

Mathematical methods for neutrino cross-section extraction

Based on [arXiv:2401.04065](https://arxiv.org/abs/2401.04065)

$$U = \bigoplus_{b=0} U_b = U_0 \oplus U_1 \oplus \dots = \begin{pmatrix} U_0 & 0 & 0 & \dots \\ 0 & U_1 & 0 & \dots \\ 0 & 0 & \ddots & \dots \\ \vdots & \vdots & \vdots & \ddots \end{pmatrix} \quad \langle \sigma \rangle_\mu = \left\langle \frac{d^n \sigma}{d\mathbf{x}} \right\rangle_\mu \cdot \Delta \mathbf{x}_\mu$$

$$\mathcal{P}_a = \beta_a n_a - D_a + D_a \log \left(\frac{D_a}{\beta_a n_a} \right) + \frac{\beta_a^2 - 1}{2\sigma_a^2}$$

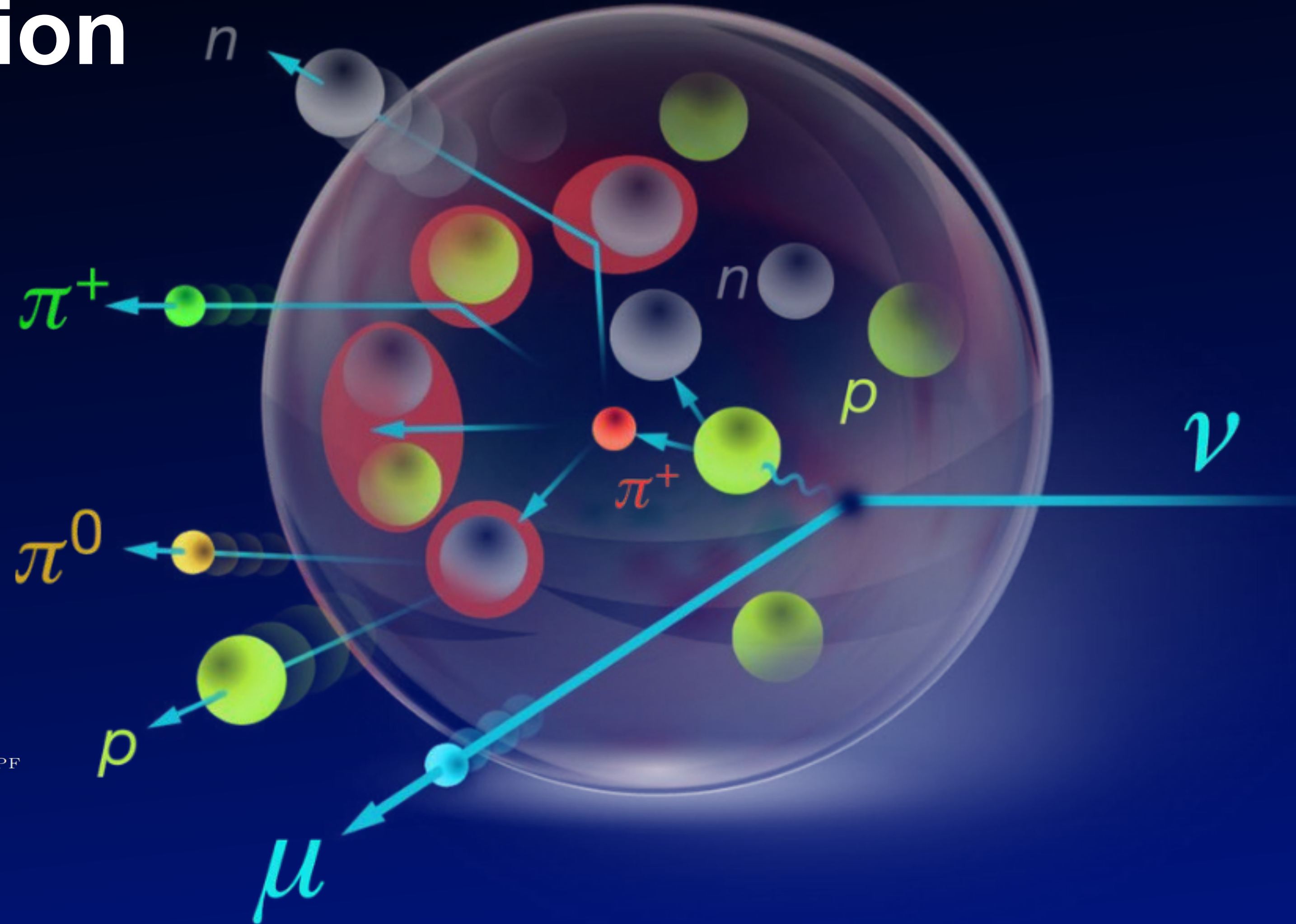
$$\mathbf{e}_{\mu a}^{i+1} = \frac{\partial \hat{\phi}_\mu^{i+1}}{\partial d_a} = U_{\mu a}^i + \frac{\hat{\phi}_\mu^{i+1}}{\hat{\phi}_\mu^i} \mathbf{e}_{\mu a}^i - \sum_{\lambda, b} \epsilon_\lambda \frac{d_b}{\hat{\phi}_\lambda^i} U_{\mu b}^i U_{\lambda b}^i \mathbf{e}_{\lambda a}^i \quad B_a = \sum_b \alpha_b B_{ab}$$

$$\mathbf{B}_S^{\text{constr, CV}} = \mathbf{B}_S^{\text{CV}} + V_{\mathbf{B}_S \mathbf{n}_C} \cdot V_{\tilde{\mathbf{n}}_C \tilde{\mathbf{n}}_C}^{-1} \cdot (\mathbf{D}_C - \mathbf{n}_C^{\text{CV}})$$

$$V_{\mathbf{B}_S \mathbf{B}_S}^{\text{constr}} = V_{\mathbf{B}_S \mathbf{B}_S} - V_{\mathbf{B}_S \mathbf{n}_C} \cdot V_{\tilde{\mathbf{n}}_C \tilde{\mathbf{n}}_C}^{-1} \cdot V_{\mathbf{B}_S \mathbf{n}_C}^T$$

$$V_{\mathbf{m}_S \mathbf{m}_S} = \text{Cov}(\mathbf{m}_S, \mathbf{m}_S) = V_{\phi_S \phi_S} + V_{\phi_S \mathbf{B}_S}^{\text{constr}} + V_{\mathbf{B}_S \phi_S}^{\text{constr}} + V_{\mathbf{B}_S \mathbf{B}_S}^{\text{constr}}$$

$$H_{\mathbf{q}_b} = - \left. \frac{\partial^2 \log(\mathcal{L})}{\partial K_{\mathbf{q}} \partial K_{\mathbf{b}}} \right|_{\mathbf{K}=\mathbf{K}^{\text{PF}}}$$



Steven Gardiner (gardiner@fnal.gov)

Measuring neutrino interactions for next-generation oscillation experiments

ECT*, 25 October 2024



Scope of the talk

- **Trailer for v2** of my recent methods paper
 - Main ideas unchanged since v1
 - Some details refined upon further thought and feedback
- Summary of approaches used in current analyses
 - Quick overview here, Lukas covered many elements already
- "Cookbook" for applying proposed innovations across experiments
- MicroBooNE "**data driven model validation**"
 - Distinct from these techniques = not a change to extraction procedure itself
 - Also likely interesting to this community

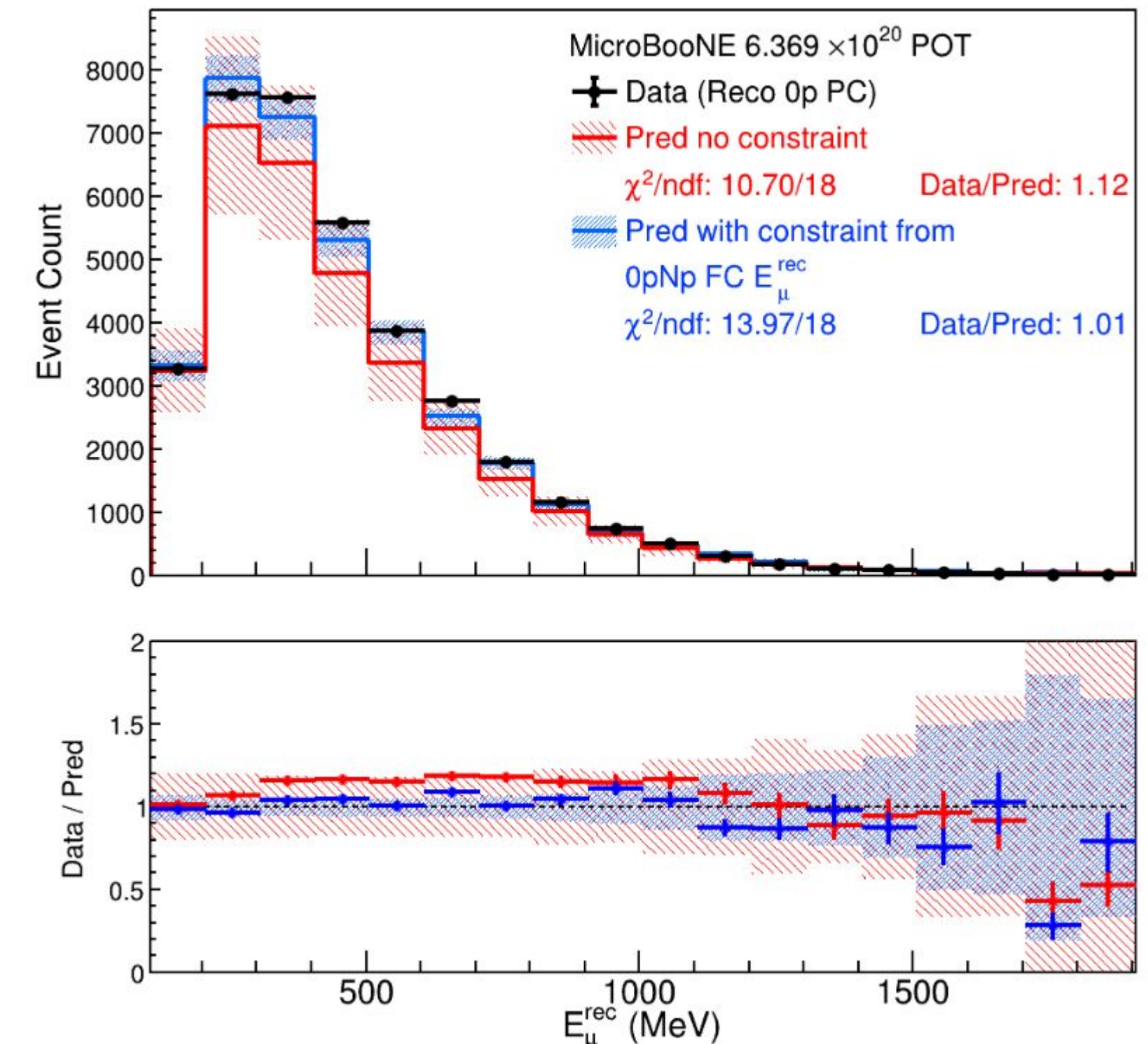
Mathematical methods for neutrino cross-section extraction

Steven Gardiner*

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(Dated: October 25, 2024)

