

Cross section Uncertainties in NOvA Analyses

Jeremy Wolcott

Tufts University

on behalf of the NOvA collaboration

ECT* Workshop on ν Interactions
October 22, 2024

Context

[NOvA physics scene & expt] — [Base cross section model]



NOvA physics scene



NOvA physics scene

Symmetry Magazine / Sandbox Studio, Chicago



Which way are the
neutrino states ordered?



Do ν_μ/ν_τ **mix equally**
into the ν mass states?



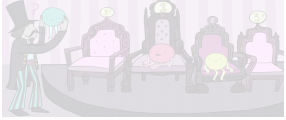
$$\Delta P_{\nu\bar{\nu}} \propto \sin \delta_{CP}$$

Do neutrinos exhibit
CP violation?

① 3-flavor
neutrino
oscillations

NOvA physics scene

Symmetry Magazine / Sandbox Studio, Chicago



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3-flavor

①

neutrino oscillations

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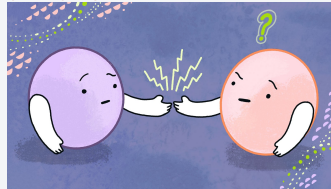
Do neutrinos exhibit CP violation?

Are there light sterile neutrinos?



particlezoo.net

Symmetry Magazine / Corinne Mucha / Sandbox Studio



$$H = UH_0U^\dagger + H_{matter} \dots + H_{NSI}??$$

Do we fully understand neutrino propagation in matter?

② Neutrino oscillations beyond PMNS

NOvA physics scene

Symmetry Magazine / Sandbox Studio, Chicago



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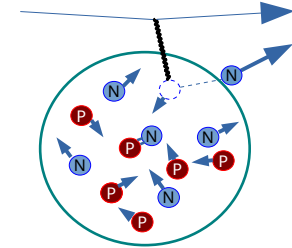
particlezoo.net

②

Neutrino
oscillations
beyond PMNS

$$H = UH_0U^\dagger + H_{matter} \\ \dots + H_{NSI}??$$

Do we fully understand
neutrino propagation in matter?



Is our picture of **neutrino scattering** sufficient?

③

Neutrino
scattering

NOvA physics scene

Symmetry Magazine / Sandbox Studio, Chicago



Which way are the neutrino states ordered?

3-flavor

①

Do ν_e/ν_μ mix equally into the mass states?

neutrino oscillations



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Neutrino oscillations beyond PMNS

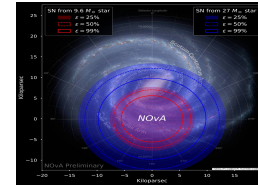
$$H = -UH_0U^\dagger + H_{matter} \dots + H_{NSI}??$$

Do we fully understand neutrino propagation in matter?

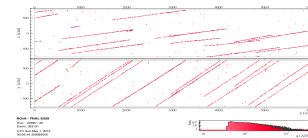
③

Neutrino scattering

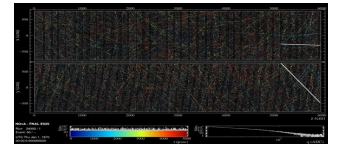
Is our picture of neutrino scattering sufficient?



Will we observe a galactic supernova?



What can we learn from cosmic muons?

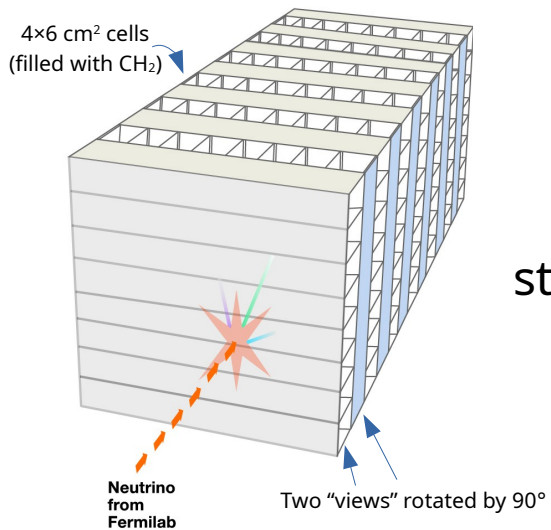
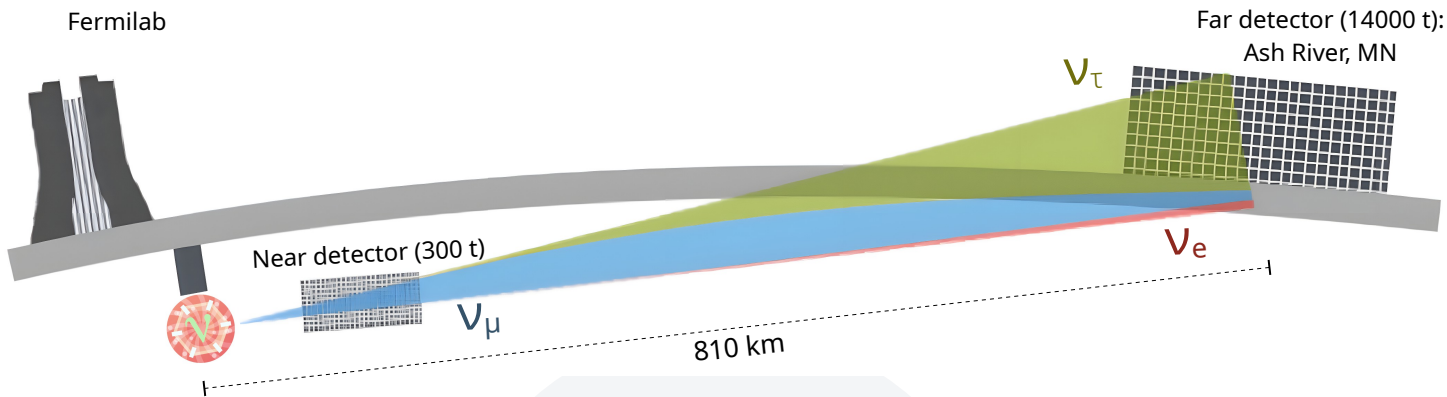


Do we see magnetic monopoles or other exotic phenomena?

