent link:

by R. Eichler

https://doi.org/10.3929/ethz-b-000526612

$\mu^+ \rightarrow e^+ \gamma$: History

Year	90% CL on $\mathcal{B}(\mu \to e\gamma)$ Collaboration/Lab		Reference
1947	$1.0 imes 10^{-1}$	Chalk River	Hincks and Pontecorvo [1948]
1948	.04	Washington University	Sard and Althaus [1948]
1955	2.0×10^{-5}	Nevis	Steinberger and Lokanathan [1955
1959	$7.5 imes 10^{-6}$	Liverpool	O'Keefe et al. [1959]
1959	2.0×10^{-6}	Nevis	Berley <i>et al.</i> [1959]
1959	1.0×10^{-5}	Rochester	Davis <i>et al.</i> [1959]
1959	1.2×10^{-6}	CERN	Ashkin <i>et al.</i> [1959]
1960	1.2×10^{-6}	LBL	Frankel <i>et al.</i> [1960]
1961	2.5×10^{-5}	Carnegie	Crittenden et al. [1961]
1962	$1.9 imes 10^{-7}$	LBL	Frankel <i>et al.</i> [1962]
1962	$6.0 imes 10^{-8}$	Nevis	Bartlett et al. [1962]
1963	$4.3 imes 10^{-8}$	LBL	Frankel <i>et al.</i> [1963]
1964	2.2×10^{-8}	Chicago	Parker <i>et al.</i> [1964]
1971	2.9×10^{-8}	Dubna	Korenchenko et al. [1971]
1977	$3.6 imes 10^{-9}$	TRIUMF	Depommier et al. [1977]
1977	1.1×10^{-9}	SIN	Povel <i>et al.</i> [1977]
1979	1.9×10^{-10}	LAMPF	Bowman <i>et al.</i> [1979]
1982	1.7×10^{-10}	LAMPF	Kinnison $et al.$ [1982]
1986	4.9×10^{-11}	LAMPF/Crystal Box	Bolton <i>et al.</i> [1986, 1988]
1999	1.2×10^{-11}	LAMPF/MEGA	Brooks <i>et al.</i> [1999]
2010	2.8×10^{-11}	PSI/MEG	Adam <i>et al.</i> [2010]
2011	2.4×10^{-12}	PSI/MEG	Adam <i>et al.</i> [2011]

What is $\mu^- \rightarrow e^-$ Conversion



Muonic Atom

- µ⁻ is stopped in material and it forms a muonic atom.
- •µ⁻ cascades down from excited states to the ground (1s) state by emitting Auger electrons and muonic X rays.
- In the ground state, μ⁻
 either decays in orbit or
 is captured by a
 nucleus.

Bound Muon Decay in Orbit (DIO) $\mu^- + N \rightarrow e^- \nu \overline{\nu} + N$ Nuclear Muon Capture (NMC) $\mu^- N(A, Z) \rightarrow \nu_\mu N(A, Z - 1)$



Lifetime of Muonic Atoms

 $\frac{1}{\tau_{\mu a tom}} = \Lambda_{\text{Cap}} + Q\Lambda_{\text{free}}$ $\Lambda_{\text{CAP}} \propto Z_{\text{eff}}^4$ Q: Huff factor $Q = 0.975 \quad \text{for Iron (Z=26)}$ $Q = 0.844 \quad \text{for lead (Z=82)}$

nucleus	Z	lifetime (ns)
С	6	2027
Al	13	864
Ti	22	330
Cu	29	164
Pb	82	74



$\mu^- \rightarrow e^-$ Conversion : Signal

Neutrinoless Muon to Electron Conversion

$$\mu^{-}N(A,Z) \to e^{-}N(A,Z)$$

$$\Delta L_{\mu} = -\Delta L_e = -1$$



Event Signature for the transition to the ground state :

a single mono-energetic electron

$$E_{\mu e} = \frac{(m_N + m_\mu - B_\mu)^2 - m_N^2 + m_e^2}{2(m_N + m_\mu - B_\mu)} \\ \sim m_\mu - B_\mu - E_{recoil}$$

nucleus	Z	$E_{\mu e}$ (MeV)
AI	13	104.97
S	16	104.70
Ti	22	104.30
Cu	82	103.50
Au	79	95.56
Pb	82	94.02

Coherency

Coherency

When the initial and final states are the same (ground state), all the nucleons can contribute to the process, and the rate of conversion is enhanced by A^2 (or Z^2).

