The Hadron Cascade

(and other related issues)

(a personal perspective)

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Nuclear response



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Modeling approximations

$$|\mathfrak{M}|^{2} \propto \frac{|\langle \Psi_{A} | \hat{i}\hat{T} | X \rangle|^{2}}{\text{Born approximation (IA)}}$$
CRPA
$$\approx \left| \sum_{\alpha} \langle \Psi_{A} | \hat{i}\hat{T} | \Psi_{\alpha} \rangle \langle \Psi_{\alpha} | X \rangle \right|^{2} \approx \sum_{\alpha} \left| \langle \Psi_{A} | \hat{i}\hat{T} | \Psi_{\alpha} \rangle \right|^{2} |\langle \Psi_{\alpha} | X \rangle|^{2}$$
+ intermediate hadronic state
+ time separation of FSI (MC factorization)
$$\approx \sum_{\alpha,h,p} \left| \langle \Psi_{h} | \hat{O}_{1b} | \Psi_{p} \rangle \langle \Psi_{A} | \hat{a}_{h}^{\dagger} \hat{a}_{p} | \Psi_{\alpha} \rangle \right|^{2} |\langle \Psi_{\alpha} | X \rangle|^{2}$$
+ one-body nuclear currents
$$\approx \sum_{h,p} \left| \langle \Psi_{h} | \hat{O}_{1b} | \Psi_{p} \rangle \right|^{2} \left| \langle \Psi_{A} | \hat{a}_{h}^{\dagger} | \Psi_{A-1} \rangle \right|^{2} |\langle \Psi_{A-1}, \psi_{p} | X \rangle|^{2}$$
+ cross section factorization (PWIA)
$$\approx \sum_{h,p} \sigma_{hp} S_{h}(E, p) P(p|X)$$
+ intranuclear cascade
$$\sum_{h,p} K_{N,MSc} A_{A.Nikolakopoulos et al., arXiv:2406.09244}$$

Monte Carlo event generators



 \rightarrow Cross sections are evaluated in a factorized scheme

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Intranuclear cascade

- Propagates particles through the nuclear medium
- **Probability** of passing a distance λ :

$$P(\lambda) = e^{-\lambda/\tilde{\lambda}}$$

where
$$\tilde{\lambda} \equiv \left(\rho\sigma\right)^{-1}$$
 and

- ρ local density
- σ cross section
- \rightarrow Implemented for nucleons, pions and kaons

T. Golan, C. Juszczak, J.T. Sobczyk, Phys.Rev. C 86 (2012) 015505



What can we rely on?



Final-state interactions in a many-body nuclear problem

Hadron propagation

Nucleon-Nucleus (NA) dynamics problem

- can be constrained by hadron scattering
- the nucleon should only propagate outwards
- \rightarrow e.g., optical potential, etc.

Built from Nucleon-Nucleon (NN) dynamics

- well-constrained nucleon-nucleon potential
- common practice across many-body nuclear problems
- \rightarrow e.g., *ab initio*, HF, etc.





Mean-field nuclear picture



\rightarrow let's try to use a realistic nucleon-nucleon potential to derive the central nuclear potential

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The iterative Hartree-Fock method

- $\rightarrow\,$ start with an initial guess for the average field or the wave functions
- $\rightarrow \mbox{ using the nucleon-nucleon potential } V(r,r')$ solve the equation

$$-\frac{\hbar^2}{2m}\nabla^2\varphi_{\mathfrak{i}}(\textbf{r})+U_{H}(\textbf{r})\varphi_{\mathfrak{i}}(\textbf{r})-\int U_{F}(\textbf{r},\textbf{r}')\varphi_{\mathfrak{i}}(\textbf{r}')d\textbf{r}'=\varepsilon_{\mathfrak{i}}\varphi_{\mathfrak{i}}(\textbf{r})$$

 $\rightarrow \,$ determine new values of $U_{H}(\textbf{r}),\,U_{F}(\textbf{r},\textbf{r}'),\,\varphi_{i}(\textbf{r}),\,\varepsilon_{i}$