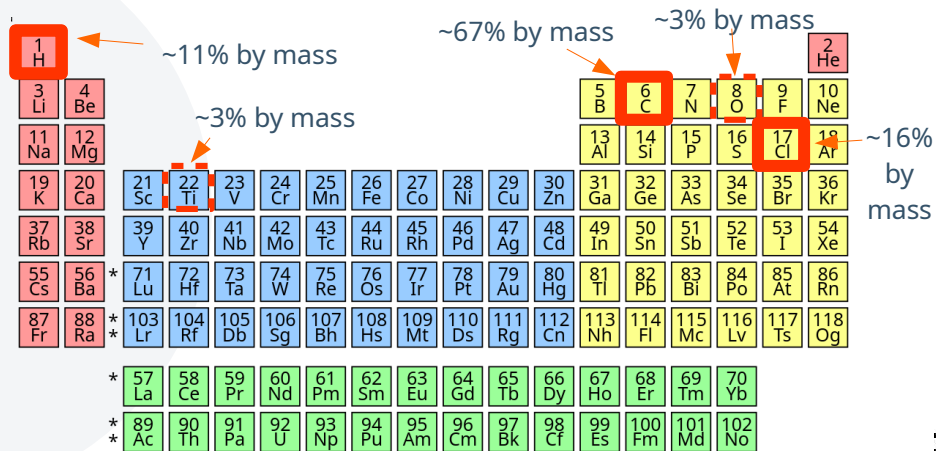
④

Opportunistic/exotic physics

NOvA: apparatus & vs

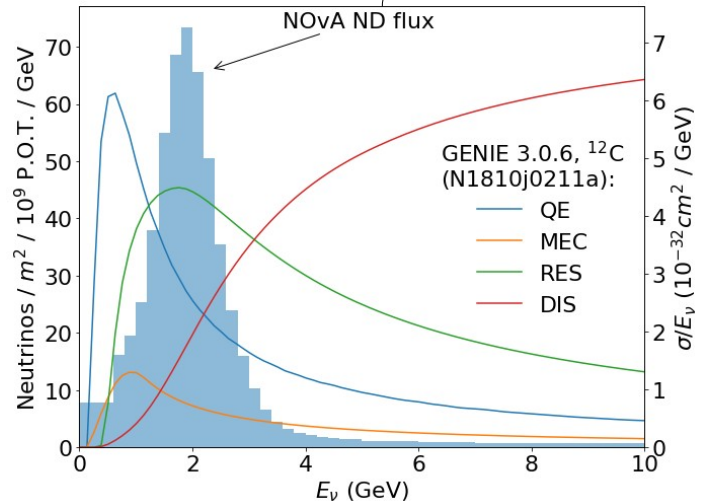
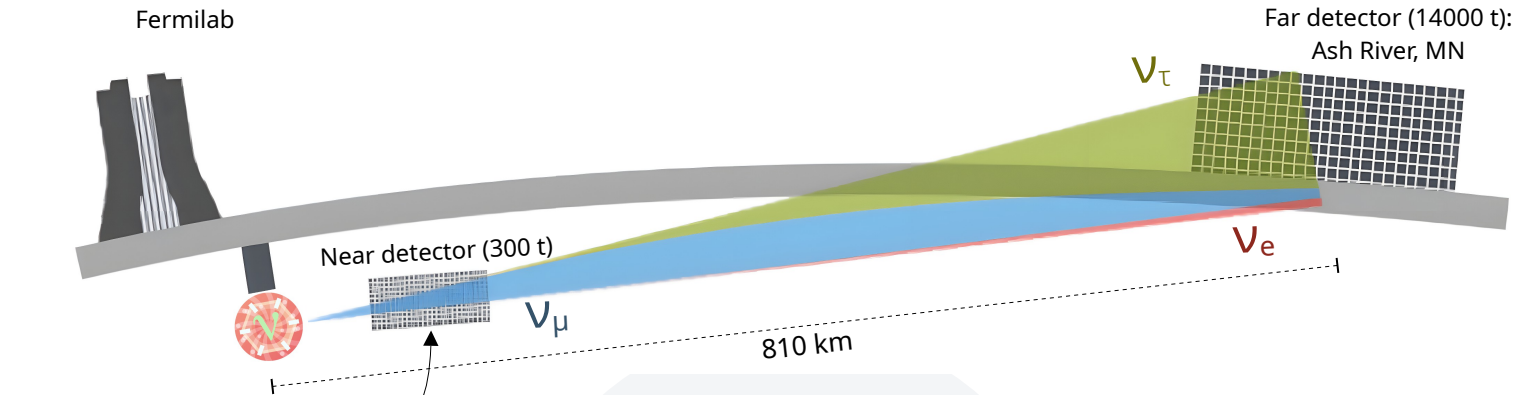


One moderate &
one *really big*
stereoscopic detector...



... made of PVC and filled with CH₂...

NOvA: apparatus & vs

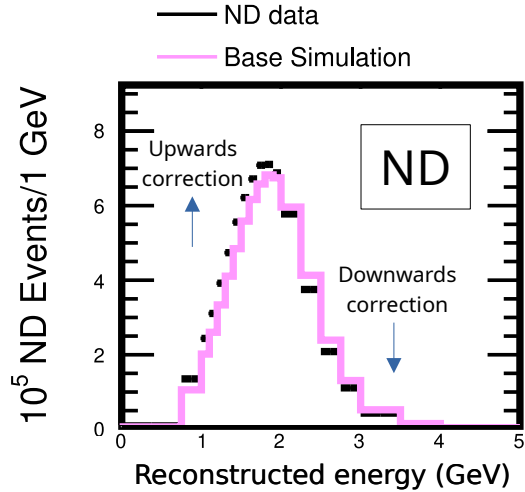


... and illuminated with
a ~ 2 GeV ν_μ beam

Analysis context

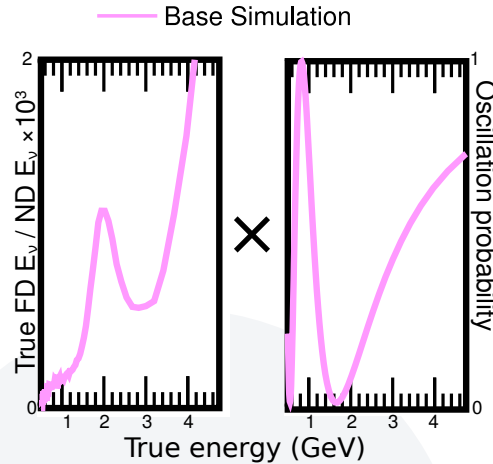
- Most uncertainties developed in context of **3-flavor oscillations**
 - Rely on **strong** ND-FD correlations/cancellation via “extrapolation” technique
→ next slide
 - **Uncertainties typically chosen conservatively** to “bracket possibilities”
 - Central value of model less important than plausible bounds from uncertainties
- Other customers (beyond-PMNS osc., cross section measurements) develop **separate uncertainties as needed**
 - ... where ND-data-driven uncertainties risk biasing results
 - ... where other ν interaction processes (e.g.: NC) become dominant

Analysis context

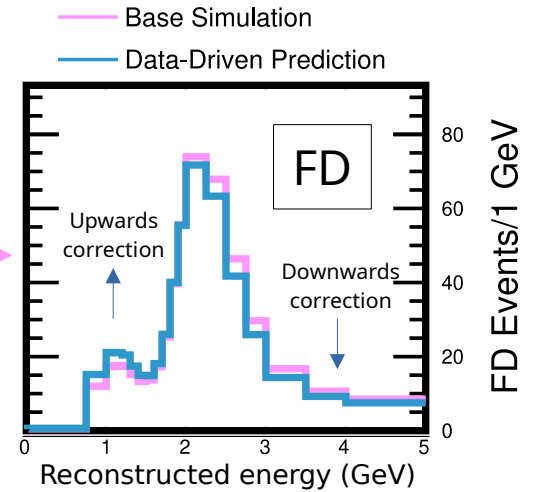


Convert to true E

×



Convert to reco E



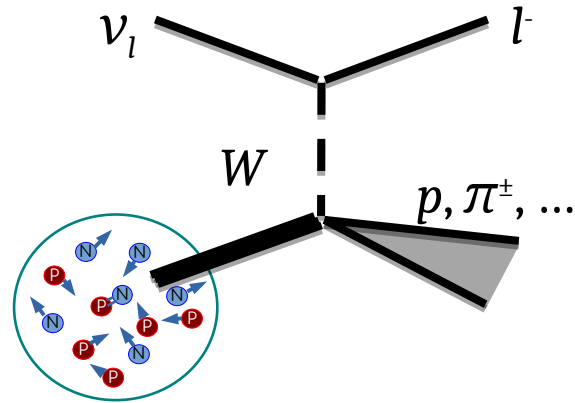
Correct ND simulation
to agree with data in reco E_ν...

... via Far/Near transformation that
comprises well understood effects
(beam divergence, detector
acceptance) + **oscillations**

... results in **constrained**
FD E_ν prediction highly correlated
with ND correction

3-flavor osc., NSI “**extrapolation**” analysis is *data-driven*
and can correct even “unknown unknowns”,
so long as the base model isn’t “too far” off

Base cross section model



Base simulation: GENIE 3.0.6

- We begin with a “theory-driven” set of models* ...

... but it agrees poorly with our data
 (as do all other stock model configurations we tested) ...
 ... so we apply *post hoc* “tuning” to get the model “close
 enough”

QE	Multinucleon	RES	DIS	FSI
València 1p1h w/ Z-expansion axial form factor	València MEC	Berger-Sehgal	Bodek-Yang + AGKY/Pythia	hN semi-classical cascade

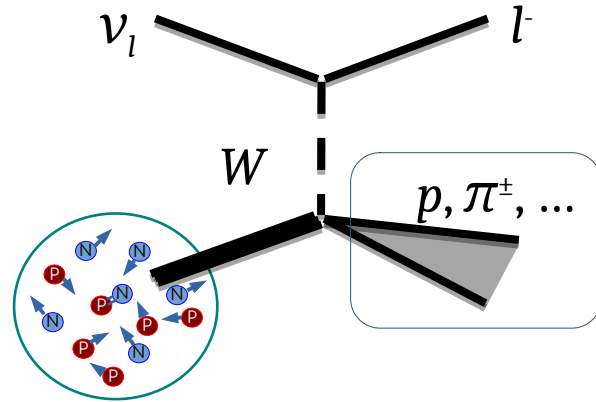
* We call our model collection **N18_10j_00_000**. It is built by starting with GENIE's **G18_10b_00_000** and substituting the Z-expansion QE axial form factor for the dipole one. This combination was not available in the 3.0.6 release, but it could be in future versions.