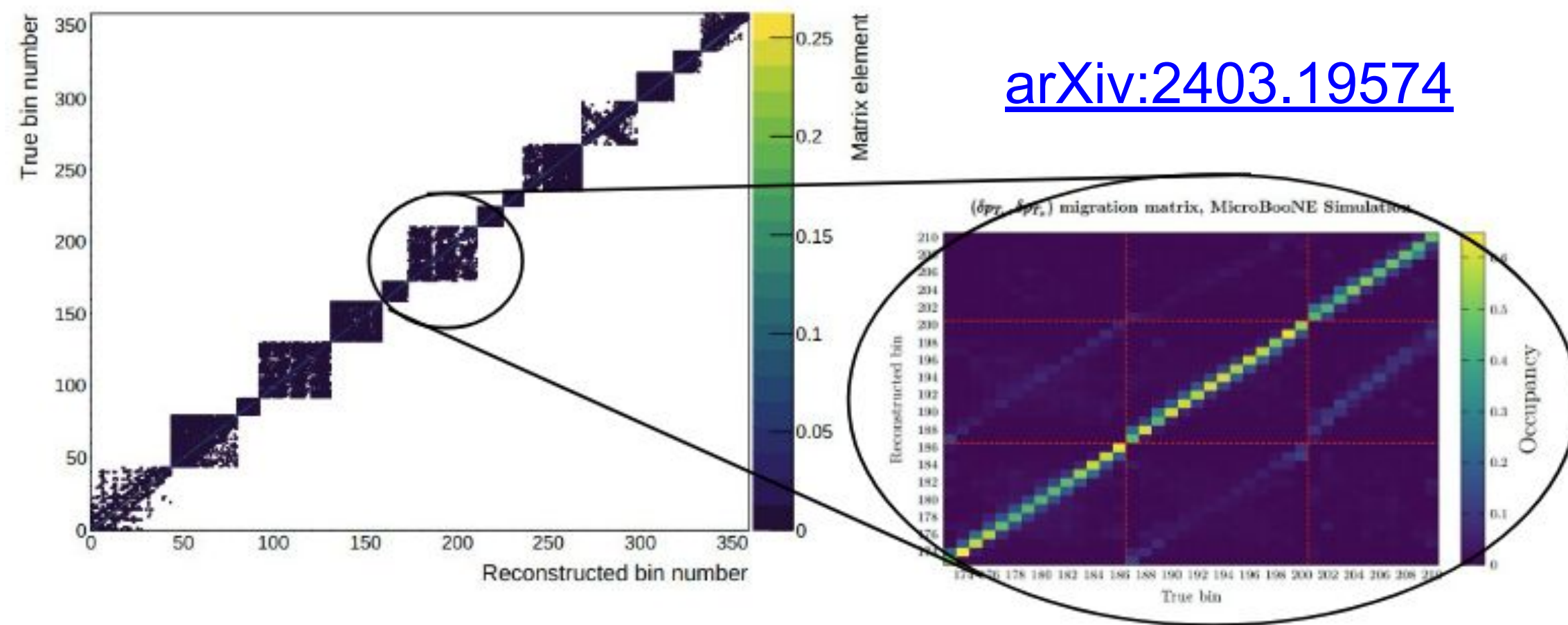
Precise modeling of neutrino-nucleus scattering is becoming increasingly important as accelerator-based oscillation experiments seek definitive answers to open questions about neutrino properties. To guide the needed model refinements, a growing number of experimental collaborations are pursuing a wide-ranging program of neutrino interaction measurements at GeV energies. A key step in most such analyses is cross-section extraction, in which measured event counts are corrected for background contamination and imperfect detector performance to yield cross-section results that are directly comparable to theoretical predictions. In this paper, I review the major approaches to cross-section extraction in the literature using representative examples from the MINERvA, MicroBooNE, and T2K experiments. I then present two mathematical techniques, blockwise unfolding and the conditional covariance background constraint, which overcome some limitations of typical cross-section extraction procedures.



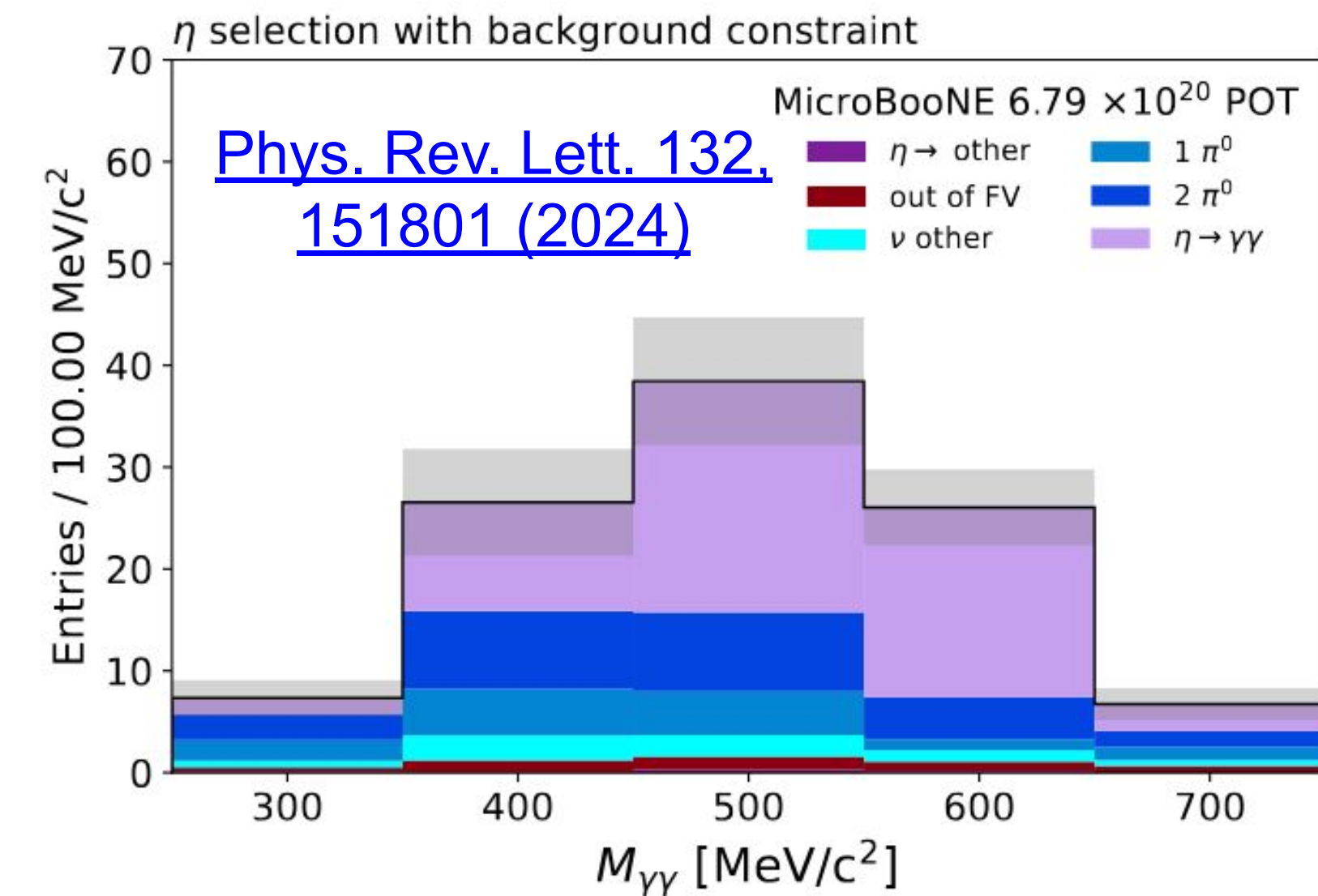
[Phys. Rev. D 110, 013006 \(2024\)](https://arxiv.org/abs/2401.04065), Sec. VI

Overview of new methods

Blockwise unfolding



Conditional Covariance Background Constraint (CCBC)

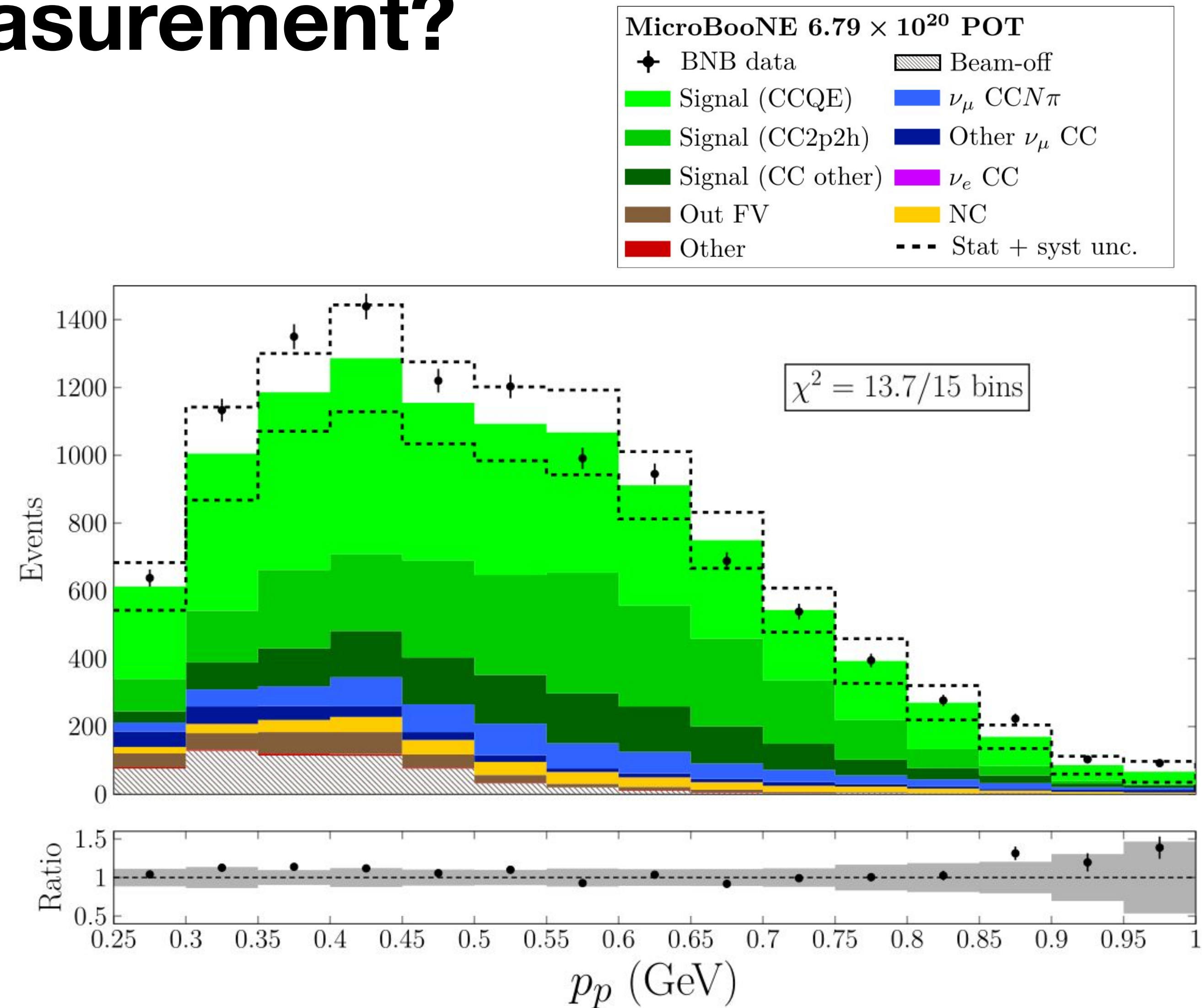


- An argument that we should report correlations much more thoroughly
- Practical advice for how to do that
- Recipes for avoiding some potential pitfalls
- Used in several recent MicroBooNE papers

- Use Gaussian statistics to refine a background prediction
- Achieves compatibility with preferred MicroBooNE extraction procedure
 - No background constraint in *any* differential μ BooNE result so far
- Inspired by μ BooNE η production analysis 3

How do we perform the measurement?

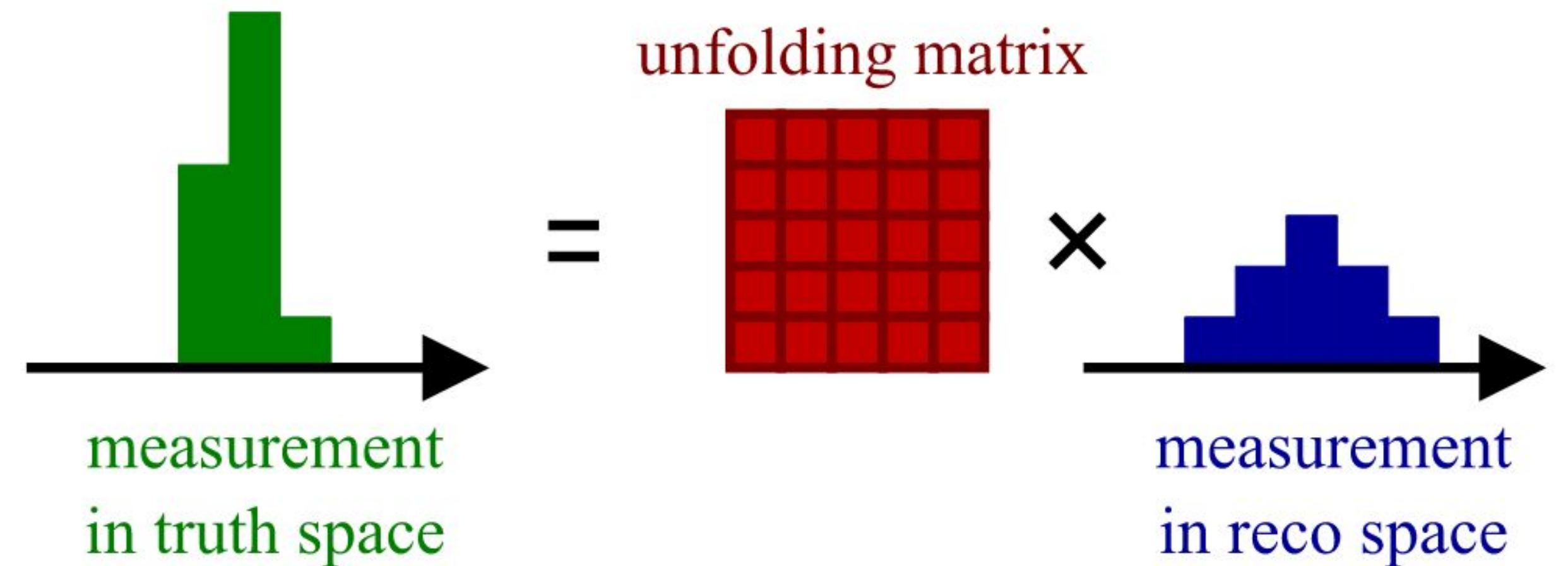
- Counting experiment: bin for variable(s) of interest
- Raw event counts comparable to simulation
 - Only feasible by the experimental collaboration
- **Cross-section extraction**
 - Converts this measurement to a result anyone can use
 - Details vary across experiments
- Many subtleties, care must be taken to avoid bias



How do we perform the measurement?

