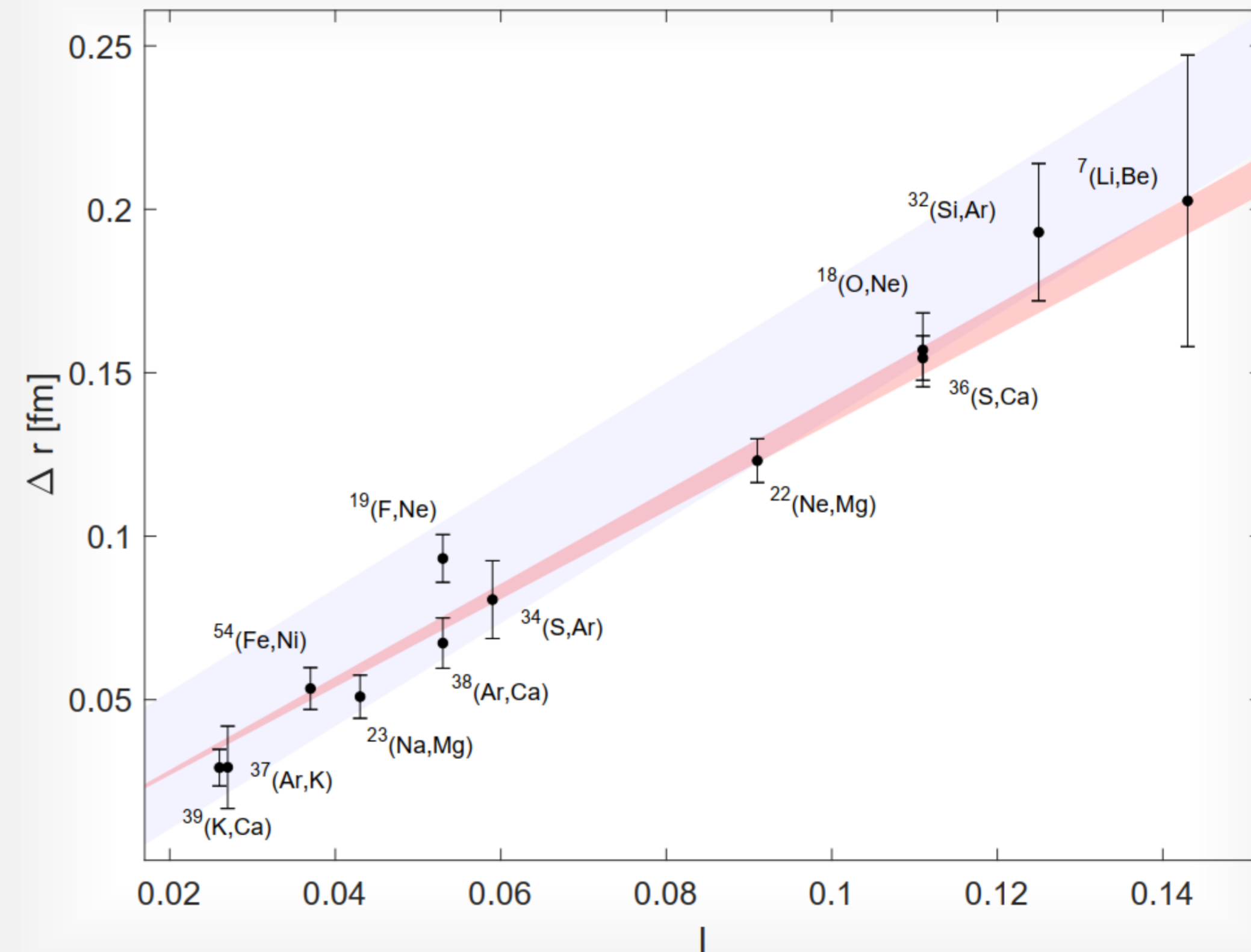
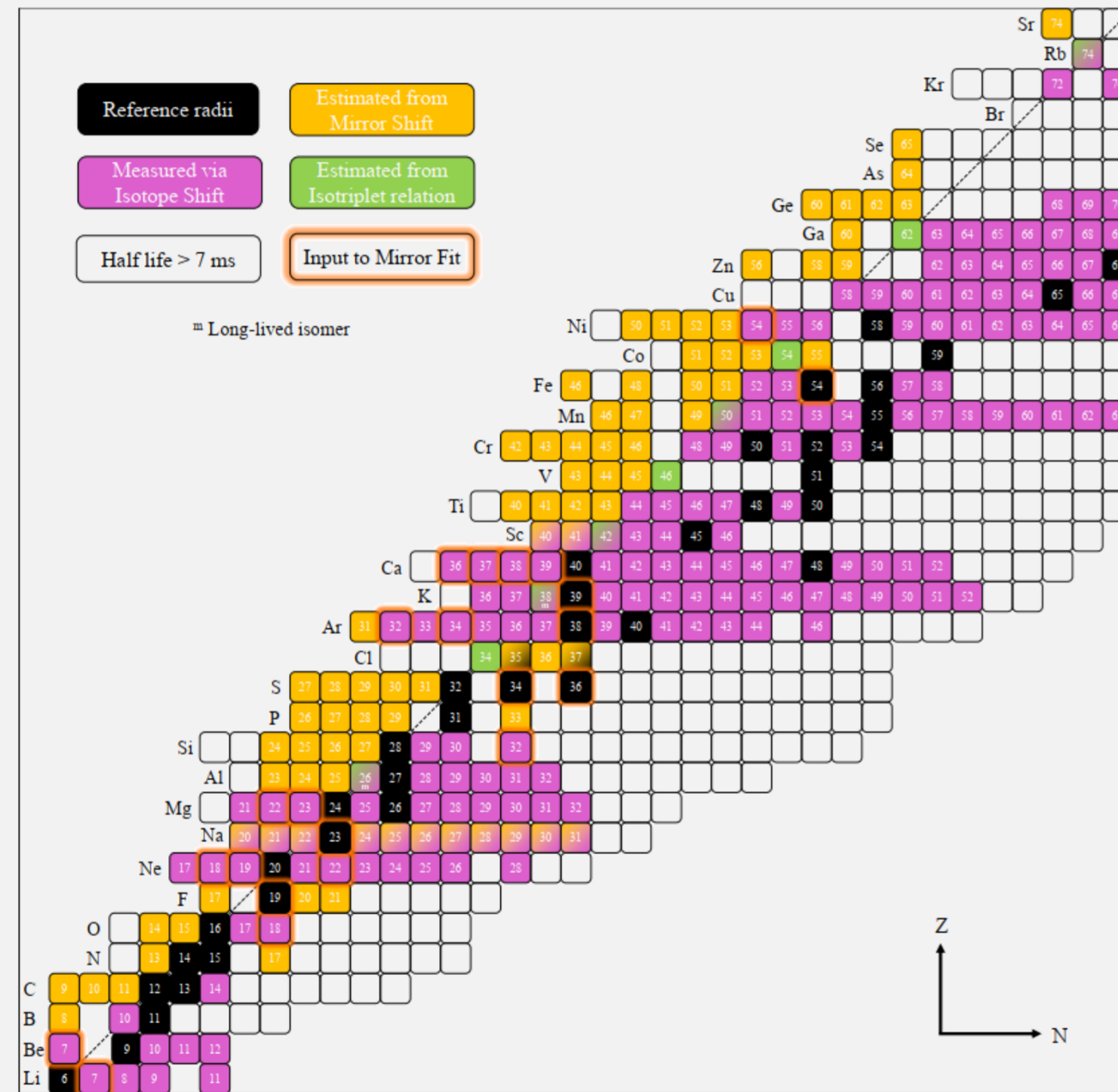


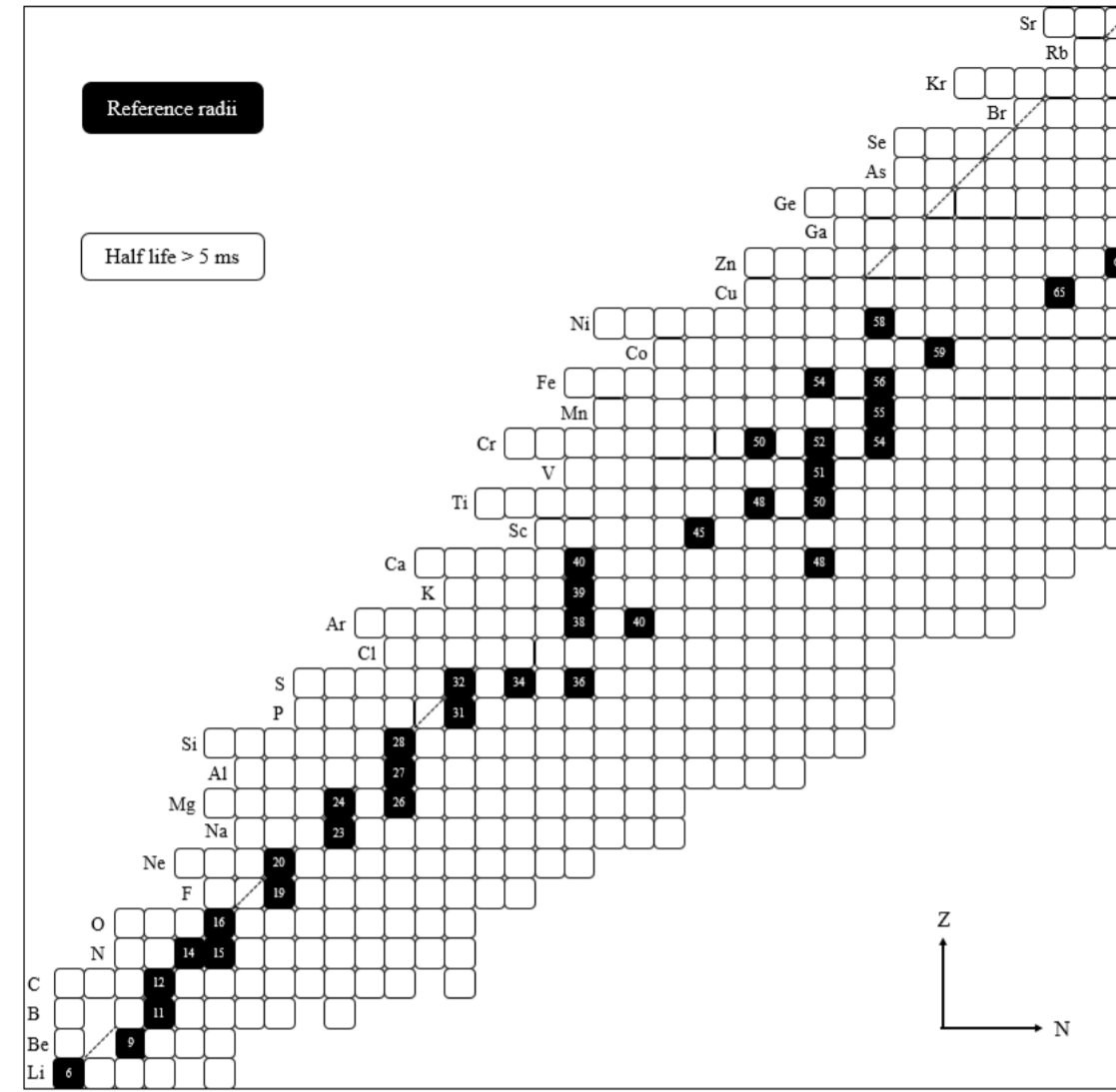
# Radii of Light Mirror Nuclei

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)



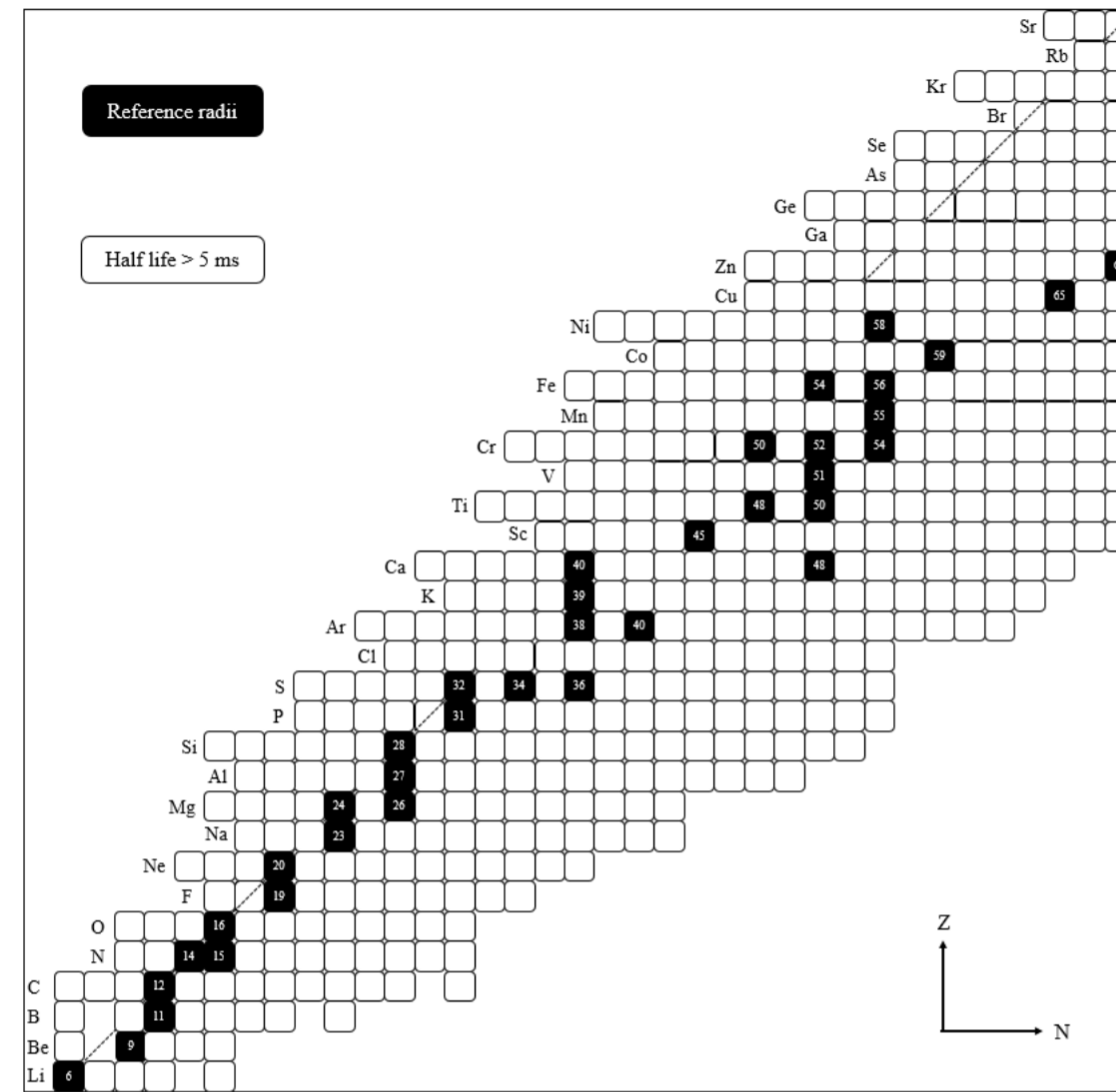
# Roadmap:

## 1. Reference radii (muonic atoms)

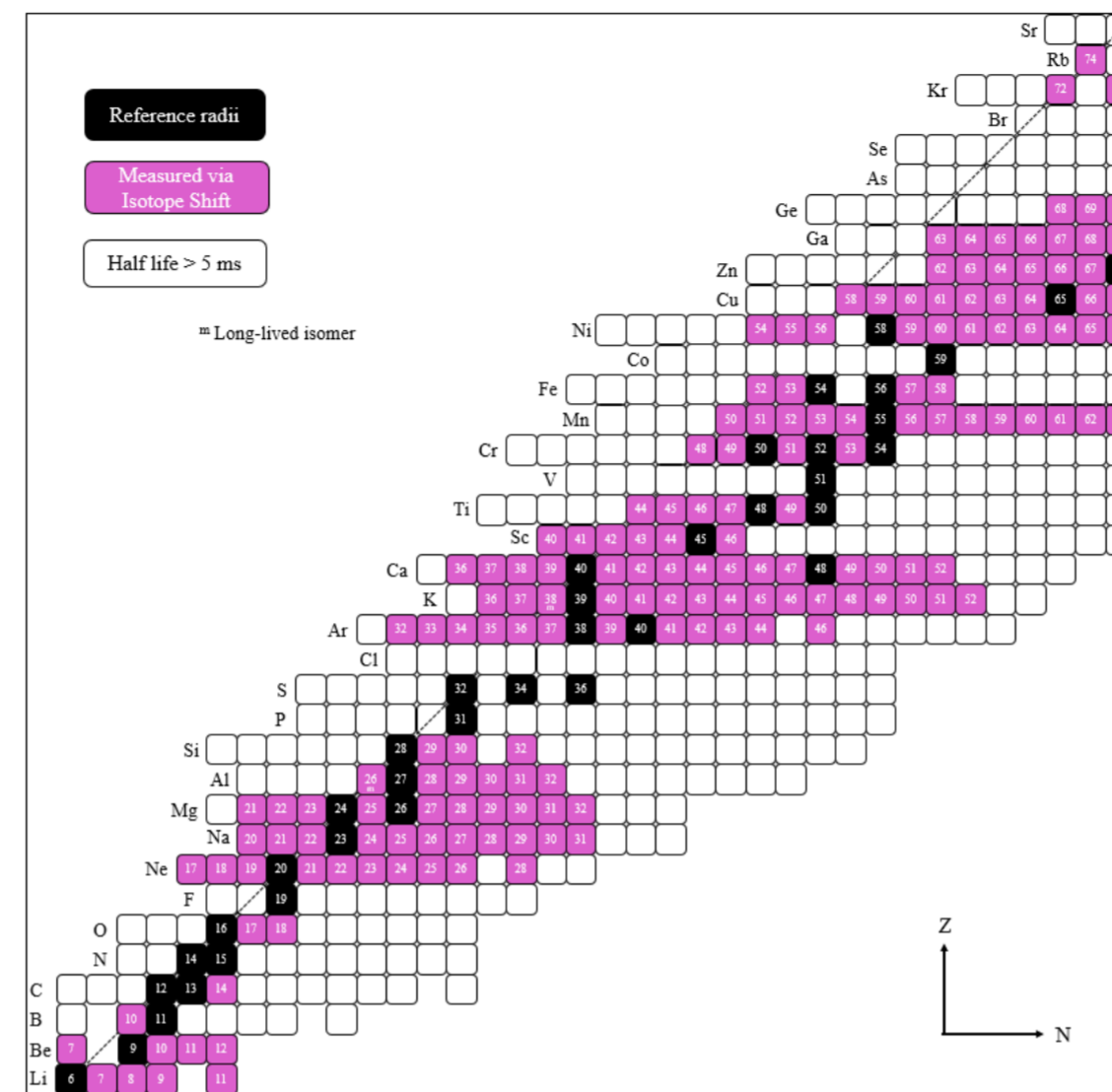


# Roadmap:

## 1. Reference radii (muonic atoms)

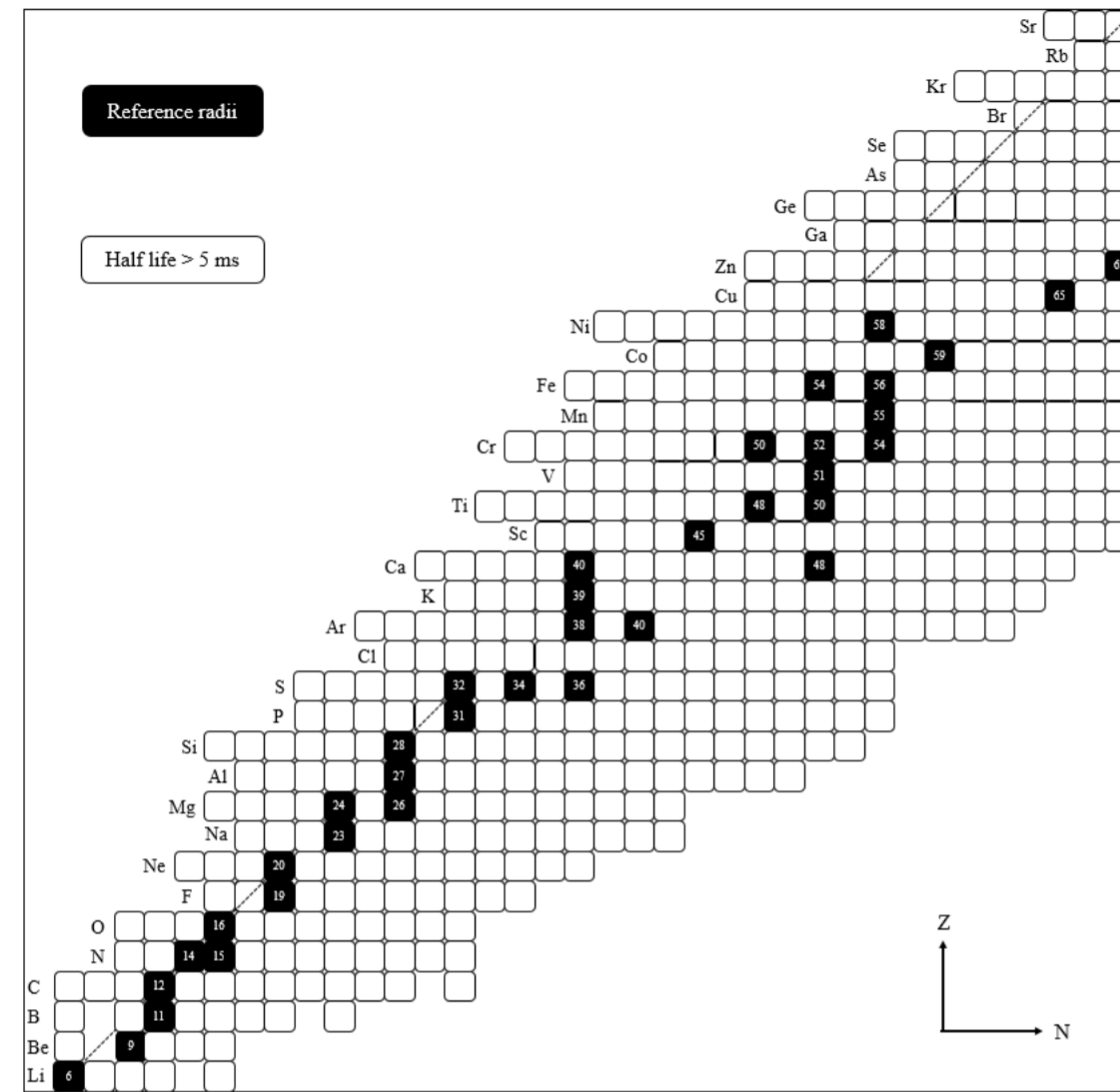


## 2. Isotope shifts (electronic atoms)

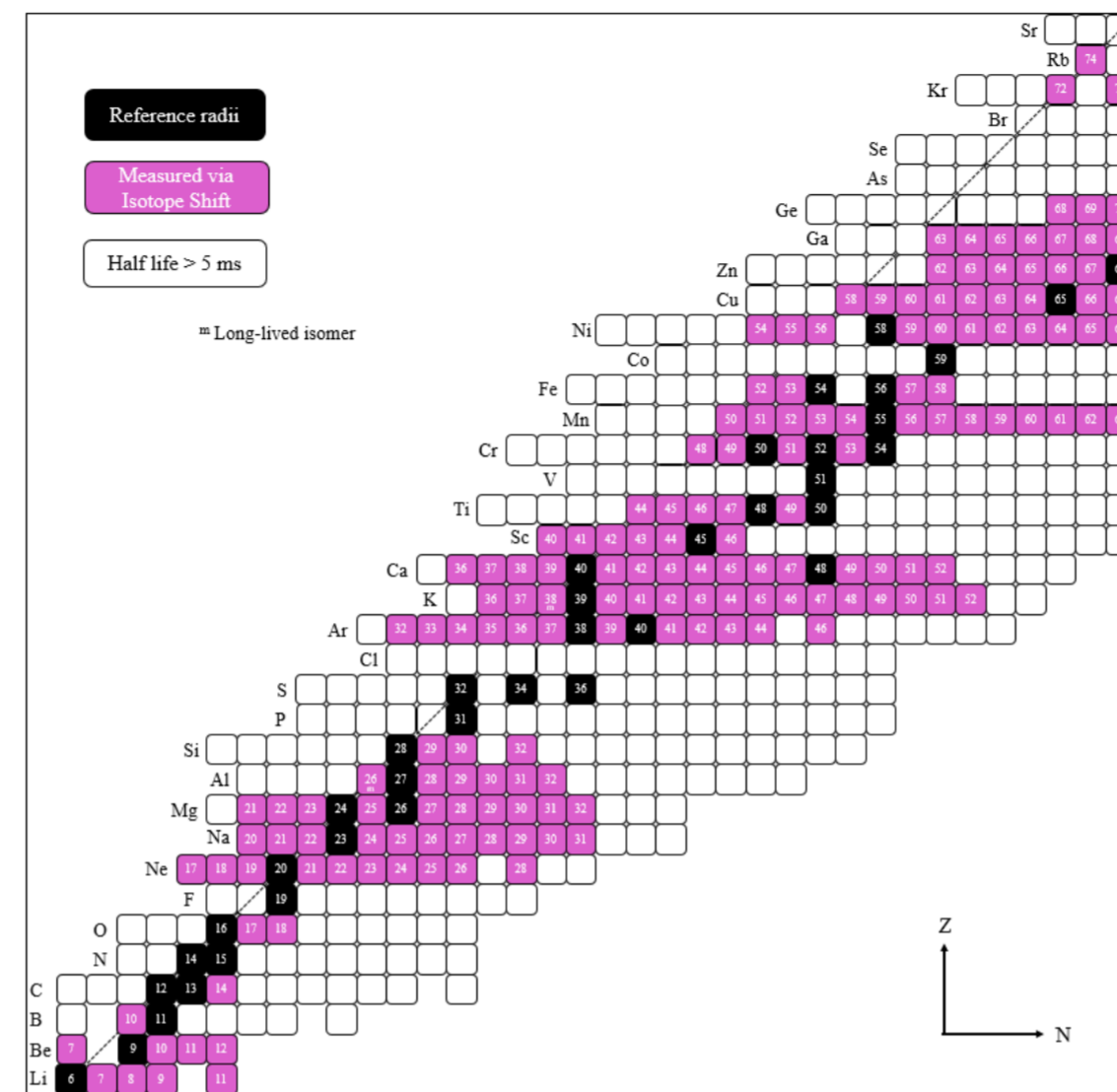


# Roadmap:

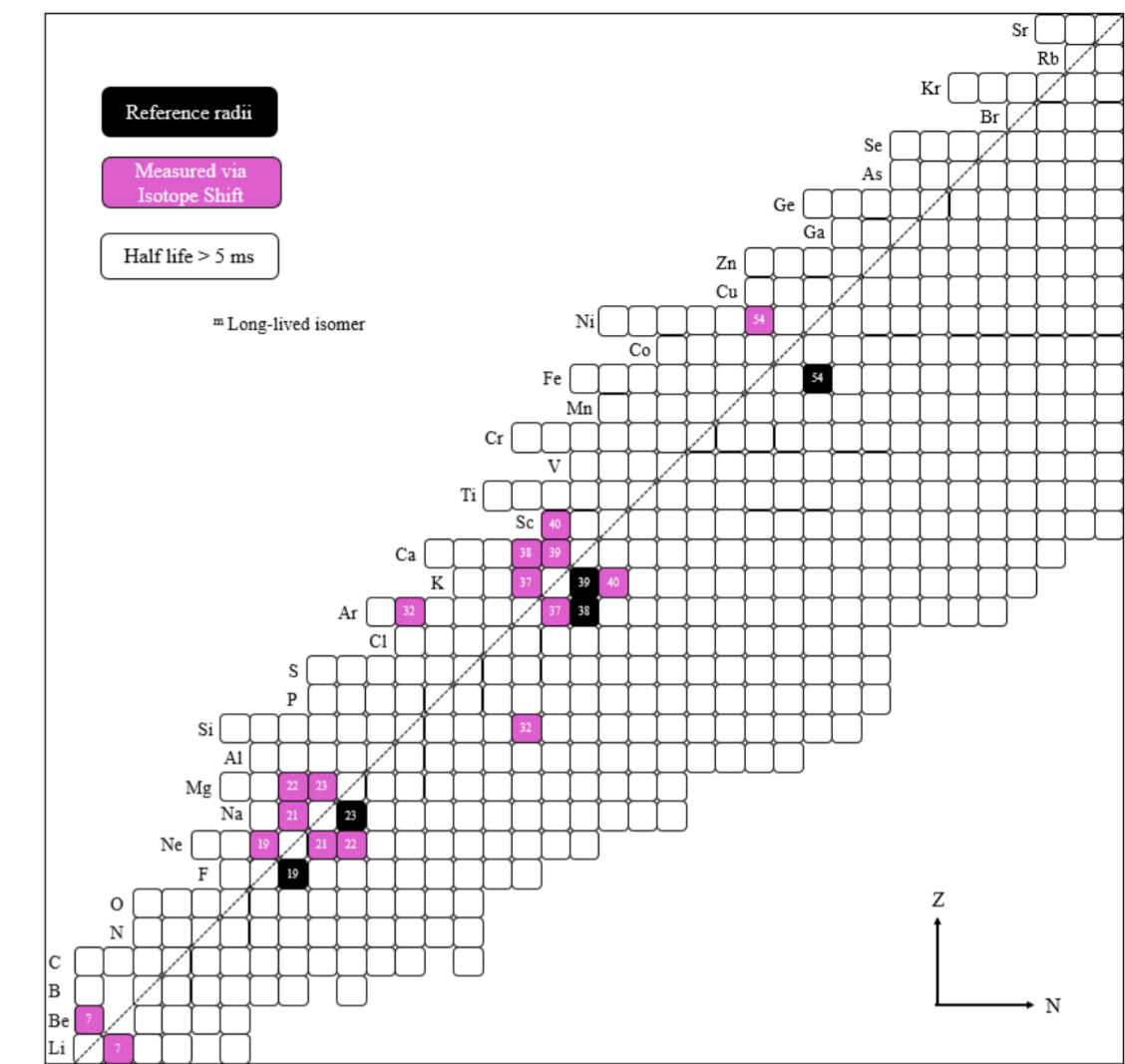
## 1. Reference radii (muonic atoms)



## 2. Isotope shifts (electronic atoms)

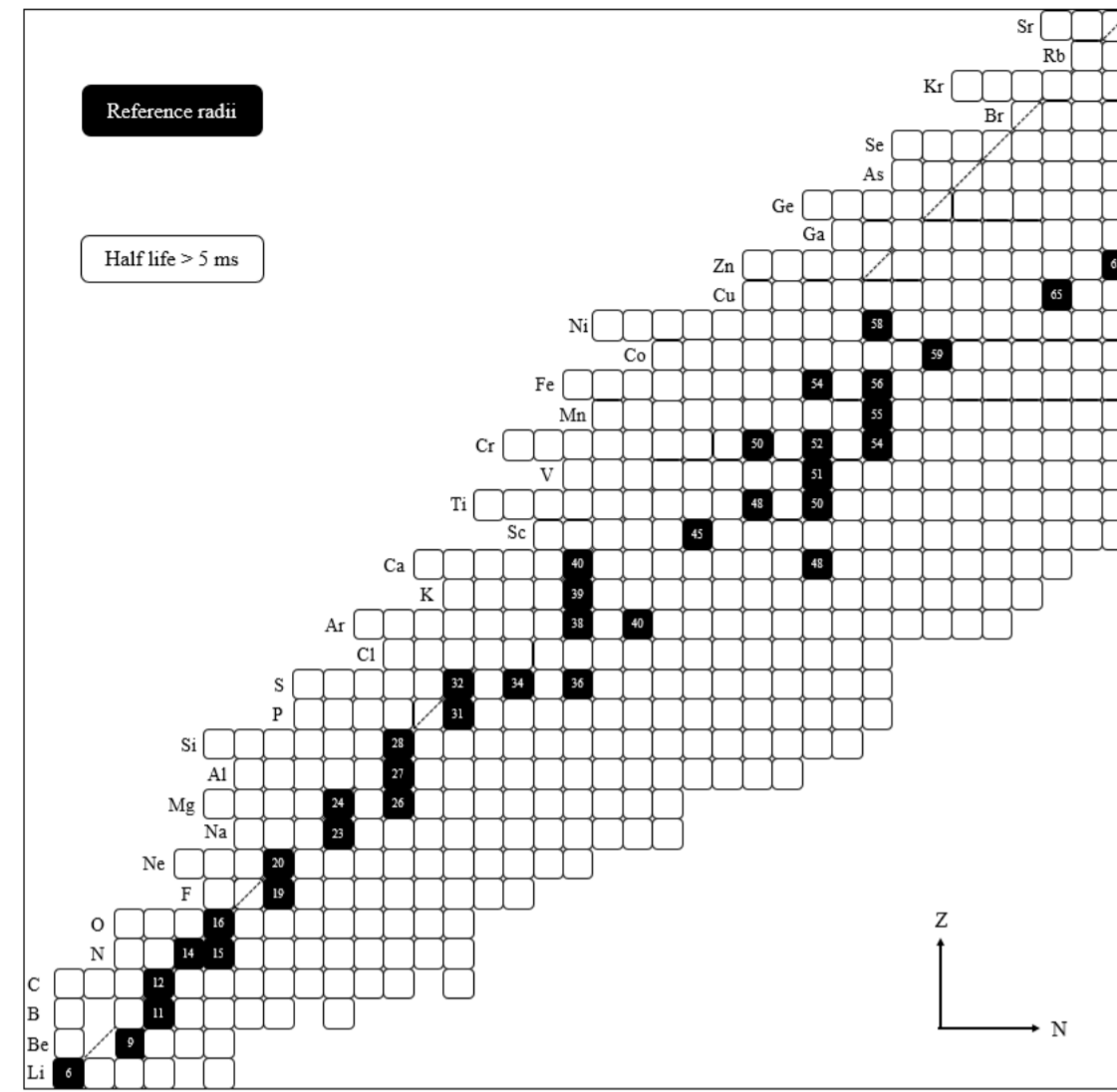


## 3. Mirror nuclei

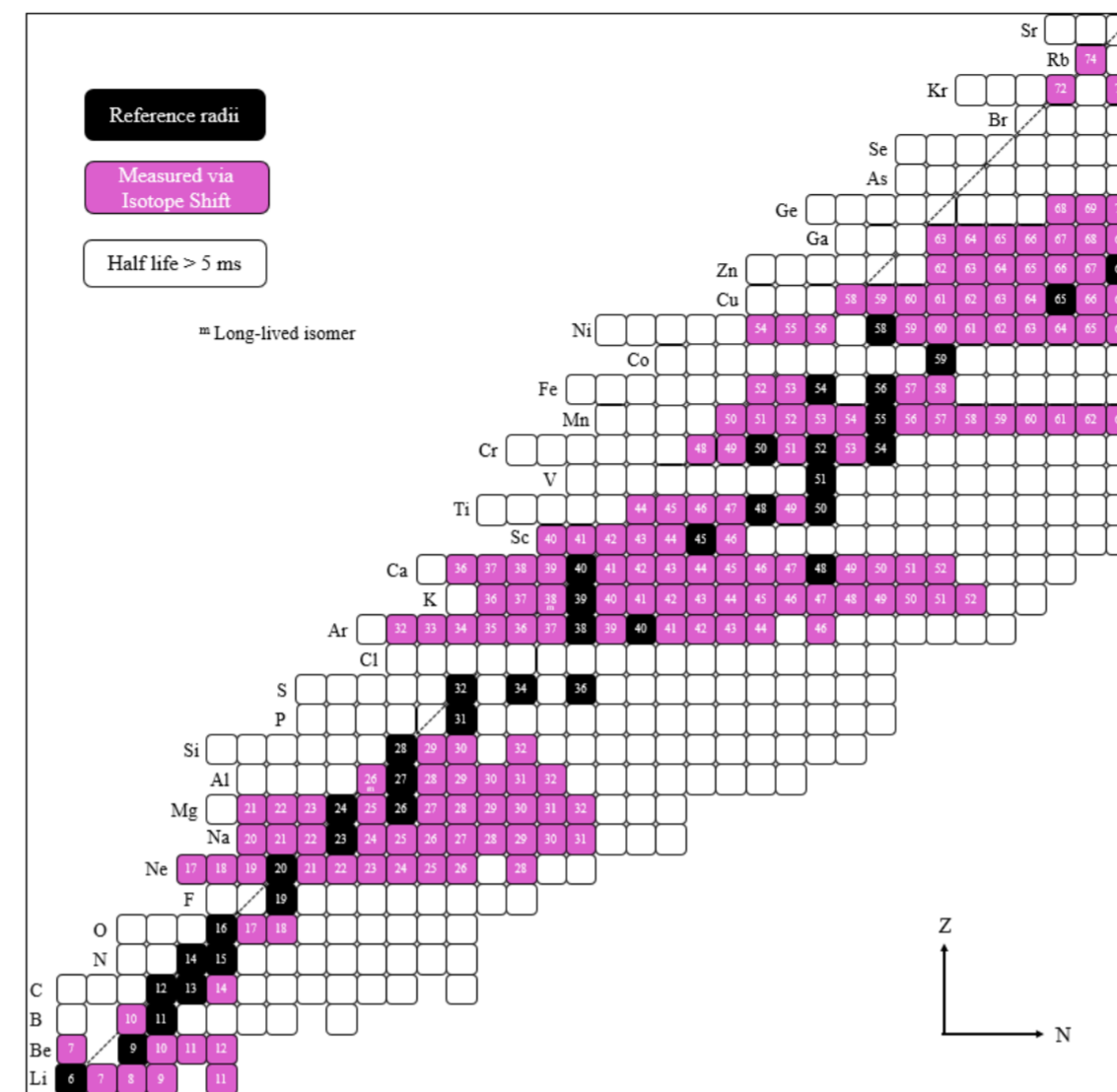


# Roadmap:

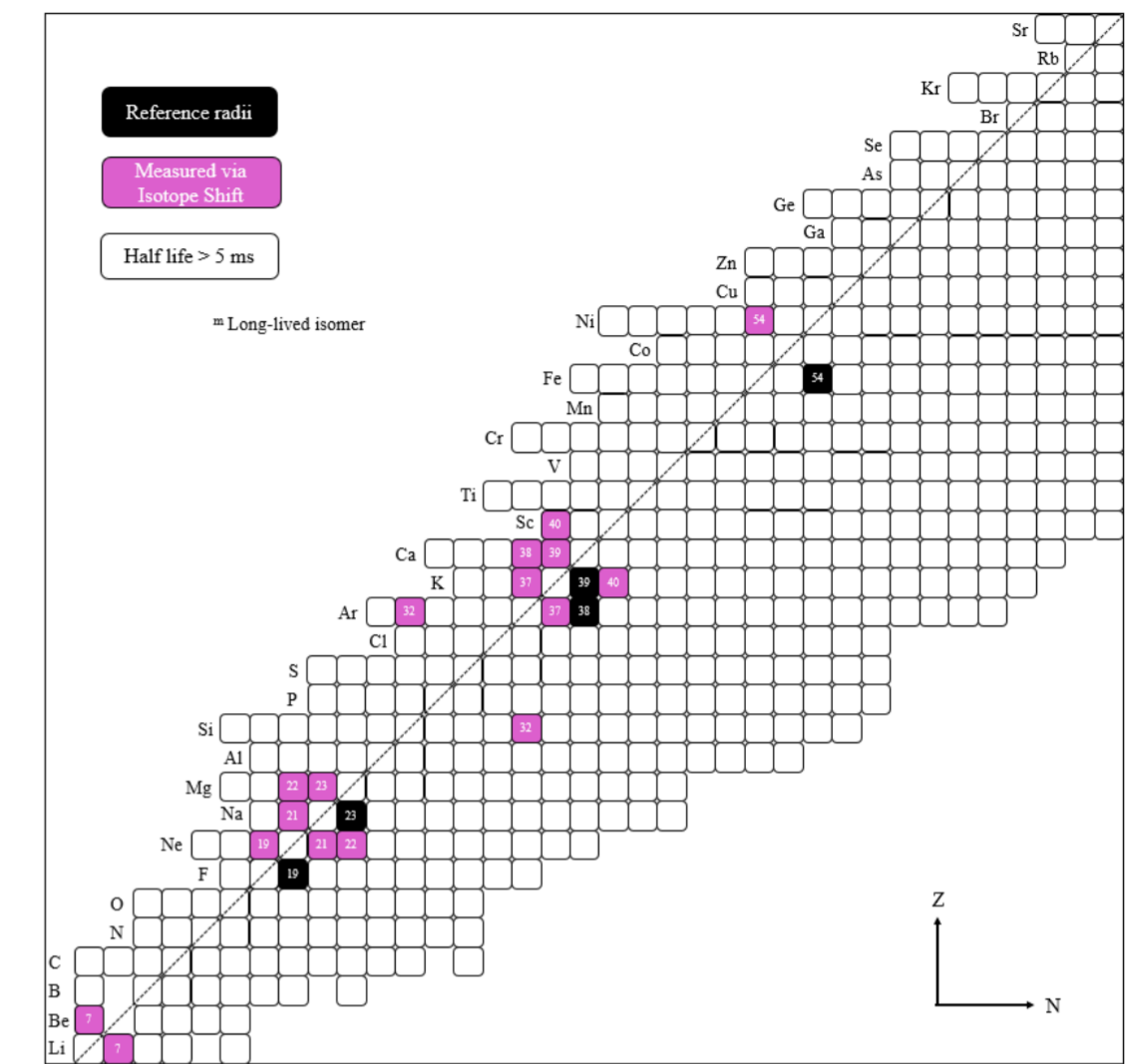
## 1. Reference radii (muonic atoms)



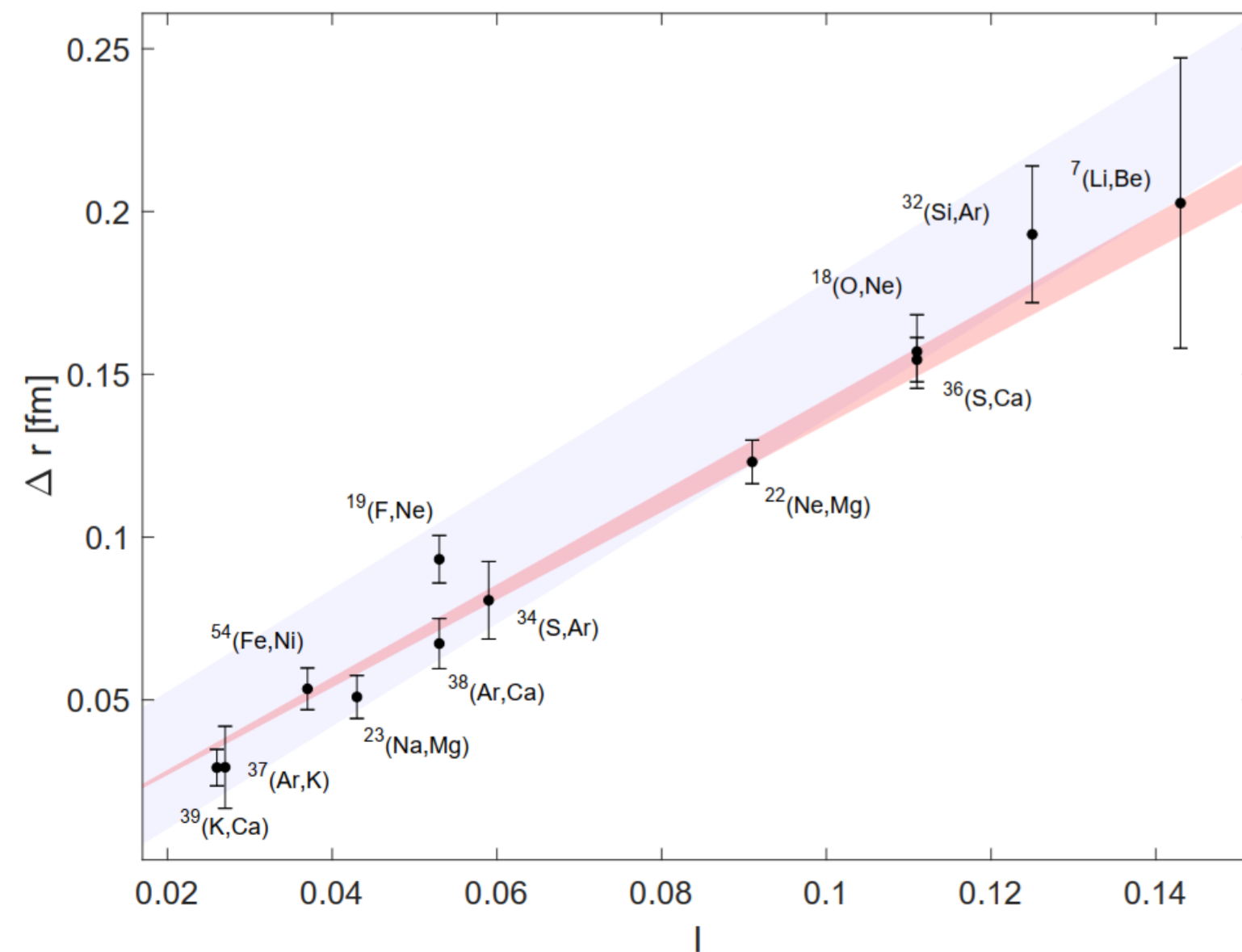
## 2. Isotope shifts (electronic atoms)