Base cross section model

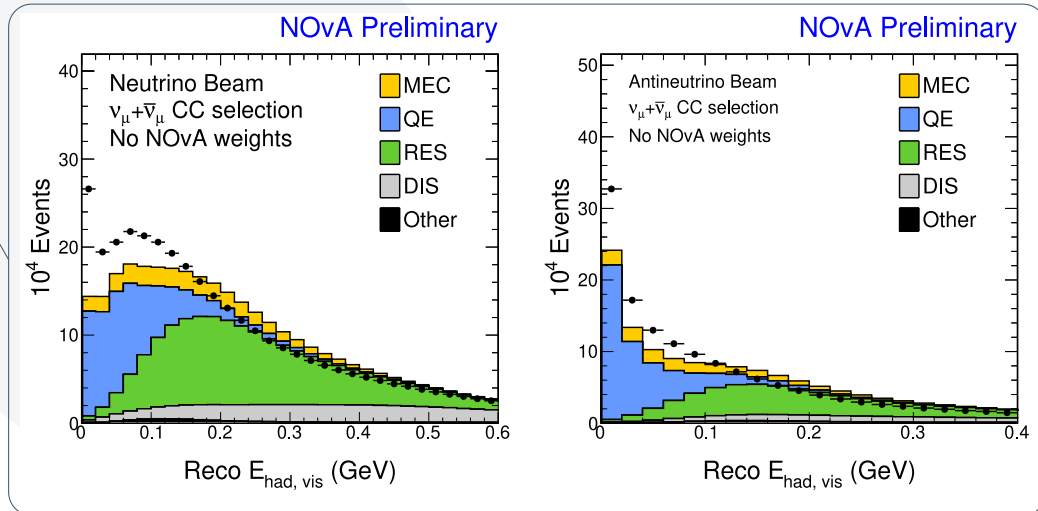
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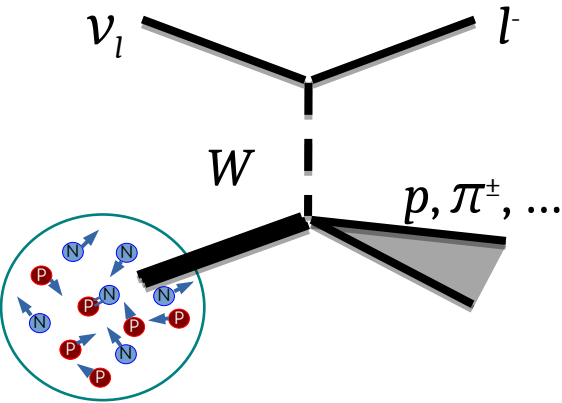
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QE

València 1p1h w/
Z-expansion axial
form factor

Multinucleon

València MEC

RES

Berger-Sehgal

DIS

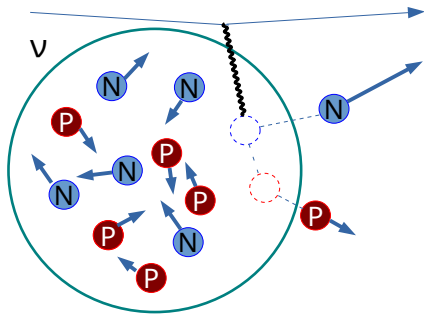
Bodek-Yang
+
AGKY/Pythia

FSI

hN semi-classical
cascade

For analyses that use “extrapolation”,
apply custom “tuning” in two places
in tandem with custom uncertainties

Tuning: multinucleon

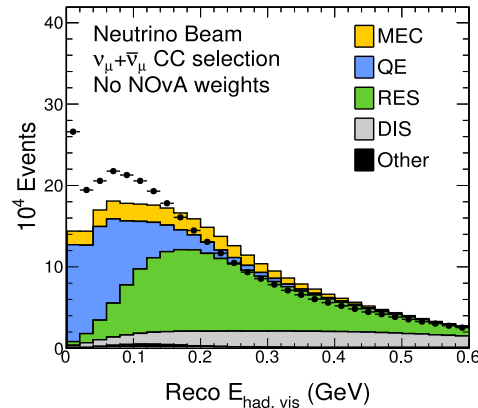


"2p2h"

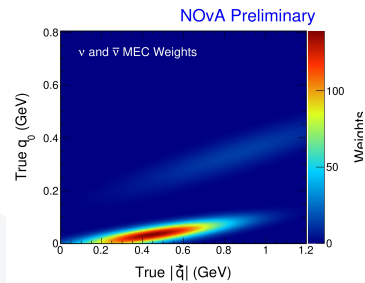
Knock out two nucleons
with an elastic-like interaction.

Most (?) uncertain part of the
model...

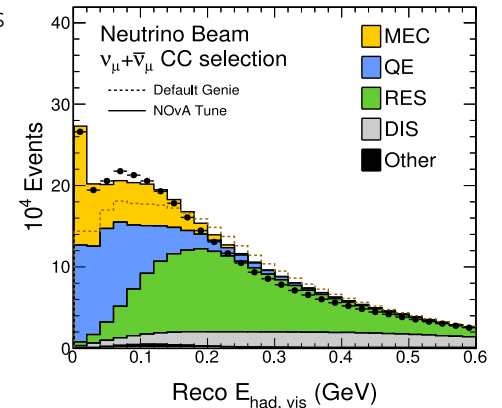
NOvA Preliminary



Fitted double-Gaussian weights
applied to true CC MEC



NOvA Preliminary



Employ **fits to NOvA ND data** to get in the ballpark

[detailed discussion of technique with different base model
in NuSTEC CTGWG Seminar [Dec. 14, 2022](#) (J. Wolcott)]

Dedicated uncertainties address dependence
on the other components of prediction (QE, RES)
(more shortly)

[Note: tuning procedure is used for **3-flavor osc., NSI.**
But **Sterile ν, cross sections** use *unaltered* Valencia base model.]

Tuning: FSI

- FSI model choice: "hN"
 - Propagates hadrons through nucleus in finite steps
 - Interaction probabilities simulated according to **Oset quantum model**
 - More rigorous foundation than "hA" effective model?...

Challenge: hN not directly reweightable

- Addressed with BDT reweighting technique adapted from DUNE (see overflow)

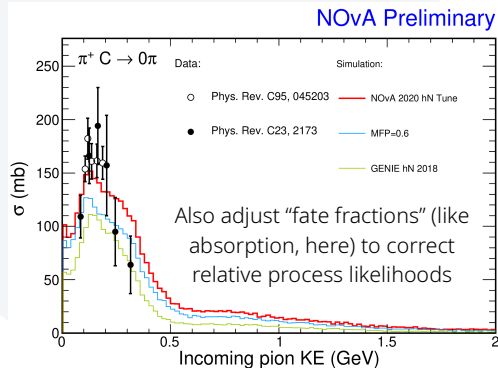
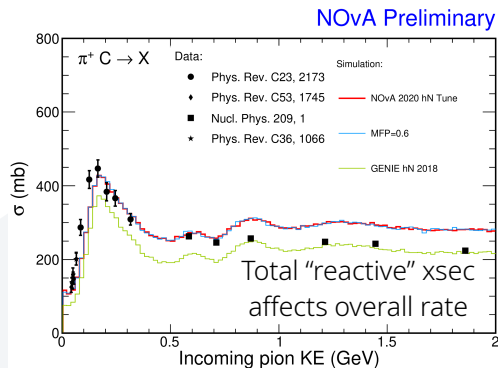
Unlike MEC tuning, FSI adjustment is used by *all* analyses since it does not depend on ν data

Adjust central value to better match

π^\pm scattering data

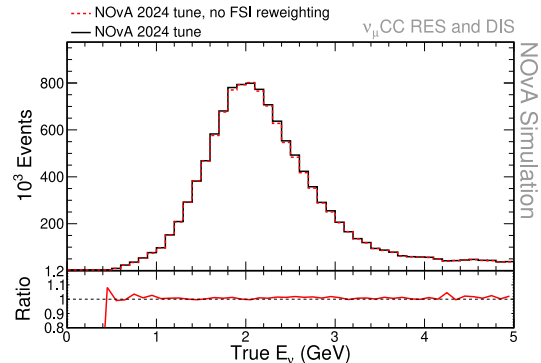
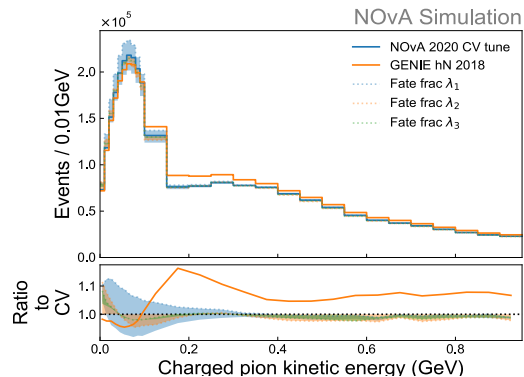
in regions most relevant to NOvA

