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DEGLI STUDI  
FIRENZE  
DIPARTIMENTO DI  
FISICA E STRONOMIA

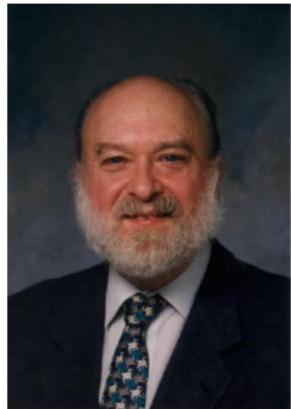
# Quantum kinetic theory and spin hydrodynamics for spin- $1/2$ and spin-1 particles

David Wagner

Spin and quantum features of QCD plasma  
20.09.2024

## Introduction to the framework

# What is quantum kinetic theory?



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# What is quantum kinetic theory?

- **Nonequilibrium** statistical description of a **dilute** gas
  - ▶ **Nonequilibrium**: Describes process of thermalization selfconsistently, concept of temperature emerges from dynamics
  - ▶ **Statistical**: Quantity of interest: single-particle distribution function
  - ▶ **Dilute**: Main assumption: System consists of essentially pointlike free particles scattering via short-range interactions



# What is quantum kinetic theory?

Ihr solltet mein Papier lesen!

L. Boltzmann, Sitz.-Ber. Akad. Wiss. Wien (II) 66, 275–370 (1872)

english translation: The Kinetic Theory of Gases, 262–349 (2003)

- **Nonequilibrium** statistical description of a dilute gas
  - ▶ **Nonequilibrium:** Describes process of thermalization selfconsistently, concept of temperature emerges from dynamics
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# Connection to macroscopic currents

- Kinetic theory is an effective microscopic description
  - ▶ Provides (infinitely) more information than macroscopic approaches, such as thermo- and hydrodynamics
  - ▶ Can be used to extract information about any macroscopic current of interest

## Kinetic representation of currents

$$N^\mu(t, \mathbf{x}) = \int dK k^\mu f(t, \mathbf{x}, \mathbf{k}),$$

$$T^{\mu\nu}(t, \mathbf{x}) = \int dK k^\mu k^\nu f(t, \mathbf{x}, \mathbf{k}),$$

$$S^\mu(t, \mathbf{x}) = - \int dK k^\mu f(t, \mathbf{x}, \mathbf{k}) [\ln f(t, \mathbf{x}, \mathbf{k}) - 1]$$

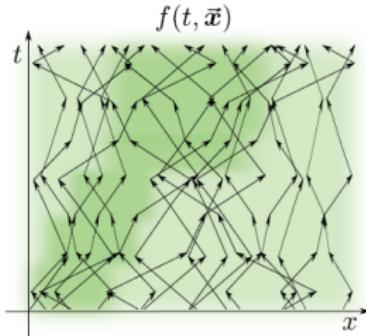
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$$dK := d^3k / [(2\pi\hbar)^3 k^0]$$

# Evolution of the distribution function

## Boltzmann equation (classical version)

$$\left( \frac{\partial}{\partial t} + \frac{\mathbf{k}}{m} \cdot \nabla + \mathbf{F} \cdot \nabla_{\mathbf{k}} \right) f(\mathbf{x}, \mathbf{k}) = C[f]$$



L. Rezzolla, O. Zanotti, 978-0-19-174650-5 (2013)

- Left-hand side: Advection through  $(\mathbf{x}, \mathbf{k})$ -phase space
- Right-hand side: Collision term
  - ▶ Depends on higher-order distribution functions, e.g.  
 $f_2(\{\mathbf{x}_1, \mathbf{k}_1\}; \{\mathbf{x}_2, \mathbf{k}_2\})$
  - ▶ Has to be truncated ( $\rightarrow$  BBGKY hierarchy)
  - ▶ Stoßzahlansatz: Replace  $f_2(\{\mathbf{x}_1, \mathbf{k}_1\}; \{\mathbf{x}_2, \mathbf{k}_2\}) \rightarrow f(\mathbf{x}_1, \mathbf{k}_1)f(\mathbf{x}_2, \mathbf{k}_2)$

## Collision term

$$C[f] = \frac{1}{2} \int dK_1 dK_2 dK' \mathcal{W} (f_1 f_2 - f f')$$

# Connecting to quantum theory

- How to translate these ideas to quantum mechanics?
  - ▶ Try to build on the conserved currents and find  $W(x, k, t)$  such that

$$\text{Tr} \left[ \hat{\rho}(\hat{\mathbf{x}}, \hat{\mathbf{p}}, t) \hat{A}(\hat{\mathbf{x}}, \hat{\mathbf{p}}) \right] = \int d^3x \int \frac{d^3k}{(2\pi\hbar)^3} A(\mathbf{x}, \mathbf{k}) W(\mathbf{x}, \mathbf{k}, t)$$

- Choice “closest” to classical kinetic theory: **Wigner function**

$$W(\mathbf{x}, \mathbf{k}, t) = \int d^3v e^{-\frac{i}{\hbar} \mathbf{k} \cdot \mathbf{v}} \left\langle \mathbf{x} + \frac{\mathbf{v}}{2} \middle| \hat{\rho} \middle| \mathbf{x} - \frac{\mathbf{v}}{2} \right\rangle$$

E. P. Wigner, Phys. Rev. 40, 749-760 (1932)

H.-W. Lee, Physics Reports 259, 147-211 (1995)

- Price to pay: Wigner function is not positive semidefinite
- Relativistic field-theoretical version (scalar field):

$$W(x, k) = \int d^4v e^{-\frac{i}{\hbar} \mathbf{k} \cdot \mathbf{v}} \left\langle \hat{\phi}^\dagger \left( x + \frac{v}{2} \right) \hat{\phi} \left( x - \frac{v}{2} \right) \right\rangle \equiv G^<(x, k)$$

# Equations of motion



$$G = \begin{pmatrix} G^{++} & G^{+-} \\ G^{-+} & G^{--} \end{pmatrix} = \begin{pmatrix} G^F & G^< \\ G^> & G^{\bar{F}} \end{pmatrix}$$

- Real-time QFT: Expectation values of operators can be represented as time evolution along a closed-time path
- EoM: (contour-ordered) Dyson-Schwinger equation

$$\begin{aligned} G_0^{-1} G^{AB}(x_1, x_2) &= -ic^{AB} \delta^{(4)}(x_1 - x_2) \\ &\quad + i \int d^4 x' \Sigma^{AC}(x_1, x') c_{CD} G^{DB}(x', x_2) \end{aligned}$$



- Wigner transform:  $-i\hbar\partial_1^\mu G^{AB}(x_1, x_2) \rightarrow \left(k^\mu - \frac{i\hbar}{2}\partial^\mu\right) G^{AB}(x, k)$

# Approximations

- $\hbar$ -gradient expansion: Assume that the two-point functions are sufficiently localized in central coordinate  $x = (x_1 + x_2)/2$ 
  - ▶ *Notion of a particle should make sense!*
- Allows to approximate memory integrals



$$\int d^4x' G(x_1, x') \Sigma(x', x_2) \longrightarrow G(x, k) \Sigma(x, k) - \frac{i\hbar}{2} \{G(x, k), \Sigma(x, k)\}_{\text{PB}}$$

## (Gradient-expanded) Kadanoff-Baym equations

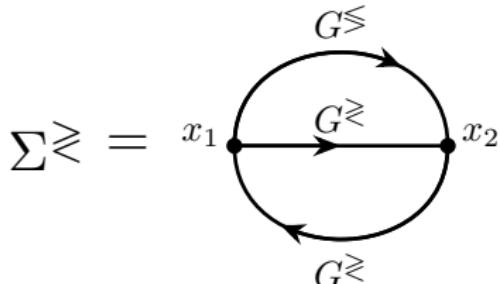
$$G_0^{-1} G^<(x, k) = \frac{i}{2} [\Sigma^>(x, k) G^<(x, k) - \Sigma^<(x, k) G^>(x, k)] \\ + \frac{\hbar}{4} \left[ \{\Sigma^>(x, k), G^<(x, k)\}_{\text{PB}} - \{\Sigma^<(x, k), G^>(x, k)\}_{\text{PB}} \right]$$

L. P. Kadanoff, G. Baym, ISBN 9780429493218 (1989)

$$\{f, g\}_{\text{PB}} := (\partial_\mu f)(\partial_k^\mu g) - (\partial_k^\mu f)(\partial_\mu g)$$

## Example: Scalar field

- Approximate self-energy to lowest nontrivial order
- Separate real and imaginary parts of KB equations



### Quantum kinetic equations (lowest order)

$$\begin{aligned} \left(k^2 - m^2\right) G^<(x, k) &= 0 \quad \Rightarrow \quad G^<(x, k) = 2\pi\hbar^2\delta(k^2 - m^2)f(x, k), \\ k \cdot \partial f(x, k) &= \frac{1}{2} \int dK_1 dK_2 dK' (2\pi\hbar)^4 \delta^{(4)}(k_1 + k_2 - k - k') \\ &\quad \times \frac{|M|^2}{16} \left(f_1 f_2 \tilde{f} \tilde{f}' - \tilde{f}_1 \tilde{f}_2 f f'\right) \end{aligned}$$

*“Tell me something new!”*  
Including interesting stuff

# Including electromagnetic fields

- Standard definition of Wigner function is not gauge invariant due to fields at different positions
  - Remedy: Include gauge link
$$U(x_1, x_2) = \exp \left[ -\frac{i}{\hbar} (x_1 - x_2) \cdot \int_{-1/2}^{1/2} dt A(x_1 + x_2 + t(x_1 - x_2)) \right]$$

## Wigner function with EM fields

$$G^<(x, k) = \int d^4v e^{-\frac{i}{\hbar} k \cdot v} \left\langle \hat{\phi}^\dagger \left( x + \frac{v}{2} \right) U \left( x + \frac{v}{2}, x - \frac{v}{2} \right) \hat{\phi} \left( x - \frac{v}{2} \right) \right\rangle$$