## 3. Mirror nuclei

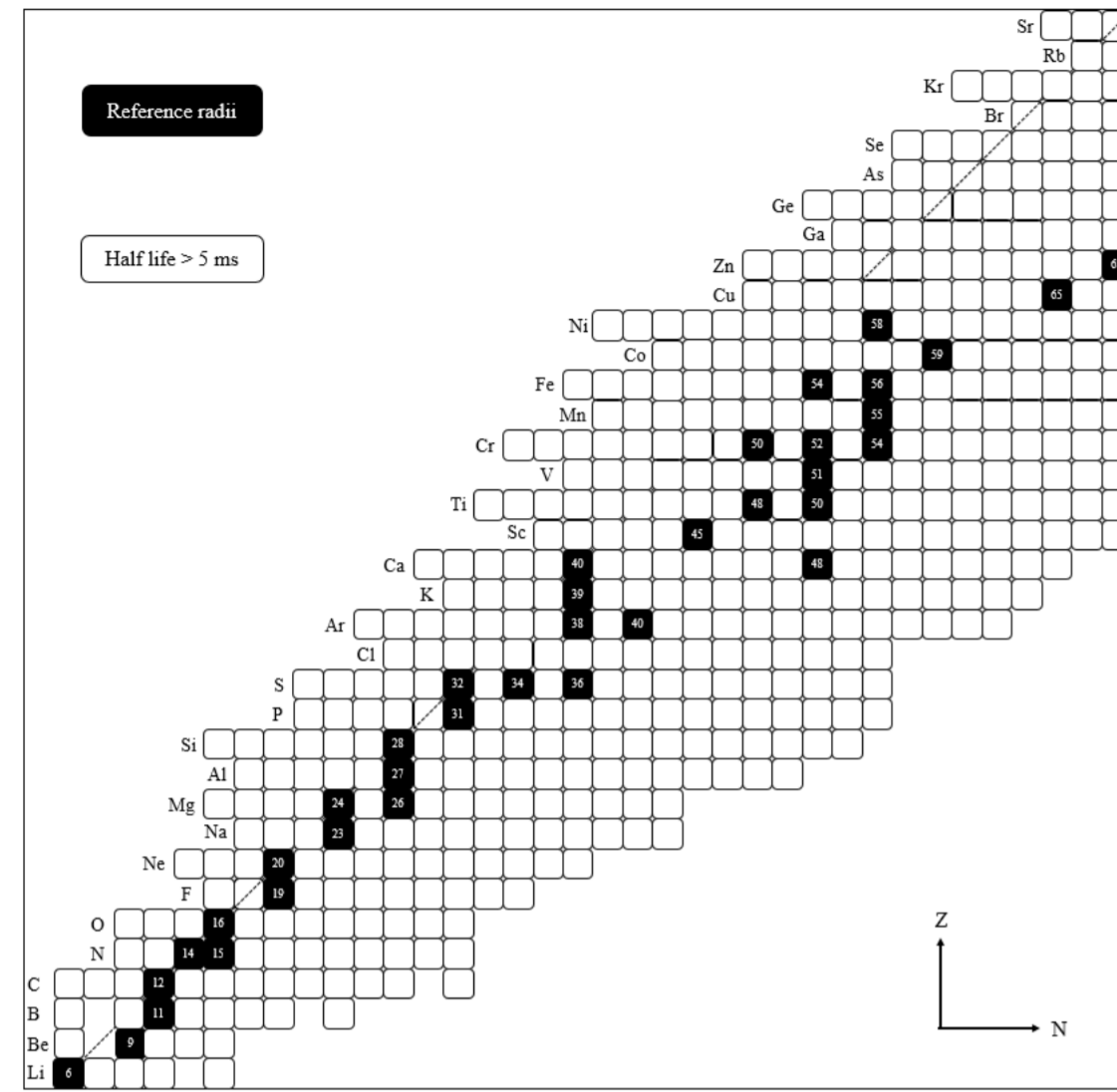


## 4. Mirror fit

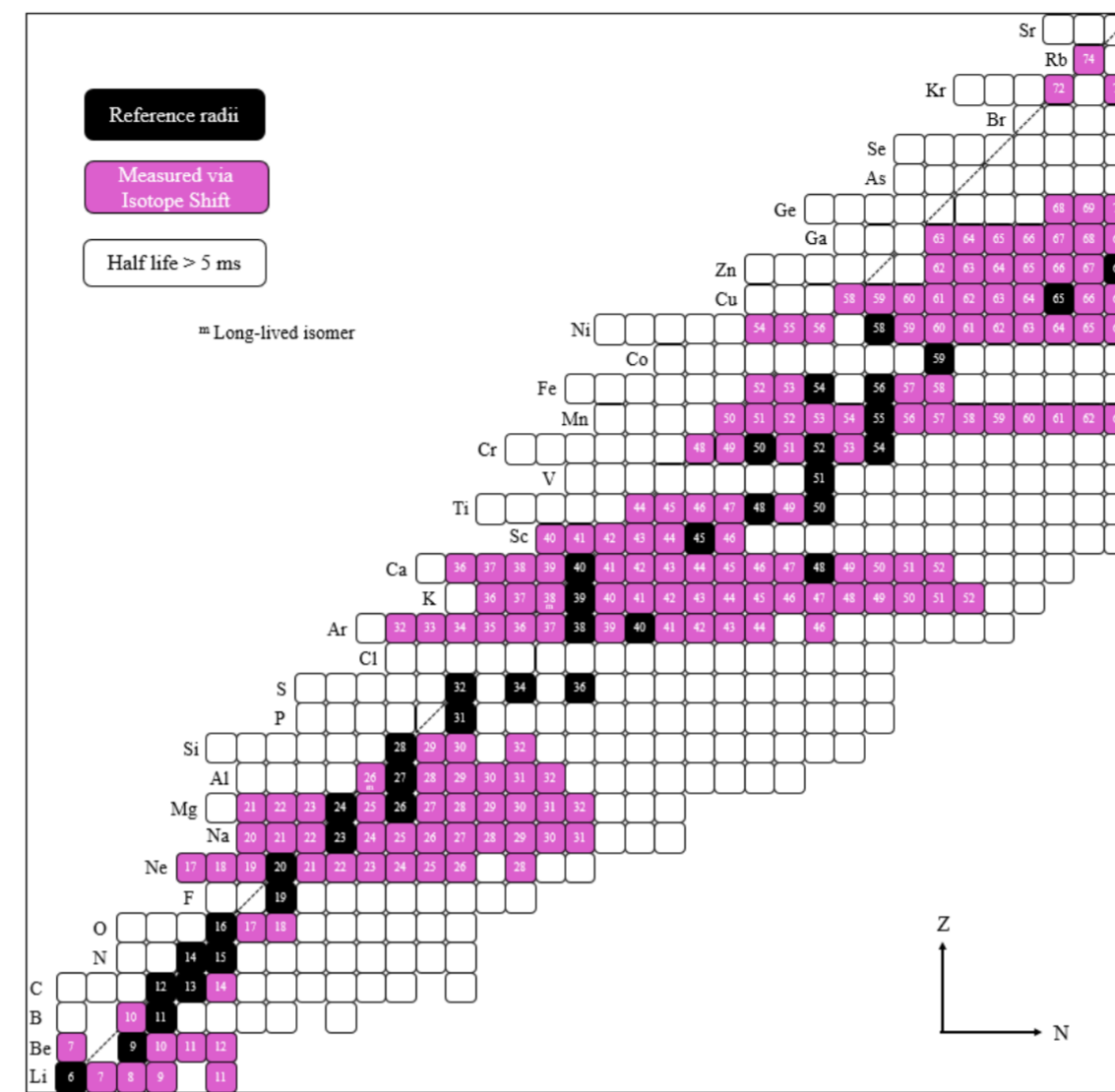


# Roadmap:

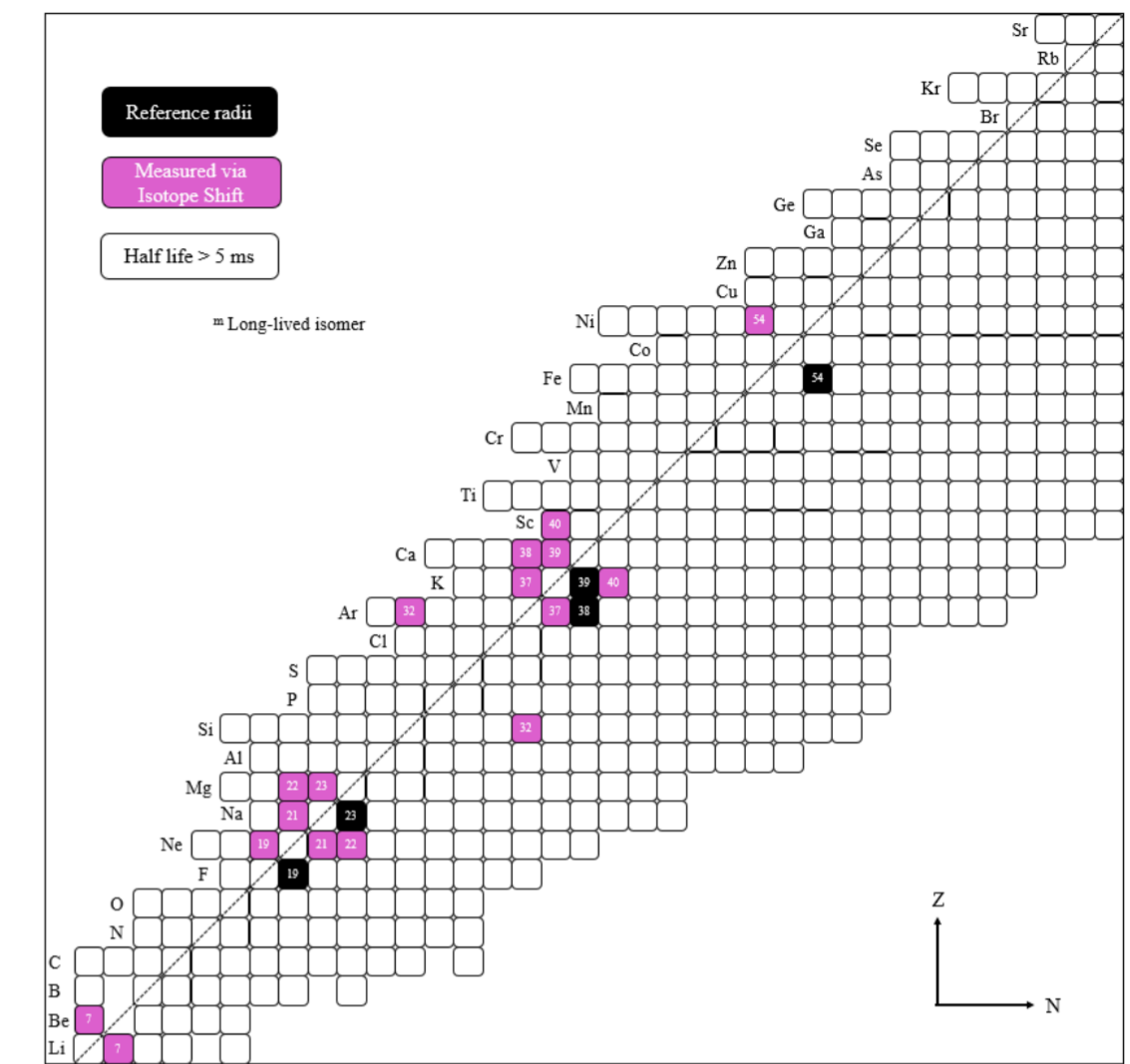
## 1. Reference radii (muonic atoms)



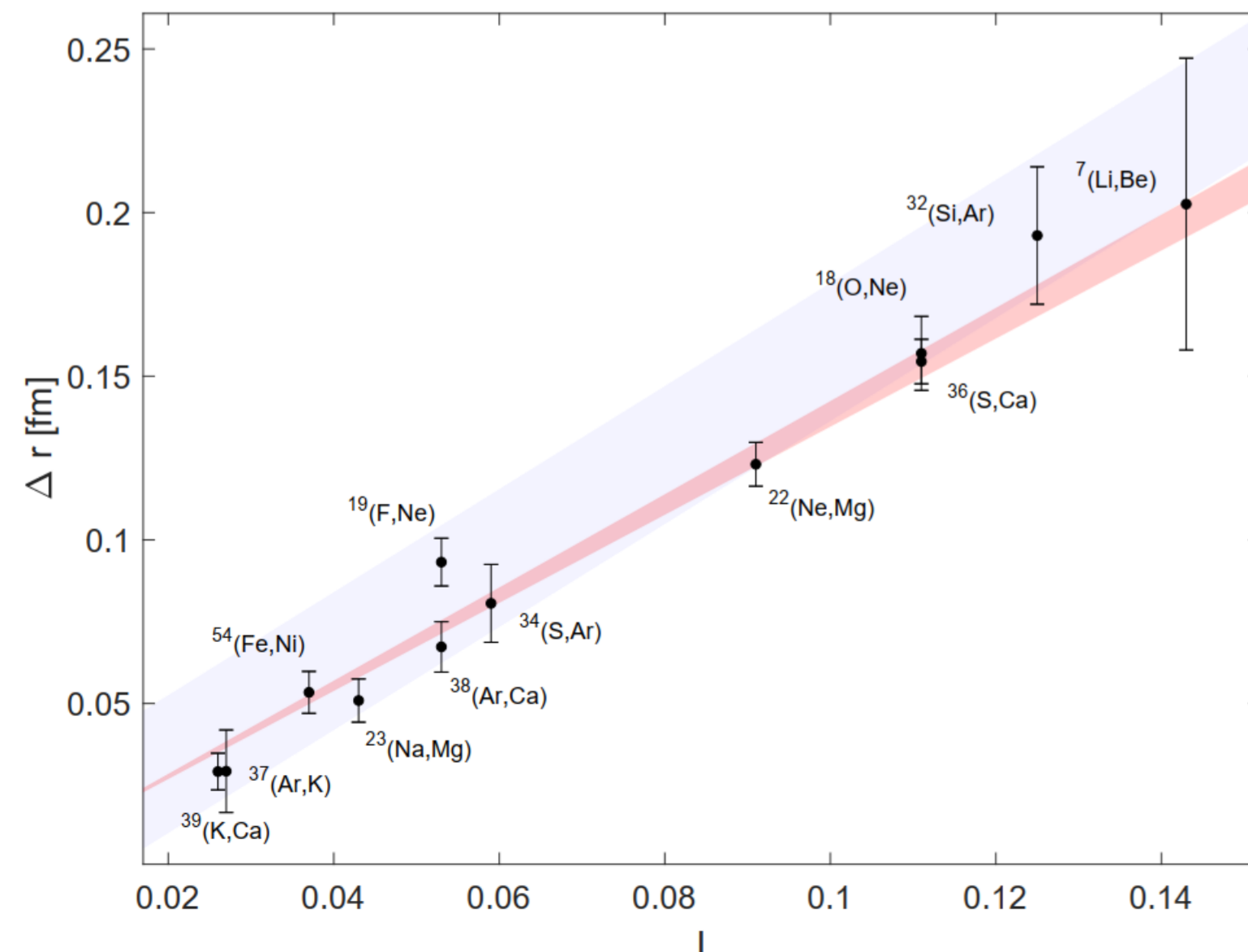
## 2. Isotope shifts (electronic atoms)



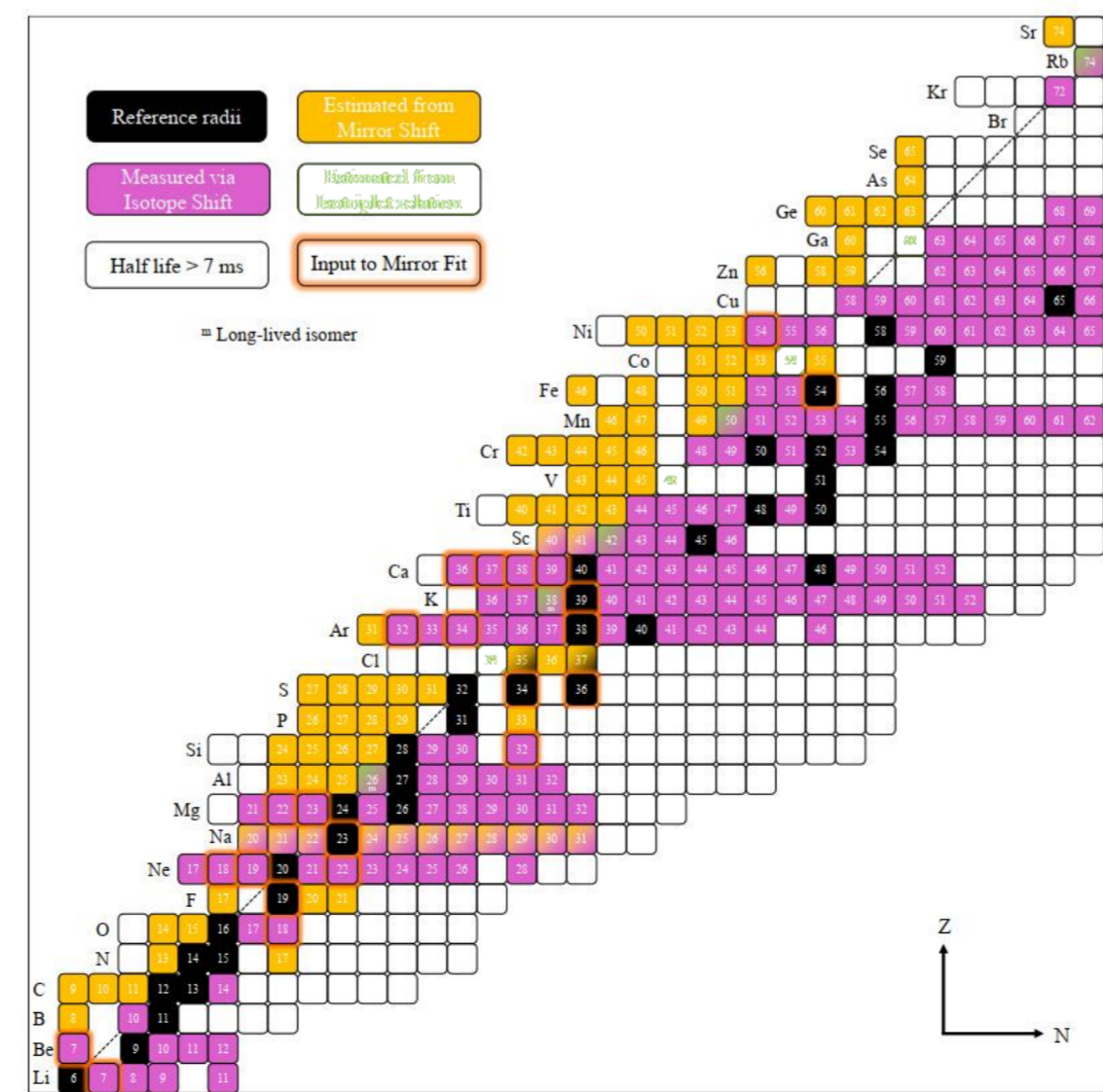
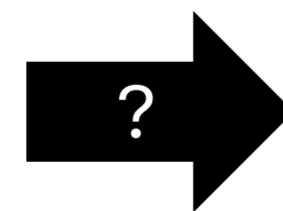
## 3. Mirror nuclei



## 4. Mirror fit

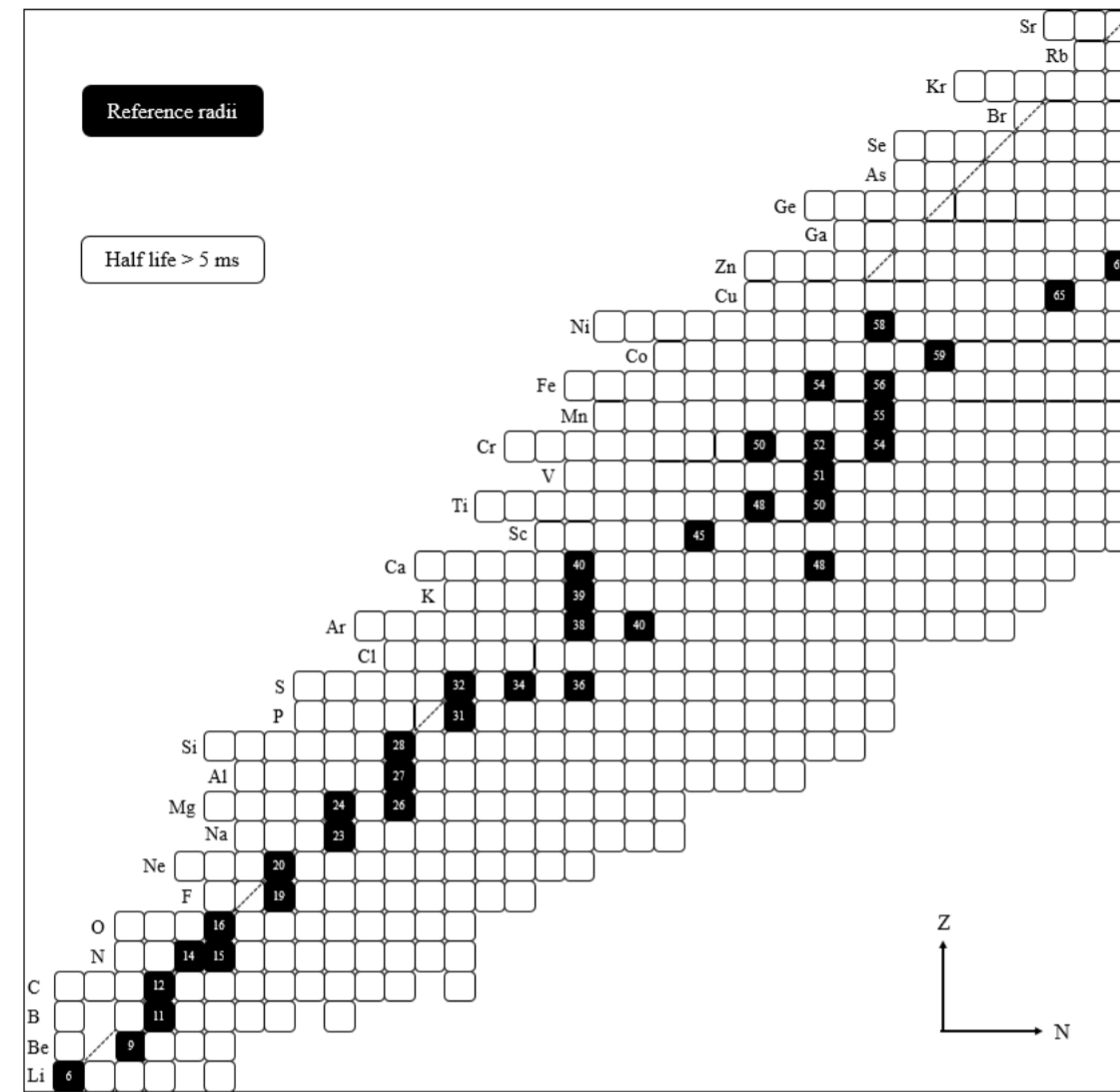


## 5. Enhanced database

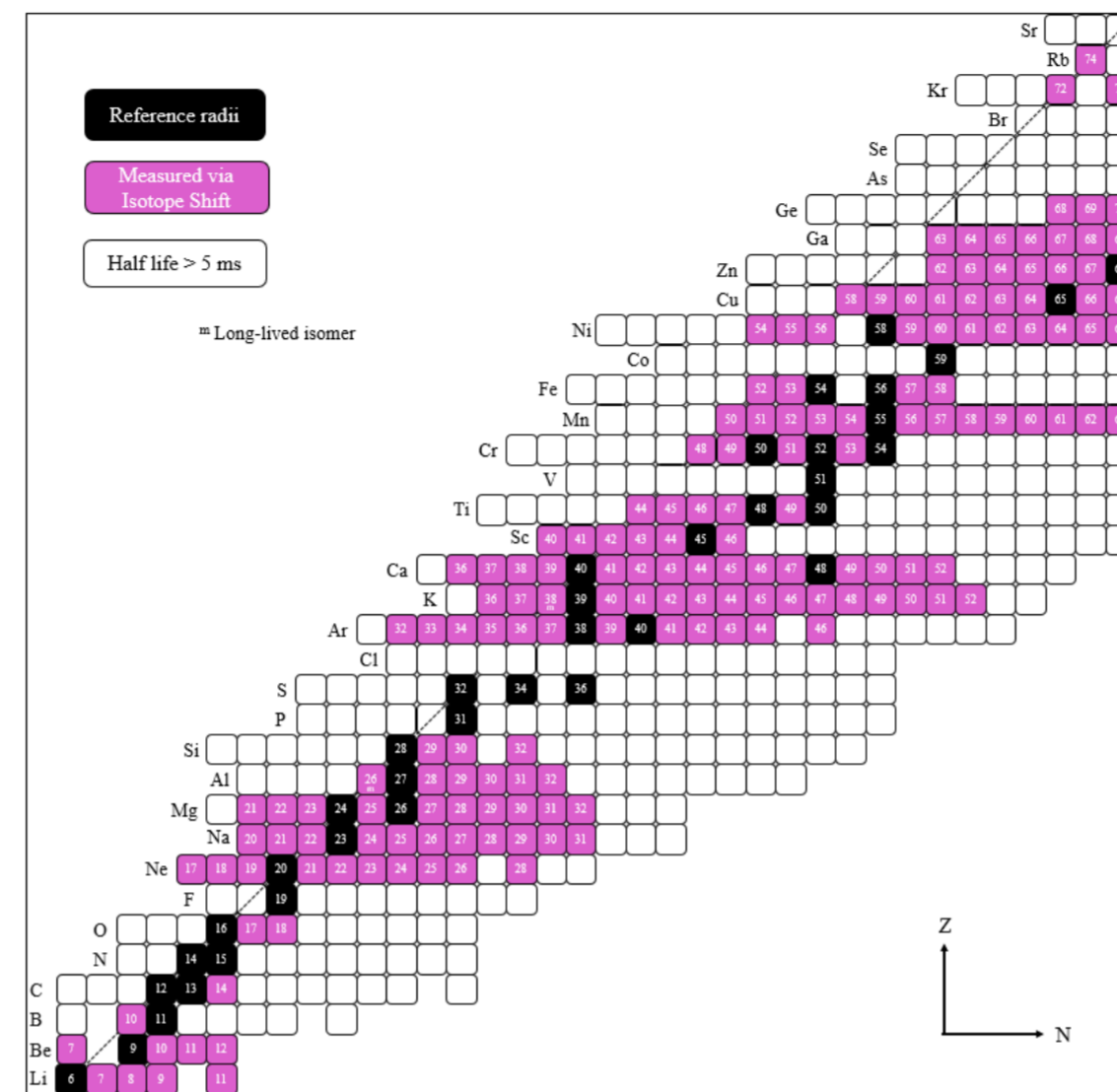


# Roadmap:

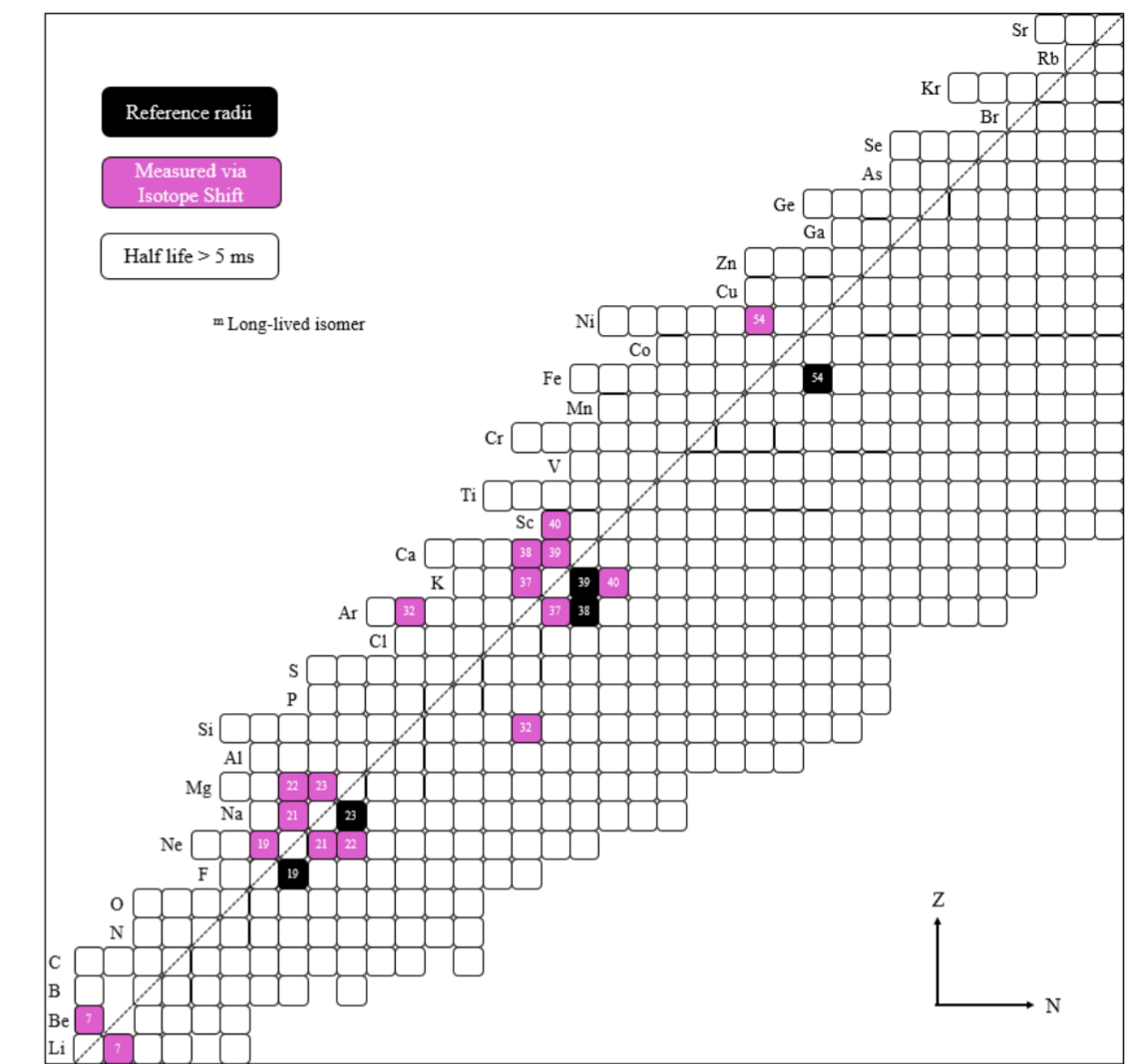
## 1. Reference radii (muonic atoms)



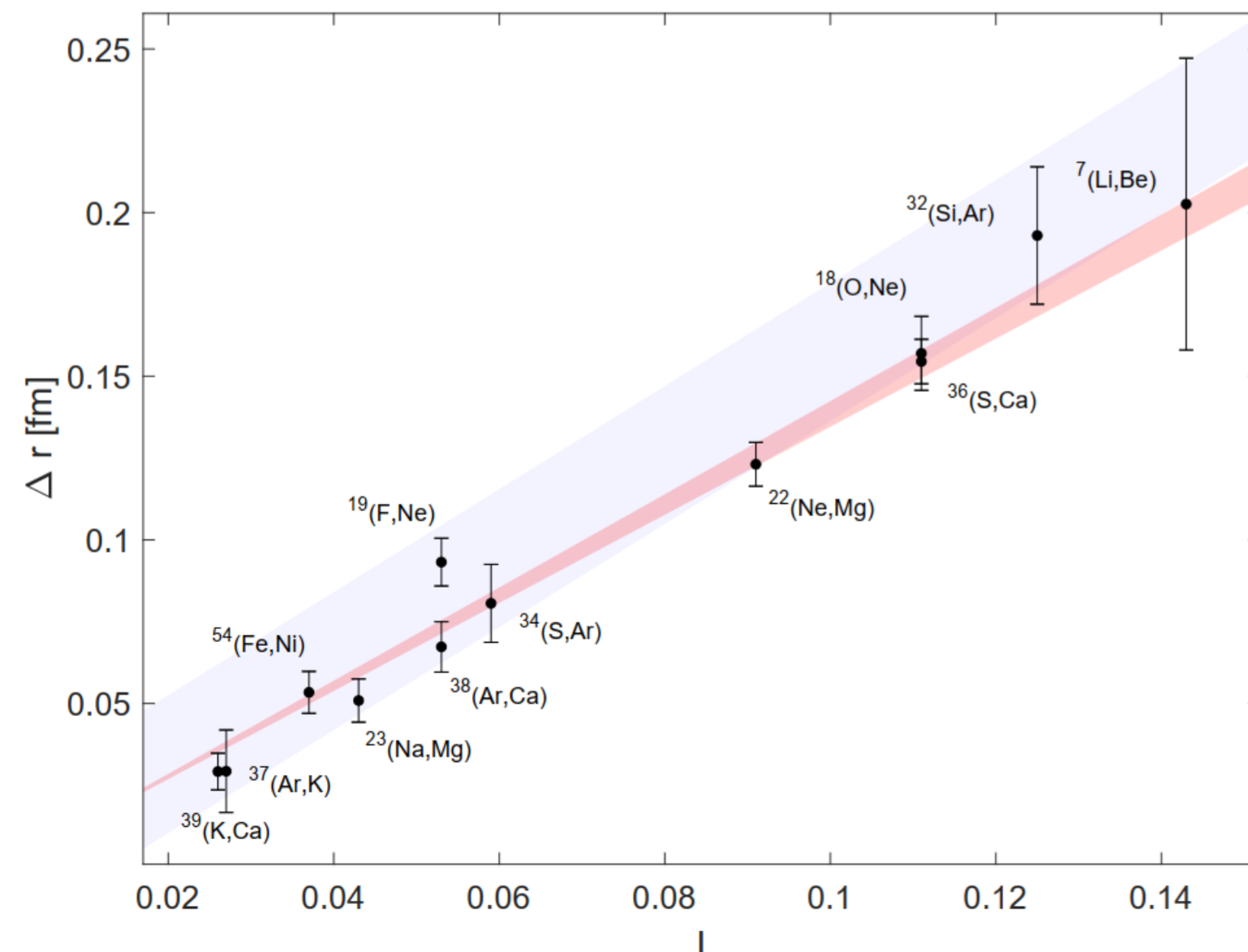
## 2. Isotope shifts (electronic atoms)



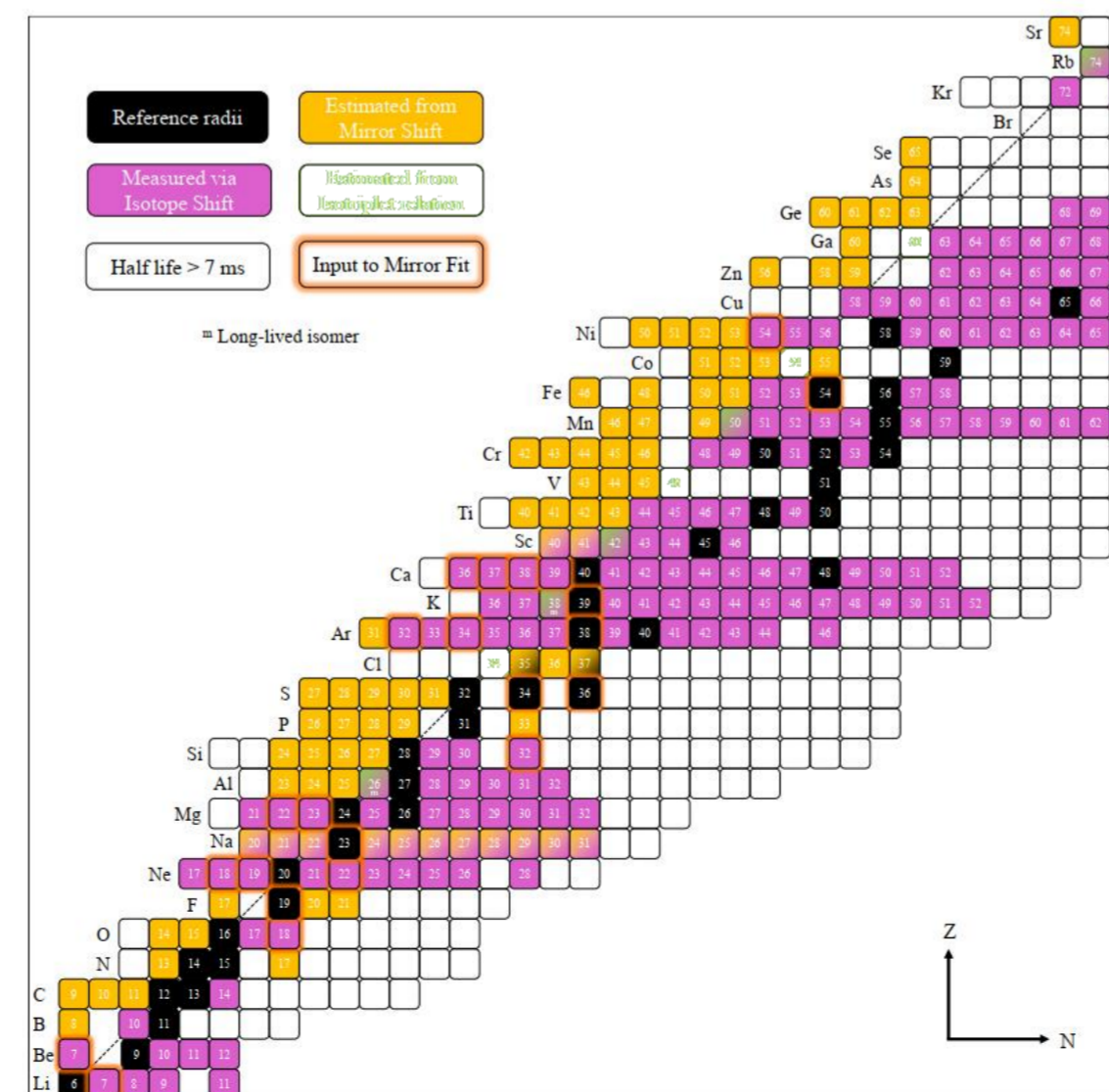
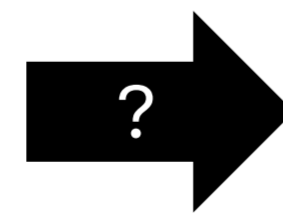
## 3. Mirror nuclei



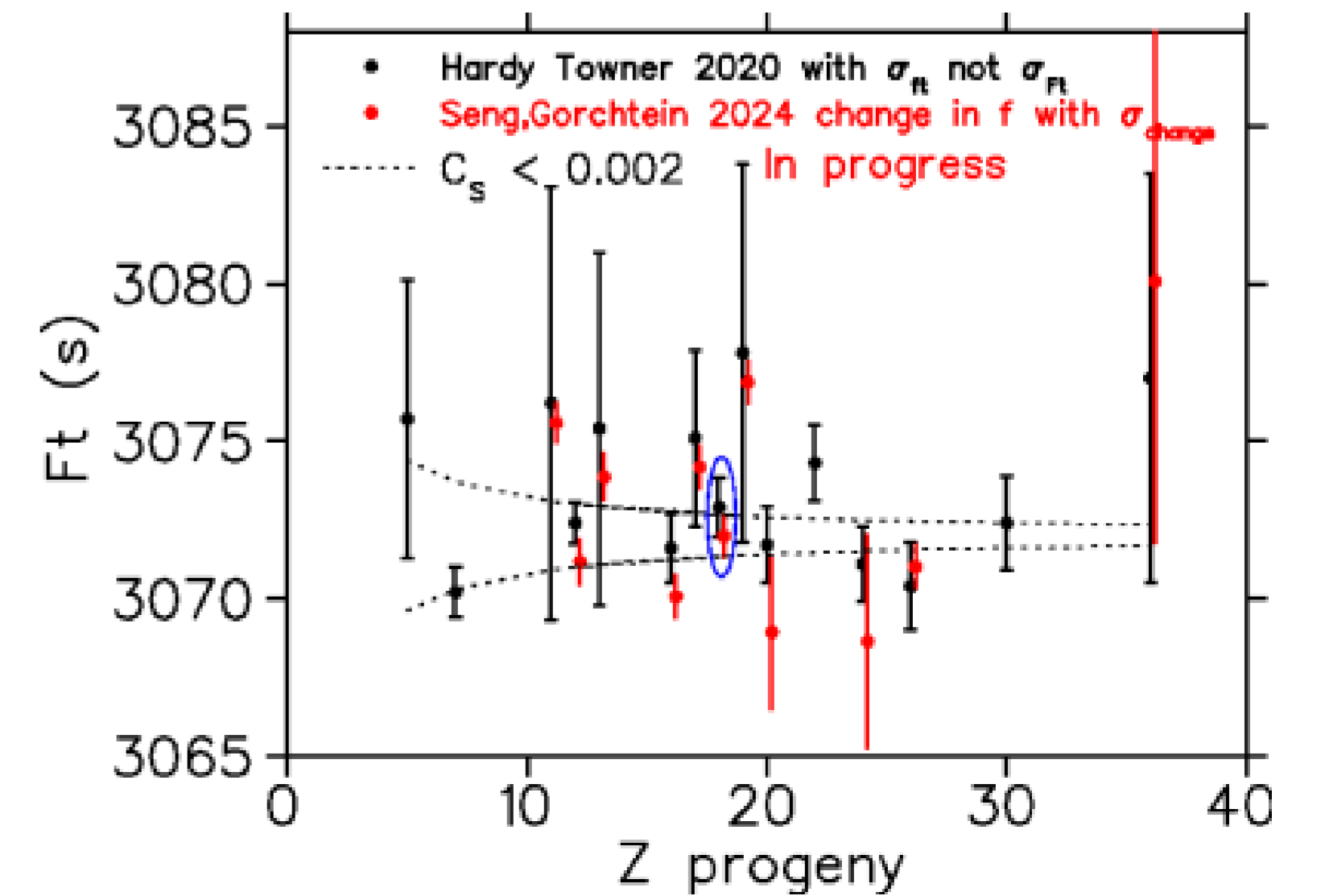
## 4. Mirror fit



## 5. Enhanced database

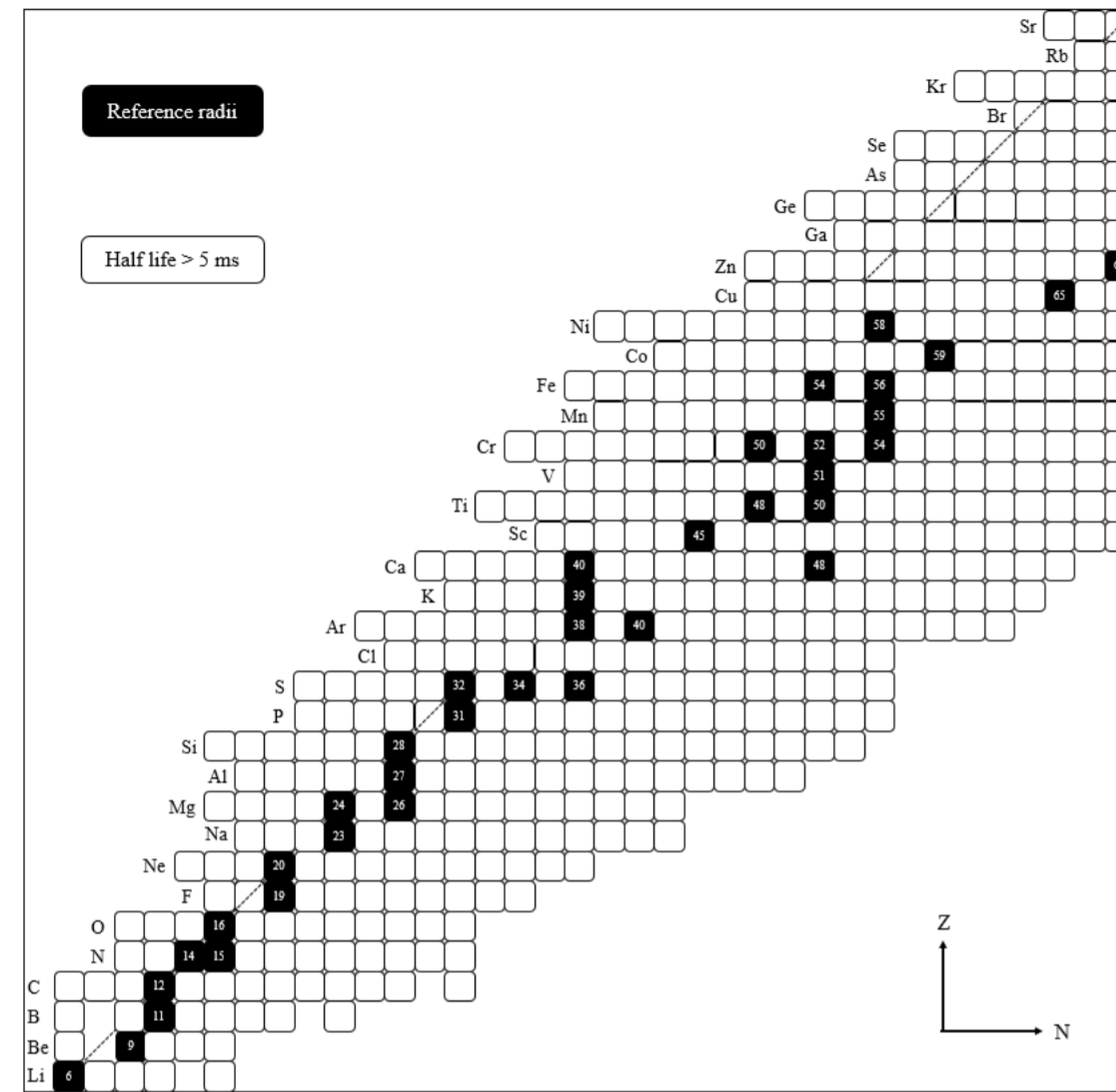


## 6. Charge and Weak radii for Vud



# Roadmap:

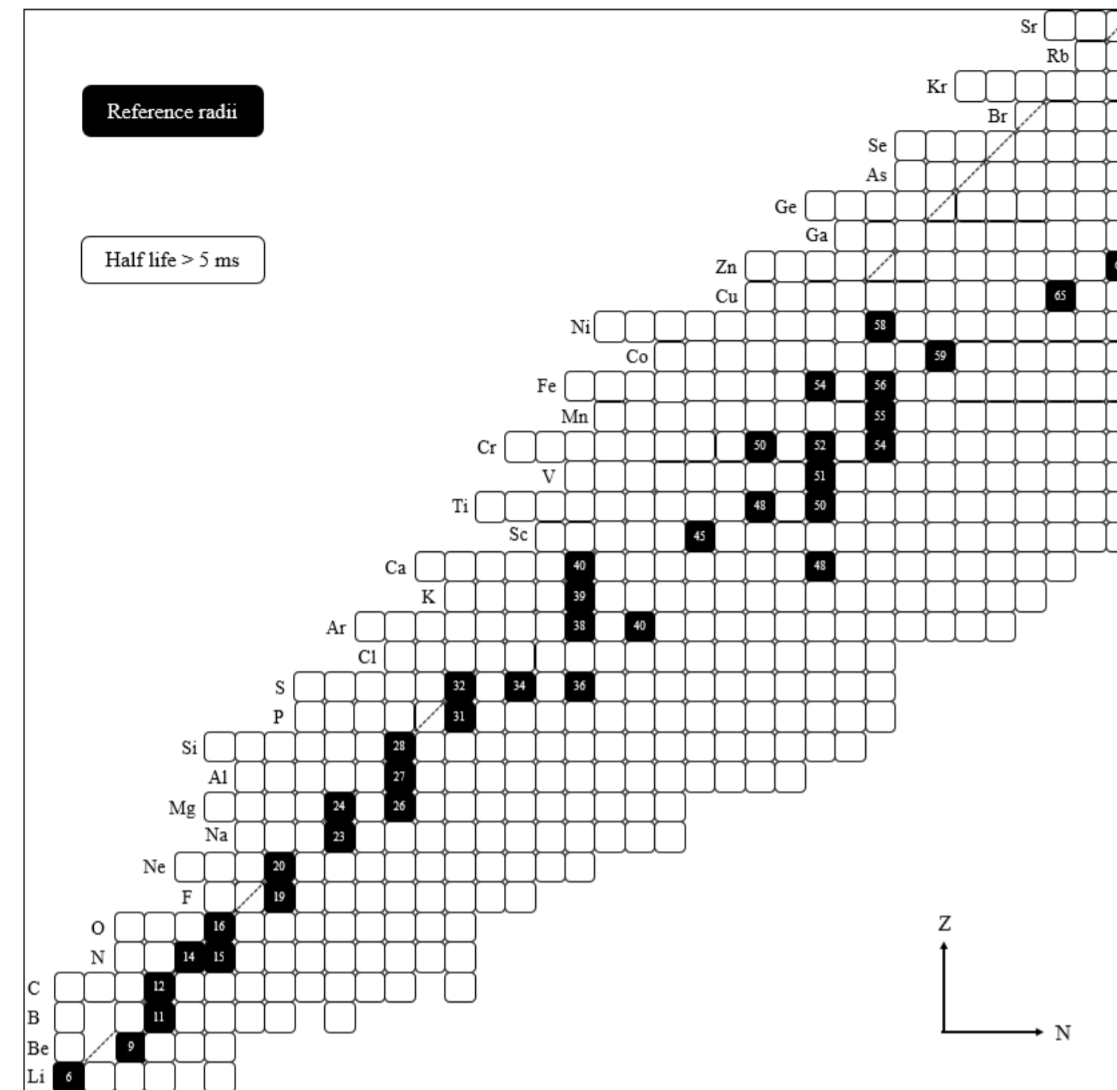
## 1. Reference radii (muonic atoms)