Transitions to excited states are suppressed.

Definition of Conversion Rate

$$\operatorname{CR}(\mu^{-}\mathrm{N} \to e^{-}\mathrm{N}) \equiv \frac{\Gamma(\mu^{-}\mathrm{N} \to e^{-}\mathrm{N})}{\Gamma(\mu^{-}\mathrm{N} \to \nu_{\mu}X)}$$

Experimental Advantages of $\mu^- \rightarrow e^-$ Conversion (1)

(1) Separation of Signals and Normal Muon Decay



Experimental Advantages of $\mu^- \rightarrow e^-$ Conversion (2)

(2) Ability to Use High Intensity Beams

	Measurements	Major background	intensity (R_{μ})
μ→eγ	coincidence of decay particles	accidentals $(\propto R_{\mu}^2)$	<10 ⁸ /sec
µ→eee	coincidence of decay particles	accidentals $(\propto R_{\mu}^2)$	<10 ⁽⁸⁻⁹⁾ /sec ?
µ-e conversion	no coincidence	beam related	10 ¹¹ /sec

accidental background = more than one muon decays at the same time to mimic the event signature

Disadvantages of $\mu^- \rightarrow e^-$ Conversion

 (R_{μ})

The cost of muon beam lines is high!

20

Backgrounds for $\mu^- \rightarrow e^-$ Conversion

Muon-induced backgrounds

Beam-related backgrounds

Other backgrounds

Bound muon decay in orbit (DIO) $(\mu^- + N \to e^- \nu_\mu \bar{\nu}_e + N)$ Radiative nuclear muon capture (RMC) $(\mu^- N \to \nu_{\mu} N' \gamma, \gamma \to e^+ e^-)$ Particles from muon nuclear capture Radiative nuclear pion capture (RPC) $(\pi^- N \to N' \gamma, \gamma \to e^+ e^-)$ Beam electrons Muon decay in flights Neutron induced background Antiproton induced background

Cosmic-ray induced background False tracking others

Bound Muon Decay in Orbit (DIO)

$$\mu^- + N(A, Z) \rightarrow e^- + \nu_\mu + \overline{\nu}_e + N(A, Z - 1)$$



Bound Muon Decay in Orbit (DIO)

DIO spectra for AI DIO endpoint spectrum for AI



A. Czarnecki and X. i Tormo and W. Marciano, Physical Review D 84, 013006 (2011)

Bound Maon Decay in Orbit (DIO)

000 1000000

880

DIO endpoint spectrum for AI B coefficient



A. Czarnecki and X. i Tormo and W. Marciano, Physical Review D 84, 013006 (2011)

Bound Muon Decay in Orbit (DIO)



from COMET Phase-I TDR



Good momentum resolution is needed.

25

Radiative Muon Capture (RMC)

$$\mu^- + N(A, Z) \to N(A, Z - 1) + \nu + \gamma; \quad \gamma \to e^+ e^-$$

Using an appropriate target

Maximum photon energy (endpoint) :

$$E_{\gamma}^{\max} = m_{\mu} - B_{\mu} - E_{\text{rec}} + M(A, Z) - M(A, Z - 1)$$

Mass Requirement of Muon target N(A, Z):

$$E_{\mu e} > E_{\gamma}^{\max} \implies M(A, Z) < M(A, Z-1)$$

	$E_{\mu e}$	E_{γ}^{\max}	k _{max}
AI	104.9	101.8	90
Au	95.56	93.81	88
Ti	104.3	102.5	93

 k_{max} is the empirical endpoint of the spectrum determined by the data that is fitted to the Primakoff (closure approximation) model?