- Main effect: Wigner representation of momentum **changes**
$$-i\hbar \partial_1^\mu G^< \rightarrow \left\{ k^\mu - \frac{\hbar}{2} j_1(\Delta) F^{\mu\nu} \partial_{k,\nu} - \frac{i\hbar}{2} [\partial^\mu - j_0(\Delta) F^{\mu\nu} \partial_{k,\nu}] \right\} G^<(x, k)$$

D. Vasak, M. Gyulassy, H. T. Elze, Annals Phys. 173, 462-492 (1987)

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$$\Delta := (\hbar/2) \partial \cdot \partial_k$$

# Including spin

- Case of nonzero spin: Wigner function becomes **matrix-valued**
  - Additional components encode spin degrees of freedom

## Wigner function (spin 1/2)

$$G^< = \frac{1}{4} \left( \mathcal{F} + i\gamma_5 \mathcal{P} + \mathcal{V} + \gamma_5 \mathcal{A} + \frac{i}{4} [\gamma_\mu, \gamma_\nu] \mathcal{S}^{\mu\nu} \right)$$

- Underlying equations (Dirac, Proca, ...) can be solved perturbatively in  $\hbar$  expansion
  - Gradient contributions appear!
    - $\mathcal{V}^\mu \sim -\frac{\hbar}{2m} \partial_\nu \mathcal{S}^{\nu\mu}$ ,  $k_\mu \mathcal{S}^{\mu\nu} \sim -\frac{\hbar}{2} \partial^\nu \mathcal{F}$ ,  $\mathcal{P} \sim -\frac{\hbar}{4m^2} \epsilon^{\mu\nu\alpha\beta} k_\mu \partial_\nu \mathcal{S}_{\alpha\beta}$ , ...
  - Responsible for a lot of the interesting transport phenomena

H. T. Elze, M. Gyulassy, D. Vasak, Nucl. Phys. B 276, 706-728 (1986)

D. Vasak, M. Gyulassy, H. T. Elze, Annals Phys. 173, 462-492 (1987)

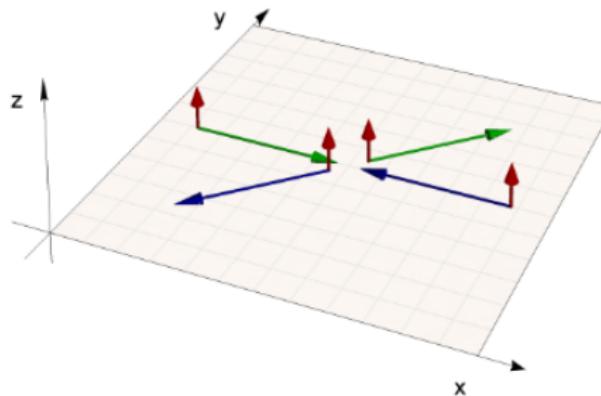
S. Mrowczynski, U. W. Heinz, Annals Phys. 229, 1-54 (1994)

Y. A. Markov, M. A. Markova, Theor. Math. Phys. 108, 977-991 (1996)

Y. A. Markov, M. A. Markova, Theor. Math. Phys. 111, 601-612 (1997)

Going nonlocal: Collisions (with spin)

# Angular momentum and collisions

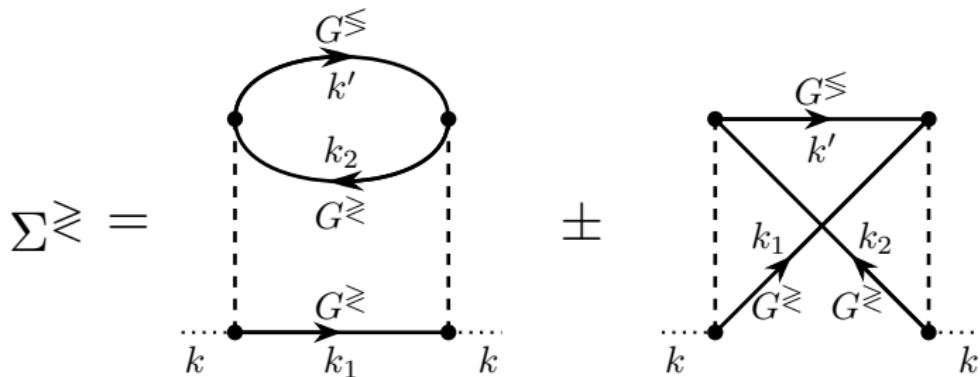


W. Florkowski, A. Kumar, R. Ryblewski, Prog. Part. Nucl. Phys. 108, 103709 (2019)

- Assume that collisions take place in a point
  - Total orbital angular momentum vanishes
  - Spin is conserved on its own
  - No exchange of spin and orbital angular momenta
- Collisions must be **nonlocal** for spin equilibration!

# Collisions from QKT

$$k \cdot \partial G^<(x, k) = \frac{1}{2} [\Sigma^<(x, k)G^>(x, k) - \Sigma^<(x, k)G^<(x, k)]$$



- Collisions determined by self-energies
- Crucial quantum enhancement: *All internal lines have to be evaluated to order  $\mathcal{O}(\hbar)$ !*  
→ Introduces gradient corrections inside the collision integral!

# Collisions from QKT

DW, NW, DHR, Phys.Rev.D 106 (2022) 11, 116021

## Boltzmann equation with collisions

$$f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) = \delta(k^2 - m^2) f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) := \frac{1}{2} [\mathcal{F}(\textcolor{red}{x}, \textcolor{blue}{k}) - \textcolor{green}{s} \cdot \mathcal{A}(\textcolor{red}{x}, \textcolor{blue}{k})]$$

$$\begin{aligned} k \cdot \partial f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) &= \frac{1}{2} \int d\Gamma_1 d\Gamma_2 d\Gamma' d\bar{S}(k) \delta^{(4)}(k_1 + k_2 - k - k') \mathcal{W} \\ &\quad \times [f(\textcolor{red}{x} + \Delta_1 - \Delta, \textcolor{blue}{k}_1, \textcolor{green}{s}_1) f(\textcolor{red}{x} + \Delta_2 - \Delta, \textcolor{blue}{k}_2, \textcolor{green}{s}_2) \\ &\quad - f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) f(\textcolor{red}{x} + \Delta' - \Delta, \textcolor{blue}{k}', \textcolor{green}{s}')] \end{aligned}$$

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$$d\Gamma := 2d^4k \delta(k^2 - m^2) dS(k)$$

# Collisions from QKT

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- Contributions inside the collision term have gradient corrections

$$f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) + \Delta^\mu \partial_\mu f(\textcolor{red}{x}, \textcolor{blue}{k}, \textcolor{green}{s}) \approx f(\textcolor{red}{x} + \Delta, \textcolor{blue}{k}, \textcolor{green}{s})$$

- A (momentum- and spin-dependent) **spacetime shift**  $\Delta^\mu$  enters  
→ Particles do not scatter at the same spacetime point!
- This enables a conversion of orbital and spin angular momenta

$$d\Gamma := 2d^4k \delta(k^2 - m^2) dS(k)$$

# Collisions from QKT

## More on collisional QKT:

- S. Mrowczynski, U. W. Heinz, Annals Phys. 229, 1-54 (1994)  
K. Morawetz, P. Lipavsky, V. Spicka, N.-H. Kwong, Phys. Rev. C 59, 3052–3059 (1999)  
D.-L. Yang, K. Hattori, Y. Hidaka, Yoshimasa JHEP 07, 070 (2020)  
K. Hattori, Y. Hidaka, N. Yamamoto, D.-L. Yang, JHEP 02, 001 (2021)  
N. Weickgenannt, E. Speranza, X.-L. Sheng, Q. Wang, D. H. Rischke, Phys. Rev. D 104, 016022 (2021)  
X.-L. Sheng, N. Weickgenannt, E. Speranza, D. H. Rischke, Q. Wang, Phys. Rev. D 104, 016029 (2021)  
J. Hu, 2110.12339 (2021)  
S. Lin, Phys. Rev. D 105, 076017 (2022)  
S. Fang, S. Pu, D.-L. Yang, Phys. Rev. D 106, 016002 (2022)  
Y. Hidaka, S. Pu, Q. Wang, D.-L. Yang, Prog. Part. Nucl. Phys. 127, 103989 (2022)  
DW, N. Weickgenannt, D. H. Rischke, Phys. Rev. D 106, 116021 (2022)  
X.-L. Sheng, Q. Wang, D. H. Rischke, Phys. Rev. D 106, L111901 (2022)  
DW, N. Weickgenannt, E. Speranza, Phys. Rev. D 108, 116017 (2023)  
N. Yamamoto, D.-L. Yang, Phys. Rev. D 109, 056010 (2024)

- A (momentum- and spin-dependent) **spacetime shift**  $\Delta^\mu$  enters  
→ Particles do not scatter at the same spacetime point!
- This enables a conversion of orbital and spin angular momenta

$$d\Gamma := 2d^4k \delta(k^2 - m^2) dS(k)$$

# Finding equilibrium

- Local-equilibrium distribution function makes local collision term vanish (without spacetime shifts  $\Delta^\mu$ )
- Has to depend on the **collisional invariants**
  - Charge, four-momentum and total angular momentum

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- Local-equilibrium distribution function makes local collision term vanish (without spacetime shifts  $\Delta^\mu$ )
- Has to depend on the **collisional invariants**
  - Charge, four-momentum and total angular momentum

## Local-equilibrium distribution function

$$f_{\text{eq}}(x, k, \mathfrak{s}) = \exp \left( \alpha_0 - \beta_0 E_{\mathbf{k}} + \frac{\hbar}{2} \Omega_{0,\mu\nu} \Sigma_{\mathfrak{s}}^{\mu\nu} \right)$$

