Guiding Center Dynamics

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Charge in magnetic field – guiding center approximation

- Matter in magnetic fields many important contexts and applications.
- Scale separation: fast gyro motion, slower guiding center motion
- Effective theory of this slower motion is a very useful tool
- Simple case B = const and $E \parallel B$: guiding center moves along a magnetic field line — 1+1d motion.
- Sut what if E
 mid B and/or $B \neq \text{const}$?





Drift



Guiding-center approximation

In (special) relativistic context, averaging over fast gyromotion, Vandervoort '60 finds a rather long and non-covariant formula:

$$\begin{split} \dot{\mathbf{R}}_{\perp} &= \frac{\mathbf{\hat{e}}_{1}}{B(1 - E_{\perp}^{2}/B^{2})} \times \left\{ -\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right) cE \right. \\ &+ \frac{M_{r}c}{\gamma e} \nabla \left[B\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right)^{1/2} \right] + \frac{m_{0}c\gamma}{e} \left(v_{\parallel} \frac{d\mathbf{\hat{e}}_{1}}{dt} + \frac{du_{E}}{dt} \right) \right. \\ &+ \frac{v_{\parallel}E_{\parallel}}{c} u_{E} + \frac{M_{r}}{\gamma e} \frac{u_{E}}{c} \frac{\partial}{\partial t} \left[B\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right)^{1/2} \right] \right\} + \mathcal{O}(\epsilon^{2}) \end{split}$$

 $\epsilon=a/L$ (orbit radius over characteristic length of B variation).

We shall begin by rederiving this formula using simpler and manifestly Lorentz (and general coordinate) covariant formalism.

$E \parallel B$ frame

- The motion is simplest in a (family of) frame(s) where $E \parallel B$. In such a frame we define $B_* \equiv |B|$ and $E_* \equiv E \cdot B/|B|$. B_* and E_* are invariants of the EM field.
- Cyclotron radius $a = p_{\perp}/B_*$. Treat both a/L and E_*/B_* as small.
- Description p_{\perp} only $p \parallel B$ remains, i.e., $p \times B = 0$.
 - Effectively 1 + 1d motion. With effective mass

$$\tilde{m}^2 = m^2 + p_\perp^2 \equiv m^2 + JB_*$$

where $J = ap_{\perp} = p_{\perp}^2 / B_*$ – adiabatic invariant.

In quantum case $J = \hbar(2n + 1 + gs_z)$.

Boost *B* and *E* fields separately back to Lab frame: $F_{\mu\nu} = B_{\mu\nu} + E_{\mu\nu}. \text{ Then } \boldsymbol{p} \times \boldsymbol{B} = 0 \text{ becomes } B^{\mu\nu}p_{\nu} = 0.$

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- Consider Lagrangian with additional constraint $B^{\mu\nu}p_{\nu} = 0$:

$$L = -\dot{x} \cdot (p+A) + \frac{\alpha}{2}(p^2 - \tilde{m}^2) + \lambda_{\mu}B^{\mu\nu}p_{\nu}.$$

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$$\dot{x}^{\mu} = \alpha p^{\mu} + \lambda_{\alpha} B^{\alpha\mu} \equiv \alpha p^{\mu} + v_{\rm D}^{\mu},$$
$$\dot{p}_{\mu} = F_{\mu\nu} \dot{x}^{\nu} + \frac{\alpha}{2} J \partial_{\mu} B_{*} - \lambda_{\alpha} p_{\beta} \partial_{\mu} B^{\alpha\beta},$$

and two constraints: $p^2 - \tilde{m}^2 = 0$, $B^{\mu\nu}p_{\nu} = 0$.

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Substitute \dot{x} into eq. for \dot{p} , use $B^{\mu\nu}p_{\nu} = 0$ and obtain:

$$v_{\rm D}^{\mu} = \alpha B_*^{-2} \left((p \cdot \partial) B^{\mu\nu} p_{\nu} + \frac{J}{2} B^{\mu\nu} \partial_{\nu} B_* \right).$$