Radiative Pion Capture (RPC)

$$\pi^- + N \to N' + \gamma; \quad \gamma \to e^- e^+ \quad (m_\pi > m_\mu)$$

Using a pulsed beam, a delayed time window is employed for measurement.

10⁻⁸

10⁻¹⁰

10⁻¹²

10⁻¹⁴

10⁻¹⁶

10⁻¹⁸

10⁻²⁰

10⁻²²

10⁻²⁴

10⁻²⁶

10⁻²⁸

10⁻³⁰

10⁻³²

stopped pions/ proton

Total proton on target (POT): 10^{21} $10^{21} \times 10^{-2} \times 10^{-3} \times 10^{-18} = 0.01$

The measurement window starts at 700 ns after the beam arrival makes total 10⁻¹⁸ suppression

High Z targets cannot be used due to their short lifetimes.



Unconvoluted

$\mu^- \rightarrow e^-$ Conversion Experiments



$\mu^- \rightarrow e^-$ Conversion : History

			-	
Year	90% Limit	Lab/Collaboration	Reference	Material
1952	1.0×10^{-1}	Cosmic Ray	Lagarrigue and Peyrou [1952]	Sn, Sb
1955	5.0×10^{-4}	Nevis	Steinberger and Wolfe [1955]	Cu
1961	4.0×10^{-6}	LBL	Sard <i>et al.</i> [1961]	Cu
1961	5.9×10^{-6}	CERN	Conversi et al. [1961]	Cu
1962	2.2×10^{-7}	CERN	Conforto et al. [1962]	Cu
1964	2.2×10^{-7}	Liverpool	Bartley et al. [1964]	Cu
1972	1.6×10^{-8}	SREL	Bryman <i>et al.</i> [1972]	Cu
1977	4.0×10^{-10}	SIN	Badertscher <i>et al.</i> [1977]	S
1982	7.0×10^{-11}	SIN	Badertscher <i>et al.</i> [1982]	S
1988	4.6×10^{-12}	TRIUMF	Ahmad <i>et al.</i> [1988]	Ti
1993	4.3×10^{-12}	SINDRUM II	Dohmen $et al.$ [1993]	Ti
1996	4.6×10^{-11}	SINDRUM II	Honecker et al. [1996]	Pb
2006	7.0×10^{-13}	SINDRUM II	Bertl <i>et al.</i> [2006]	Au

Summery of Current Limits on $\mu^- \rightarrow e^-$ Conversion

	Z	spin	CR upper limit
sulfur	16	0	7 x 10 ⁻¹¹
titanium	22	0, 5/2, 7/2	4.3 x 10 ⁻¹²
copper	29	3/2	1.6 x 10 ⁻⁸
gold	79	5/2	7 x 10 ⁻¹³
lead	82	0, 1/2	4.6 x 10 ⁻¹¹

Current Limits on $\mu^- \rightarrow e^-$ Conversion (2004)

SINDRUM-II (PSI) 1m A exit beam solenoid F inner drift chamber B gold target G outer drift chamber H superconducting coil C vacuum wall D scintillator hodoscope I helium bath (J) E Cerenkov hodoscope J magnet yoke G (\mathbf{A}) configuration 2000 SINDRUM II

$$B(\mu^{-} + Au \to e^{-} + Au) < 7 \times 10^{-13}$$

one e⁻ and one e⁺ background at 98 MeV/c ?
e⁺ peak at 90 MeV/c ?



Improvement of $\mu^- \rightarrow e^-$ Conversion Sensitivity



$B(\mu N \rightarrow eN) \leq 10^{-16}$

with a factor of 10,000 improvement

Improvements for Signal Sensitivity



 $10^{11} \mu$ /s for 50 kW proton beam power or 10^{18} muons in total

MuSIC at RCNP, Osaka University

MuSIC (Muon Science Intense Channel) since 2011



PCS : 3.5T solenoid field and graphite target

MuSIC: 10⁵ muons/sec/W PSI: 3x10² muons/sec/W



(RCNP proton cyclotron 400 W) (PSI proton cyclotron 1.2 MW)