# Roadmap:

## 1. Reference radii (muonic atoms)



## Table of experimental nuclear ground state charge radii: An update

I. Angeli<sup>a</sup>, K.P. Marinova<sup>b,\*</sup>

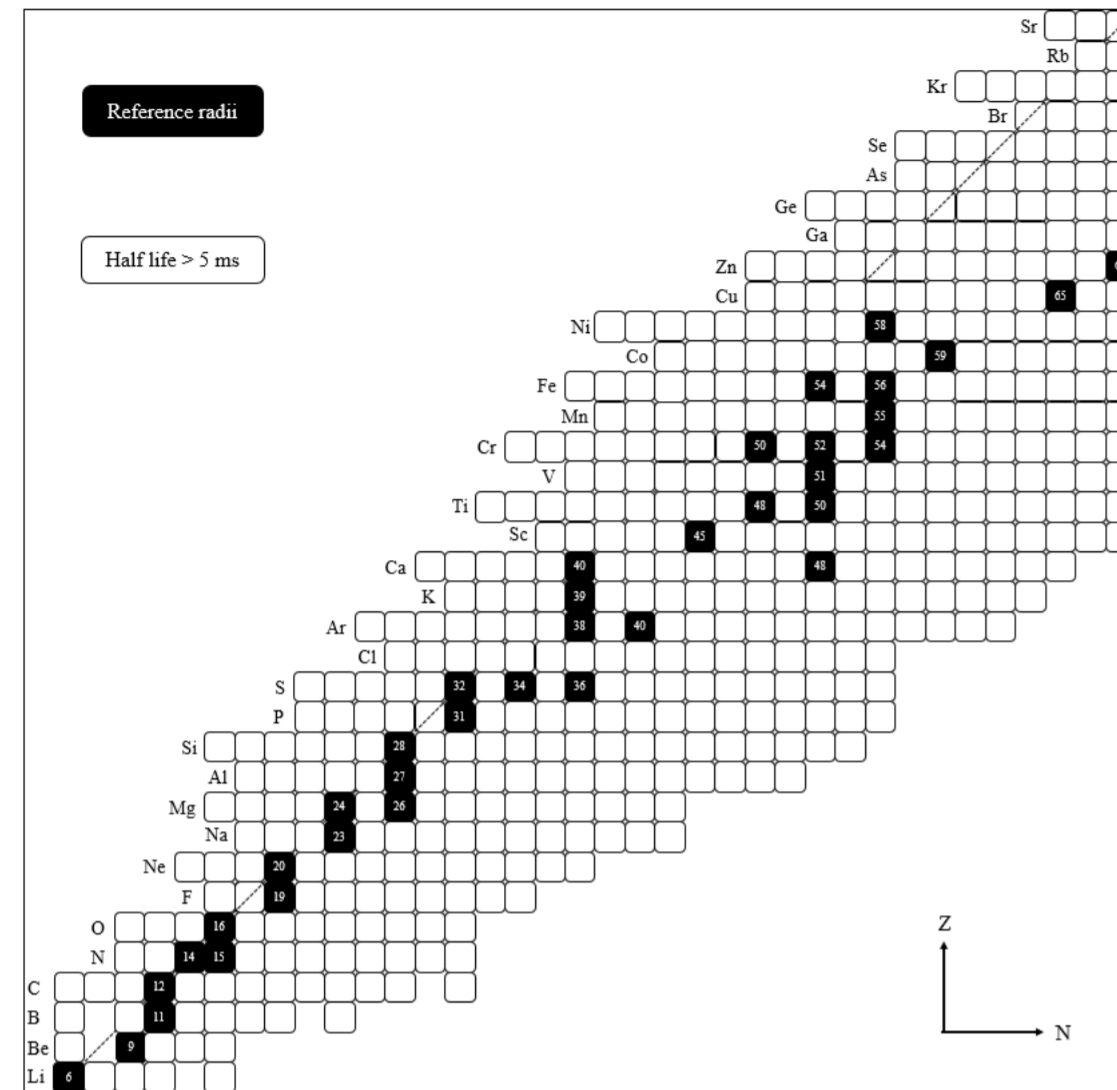
<sup>a</sup> Institute of Experimental Physics, University of Debrecen, H-4010 Debrecen Pf. 105, Hungary

<sup>b</sup> Joint Institute for Nuclear Research, 141980 Dubna, Moscow Region, Russia



# Roadmap:

## 1. Reference radii (muonic atoms)



## Table of experimental nuclear ground state charge radii: An update

I. Angeli<sup>a</sup>, K.P. Marinova<sup>b,\*</sup>

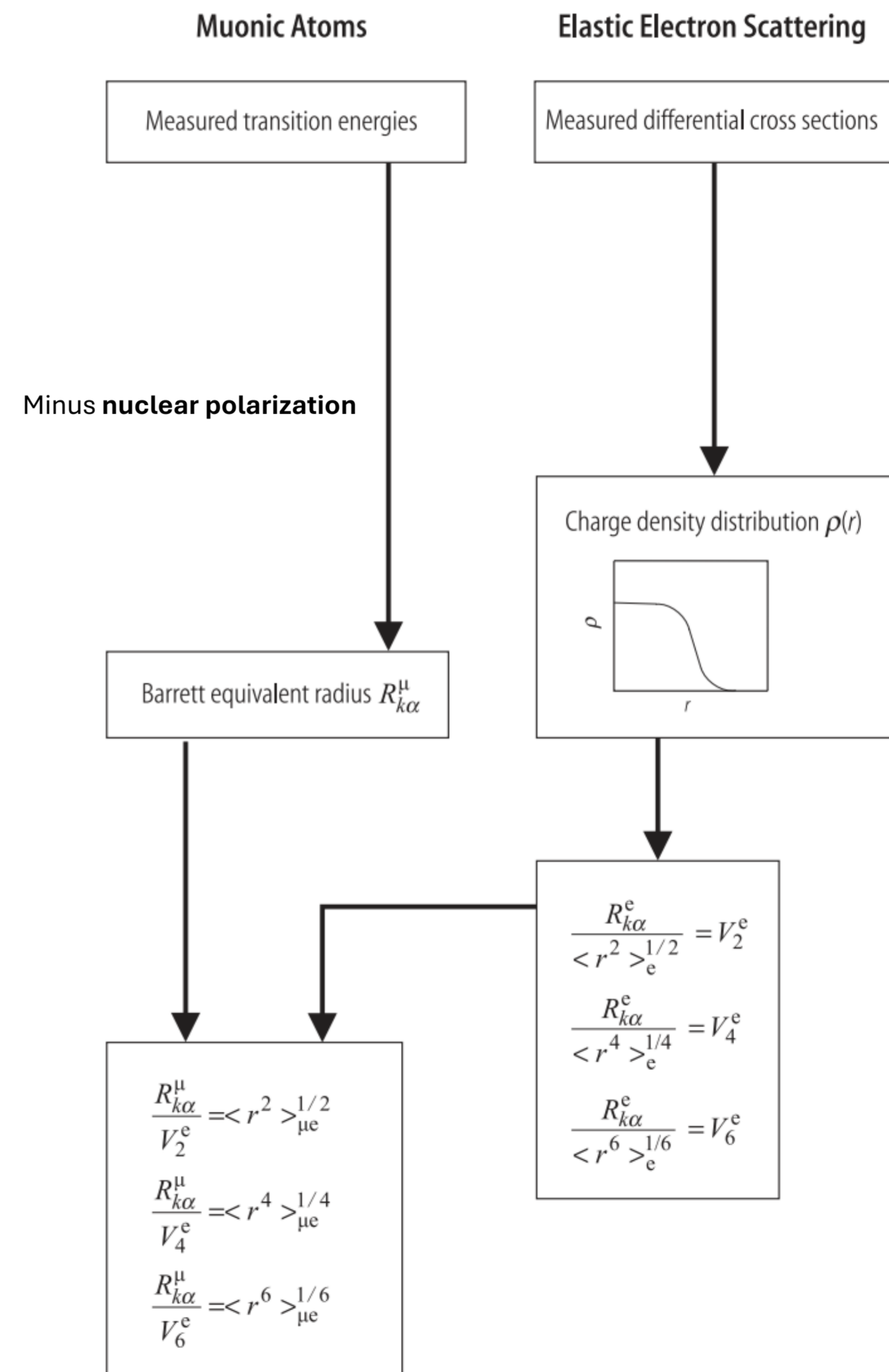
<sup>a</sup> Institute of Experimental Physics, University of Debrecen, H-4010 Debrecen Pf. 105, Hungary

<sup>b</sup> Joint Institute for Nuclear Research, 141980 Dubna, Moscow Region, Russia



- Step 1 : Update reference radii to reflect uncertainties

# “Old school” combined analysis of muonic atoms and electron scattering:

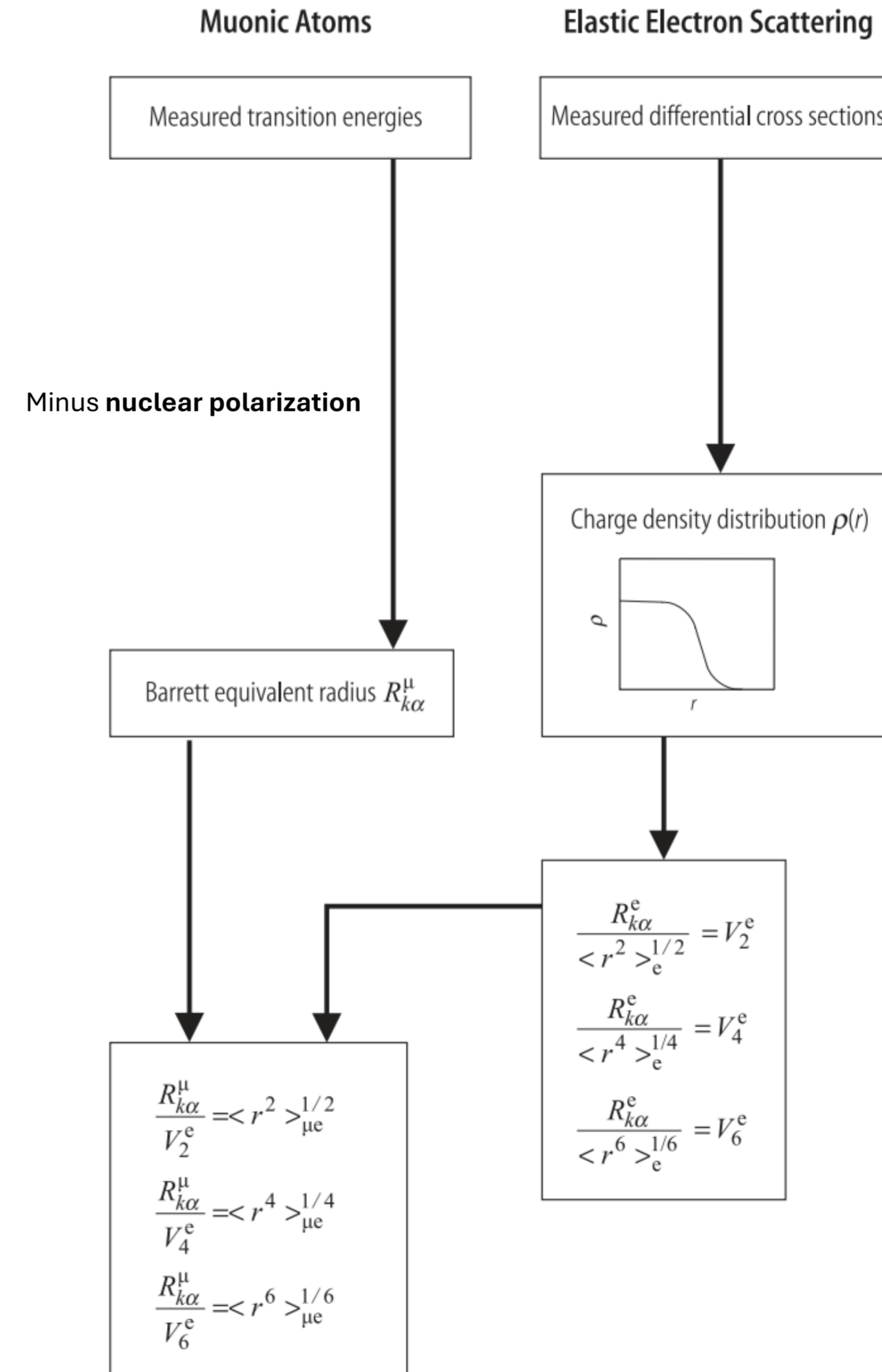


# “Old school” combined analysis of muonic atoms and electron scattering:

- What is the limitation of the “Barret recipe?”

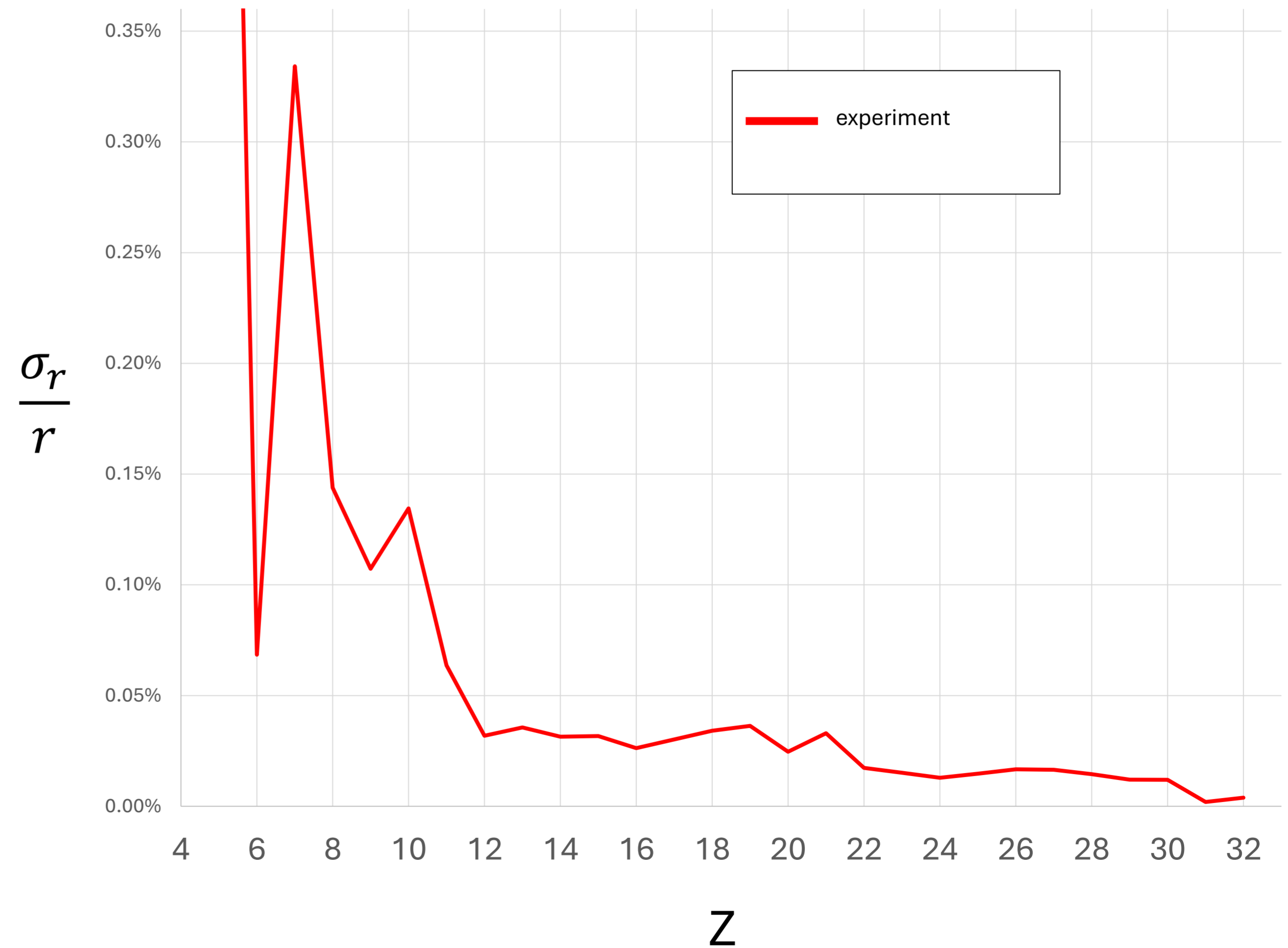
Three sources of uncertainty:

1. Experiment (muonic atom energies)
2. Theory (nuclear polarization)
3. **Charge distribution (scattering)**



# Sources of uncertainty :

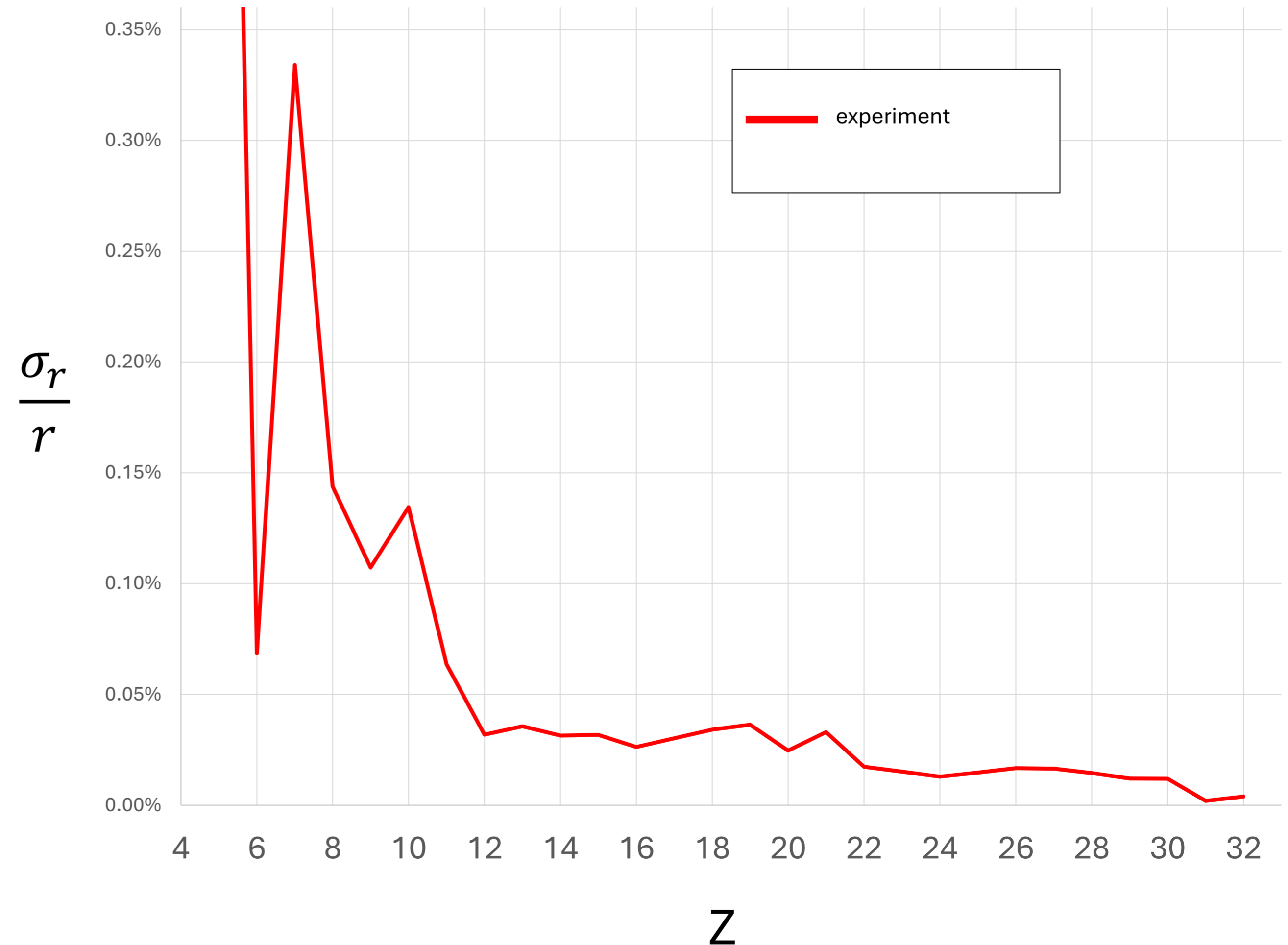
Radii of light nuclei from muonic atom x-ray spec.



# Sources of uncertainty :

Radii of light nuclei from muonic atom x-ray spec.

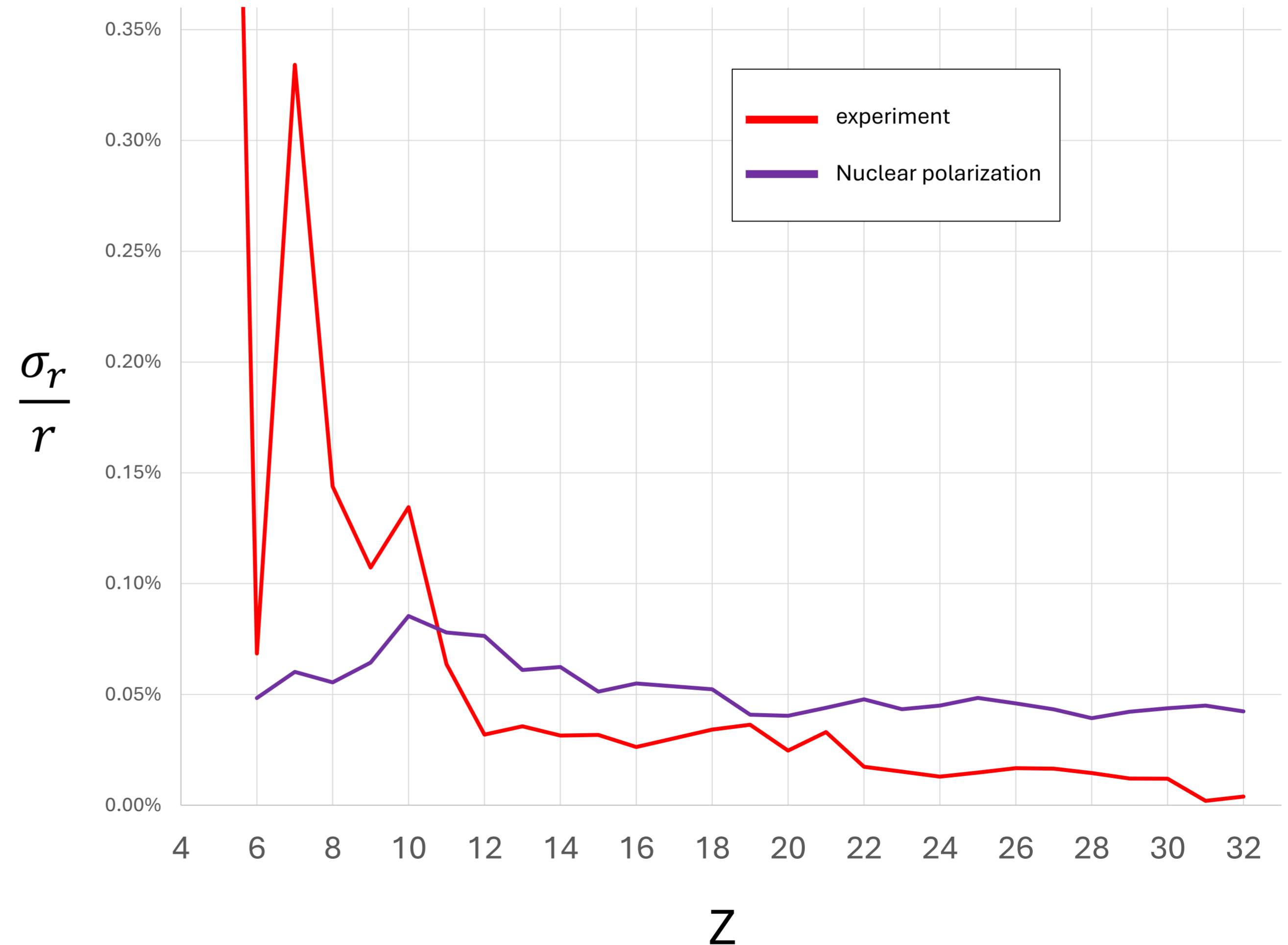
- Notice: Fricke & Heilig only include statistical unc. -> consult original papers



# Sources of uncertainty :

Radii of light nuclei from muonic atom x-ray spec.

- Notice: nuclear polarization from Rinker & Speth 1970s, assumed 30% unc. Reliable?



# How about electron scattering?

Idea by Ingo Sick, applied [here](#)

- Experimental uncertainty (normalization) goes away for ratios
- Residual model-dependency from finite momentum transfer. Hard to quantify!

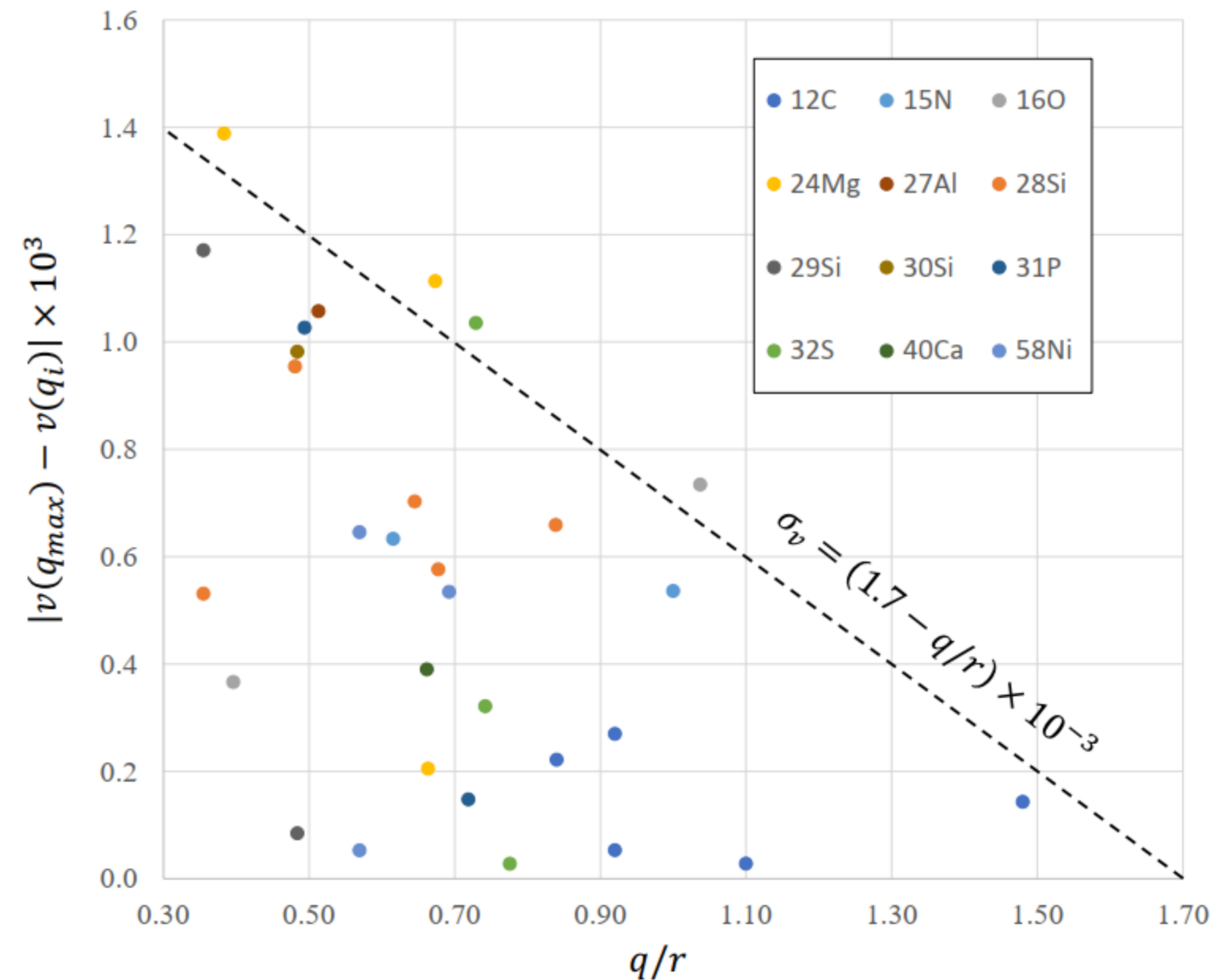


# Uncertainty in ratios of moments?

- Experimental uncertainty (normalization) goes away for ratios
- Residual model-dependency from finite momentum transfer
- How much does the second-best scattering measurement deviates from the best one?

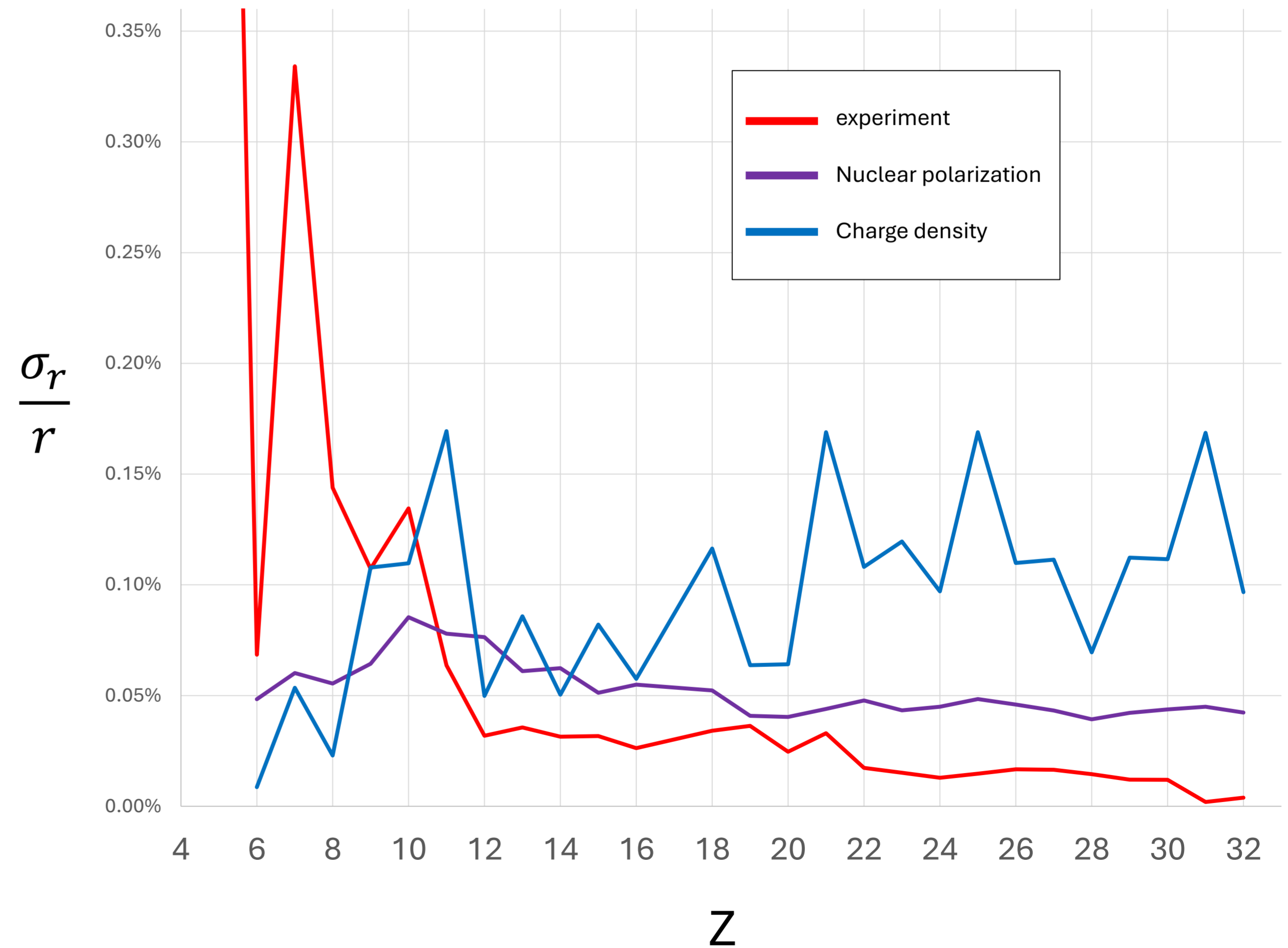
Quantify our intuition:

Best experiment has broad momentum transfer compared to nuclear size.



# Sources of uncertainty :

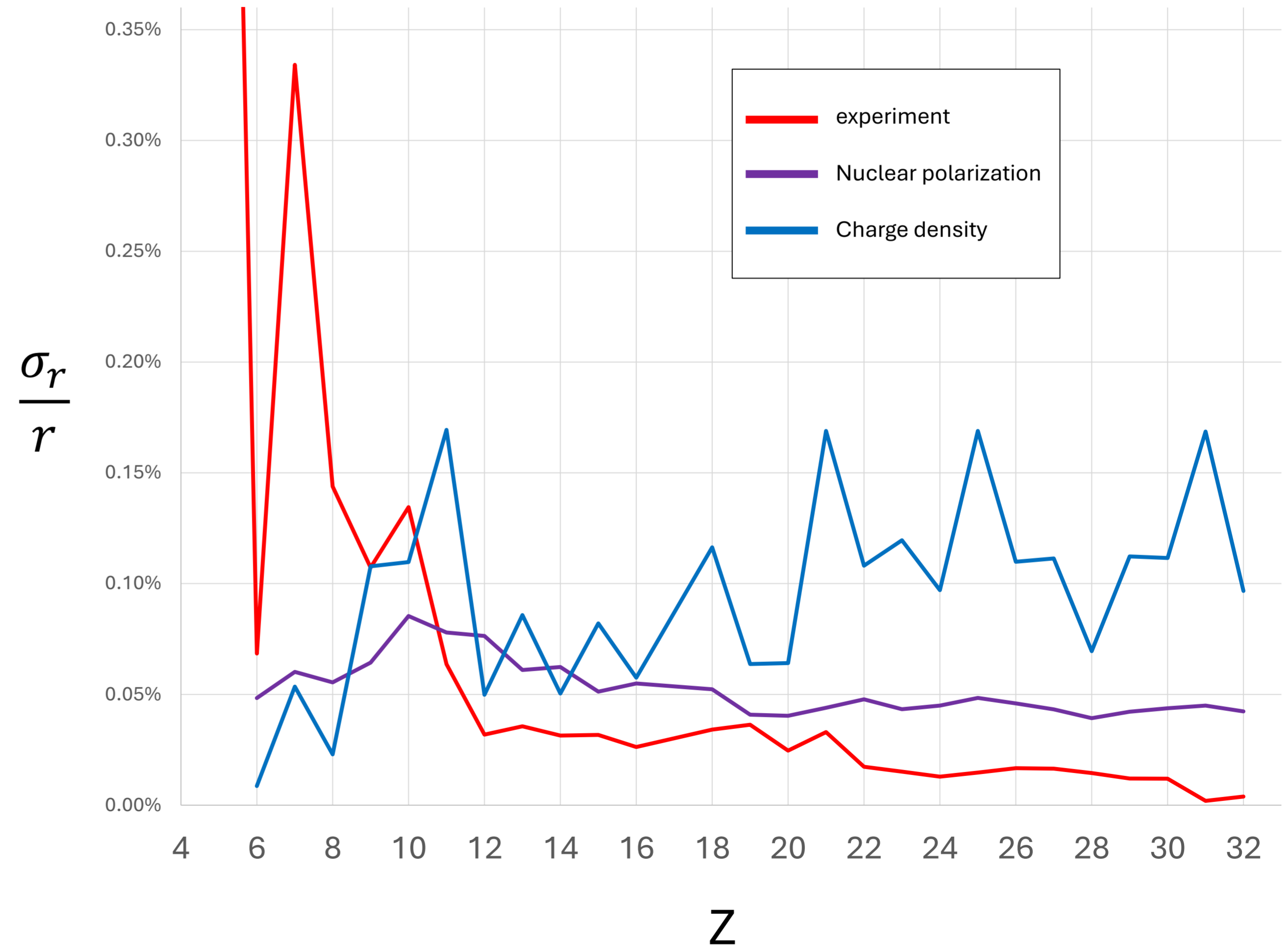
Radii of light nuclei from muonic atom x-ray spec.



# Sources of uncertainty :

Radii of light nuclei from muonic atom x-ray spec.

- Within Barret-recipe: scattering model dependency matters !



# Transparent tabulation (inviting your input!)

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)

**Table 2**

Reference radii used in this work. Unless stated otherwise in the note, they are determined via Eq. 1 and 2 with the  $2P_{3/2} - 1S$  Barret radii given in [4] and the  $\nu$  factors from tab. 1. Uncertainties are denoted by  $\sigma$  and correspond to statistics and energy calibration (exp), nuclear polarization (NP), and charge distribution (CD) as resulting from the  $\nu$  factors of Tab. 1

el.	Z	A	$r_{\text{ch}}$	$\sigma_{\text{exp}}$	$\sigma_{\text{NP}}$	$\sigma_{\text{CD}}$	$\sigma_{\text{tot}}$	Note
Li	3	6	2.589	0.039			0.039	A
Be	4	9	2.519	0.012		0.030	0.032	B
B	5	11	2.411	0.021			0.021	C
C	6	12	2.483	0.002	0.001	0.000	0.002	D
N	7	14	2.556	0.009	0.002	0.001	0.009	
	7	15	2.612	0.009			0.009	E
O	8	16	2.701	0.004	0.001	0.001	0.004	F
F	9	19	2.902	0.003	0.002	0.003	0.005	†
Ne	10	20	3.001	0.004	0.003	0.003	0.006	†
Na	11	23	2.992	0.002	0.002	0.005	0.006	†
Mg	12	24	3.056	0.001	0.002	0.002	0.003	†
	12	26	3.030	0.001	0.002	0.002	0.003	†
Al	13	27	3.061	0.001	0.002	0.003	0.003	†G
Si	14	28	3.123	0.001	0.002	0.002	0.003	†
P	15	31	3.190	0.001	0.002	0.002	0.003	
S	16	32	3.262	0.001	0.002	0.003	0.003	
	16	34	3.284	0.001	0.002	0.003	0.004	
	16	36	3.298	0.001	0.001	0.003	0.004	
Cl	17	35	3.388	0.015			0.015	H
Cl	17	37	3.384	0.015			0.015	H
Ar	18	38	3.402	0.002	0.003	0.005	0.006	
	18	40	3.427	0.001	0.002	0.003	0.004	
K	19	39	3.435	0.001	0.001	0.003	0.004	
Ca	20	40	3.481	0.001	0.001	0.004	0.004	
	20	48	3.475	0.001	0.001	0.002	0.002	
Sc	21	45	3.548	0.001	0.002	0.006	0.007	

# Transparent tabulation (inviting your input!)

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)

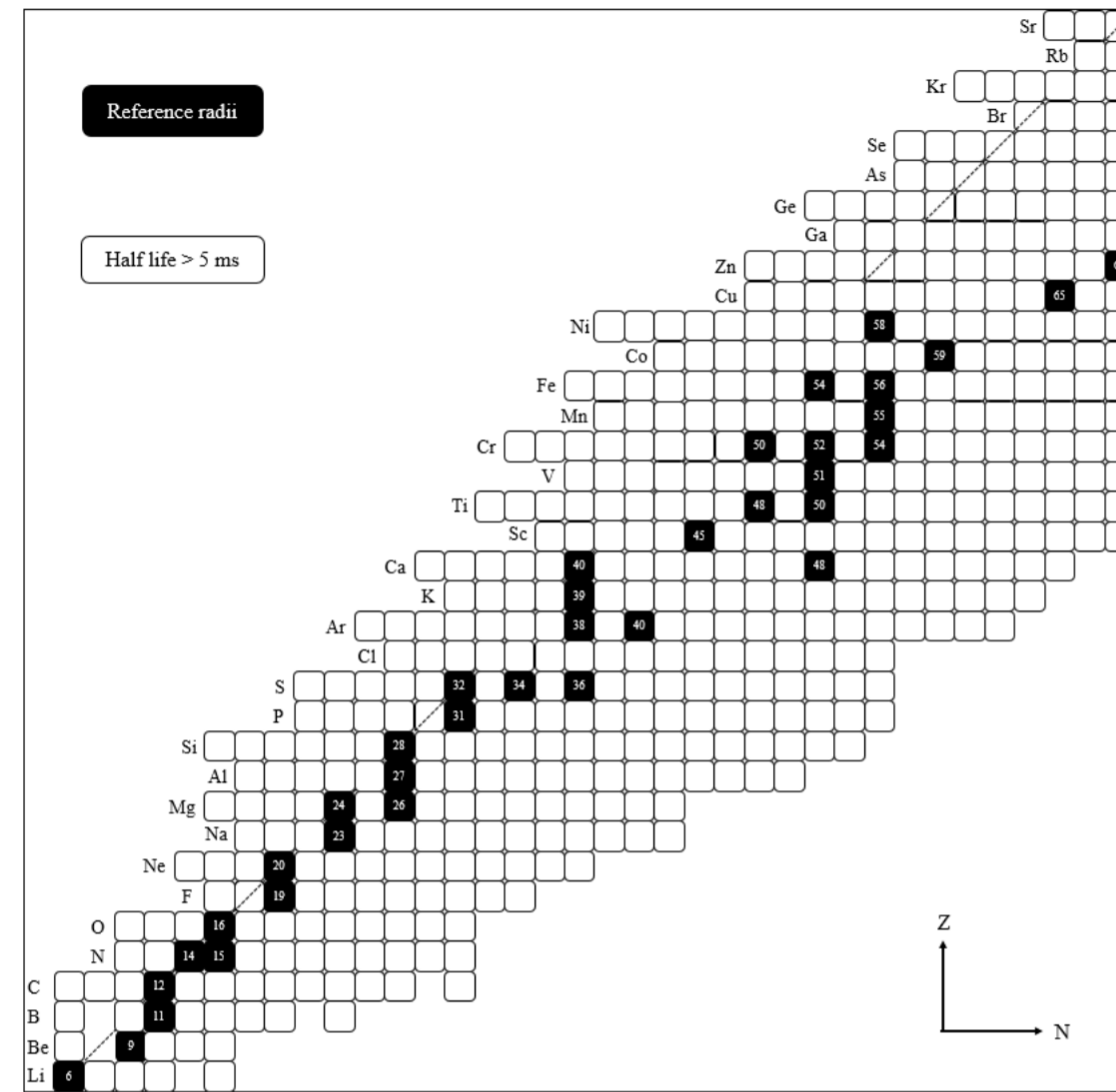
**Table 2**

Reference radii used in this work. Unless stated otherwise in the note, they are determined via Eq. 1 and 2 with the  $2P_{3/2} - 1S$  Barret radii given in [4] and the  $\nu$  factors from tab. 1. Uncertainties are denoted by  $\sigma$  and correspond to statistics and energy calibration (exp), nuclear polarization (NP), and charge distribution (CD) as resulting from the  $\nu$  factors of Tab. 1

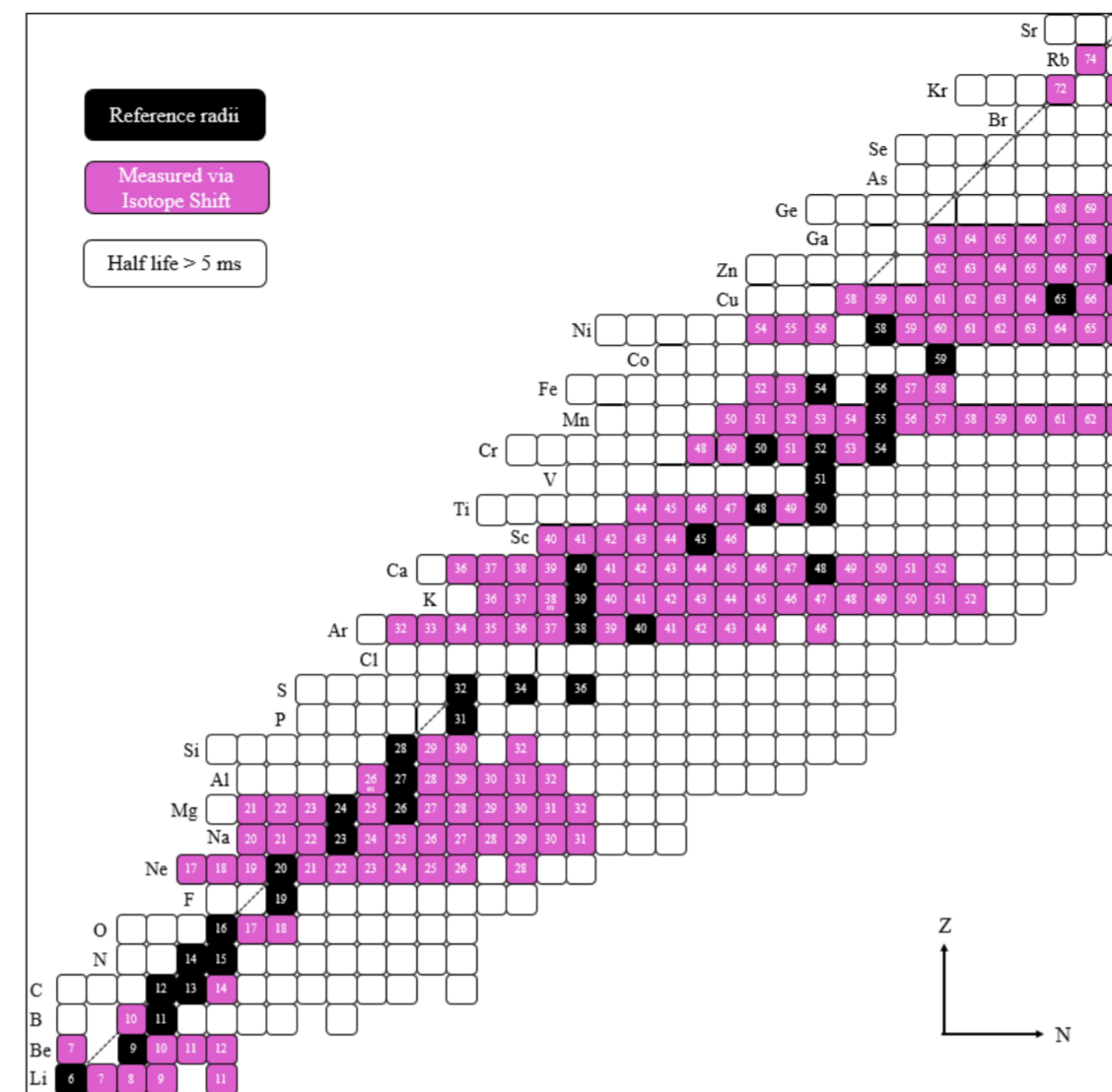
el.	Z	A	$r_{\text{ch}}$	$\sigma_{\text{exp}}$	$\sigma_{\text{NP}}$	$\sigma_{\text{CD}}$	$\sigma_{\text{tot}}$	Note
Li	3	6	2.589	0.039			0.039	A
Be	4	9	2.519	0.012		0.030	0.032	B
B	5	11	2.411	0.021			0.021	C
C	6	12	2.483	0.002	0.001	0.000	0.002	D
N	7	14	2.556	0.009	0.002	0.001	0.009	
	7	15	2.612	0.009			0.009	E
O	8	16	2.701	0.004	0.001	0.001	0.004	F
F	9	19	2.902	0.003	0.002	0.003	0.005	†
Example: Ne	10	20	3.001	0.004	0.003	0.003	0.006	AM: 3.006(2) fm
Na	11	23	2.992	0.002	0.002	0.005	0.006	†
Mg	12	24	3.056	0.001	0.002	0.002	0.003	†
	12	26	3.030	0.001	0.002	0.002	0.003	†
Al	13	27	3.061	0.001	0.002	0.003	0.003	†G
Si	14	28	3.123	0.001	0.002	0.002	0.003	†
P	15	31	3.190	0.001	0.002	0.002	0.003	
S	16	32	3.262	0.001	0.002	0.003	0.003	
	16	34	3.284	0.001	0.002	0.003	0.004	
	16	36	3.298	0.001	0.001	0.003	0.004	
Cl	17	35	3.388	0.015			0.015	H
Cl	17	37	3.384	0.015			0.015	H
Ar	18	38	3.402	0.002	0.003	0.005	0.006	
	18	40	3.427	0.001	0.002	0.003	0.004	
K	19	39	3.435	0.001	0.001	0.003	0.004	
Ca	20	40	3.481	0.001	0.001	0.004	0.004	
	20	48	3.475	0.001	0.001	0.002	0.002	
Sc	21	45	3.548	0.001	0.002	0.006	0.007	