- **Flux-averaged differential cross section**
 - true bins μ , reco bins a
 - Average value in true bin μ
- **Unfolding matrix U** accounts for inefficiency and bin migrations
- **Unfolded space \approx true space**
 - Systematics must be considered carefully

$$\left\langle \frac{d^n \sigma}{d\mathbf{x}} \right\rangle_{\mu} = \frac{\sum_a U_{\mu a} (D_a - B_a)}{\Phi T \Delta \mathbf{x}_{\mu}}$$



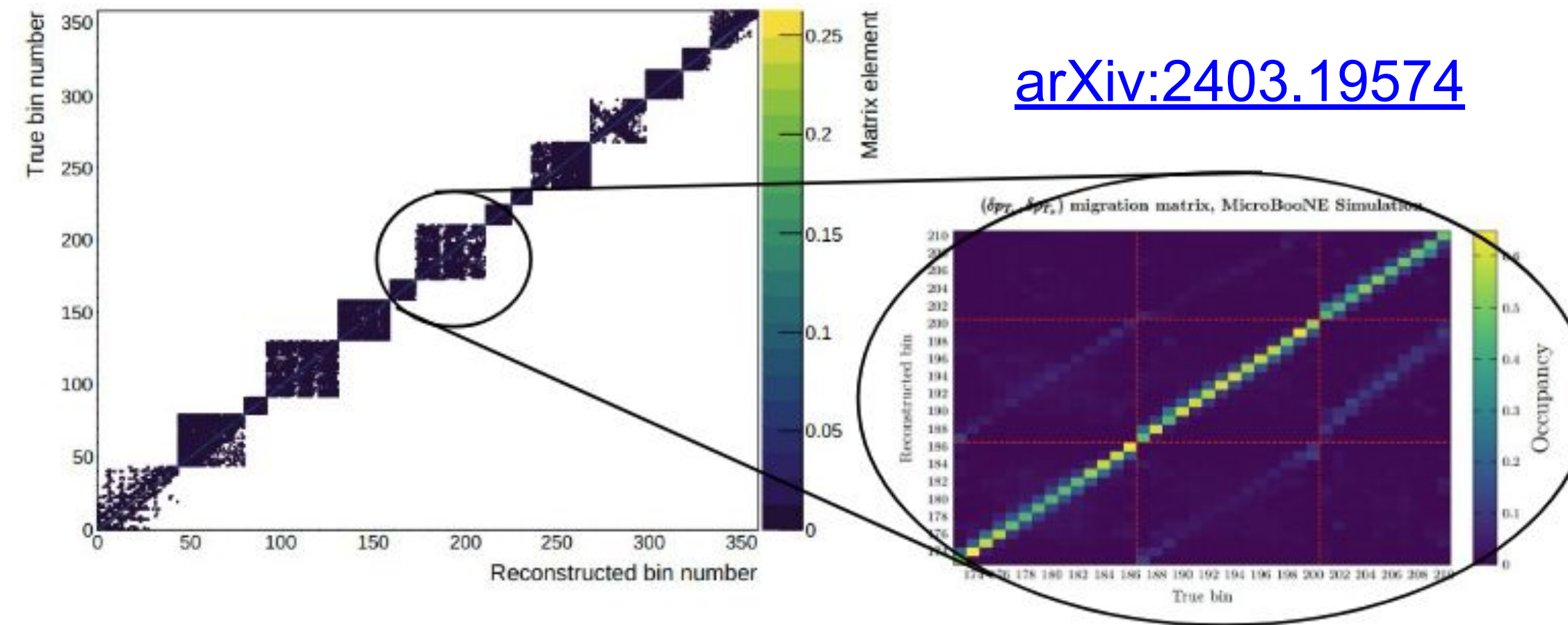
$$U_{\mu a} = \frac{P_{\mu a}}{\epsilon_{\mu}}$$

"Styles" of cross-section extraction

- Superficially, everyone plays the same game, but differently
 - **3 major approaches** at GeV scale, the rest are perturbations
 - Details are often not spelled out, especially for *Phys. Rev. Lett.*
- **MINERvA**
 - [D'Agostini iterative](#) recipe for building unfolding matrix U
 - Uncertainties: repeat extraction, take spread between "universes"
- **MicroBooNE**
 - [Wiener-SVD](#) unfolding
 - Compute total covariance on event counts, propagate through unfolding
- **T2K**
 - Perform likelihood fit to event counts (huge number of parameters)
 - Uncertainties can be treated two ways
 - Repeat the fit across many universes (MINERvA-esque)
 - Vary parameters according to post-fit covariance matrix

Blockwise unfolding

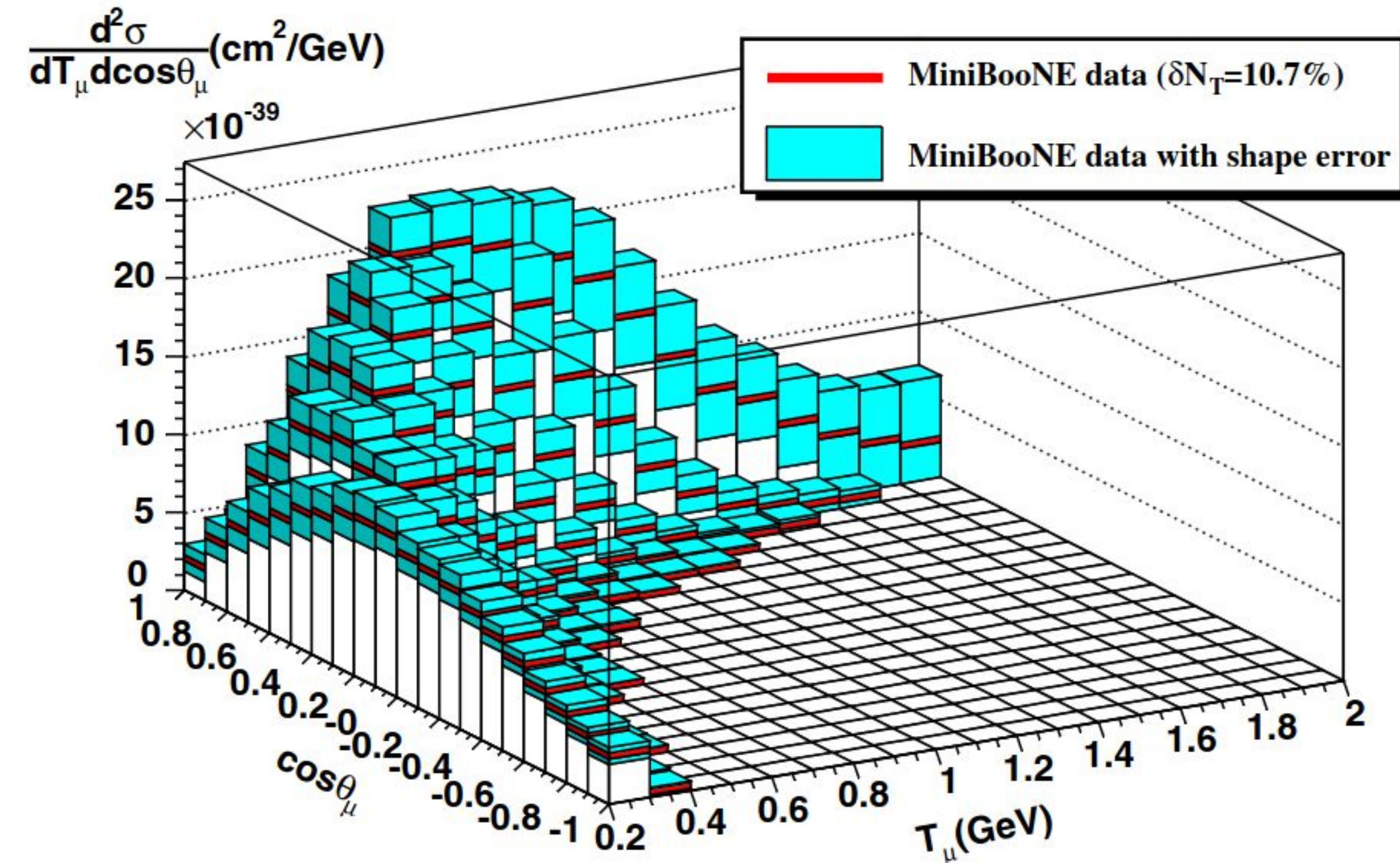
[arXiv:2403.19574](https://arxiv.org/abs/2403.19574)



"Blockwise unfolding": motivation

[Phys. Rev. D 81, 092005 \(2010\)](#)

- **MiniBooNE**: pioneering neutrino experiment at Fermilab
 - Many cross-section analysis practices established
 - Key early measurements
- Several data releases report binwise uncertainties but **not correlations**
 - Large & important
 - Both systematic (e.g., flux) and statistical (unfolding)



2D result for CH target

Problematic for quantitative comparisons (χ^2 , etc.)