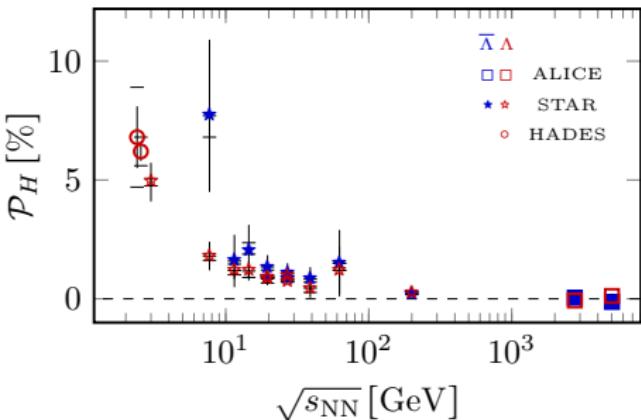
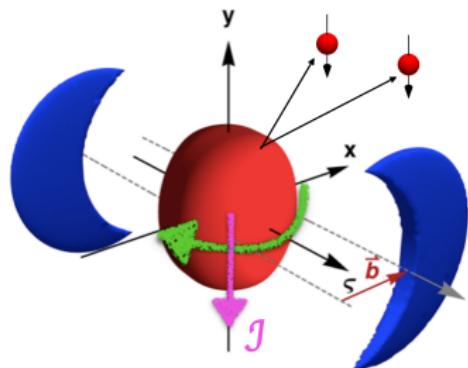
- Necessary conditions on Lagrange multipliers for a vanishing **nonlocal** collision term:  $\partial^\mu \alpha_0 = 0$ ,  $\partial^{(\mu} (\beta_0 u^{\nu)}) = 0$ ,  $\Omega_0^{\mu\nu} = -\frac{1}{2} \partial^{[\mu} (\beta_0 u^{\nu}])$
- Same conditions as for **global** equilibrium, where  $k \cdot \partial f_{\text{eq}} = 0$

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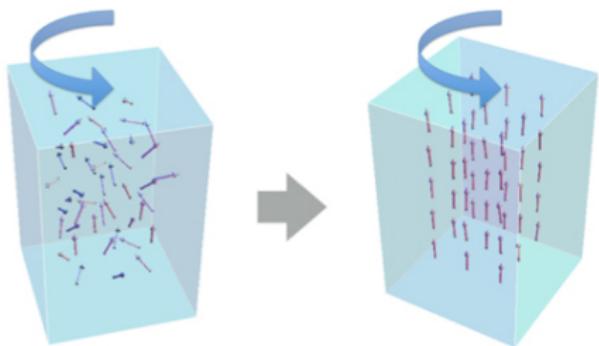
$$\Sigma_{\mathfrak{s}}^{\mu\nu} := -\frac{1}{m} \epsilon^{\mu\nu\alpha\beta} k_\alpha \mathfrak{s}_\beta, \quad E_{\mathbf{k}} := k \cdot u$$

# Polarization & spin hydrodynamics

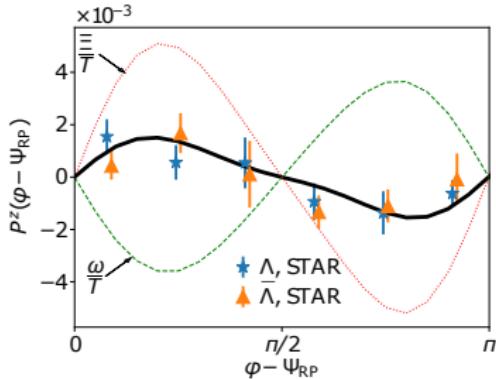
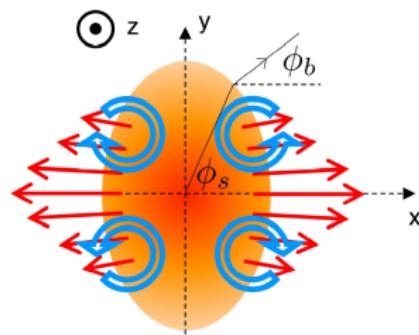
# Global $\Lambda$ -Polarization



- “Global”: Integrated polarization along the direction of orbital angular momentum
- Can be explained by assuming spins in equilibrium
- “Polarization through rotation”
  - Analogous to Barnett effect

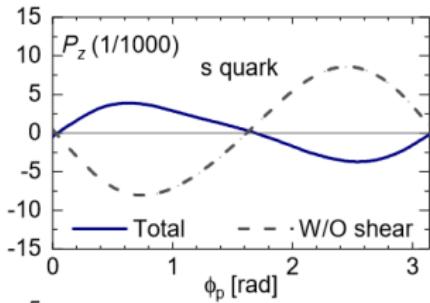


# Local $\Lambda$ -Polarization



F. Becattini, M. Buzzegoli, G. Inghirami, I. Karpenko, A. Palermo,  
PRL 127, 272302 (2021)

- “Local”: Angle-dependent polarization along beam-direction
- Could only be explained recently by incorporating shear effects
- Simple picture of equilibrated spins not complete



B. Fu, S. Y. F. Liu, L. Pang, H. Song, Y. Yin, Phys. Rev. Lett. 127, 142301 (2021)

# Open questions

## Pauli-Lubanski vector in (global) equilibrium

$$S_{\varpi}^{\mu} = -\epsilon^{\mu\nu\alpha\beta} k_{\nu} \frac{\int d\Sigma_{\lambda} k^{\lambda} f_0 (1 - f_0) \varpi_{\alpha\beta}}{8m \int d\Sigma_{\lambda} k^{\lambda} f_0}$$

- Traditional approaches to computing the polarization
  - ▶ assume equilibrated spin degrees of freedom
  - ▶ neglect dissipative terms

$$\varpi_{\mu\nu} := \frac{1}{2} [\partial_{\mu} (u_{\nu}/T) - \partial_{\nu} (u_{\mu}/T)]$$

# Open questions

## Pauli-Lubanski vector in (global) equilibrium

$$S_{\varpi}^{\mu} = -\epsilon^{\mu\nu\alpha\beta} k_{\nu} \frac{\int d\Sigma_{\lambda} k^{\lambda} f_0 (1 - f_0) \varpi_{\alpha\beta}}{8m \int d\Sigma_{\lambda} k^{\lambda} f_0}$$

- Traditional approaches to computing the polarization
    - ▶ assume equilibrated spin degrees of freedom
    - ▶ neglect dissipative terms
  - Not clear so far:
    - (I) How fast do spin degrees of freedom equilibrate?
    - (II) How do dissipative effects influence polarization?
- Can be answered through spin hydrodynamics from quantum kinetic theory

$$\varpi_{\mu\nu} := \frac{1}{2} [\partial_{\mu} (u_{\nu}/T) - \partial_{\nu} (u_{\mu}/T)]$$

# (I) Timescales: Magnetic resonance imaging (MRI)

- MRI: Large constant  $B$ -field in  $z$ -direction and short-lived alternating field in  $x, y$ -plane
- Identify materials by relaxation times  $T_1, T_2$



[https://en.wikipedia.org/wiki/Bloch\\_equations](https://en.wikipedia.org/wiki/Bloch_equations)

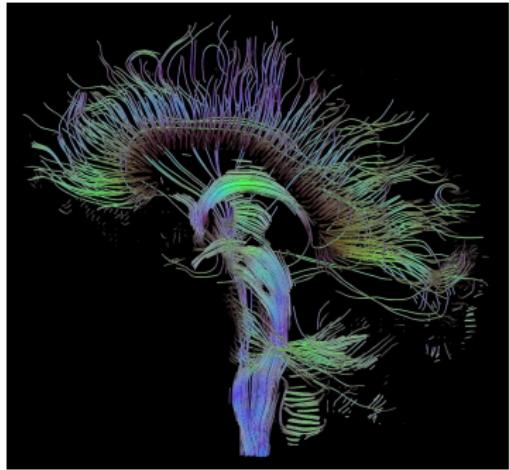
## Bloch equations

$$\begin{aligned}T_2 \dot{M}_{x,y} + M_{x,y} &= \mu_2 (\mathbf{M} \times \mathbf{B})_{x,y} , \\T_1 \dot{M}_z + M_z &= \mu_1 (\mathbf{M} \times \mathbf{B})_z + M_0 .\end{aligned}$$

$$\mu_1 := T_1 \frac{gq}{2m}, \quad \mu_2 := T_2 \frac{gq}{2m}$$

## (II) Dissipation: Diffusion-MRI

- Apply additional  $\vec{B}$ -gradients to make Larmor frequencies position-dependent
- Allows to additionally track diffusion of molecules



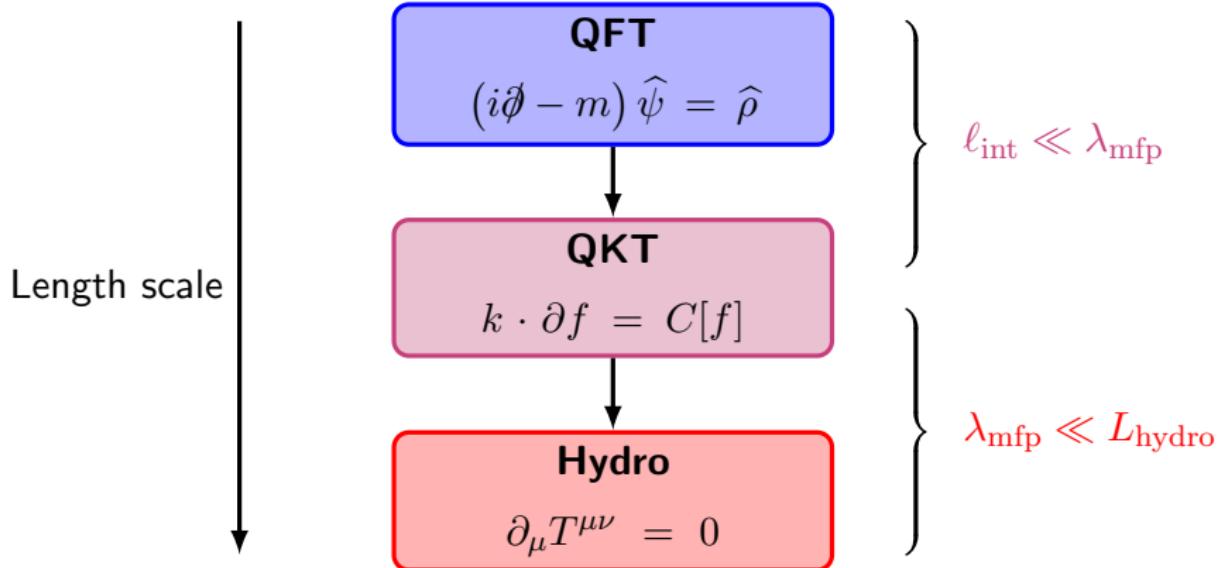
[https://en.wikipedia.org/wiki/Diffusion\\_MRI](https://en.wikipedia.org/wiki/Diffusion_MRI)