 $\rightarrow\,$ at convergence: the final field $U_H(\textbf{r}),$ wave function $\varphi_i(\textbf{r}),$ and single-particle energy ε_i

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Nucleons in the mean-field potential



W. H. Dickhoff, D. Van Neck, Many-body Theory Exposed! (2005)

 \rightarrow nucleon lines are dressed according to the Hartree-Fock procedure

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Charge densities from the mean-field framework



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Relativistic mean-field

All of this can be also done in a **relativistic** framework:

- $\circ~$ Schrödinger equation \rightarrow Dirac equation,
- Wave functions \rightarrow **Dirac spinors**,
- Spin-orbit term comes for free!

$$\begin{split} \left(\tilde{E}\gamma_{0}-\vec{p}\cdot\vec{\gamma}-\tilde{M}\right)\psi &= 0\\ \tilde{E} &= E-V(r)\\ \tilde{M} &= M-S(r) \end{split}$$



Final-state interactions

• RPWIA:

- → plane-wave outgoing nucleon
- RDWIA:
- \rightarrow optical potential from NA
- RMSGA:
- → Glauber approximation from NN

(notice the low energy behavior)

 \rightarrow for more results see the talks by Raúl or Juanma!



M.C. Martinez et al., Phys.Rev. C 73 (2006) 024607

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Modeling the fate of hadrons in final-state interactions

Classical particle propagating through a medium

Probability for a particle to propagate over a distance x with **no interactions** is

$$P(x) = \frac{1}{\lambda} \exp(-x/\lambda)$$

where $\lambda = (\rho \sigma)^{-1}$ is the mean free path, while ρ is target density and σ is interaction cross section

We can try to apply it to nucleons in nuclei because:

$$\tilde{\lambda} \ll d < \lambda < R$$

where $\tilde{\lambda}$ is the de Broglie wavelength, d is the distance between targets, and R is the nuclear radius

N. Metropolis et al., Phys.Rev. 188 (1958) 185, Phys.Rev. 188 (1958) 204

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Application to nuclei (space-like approach)

- Pick a random starting point in the nucleus
- $\circ~$ Propagate the nucleon in discrete steps, e.g., $\Delta x=0.2~\text{fm}$
- At every step, we sample x from $P(x) = \lambda^{-1} \exp(-x/\lambda)$
- \rightarrow If x < Δ x, then the nucleon-nucleon interaction happens
- The probability that the nucleon leaves the nucleus with no re-interactions is called **transparency**
- \rightarrow Our procedure solves an integral

$$\mathsf{T} = \int_0^{2R} \mathsf{f}_{\mathsf{R}}(z) e^{-z/\lambda} \, \mathrm{d}z$$

where $f_R(z)$ is the distribution of the starting points



• For fixed density (uniform ball), the solution is given analytically

$$\mathsf{T} = 3e^{-\mathsf{A}} \left(\frac{1}{\mathsf{A}^2} + \frac{1}{\mathsf{A}^3} \right) + 3 \left(\frac{1}{2\mathsf{A}} - \frac{1}{\mathsf{A}^3} \right)$$

where $A=2R/\lambda=2R\rho\sigma$

 $\rightarrow\,$ e.g., for $\rho=0.16~fm^{-3},\,\sigma=40$ mb, and R=6 fm, we get $T\simeq0.189$



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Time-like approach

• tracking all particles

evolution over dt

 calculating collision probability in a (dt, d³x) box



Cascade models

	algorithm	nuclear model	cross sections	potential
GENIE hA	space-like	LFG	SAID	-
GENIE hN	space-like	LFG	data + ?	-
NEUT	space-like	LFG	data + ?	-
NuWro	space-like	LFG	data + in-medium	-
Achilles	time-like	ab initio	data	-
INCL	time-like	HF	data + ?	constant
GiBUU	time-like	RMF and HF (LFG)	off-shell	RMF and HF

 \rightarrow see the talk by Patrick for an overview of MC generators' efforts!