Standard practice is now to provide a full covariance matrix

"Blockwise unfolding": motivation

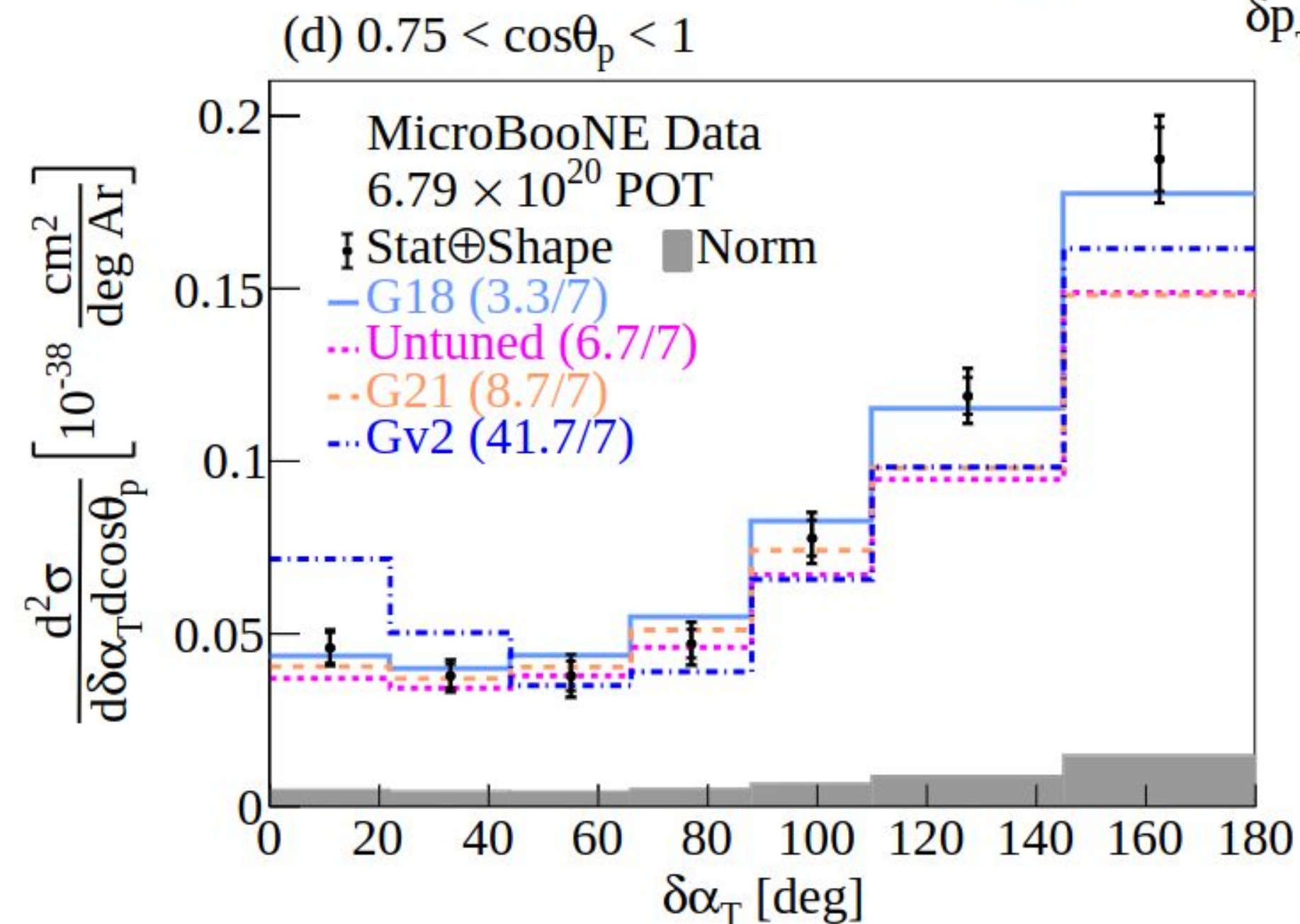
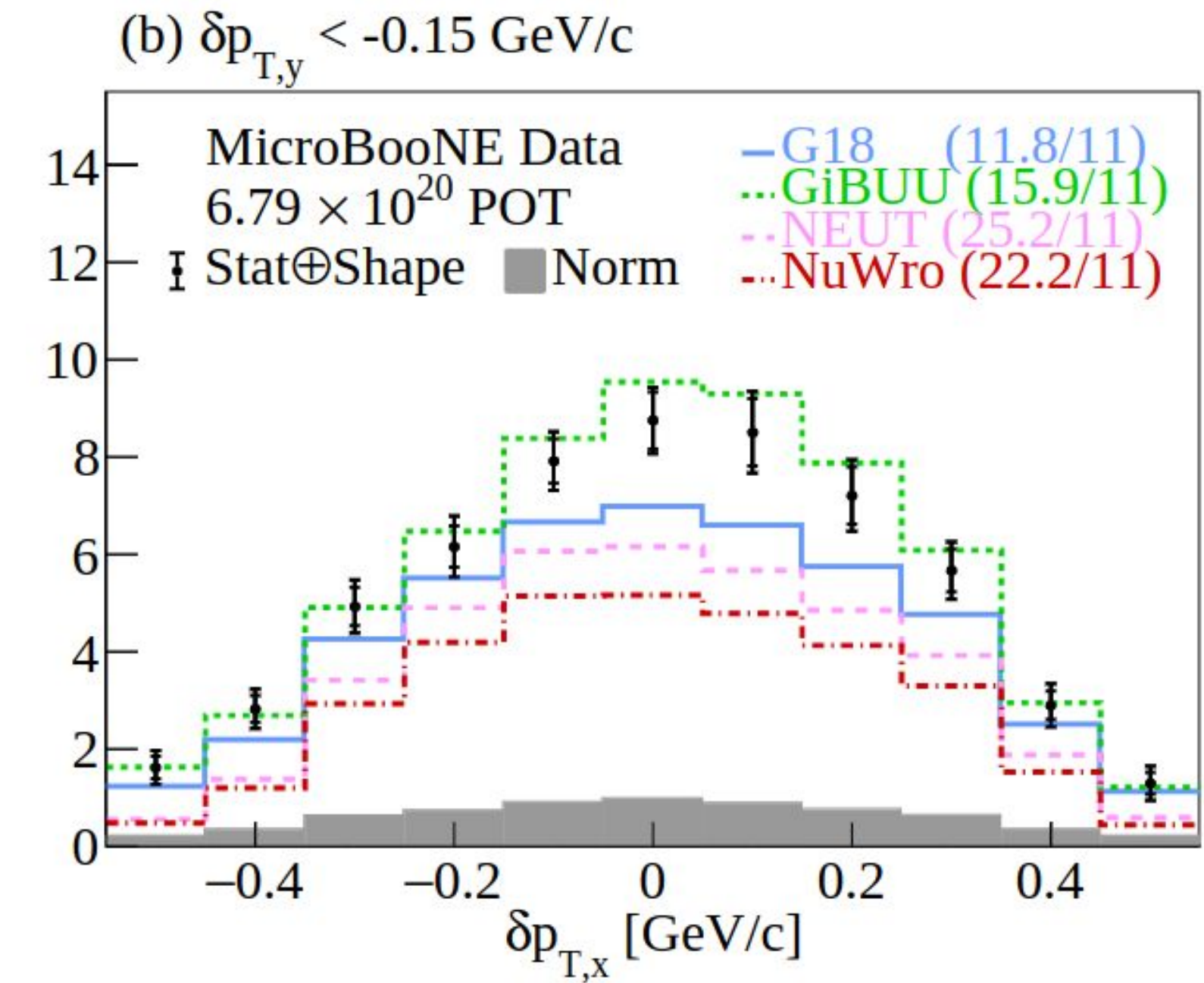
- Experiments often report **multiple kinematic distributions**

- Same analysis or complementary ones

- Correlated uncertainties** between distributions are still not typically reported

- All the same drawbacks as before

ν_μ CC 1p0 π
data from
MicroBooNE

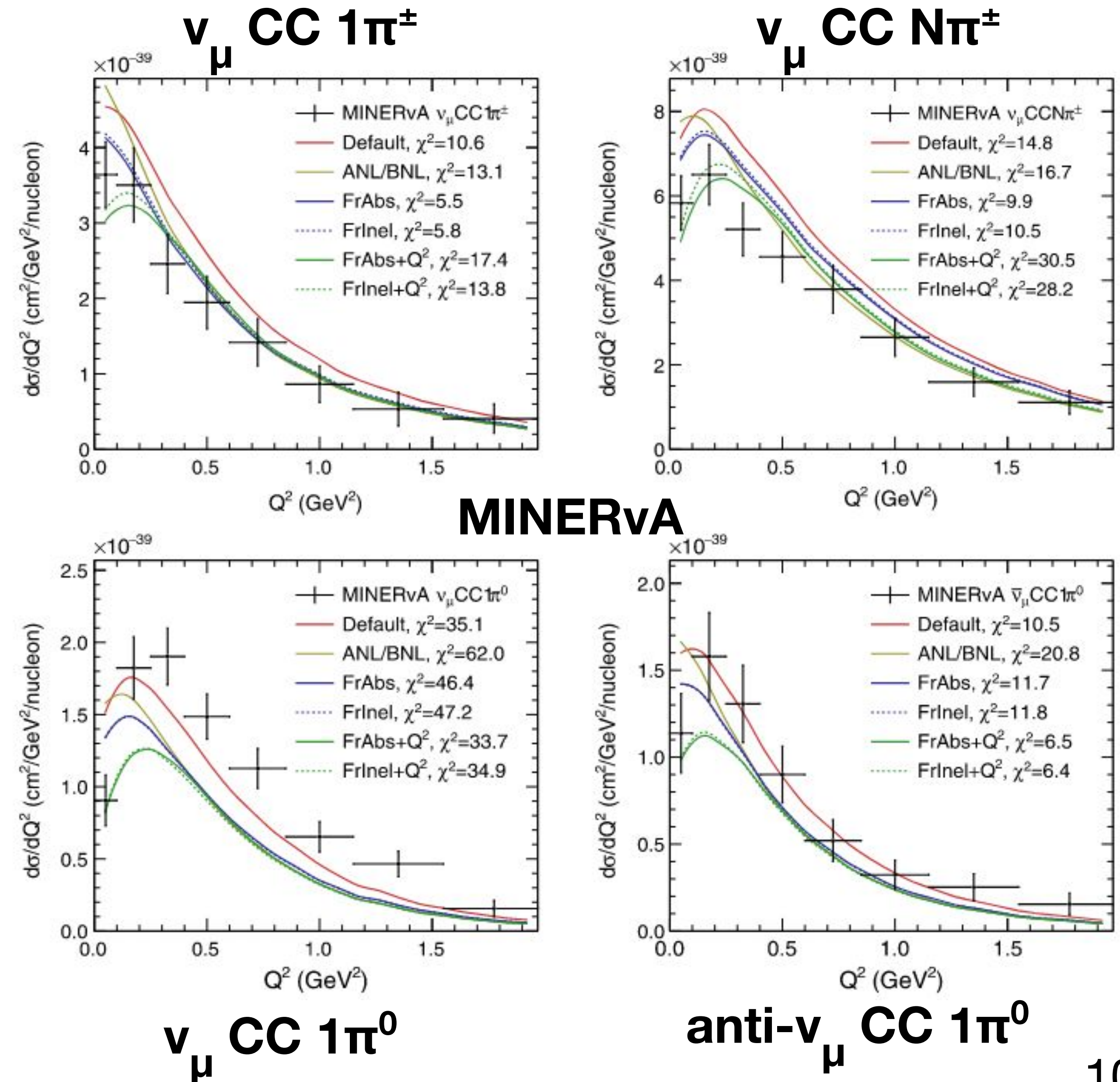


Measurements use same set of ~9000 events

"Blockwise unfolding": motivation

[Phys. Rev D. 100, 072005 \(2019\)](#)

- Experiments often report **multiple kinematic distributions**
 - Same analysis or complementary ones
- **Correlated uncertainties** between distributions are still not typically reported
 - All the same drawbacks as before
- **Limitations** discussed in MINERvA paper tuning GENIE to π production data



"Blockwise unfolding": motivation

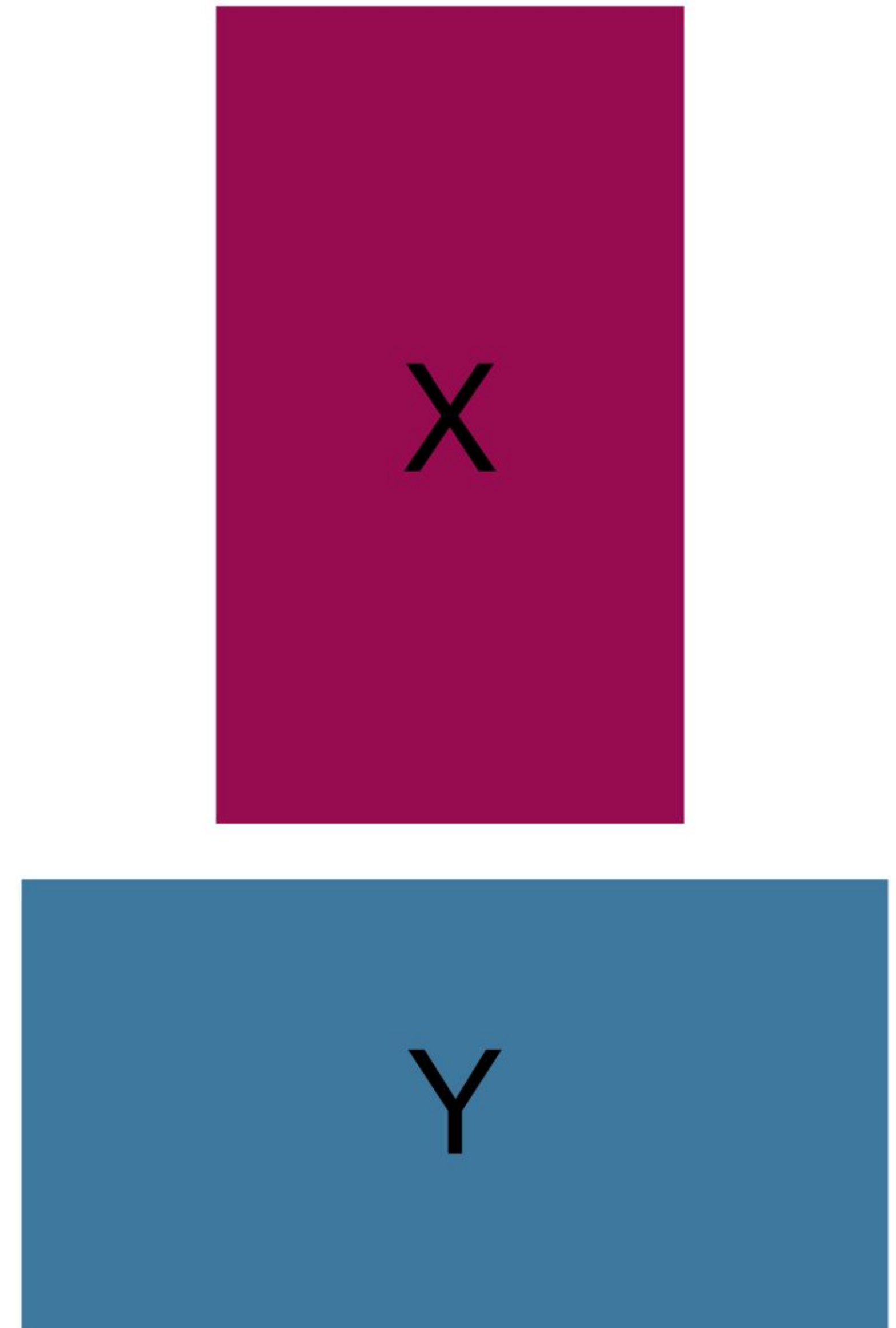
The published cross sections are one dimensional with correlations provided between the bins within each distribution. No correlations are provided between measurements of different final states, or between different one-dimensional projections of the same measurement. These correlations are expected to be large, coming predominantly from flux and detector uncertainties. Additionally, the $\nu_\mu \text{CC}1\pi^\pm$ event sample is a subset ($\sim 64\%$) of the $\nu_\mu \text{CC}N\pi^\pm$ event sample, and including both channels introduces a statistical correlation. Not assessing correlations between the distributions, while a common practice in this field, is a limitation when tuning models to multiple datasets. It introduces a bias in the χ^2 statistic that is difficult to quantify, and requires imposing *ad hoc* uncertainties [4] as the test statistic is not expected to follow a χ^2 distribution for the given degrees of freedom.