S. Cook et al., Phys. Rev. Accel. Beams, vol. 20, no. 3, p. 030101, 2017.

New Experiments of $\mu^- \rightarrow e^-$ Conversion




Mu2e at Fermilab



6x10¹⁰ muons/s from 8 kW, 8 GeV proton beam



From MELC to MECO (1992 - 2005)

MELC

MEC



Proposal (1992) at Moscow Meson Factory

R. M. Dzhilkibaev and V. M. Lobashev, Sov. J. Nucl. Phys. 49, 384 (1989)



BNL E940 (1997) one of the RSVP (rare symmetry violating processes) with KOPIO

terminated in 2005

$\mu^- \rightarrow e^-$ Conversion Phenomenology



CLFV in EFT : RGE

RGE mixes different operators.



A. Crivellin, S. Davidson, at al., JHEP 117 (2017) 5 S. Davidson, Eur. Phys. J. C76 (2016) 370

CLFV in EFT : Dipole vs. Contact Interactions



S. Davidson and B. Echenard, Eur. Phys. J. C 82 (2022) 9, 836

43



Rate of $\mu^- \rightarrow e^-$ Conversion



nucleon form factors overlap integral (leptons and nucleus)

45

"One of the major topics in this workshop"

Model Discrimination with Different Muon Targets

5% for light targets, 20% for heavy targets



R. Kitano, M. Koike and Y. Okada, Phys.Rev. D66 (2002) 096002; D76 (2007) 059902V. Cirigliano, R. Kitano, Y. Okada, and P. Tuzon, Phys. Rev. D80 (2009) 013002

Model Discrimination with Different Muon Targets (2)

Scalar (n,p)+Vector (n,p)(+Dipole)





choose a target to maximize the constraints on the operators, represented by a large misalignment angle.

S. Davidson, YK, M. Yamanaka, Phys. Lett. B790 (2019) 380-388 J. Heeck, R. Szafron, and Y. Uesaka , Nucl.Phys.B 980 (2022) 115833



Inelastic $\mu^- \rightarrow e^-$ Conversion

"Inelastic" $\mu^- \rightarrow e^-$ Conversion $\mu^- N(A, Z) \rightarrow e^- N(A, Z)^*$

Conversion to excited states.incoherent processes

Event Signature for the transition to the excited state : a single mono-energetic electron



$$E_{\mu e}^{\ast} = E_{\mu e}^{\ast}(gs) - E_{es}^{\ast}$$

when the energy levels of excited states are well separated.

Inelastic $\mu^- \rightarrow e^-$ Conversion

 $\mu^- \rightarrow e^-$ conversion to the excited states

aluminum



from COMET Phase-I TDR: (notes, DIO at 10-15)



Inelastic $\mu^- \rightarrow e^-$ Conversion (Aluminium)



Mu2e spectrum response

default signal region (103.6-105 MeV/c) DIO (green)

The peak of elastic conversion can be either coherent or incoherent?

W. C. Hax

Rule, Physical R

, 025501 (2025)



Experimental Observation of Inelastic $\mu^- \rightarrow e^-$ Conversion

Experimental considerations for potential observation?

(1) Suppression of low energy tail in the spectrum by GBDT.

(2) Suppression of DIO electrons

Use the light target with low Z for the small "B" coefficient (p.24).





Polarized $\mu^- \rightarrow e^-$ Conversion

 $\mu^- \rightarrow e^-$ conversion with spin polarized muons could be used to determine the chirality of the outgoing electron.

Creating a highly spin-polarized muonic atom for CLFV measurements could be extremely challenging!