# Roadmap:

## 1. Reference radii (muonic atoms)

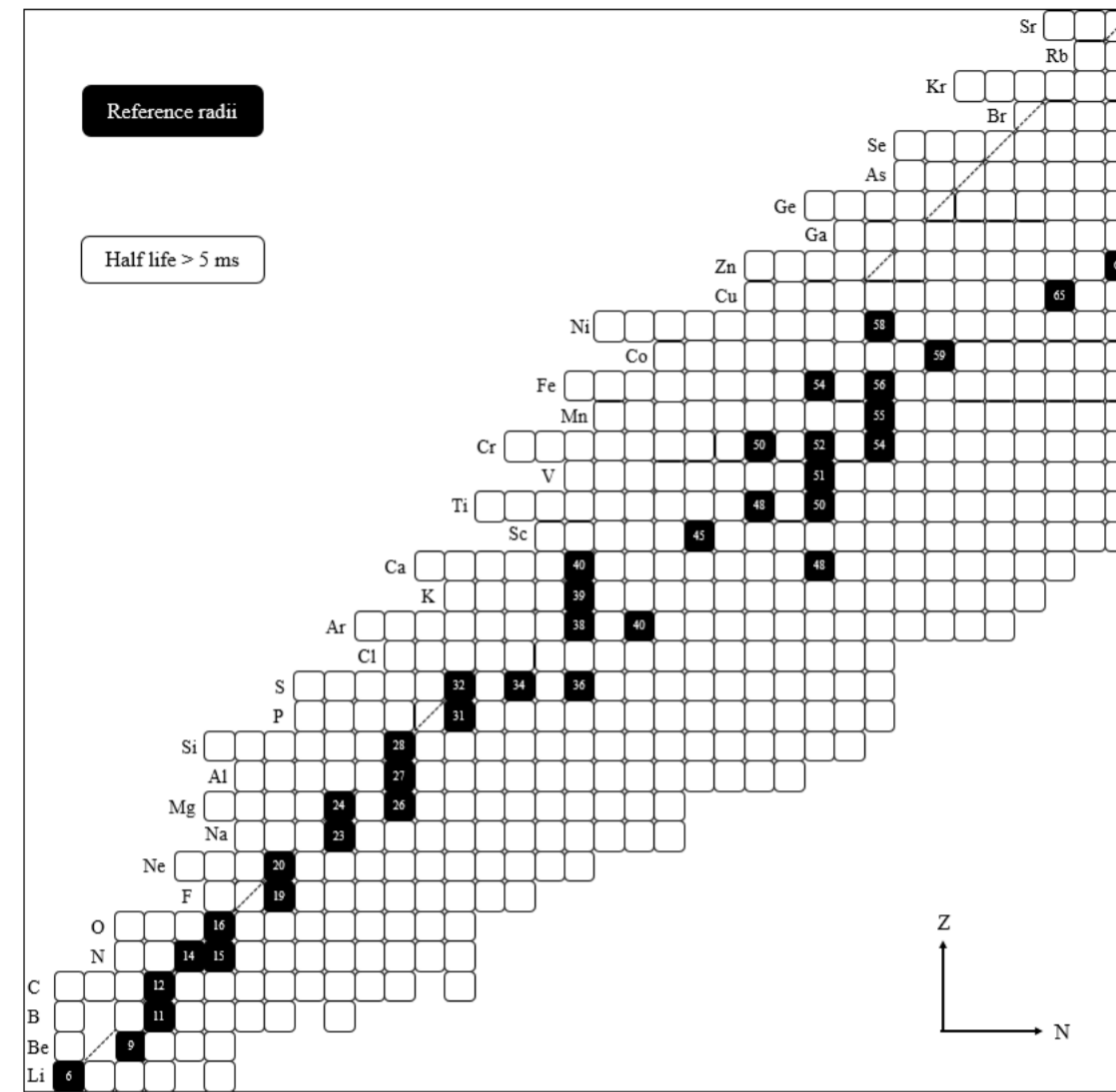


## 2. With Isotope shifts (electronic atoms)

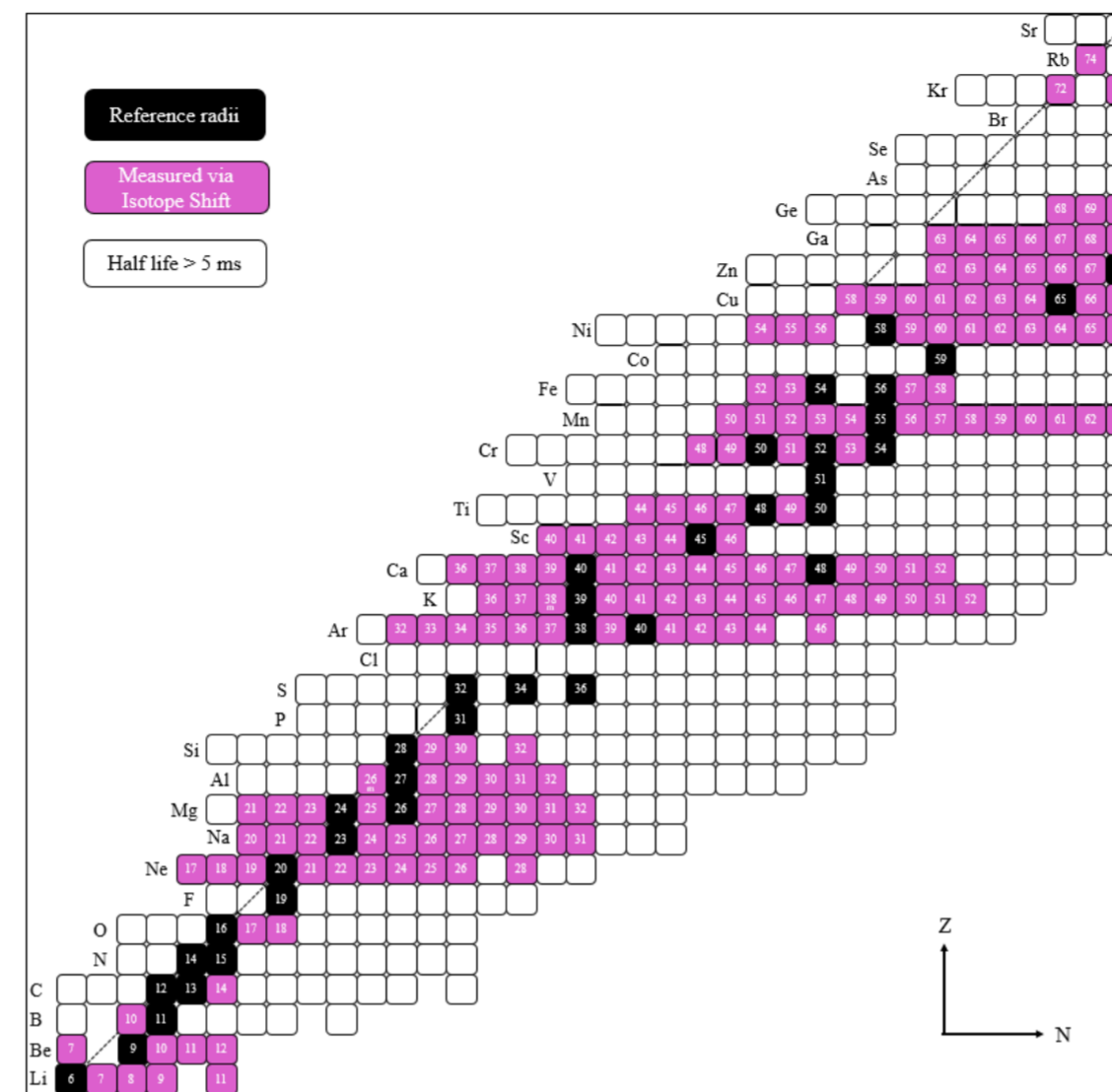


# Roadmap:

## 1. Reference radii (muonic atoms)



## 2. With Isotope shifts (electronic atoms)



$$r_x^2 = r_{ref}^2 + \delta r_{a,x}^2 \approx \frac{IS_{a,x} - K\mu_{a,x}}{F}$$

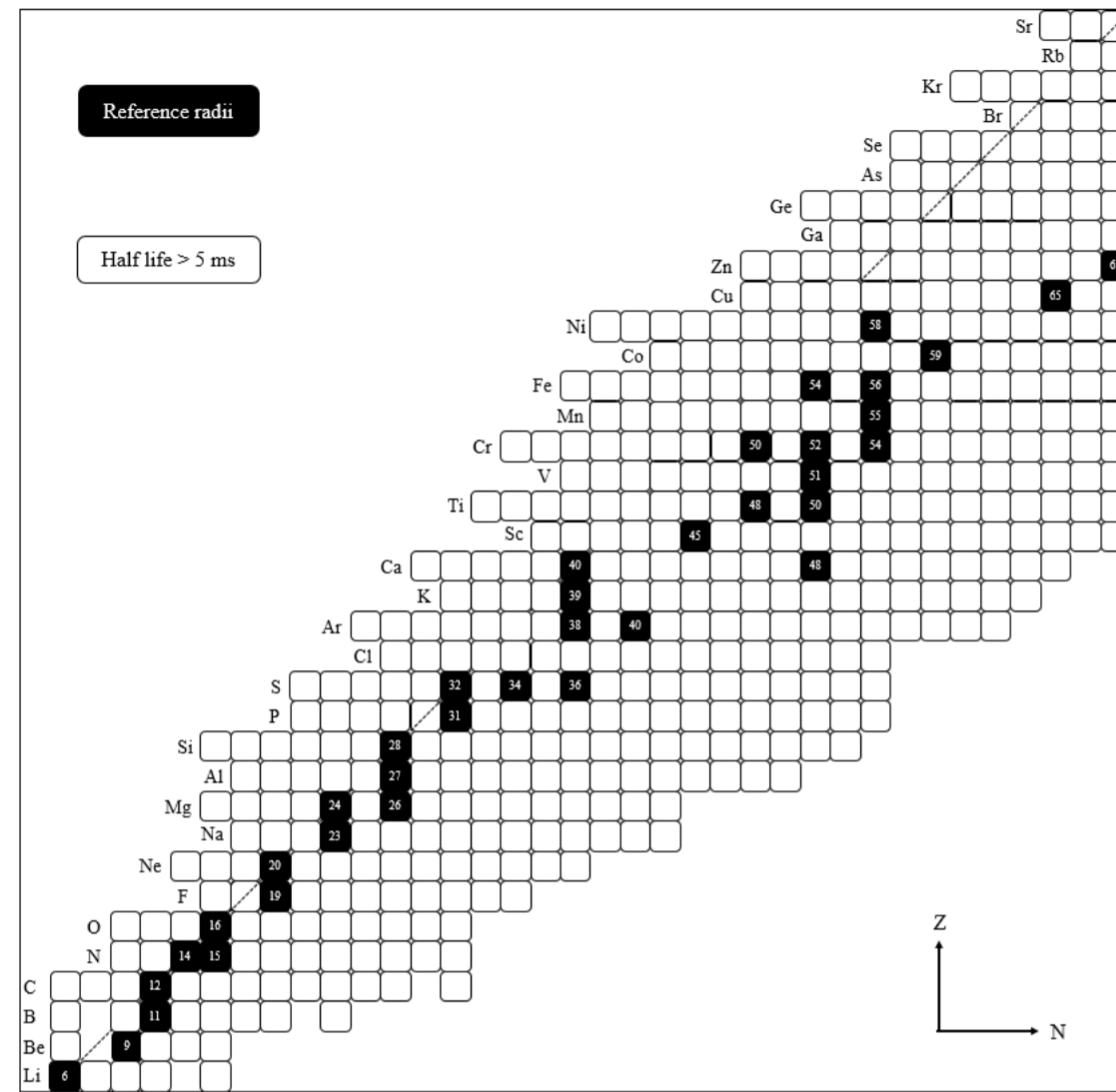
Isotope shift

Mass factor - Penning traps

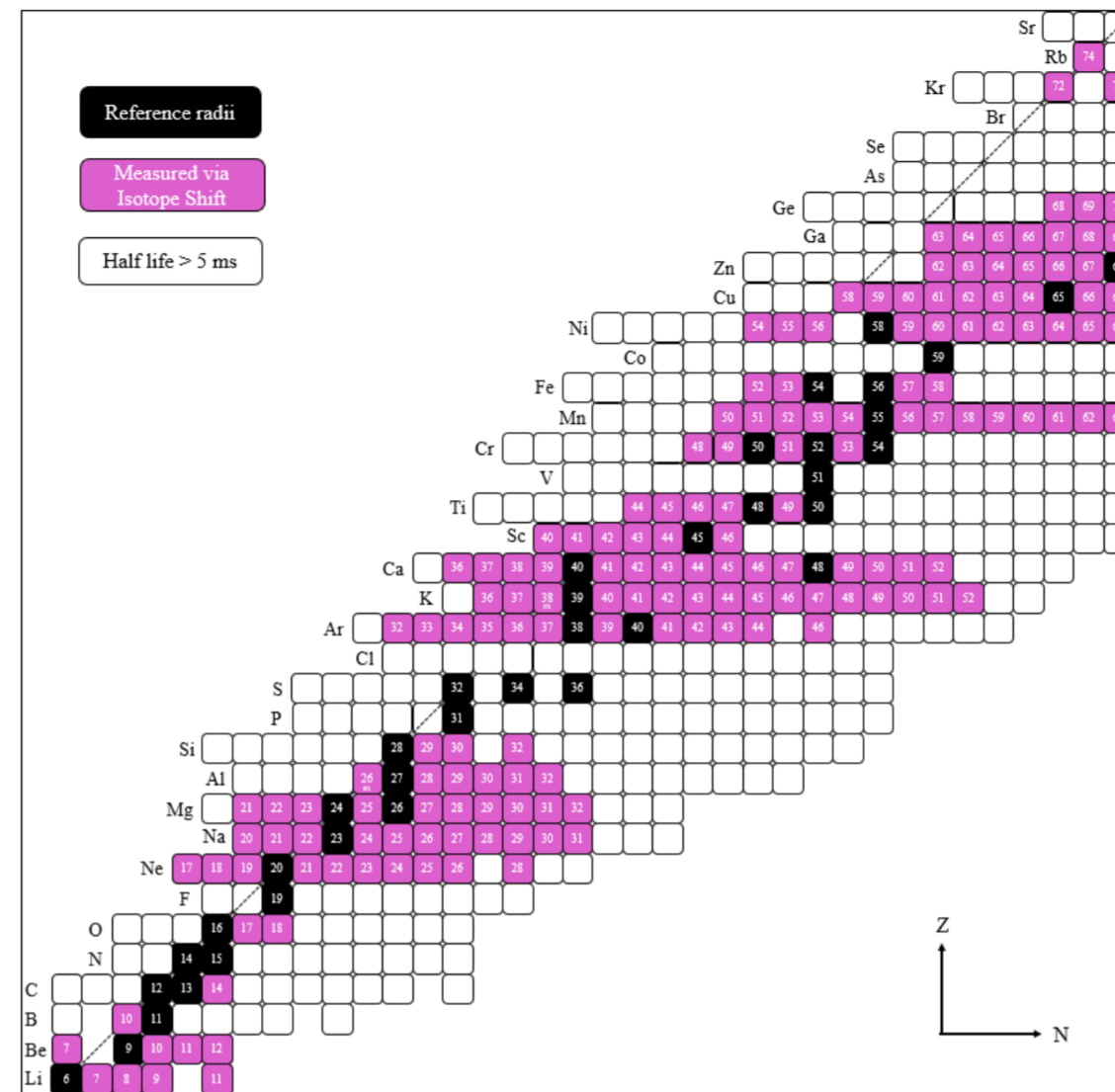
Atomic Factors (King plot / Calc.)

# Roadmap:

## 1. Reference radii (muonic atoms)



## 2. With Isotope shifts (electronic atoms)



$$r_x^2 = r_{ref}^2 + \delta r_{a,x}^2 \approx \frac{IS_{a,x} - K\mu_{a,x}}{F}$$

Isotope shift

Mass factor - Penning traps

Atomic Factors (King plot / Calc.)

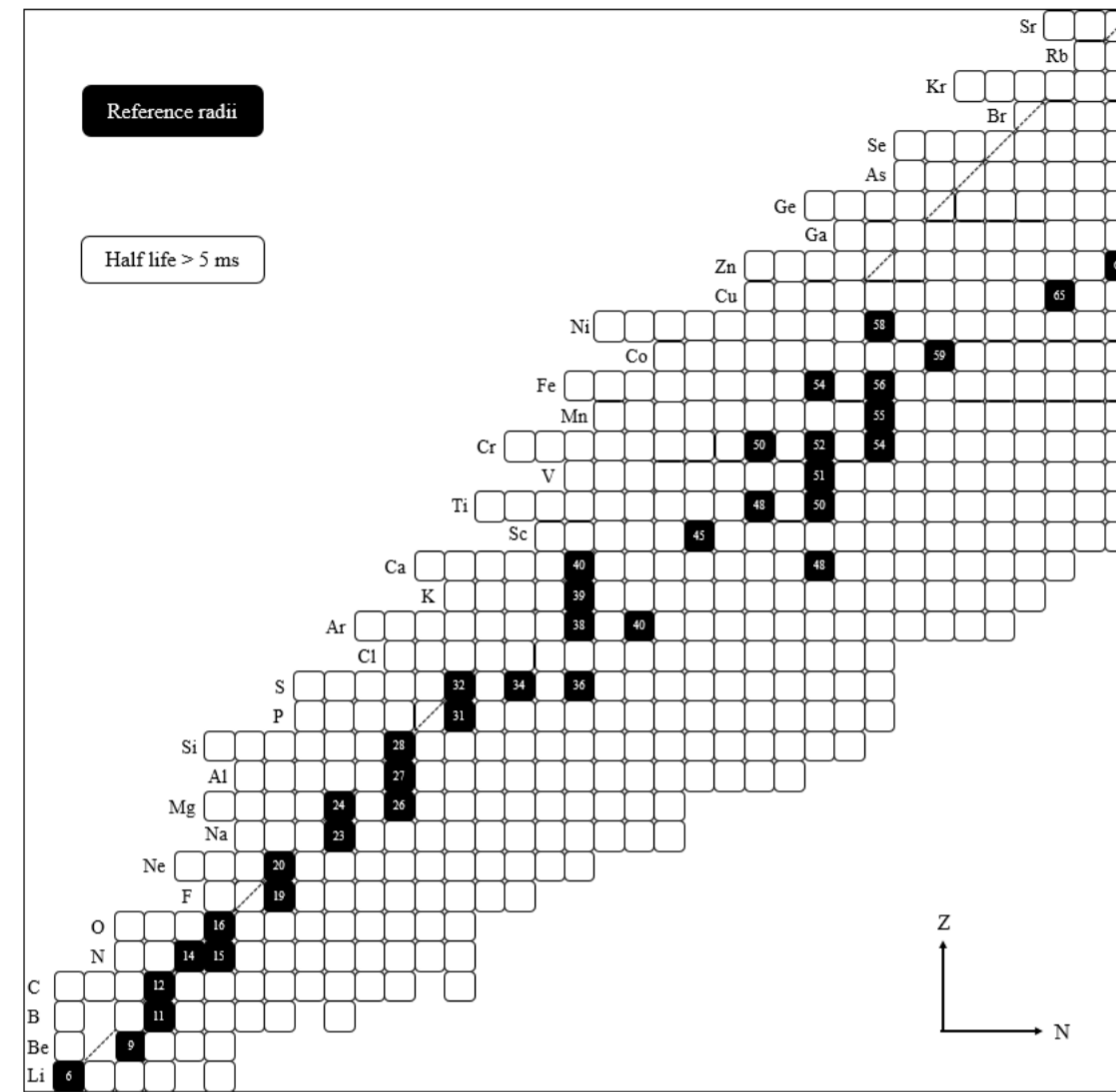
Step 2 : Update differential radii:

- Go over original publications (no compilations)
- Include some missing atomic theory uncertainties.
- Update neon chain using novel g-factor difference [measurement](#) at MPK

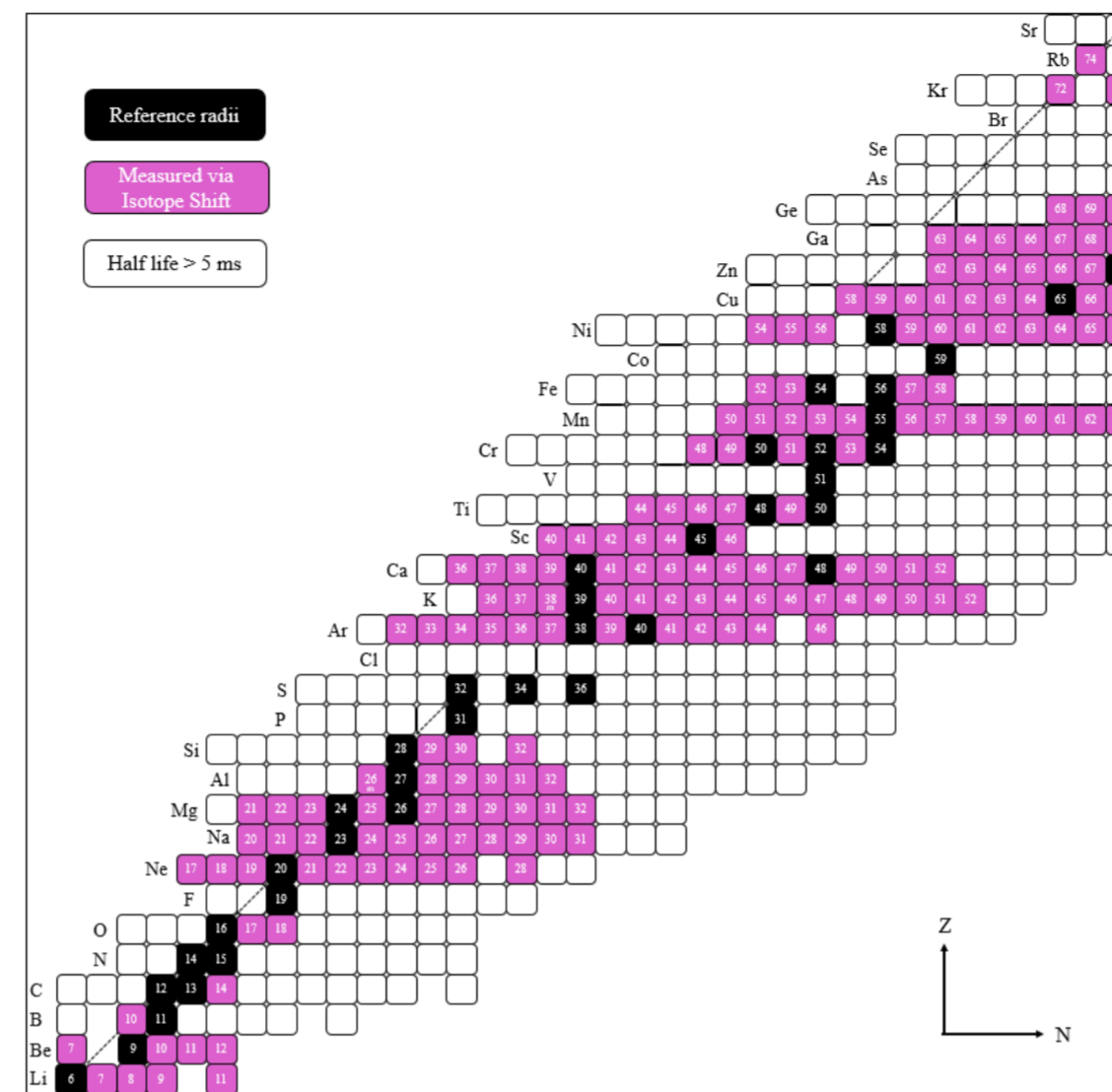


# Roadmap:

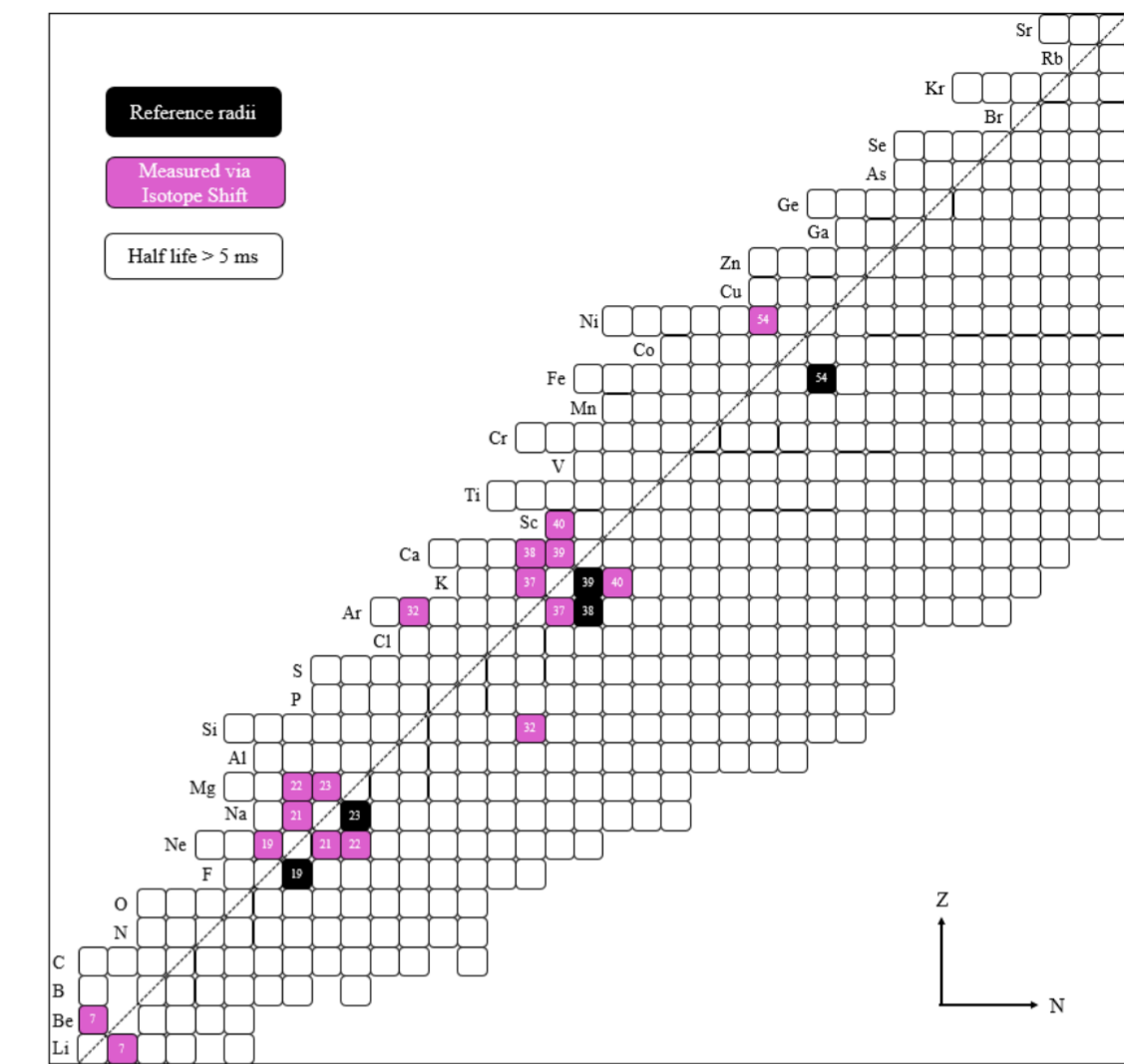
## 1. Reference radii (muonic atoms)



## 2. With Isotope shifts (electronic atoms)

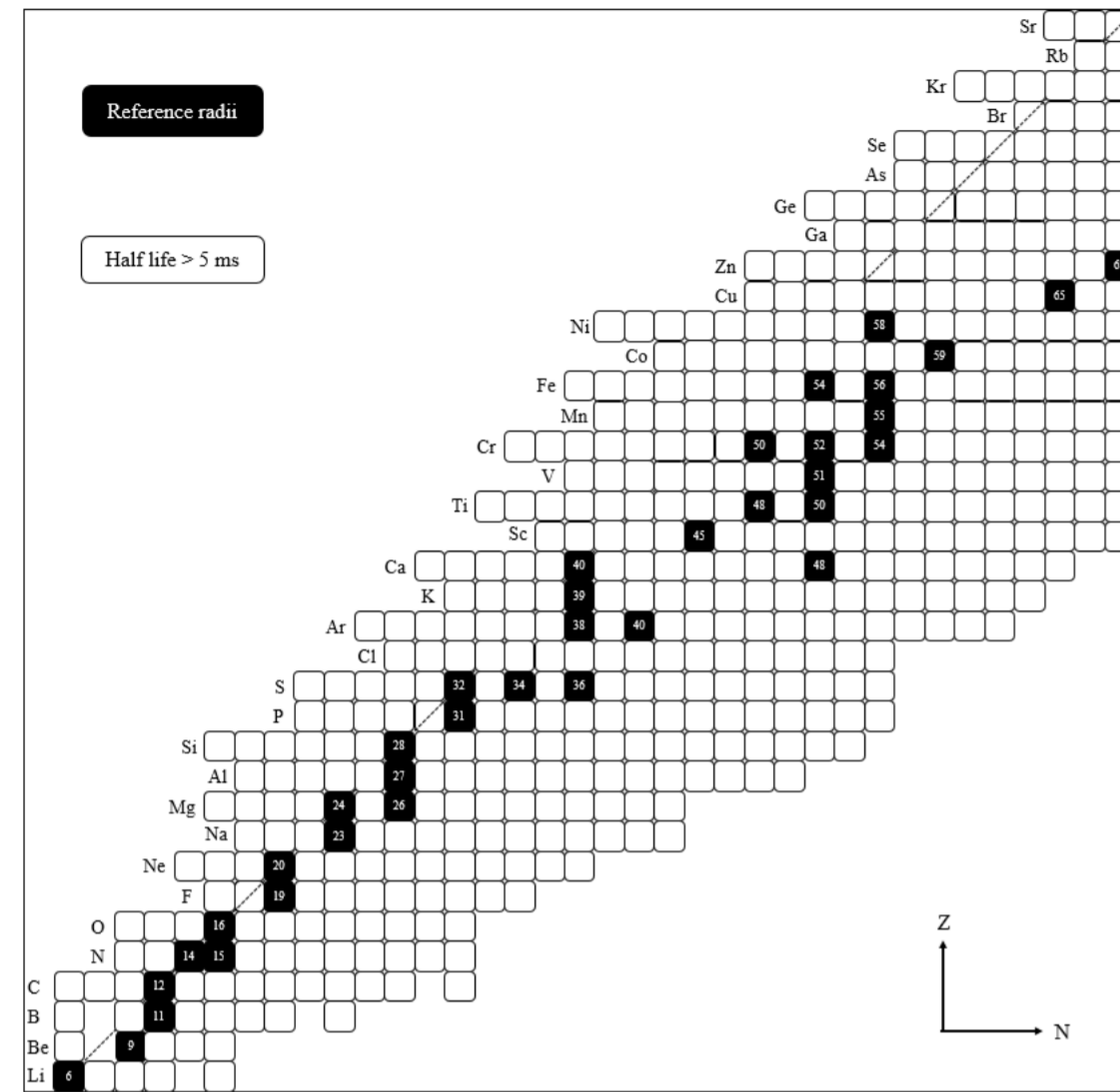


## 3. Mirror nuclei

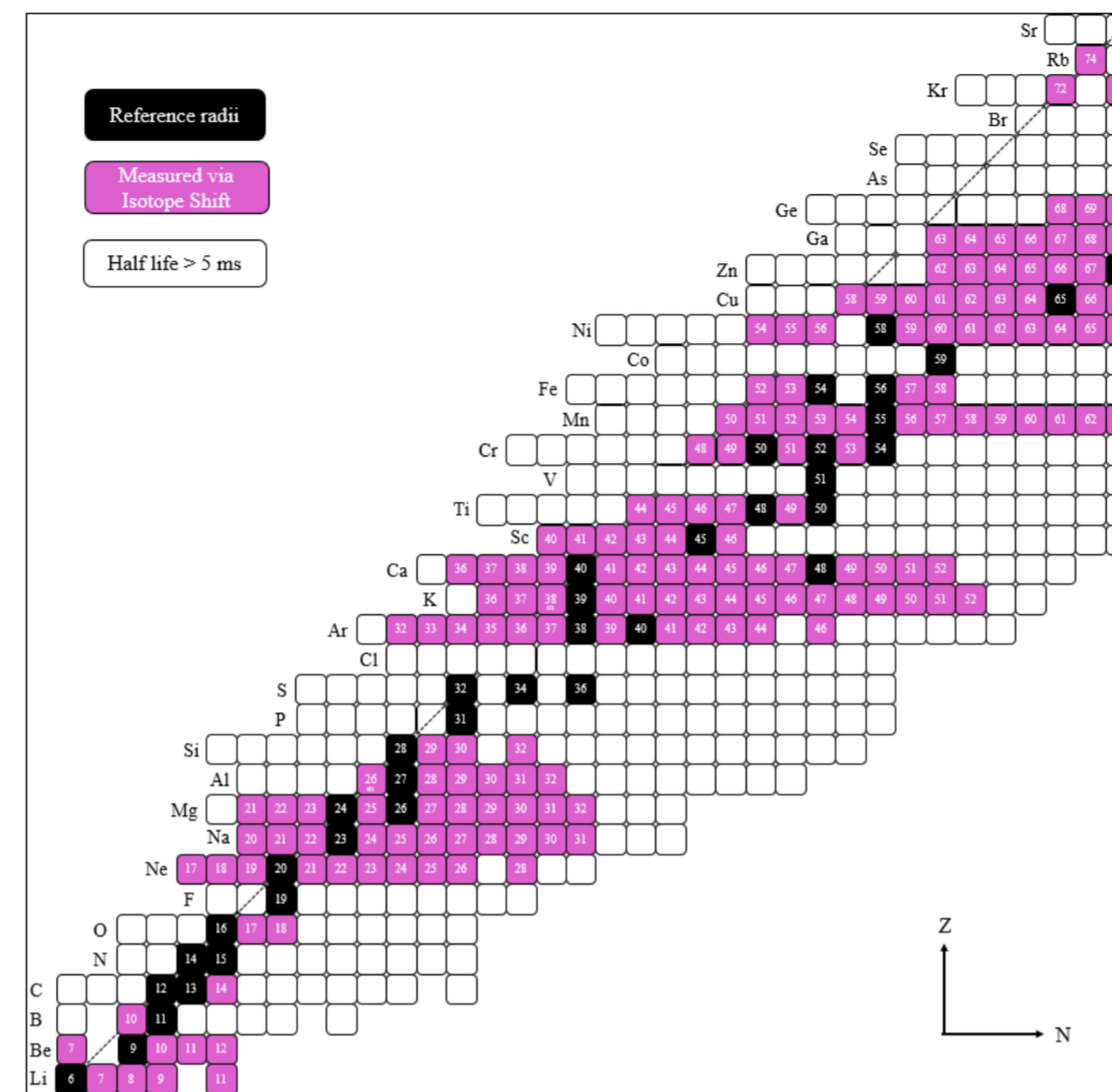


# Roadmap:

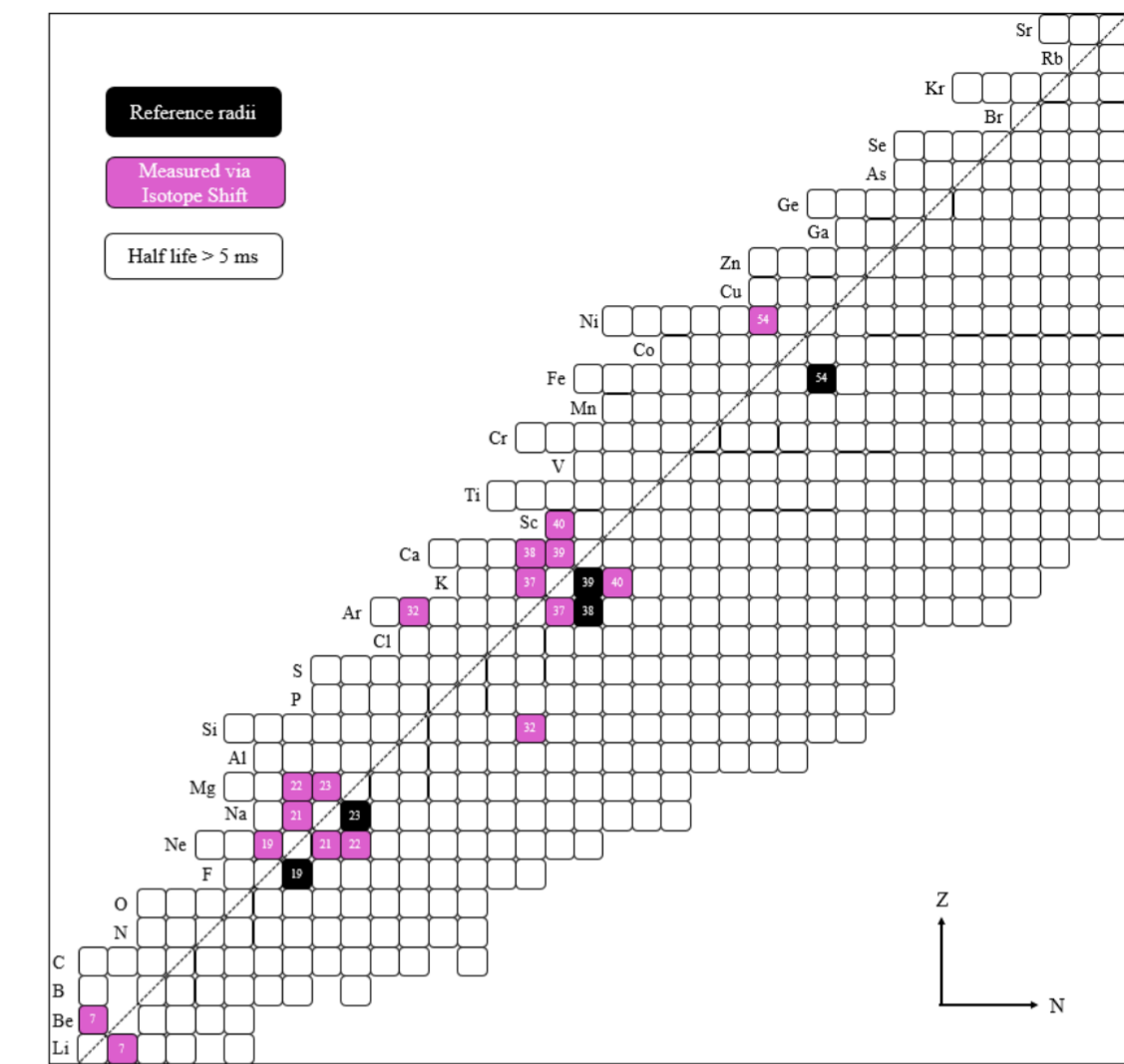
## 1. Reference radii (muonic atoms)



## 2. With Isotope shifts (electronic atoms)



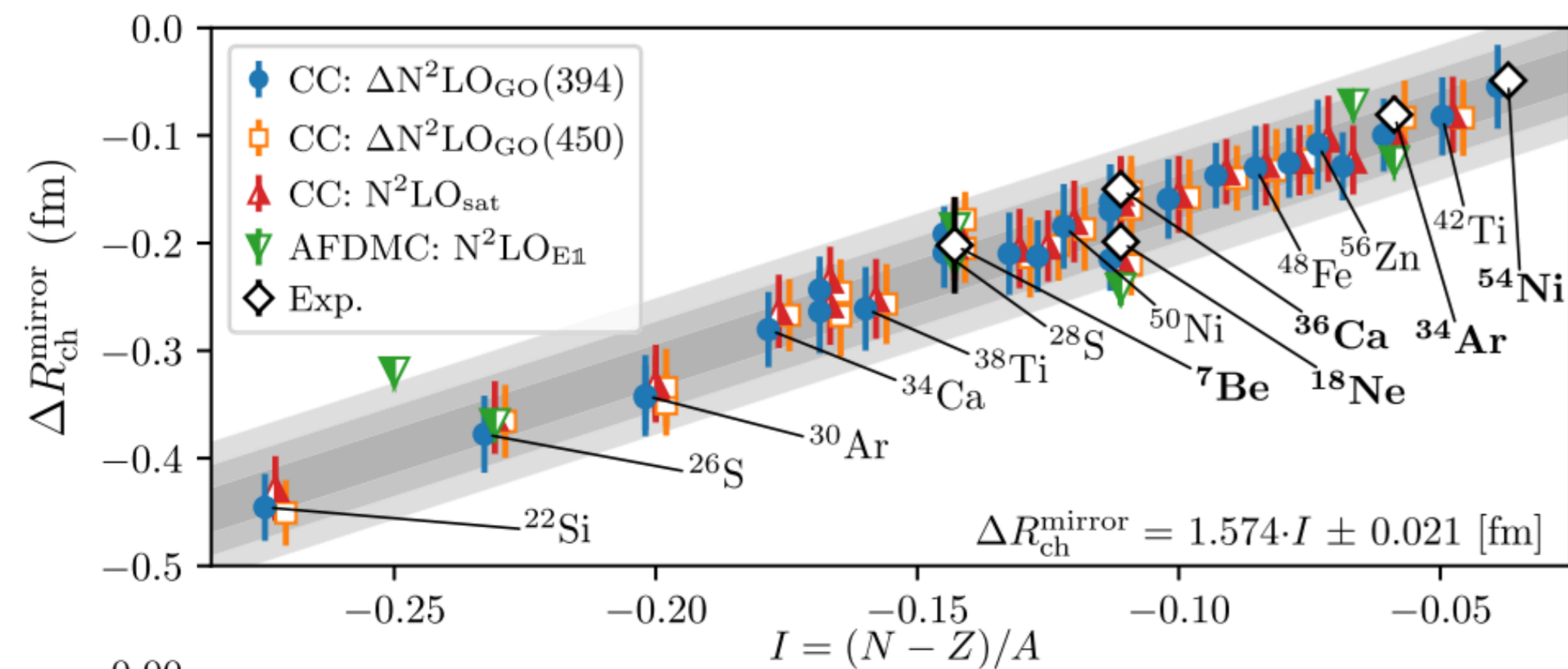
## 3. Mirror nuclei



PHYSICAL REVIEW LETTERS 130, 032501 (2023)

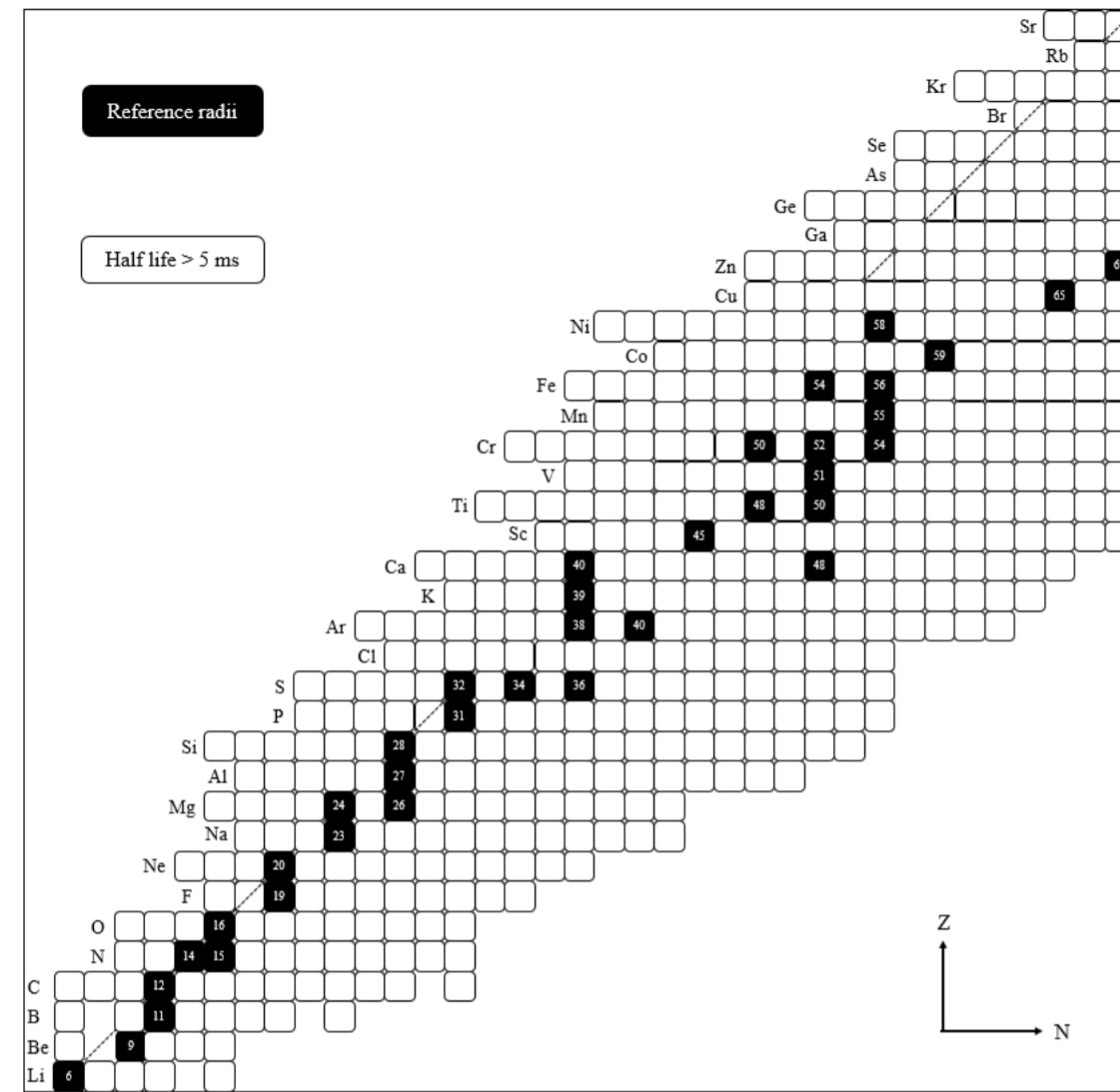
### Trends of Neutron Skins and Radii of Mirror Nuclei from First Principles