($T_\pi \lesssim 500$ MeV)



Impact

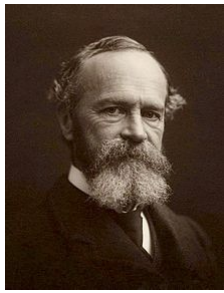
- 5-10% effect on pion kinematics (more on uncertainties shortly)
- Ultimately subdominant for calorimetric E_ν reco. used in NOvA



Current cross section uncertainties

[Philosophy] — [GENIE knobs] — [Bespoke knobs]

Cross section philosophy



“There is only one thing a philosopher can be relied upon to do, and that is to contradict other philosophers”
– William James, philosopher

- We have not sought “the right” cross section model – just a *functional* model
 - “Extrapolation” insulates against many potential pitfalls, only requires being “close enough”
 - Viable theoretical underpinnings are not always needed (so long as data description is reasonable), though unquestionably preferred
 - NOvA detector design is relative coarse (by modern standards) → hadronic systems are poorly resolved
 - Cross section measurements help us gauge how far off we are
- We typically choose *representative* uncertainties
 - Uncertainties to gauge whether “extrapolation” has *weaknesses*—don’t need an exhaustive collection to get the right answer
 - “Bracketing” space of possibilities is usually sufficient
 - Fake data studies with out-of-model warping used to test for possible issues
- We have taken a **NOvA-centric view** to which “holes” to prioritize plugging first
 - $1p1h/2p2h \ 0\pi \supset \text{RES } 1\pi \gg \text{DIS}/N\pi$
 - $\nu \supset \bar{\nu}$

Types of uncertainties

“Inherited”

GENIE-Reweight

“NOvA Bespoke”

Based on external info

Based on NOvA data

GENIE-Reweight uncertainties

CC QE

Dipole FF masses
 M_A, M_V

Coulomb
off/on

Z-exp. FF coeffs
 a_0, a_1, a_2, a_3, a_4

RPA
off/on

CC MEC

np fraction
of init. state

NN pair
"decay" angle

fraction
of MEC from
intermediate Δ

kinematic
shape

CC RES

Dipole FF masses
 M_A, M_V

$\Delta \rightarrow N\gamma$
branching
fraction

Δ polarization

$\Delta \rightarrow N\eta$
branching
fraction

COH

Axial mass M_A

NC norm

Effective
nuclear
radius R_0

CC norm

CC "SIS"

$\{\bar{\nu}, \nu\} \times \{n, p\} \times \{1, 2\}\pi$
normalization

CC DIS

Bodek-Yang
 $A_{ht}, B_{ht}, C_{V1u}, C_{V2u}$

AGKY
 $x_F^{1\pi}, p_T^{1\pi}$

hA FSI

Mean
free paths

"Fate fractions"
 $f_{Inel}^\pi, f_{CEX}^\pi, f_{ABS}^\pi, f_{\pi prod}^\pi$
 $f_{Inel}^N, f_{CEX}^N, f_{ABS}^N, f_{\pi prod}^N$

... plus NC variants
of many of these

GENIE comes bundled with a large suite of reweightable uncertainties

GENIE-Reweight uncertainties

CC QE

Dipole FF masses
 M_A, M_V

Coulomb
lepton-nucl
off/on

Z-exp. FF coeffs
 a_0, a_1, a_2, a_3, a_4

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hA FSI

Mean "Fate fractions"
free paths $f_{\text{Inel}}^\pi, f_{\text{CEX}}^\pi, f_{\text{ABS}}^\pi, f_{\pi \text{ prod}}^\pi$
 λ_π, λ_N $f_{\text{Inel}}^N, f_{\text{CEX}}^N, f_{\text{ABS}}^N, f_{\pi \text{ prod}}^N$

... plus NC variants
of many of these

For NOvA:

- We use **some** of these knobs as-is
- Some **others** are irrelevant for our chosen configuration
- We've implemented **our own knobs** that provide alternate behavior to many others

GENIE comes bundled with a large suite of reweightable uncertainties

Bespoke NOvA uncertainties

NOvA focus areas

- CCQE
 - Reimplementation of z-expansion a_1 - a_4 to preserve correlations
 - RPA [predates similar GENIE knob]
- MEC
 - Kinematic shape [coupled to fitting]
 - E_ν dependence
 - np fraction [predates similar GENIE knob]
- RES
 - Low Q^2 suppression
 - Δ vs. non- Δ resonance composition
- Pion production [RES + {D/S}IS]
 - Charged vs. neutral baryon production
 - "SIS" channel normalizations

Making non-reweightable GENIE parameters reweightable

- hN FSI
 - Hadronization formation zone
- No comparable
GENIE knob(s)**
- V_e / V_μ
 - Radiative corrections
 - 2nd class currents

Replacing GENIE knobs that used to be broken

- Coherent
 - CC, NC norms

Bespoke NOvA uncertainties

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Making non-reweightable GENIE parameters reweightable

- hN FSI
- Hadronization formation zone
- Coherent

Replacing GENIE knobs that used to be broken

I'm going to explore these 3 further,
where NOvA choices are more unique

(but feel free to ask about any others
you're interested in as well)

Multinucleon uncertainties

Two separate use cases

ND provides *systematic control*:
3-flavor osc., NSI measurements



Fit-based uncertainty

Based on NOvA data

ND provides *signal*:
Sterile ν , cross section measurements

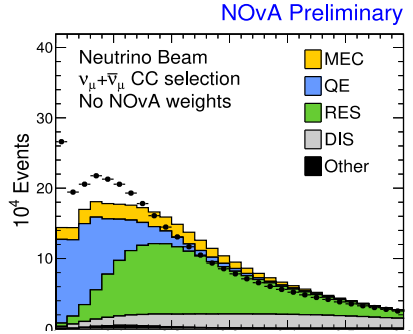


Model-spread uncertainty

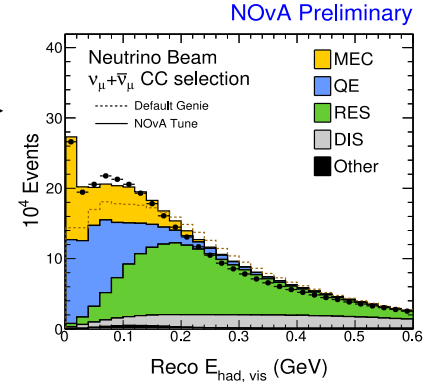
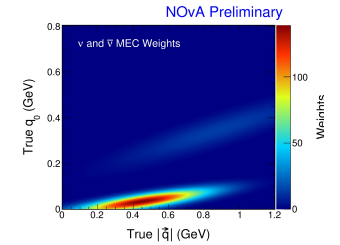
Based on external info

Multinucleon: fit-based unc.

Recall:
central value
fitted to ND data



Fitted double-Gaussian weights
applied to true CC MEC



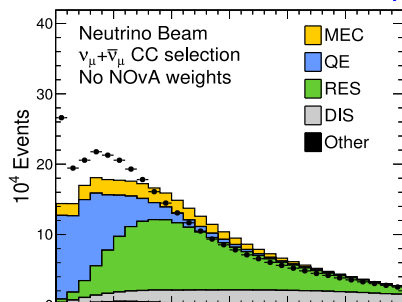
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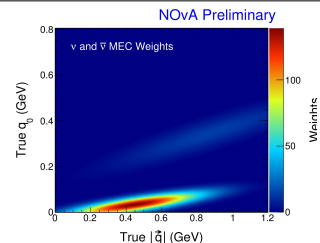
Re-perform fit,
with alternate
base models,
for uncertainties
(interpolate between these extremes)

Based on NOvA data

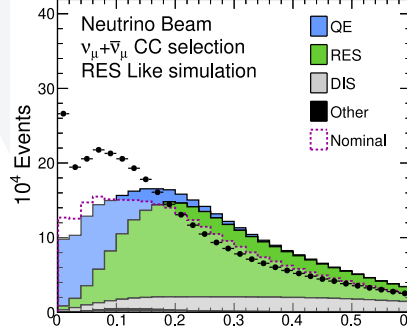
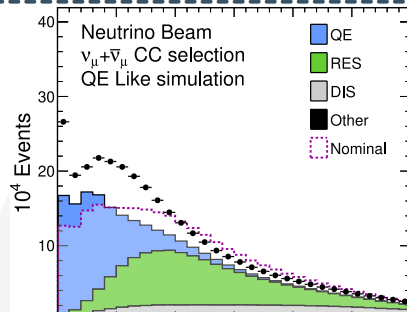
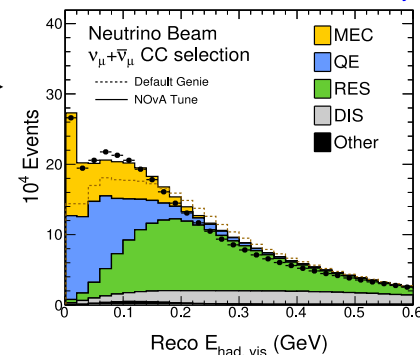
NOvA Preliminary