### Bloch-Torrey equations

$$\begin{aligned} T_2 \dot{M}_{x,y} + M_{x,y} &= \mu_2 (\mathbf{M} \times \mathbf{B})_{x,y} + D \Delta M_{x,y} , \\ T_1 \dot{M}_z + M_z &= \mu_1 (\mathbf{M} \times \mathbf{B})_z + D \Delta M_z + M_0 . \end{aligned}$$

$$\mu_1 := T_1 \frac{gq}{2m}, \quad \mu_2 := T_2 \frac{gq}{2m}$$

# Spin hydrodynamics: Procedure



# Spin hydrodynamics: Ideal case

## Conservation equations

$$\partial_\mu T^{(\mu\nu)} = 0 + \mathcal{O}(\hbar^2) ,$$

$$\partial_\lambda S^{\lambda\mu\nu} = \frac{1}{\hbar} T^{[\nu\mu]} + \mathcal{O}(\hbar^2) .$$

- No backreaction of spin on fluid evolution, fluid profile serves as input for spin potential
  - ▶ Simplest configuration: fluid at rest,  $u^\mu = \text{const}$ ,  $T = \text{const}$
- Ideal fluid (kinetic language): Assume local equilibrium, i.e., set  $f = f_{\text{eq}}$ 
  - ▶ Express  $S^{\lambda\mu\nu}$  and  $T^{[\nu\mu]}$  in terms of  $\kappa_0^\mu := \Omega_0^{\nu\mu} u_\nu$  and  $\omega_0^\mu := \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} u_\nu \Omega_{0,\alpha\beta}$

## Form of $T^{[\mu\nu]}$

$$T^{[\mu\nu]} = -\hbar^2 \Gamma^{(\kappa)} \left( u^{[\mu} \kappa_0^{\nu]} - u_\alpha \varpi^{\alpha[\nu} u^{\mu]} \right) + \hbar^2 \Gamma^{(\omega)} \left( \epsilon^{\mu\nu\alpha\beta} u_\alpha \omega_{0,\beta} - \varpi^{\langle\mu\rangle\langle\nu\rangle} \right)$$

# Spin equations of motion in a fluid at rest

- Project equations of motion for spin tensor to obtain evolution equations for the components of the spin potential

## Equations of motion for the spin potential

$$\tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu = \mu_\kappa \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \omega_{0,\beta} ,$$
$$\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu = -\mu_\omega \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \kappa_{0,\beta} .$$

- Go to fluid rest frame,  $\kappa_0^\mu \equiv (0, \kappa)$ ,  $\omega_0^\mu \equiv (0, \omega)$

$$\tau_\kappa \dot{\kappa} + \kappa = \mu_\kappa \nabla \times \omega ,$$

$$\tau_\omega \dot{\omega} + \omega = -\mu_\omega \nabla \times \kappa ,$$

- Relaxation times  $\tau_\kappa$ ,  $\tau_\omega$  determined by **nonlocal** collisions

$$\tau_\kappa := \frac{I_{31}}{2m^2 \Gamma(\kappa)}, \quad \tau_\omega := \frac{I_{30} - I_{31}}{4m^2 \Gamma(\omega)}, \quad \mu_\kappa := \frac{\tau_\kappa}{2}, \quad \mu_\omega := \frac{I_{31}}{4m^2 \Gamma(\omega)}$$

# Spin equations of motion in a fluid at rest

- Project equations of motion for spin tensor to obtain evolution equations for the components of the spin potential

## Equations of motion for the spin potential

$$\tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu = \mu_\kappa \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \omega_{0,\beta},$$

$$\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu = -\mu_\omega \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \kappa_{0,\beta}.$$

- Go to fl

Let  $\tau_\kappa \rightarrow 0$  for (boring) version of Bloch equations:

$$\tau_\omega \dot{\omega} + \omega = 0$$

- Relaxation times  $\tau_\kappa, \tau_\omega$  determined by **nonlocal** collisions

$$\tau_\kappa := \frac{I_{31}}{2m^2\Gamma(\kappa)}, \quad \tau_\omega := \frac{I_{30}-I_{31}}{4m^2\Gamma(\omega)}, \quad \mu_\kappa := \frac{\tau_\kappa}{2}, \quad \mu_\omega := \frac{I_{31}}{4m^2\Gamma(\omega)}$$

# Spin waves

- $\kappa$  and  $\omega$  follow coupled relaxation equations  
→ Disentangle longitudinal and transverse components

## Longitudinal components: Decay

$$\tau_\kappa \frac{d}{dt} (\nabla \cdot \kappa) = -\nabla \cdot \kappa ,$$
$$\tau_\omega \frac{d}{dt} (\nabla \cdot \omega) = -\nabla \cdot \omega ,$$

## Transverse components: Damped waves

$$\ddot{\kappa} + a\dot{\kappa} + b\kappa - c_s^2 \Delta \kappa = 0 ,$$
$$\ddot{\omega} + a\dot{\omega} + b\omega - c_s^2 \Delta \omega = 0 ,$$

V. E. Ambrus, R. Ryblewski, R. Singh, Phys. Rev. D 106, 014018(2022)

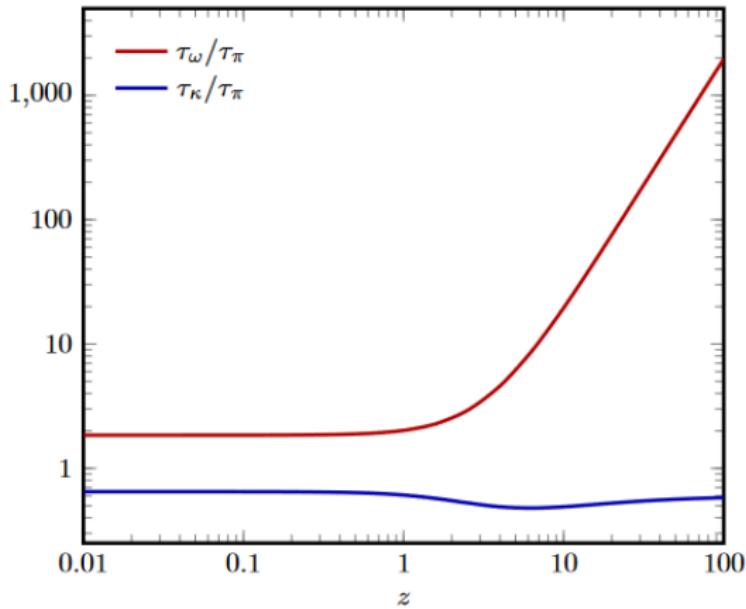
J. Hu, Z. Xu, Phys. Rev. D 107, 016010 (2023)

DW, M. Shokri, D. H. Rischke, 2405.00533 (2024)

$$a := \frac{\tau_\kappa + \tau_\omega}{\tau_\kappa \tau_\omega} , \quad b := \frac{1}{\tau_\kappa \tau_\omega} , \quad c_s^2 := \frac{\mu_\kappa \mu_\omega}{\tau_\kappa \tau_\omega} .$$

# Spin relaxation timescales

- Relaxation time depends on interaction model, can potentially get large
  - Spin relaxation can be a slow process, dynamics cannot be neglected!

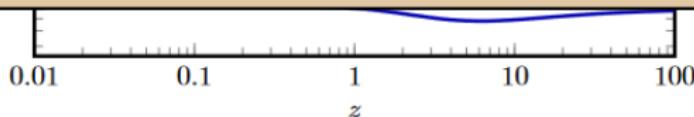


# Spin relaxation timescales

- Relaxation time depends on interaction model, can potentially get large
  - Spin relaxation can be a slow process, dynamics cannot be neglected!

## More on spin relaxation:

- Y. Hidaka, M. Hongo, M. A. Stephanov, H.-U. Yee, Phys. Rev. C 109, 054909 (2024)  
M. Hongo, X.-G. Huang, M. Kaminski, M. Stephanov, H.-U. Yee, JHEP 08, 263 (2022)  
A. Ayala et al, Phys. Rev. D 109, 074018 (2024)  
A. Ayala, D. de la Cruz, L. A. Hernández, J. Salinas, Phys. Rev. D 102, 056019 (2020)  
A. Ayala, D. de la Cruz, S. Hernández-Ortíz, L. A. Hernández, J. Salinas, Phys. Lett. B 801, 135169 (2020)  
J. I. Kapusta, E. Rrapaj, S. Rudaz, Phys. Rev. C 101, 024907 (2020)  
J. Hu, Phys. Rev. D 105, 096021 (2022)  
J. Hu, Z. Xu, Phys. Rev. D 107, 016010 (2023)  
DW, M. Shokri, D. H. Rischke, 2405.00533 (2024)  
S. Lin, H. Tang, 2406.17632 (2024)  
G. Torrieri, D. Montenegro, Phys. Rev. D 107, 076010 (2023)



DW, M. Shokri, D. H. Rischke, 2405.00533 (2024)

# Beyond equilibrium: Moment method

- Split distribution function  $f = f_{\text{eq}} + \delta f$
- Perform moment expansion including spin degrees of freedom

## Irreducible moments

$$\begin{aligned}\rho_{\textcolor{red}{r}}^{\mu_1 \dots \mu_\ell}(x) &:= \int d\Gamma \textcolor{blue}{E}_{\mathbf{k}}^r k^{\langle \mu_1} \dots k^{\mu_\ell \rangle} \delta f(x, k, \mathfrak{s}) \\ \tau_{\textcolor{red}{r}}^{\mu, \mu_1 \dots \mu_\ell}(x) &:= \int d\Gamma \textcolor{red}{s}^\mu \textcolor{blue}{E}_{\mathbf{k}}^r k^{\langle \mu_1} \dots k^{\mu_\ell \rangle} \delta f(x, k, \mathfrak{s})\end{aligned}$$

- Equations of motion can be derived from Boltzmann equation
- Knowing the evolution of all moments is equivalent to solving the Boltzmann equation

$$k^{\langle \mu_1 \dots \mu_\ell \rangle} := \Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} k^{\nu_1} \dots k^{\nu_\ell}$$