Manifestly GR covariant drift formula

$$v_{\rm D}^{\mu} = \alpha B_*^{-2} \Big(\underbrace{(p \cdot \nabla) B^{\mu\nu} p_{\nu}}_{B\text{-curvature, gravity/inertia}} + \underbrace{\frac{J}{2} B^{\mu\nu} \nabla_{\nu} B_*}_{\text{gradient drift}} \Big).$$

written in components contains all of this

$$\begin{split} \dot{\mathbf{R}}_{\perp} &= \frac{\mathbf{\hat{e}}_{1}}{B(1 - E_{\perp}^{2}/B^{2})} \times \left\{ -\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right) cE \right. \\ &+ \frac{M_{r}c}{\gamma e} \nabla \left[B\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right)^{1/2} \right] + \frac{m_{0}c\gamma}{e} \left(v_{\parallel} \frac{d\mathbf{\hat{e}}_{1}}{dt} + \frac{du_{B}}{dt} \right) \\ &+ \frac{v_{\parallel}E_{\parallel}}{c} u_{B} + \frac{M_{r}}{\gamma e} \frac{u_{B}}{c} \frac{\partial}{\partial t} \left[B\left(1 - \frac{E_{\perp}^{2}}{B^{2}}\right)^{1/2} \right] \right\} + \mathcal{O}(\epsilon^{2}) \end{split}$$

but also gravity effects.

Mainfestly general coordinate covariant and much shorter.

Guiding center kinetic theory

$$\mathcal{L}[f] \equiv \dot{x} \cdot \frac{\partial f}{\partial x} + \dot{p} \cdot \frac{\partial f}{\partial p} = \mathcal{C}[f],$$

$$\dot{x}^{\mu} = \alpha p^{\mu} + \lambda_{\alpha} B^{\alpha\mu} \equiv \alpha p^{\mu} + v_{\rm D}^{\mu},$$

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The invariant phase volume is Wd^4p , where

$$W = \underbrace{\theta(p_0)\delta(p^2 - \tilde{m}^2)}_{\text{mass shell}} \underbrace{\frac{\delta^2 \left(\Delta_{\nu}^{\mu} p^{\nu}\right)}{p \parallel B}}_{p \parallel B} \underbrace{\frac{B_*}{2\pi^2 \alpha}}_{\text{LL density}} \left(1 + \frac{p_{\mu}\partial_{\nu}B^{\mu\nu}}{B_*^2}\right)$$

 $\Delta^{\mu}_{\nu} \equiv B^{-2}_{*}B^{\mu\alpha}B_{\nu\alpha}$ — projector onto $\perp B$ plane (in $E \parallel B$ frame).

The phase space has only one momentum direction: $p \parallel B$.

B-field line curvature correction is important.

(more) $T^{\mu\nu}$

Ideal hydrodynamics

9 To lowest non-trivial order in a/L and E_*/B_* :

 $\partial_{\mu}\tilde{N}^{\mu} = 0$ — charge conservation,

 $\tilde{\Delta}^{\lambda}_{\nu}\partial_{\mu}T^{\mu}_{\ \lambda} = E_{\nu\lambda}\tilde{N}^{\lambda}$ — 1+1 energy-momentum conservation.

where $\tilde{\Delta}^{\mu}_{\nu} = \delta^{\mu}_{\nu} - \Delta^{\mu}_{\nu}$ projector onto 1+1 $E \parallel B$ space.

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Constitutive equations:

$$\begin{split} \tilde{N}^{\mu} &= n u^{\mu} \,, \\ T^{\mu\nu} &= \underbrace{(\epsilon + P) u^{\mu} u^{\nu} - P \tilde{\Delta}^{\mu\nu}}_{\tilde{T}^{\mu\nu} - 1 + \mathrm{1d \ tensor}} - P_{\perp} \Delta^{\mu\nu} \,, \end{split}$$

and $\Delta^{\mu}_{\nu}(x)u^{\nu}(x) = 0$ — i.e., fluid rest frame is an $E \parallel B$ frame.

• The flow is (locally) 1+1-dimensional (along B).

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• The flow is (locally) 1+1-dimensional (along B).

But there is transverse pressure P_{\perp} (due to gyromotion).

Full EM current

The full (unprojected) divergence of the full T^{μ}_{ν} :

$$\partial_{\mu}T^{\mu}_{\ \nu} = F_{\nu\mu}J^{\mu}, \quad \text{with} \quad J^{\mu} \equiv N^{\mu} + \partial_{\lambda}M^{\lambda\mu}.$$

Transport current $N^{\mu} = \tilde{N}^{\mu} + N^{\mu}_{D}$ contains "longitudinal" (|| u) + "drift"

$$N_{\mathsf{D}}^{\mu} \equiv \int_{p} W v_{\mathsf{D}}^{\mu} f = B_{*}^{-2} (\tilde{T}_{\nu}^{\lambda} \partial_{\lambda} B^{\mu\nu} - M B^{\mu\nu} \partial_{\nu} B_{*}) \,.$$

currents. The "magnetization current" is $\partial_{\lambda}M^{\lambda\mu}$, where

$$M^{\lambda\mu} = MB^{\lambda\mu}/B_*$$

is the magnetization density due to the fast transverse gyro motion.