Fitted double-Gaussian weights
applied to true CC MEC



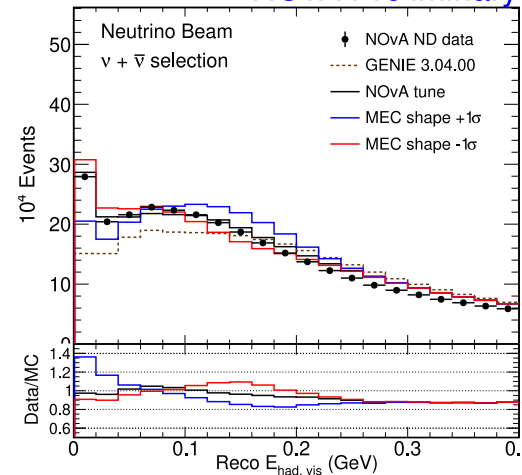
NOvA Preliminary



Re-fit using “extreme”
variations of QE, RES
systematics to bracket
impact on fitted MEC

(these alternate MEC
predictions used as
systematic uncertainties)

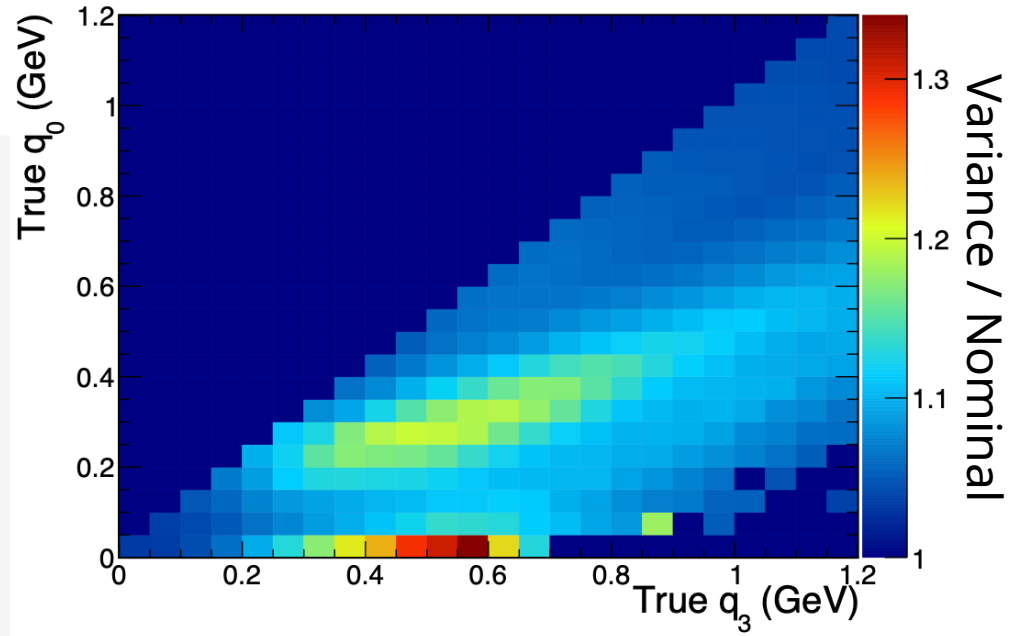
NOvA Preliminary



Multinucleon: model spread unc.

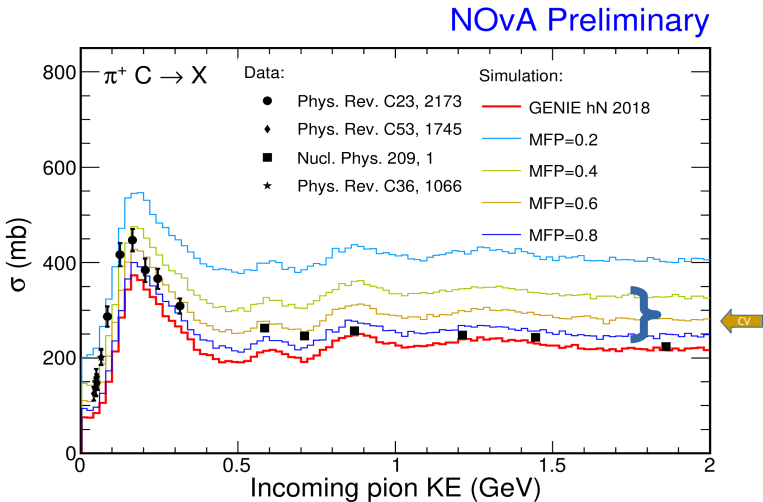
Based on external info

- Variance in $(q_0, |\mathbf{q}|)$ over 3 MEC models:
SuSA, Valencia, GENIE “Empirical”
- Result just touches NOvA ND data at “+1 σ ”
→ but note less flexibility in “moderate q_0 ” region compared to “fitted” unc. (previous)



FSI: uncertainties

Based on external info

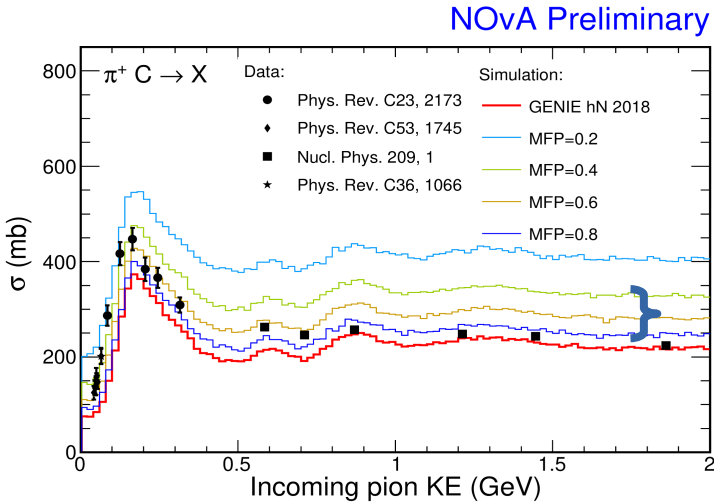


- It's straightforward to bracket
- 1 the **mean free path** variations:
 $f_{\text{MFP}}=0.4$ and $f_{\text{MFP}}=0.8$

Use same “BDT reweight” technique as CV to implement in analysis

FSI: uncertainties

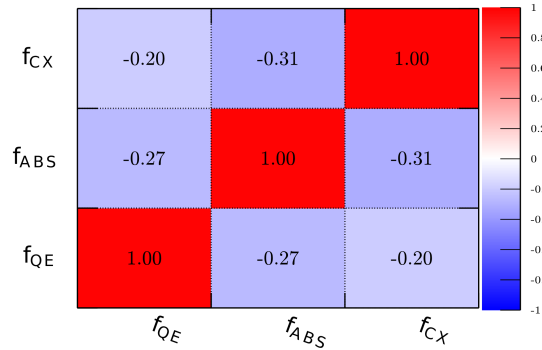
Based on external info



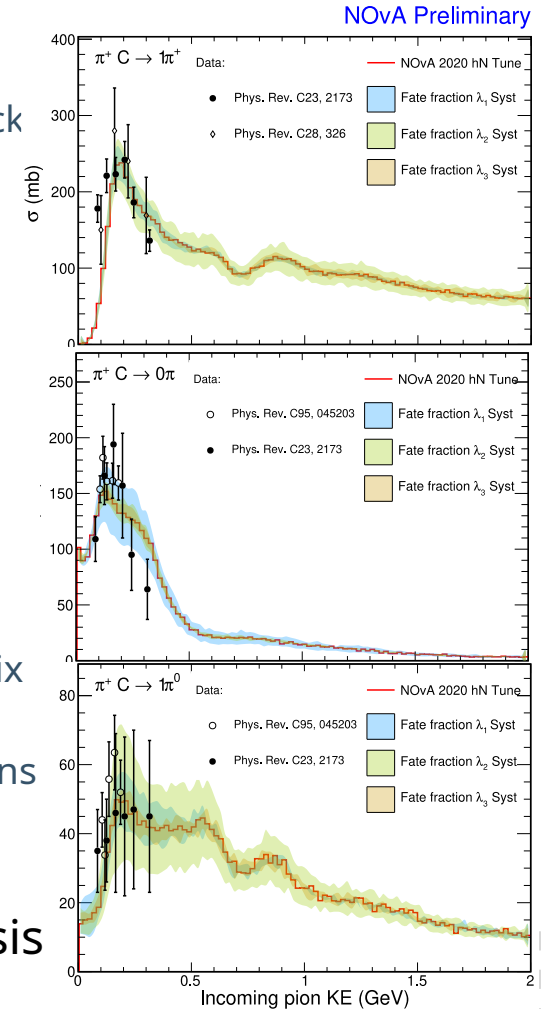
- ① It's straightforward to bracket the **mean free path** variations:
 $f_{MFP}=0.4$ and $f_{MFP}=0.8$