- High-intensity muon sources, such as COMET/Mu2e
- Atomic capture
- Hyperfine interaction, when a nucleus has a spin

Re-polarization of muonic atom

$\mu^- \rightarrow e^+$ Conversion



$\mu^- \rightarrow e^+$ Conversion in a Muonic Atom

$\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z - 2)$

- Lepton number violation (LNV) and charged lepton flavour violation (CLFV)
- Long range interaction
 - Exchange of light Majorana neutrino is small (<10⁻⁴⁰)

•
$$< m_{\mu e} > = |\sum U_{\mu i} U_{ei} m_{\nu_i}|$$

- Short range interaction
 - TeV LNV physics

$\mu^- \rightarrow e^+$ Conversion in a Muonic Atom

$\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z-2)$

The final state is either ground state or excited states. Event Signature for the transition to the ground state :

$$E_{\mu e^+} = m_{\mu} - B_{\mu} - E_{rec} + M(A, Z) - M(A, Z - 2)$$



Incoherent

Background for $\mu^- \rightarrow e^+$ Conversion : RMC

$$\mu^- + N(A, Z) \to N(A, Z - 1) + \nu + \gamma; \quad \gamma \to e^+ e^-$$

Using an appropriate target

Maximum photon energy (endpoint) :

$$E_{\gamma}^{\max} = m_{\mu} - B_{\mu} - E_{\text{rec}} + M(A, Z) - M(A, Z - 1)$$

Mass Requirement of Muon target N(A, Z):

$$\begin{split} E_{\mu e^+} > E_{\gamma}^{\max} & \Rightarrow \quad M(A, Z-2) < M(A, Z-1) \\ E_{\mu e} > E_{\gamma}^{\max} & \Rightarrow \quad M(A, Z) < M(A, Z-1) \end{split}$$

ex. ³²S, ⁴⁰Ca, ⁴⁸Ti, ⁵⁰Cr, ⁵⁴Fe, ⁵⁸Ni, ⁶⁴Zn, ⁷⁰Ge (for light and medium nuclei)



$\mu^- \rightarrow e^+$ Conversion : Current Limits (1998)



Current limits

 $\mu^{-} + \text{Ti} \rightarrow e^{+} + \text{Ca}(\text{gs}) \le 1.7 \times 10^{-12}$ $\mu^{-} + \text{Ti} \rightarrow e^{+} + \text{Ca}(\text{ex}) \le 3.6 \times 10^{-11}$

(a) positron spectrum

- (b) prompt and delayed timing
- (c) cosmic-ray background

RMC endpoint is 93 MeV/c, instead of maximum off 99 MeV/c (maybe ⁴⁸Sc(0+;6.7MeV) excited ?)

The sensitivity of COMET/Mu2e should be comparable to $\mu^- \rightarrow e^-$ Conversion.

J. Kaulard et al. (SINDRUM-II), Phys. Lett. B422 (1998) 334.

$\mu^- \rightarrow e^+$ Conversion in Gold ????

e^+ peak at 90 MeV/c ?

- The signal energy of $\mu^- + Au \rightarrow e^+ + Ir$ is 91.96 MeV.
- The energy shift (from the material in front) can be 0.6 MeV.
- The expected peak energy is 91.36 MeV.
- It is about 1 MeV higher than the observed peak.
- $k_{\rm max}$ for RMC in the Primakoff model is 88 MeV.

arXiv:2009.00214v1



$\mu^- \rightarrow e^+$ Conversion : History

•

.

Process	90%-C.L. upper limit	Place	Year	Reference	
μ^- + Cu $\rightarrow e^+$ + Co	2.6×10^{-8}	SREL	1972	Bryman <i>et al.</i> (1972)	
μ^- + S $\rightarrow e^+$ + Si	9×10^{-10}	SIN	1982	Badertsher <i>et al.</i> (1982)	
μ^- + Ti $\rightarrow e^+$ + Ca(gs)	9×10^{-12}	TRIUMF	1988	Ahmad <i>et al.</i> (1988)	
μ^- + Ti $\rightarrow e^+$ + Ca(ex)	1.7×10^{-10}	TRIUMF	1988	Ahmad <i>et al.</i> (1988)	
μ^- + Ti $\rightarrow e^+$ + Ca(gs)	4.3×10^{-12}	PSI	1993	Dohmen <i>et al.</i> (1993)	
μ^- + Ti $\rightarrow e^+$ + Ca(ex)	8.9×10^{-11}	PSI	1993	Dohmen et al. (1993)	
μ^- + Ti $\rightarrow e^+$ + Ca(gs)	1.7×10^{-12}	PSI	1998	Kaulard <i>et al.</i> (1998)	
$\mu^- + \mathrm{Ti} \rightarrow e^+ + \mathrm{Ca}(\mathrm{ex})$	3.6×10^{-11}	PSI	1998	Kaulard <i>et al.</i> (1998)	