S. J. Novario<sup>1</sup>, D. Lonardoni<sup>1,\*</sup>, S. Gandolfi<sup>1</sup> and G. Hagen<sup>2,3</sup>

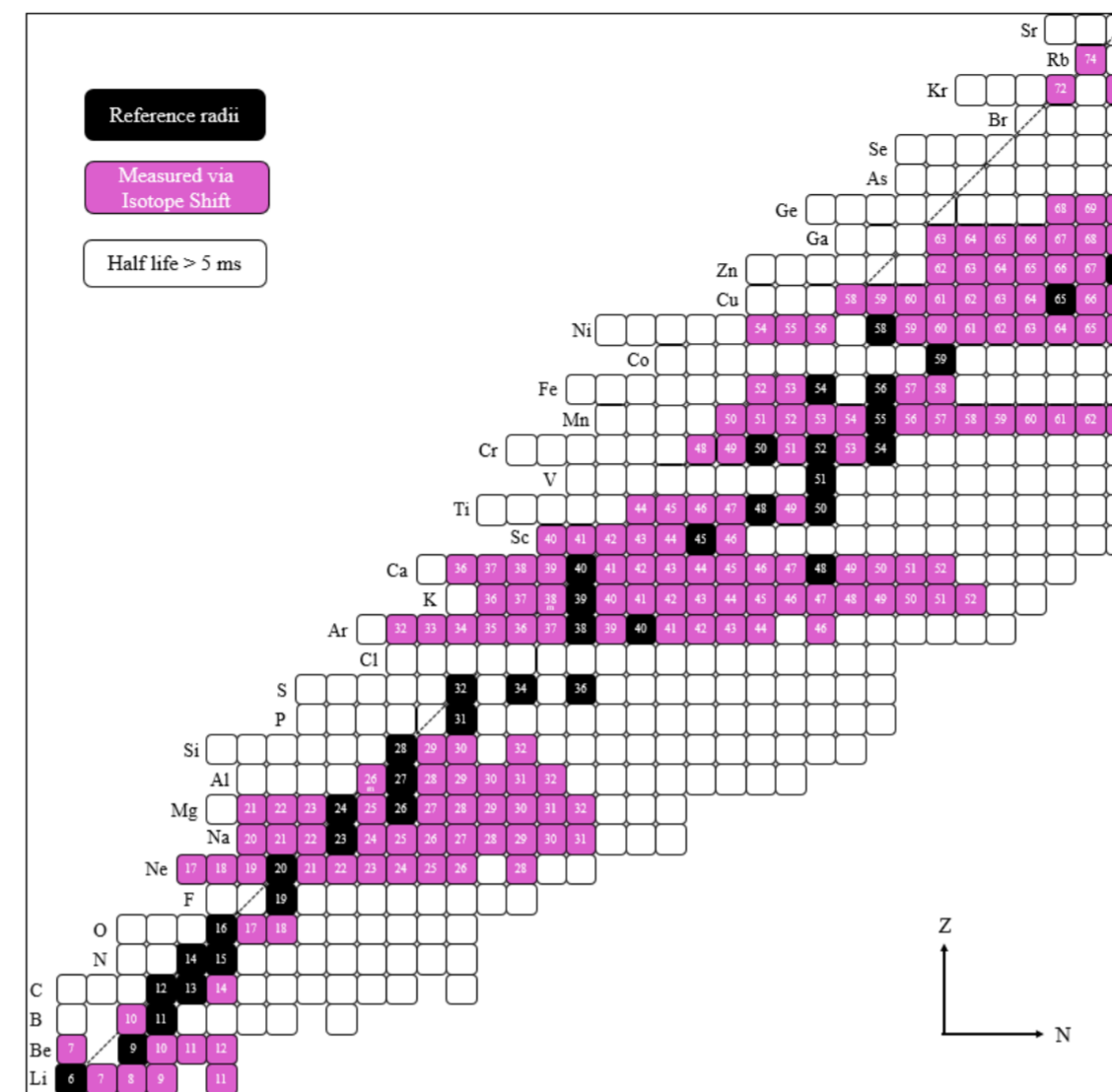


# Roadmap:

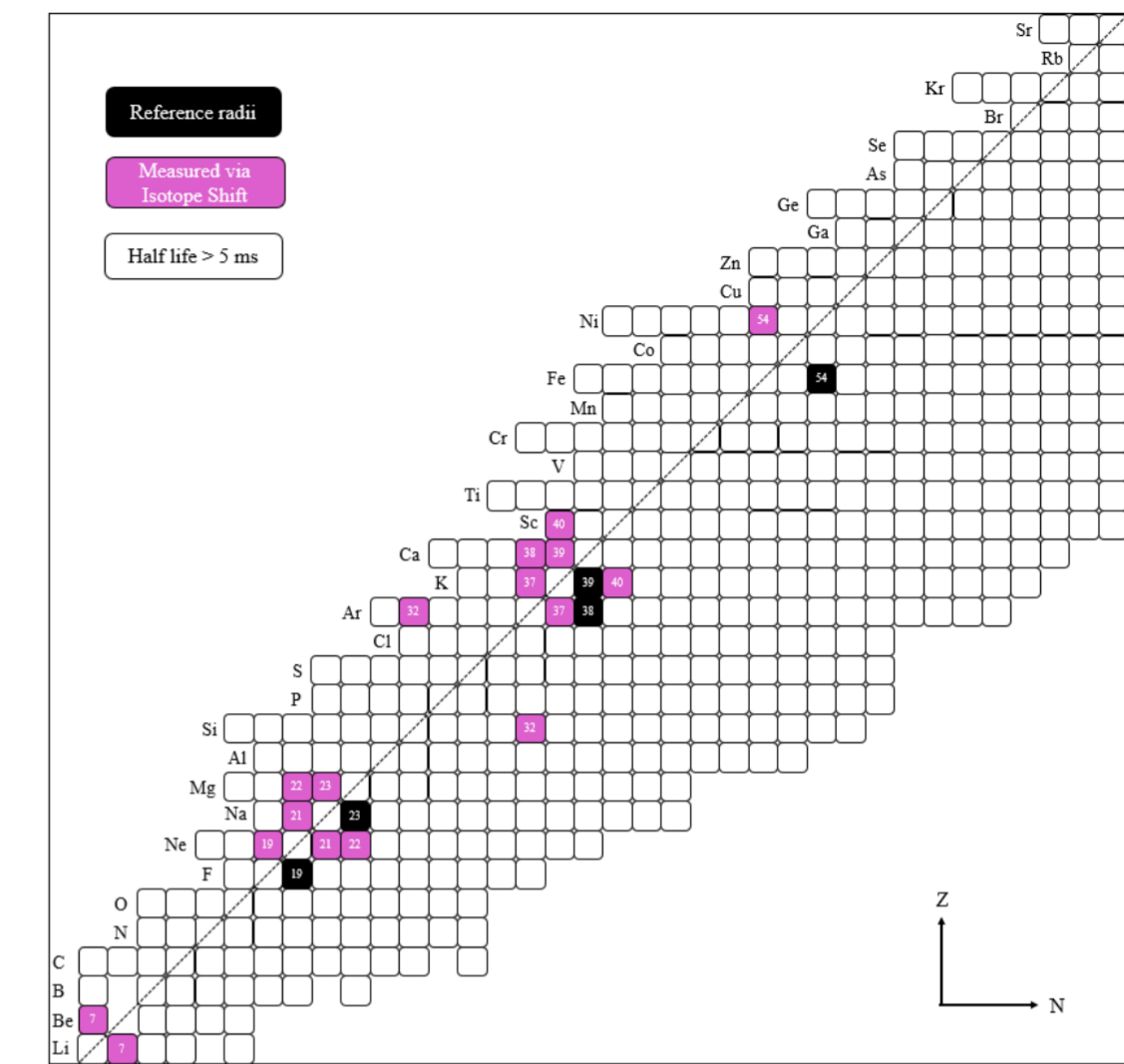
## 1. Reference radii (muonic atoms)



## 2. With Isotope shifts (electronic atoms)



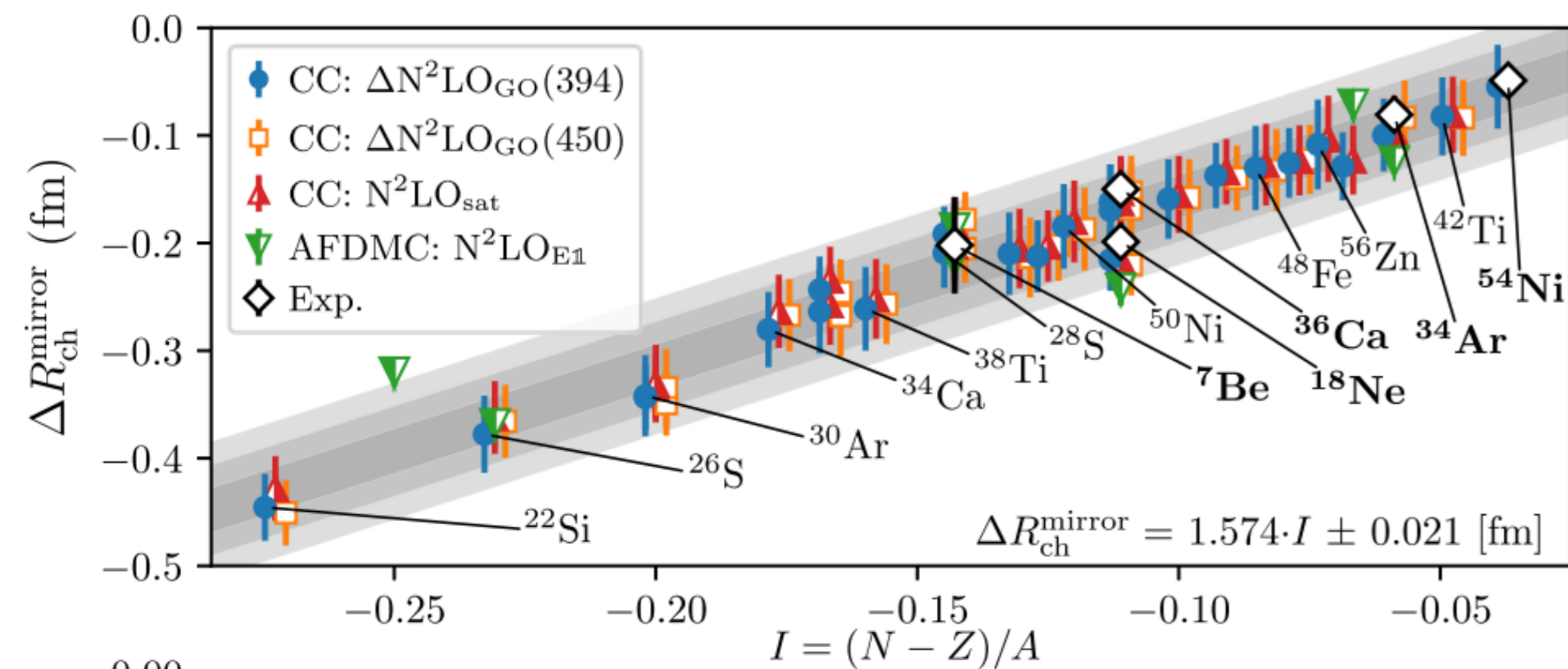
## 3. Mirror nuclei



PHYSICAL REVIEW LETTERS 130, 032501 (2023)

### Trends of Neutron Skins and Radii of Mirror Nuclei from First Principles

S. J. Novario<sup>1</sup>, D. Lonardoni<sup>1,\*</sup>, S. Gandolfi<sup>1</sup> and G. Hagen<sup>2,3</sup>



Step 3 : Update experimental data on mirror nuclei

# Updated mirror shift fit:

**Table 4**

Input mirror shifts  $\Delta_I$  to the mirror fit, based on the reference radii of Table 2, with radii differences taken from the references given in Table 5

$A$	$I$	el.	$Z$	$N$	$r$ fm	el.	$Z$	$N$	$r$ fm	$\Delta_I$ fm	$\Delta_I/I$ fm	Weight fm <sup>-2</sup>	$n\sigma$
7	0.14	Li	3	4	2.449(41)	Be	4	3	2.646(33)	0.197(53)	1.42(37)	7	0.0
18	0.11	O	8	10	2.777(07)	Ne	10	8	2.934(09)	0.161(15)	1.45(13)	96	0.3
19	0.05	F	9	10	2.902(04)	Ne	10	9	2.995(06)	0.093(07)	1.77(14)	52	2.7
22	0.09	Ne	10	12	2.948(04)	Mg	12	10	3.071(05)	0.123(07)	1.35(07)	186	0.5
23	0.04	Na	11	12	2.992(05)	Mg	12	11	3.043(04)	0.051(07)	1.17(15)	43	1.4
32	0.13	Si	14	18	3.154(13)	Ar	18	14	3.346(17)	0.193(21)	1.54(17)	36	0.9
34	0.06	S	16	18	3.284(04)	Ar	18	16	3.365(11)	0.081(12)	1.37(20)	24	0.1
36	0.11	S	16	20	3.298(04)	Ca	20	16	3.452(06)	0.155(07)	1.39(06)	270	0.0
37	0.03	Ar	18	19	3.390(07)	K	19	18	3.419(10)	0.029(13)	1.09(47)	5	0.7
38	0.05	Ar	18	20	3.402(06)	Ca	20	18	3.469(04)	0.067(08)	1.28(15)	47	0.8
39	0.03	K	19	20	3.435(04)	Ca	20	19	3.464(04)	0.029(06)	1.14(22)	21	1.2
54	0.04	Fe	26	28	3.688(04)	Ni	28	26	3.741(05)	0.053(06)	1.44(17)	34	0.3

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)

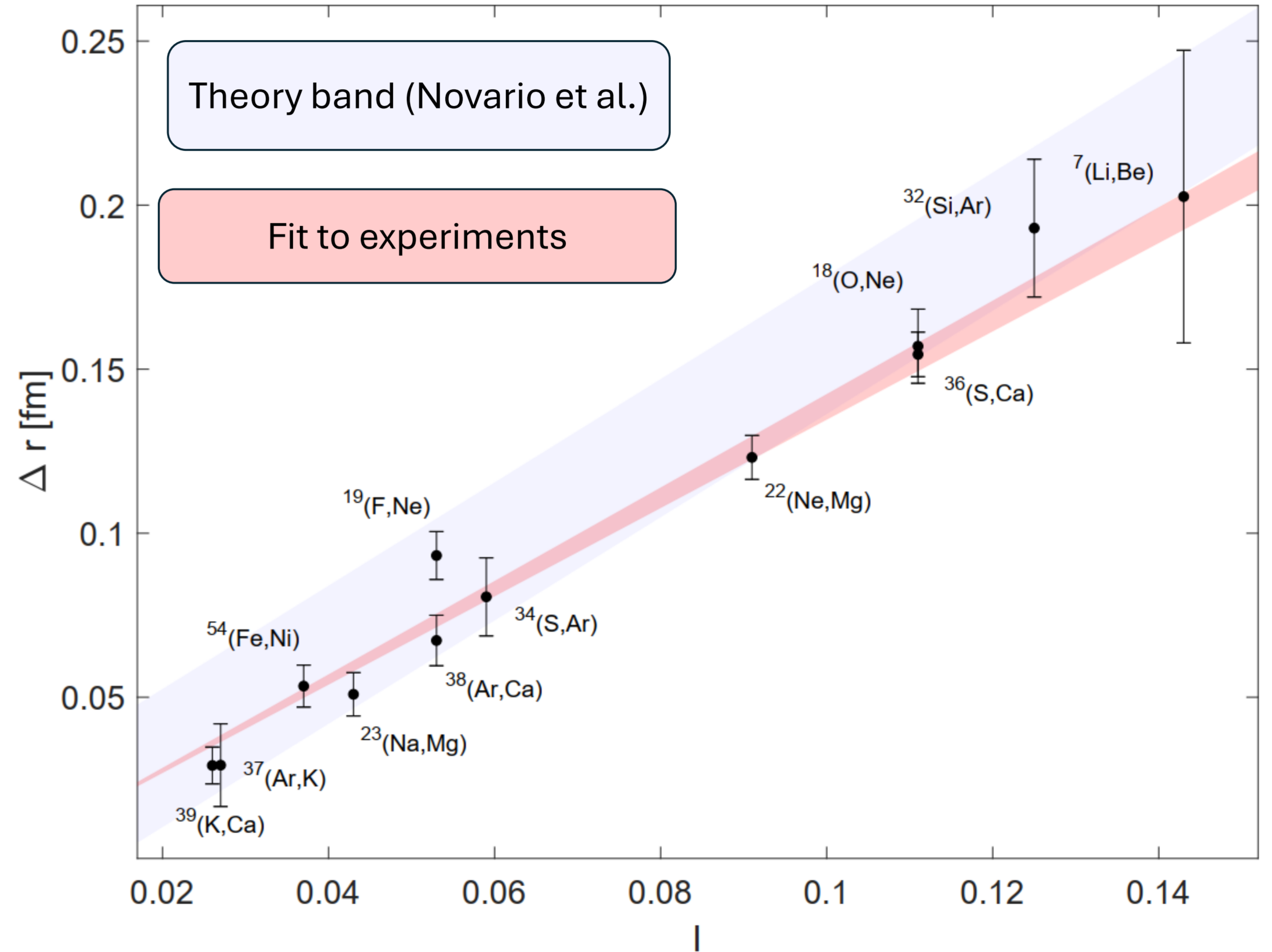
# Updated mirror shift fit:

**Table 4**

Input mirror shifts  $\Delta_I$  to the mirror fit, based on the reference radii of Table 2, with radii differences taken from the references given in Table 5

$A$	$I$	el.	$Z$	$N$	$r$ fm	el.	$Z$	$N$	$r$ fm	$\Delta_I$ fm	$\Delta_I/I$ fm	Weight fm <sup>-2</sup>	$n\sigma$
7	0.14	Li	3	4	2.449(41)	Be	4	3	2.646(33)	0.197(53)	1.42(37)	7	0.0
18	0.11	O	8	10	2.777(07)	Ne	10	8	2.934(09)	0.161(15)	1.45(13)	96	0.3
19	0.05	F	9	10	2.902(04)	Ne	10	9	2.995(06)	0.093(07)	1.77(14)	52	2.7
22	0.09	Ne	10	12	2.948(04)	Mg	12	10	3.071(05)	0.123(07)	1.35(07)	186	0.5
23	0.04	Na	11	12	2.992(05)	Mg	12	11	3.043(04)	0.051(07)	1.17(15)	43	1.4
32	0.13	Si	14	18	3.154(13)	Ar	18	14	3.346(17)	0.193(21)	1.54(17)	36	0.9
34	0.06	S	16	18	3.284(04)	Ar	18	16	3.365(11)	0.081(12)	1.37(20)	24	0.1
36	0.11	S	16	20	3.298(04)	Ca	20	16	3.452(06)	0.155(07)	1.39(06)	270	0.0
37	0.03	Ar	18	19	3.390(07)	K	19	18	3.419(10)	0.029(13)	1.09(47)	5	0.7
38	0.05	Ar	18	20	3.402(06)	Ca	20	18	3.469(04)	0.067(08)	1.28(15)	47	0.8
39	0.03	K	19	20	3.435(04)	Ca	20	19	3.464(04)	0.029(06)	1.14(22)	21	1.2
54	0.04	Fe	26	28	3.688(04)	Ni	28	26	3.741(05)	0.053(06)	1.44(17)	34	0.3

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)

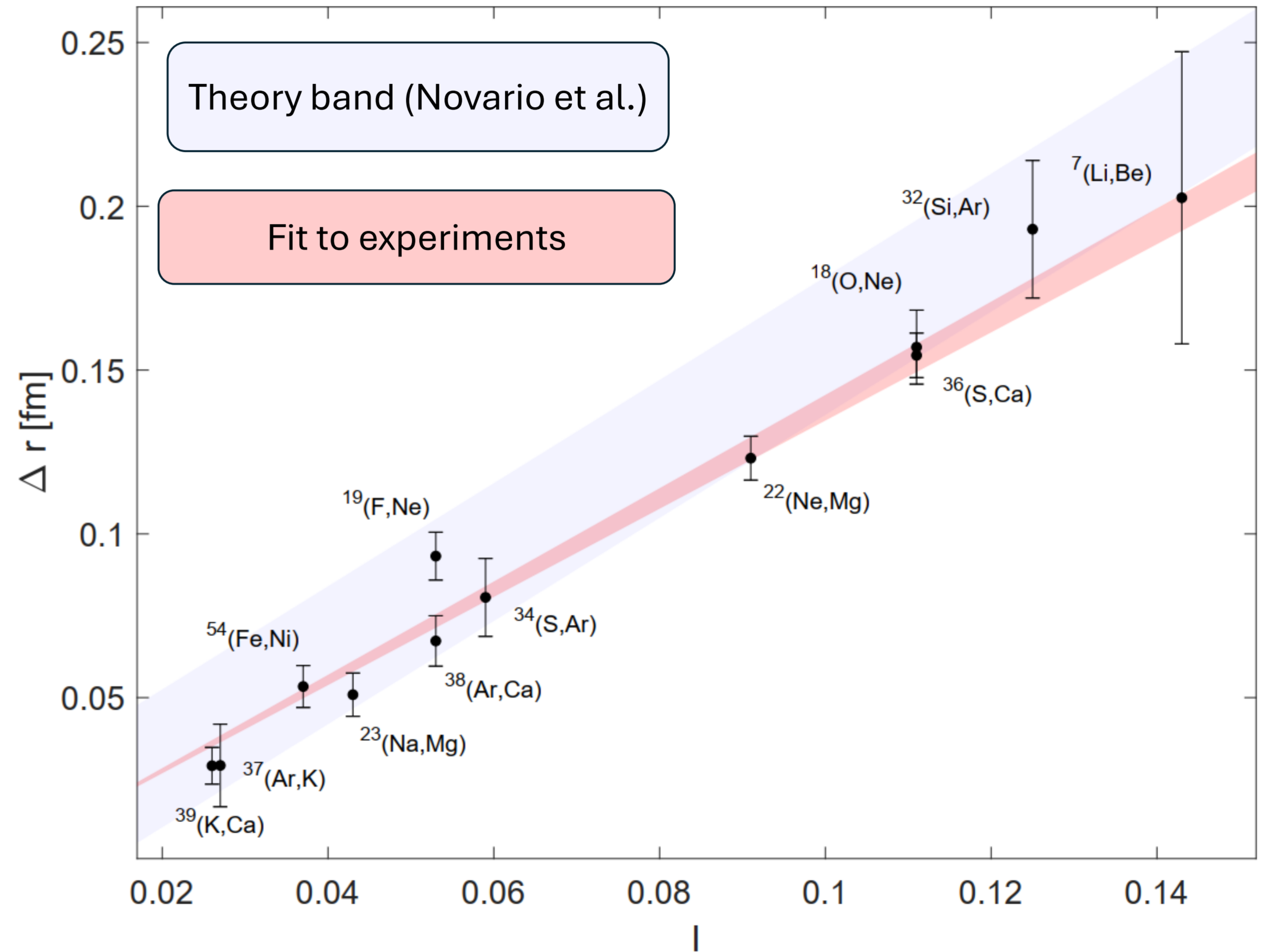


# Step 4 - Updated mirror shift fit:

Fit result:

$$\Delta R_{ch}^{mirr}(I) = 1.38(3) \times I$$

- Agrees with theory band where these is accurate data, up to  $I = 0.11$
- But missing accuracy for  $0.11 < I$

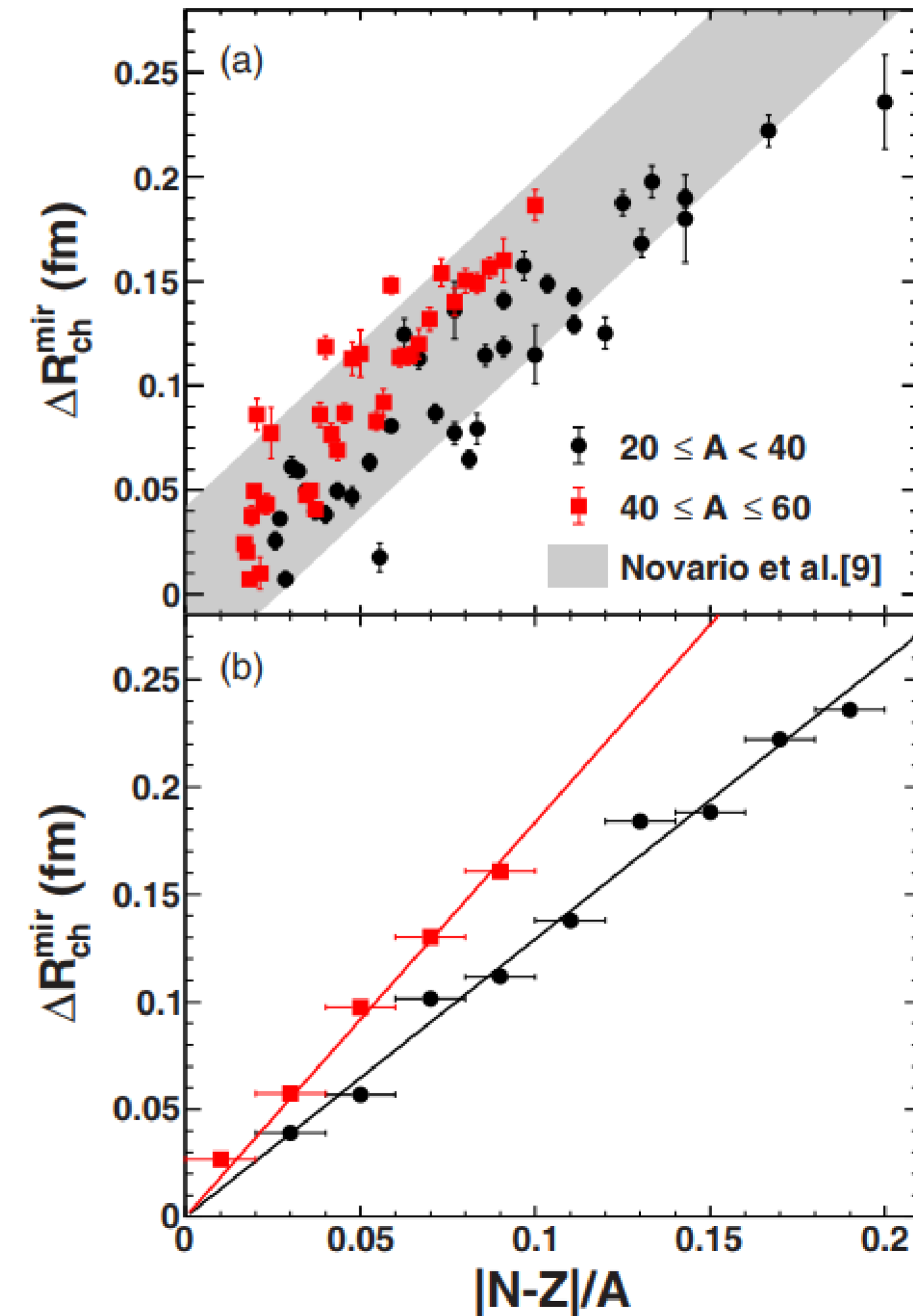


# Open questions:

- Does linearity hold at large asymmetry  $I$  ?
- Does the slope depend on mass?

## Investigation of the difference in charge radii of mirror pairs with deep Bayesian neural networks

X. Zhang (张鑫)<sup>1</sup>, H. He (何红斌)<sup>2</sup>, G. Qu (曲国峰)<sup>1</sup>, X. Liu (刘星泉)<sup>1,\*</sup>, H. Zheng (郑华)<sup>2,†</sup>,  
W. Lin (林炜平)<sup>1</sup>, J. Han (韩纪锋)<sup>1</sup>, P. Ren (任培培)<sup>1</sup> and R. Wada<sup>3</sup>

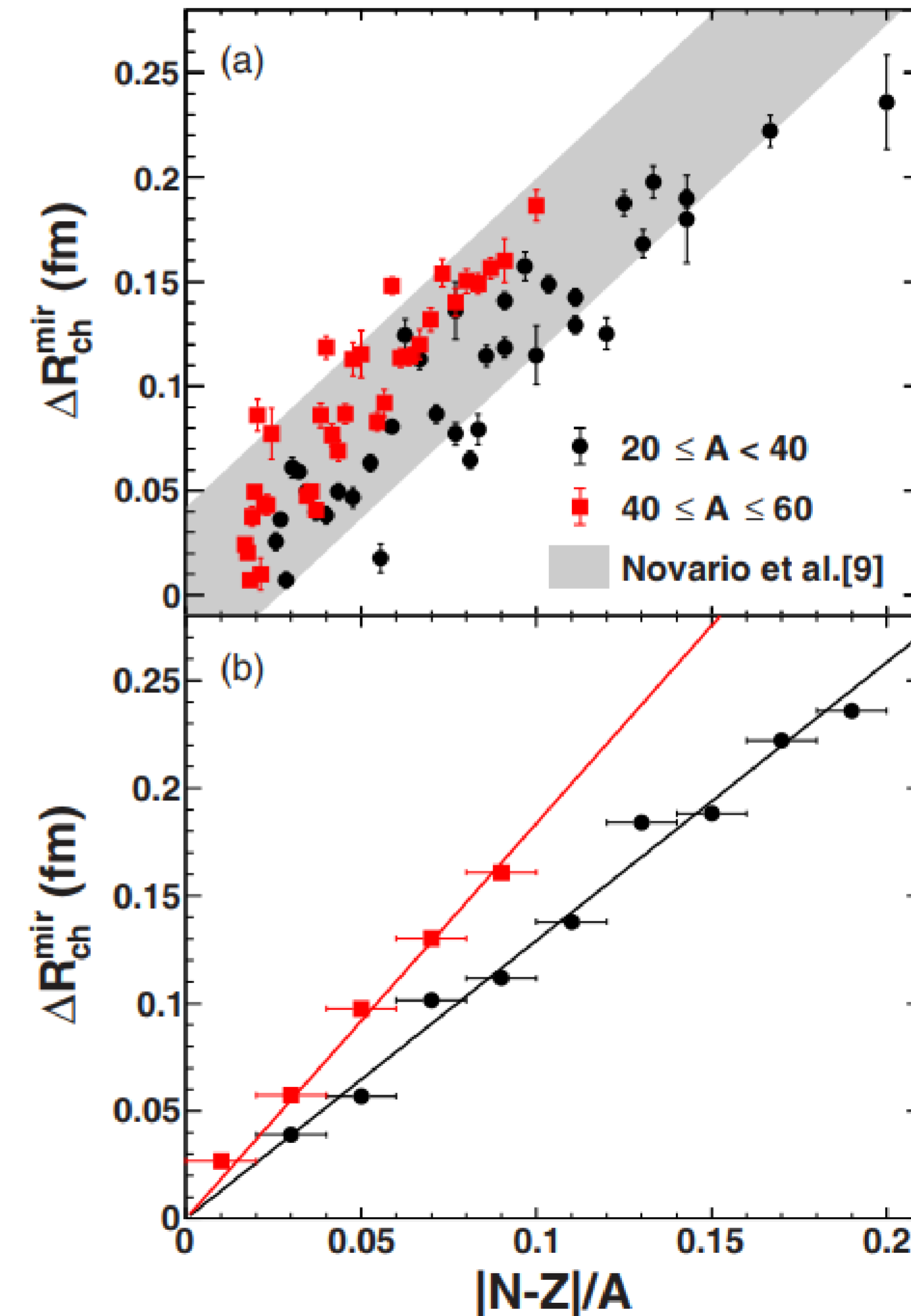


# Open questions:

- Does linearity hold at large asymmetry  $I$  ?
- Does the slope depend on mass?
- **Light nuclei are key**

## Investigation of the difference in charge radii of mirror pairs with deep Bayesian neural networks

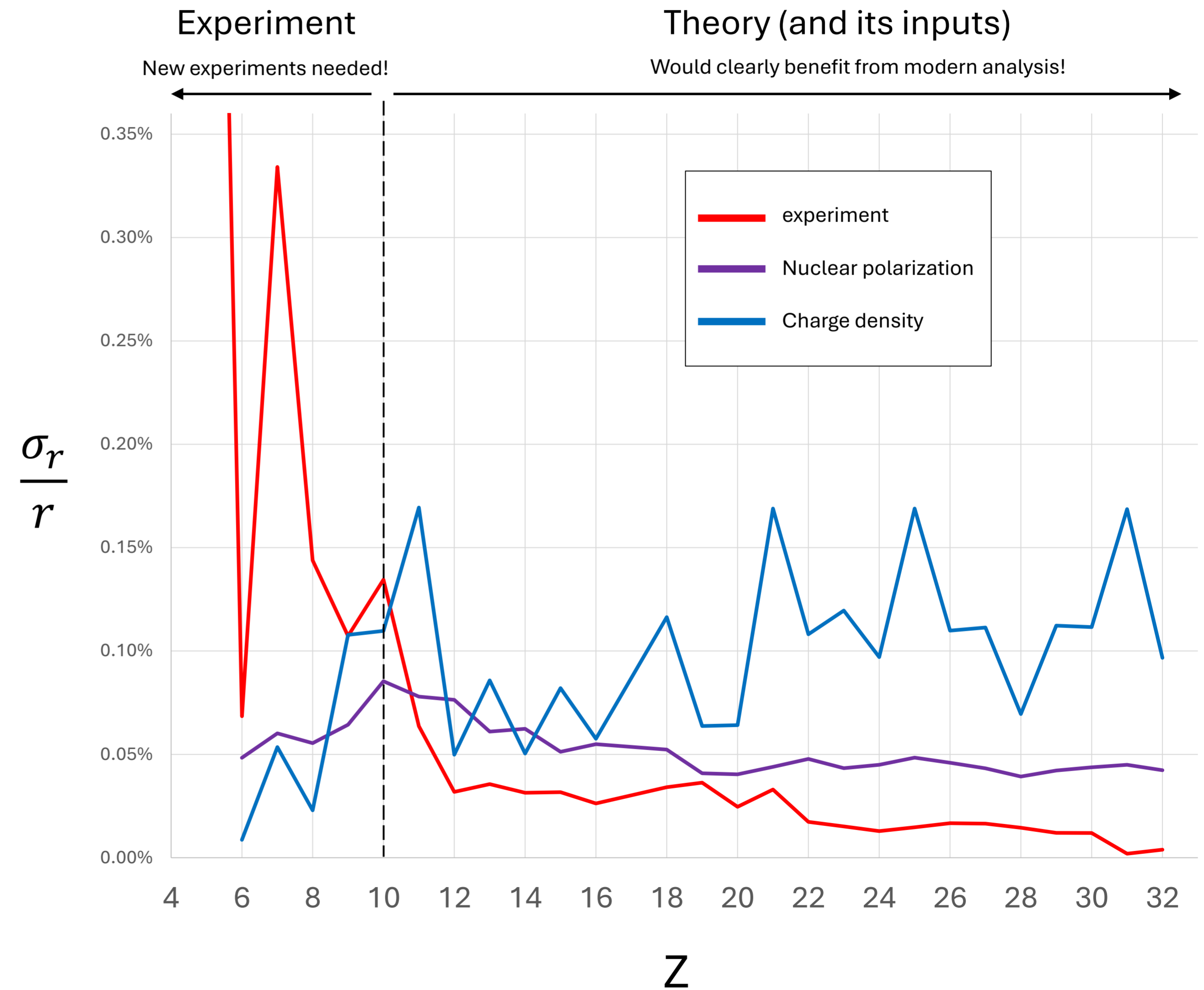
X. Zhang (张鑫)<sup>1</sup>, H. He (何红斌)<sup>2</sup>, G. Qu (曲国峰)<sup>1</sup>, X. Liu (刘星泉)<sup>1,\*</sup>, H. Zheng (郑华)<sup>2,†</sup>,  
W. Lin (林炜平)<sup>1</sup>, J. Han (韩纪锋)<sup>1</sup>, P. Ren (任培培)<sup>1</sup> and R. Wada<sup>3</sup>





# Sources of uncertainty :

Radii of light nuclei from muonic atom x-ray spec.

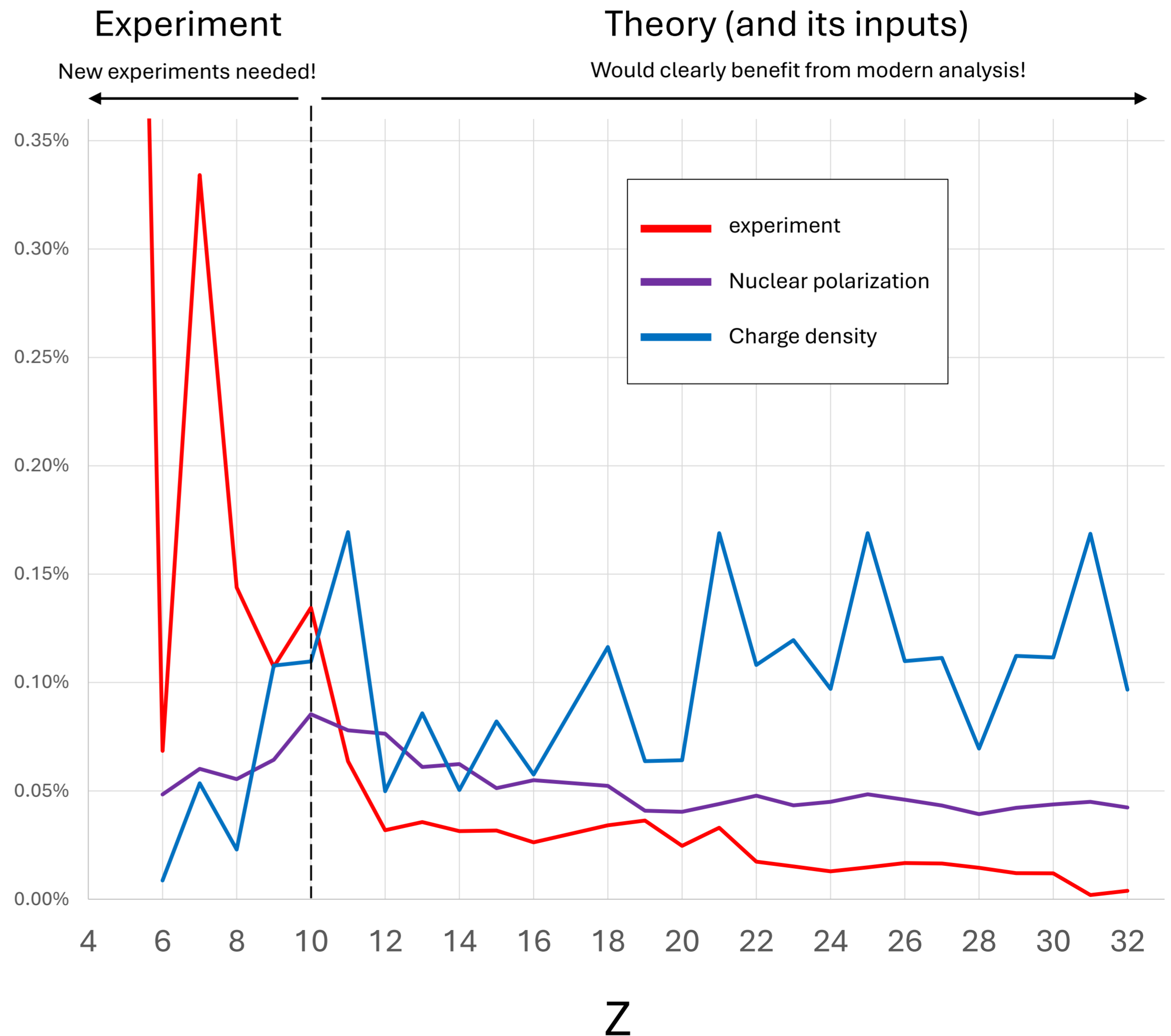


# Sources of uncertainty :

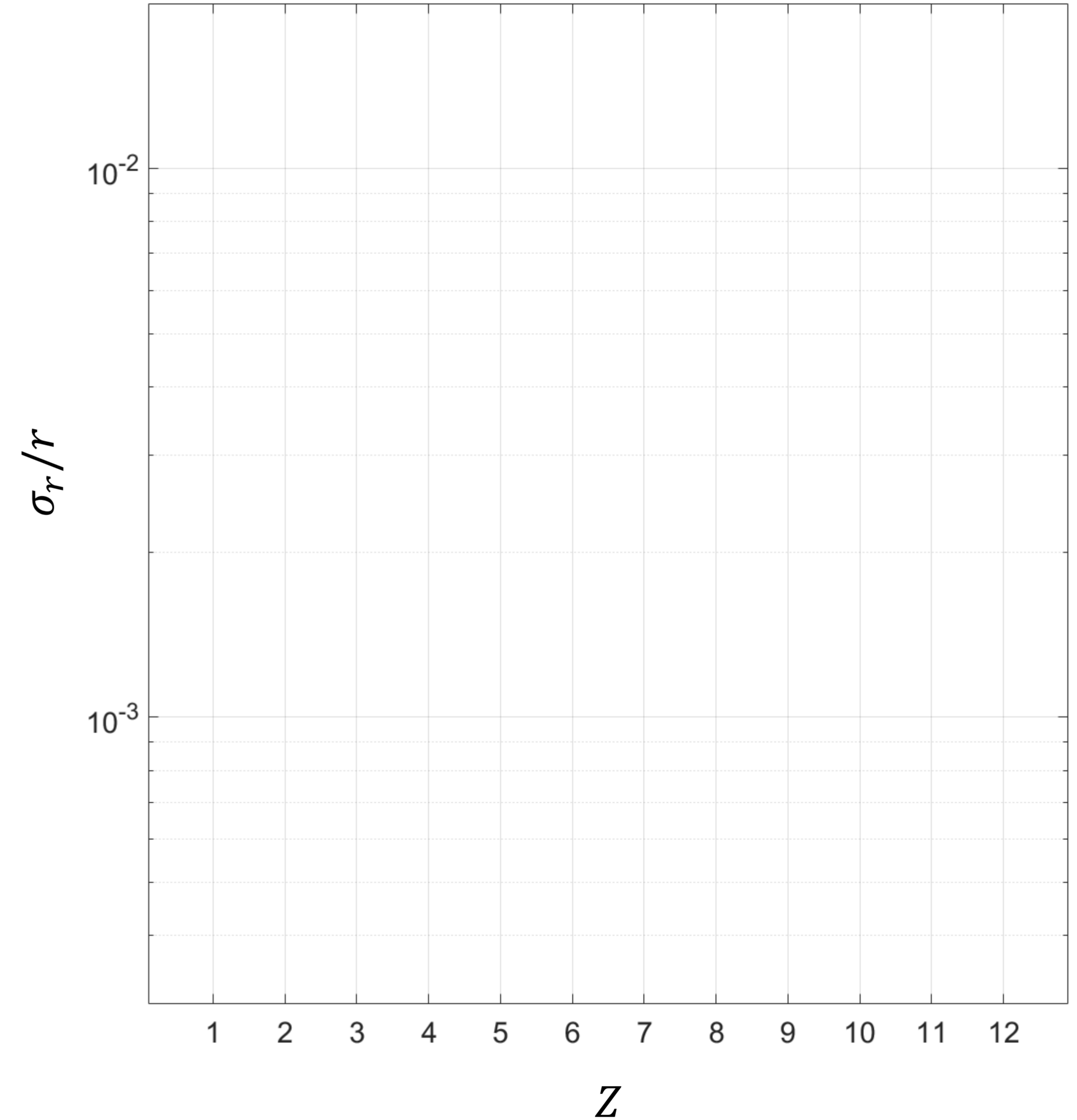
Radii of light nuclei from muonic atom x-ray spec.