# Beyond equilibrium: Moment method

- Split distribution function  $f = f_{\text{eq}} + \delta f$
- Perform moment expansion including spin degrees of freedom

## Irreducible moments

### Standard dissipation

$$\rho_r^{\mu_1 \dots \mu_\ell}(x) := \int d\Gamma E_k^r k^{\langle \mu_1 \dots k^{\mu_\ell} \rangle} \delta f(x, k, \mathfrak{s})$$

$$\tau_r^{\mu, \mu_1 \dots \mu_\ell}(x) := \int d\Gamma \mathfrak{s}^\mu E_k^r k^{\langle \mu_1 \dots k^{\mu_\ell} \rangle} \delta f(x, k, \mathfrak{s})$$

### Spin dissipation

- Equations of motion can be derived from Boltzmann equation
- Knowing the evolution of all moments is equivalent to solving the Boltzmann equation

---

$$k^{\langle \mu_1 \dots k^{\mu_\ell} \rangle} := \Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} k^{\nu_1} \dots k^{\nu_\ell}$$

# Resumming (spin) hydrodynamics: IReD

DW, A. Palermo, V. E. Ambruş, Phys. Rev. D **106**, 016013 (2022)

DW, 2409.07143 (2024)

- Basic idea: Power-counting scheme to second order in
  - ▶ Knudsen number  $\text{Kn} := \lambda_{\text{mfp}}/L_{\text{hydro}}$
  - ▶ inverse Reynolds numbers  $\text{Re}^{-1} \sim \delta f/f_{\text{eq}}$
- Derive asymptotic (Navier-Stokes) relations to close the system

## Asymptotic matching (example)

$$\rho_r^{\mu\nu} = \eta_r \sigma^{\mu\nu} + \mathcal{O}(\text{KnRe}^{-1}) = \frac{\eta_r}{\eta_0} \pi^{\mu\nu} + \mathcal{O}(\text{KnRe}^{-1})$$

- The same procedure can be done for the moments  $\tau_r^{\mu,\mu_1\dots\mu_\ell}$
- Many moments can be related to  $\omega_0^\mu$  and  $\kappa_0^\mu$ 
  - ▶ No need to introduce more dynamical quantities
- Exception: tensor-valued moments  $t_r^{\mu\nu} := \tau_{r,\alpha,\beta}^{\langle\mu} \epsilon^{\nu\rangle\alpha\beta\rho} u_\rho$ 
  - ▶ Additional dynamical quantity  $t^{\mu\nu}$  is needed,  $S^{\lambda\mu\nu} \sim t^{\lambda[\mu} u^{\nu]}$

# Dissipative spin hydrodynamics

DW, 2409.07143 (2024)

$$\begin{aligned}\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu &= -\beta_0 \omega^\mu + \delta_{\omega\omega} \omega_0^\mu \theta + \lambda_{\omega\omega} \sigma^{\mu\nu} \omega_{0,\nu} + \lambda_{\omega t} \mathbf{t}^{\mu\nu} \omega_\nu \\ &\quad + \epsilon^{\mu\nu\alpha\beta} u_\nu (\ell_{\omega\kappa} \nabla_\alpha \kappa_{0,\beta} - \tau_\omega \dot{u}_\alpha \kappa_{0,\beta} + \lambda_{\omega\kappa} I_\alpha \kappa_{0,\beta} \\ &\quad + \ell_{\omega n} \nabla_\alpha n_\beta + \tau_{\omega n} \dot{u}_\alpha n_\beta + \lambda_{\omega n} I_\alpha n_\beta) \\ \tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu &= -\beta_0 \dot{u}^\mu + \mathfrak{b} I^\mu + \delta_{\kappa\kappa} \kappa_0^\mu \theta + \left( \lambda_{\kappa\kappa} \sigma^{\mu\nu} + \frac{\tau_\kappa}{2} \omega^{\mu\nu} \right) \kappa_{0,\nu} \\ &\quad + \epsilon^{\mu\nu\alpha\beta} u_\nu \left( \frac{\tau_\kappa}{2} \nabla_\alpha \omega_{0,\beta} + \tau_\kappa \dot{u}_\alpha \omega_{0,\beta} + \lambda_{\kappa\omega} I_\alpha \omega_{0,\beta} \right) \\ &\quad + \mathbf{t}^{\mu\nu} (\tau_{\kappa t} \dot{u}_\nu + \lambda_{\kappa t} I_\nu) + \ell_{\kappa t} \Delta_\lambda^\mu \nabla_\nu \mathbf{t}^{\nu\lambda} \\ &\quad + \ell_{\kappa n} \dot{n}^{\langle\mu\rangle} + \delta_{\kappa n} n^\mu \theta + (\lambda_{\kappa n} \sigma^{\mu\nu} + \tau_{\kappa n} \omega^{\mu\nu}) n_\nu \\ \tau_t \dot{\mathbf{t}}^{\langle\mu\nu\rangle} + \mathbf{t}^{\mu\nu} &= \mathfrak{d} \beta_0 \sigma^{\mu\nu} + \delta_{tt} \mathbf{t}^{\mu\nu} \theta + \lambda_{tt} \mathbf{t}_\lambda^{\langle\mu} \sigma^{\nu\rangle\lambda} + \frac{5}{3} \tau_t \mathbf{t}_\lambda^{\langle\mu} \omega^{\nu\rangle\lambda} + \ell_{t\kappa} \nabla^{\langle\mu} \kappa_0^{\nu\rangle} \\ &\quad + \lambda_{t\kappa} I^{\langle\mu} \kappa_0^{\nu\rangle} + \tau_{t\omega} \omega^{\langle\mu} \omega_0^{\nu\rangle} + \lambda_{t\omega} \sigma_\lambda^{\langle\mu} \epsilon^{\nu\rangle\lambda\alpha\beta} u_\alpha \omega_{0,\beta} \\ &\quad + \ell_{tn} \nabla^{\langle\mu} n^{\nu\rangle} + \tau_{tn} \dot{u}^{\langle\mu} n^{\nu\rangle} + \lambda_{tn} I^{\langle\mu} n^{\nu\rangle}\end{aligned}$$

# Dissipative spin hydrodynamics (simpler)

DW, 2409.07143 (2024)

$$\begin{aligned}\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu &= -\beta_0 \omega^\mu + \delta_{\omega\omega} \omega_0^\mu \theta + \lambda_{\omega\omega} \sigma^{\mu\nu} \omega_{0,\nu} + \lambda_{\omega t} \mathbf{t}^{\mu\nu} \omega_\nu \\ &\quad + \epsilon^{\mu\nu\alpha\beta} u_\nu (\ell_{\omega\kappa} \nabla_\alpha \kappa_{0,\beta} - \tau_\omega \dot{u}_\alpha \kappa_{0,\beta} + \lambda_{\omega\kappa} I_\alpha \kappa_{0,\beta}) \\ \tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu &= -\beta_0 \dot{u}^\mu + \mathfrak{b} I^\mu + \delta_{\kappa\kappa} \kappa_0^\mu \theta + \left( \lambda_{\kappa\kappa} \sigma^{\mu\nu} + \frac{\tau_\kappa}{2} \omega^{\mu\nu} \right) \kappa_{0,\nu} \\ &\quad + \epsilon^{\mu\nu\alpha\beta} u_\nu \left( \frac{\tau_\kappa}{2} \nabla_\alpha \omega_{0,\beta} + \tau_\kappa \dot{u}_\alpha \omega_{0,\beta} + \lambda_{\kappa\omega} I_\alpha \omega_{0,\beta} \right) \\ &\quad + \mathbf{t}^{\mu\nu} (\tau_{\kappa t} \dot{u}_\nu + \lambda_{\kappa t} I_\nu) + \ell_{\kappa t} \Delta_\lambda^\mu \nabla_\nu \mathbf{t}^{\nu\lambda} \\ \tau_t \dot{\mathbf{t}}^{\langle\mu\nu\rangle} + \mathbf{t}^{\mu\nu} &= \mathfrak{d} \beta_0 \sigma^{\mu\nu} + \delta_{tt} \mathbf{t}^{\mu\nu} \theta + \lambda_{tt} \mathbf{t}_\lambda^{\langle\mu} \sigma^{\nu\rangle\lambda} + \frac{5}{3} \tau_t \mathbf{t}_\lambda^{\langle\mu} \omega^{\nu\rangle\lambda} + \ell_{t\kappa} \nabla^{\langle\mu} \kappa_0^{\nu\rangle} \\ &\quad + \lambda_{t\kappa} I^{\langle\mu} \kappa_0^{\nu\rangle} + \tau_{t\omega} \omega^{\langle\mu} \omega_0^{\nu\rangle} + \lambda_{t\omega} \sigma_\lambda^{\langle\mu} \epsilon^{\nu\rangle\lambda\alpha\beta} u_\alpha \omega_{0,\beta}\end{aligned}$$

---

$$I^\mu := \nabla^\mu \alpha_0$$

# Dissipative spin hydrodynamics (simpler)

DW, 2409.07143 (2024)

$$\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu$$

Without fluid gradients:

$$\tau_\omega \dot{\omega}_0^{\langle\mu\rangle} + \omega_0^\mu = \ell_{\omega\kappa} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \kappa_{0,\beta}$$

$$\tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu$$

$$\tau_\kappa \dot{\kappa}_0^{\langle\mu\rangle} + \kappa_0^\mu = \frac{\tau_\kappa}{2} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \omega_{0,\beta} + \ell_{\kappa t} \Delta_\lambda^\mu \nabla_\nu t^{\nu\lambda}$$

$$\tau_t \dot{t}^{\langle\mu\nu\rangle} + t^{\mu\nu} = \ell_{t\kappa} \nabla^{\langle\mu} \kappa_0^{\nu\rangle}$$

$$+ \epsilon^{\mu\nu\alpha\beta} u_\nu \left( \frac{1}{2} \nabla_\alpha \omega_{0,\beta} + \tau_\kappa u_\alpha \omega_{0,\beta} + \lambda_{\kappa\omega} I_\alpha \omega_{0,\beta} \right)$$

$$+ t^{\mu\nu} (\tau_{\kappa t} \dot{u}_\nu + \lambda_{\kappa t} I_\nu) + \ell_{\kappa t} \Delta_\lambda^\mu \nabla_\nu t^{\nu\lambda}$$

$$\tau_t \dot{t}^{\langle\mu\nu\rangle} + t^{\mu\nu} = \delta\beta_0 \sigma^{\mu\nu} + \delta_{tt} t^{\mu\nu} \theta + \lambda_{tt} t_\lambda^{\langle\mu} \sigma^{\nu\rangle\lambda} + \frac{5}{3} \tau_t t_\lambda^{\langle\mu} \omega^{\nu\rangle\lambda} + \ell_{t\kappa} \nabla^{\langle\mu} \kappa_0^{\nu\rangle}$$

$$+ \lambda_{t\kappa} I^{\langle\mu} \kappa_0^{\nu\rangle} + \tau_{t\omega} \omega^{\langle\mu} \omega_0^{\nu\rangle} + \lambda_{t\omega} \sigma_\lambda^{\langle\mu} \epsilon^{\nu\rangle\lambda\alpha\beta} u_\alpha \omega_{0,\beta}$$