[Phys. Rev D. 100, 072005 \(2019\)](#)

- Not trivial to add this information after the fact
- Correlations calculable with **suitable planning ahead**
 - Maximize impact from cross-section analyses
- **Two issues**
 - Event overlaps (statistical covariances)
 - Unfolding treatment
- **Methods paper** ([arXiv:2401.04065](#)) gives recipes for solving these problems

Statistical covariances

- Events belong to multiple bins
⇒ correlated stat uncertainties
- **Easily calculable** if the problem is framed properly
- Arbitrary bins X and Y
 - Event count n_x in bin X follows a Poisson distribution
- Estimator for the mean: n_x
- Estimator for the variance: n_x
- Bin Y is similar. How to get the covariance?

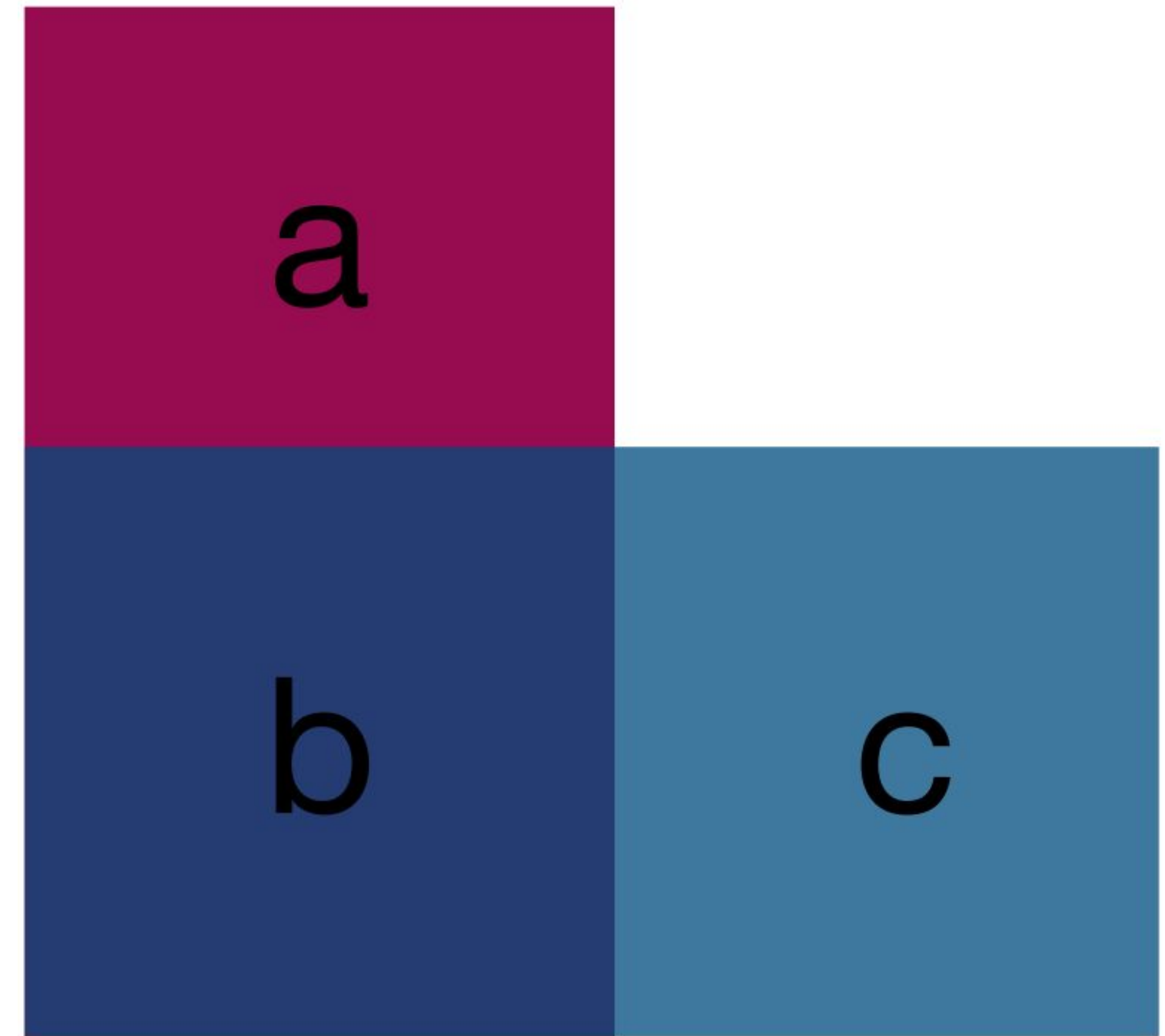


Statistical covariances

- The trick: one may always rebin 2 \rightarrow 3
- Bins a, b, and c are **non-overlapping**
- Independent Poisson distributions

$$\begin{aligned}\text{cov}(X, Y) &= \text{cov}(a + b, b + c) \\ &= \text{cov}(a, b) + \text{cov}(a, c) + \text{cov}(b, b) + \text{cov}(b, c) \\ &= 0 + 0 + \text{var}(b) + 0 \\ &\approx n_b\end{aligned}$$

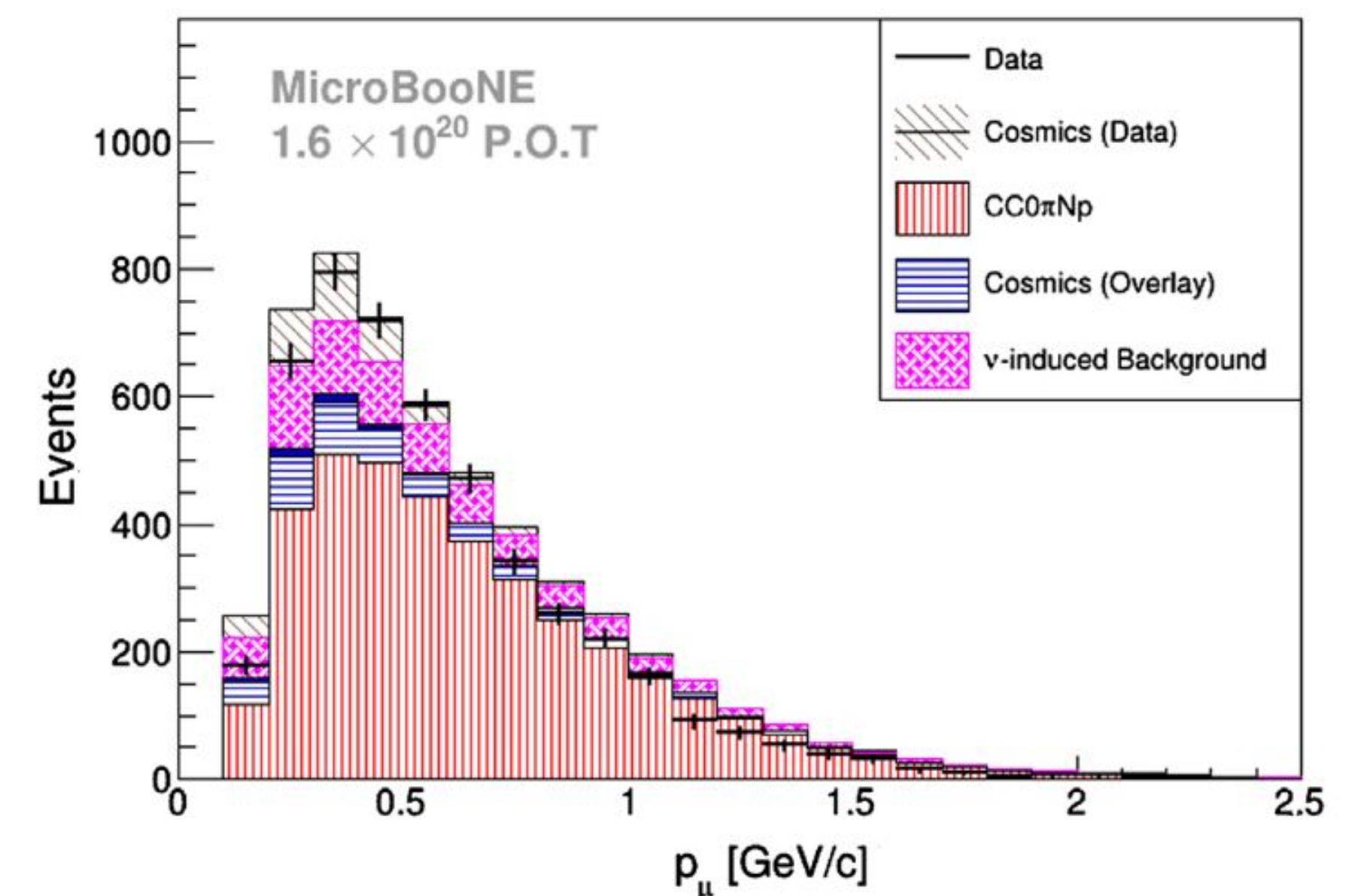
- Estimator for statistical covariance is just the **number of events that bins X and Y have in common**
- Details can change for MINERvA/T2K, but solution is conceptually similar



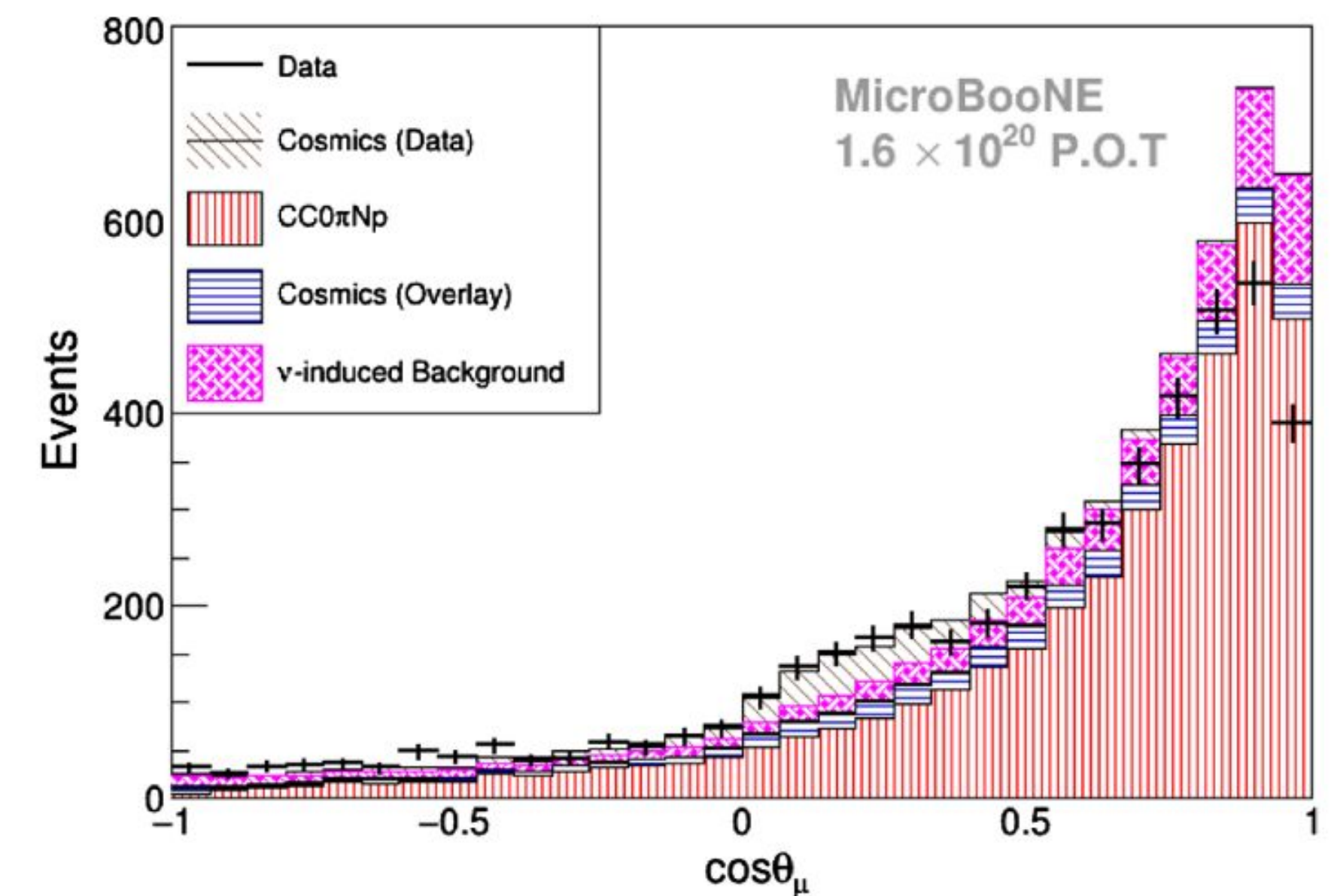
Note that this behaves as expected for $X = Y$ as well as disjoint bins