$\mu^- \rightarrow e^-$ conversion						$\mu^- \rightarrow e^+$ experiments were
1972	1.6×10^{-8}	SREL	Bryman et al. [1972]	Cu	$\left(1\right)$	made at the same time as
1977	4.0×10^{-10}	SIN	Badertscher <i>et al.</i> [1977]	S	\bigcirc	$\mu^- \rightarrow e^-$ experiments.
1982	7.0×10^{-11}	SIN	Badertscher <i>et al.</i> [1982]	S	$\begin{pmatrix} 2 \\ \end{pmatrix}$	
1988	4.6×10^{-12}	TRIUMF	Ahmad <i>et al.</i> [1988]	Ti	(3)	
1993	4.3×10^{-12}	SINDRUM II	Dohmen <i>et al.</i> [1993]	Ti	$\left(4\right)$	Where is (5) for $\mu^- \rightarrow e^-$?
1996	4.6×10^{-11}	SINDRUM II	Honecker et al. [1996]	Pb		
2006	7.0×10^{-13}	SINDRUM II	Bertl <i>et al.</i> [2006]	Au		61

SINDRUM II $\mu^- \rightarrow e^-$ Conversion on Ti (1998)

Wintz, P., 1998, in Proceedings of the First International Symposium on Lepton and Baryon Number Violation, edited by H.V. Klapdor-Kleingrothaus and I.V. Krivosheina (Institute of Physics, Bristol/Philadelphia), p. 534.

$$CR(\mu^{-} + Ti \rightarrow e^{-} + Ti) < 6.1 \times 10^{-13}$$

Lepton-Baryon'98, Trento, ECT*, Italy, April 20-25,1998

Future : PRISM/PRIME



$B(\mu N \rightarrow eN) \leq 10^{-18}$

with a factor of 1,000,000 improvement

Muon Storage Ring: Merit (1)

Allowing the use of high-Z target material

Flight length of 17 m for COMET/Mu2



Flight length of 200 m

Pion survival rate ~5x10-4

Pion survival rate $< 10^{-39}$

The measurement can start from about the beam arrival time, since no pions remain,





diameter 13 m, 5 turns

Muon Storage Ring : Merit (2)

2 Allowing the use of thinner targets

Phase rotation (synchrotron oscillation)

By accelerating slow muons and decelerating fast muons, a beam energy spread can be narrowed.





smaller number of RF's

Muon Storage Ring

PRISM=Phase Rotated Intense Slow Muon source

PRISM/PRIME (2003)





PRISM FFA R&D at Osaka University (2003 - 2007)



Fixed Field Alternating Gradient Synchrotron (FFA)

- FFA is suitable for acceptation low-energy muons
 - large beam acceptance
 - fast beam acceleration
 - synchrotron oscillation



FFA R&D at Osaka

• Scaling FFA with DFD triplet magnets $B(r) \propto \left(\frac{r}{-}\right)^5$



Advanced Muon Facility at Fermilab in the US (2022)

One Concept for $\mu^- N \rightarrow e^- N$

 Spiral Detector Solenoid greatly reduces rate seen by detector, opens up new detector designs (from PRISM)



Future advanced muon facility at Fermilab discussed in Snowmass 2022 workshop by the FNAL people.

Snowmass RPF5

PRISM FFA Phase Rotation at Osaka University (2003 - 2007)





Demonstration of phase rotation has been made.

Global Timeline of Muon CLFV



modified from the muon CLFV white paper for the 2020 update of European Strategy of Particle Physics

 $\times 100,000$



Summary


Summary

The history, experiments, and phenomenology of $\mu \rightarrow e$ conversion are presented, along with the planned experiments and long-term improvements.

Although the $\mu \rightarrow e$ conversion is an indirect search for BSM, it would be crucial to develop the necessary tools to understand the structure of BSM if it were discovered.



Backup