What is the spike in  
exp. uncertainty < Z=11?

$$\frac{\sigma_r}{r}$$

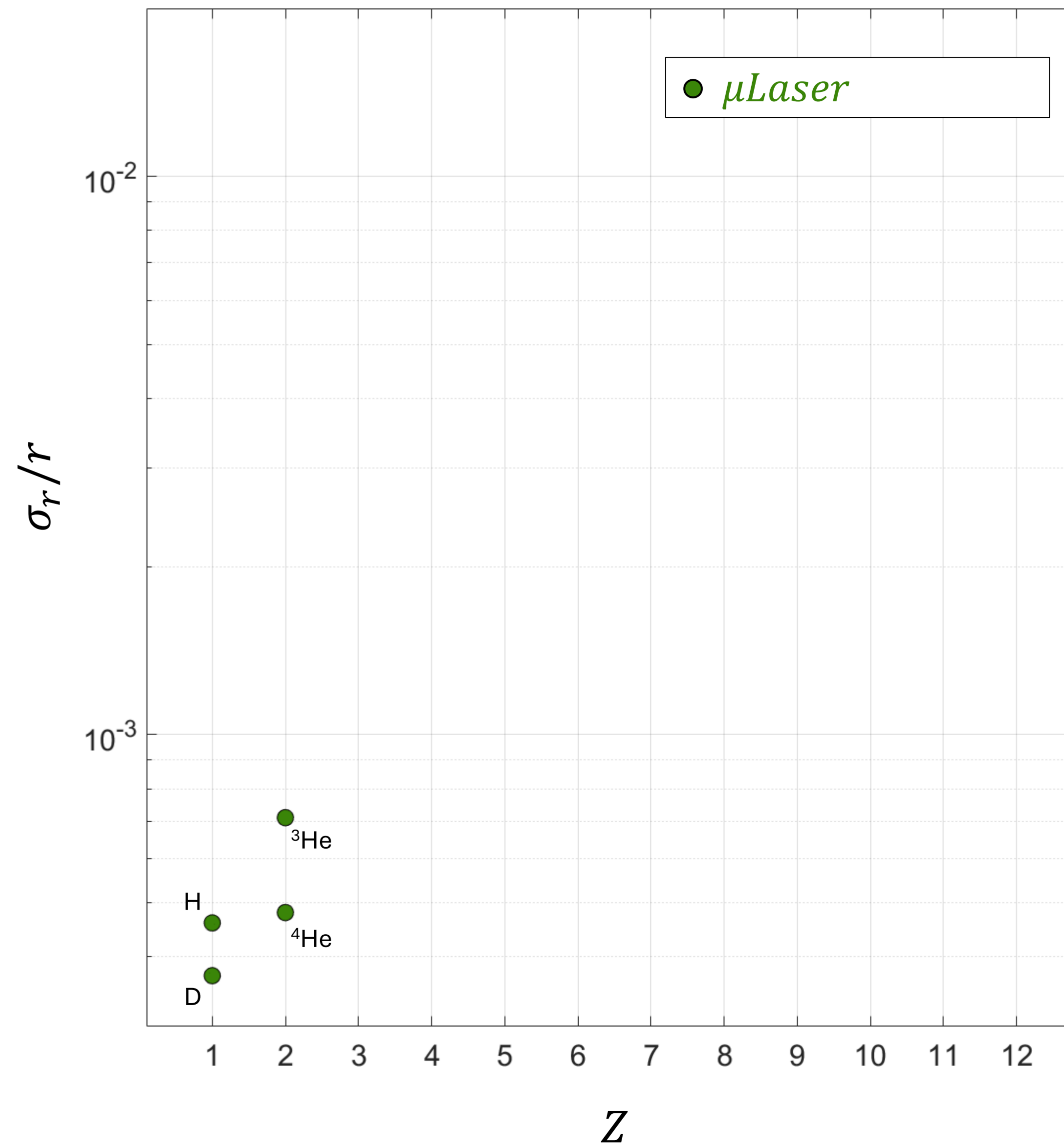


# The radius gap



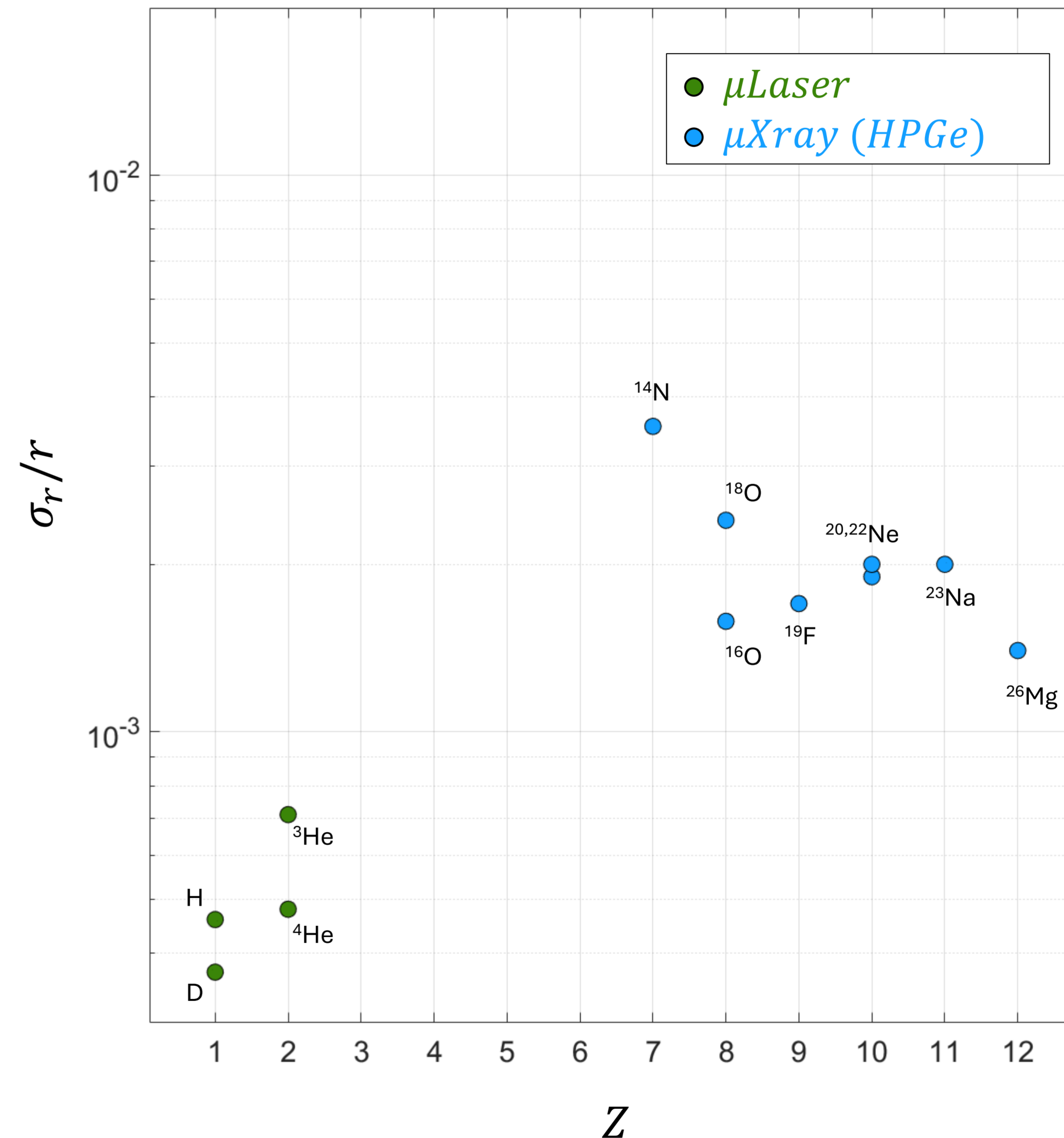
# The radius gap

- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory



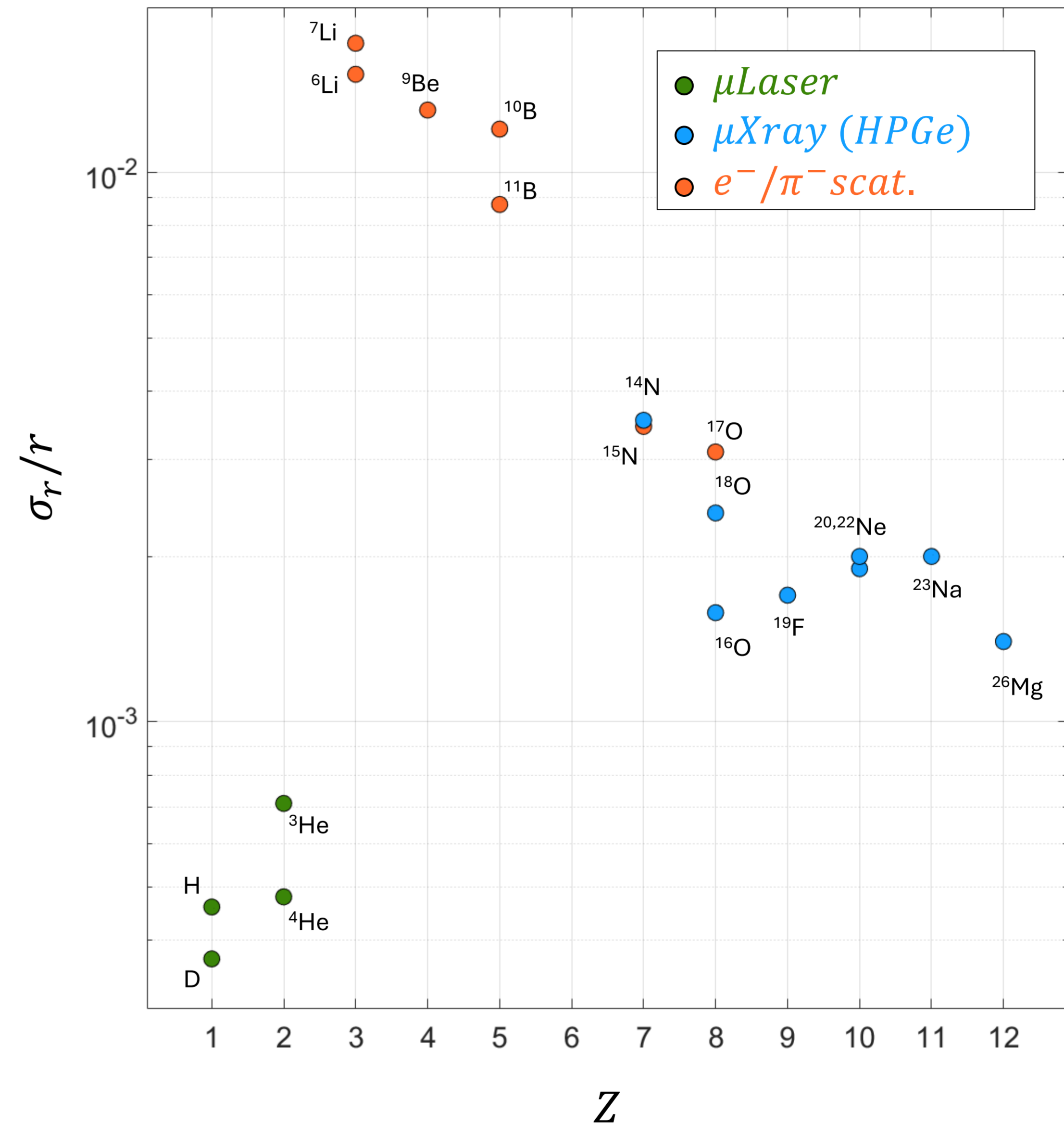
# The radius gap

- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory
- **For  $Z > 6$ :**  
Measured x-rays from muonic atoms using solid-state detectors.



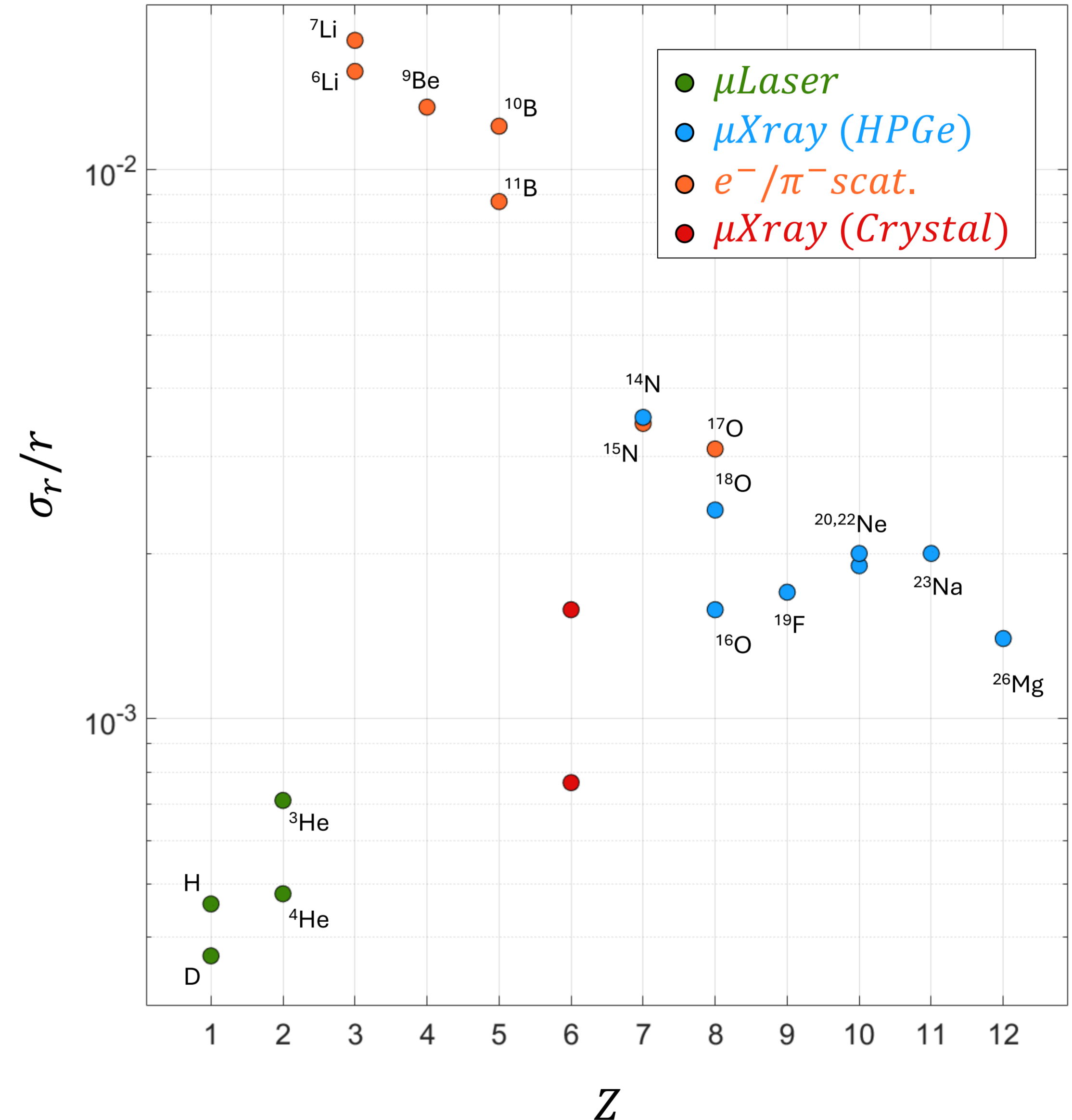
# The radius gap

- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory
- **For  $Z > 6$ :**  
Measured x-rays from muonic atoms using solid-state detectors.
- **For  $Z = 3 - 5$ , and others:**  
Electron scattering



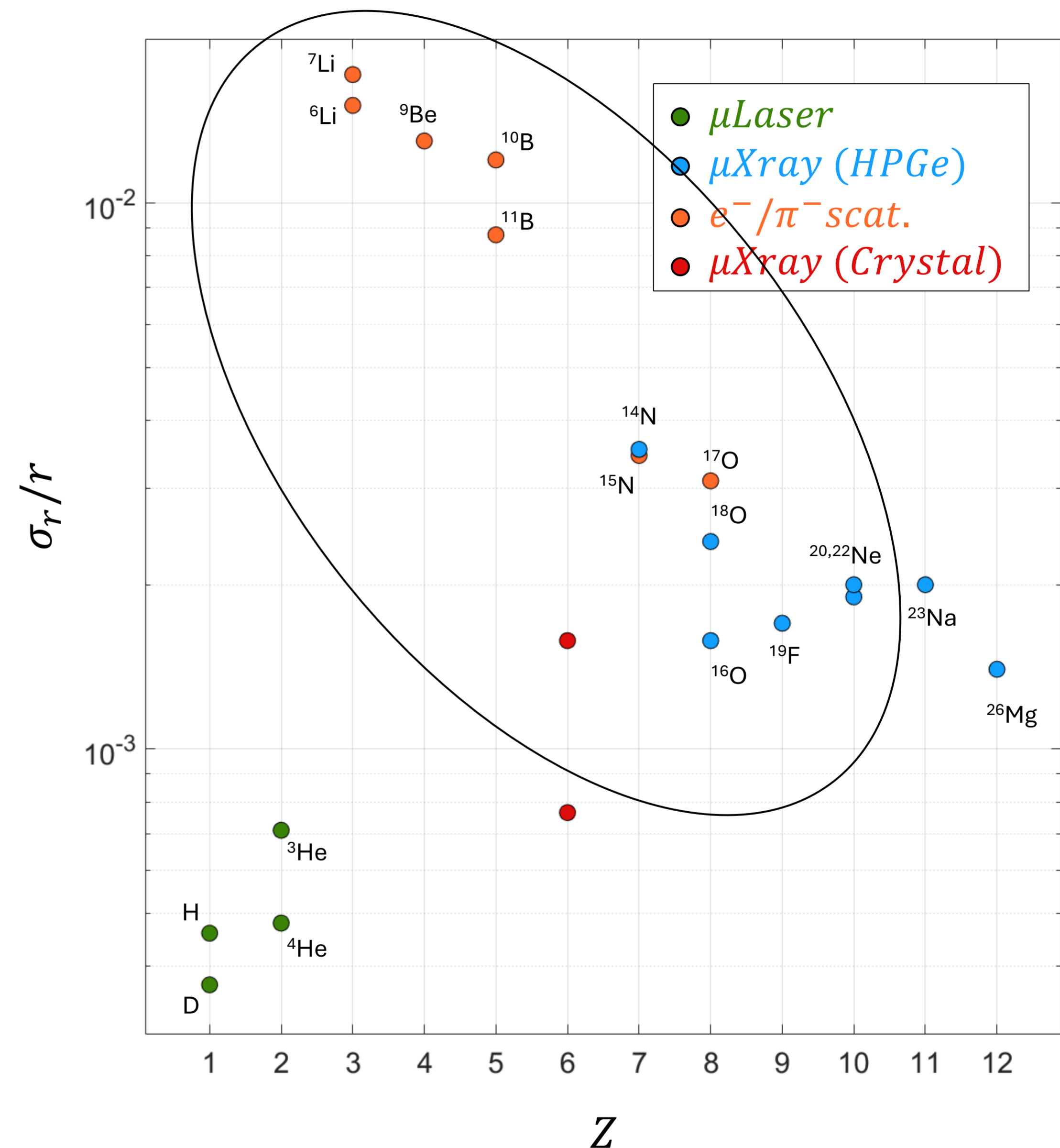
# The radius gap

- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory
- **For  $Z > 6$ :**  
Measured x-rays from muonic atoms using solid-state detectors.
- **For  $Z = 3 - 5$ , and others:**  
Electron scattering
- **For  $Z = 6$**   
Measured with crystal spectrometer. Not widely applicable



# The radius gap

- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory
- **For  $Z > 6$ :**  
Measured x-rays from muonic atoms using solid-state detectors.
- **For  $Z = 3 - 5$ , and others:**  
Electron scattering
- **For  $Z = 6$**   
Measured with crystal spectrometer. Not widely applicable

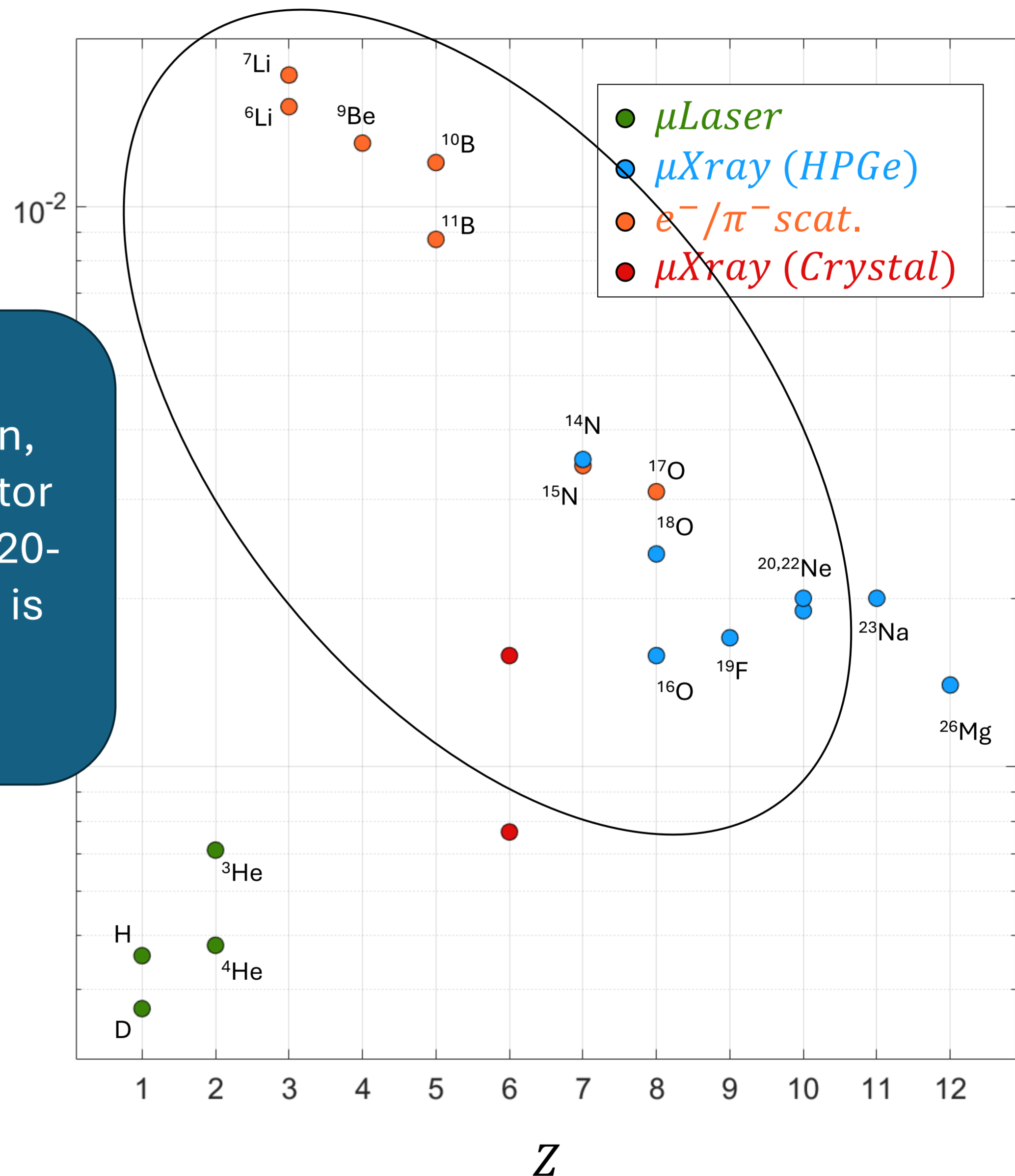




# The radius gap

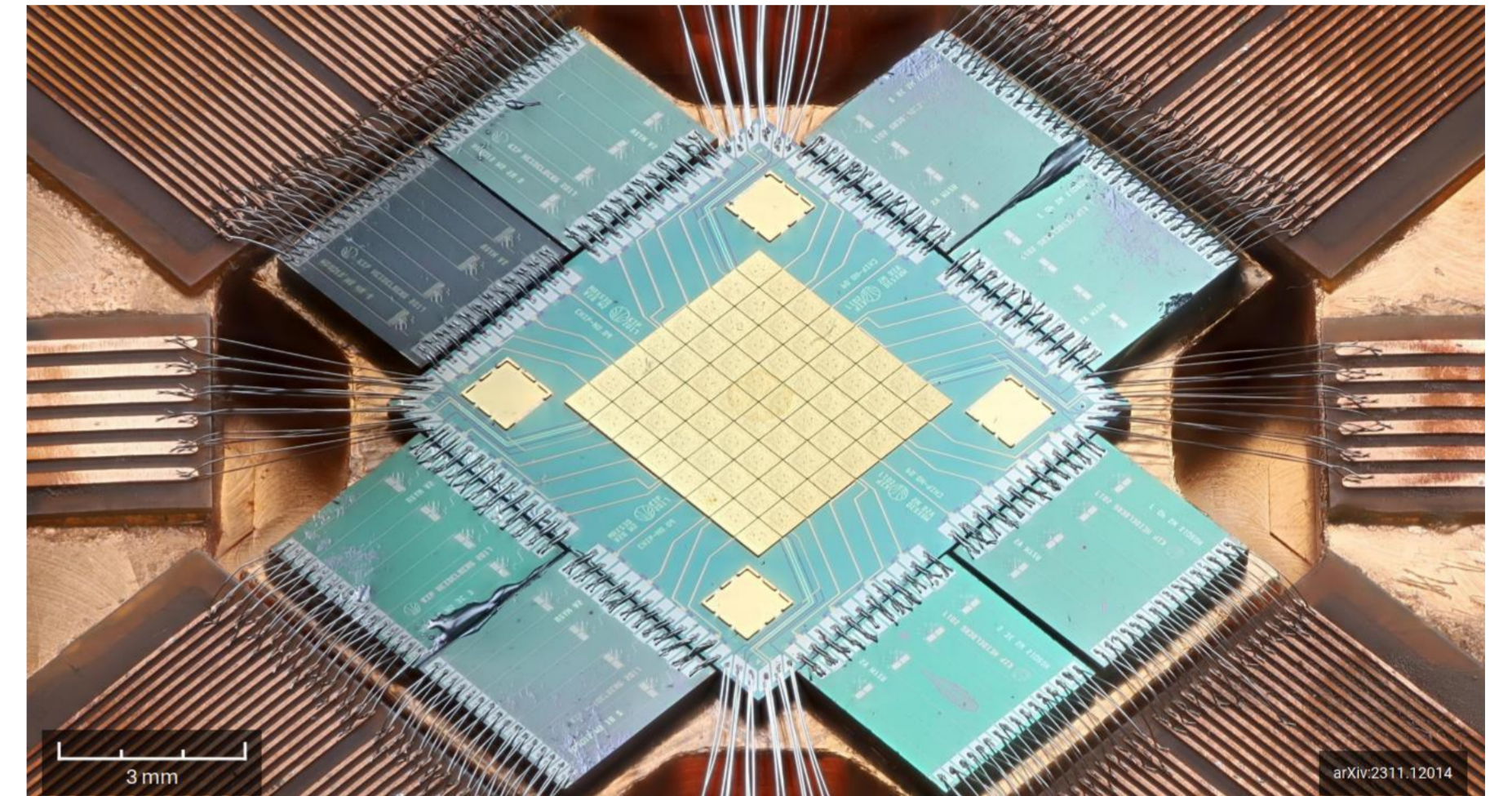
- **For  $Z < 3$ :**  
Laser spectroscopy of muonic atoms, limited by nuclear theory
- **For  $Z > 6$ :**  
Measured x-rays from muonic atoms using solid-state detectors
- **For  $Z = 3 - 5$ , and others:**  
Electron scattering
- **For  $Z = 6$**   
Measured with crystal spectrometer. Not widely applicable

High-resolution, efficient, detector for low-energy (20-200 keV) x-rays is needed



# Enter microcalorimeters

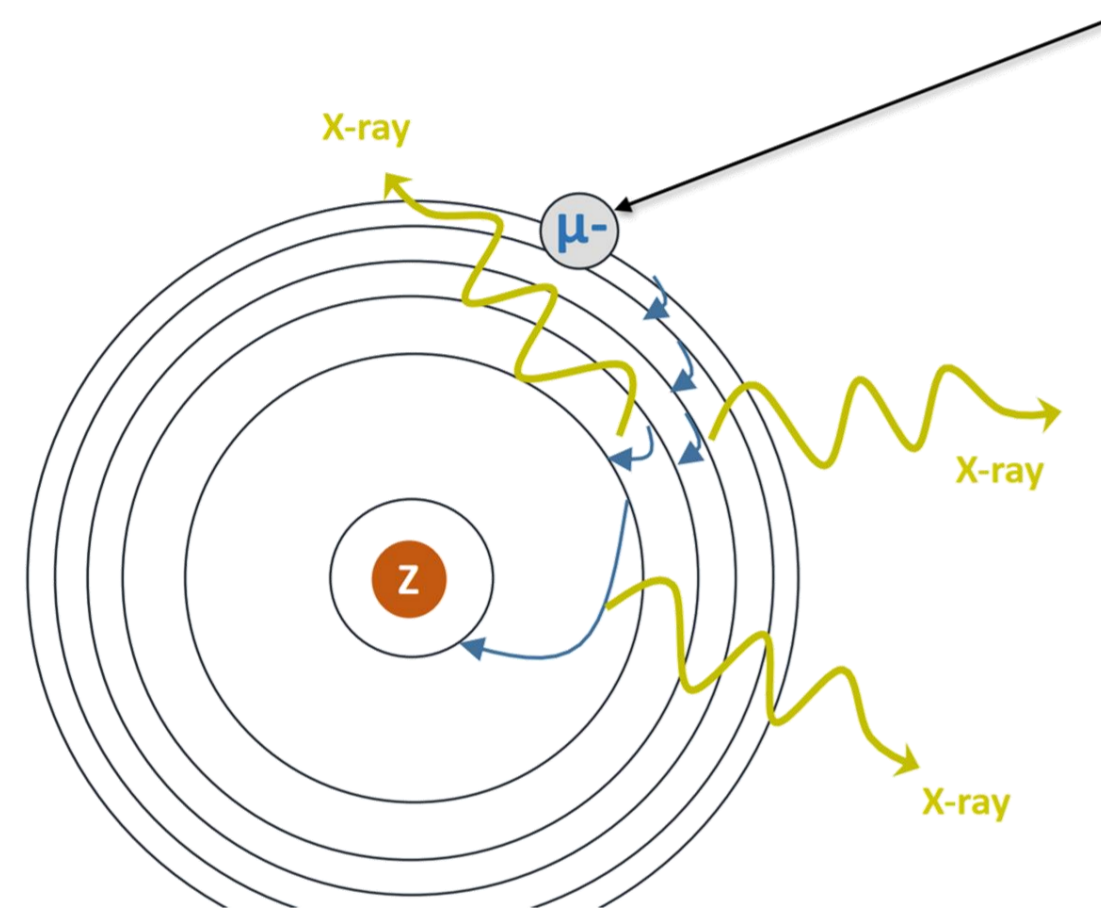
## Cryogenic microcalorimeters



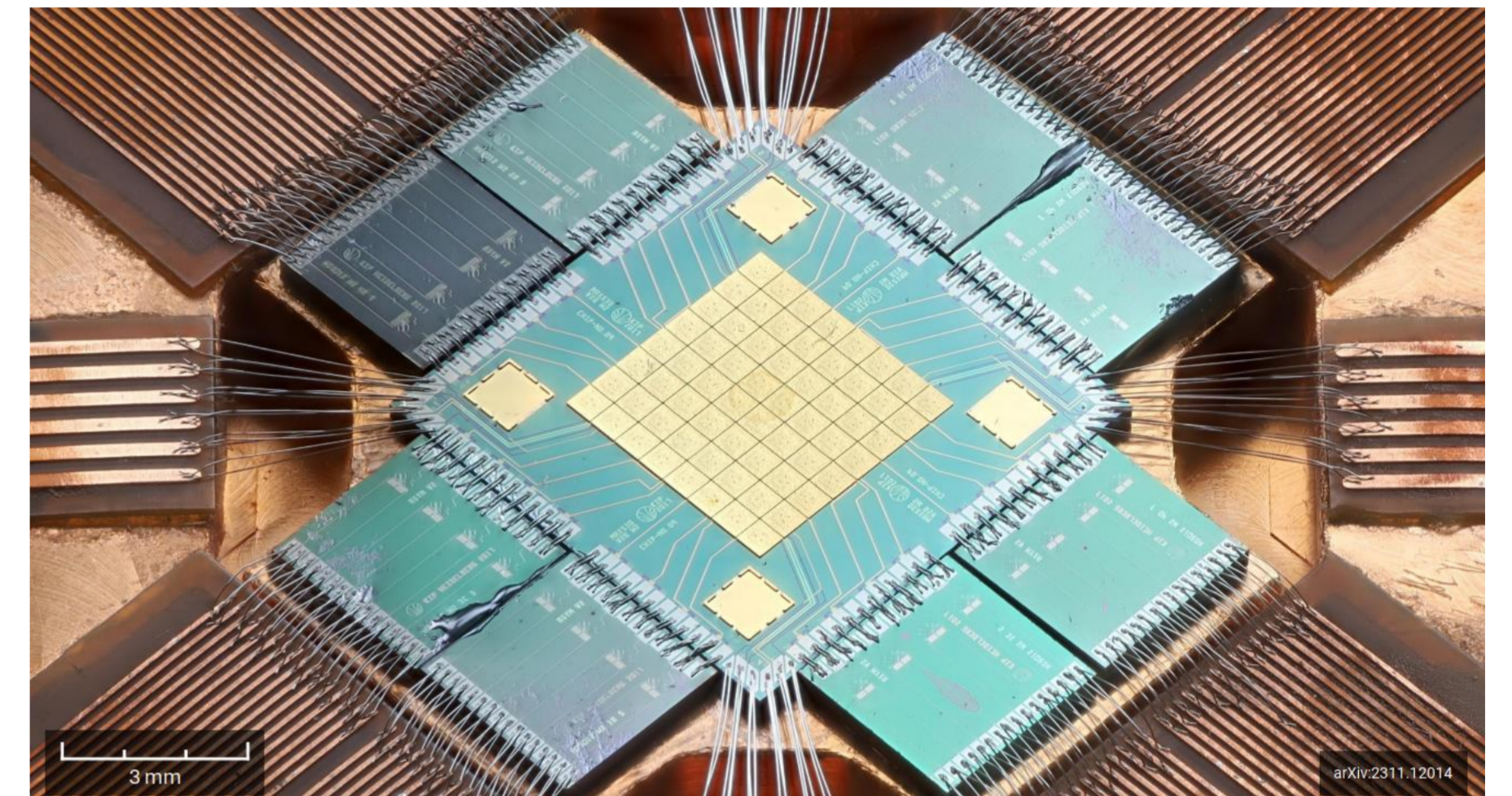
- High quantum efficiency
- Broadband (important for calibration)
- **Superb resolution**  $\left(\frac{E}{\Gamma_E} > 10^3\right)$
- Fast rise time

## Cryogenic microcalorimeters

### Muonic atoms



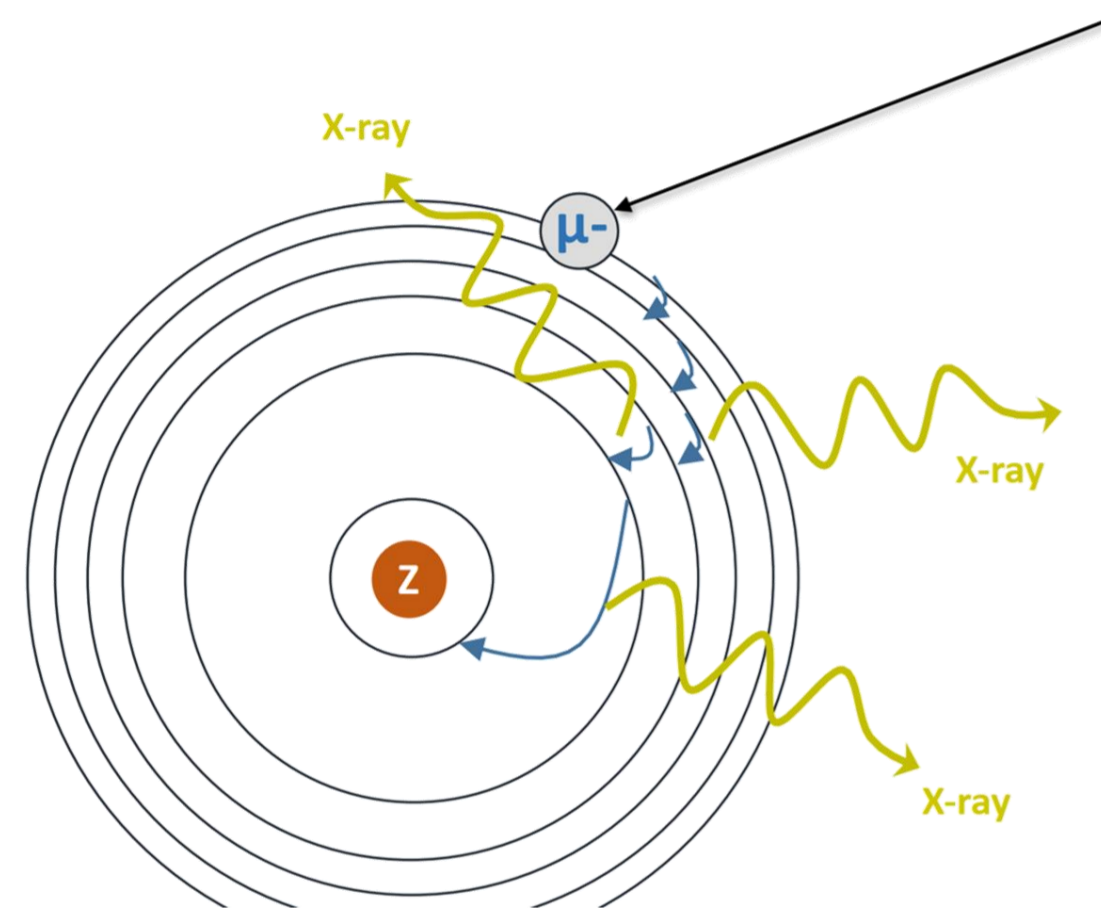
$$\frac{\delta E_{FNS}}{E_0} \sim Z^2 \left( \frac{r_c}{a_0} \right)^2 \left( \frac{m_\mu}{m_e} \right)^2 \sim 10^{-4} Z^2$$



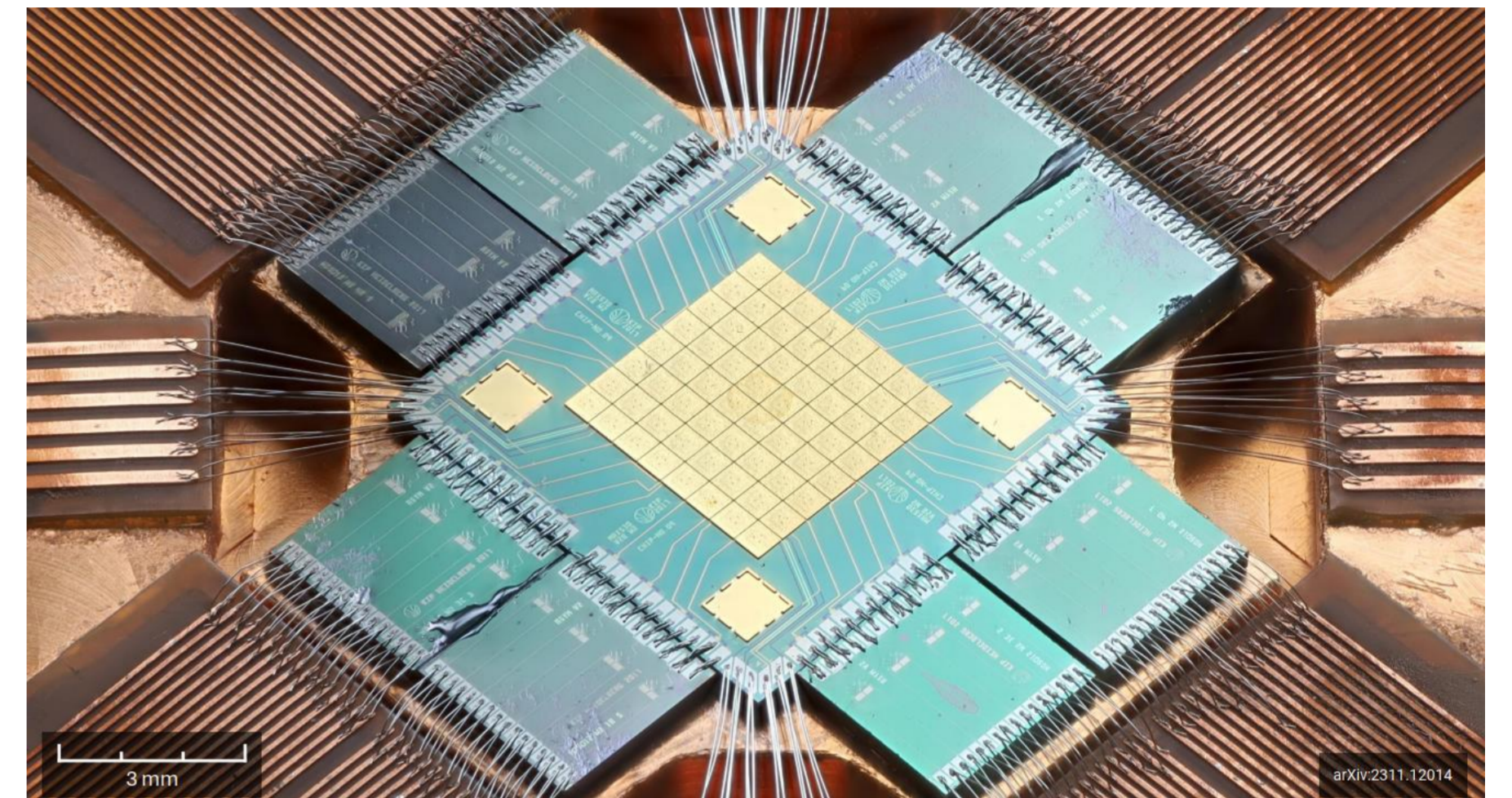
- High quantum efficiency
- Broadband (important for calibration)
- **Superb resolution**  $\left( \frac{E}{\Gamma_E} > 10^3 \right)$
- Fast rise time

## Cryogenic microcalorimeters

### Muonic atoms



$$\frac{\delta E_{FNS}}{E_0} \sim Z^2 \left( \frac{r_c}{a_0} \right)^2 \left( \frac{m_\mu}{m_e} \right)^2 \sim 10^{-4} Z^2$$



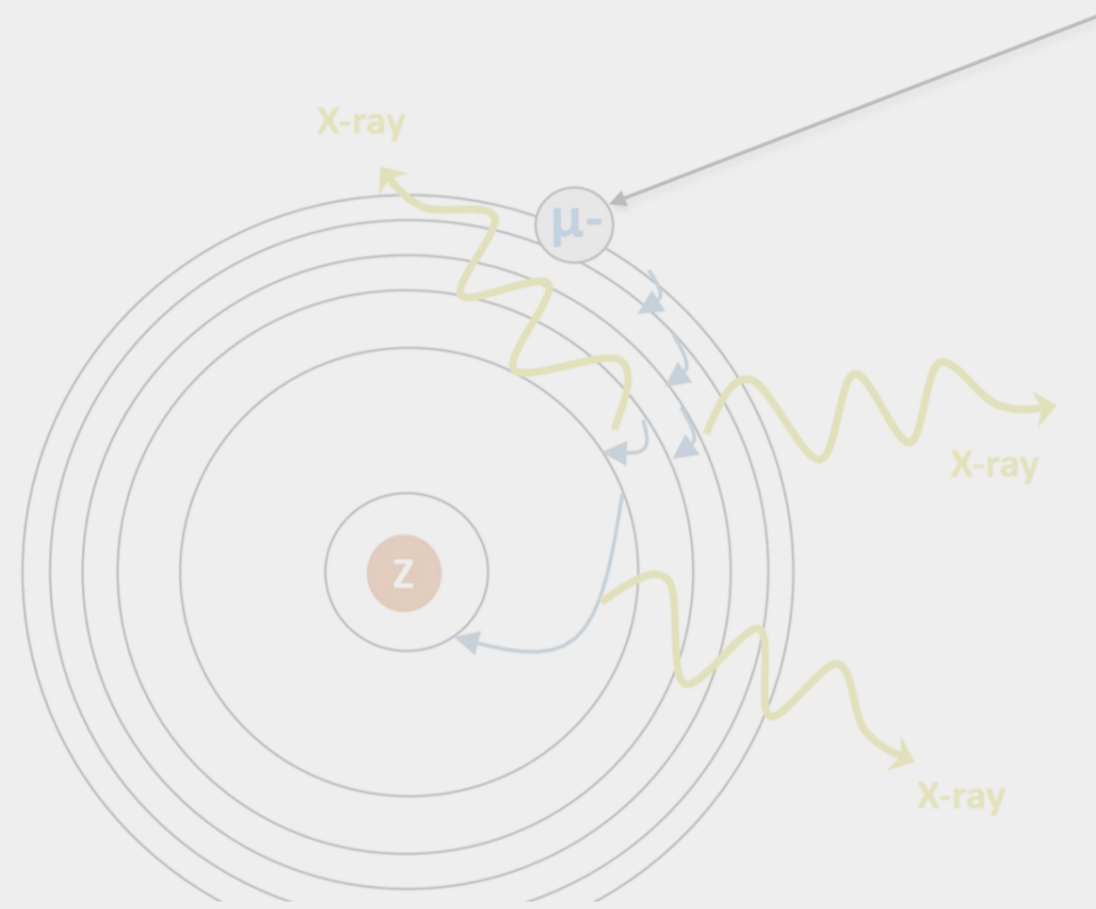
- High quantum efficiency
- Broadband (important for calibration)
- **Superb resolution**  $\left( \frac{E}{\Gamma_E} > 10^3 \right)$
- Fast rise time