---

$$I^\mu := \nabla^\mu \alpha_0$$

# Dissipative spin hydrodynamics (simpler)

DW, 2409.07143 (2024)

$$\tau_\omega \dot{\omega}_0^{(\mu)} + \omega_0^\mu$$

Without fluid gradients:

$$\tau_\omega \dot{\omega}_0^{(\mu)} + \omega_0^\mu = \ell_{\omega\kappa} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \kappa_{0,\beta}$$

$$\tau_\kappa \dot{\kappa}_0^{(\mu)} + \kappa_0^\mu$$

$$\tau_\kappa \dot{\kappa}_0^{(\mu)} + \kappa_0^\mu = \frac{\tau_\kappa}{2} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \omega_{0,\beta} + \ell_{\kappa t} \Delta_\lambda^\mu \nabla_\nu t^{\nu\lambda}$$

$$\tau_t \dot{t}^{(\mu\nu)} + t^{\mu\nu} = \ell_{t\kappa} \nabla^{(\mu} \kappa_0^{\nu)}$$

$$+ \epsilon^{\mu\nu\alpha\beta} u_\nu \left( \frac{1}{2} \nabla_\alpha \omega_{0,\beta} + \tau_\kappa u_\alpha \omega_{0,\beta} + \lambda_{\kappa\omega} I_\alpha \omega_{0,\beta} \right)$$

Let  $\tau_t \rightarrow 0$  to get Bloch–Torrey-type contributions:

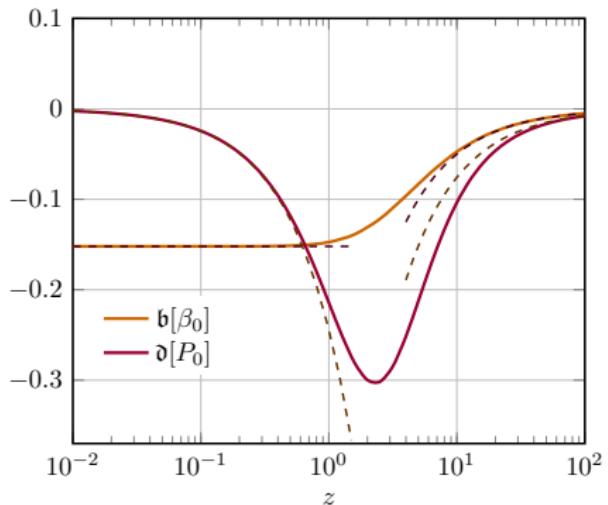
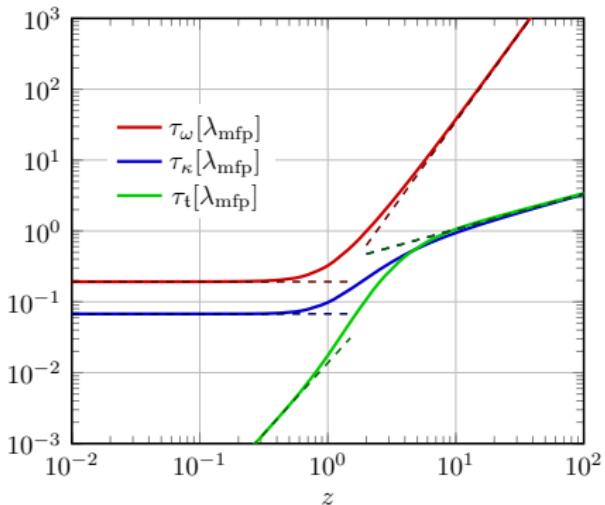
$$\tau_\omega \dot{\omega}_0^{(\mu)} + \omega_0^\mu = \ell_{\omega\kappa} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \kappa_{0,\beta}$$

$$\tau_\kappa \dot{\kappa}_0^{(\mu)} + \kappa_0^\mu = \frac{\tau_\kappa}{2} \epsilon^{\mu\nu\alpha\beta} u_\nu \nabla_\alpha \omega_{0,\beta} + \frac{\ell_{\kappa t} \ell_{t\kappa}}{2} \Delta_\lambda^\mu \Delta \kappa_0^\lambda + \frac{\ell_{\kappa t} \ell_{t\kappa}}{6} \nabla^\mu \nabla_\lambda \kappa_0^\lambda$$

$$I^\mu := \nabla^\mu \alpha_0$$

# Relaxation times and first-order coefficients

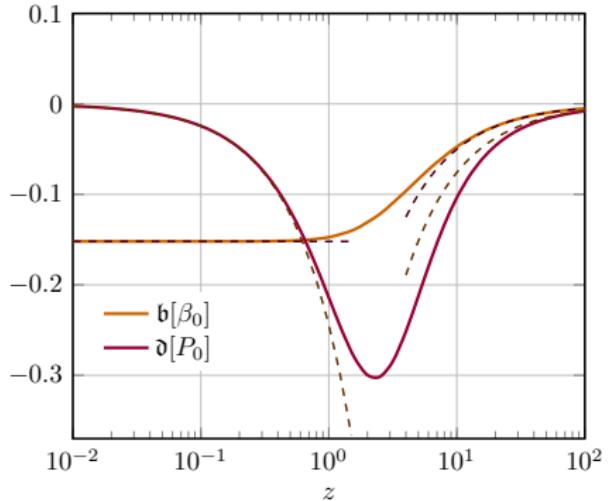
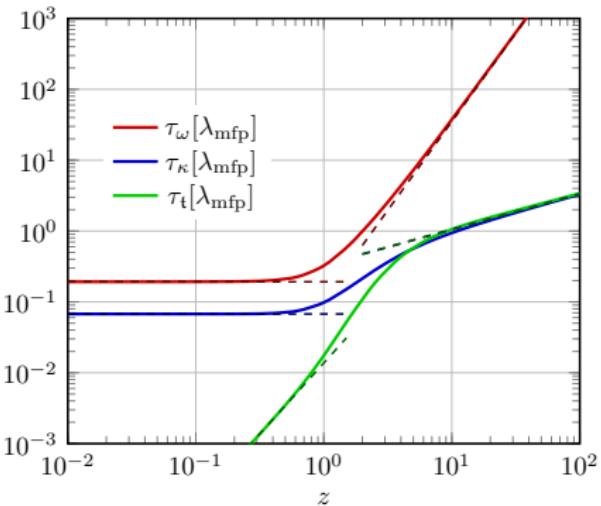
DW, 2409.07143 (2024)



- $\tau_\omega$  grows with  $z^2$  compared to  $\tau_\kappa$  and  $\tau_t$
- $\tau_t$  vanishes for  $z \rightarrow 0$

# Relaxation times and first-order coefficients

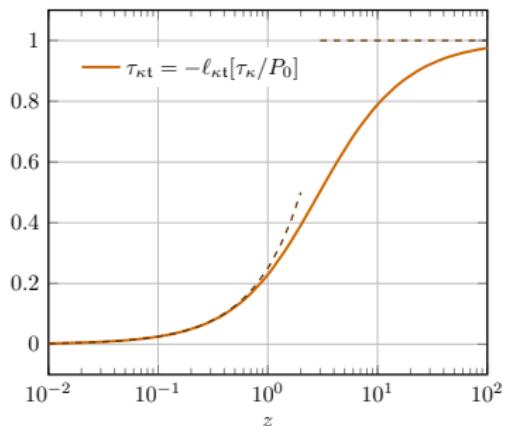
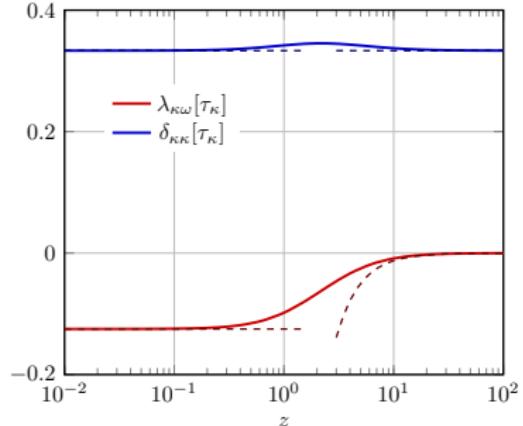
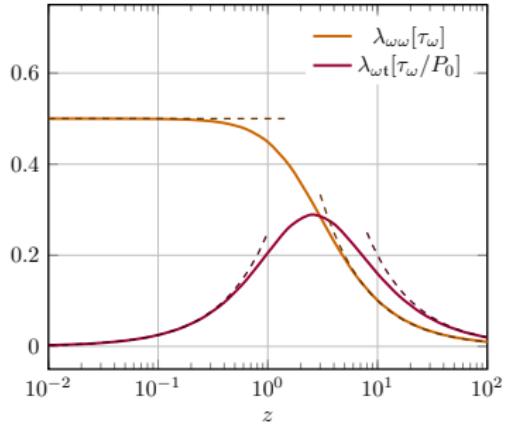
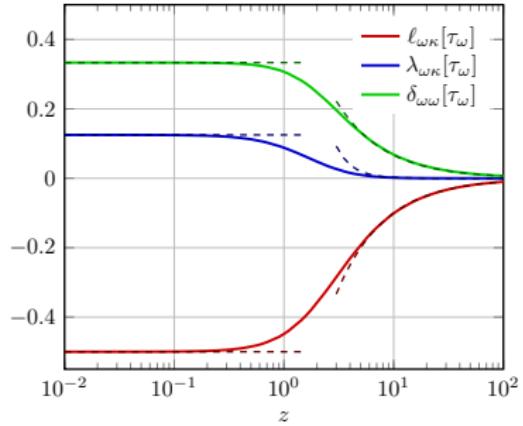
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	$\tau_\omega[\lambda_{\text{mfp}}]$	$\tau_\kappa[\lambda_{\text{mfp}}]$	$\tau_t[\lambda_{\text{mfp}}]$	$b[\beta_0]$	$d[P_0]$
$z \rightarrow 0$	$5/26$	$5/74$	$4z^2/291$	$-45/296$	$-592z/2425$
$z \rightarrow \infty$	$\sqrt{\pi}z^{5/2}/16$	$3\sqrt{\pi}\sqrt{z}/16$	$60\sqrt{\pi}\sqrt{z}/317$	$-1/(2z)$	$-240/(317z)$