The drift and the magnetization currents conspire to make the 3+1 energy-momentum conservation equations linearly dependent, reducible to 1+1 (projected) equations.

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Guiding Center Dynamics

Constitutive relations

Can be obtained in three different ways, with the same result:

$$n = \frac{\partial P}{\partial \mu}, \quad \epsilon = T \frac{\partial P}{\partial T} + \mu \frac{\partial P}{\partial \mu} - P, \quad P_{\perp} = P - B_* \frac{\partial P}{\partial B_*}.$$

Directly from kinetic theory in terms of feq:

$$\{n,\epsilon,P,P_{\perp}\} = \frac{B_*}{2\pi} \int \frac{dp_z}{2\pi} f_{\text{eq}}(\beta u \cdot p - \alpha) \bigg\{ 1, p_0, \frac{p_z^2}{p_0}, \underbrace{\frac{JB_*}{2p_0}}_{JB_* = p_{\perp}^2} \bigg\}.$$

Pressure depends on f_{eq} . Relations between P, n, ϵ , P_{\perp} do not.

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Pressure depends on f_{eq} . Relations between P, n, ϵ , P_{\perp} do not.

- By applying the 2nd law of thermodynamics.
- From the partition function $\ln Z = \int d^4x \sqrt{-g} P(\mu, T, B_*)$ in external gauge field and metric backgrounds.

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Beyond kinetic theory

• With u^{μ} defined via $B_{\mu\nu}u^{\nu} = 0$ and $\tilde{\Delta}^{\mu}_{\lambda}T^{\lambda}_{\nu}u^{\nu} = \epsilon u^{\mu}$ (1+1 projected Landau condition), the constitutive relations

$$\begin{split} \tilde{N}^{\mu} &= nu^{\mu} \,, \\ T^{\mu\nu} &= (\epsilon + P)u^{\mu}u^{\nu} - P\tilde{\Delta}^{\mu\nu} - P_{\perp}\Delta^{\mu\nu} \,, \\ &= \frac{\partial P}{\partial \mu} \,, \ \epsilon = T\frac{\partial P}{\partial T} + \mu\frac{\partial P}{\partial \mu} - P \,, \ P_{\perp} = P - B_{*}\frac{\partial P}{\partial B_{*}} \,, \end{split}$$

are the most general covariant relations at zeroth order in derivatives satisfying second law constraints.

- Thus, such hydrodynamic description would also apply to strongly coupled QGP (as found in, e.g., Hattori et al '22).
- The condition $B_{\mu\nu}u^{\nu} = 0$ eliminates non-hydro modes relaxing fast: $\Gamma_{\perp} = \sigma B^2/(\epsilon + P)$ flow \perp to *B* generates ohmic heat.

n

- Action principle with a momentum constraint offers a simple and manifestly covariant derivation of the guiding-center dynamics.
- This approach leads to a number of new results such as the mainfestly covariant expression for v_D, invariant phase space volume W, covariant kinetic theory and hydrodynamics.
- More work is needed to incorporate collisions and dissipation.
- Did not consider feedback of the particles onto the EM fields. But we found the full EM current (longitudinal, drift and magnetization) – a crucial ingredient for MHD simulations.

More

Properties of the $E \parallel B$ frame(s)

Frame velocity u^{μ} is an eigenvector: $(F^{\mu\lambda}F_{\lambda\nu})u^{\nu} = E_*^2 u^{\mu}$.

In such a frame $\boldsymbol{E} \times \boldsymbol{B} = 0$.

E and B are invariant under boosts along E and B.

 $2(B_*^2 - E_*^2) = F_{\mu\nu}F^{\mu\nu} \quad \text{ and } \quad -4E_*B_* = F_{\mu\nu}\tilde{F}^{\mu\nu}.$

Unique decomposition $F_{\mu\nu} = E_{\mu\nu} + B_{\mu\nu}$, where

$$B_{\mu\nu} = \frac{F_{\mu\nu} - \epsilon_* \tilde{F}_{\mu\nu}}{1 + \epsilon_*^2}, \quad E_{\mu\nu} = \epsilon_* \frac{\tilde{F}_{\mu\nu} + \epsilon_* F_{\mu\nu}}{1 + \epsilon_*^2}, \qquad \epsilon_* = E_*/B_*.$$

such that $B_{\mu\nu}u^{\nu} = 0$, $E^{\mu\nu}B_{\lambda\nu} = 0$, $B^{\mu\nu}B_{\mu\nu} = 2B_*^2$, $E^{\mu\nu}E_{\mu\nu} = -2E_*^2$.