Unfolding with correlated uncertainties

- Group bins belonging to the same kinematic distribution in a "**block**"
- An event should belong to a maximum of one reco bin and one true bin in each block → avoids double-counting
- Observables can be abstracted away by working in "bin number space"
 - Trivially generalizes to 2D, 3D, etc.
- Example:
 - Bins 0-19 represent $p_\mu \rightarrow$ block #0
 - Bins 20-49 represent $\cos\theta_\mu \rightarrow$ block #1



[Phys. Rev. D 102, 112013 \(2020\)](#)



A "blockwise" unfolding matrix

- Build an unfolding matrix U_b for the b -th block according to one's preferred approach
- Overall unfolding matrix U is block-diagonal
- Results for individual blocks are the **same** as for stand-alone measurements of each
- This organization allows reporting of correlated uncertainties between all bins in all blocks
 - Details depend on extraction style, but fully documented in paper

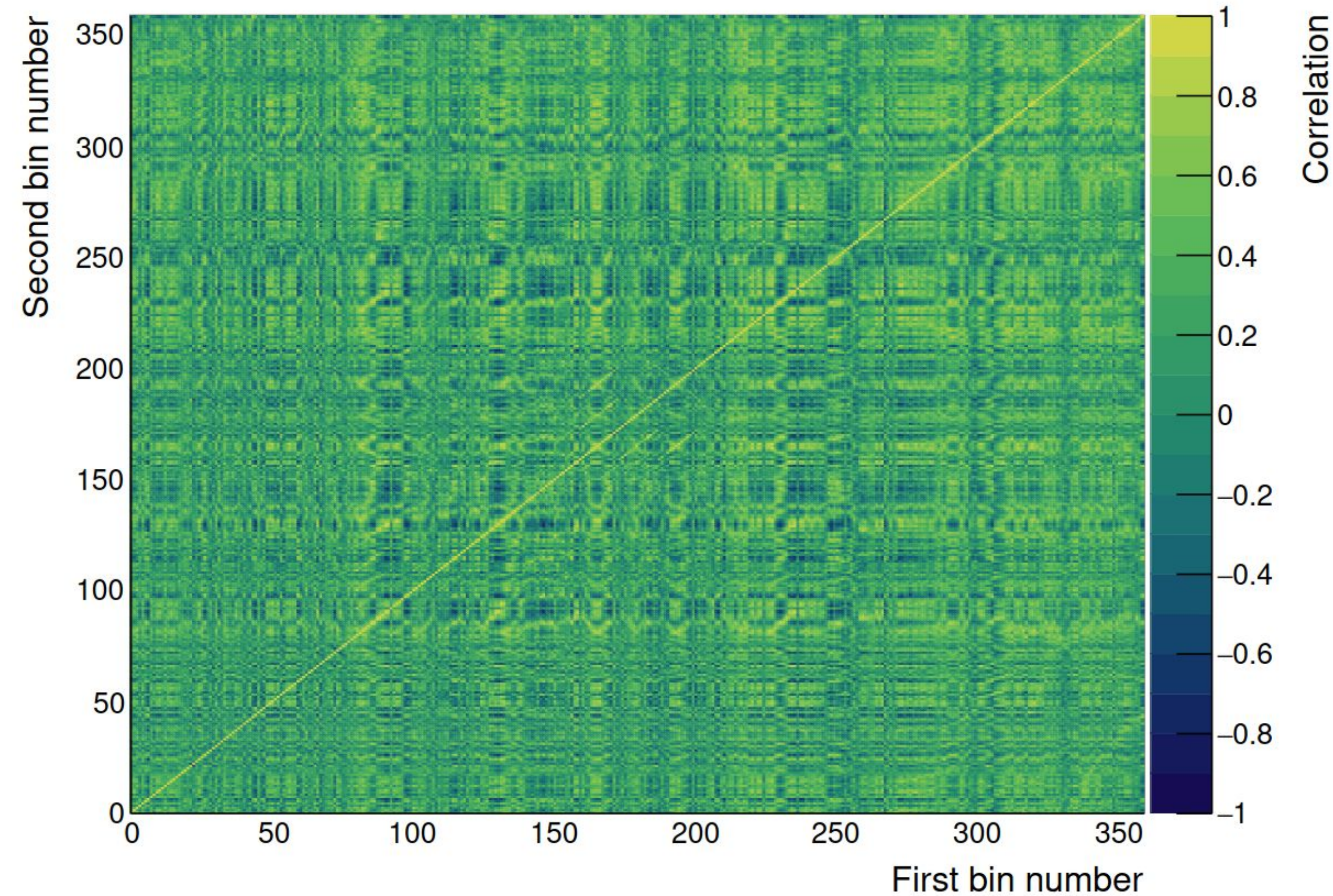
$$U = \bigoplus_{b=0} U_b = U_0 \oplus U_1 \oplus \dots = \begin{pmatrix} U_0 & 0 & 0 & \dots \\ 0 & U_1 & 0 & \dots \\ 0 & 0 & \ddots & \dots \\ \vdots & \vdots & \vdots & \ddots \end{pmatrix}$$

Inter-distribution correlations

MicroBooNE, [arXiv:2403.19574](https://arxiv.org/abs/2403.19574)

- Enable χ^2 comparisons to entire data set!
- **Annoying detail:** differential cross sections vary in their units
 - Can lead to confusion when reporting covariances
- **My recommendation:**
 - Re-express as *total* cross sections per bin
 - Makes inclusion of under/overflow bins easy

Total correlation matrix for measured CC0 π Np cross sections



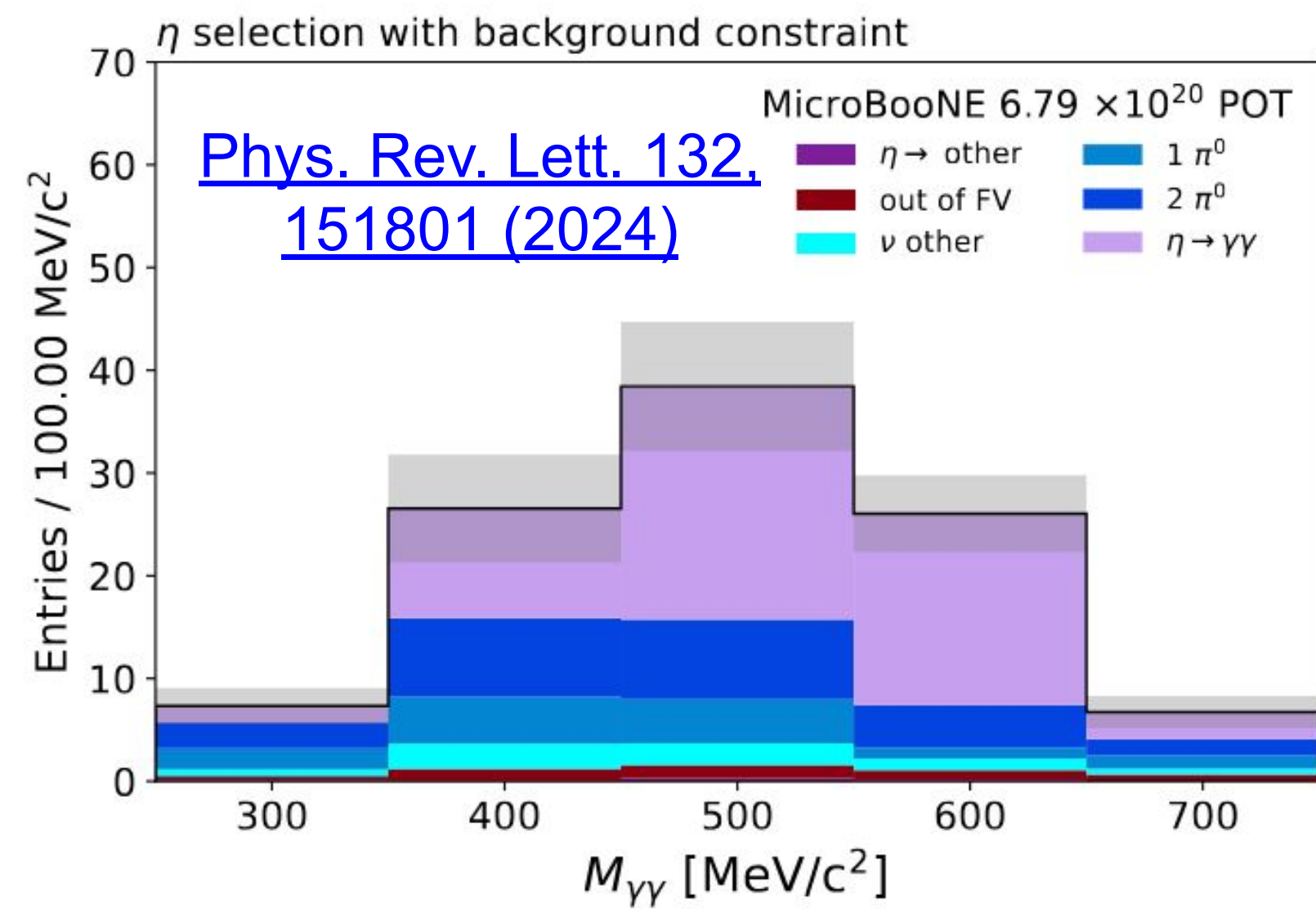
$$\langle \sigma \rangle_\mu = \left\langle \frac{d^n \sigma}{d\mathbf{x}} \right\rangle_\mu \cdot \Delta \mathbf{x}_\mu \quad \langle \sigma \rangle_\mu = \frac{\hat{\phi}_\mu}{\Phi_\mu T_\mu} = \frac{\sum_a U_{\mu a} (D_a - B_a)}{\Phi_\mu T_\mu}$$

$$\text{Cov}(\langle \sigma \rangle_\mu, \langle \sigma \rangle_\lambda) = \frac{\sum_{a,b} \mathbf{E}_{\mu a} \text{Cov}(D_a, D_b) \mathbf{E}_{\lambda b}}{\Phi_\mu \Phi_\lambda T_\mu T_\lambda}$$

Outlook for the blockwise unfolding technique

- Theorists and generator developers can fit to all measured distributions simultaneously
 - Increases discrimination power of the data: can the model describe the correlations as well as each individual block?
- No need for ad hoc estimates of flux-related covariances, etc.
 - All uncertainties come from the experiment itself
- Potential for **inter-analysis covariances** with two ingredients:
 - Bookkeeping for event overlaps (statistical uncertainties)
 - Consistent systematic variations
- **Latest MicroBooNE analyses** report model goodness-of-fit χ^2 over hundreds of bins in this way
 - Other experiments can do this too!