# Enter microcalorimeters

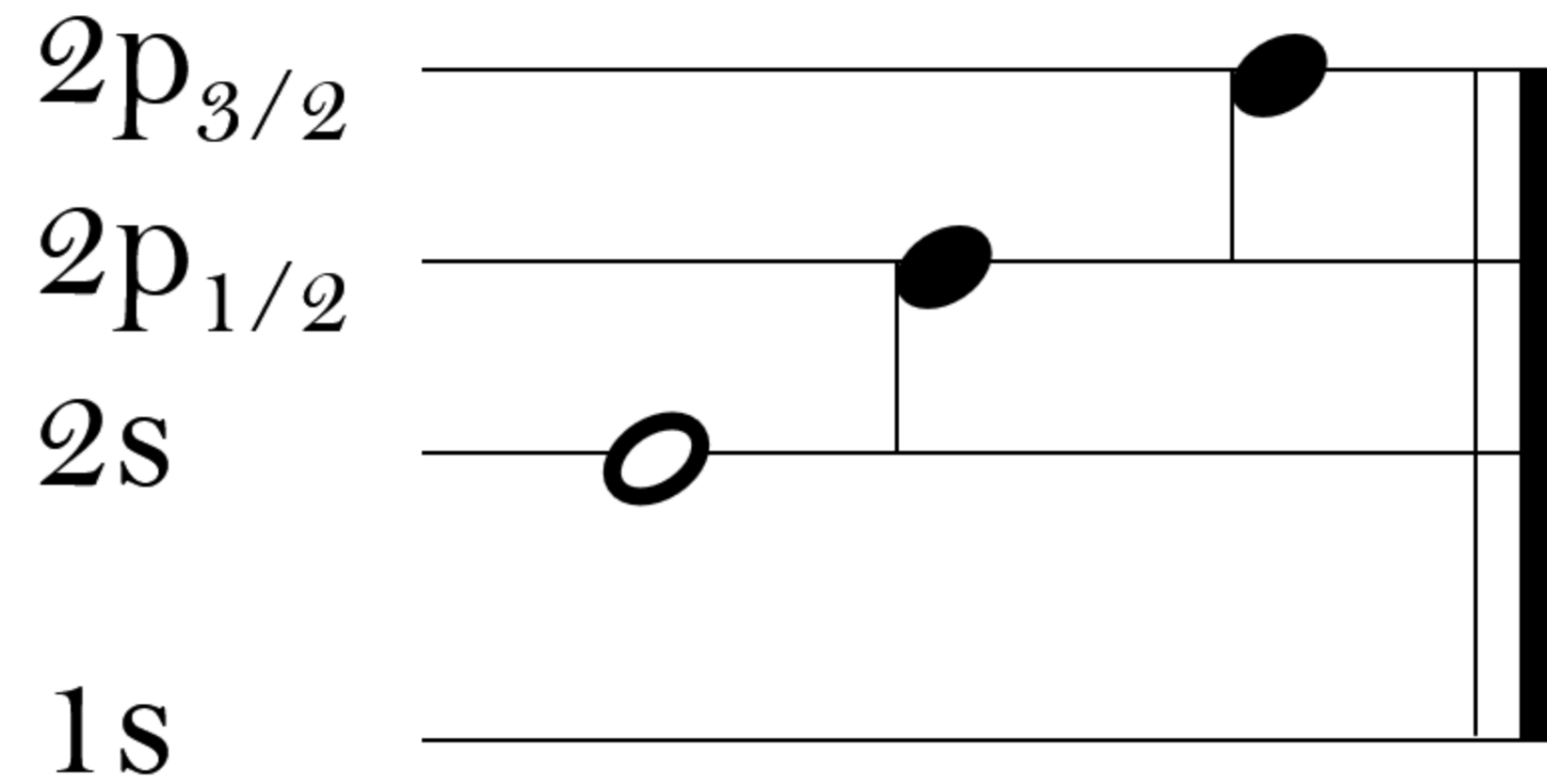
Cryogenic microcalorimeters

## Quantum Interactions with Exotic Atoms

Muonic atoms



# QUARTET

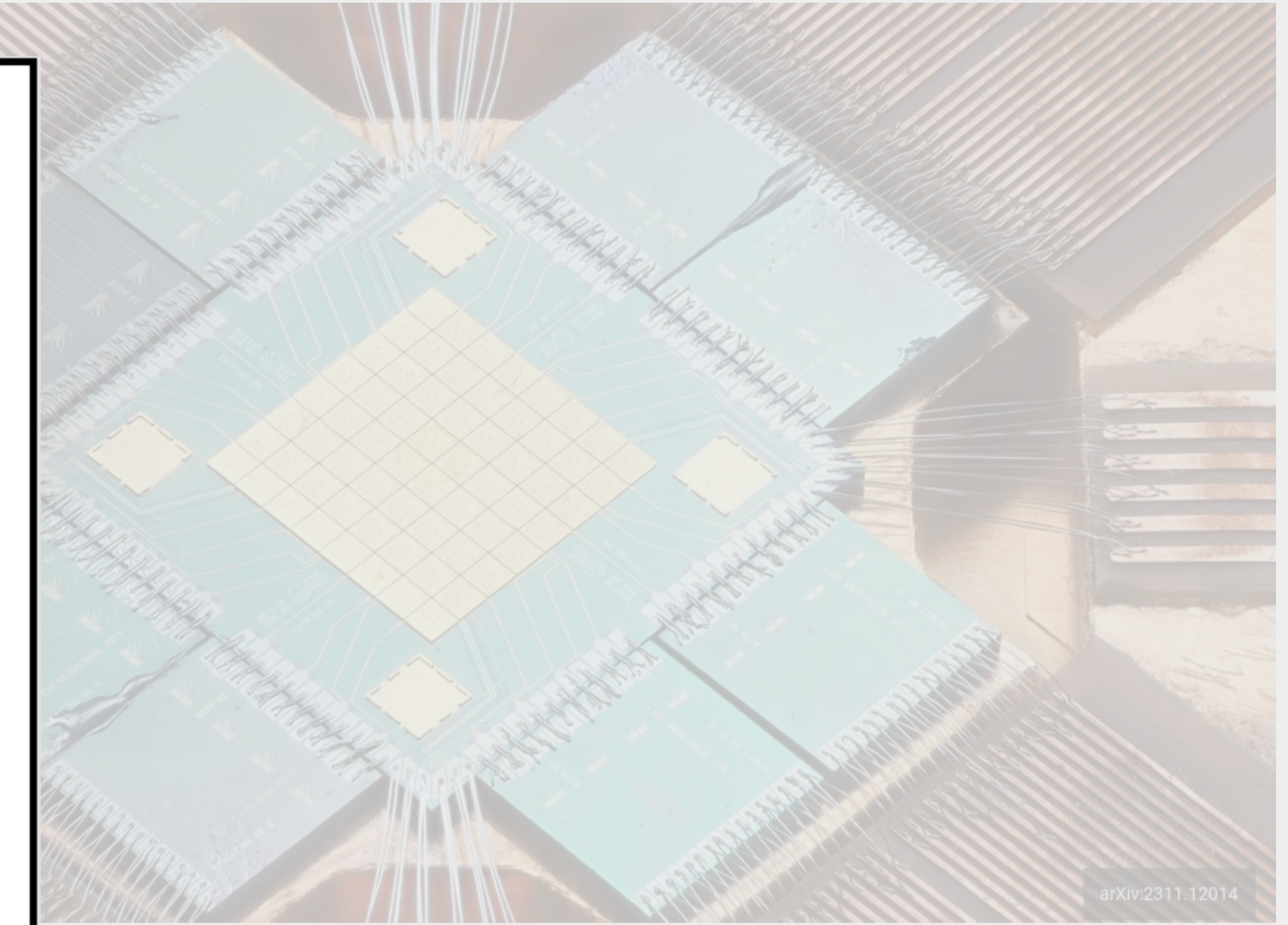


$$\frac{\delta E_{FNS}}{E_0} \sim Z^2 \left( \frac{r_c}{a_0} \right)^2 \left( \frac{m_\mu}{m_e} \right)^2 \sim 10^{-4} Z^2$$

More info:

arXiv:2311.12014

arXiv:2310.03846



High quantum efficiency

- Broadband (important for calibration)
- **Superb resolution**  $\left( \frac{E}{\Gamma_E} > 10^3 \right)$
- Fast rise time (important for background suppression)

# Enter microcalorimeters

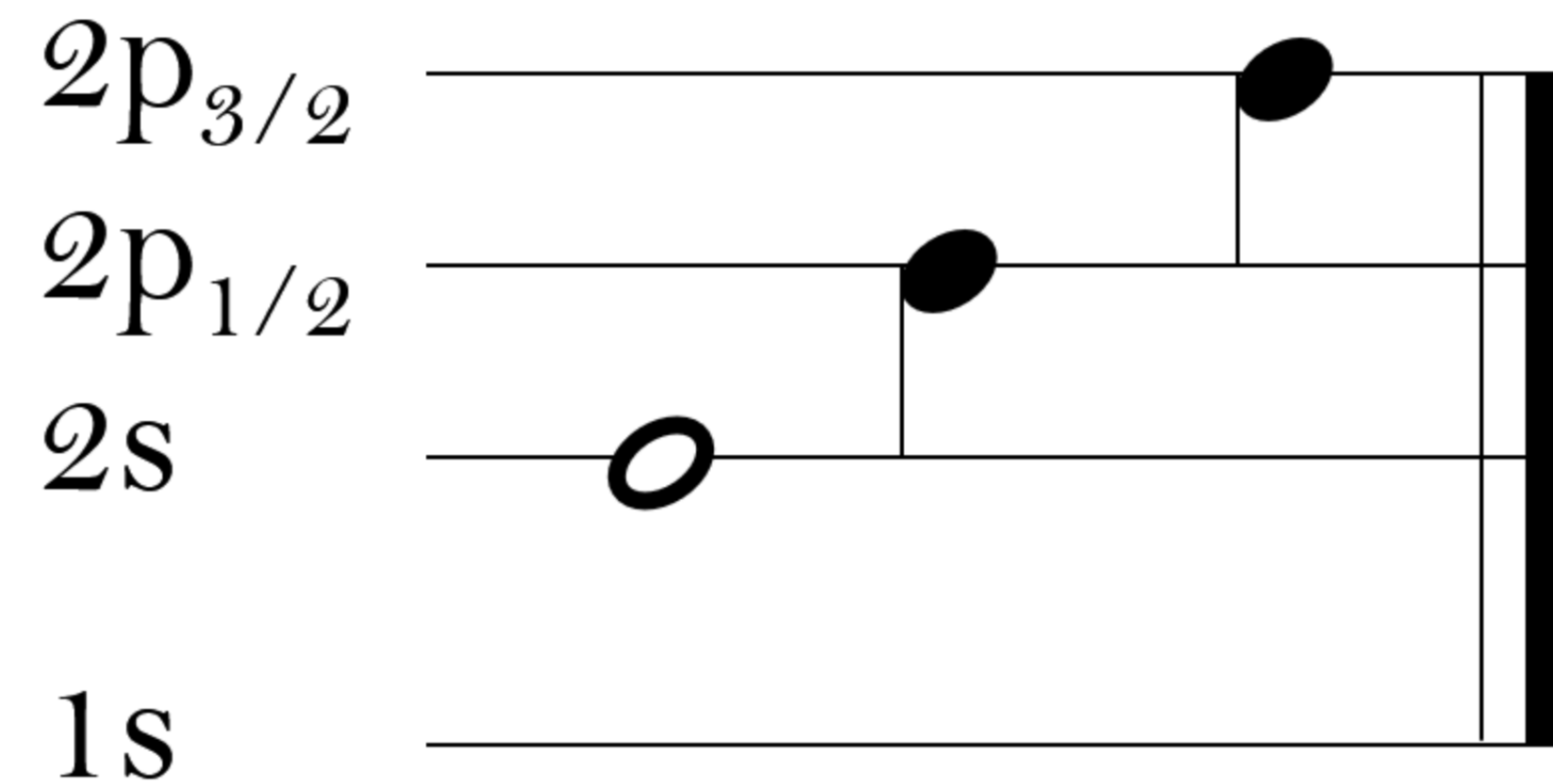
Cryogenic microcalorimeters

## Quantum Interactions with Exotic Atoms

Muonic atoms



# QUARTET

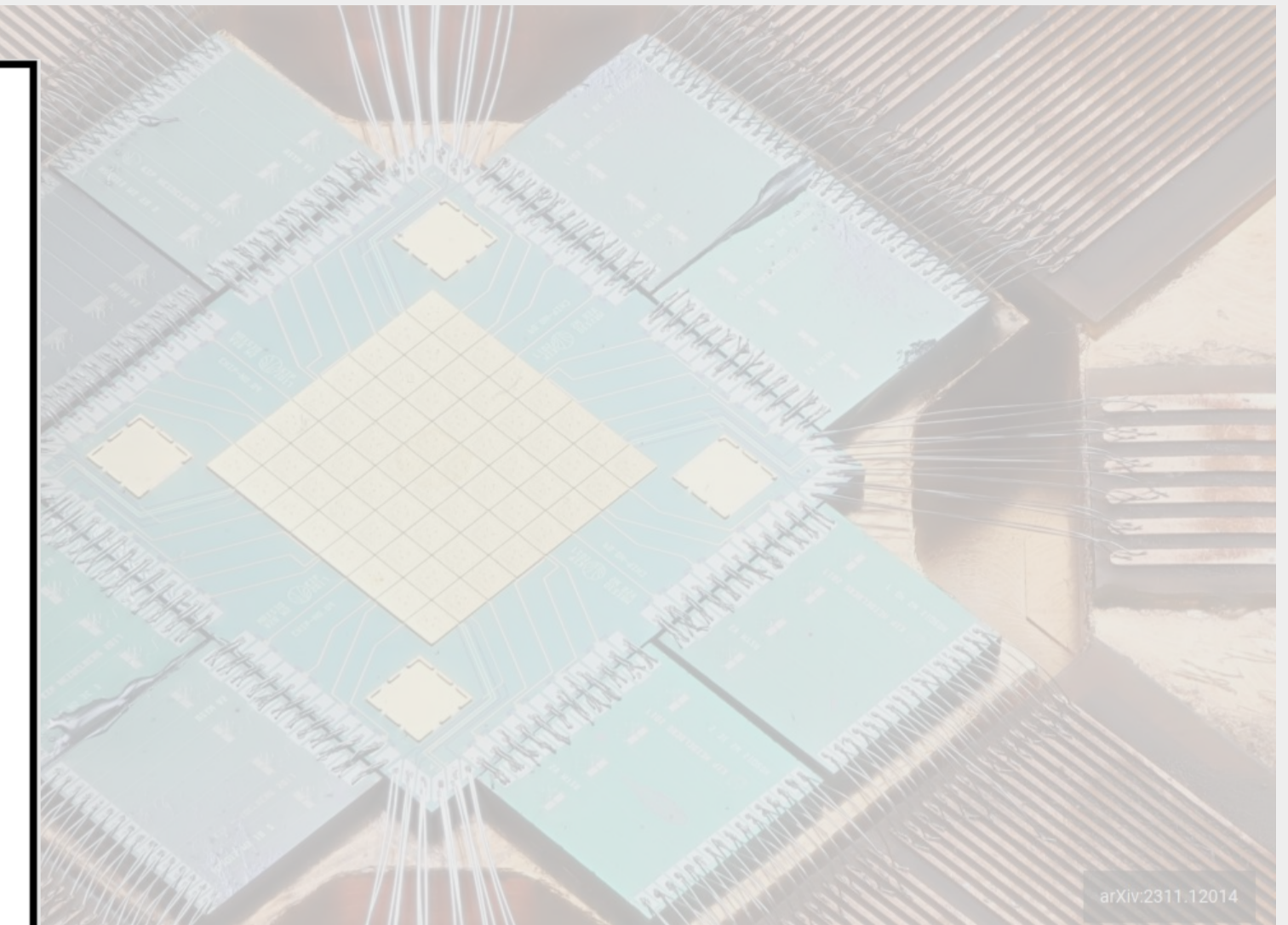


$$\frac{\delta E_{FNS}}{E_0} \sim Z^2 \left( \frac{r_c}{a_0} \right)^2 \left( \frac{m_\mu}{m_e} \right)^2 \sim 10^{-4} Z^2$$

More info:

arXiv:2311.12014

arXiv:2310.03846



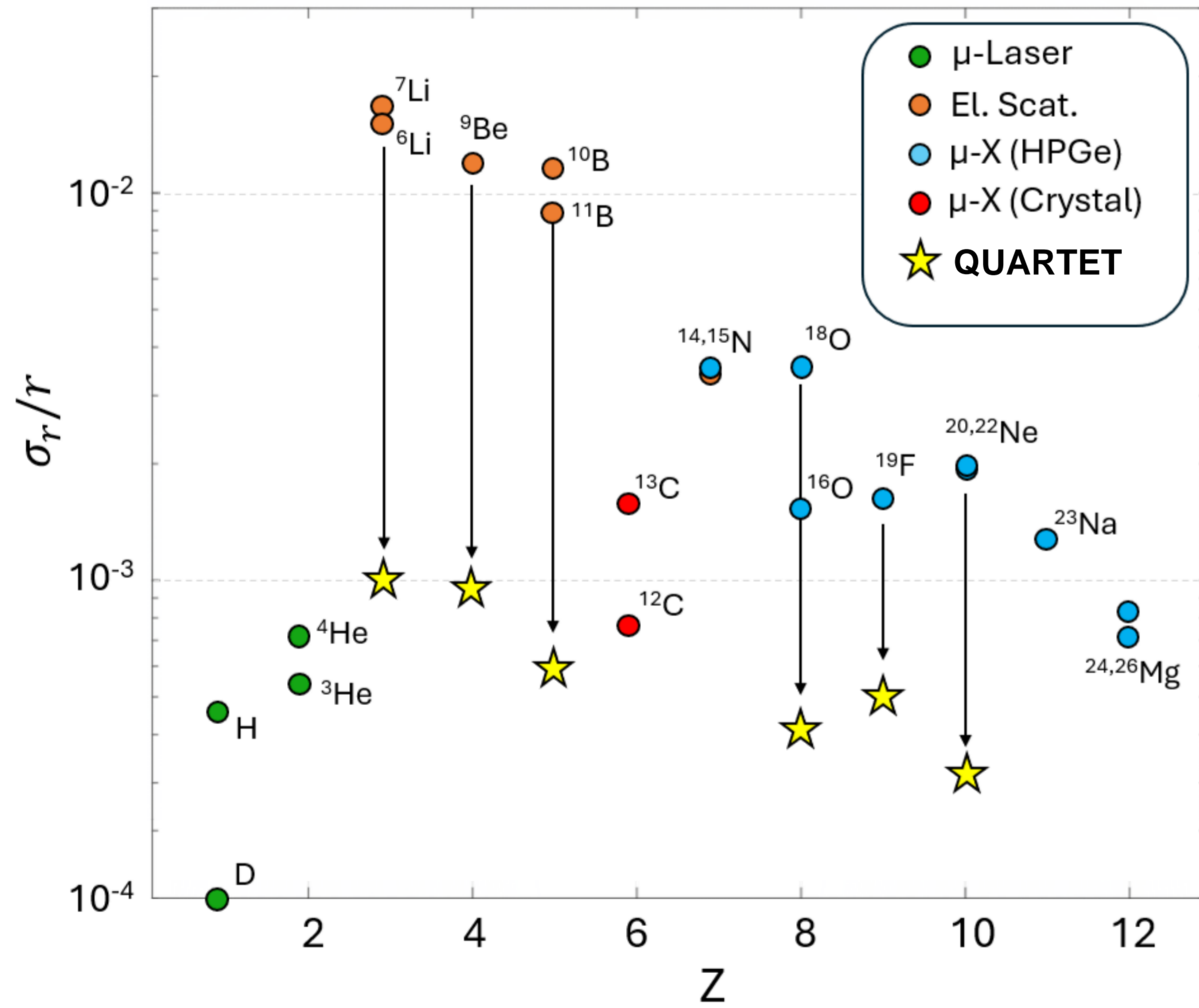
High quantum efficiency

- Broadband (important for calibration)

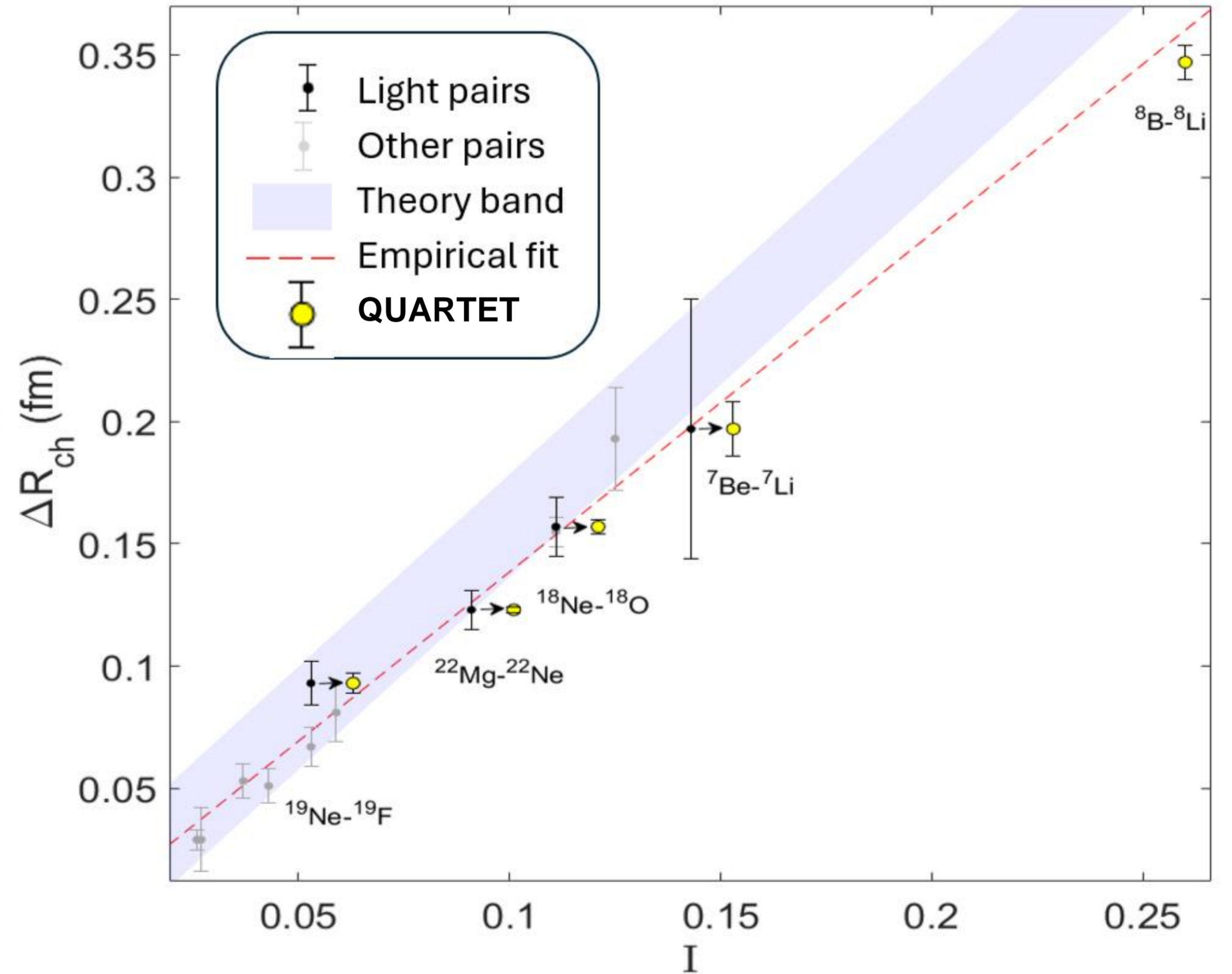
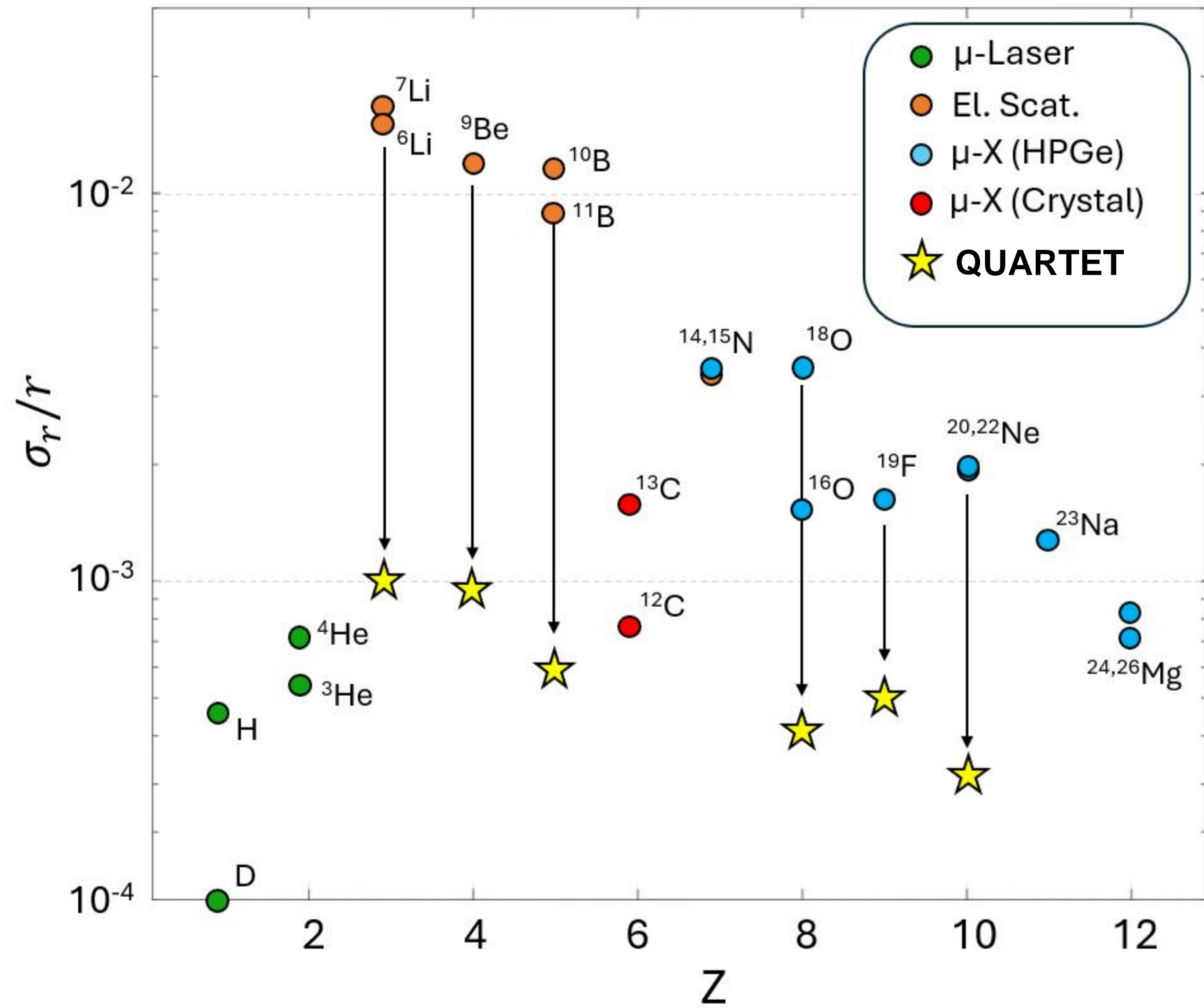
- **Superb resolution**  $\left( \frac{E}{\Gamma_E} > 10^3 \right)$

**Stay tuned for Loredana's talk !!!** Fast rise time (important for background suppression)

# What can better radii of light nuclei do for the mirror fit?

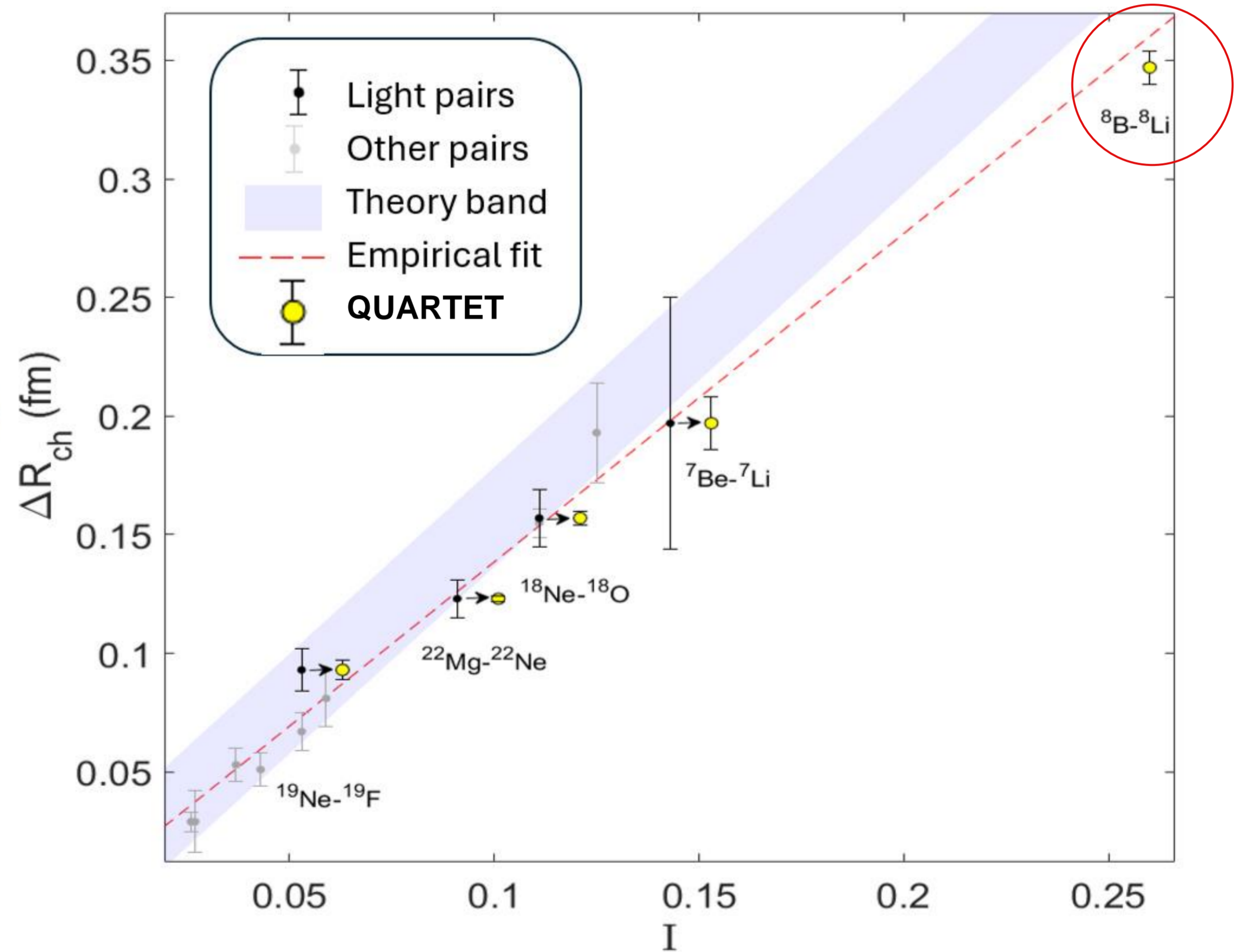
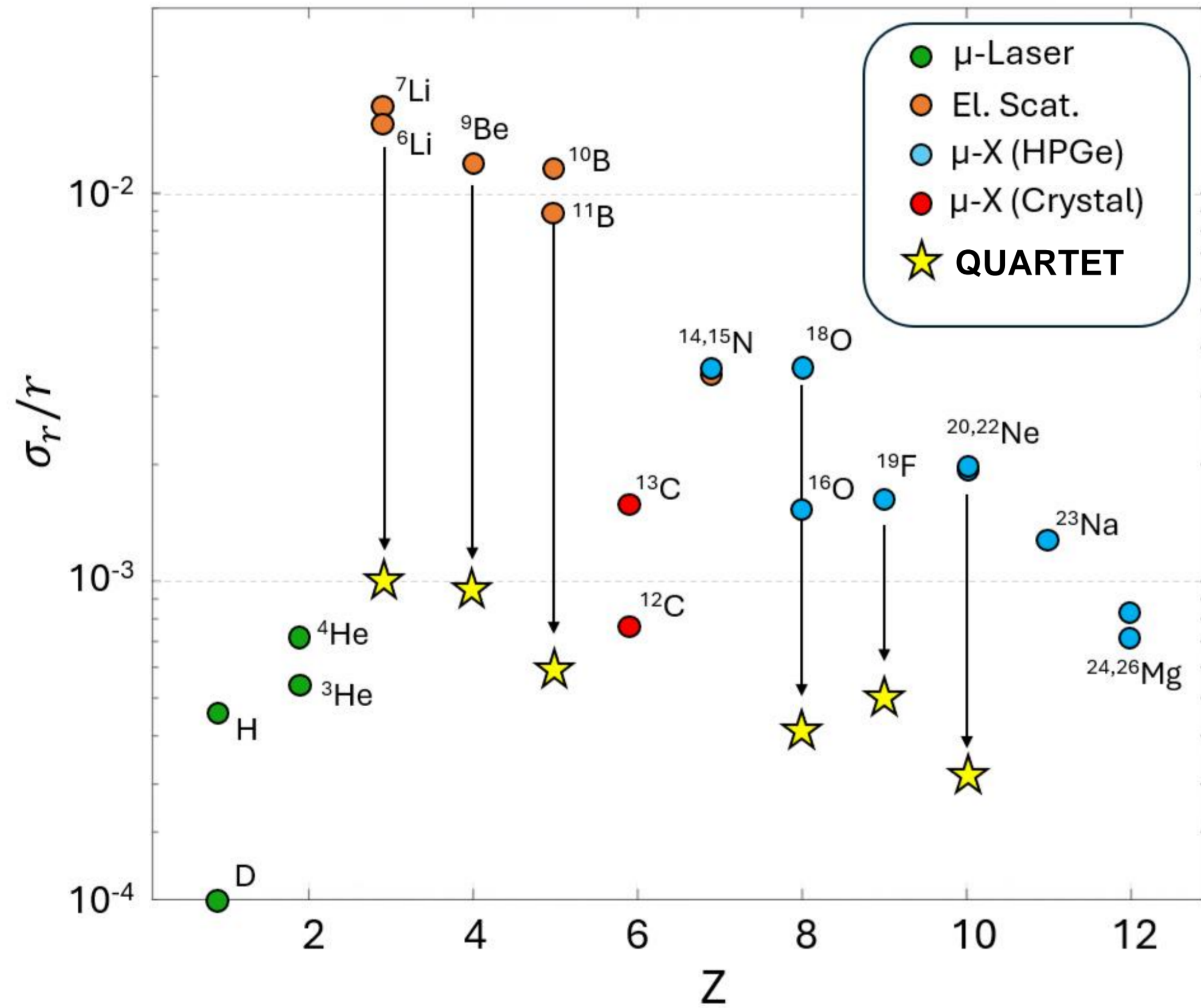