Highlights: GENIE



J. Plows, NuFACT 2024

→ variety of models available: homegrown hA and hN, INCL++, Bertini (Geant4)

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Highlights: NEUT

Relativistic Mean Field

- First implementation of a macroscopic model based on a Relativistic Mean Field optical models into NEUT:
 - Jake McKean, Raul González-Jiménez, Minoo Kabirnezhad
- Potential for new theory-motivated systematic uncertainty studies in NEUT.
- Possible consideration of alternative operators for different processes.
 - E.g. Kabirnezhad inelastic pion production model operator.





L. Pickering, NuFACT 2024

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Highlights: NuWro



KN, J. Sobczyk, Phys.Rev. C 100 (2019) 015505

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Highlights: Achilles



J. Isaacson et al., Phys.Rev. C 103 (2021) 015502

\rightarrow time-like cascade based on *ab initio* nucleon distributions

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Highlights: INCL



A. Ershova et al., Phys.Rev. D 106 (2022) 032009

 \rightarrow softer interactions with higher ingredient complexity, e.g. cluster emission

Highlights: GiBUU

Nuclear ground state:

- **density distribution**: Woods-Saxon
- particle momenta:

 $f(\vec{r},\vec{p}) = \Theta\left[p_F(\vec{r}) - |\vec{p}|\right]$

• Fermi-momentum:

$$p_{\rm F}(\vec{r}) = (3\pi^2 \rho(\vec{r}))^{1/3}$$

• Fermi-energy:

$$\mathsf{E}_{\mathsf{F}} = \sqrt{\mathsf{p}_{\mathsf{F}}^2 + \mathsf{M}_{\mathsf{N}}^2} + \mathsf{U}_{\mathsf{M}\mathsf{F}}(\vec{\mathsf{r}},\mathsf{p}_{\mathsf{F}})$$



K. Gallmeister, NuSTEC School 2017

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Highlights: GiBUU



K. Gallmeister, NuSTEC School 2017

Constraining through electron scattering (and other probes)

Exclusive electron scattering

Exclusive QE proton knockout at **fixed kinematics**:

- beam: E_e
- electron: $E_{e'}, \theta_{e'}, \phi_{e'}$
- proton: E_p, θ_p, ϕ_p

With provided:
$$\frac{\Delta p}{p}$$
, $\Delta \theta$, $\Delta \phi$

Cuts on "missing" variables:

- energy: $E_m = \omega T_p T_{A-1}$
- momentum: $\vec{p}_m = \vec{p}_p \vec{q}$

 $E_m < 80~{\rm MeV}, ~|\vec{p}_m| < 300~{\rm MeV/c}$



Transparency:

$$\langle T \rangle_{\theta_p} = \frac{\sigma_{\exp}}{\sigma_{PWIA}} = \frac{[after FSI]}{[total]}$$

 $\sigma_{\rm PWIA}$ - expected value without FSI (model dependent)

Definition

Nuclear transparency is the average **probability** for a knocked-out **proton** to **escape** the nucleus **without significant reinteraction**.

e.g. measured for Carbon: T \simeq 0.60 [D. Abbott *et al.*, PRL 80 (1998), 5072]



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KN, J. Sobczyk, Phys.Rev. C 100 (2019) 015505



KN, J. Sobczyk, Phys.Rev. C 100 (2019) 015505



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Exclusive electron scattering



JLab E91013 (D. Dutta et al.), Phys.Rev. C 68 (2003) 064603



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Issues with transparency measurements

• Performed at high energies

(color transparency)

• Valid up to the **first inelastic interaction**

(imaginary part of the optical potential)

• Results are **model-dependent**

(spectroscopic factors)

• Issues with consistency across experimental analyses



K. G. Fissum et al., Phys.Rev. C 70 (2004) 034606

E_{miss} [MeV]

80

60

40

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20

10-4

100

120

Merging together the optical potential and intranuclear cascade



A. Nikolakopoulos et al., Phys.Rev. C 105 (2022) 054603

 \rightarrow see the talk by Anna for more results and comparisons!