For the “**fate fractions**,” we piggyback on analogous work done by T2K:

[adapted from PRD 99, 052007]



- ② We diagonalize the covariance matrix they obtain from fitting world data to obtain 3 independent $\pm 1\sigma$ variations

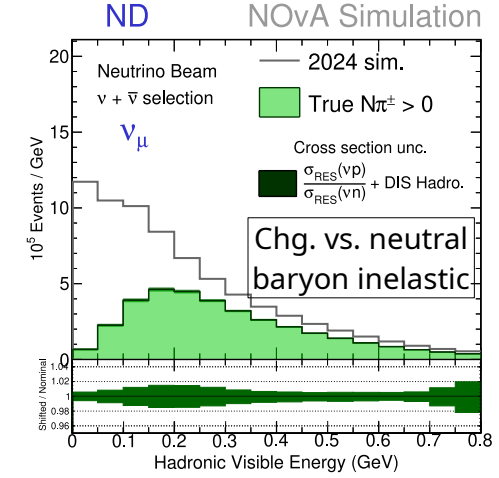
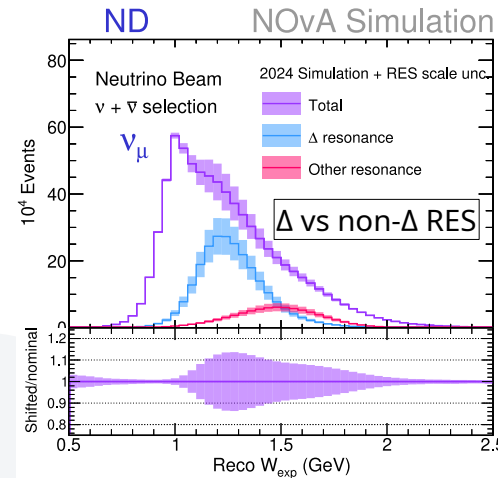


Use same “BDT reweight” technique as CV to implement in analysis

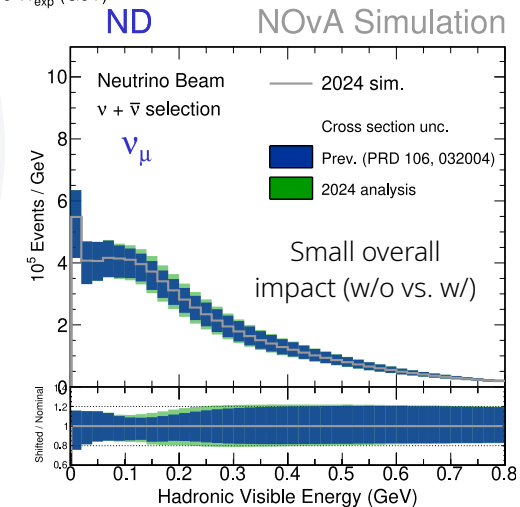
Pion production uncertainties

Two new categories of syst:

- Relative strength of Δ vs non- Δ resonant production (1 uncertainty)
 - Shift Δ and non- Δ resonances $\pm 20\%$ independently
 - Default GENIE: all resonances affected together
 - Overcounts Δ -specific uncertainty somewhat with other GENIE knobs, but “conservative”
- Charged vs. neutral baryon production in RES, DIS (2 uncertainties)
 - Shift ratio of RES Δ -channel proton/neutron final states by $\pm 5\%$
 - Shift composition of proton/neutron final states in DIS
- Moderate impacts on pion-rich subsamples, but overall effect on uncertainty budget is minor



Based on external info

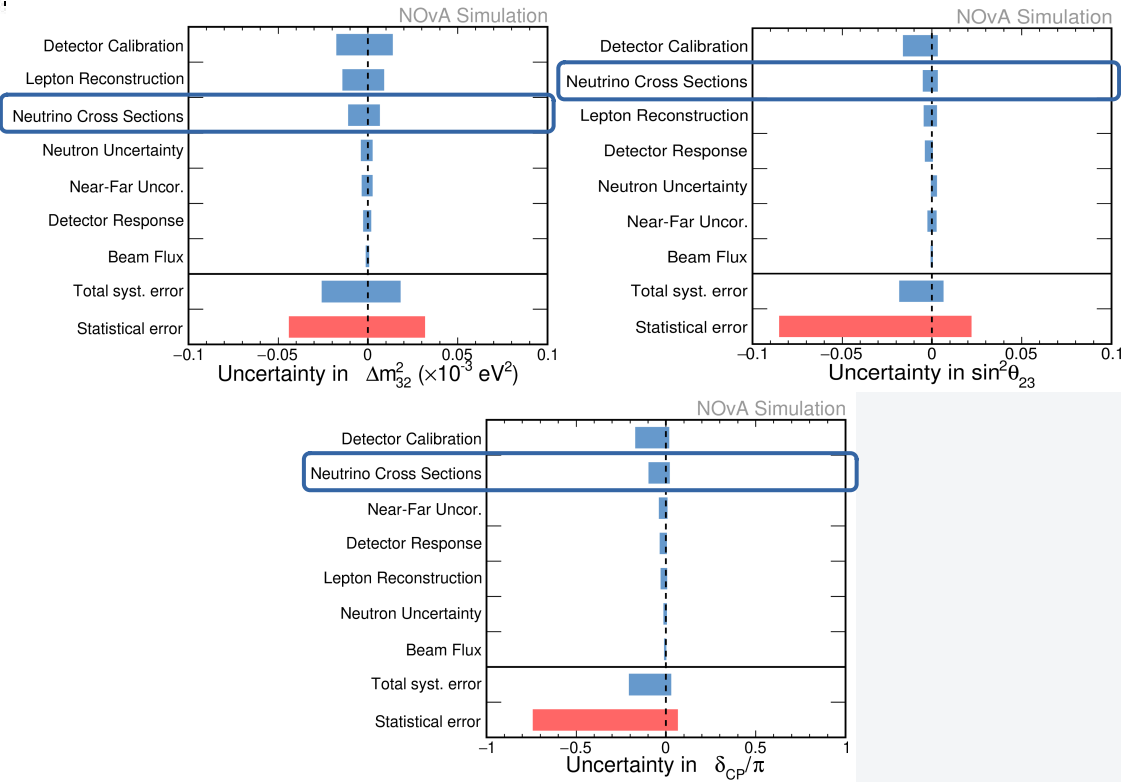


[see **Neutrino '24 poster** by M. Dolce & M. Martinez-Casales for more]

What matters the most?



What matters the most?