# What can better radii of light nuclei do for the mirror fit?



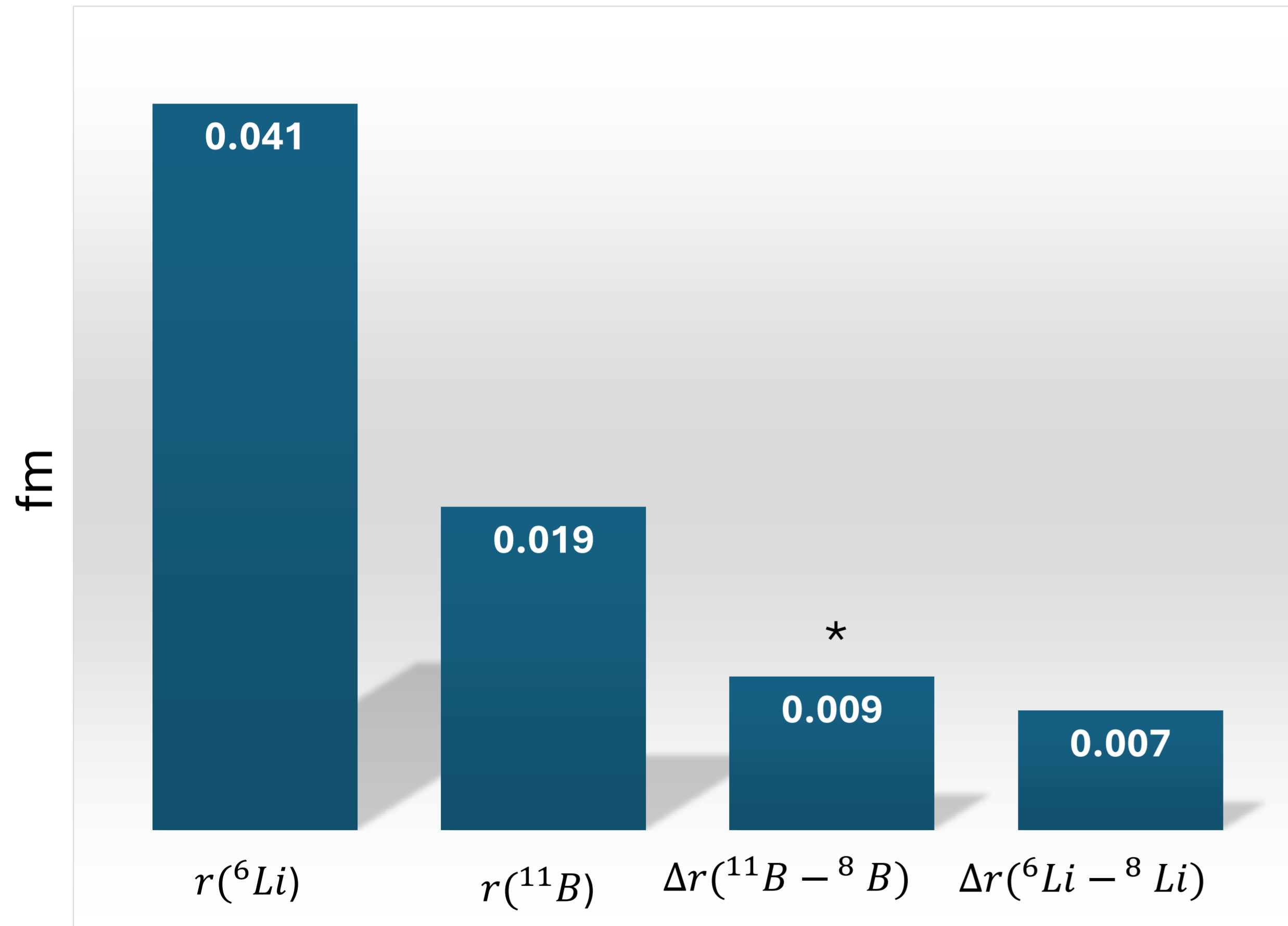


# What can better radii of light nuclei do for the mirror fit?

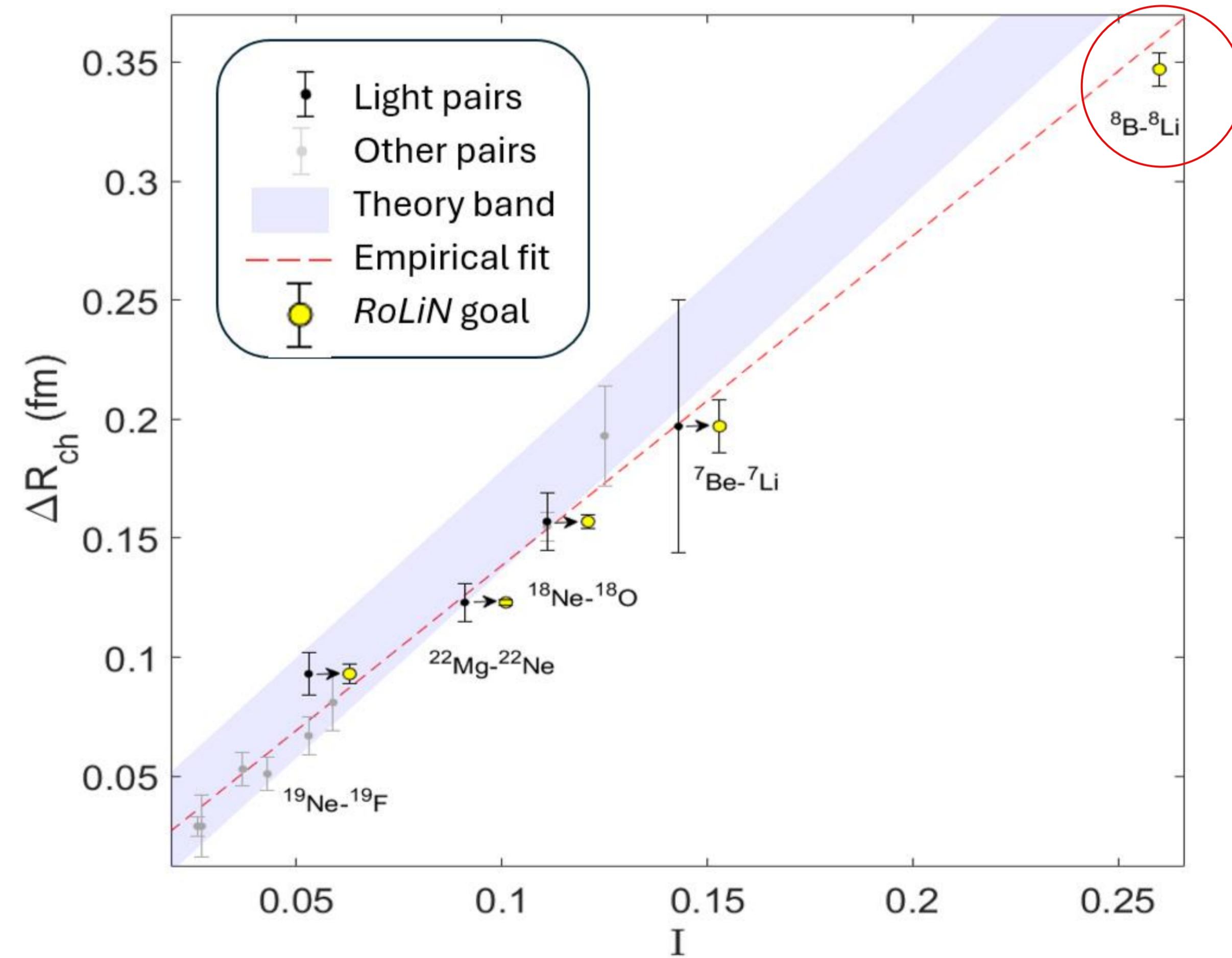


# Golden case: A=8

Uncertainty contributions to  $\Delta_{ch}^{mirr}(A=8)^*$

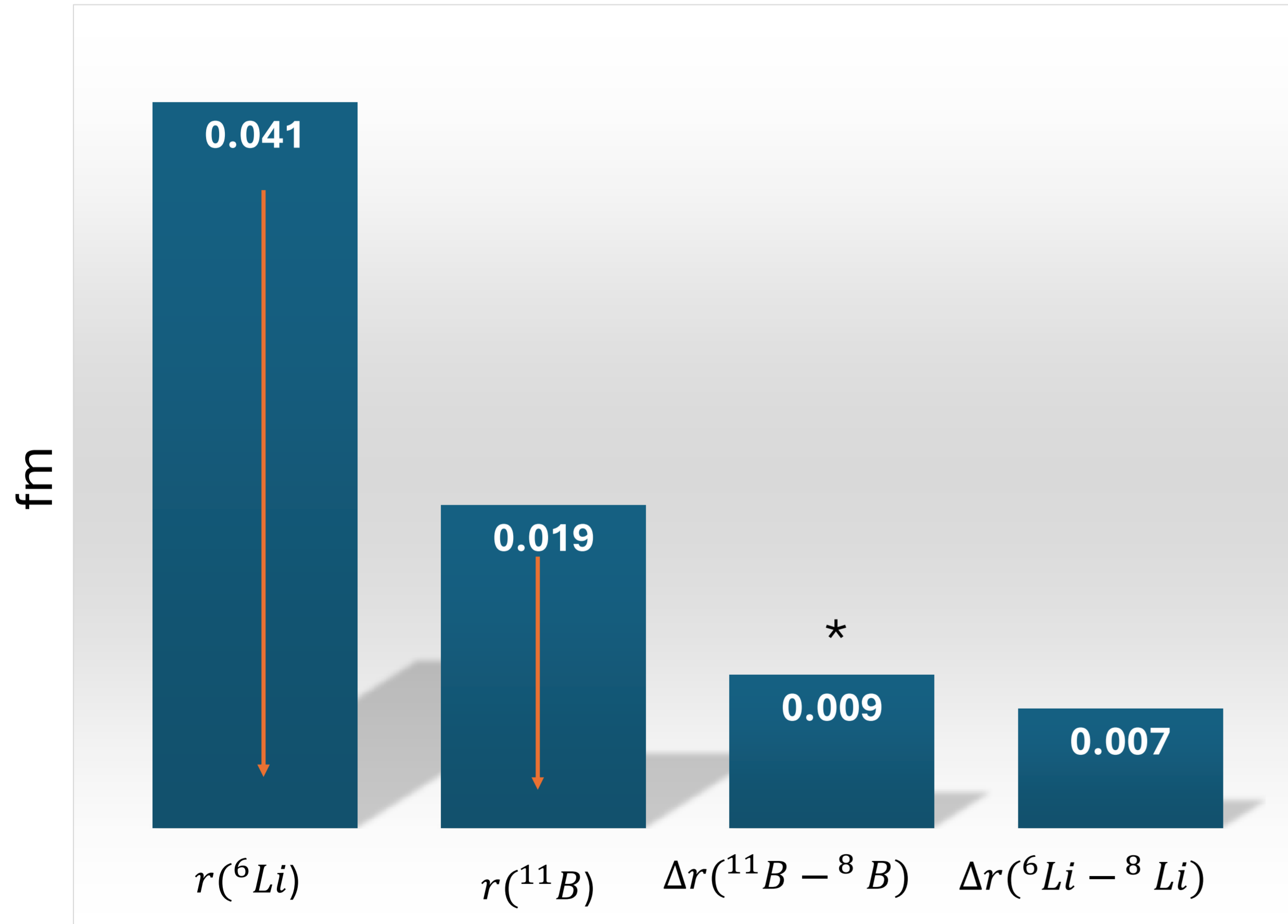


\* Private com. Wilfried Nörtershäuser

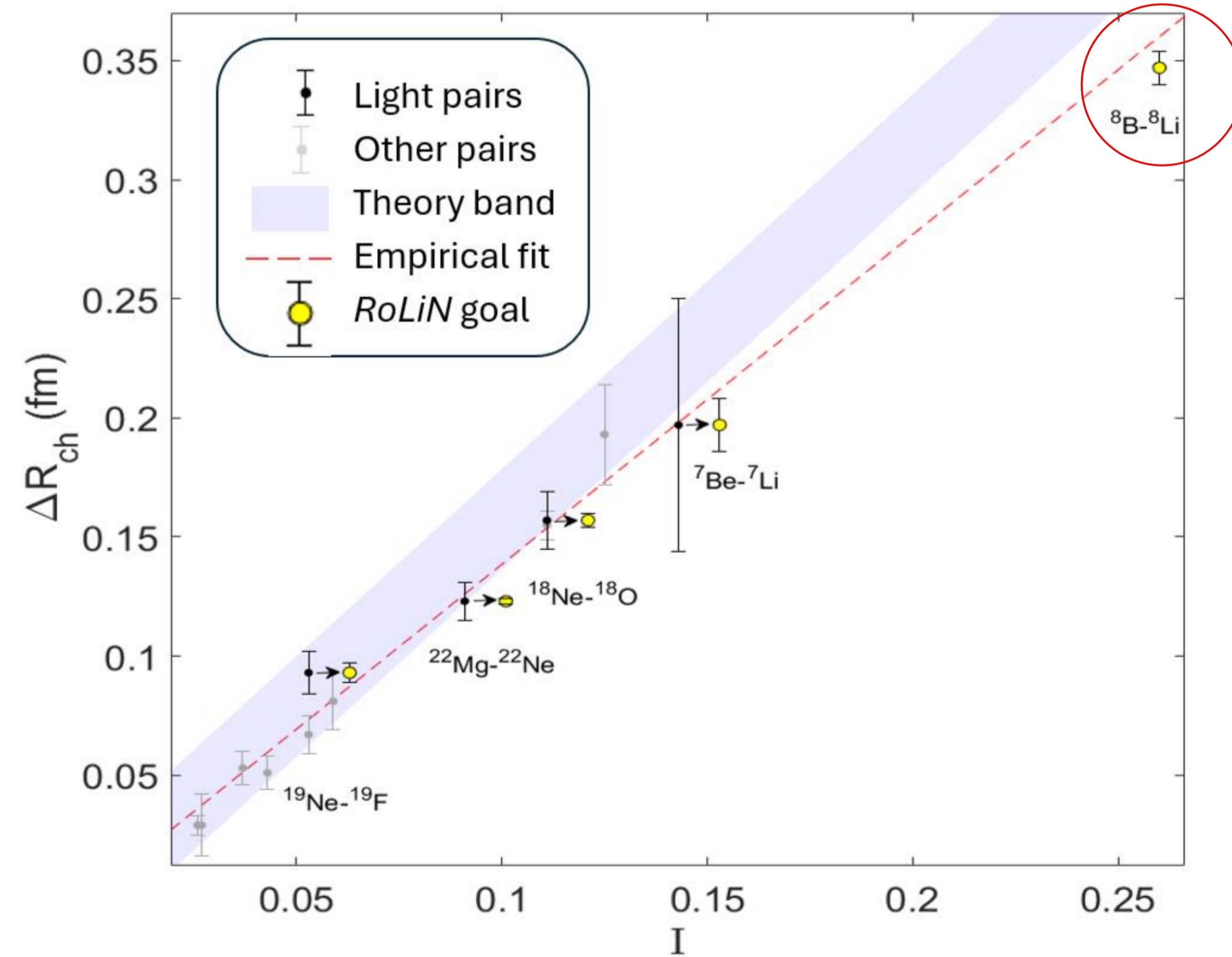


# Golden case: A=8

Uncertainty contributions to  $\Delta_{ch}^{mirr}(A=8)^*$



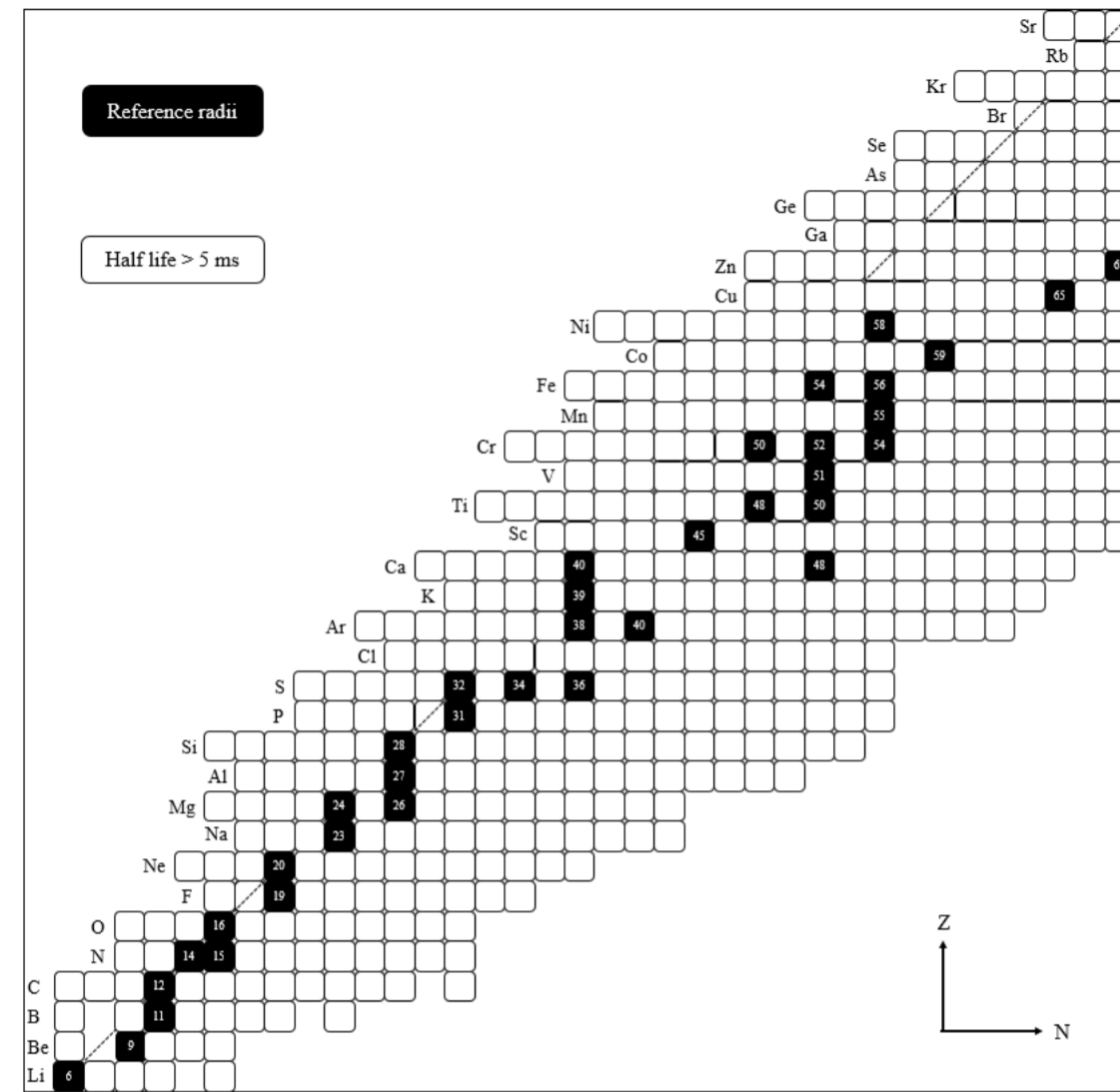
\* Private com. Wilfried Nörtershäuser



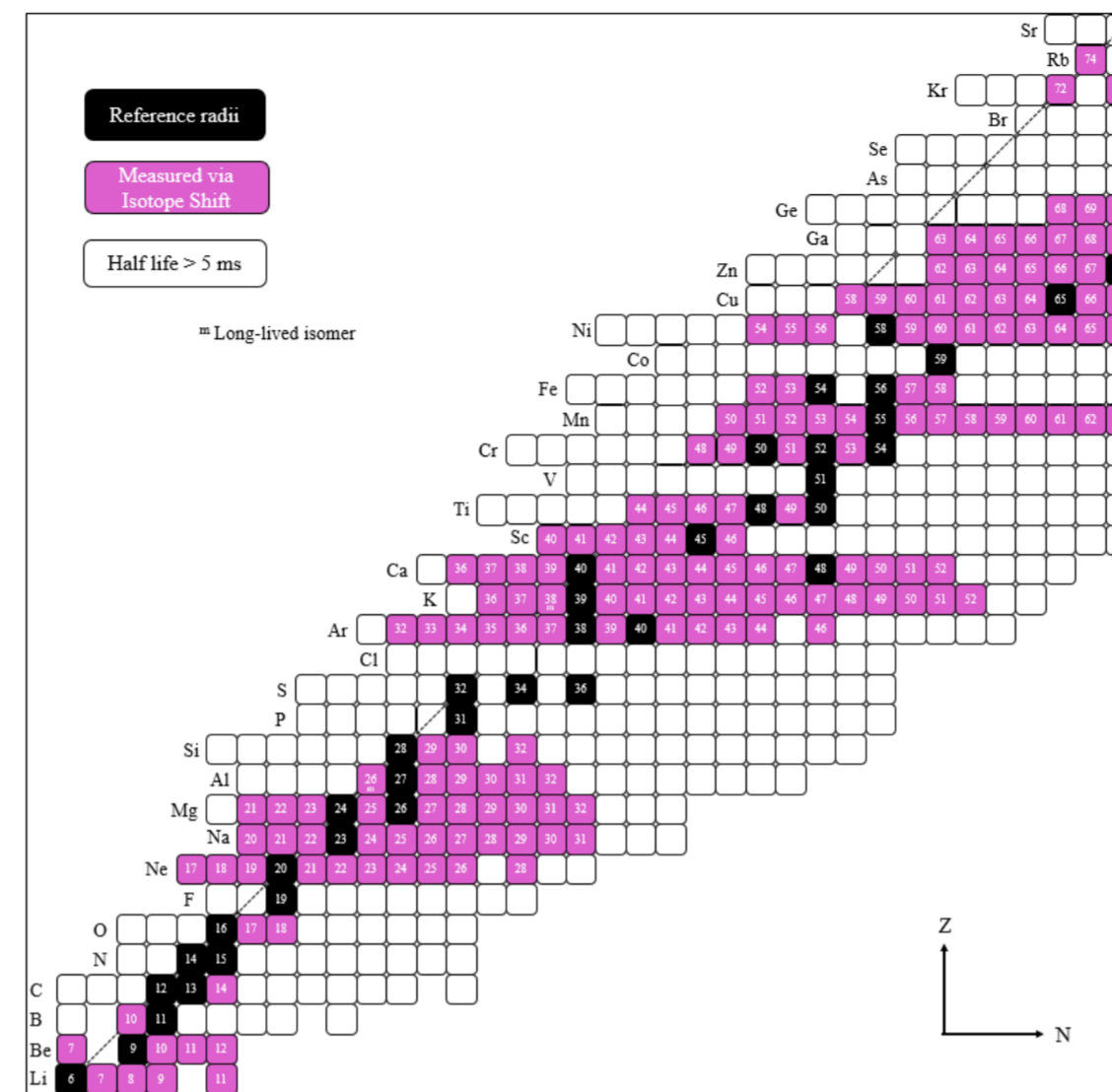
What can we do if the mirror fit holds?

# Roadmap:

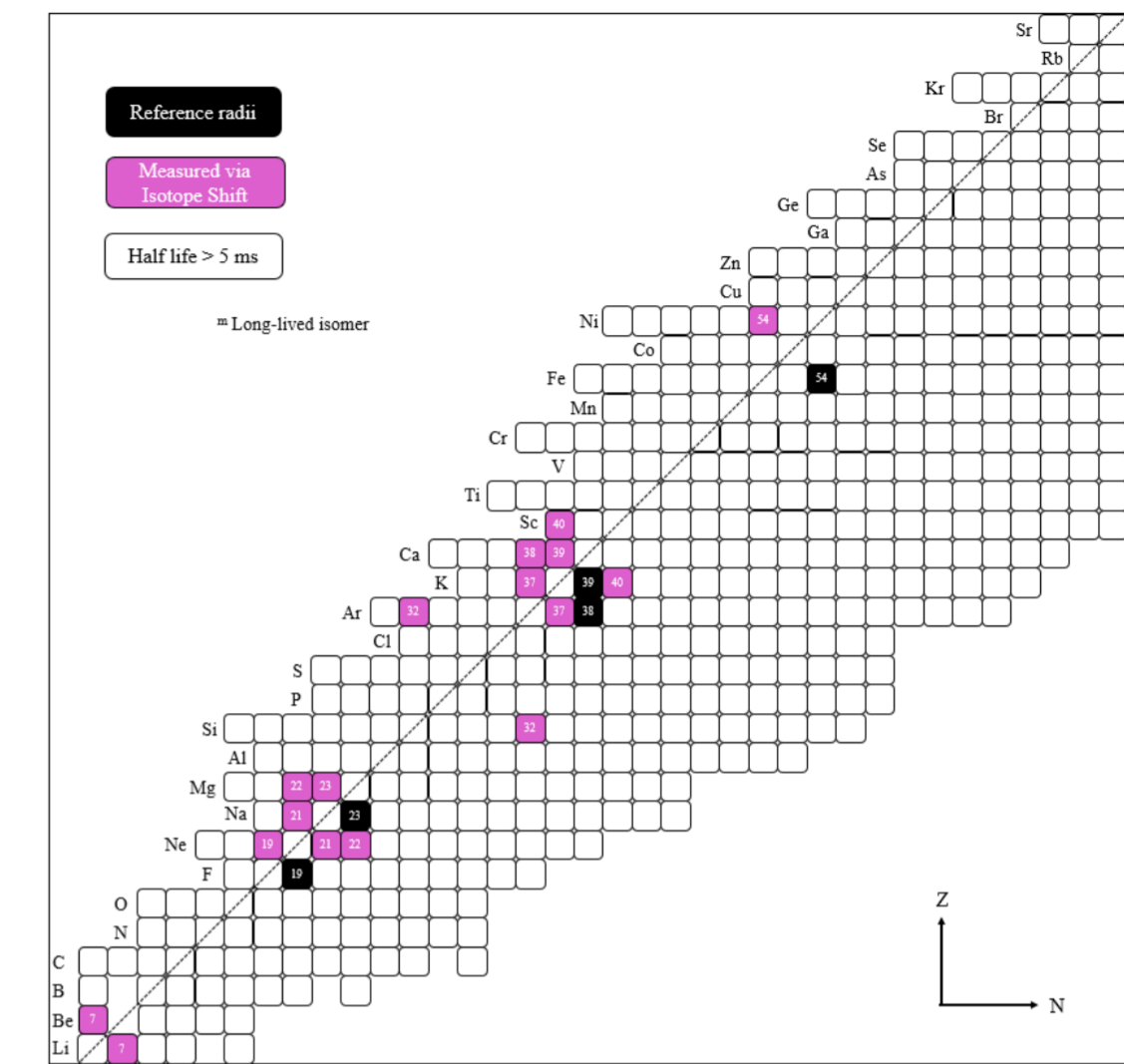
## 1. Reference radii (muonic atoms)



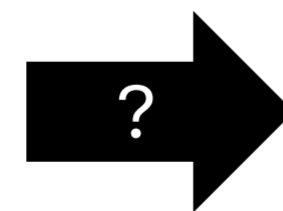
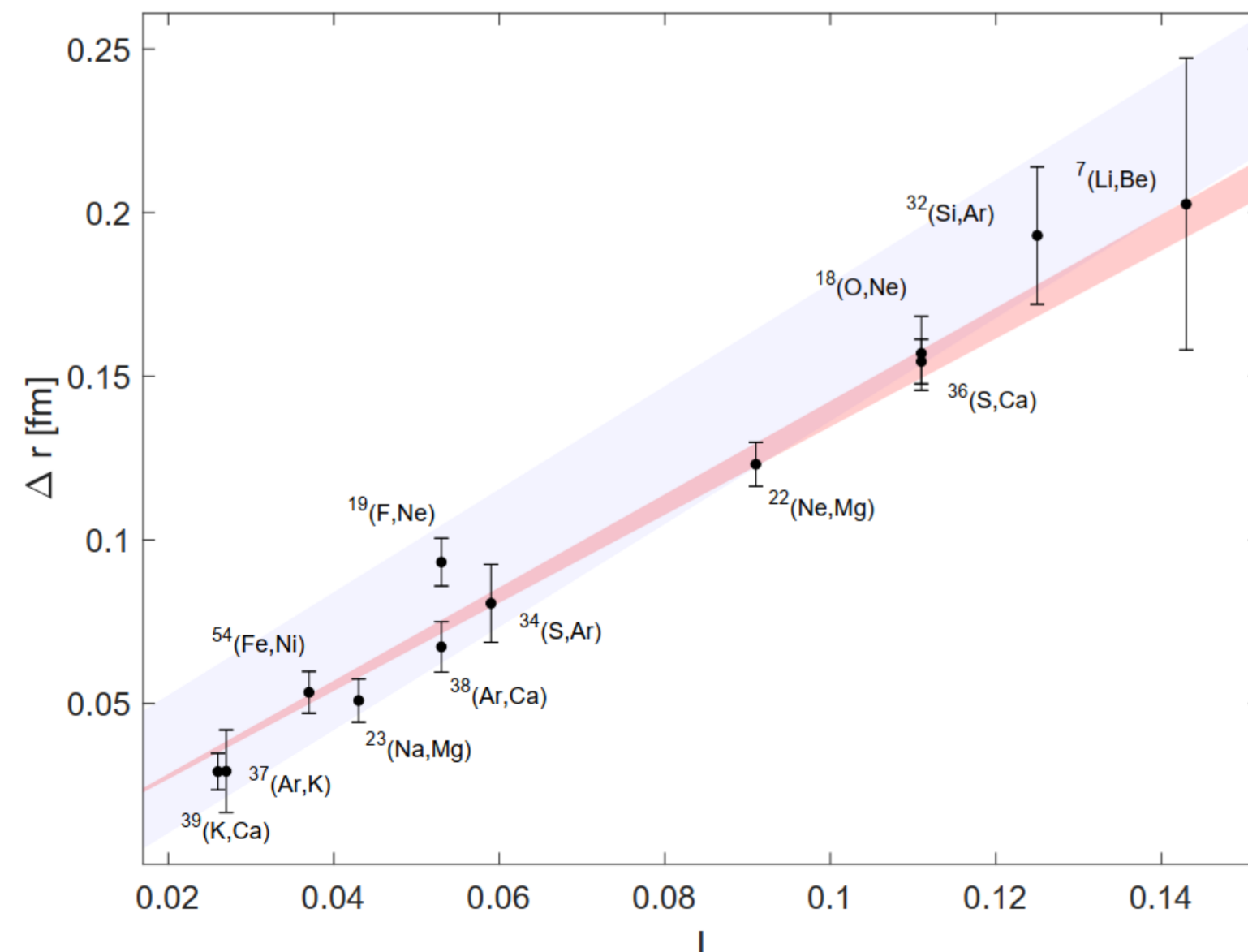
## 2. With Isotope shifts (electronic atoms)



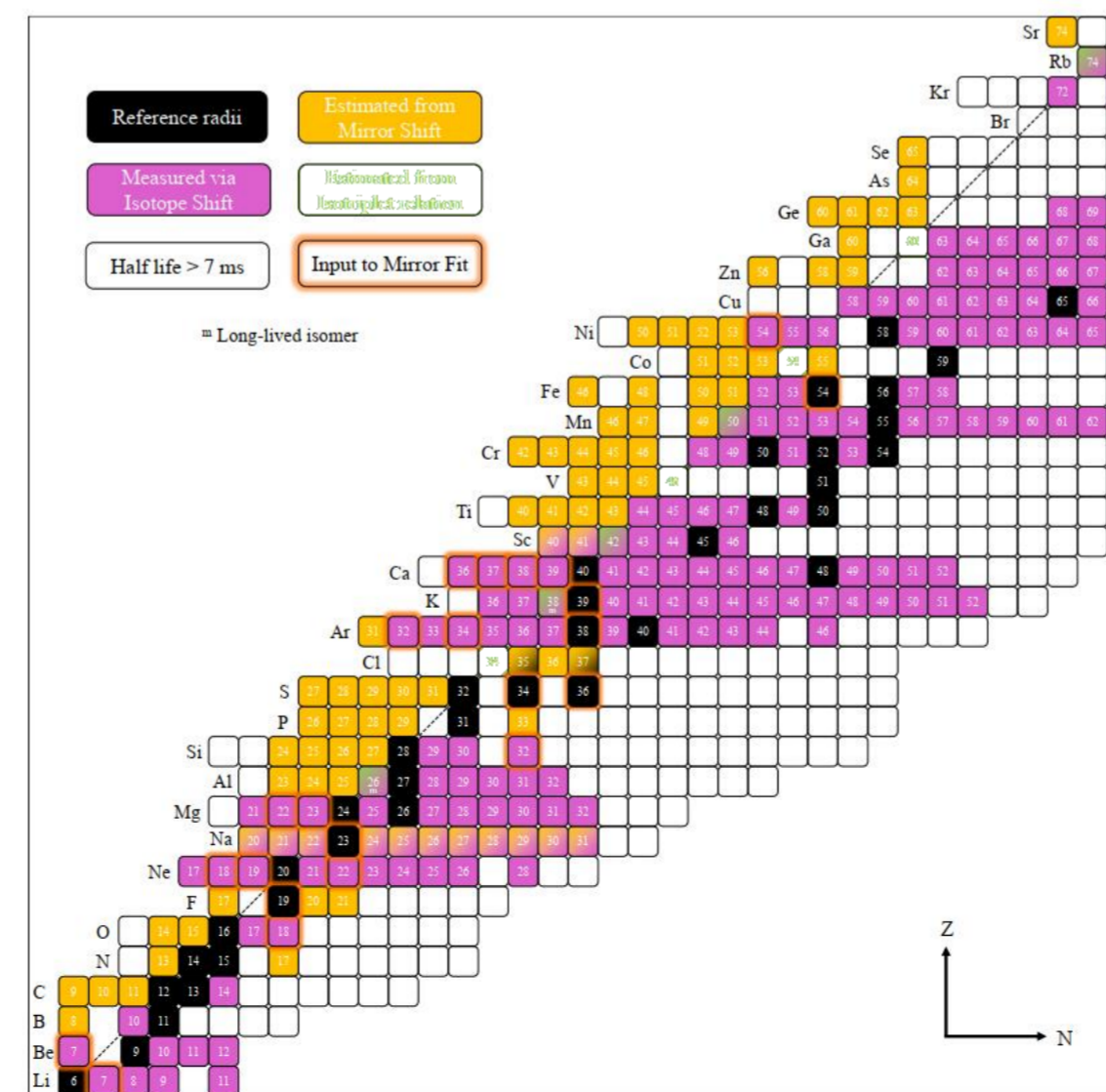
## 3. Mirror nuclei



## 4. Mirror fit



## 5. Predict 74 charge radii



# Transparent tabulation of measured and predicted radii of mirror nuclei

**Table 5**

Charge radii of mirror nuclei. Known radii are taken from Tab. 2 along with isotopic differences taken from the indicated references. Predicted radii relying on the validity of Eq. 6 are given in italics.

A	I	el.	Z	N	r	Ref./Note	el.	Z	N	r	Ref./Note
7	0.14	Be	4	3	2.646(33)	[58]	Li	3	4	2.449(41)	[59, 60]
8	0.25	B	5	3	<i>2.685(45)</i>		Li	3	5	2.339(44)	[48, 61]
9	0.33	C	6	3	<i>2.705(47)</i>		Li	3	6	2.245(46)	[48, 61]
10	0.20	C	6	4	<i>2.638(36)</i>		Be	4	6	2.361(36)	[58]
11	0.09	C	6	5	<i>2.536(21)</i>		B	5	6	2.411(21)	
13	0.08	N	7	6	<i>2.564(04)</i>		C	6	7	2.458(02)	[62]
14	0.14	O	8	6	<i>2.705(10)</i>		C	6	8	2.508(09)	[63] A
15	0.07	O	8	7	<i>2.704(09)</i>		N	7	8	2.612(09)	
17	0.06	F	9	8	<i>2.774(08)</i>		O	8	9	2.693(08)	[64] B
17	0.18	Ne	10	7	3.015(10)	Tab. 3	N	7	10	<i>2.771(12)</i>	
18	0.11	Ne	10	8	2.934(10)	Tab. 3	O	8	10	2.777(07)	[65] C
19	0.05	Ne	10	9	2.995(07)	Tab. 3	F	9	10	2.902(05)	
20	0.10	Na	11	9	<i>2.983(25)</i>	Tab. 6	F	9	11	<i>2.845(25)</i>	
21	0.05	Na	11	10	<i>3.029(07)</i>		Ne	10	11	2.963(07)	Tab. 3
21	0.14	Mg	12	9	3.067(07)	[37, 66]	F	9	12	<i>2.869(09)</i>	
22	0.09	Mg	12	10	3.071(05)	[37, 66]	Ne	10	12	2.948(06)	[33]
23	0.04	Mg	12	11	3.043(04)	[37, 66]	Na	11	12	2.992(06)	
23	0.13	Al	13	10	<i>3.080(11)</i>		Ne	10	13	2.899(10)	Tab. 3
24	0.08	Al	13	11	<i>3.078(11)</i>		Na	11	13	<i>2.963(11)</i>	Tab. 6
24	0.17	Si	14	10	<i>3.124(10)</i>		Ne	10	14	2.894(08)	Tab. 3
25	0.04	Al	13	12	<i>3.082(04)</i>		Mg	12	13	3.026(03)	[37, 66]
25	0.12	Si	14	11	<i>3.129(15)</i>		Na	11	14	<i>2.963(14)</i>	Tab. 6
26	0.08	Si	14	12	<i>3.136(04)</i>		Mg	12	14	3.030(03)	

...

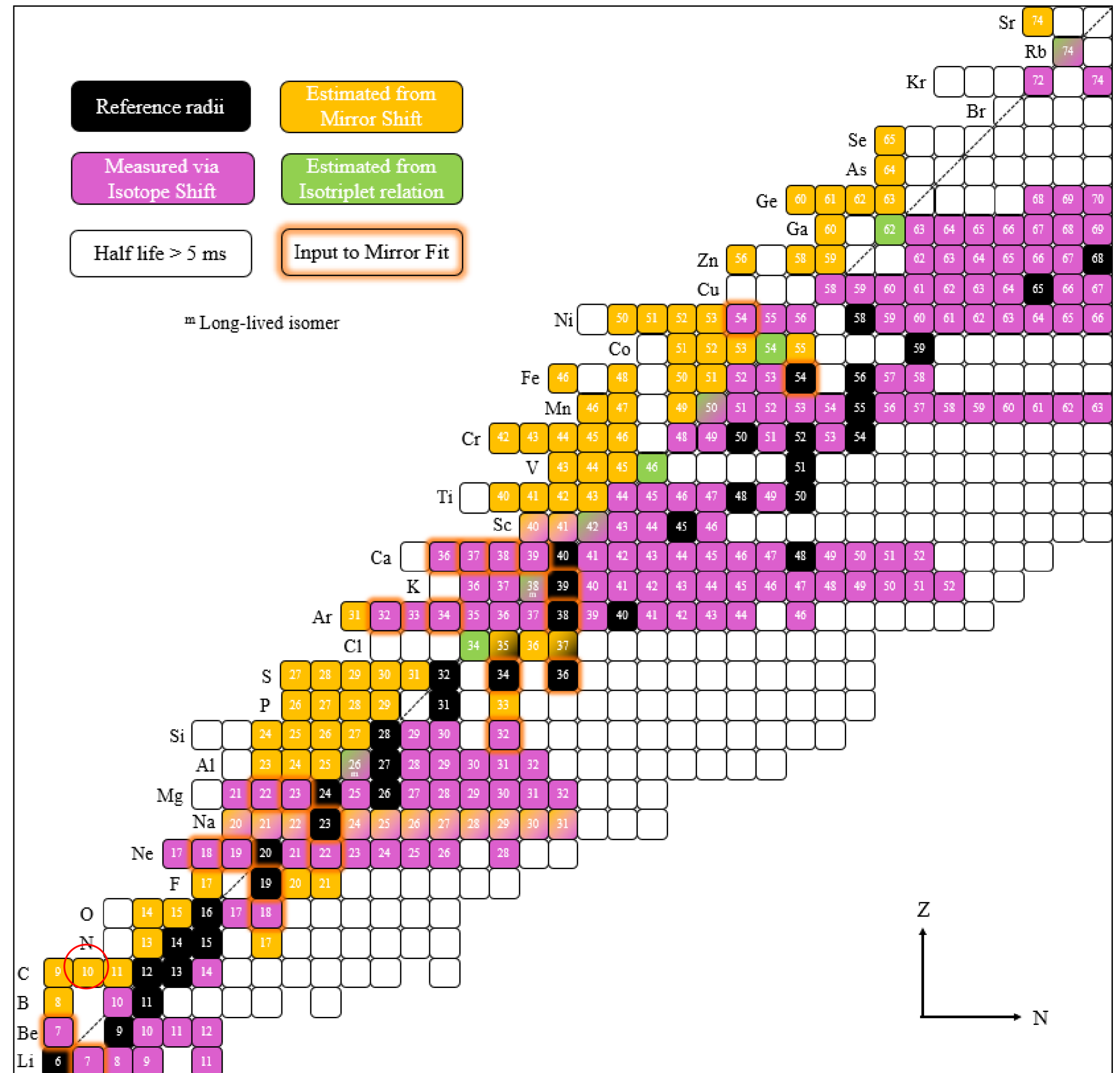
What can we do with this table?

What can we do with this table?

Much! I'll focus on  $V_{ud}$

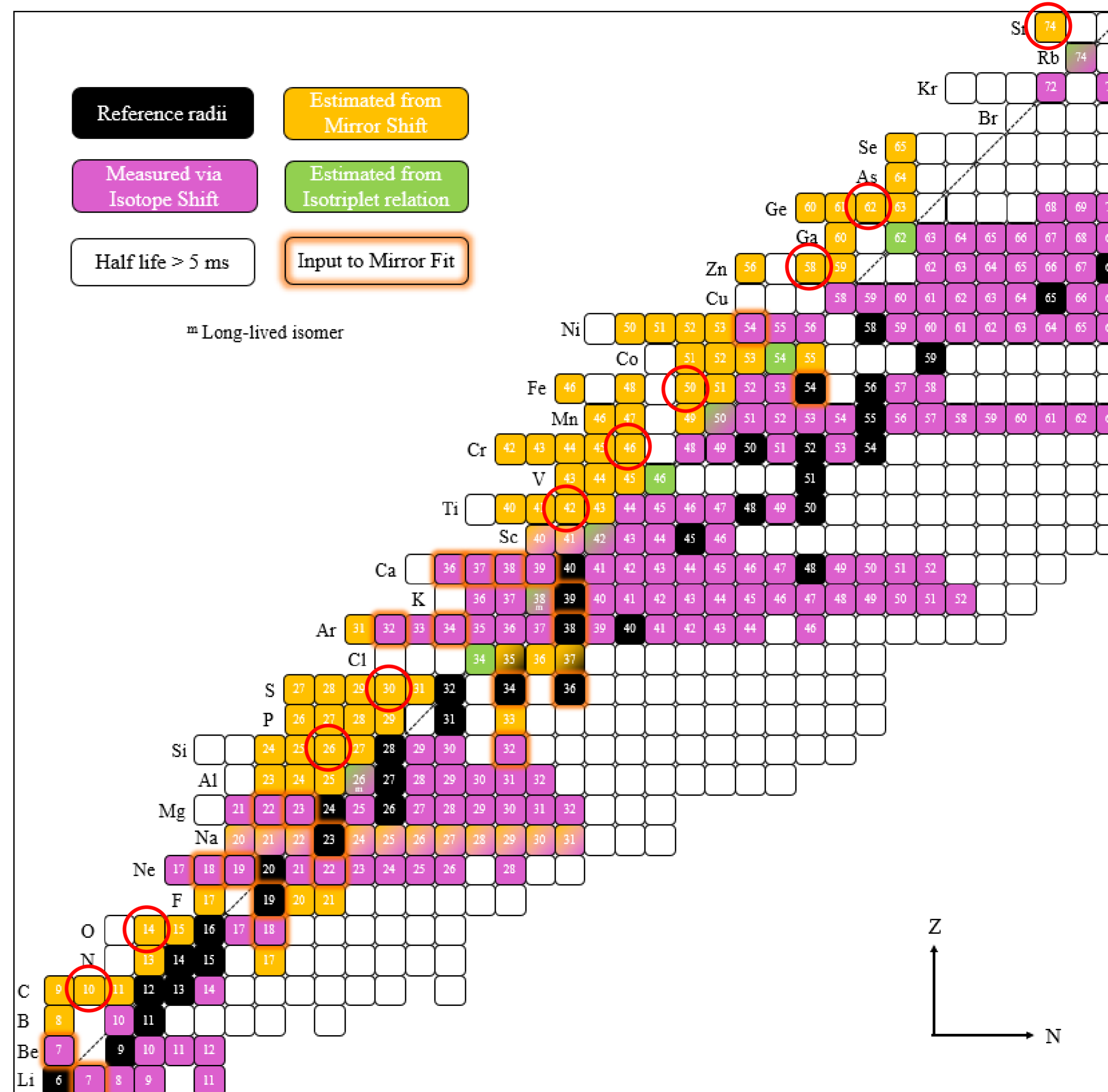


# Example 1: Missing charge radii for CKM



# Example 1: Missing charge radii for CKM

- Important for analysis of superallowed beta-decays:



# Example 1: Missing charge radii for CKM

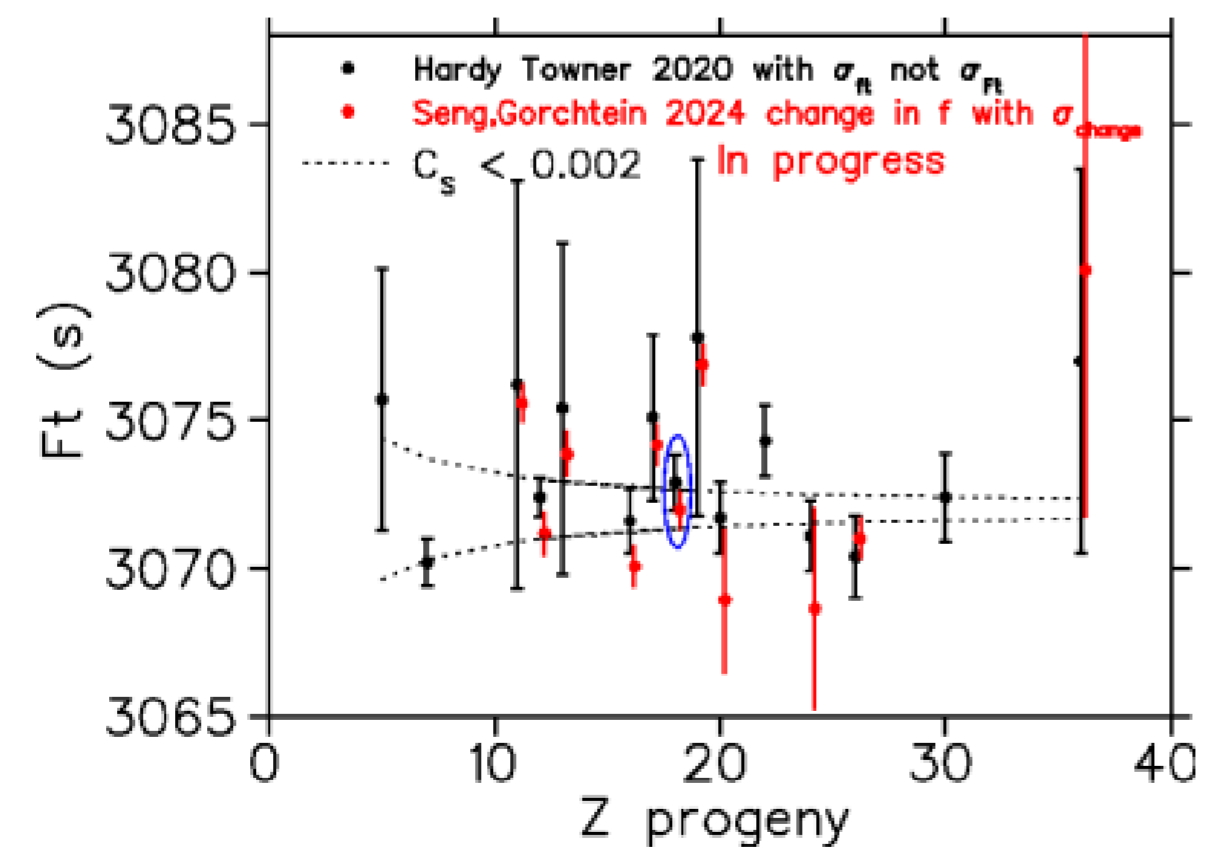
- Important for analysis of superallowed beta-decays:

PHYSICAL REVIEW C **109**, 045501 (2024)

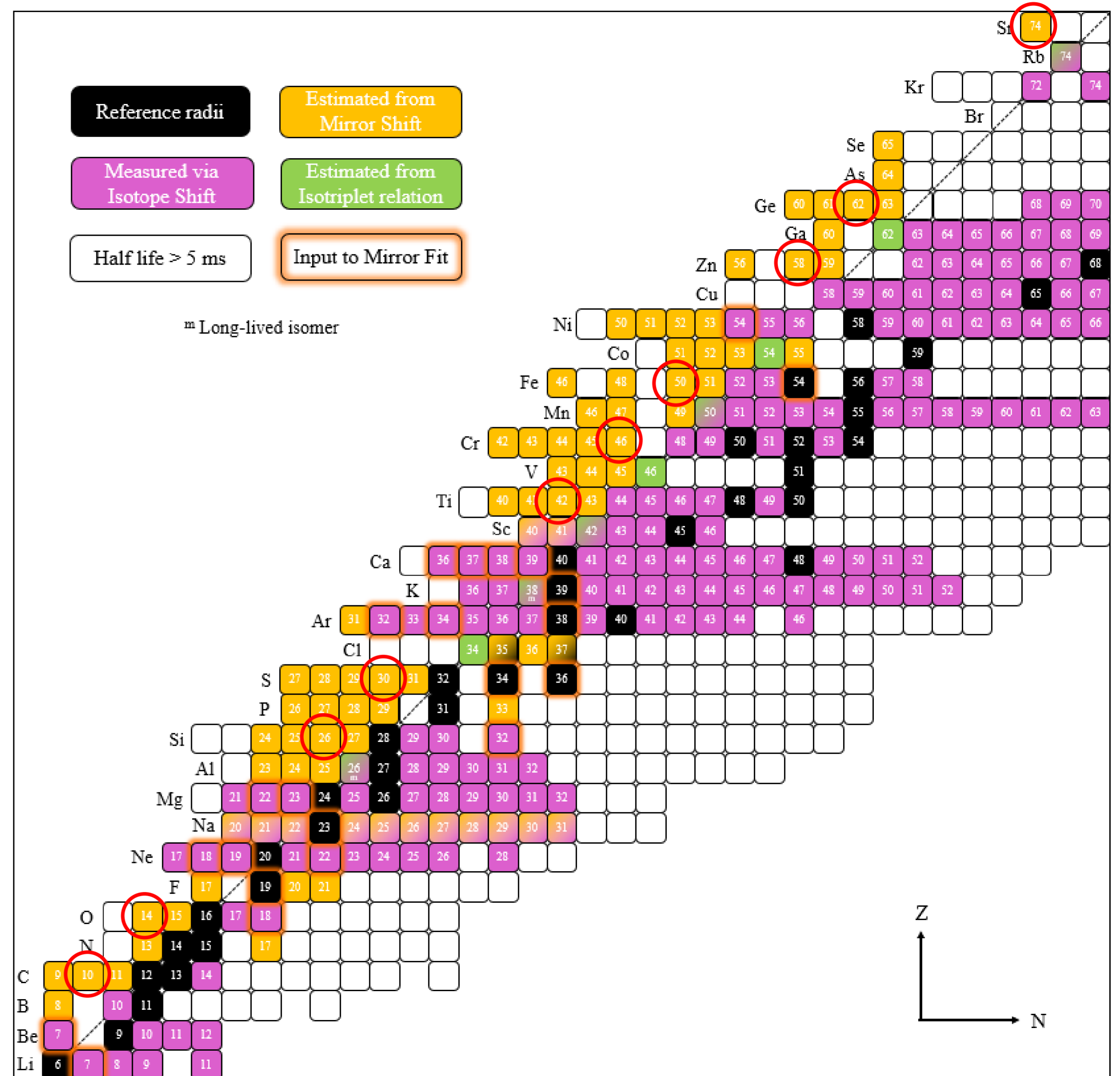
## Data-driven reevaluation of $ft$ values in superallowed $\beta$ decays

Chien-Yeah Seng<sup>1,2</sup> and Mikhail Gorchtein<sup>3,4</sup>

- Figure by J. Behr:



- New values not implemented (yet?)



# Example 2: Test of isospin-symmetry breaking (ISB)



ELSEVIER

Contents lists available at ScienceDirect

Physics Letters B

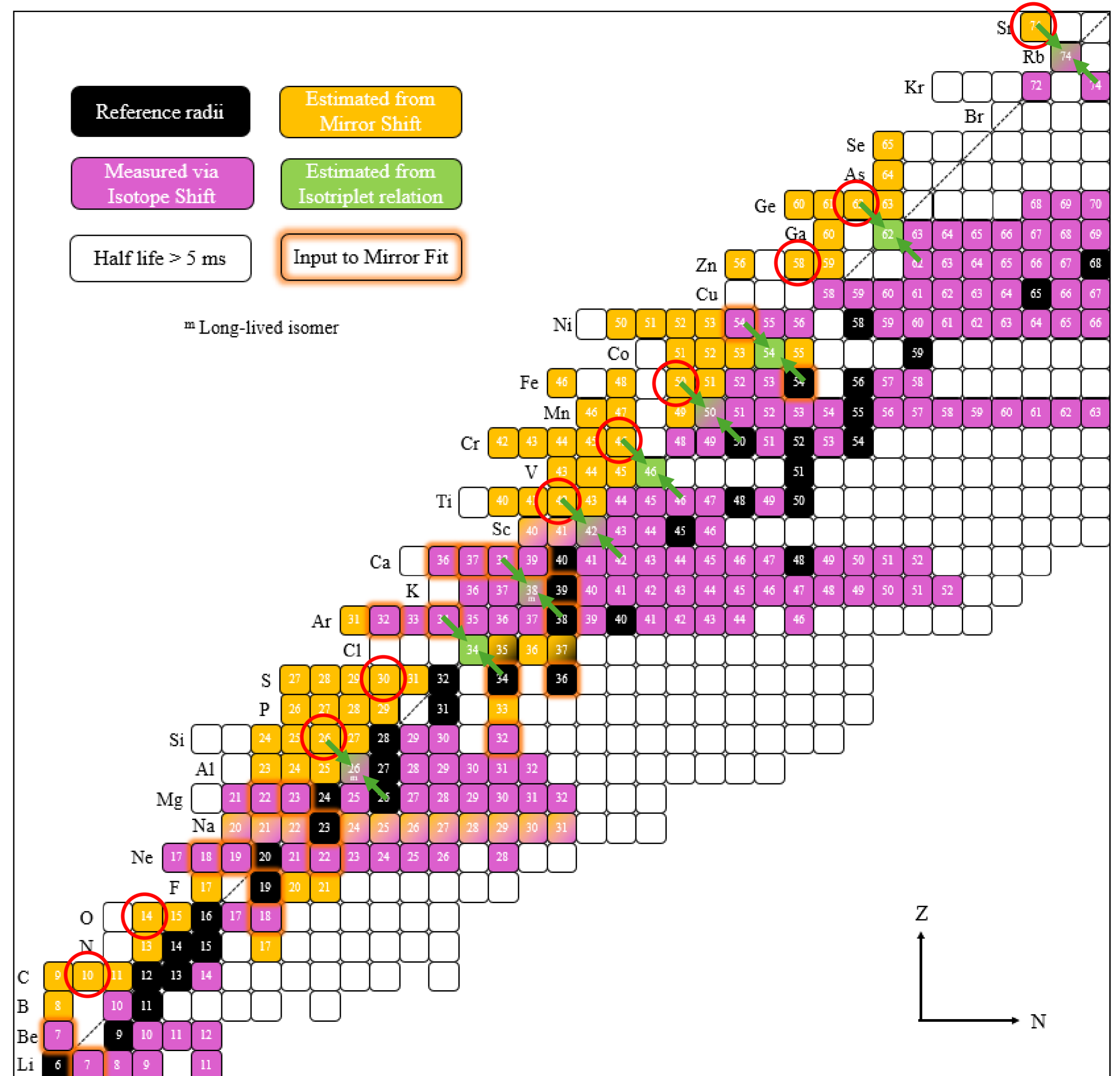
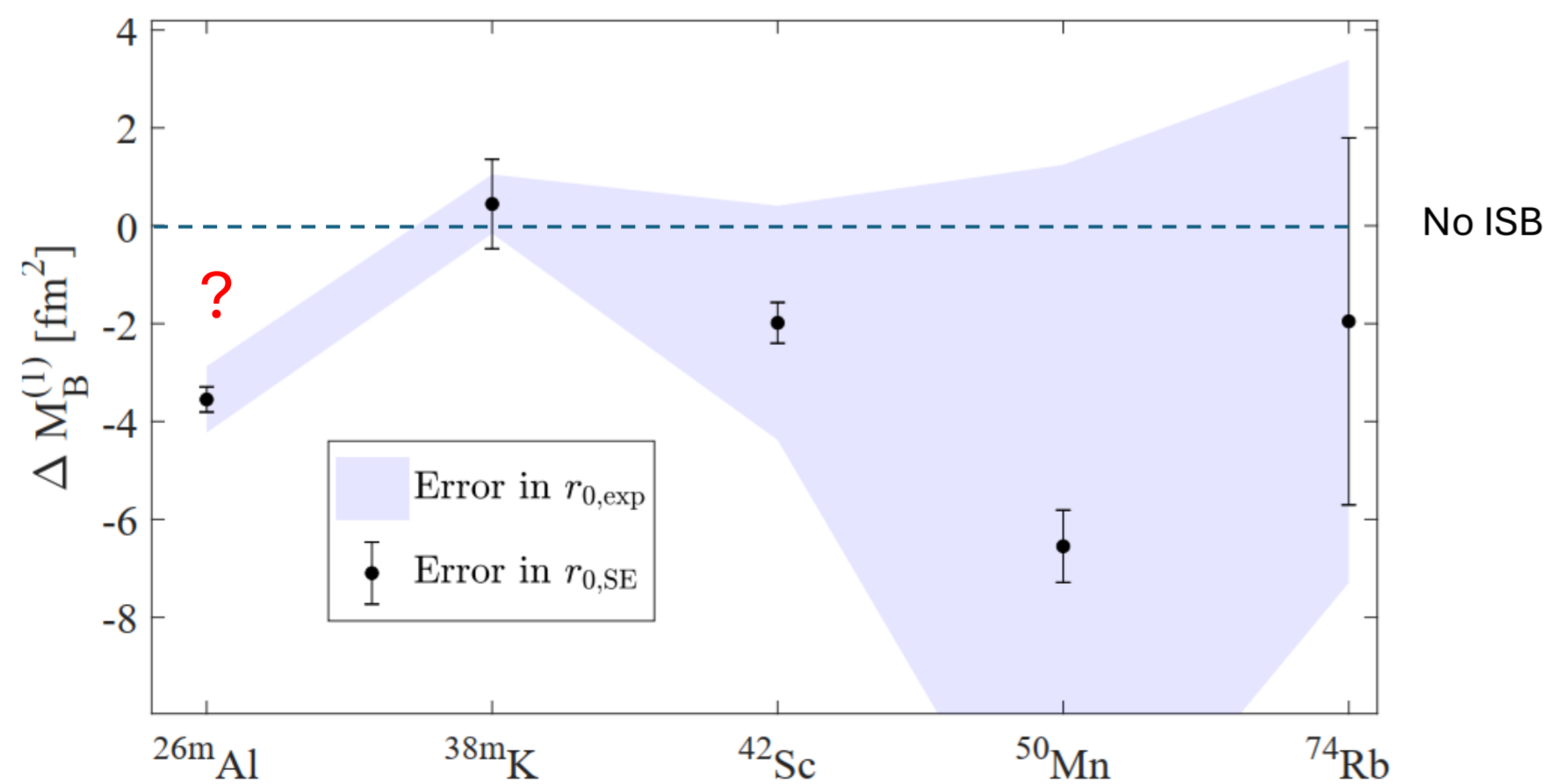
journal homepage: [www.elsevier.com/locate/physletb](http://www.elsevier.com/locate/physletb)

Electroweak nuclear radii constrain the isospin breaking correction to  $V_{ud}$

Chien-Yeah Seng<sup>a,b,c,\*</sup>, Mikhail Gorchtein<sup>d,e</sup>

Interpolated radii without ISB: 
$$r_{0,SE}^2 = r_{+1}^2 + \frac{Z-1}{2Z_0} \Delta_I (2r_{+1} + \Delta_I).$$

ISB test: 
$$\Delta M_B^{(1)} = Z_0 (r_{0,SE}^2 - r_{0,exp}^2)$$



# Example 3: Weak radii for (recoil corrections to) $V_{ud}$

[arXiv:2409.08193](https://arxiv.org/abs/2409.08193)

PHYSICAL REVIEW LETTERS **130**, 152501 (2023)

**Model-Independent Determination of Nuclear Weak Form Factors and Implications for Standard Model Precision Tests**

Chien-Yeah Seng 

$$R_{CW}^2 = R_{Ch,1}^2 + \frac{Z_{-1}}{2} (R_{Ch,-1}^2 - R_{Ch,1}^2),$$



With mirror fit

$$R_{CW}^2 = r_{+1}^2 + \frac{Z_{-1}}{2} \Delta_I (2r_{+1} + \Delta_I)$$



A	$r_{CW}^2$ fm <sup>2</sup>	Seng 2023
10	9.72(25)	N/A
14	10.41(12)	N/A
18	12.08(12)	13.4(5)
22	13.24(12)	12.9(7)
26	13.77(12)	N/A
30	14.50(13)	N/A
34	15.66(13)	15.6(5)
38	16.58(13)	16.0(3)
42	17.46(13)	21.5(3.6)
46	18.29(14)	N/A
50	18.73(14)	23.2(3.8)
54	18.93(14)	18.3(9)
58	19.66(14)	N/A
62	20.65(15)	N/A
74	23.32(19)	19.5(5.5)

# Summary

1. Charge Radii of mirror nuclei crucial to determining  $V_{ud}$
2. If mirror relation holds, then can be predicted by global analysis
3. Missing information on light nuclei, and asymmetric pairs
4. QUARTET focused on improved Reference radii from muonic atoms using MMC detectors (next talk!)