Summary

- We have many independent models capable of propagating hadrons through nuclear matter
- Although sometimes simplistic, these solutions work remarkably well
- Comparisons to hadron and electron scattering provide vital information
- We need more dedicated, neutrino-like experimental analyses to test our models:
- \rightarrow Data should focus on clear experimental signals to test various dynamics
- \rightarrow This requires the best possible knowledge of the **primary interaction**

Supplementary material

Mean-field potential

Single-particle radial Schrödinger equation:

$$(T + U(r)) \varphi_{\alpha}(r) = \varepsilon_{\alpha} \varphi_{\alpha}(r)$$

Nuclear Hamiltonian:

$$H_0 = \sum_{i=1}^A \left(T_i + U(r_i) \right) = \sum_{i=1}^A h_0(i), \quad E_0 = \sum_{i=1}^A \varepsilon_{\alpha_i}(r_i)$$

Nuclear wave function is a Slater determinant:

$$\Phi_{\alpha_1,...,\alpha_A}(\mathbf{r}_1,...,\mathbf{r}_A) = \frac{1}{\sqrt{A!}} \begin{vmatrix} \phi_{\alpha_1}(\mathbf{r}_1) & \dots & \phi_{\alpha_1}(\mathbf{r}_A) \\ \vdots & \ddots & \vdots \\ \phi_{\alpha_A}(\mathbf{r}_1) & \dots & \phi_{\alpha_A}(\mathbf{r}_A) \end{vmatrix}$$

Let's restrict ourselves to two-body interactions only and evaluate the mean-field potential:

$$H = \sum_{i=1}^{A} T_i + \frac{1}{2} \sum_{i,j=1}^{A} V_{ij}$$
(4)

$$H = \sum_{i=1}^{A} (T_i + U(r_i)) + \left(\frac{1}{2} \sum_{i,j=1}^{A} V_{ij} - \sum_{i=1}^{A} U(r_i)\right) = H_0 + H_{res} = \sum_{i=1}^{A} h_0(i) + H_{res},$$
(5)

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(1)

(2)

(3)

Hartree-Fock methods

Let's consider a density in terms of the occupied single-particle states:

$$\rho(\textbf{r}) = \sum_b \varphi_b^*(\textbf{r}) \varphi_b(\textbf{r})$$

The Hartree potential at a given point generated by the two-body interaction:

$$U_{H}(\textbf{r}) = \sum_{b} \int \varphi_{b}^{*}(\textbf{r}') V(\textbf{r},\textbf{r}') \varphi_{b}(\textbf{r}') d\textbf{r}'$$

The Schrödinger equation becomes:

$$-\frac{\hbar^{2}}{2m}\nabla^{2}\phi_{i}(\mathbf{r}) + \sum_{b}\int\phi_{b}^{*}(\mathbf{r}')V(\mathbf{r},\mathbf{r}')\phi_{b}(\mathbf{r}')d\mathbf{r}'\cdot\phi_{i}(\mathbf{r}) -\sum_{b}\int\phi_{b}^{*}(\mathbf{r}')V(\mathbf{r},\mathbf{r}')\phi_{b}(\mathbf{r})\phi_{i}(\mathbf{r}')d\mathbf{r}' = \varepsilon_{i}\phi_{i}(\mathbf{r}) -\frac{\hbar^{2}}{2m}\nabla^{2}\phi_{i}(\mathbf{r}) + U_{H}(\mathbf{r})\phi_{i}(\mathbf{r}) - \int U_{F}(\mathbf{r},\mathbf{r}')\phi_{i}(\mathbf{r}')d\mathbf{r}' = \varepsilon_{i}\phi_{i}(\mathbf{r})$$
(9)

where the exchange term is driven by the Fock potential:

$$U_{F}(\mathbf{r}) = \sum_{b} \phi_{b}^{*}(\mathbf{r}') V(\mathbf{r}, \mathbf{r}') \phi_{b}(\mathbf{r})$$
(10)

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(6)

(7)