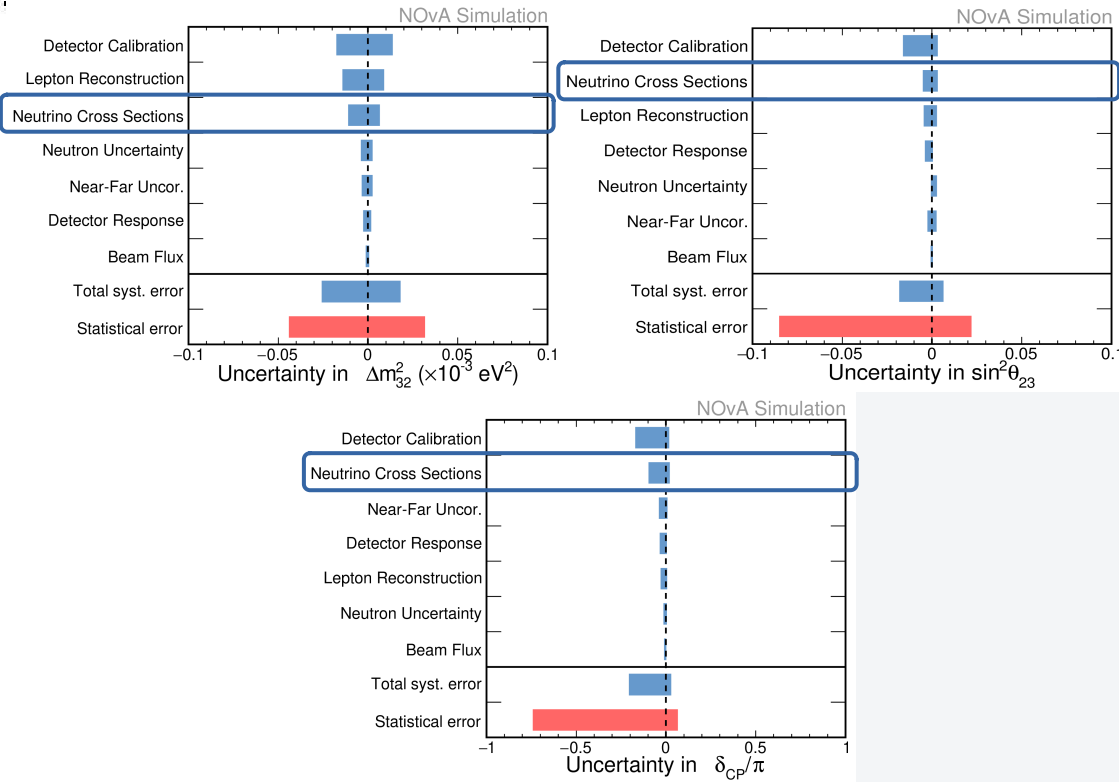


The **largest contributors** are not surprising:

- CCQE z-expansion FF
- CCQE RPA
- CC RES M_A and M_V
- ν_e / ν_μ (unconstrained by ND)

① For **3-flavor oscillations** (and NSI),
 ν interaction uncertainties do not dominate
(and we don't expect this to change even when data-taking is completed)

What matters the most?



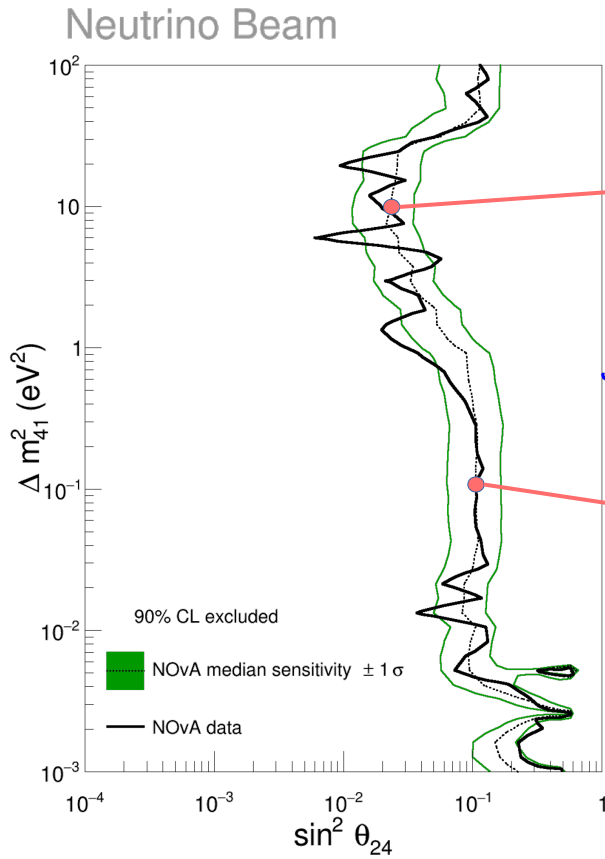
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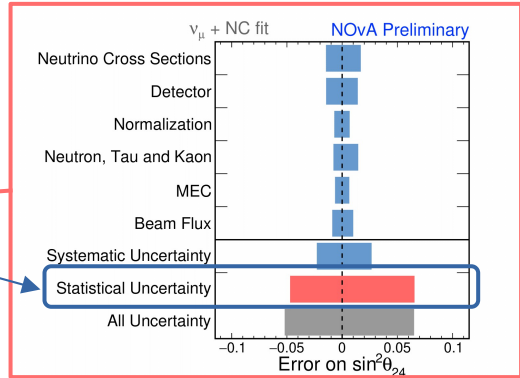
Recent theory work has concentrated on QE, but only ~20% of NOvA sample :(

① For **3-flavor oscillations** (and NSI), **ν interaction uncertainties do not dominate** (and we don't expect this to change even when data-taking is completed)

What matters the most?

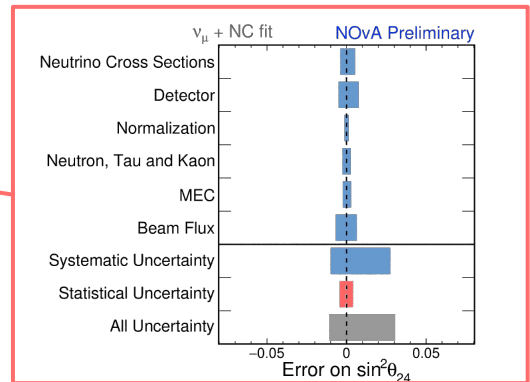


At high Δm^2 , sensitivity driven by FD (low stats), thus stat. unc. dominates



At low Δm^2 , sensitivity driven by ND (large stats), but syst. dependence is spread over many types of syst.

(+add CCMEC, NC to significant xsec systs)



② For sterile ν searches, the situation is more interesting, but xsec systs still are not dominant

What matters the most?

	ν_μ CC <small>[PRD 107, 052011]</small>	$\bar{\nu}_\mu$ CC <small>[FNAL W&C seminar]</small>	ν_e CC <small>[PRL 130, 051802]</small>	$ q -E_{\text{avail}}$ <small>[arXiv:2410.05526]</small>	Low E_{had} <small>[FNAL W&C seminar]</small>	ν_μ CC π^0 <small>[PRD 107, 112008]</small>
Flux	9.1	9.9	10.3	11.4	~11%	8.3
E-Scale + Det Model	6.1	6.1	8.6	3.8	~5-6%	7.6
Cross Section Model	1.9	2.5	9.8	5.6	~2%	4.6
Neutron Modeling	1.5	2.3			~3%	
Statistical			7.4			
2p2h Model				7.1		
Pi Charge Exchange						3.8

Predominantly ν_μ CC π^0 channels (major bkgd) and ν_e MEC (in signal; not called out separately)

Observables deliberately in hadronic system, thus inevitable

FSI is major production mechanism for π^0

For **neutrino scattering** measurements,
3 **xsec syts** are sometimes important,
 but *which depends a lot on the channel*



Future directions

2-detector fitting

FERMILAB-TD938-2023-04

Constraining neutrino oscillation and interaction parameters with the NOvA Near Detector and Far Detector data using Markov Chain Monte Carlo

A dissertation submitted by
Michael Dolce
in partial fulfillment of the requirements for the degree of
Doctor of Philosophy
in
Physics
Tufts University
August 2023
Adviser: Prof. Hugh Gallagher

[M. Dolce, Tufts Univ., 2023]

Constraining neutrino interaction uncertainties for oscillation measurements in the NOvA experiment using Near Detector data

by
Maria Martinez Casales

A dissertation submitted to the graduate faculty
in partial fulfillment of the requirements for the degree of
DOCTOR OF PHILOSOPHY

Major: High Energy Physics

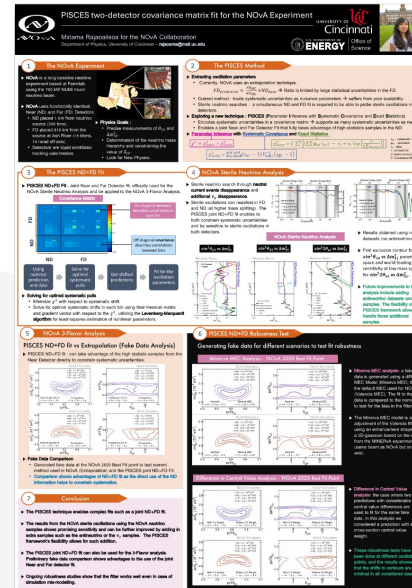
Program of Study Committee:
Mayly Sanchez, Co-major Professor
Matthew Weirum, Co-major Professor
Rebecca Flint
Soren A. Prell
Kiriell Tuckin
Amanda Weinstein

The student author, whose presentation of the scholarship herein was approved by the program of study committee, is solely responsible for the content of this dissertation. The Graduate College will ensure this dissertation is globally accessible and will not permit alterations after a degree is conferred.