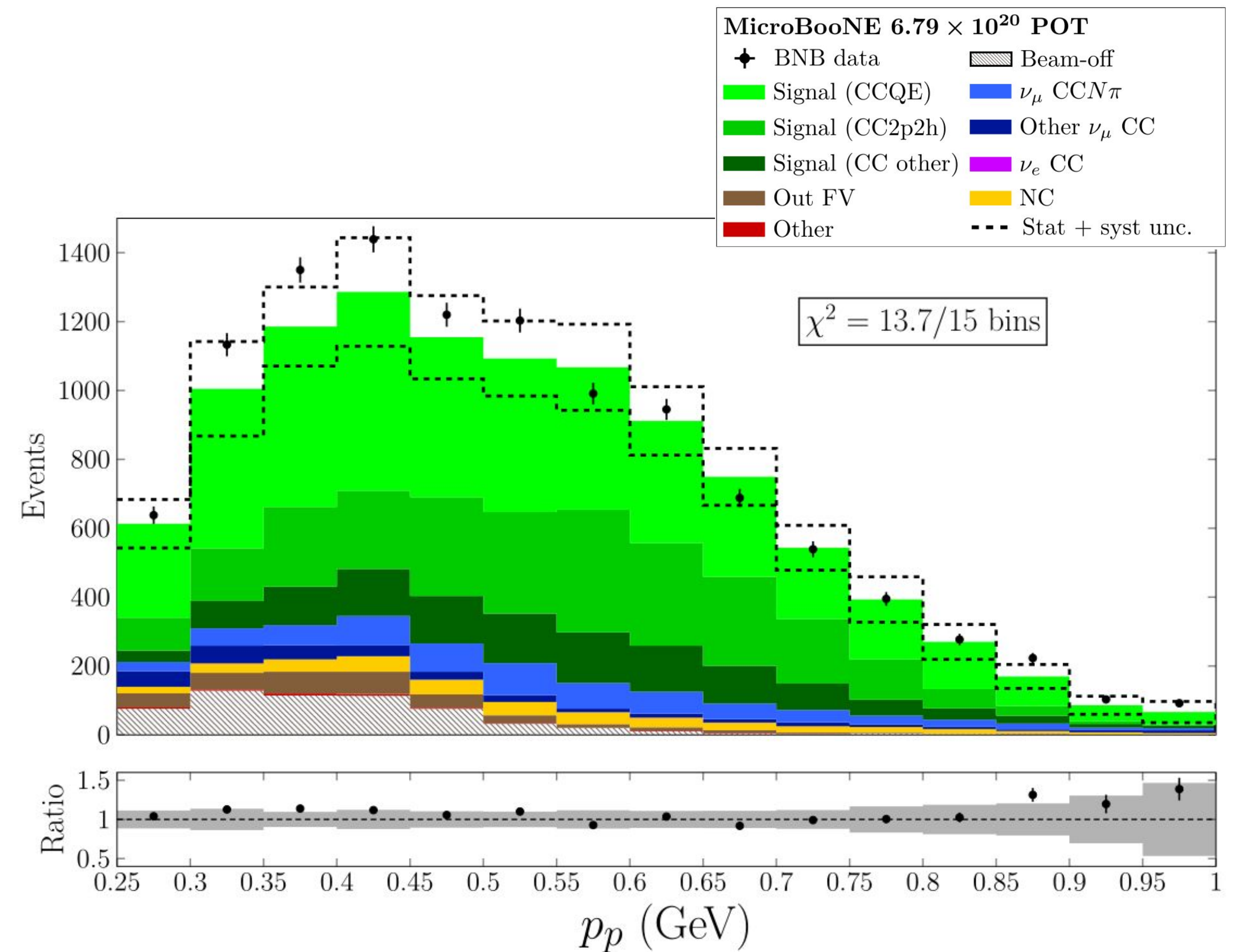
Conditional Covariance Background Constraint (CCBC)



Background control samples

- Minimizing model dependence is critical
 - We want to learn about Nature, not our simulation!
- Risk of biasing the measurement in both the unfolding (U) and **background subtraction (B)**
 - Sometimes we have to rely on the prediction
 - **Is it good enough to do this?** If not, how do we fix it?

$$\left\langle \frac{d^n \sigma}{d\mathbf{x}} \right\rangle_{\mu} = \frac{\sum_a U_{\mu a} (D_a - B_a)}{\Phi T \Delta \mathbf{x}_{\mu}}$$

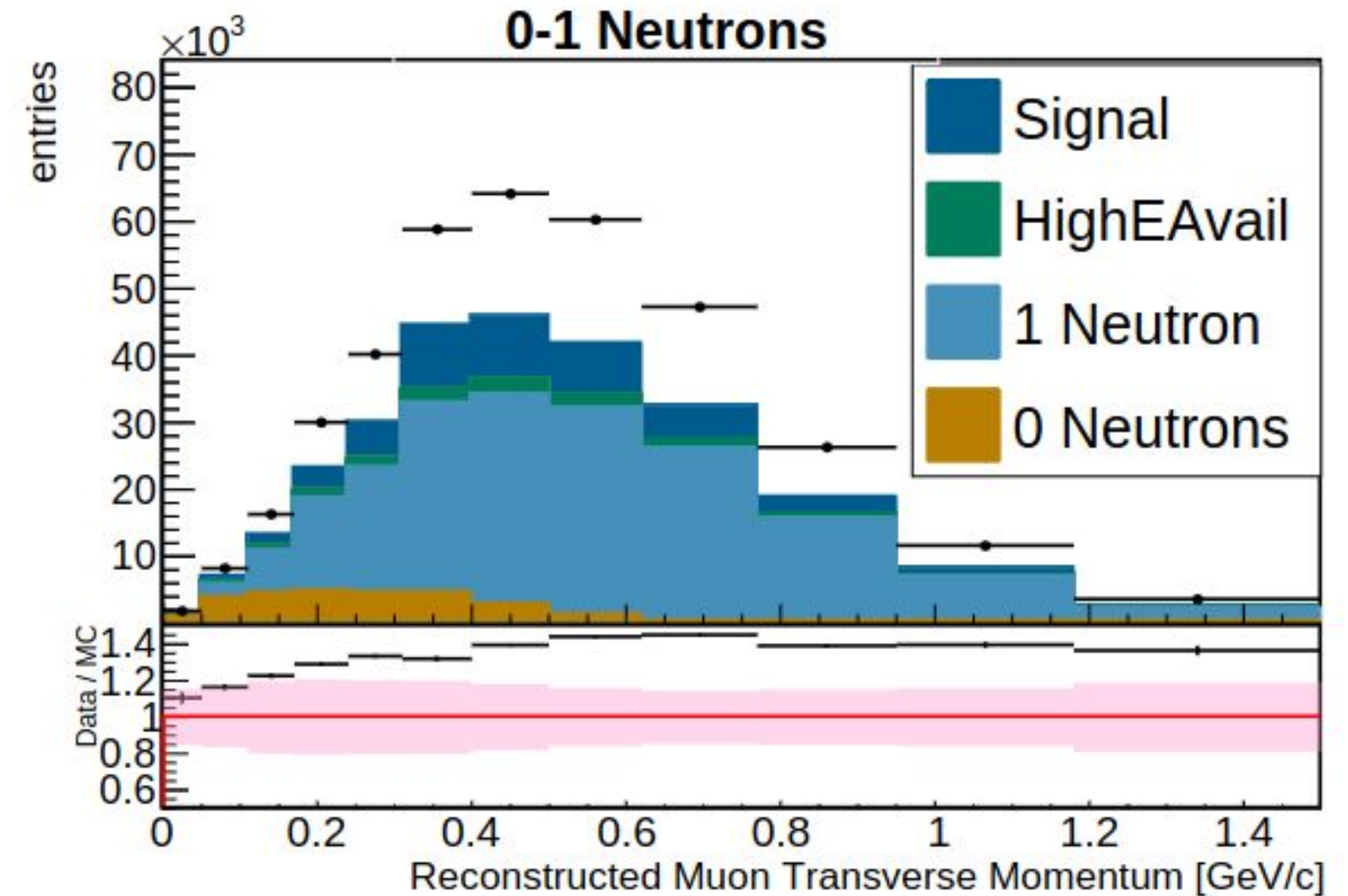


Background control samples

- **Control samples: check/correct background model** based on parallel measurement
 - Background-enhanced selection
- Also often referred to as "**sidebands**"
 - I use the terms interchangeably in the paper
- I propose a semi-new way of using these for cross-section analyses

[Phys. Rev. D 108, 112010 \(2023\)](#)

anti- ν_{μ} CC 2+ neutrons (MINERvA)



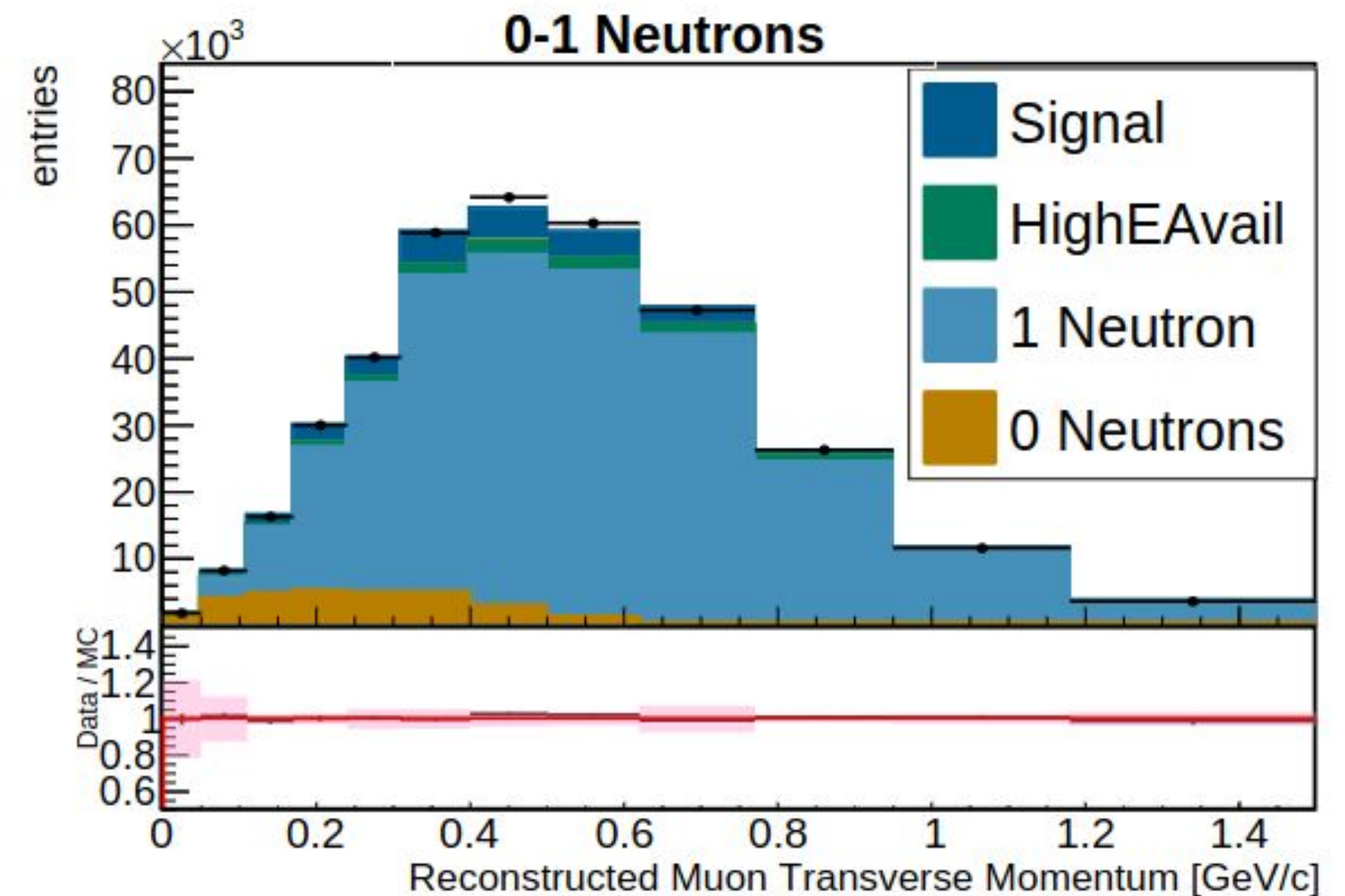
Few-neutron sideband (**pre-fit**)

Background control samples

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[Phys. Rev. D 108, 112010 \(2023\)](#)

anti- ν_{μ} CC 2+ neutrons (MINERvA)



Few-neutron sideband (**post-fit**)

Use by experiments

- **T2K** gets background model constraints "for free"
 - Just include bins from the sideband(s) in the fit!