For $u = \gamma(1, 0, 0, v)$, $E_{0z} = E_*$ and $B_{xy} = -B_*$.

Invariant phase-space volume

Invariant phase-space volume = flux of the state density current

$$J^M = (\dot{x}^\mu, \dot{p}_\nu)W,$$

in phase space through a given equal-time surface:

$$d^{7}\Gamma = J^{M}d^{7}\Sigma_{M} = (n \cdot \dot{x})Wd^{3}xd^{4}p.$$

Solving conservation equation $\partial_M J^M = 0$ we find

$$W = \frac{B_* + B_*^{-1} p_{\mu} \partial_{\nu} B^{\mu\nu}}{2\pi^2 \alpha} \theta(p_0) \delta(p^2 - \tilde{m}^2) \delta^2(\Delta^{\mu\nu} p_{\nu}).$$

 $\Delta^{\mu}_{\nu} = B^{-2}_{*}B^{\mu\alpha}B_{\nu\alpha}$ – projector onto the plane orthogonal to **B**.

The overall (quantum) coefficient is fixed by Landau-level degeneracy.

back

From kinetic theory to hydrodynamics

Using conservation of charge and energy/momentum in collisions:

$$\int_{p} W\{1, p_{\nu}\} \mathcal{L}[f] = \{\partial_{\mu} N^{\mu}, \partial_{\mu} \tilde{T}^{\mu}_{\ \nu} - F_{\nu}\} = 0.$$

where current, energy-momentum and force density:

$$\{N^{\mu}, \tilde{T}^{\mu}_{\ \nu}, F_{\nu}\} = \int_{p} W\{\dot{x}^{\mu}, \dot{x}^{\mu}p_{\nu}, \dot{p}_{\nu}\}f$$

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Using equations of motion, to leading order:

$$\tilde{T}^{\mu\nu} = \int_p W \alpha \, p^\mu p^\nu f,$$

$$F_{\nu} = F_{\nu\lambda}N^{\lambda} - M\partial_{\nu}B_{*}, \ M \equiv -\int_{p}W\frac{\alpha J}{2}f.$$

Charge transport current

"Transport" current $N^{\mu} = \tilde{N}^{\mu} + N^{\mu}_{D}$ consists of "longitudinal"

$$\tilde{N}^{\mu} \equiv \int_{p} W \alpha p^{\mu} f \,,$$

and "drift" components (higher order in a/L, due to ∂s)

$$N_{\mathsf{D}}^{\mu} \equiv \int_{p} W v_{\mathsf{D}}^{\mu} f = B_{*}^{-2} (\tilde{T}_{\nu}^{\lambda} \partial_{\lambda} B^{\mu\nu} - M B^{\mu\nu} \partial_{\nu} B_{*}) \,.$$

 $N_{\rm D}$ is parametrically negligible in current conservation equation. Hence $\partial \cdot \tilde{N} = 0$ to leading order.

Stress tensor

But $N_{\rm D}$ is *not* negligible in energy-momentum (non-)conservation (because leading term $B_{\nu\lambda}\tilde{N}^{\lambda}=0$). We can rewrite

$$\partial_{\mu}\tilde{T}^{\mu}_{\ \nu} = F_{\nu\lambda}N^{\lambda} - M\partial_{\nu}B_{*}$$

in the final form (1+1 projected equations)

$$\tilde{\Delta}^{\lambda}_{\nu}\partial_{\mu}T^{\mu}_{\lambda} = E_{\nu\lambda}\tilde{N}^{\lambda}$$

if we define full stress tensor as

$$T^{\mu\nu} \equiv \tilde{T}^{\mu\nu} + T^{\mu\nu}_{\perp}, \quad \text{where} \quad T^{\mu\nu}_{\perp} = M B_* \Delta^{\mu}_{\nu} \,.$$

Since

$$T_{\perp}^{\mu\nu} \equiv \int_{p} W \alpha \, \overline{p_{\perp}^{\mu} p_{\perp}^{\nu}} f.$$

this is the contribution to pressure from the fast transverse gyro motion.

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