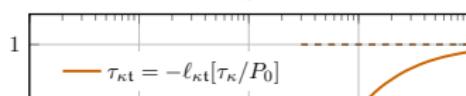
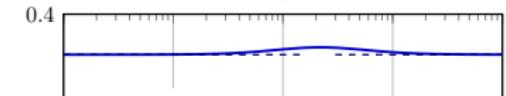
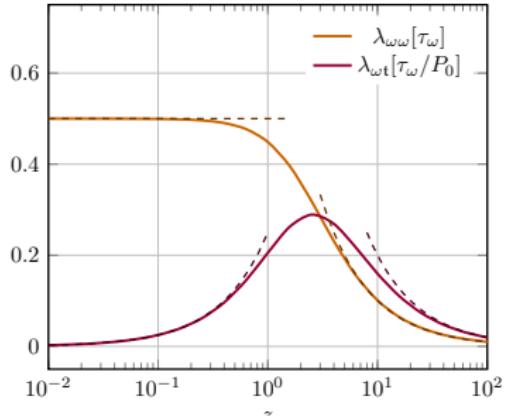
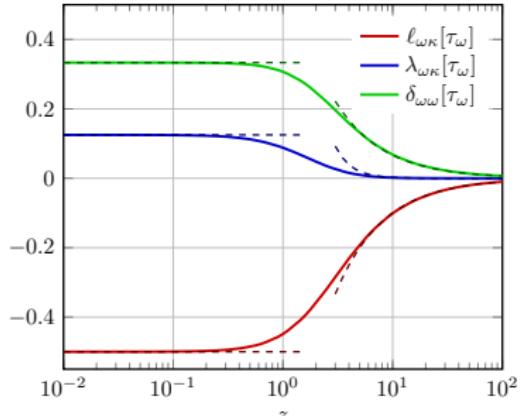
# Second-order coefficients for $\omega_0^\mu$ and $\kappa_0^\mu$

DW, 2409.07143 (2024)

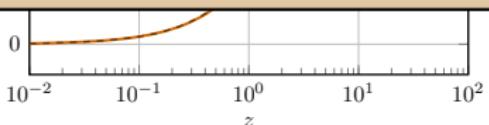
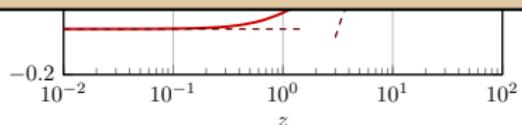


# Second-order coefficients for $\omega_0^\mu$ and $\kappa_0^\mu$

DW, 2409.07143 (2024)

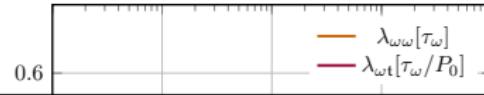
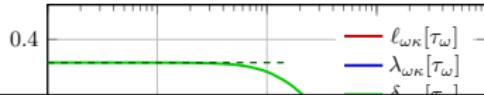


	$\ell_{\omega\kappa}[\tau_\omega]$	$\lambda_{\omega\kappa}[\tau_\omega]$	$\delta_{\omega\omega}[\tau_\omega]$	$\lambda_{\omega\omega}[\tau_\omega]$	$\lambda_{\omega t}[\tau_\omega/P_0]$
$z \rightarrow 0$	$-1/2$	$1/8$	$1/3$	$1/2$	$z/4$
$z \rightarrow \infty$	$-1/z$	$5/(2z^3)$	$2/(3z)$	$1/z$	$2/z$

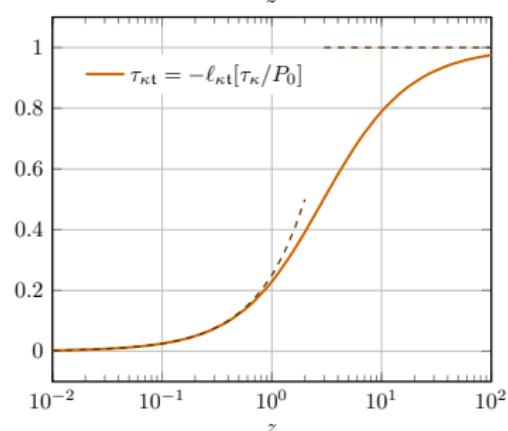
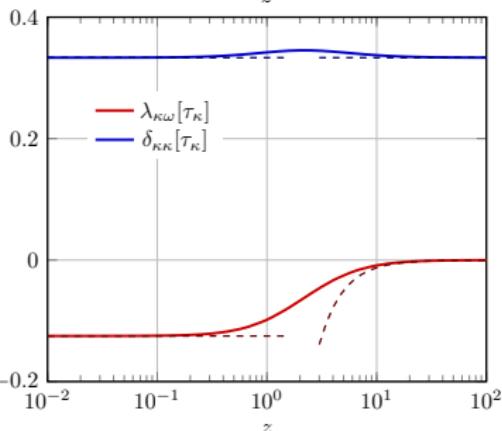
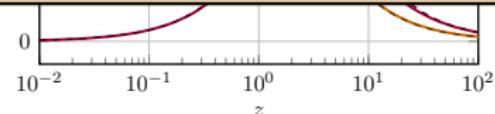
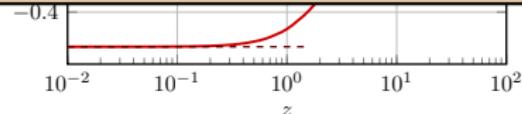


# Second-order coefficients for $\omega_0^\mu$ and $\kappa_0^\mu$

DW, 2409.07143 (2024)

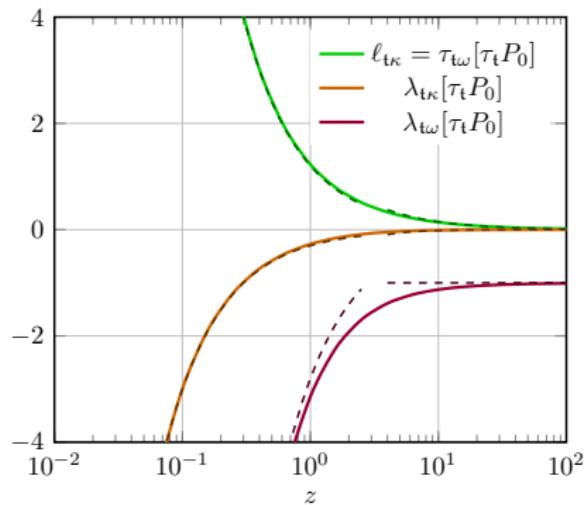
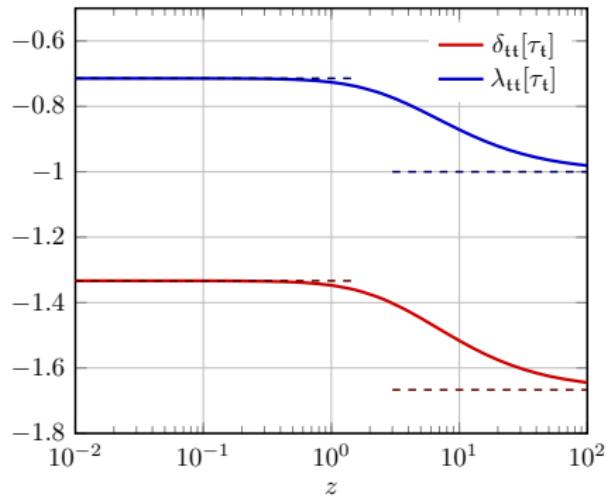


	$\lambda_{\kappa\omega}[\tau_\kappa]$	$\delta_{\kappa\kappa}[\tau_\kappa]$	$\lambda_{\kappa\kappa}[\tau_\kappa]$	$\tau_{\kappa t}[\tau_\kappa/P_0]$	$\ell_{\kappa t}[\tau_\kappa/P_0]$
$z \rightarrow 0$	$-1/8$	$1/3$	$1/2$	$z/4$	$-z/4$
$z \rightarrow \infty$	$-5/(4z^2)$	$1/3$	$1/2$	$1$	$-1$



# Second-order coefficients for $t^{\mu\nu}$

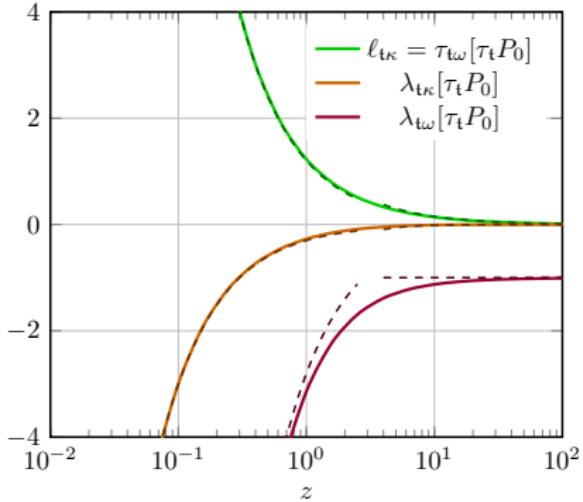
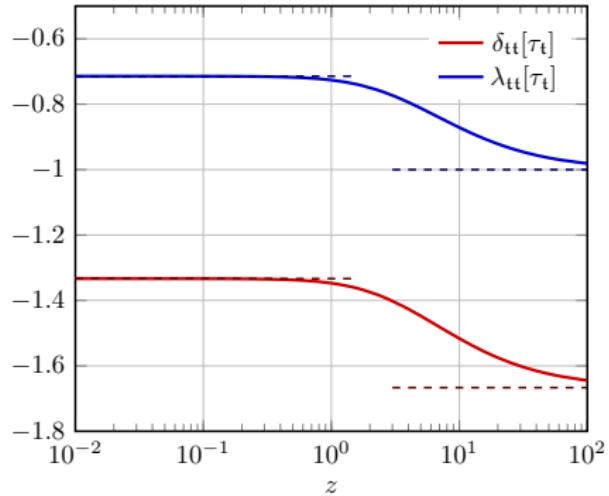
DW, 2409.07143 (2024)