Iowa State University
Ames, Iowa
2023

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[M. Martinez Casales, Iowa State U., 2023]



[M. Rajaoalisoa, Neutrino '24 poster]

Not being able to depend on “extrapolation” forcing us to grapple with **need for more exhaustive ν int systematics**

- Extending work begun for sterile ν search & xsec measurements (model spread MEC)
- New π -production systems emerged from this work

Strong collaboration interest in **exploring simultaneous 2-det fits** in lieu of “extrapolation” for 3-flavor oscillations

GENIE “AR23” and ‘nusystematics’

- Recent flurry of effort from DUNE/SBN to produce models usable in multiple expt contexts
 - GENIE configuration AR23_20i_00_000 encapsulates base model choices
 - `nusystematics` contains publicly-accessible cross section uncertainty knobs (like `NOvARwgt`)
- We are seriously exploring leveraging this effort for our last major simulation production campaign
 - Base model has desirable flexibility for *post hoc* reweighting
 - Many systematics are in same spirit as, but more sophisticated than, our custom knobs
- Some technical hurdles to overcome
 - Legacy choices about MC format & persistence from 10+ years ago make this ... creaky, but doable with effort



Summary thoughts

NOvA experience with xsec systs

- “Functionally equivalent” detector design & “extrapolation” paradigm a powerful combination
 - Blunted impact of intrigue over 1p1h & 2p2h
 - Enabled *staged* improvements to the model with limited personnel
 - Even at full exposure don’t anticipate xsec systs limiting 3-flavor results
- “Owning” complete xsec model requires critical mass of interested generator experts
 - Our targeted “bespoke” uncertainties approach only possible against the larger backdrop of centralized GENIE library
 - We have profited extensively from cross-collaboration ... collaboration (MINERvA, T2K, DUNE)
- Effort required to convert non-reweightable to reweightable worth the investment
 - Expanded usable model space in regions important to NOvA (π FSI, hadronization)
 - Interesting intellectual / technical challenge, fun for students / postdocs

Lessons for future experiments

- Experiment & analysis design sculpt uncertainty needs
 - Different weaknesses of global cross section picture are thrown into relief at 600 MeV vs. 2 GeV vs. 5+ GeV
 - “Extrapolation” vs. “simultaneous 2-detector fit” (vs. PRISM!) impose different demands on model
- Real data is a blessing and a curse
 - Don't always have the luxury of fully rigorous model-based uncertainties
 - Judicious “blends” of model- and data-based uncertainties can be effective stand-ins when all else fails
- Technical model synergy with the wider community is extremely profitable, but requires some “homework”
 - Choices made in early stages of persistence & large-scale simulation infrastructure can later become obstacles

Overflow slides follow

BDT reweights

We use a boosted decision tree technique....

$$f_{\text{BDT}} = \alpha_1 \text{ (tree 1)} + \dots + \alpha_N \text{ (tree N)}$$

$$\approx \frac{1}{N} \sum_{\text{tree } i=1}^N \alpha_i \Theta \left(\underset{\substack{\uparrow \\ \text{event} \\ \text{values}}}{\vec{X}} - \underset{\substack{\uparrow \\ \text{trained} \\ \text{"cut"} \\ \text{for} \\ \text{tree } i}}{\vec{X}_0^i} \right)$$

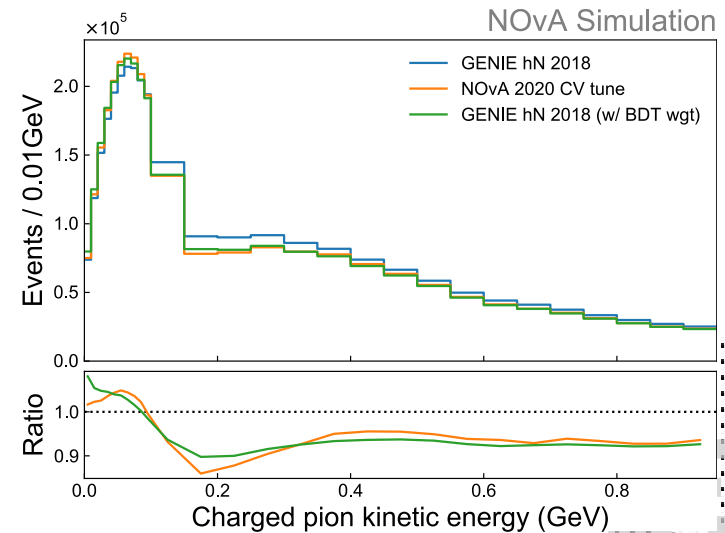
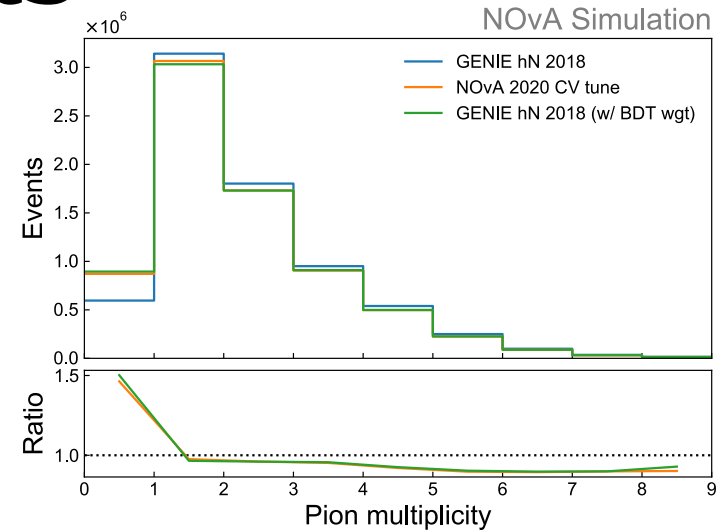
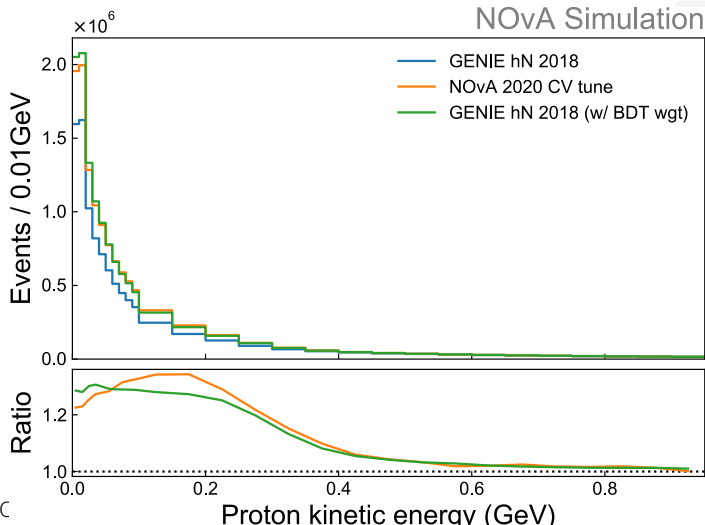
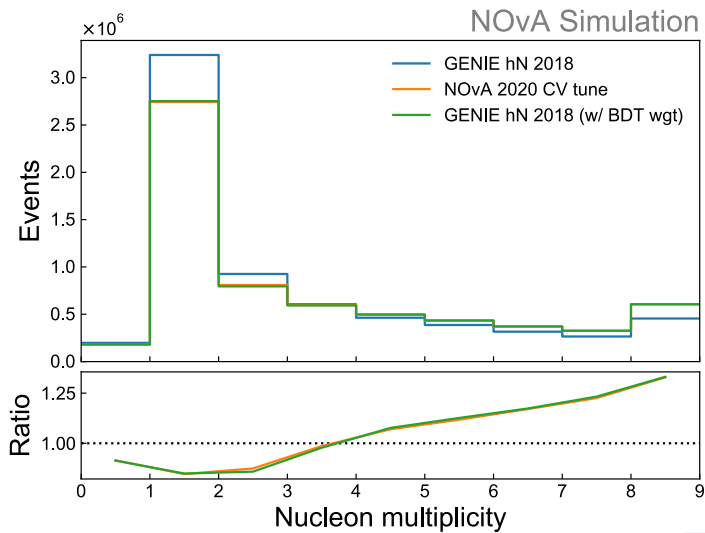
BDT trained using binary logistic loss
to distinguish between simulated GENIE CV
and desired "alternate" sample

... to create reweights
from the nominal to the
uncertainty samples

$$w(\vec{X}) = \frac{f_{\text{BDT}}(\vec{X})}{1 - f_{\text{BDT}}(\vec{X})}$$

(detailed discussion of technique
in NuSTEC CTGWG Seminar [Dec. 14, 2022](#))

BDT reweights



While not perfect,
performance is
adequate to use
as a computing
time-saver