- **MINERvA**: normalization scale factor approach

- **Pre-fit**: $\alpha_p = 1$ for all background classes p

- **Post-fit** values obtained from sidebands

- Details vary widely

- Shape from simulation unaltered*

- Implicit 100% correlation between α_p in sidebands and **signal region**

- **MicroBooNE**: no sidebands used as a constraint for any multi-bin cross-section result so far

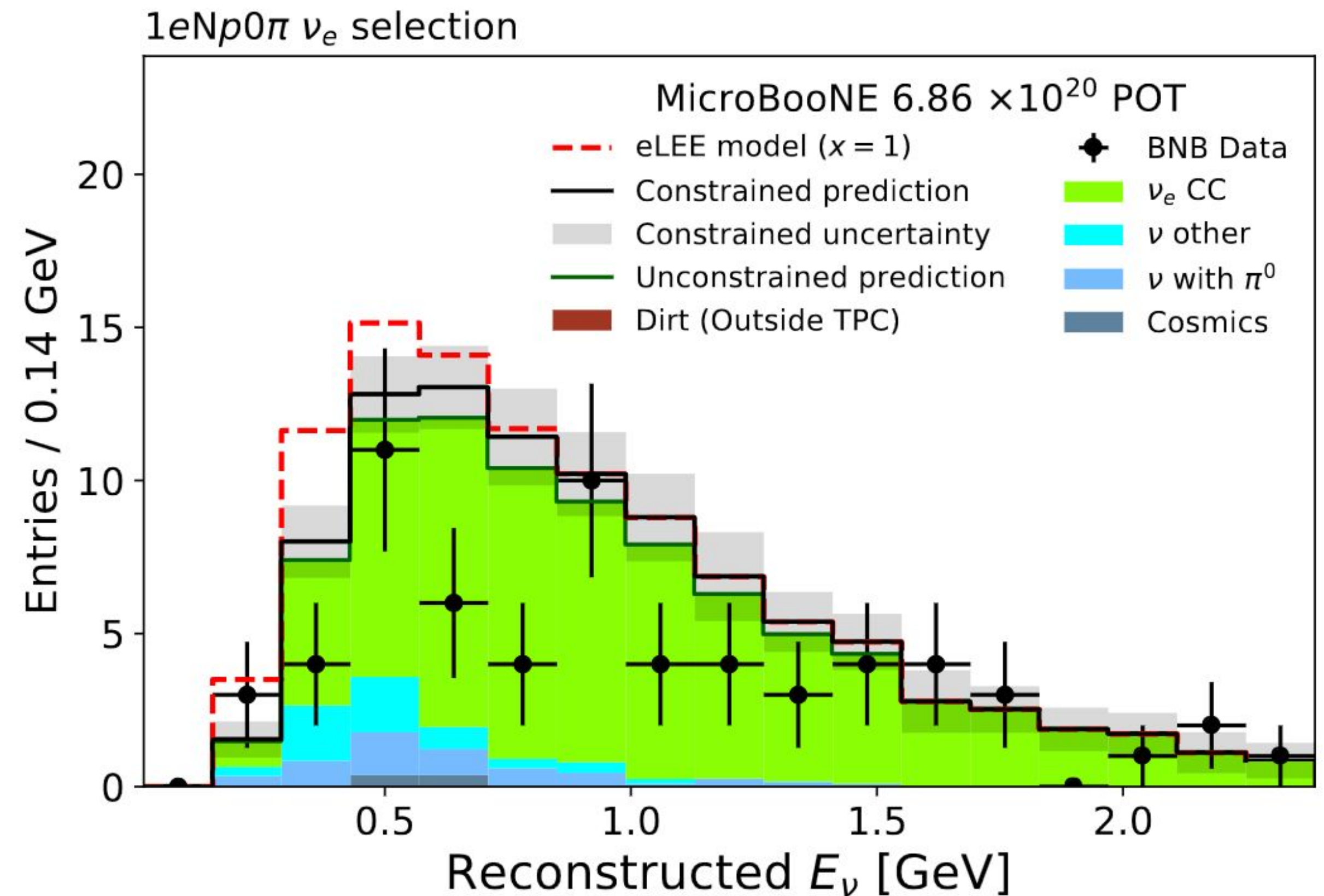
- I **generalize and improve** a method used for single-bin η analysis

$$B_a = \sum_p \alpha_p B_{ap}$$

Data-driven constraint in MicroBooNE LEE analyses

- MicroBooNE built to investigate anomalous excess of ν_e -like events seen by MiniBooNE at low energies ("**LEE**")
- First results October 2021
 - Data prefer **no excess**
- Judged relative to prediction of "MicroBooNE GENIE tune" with **data-driven, analysis-specific adjustments**
- All based on a **conditional covariance** treatment

[Phys. Rev. D 105, 112004 \(2022\)](#)



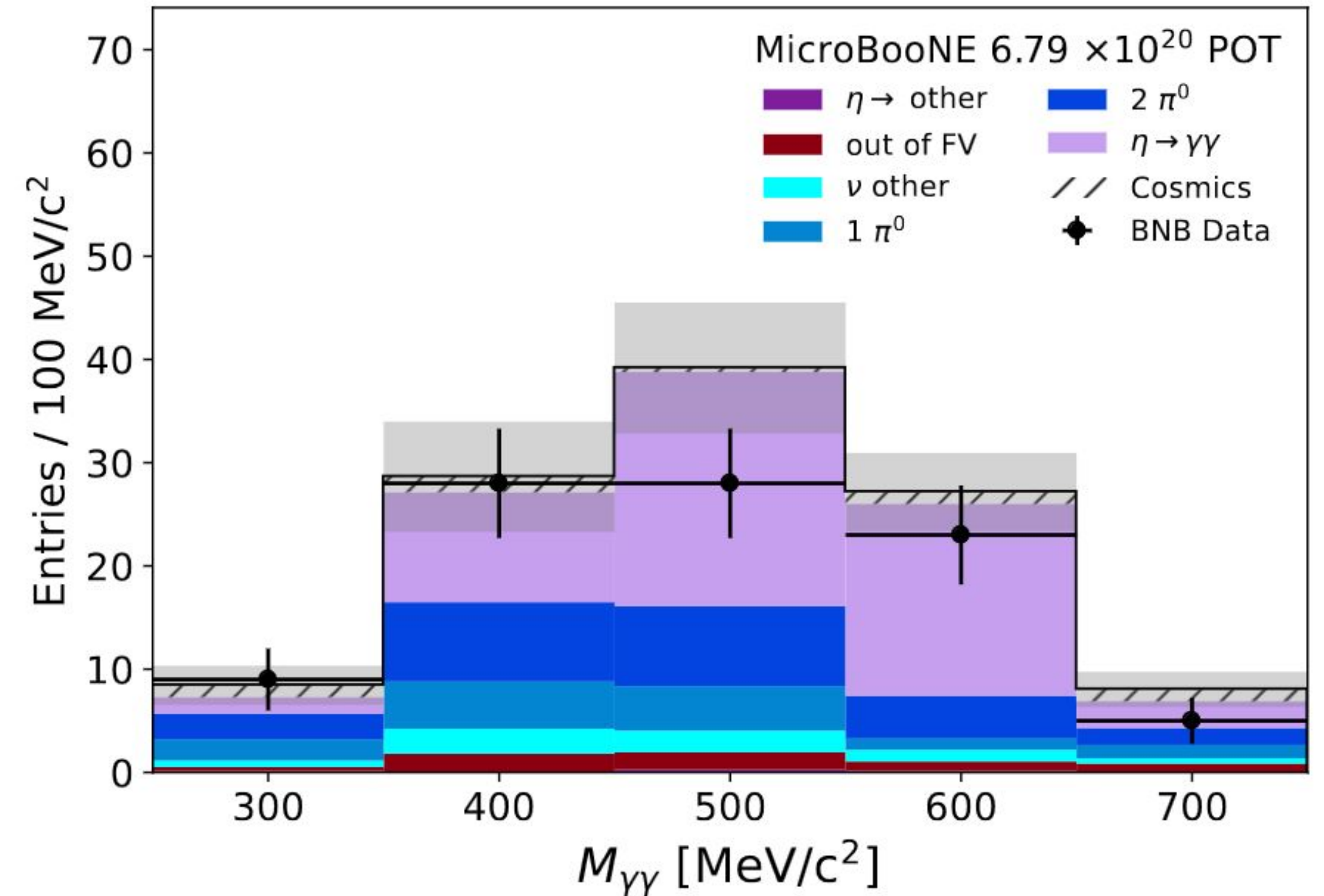
$$m^e \text{ constrained} = m^e + C^{e\mu} (C^{\mu\mu})^{-1} (n^\mu - m^\mu)$$

$$C^{ee} \text{ constrained} = C^{ee} - C^{e\mu} (C^{\mu\mu})^{-1} C^{\mu e}$$

Use for a background model constraint

- MicroBooNE η production study
 - Signal is two photons with the η invariant mass
- Dominant backgrounds are single- and multi- π^0 production
 - Each constrained separately with a single sideband bin
- I **generalize this procedure** for multiple bins and simultaneous fits to multiple backgrounds
 - Treatment suitable for MicroBooNE-style extraction

[Phys. Rev. Lett. 132, 151801 \(2024\)](#)



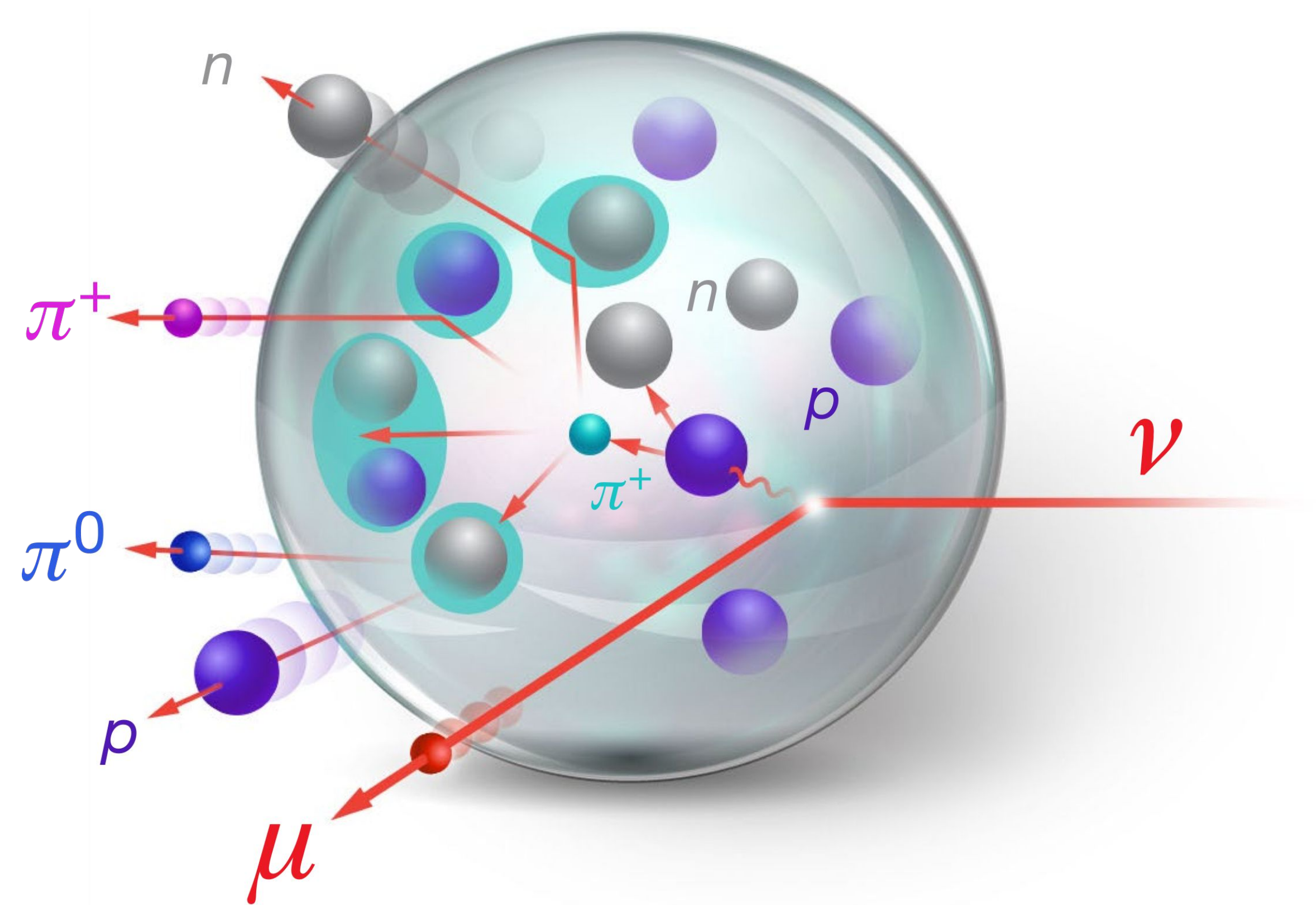
$$N_{MC}^{S, \text{constrained}} = N_{MC}^S + \frac{\sigma^{corr}}{(\sigma^B)^2} \times (N_{data}^B - N_{MC}^B)$$

$$(\sigma^{S, \text{constrained}})^2 = (\sigma^S)^2 - \frac{(\sigma^{corr})^2}{(\sigma^B)^2}$$

Can also be adapted to MINERvA's style (no 100% correlation assumption)

Conclusion

- Recent paper ([arXiv:2401.04065](https://arxiv.org/abs/2401.04065)) proposes some adjustments to how we extract neutrino cross section data
- "**Blockwise unfolding**" enables full reporting of correlated uncertainties
 - Make our hard work even more informative!
- **CCBC** provides somewhat new way of refining background predictions with data
 - Basic ideas have existed for some time, now applied to cross-section extraction



Backup