- Some coefficients diverge for  $z \rightarrow 0$  in units of  $\tau_t$ , but since  $\tau_t \sim z^2$  they actually vanish
- Equation becomes trivial in the ultrarelativistic case,  $t^{\mu\nu} = 0$ 
  - ▶ Spin dynamics solely described by  $\omega_0^\mu$  and  $\kappa_0^\mu$

# Second-order coefficients for $t^{\mu\nu}$

DW, 2409.07143 (2024)



	$\delta_{tt}[\tau_t]$	$\lambda_{tt}[\tau_t]$	$\ell_{tk}[\tau_t P_0]$	$\lambda_{tk}[\tau_t P_0]$	$\tau_{tw}[\tau_t P_0]$	$\lambda_{tw}[\tau_t P_0]$
$z \rightarrow 0$	$-4/3$	$-5/7$	$6/(5z)$	$-3/(10z)$	$6/(5z)$	$-14/(5z)$
$z \rightarrow \infty$	$-5/3$	$-1$	$3/(2z)$	$-3/(2z^2)$	$3/(2z)$	$-1$

# Summary and outlook

- Quantum kinetic theory is a versatile effective microscopic theory
  - ▶ At local equilibrium: provides a basis for ideal (spin) hydrodynamics
  - ▶ Near local equilibrium: provides a basis for dissipative (spin) hydrodynamics
  - ▶ Away from equilibrium: can be studied as a full-fledged nonequilibrium transport theory
- Future applications & developments
  - ▶ Study of spin hydro from QKT in various setups, in particular polarization dynamics
    - Allows to provide numerical answers to points (I) and (II) in the context of heavy-ion collisions
  - ▶ Numerical implementation of kinetic equations to first order in  $\hbar$

## Appendix

# Conserved currents in QKT

## Conserved currents

$$\frac{1}{2} T^{(\mu\nu)} = \int d\Gamma k^\mu k^\nu f ,$$

$$S^{\lambda\mu\nu} = \frac{1}{2m} \int d\Gamma k^\lambda \epsilon^{\mu\nu\alpha\beta} k_\alpha s_\beta f .$$

$$T^{[\mu\nu]} = \frac{1}{2} \int [d\Gamma] \widetilde{\mathcal{W}} \Delta^{[\mu} k^{\nu]} (f_1 f_2 - f f')$$

## Conservation laws

$$\int d\Gamma k^\mu C[f] = 0$$

$$\frac{\hbar}{2m} \int d\Gamma \epsilon^{\mu\nu\alpha\beta} k_\alpha s_\beta C[f] = \frac{\hbar}{m} \int \frac{d^4 k}{(2\pi\hbar)^4} k^{[\mu} \mathcal{D}_{\nu]}^\nu$$

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$$[d\Gamma] := d\Gamma_1 d\Gamma_2 d\Gamma d\Gamma'$$

# Polarization observables in kinetic theory

## Vector Polarization (Pauli-Lubanski Pseudovector)

$$S^\mu(\mathbf{k}) := \text{Tr} \left[ \hat{S}^\mu \hat{\rho}(\mathbf{k}) \right] = \frac{1}{N(\mathbf{k})} \int d\Sigma_\lambda k^\lambda \int dS(\mathbf{k}) \mathfrak{s}^\mu f(\mathbf{x}, \mathbf{k}, \mathfrak{s})$$

## Tensor Polarization

$$\begin{aligned}\rho_{00}(\mathbf{k}) &= \frac{1}{3} - \sqrt{\frac{2}{3}} \epsilon_\mu^{(0)}(\mathbf{k}) \epsilon_\nu^{(0)}(\mathbf{k}) \Theta^{\mu\nu}(\mathbf{k}) \\ \Theta^{\mu\nu}(\mathbf{k}) &:= \frac{1}{2} \sqrt{\frac{3}{2}} \text{Tr} \left[ \left( \hat{S}^{(\mu} \hat{S}^{\nu)} + \frac{4}{3} K^{\mu\nu} \right) \hat{\rho}(\mathbf{k}) \right] \\ &= \frac{1}{2} \sqrt{\frac{3}{2}} \frac{1}{N(\mathbf{k})} \int d\Sigma_\lambda k^\lambda \int dS(\mathbf{k}) K_{\alpha\beta}^{\mu\nu} \mathfrak{s}^\alpha \mathfrak{s}^\beta f(\mathbf{x}, \mathbf{k}, \mathfrak{s})\end{aligned}$$

$$N(\mathbf{k}) := \int d\Sigma_\gamma k^\gamma \int dS(\mathbf{k}) f(\mathbf{x}, \mathbf{k}, \mathfrak{s}), \quad \hat{S}^\mu := -(1/2m) \epsilon^{\mu\nu\alpha\beta} \hat{J}_{\nu\alpha} \hat{P}_\beta$$

# Polarization in spin hydrodynamics

## Local Polarization

$$\begin{aligned} S_0^\mu &= \frac{2\sigma^2 \hbar}{N(k)m} \int d\Sigma_\lambda k^\lambda \left( u^\mu \omega_0^\nu k_\nu - E_{\mathbf{k}} \omega_0^\mu + \epsilon^{\mu\nu\alpha\beta} u_\nu k_\alpha \kappa_{0,\beta} \right) f_0 \tilde{f}_0 \\ \delta S^\mu &= -\frac{2\sigma}{N(k)} \int d\Sigma_\lambda k^\lambda K^{\mu\gamma} \Xi_{\gamma\alpha} f_0 \tilde{f}_0 \\ &\quad \times \left( \mathfrak{x}_n \epsilon^{\alpha\beta\rho\sigma} u_\beta k_\rho n_\sigma + \mathfrak{x}_{\mathfrak{t}} \mathfrak{t}_\rho^{\langle\beta} \epsilon^{\gamma\rangle\alpha\sigma\rho} u_\sigma k_{\beta} k_{\gamma\rangle} \right) \end{aligned}$$

## Global Polarization

$$\begin{aligned} \overline{S}_0^\mu &= -\frac{2\sigma^2 \hbar}{\overline{N}m} \int d\Sigma_\lambda \left( J_{21} u^\mu \omega_0^\lambda + J_{20} \omega_0^\mu u^\lambda + J_{21} \epsilon^{\mu\nu\lambda\beta} u_\nu \kappa_{0,\beta} \right) \\ \delta \overline{S}^\mu &= \frac{\sigma}{\overline{N}} \frac{1}{2} \int d\Sigma_\lambda B_0 \epsilon^{\mu\lambda\alpha\beta} u_\alpha n_\beta \end{aligned}$$

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$$\mathfrak{x}_n := \frac{1}{2} \sum_n \mathcal{H}_{\mathbf{k}n}^{(1,1)} \frac{\mathfrak{b}_n^{(1)}}{\varkappa} , \quad \mathfrak{x}_{\mathfrak{t}} := \frac{2}{3} \sum_n \mathcal{H}_{\mathbf{k}n}^{(1,2)} \frac{\mathfrak{d}_n}{\mathfrak{d}_0}$$

# Nonlocal collisions

DW, NW, ES, 2306.05936 (2023)

## Spacetime shifts

$$\Delta^\mu := -\frac{i\hbar}{8m} \frac{m^4}{\mathcal{W}} M^{\gamma_1 \gamma_2 \delta_1 \delta_2} M^{\zeta_1 \zeta_2 \eta_1 \eta_2} h_{1,\gamma_1 \eta_1} h_{2,\gamma_2 \eta_2} h'_{\zeta_2 \delta_2} [h, \gamma^\mu]_{\zeta_1 \delta_1}$$

- Depend on the transfer-matrix elements

$$\langle 11' | \hat{t} | 22' \rangle = \bar{u}_{1,\alpha} \bar{u}_{1',\beta} u_{2,\gamma} u_{2',\delta} M^{\alpha \beta \gamma \delta}$$

- Manifestly covariant
  - no “no-jump” frame

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$$h := \frac{1}{4} (\mathbb{1} + \gamma_5 \not{s})(\not{k} + m)$$

# Moment equations: Spin-rank 1

- Same procedure as for the moments of spin-rank 0

Moment equation for  $\ell = 2$

$$\dot{\tau}_r^{\langle\mu\rangle,\nu\lambda} - \mathfrak{C}_{r-1}^{\langle\mu\rangle,\nu\lambda} = \dots$$

- Navier-Stokes limit:  $\mathfrak{C}_{r-1}^{\langle\mu\rangle,\nu\lambda} = 0$
- Contains local and nonlocal contributions
  - ▶  $\mathfrak{C}_{r,\text{local}}^{\langle\mu\rangle,\nu\lambda} \sim \tau_r^{\mu,\nu\lambda}$
  - ▶  $\mathfrak{C}_{r,\text{nonlocal}}^{\langle\mu\rangle,\nu\lambda} \sim \sigma_\rho^{\langle\nu} \epsilon^{\lambda\rangle\mu\alpha\rho} u_\alpha$
- Leads to shear-induced polarization, coefficient independent of total cross-section
- Magnitude not yet clear

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# Moment equations: Spin-rank 2

## Moment equation for $\ell = 0$

$$\begin{aligned}\dot{\psi}_r^{\langle\mu\nu\rangle} - \mathfrak{C}_{r-1}^{\langle\mu\nu\rangle} &= -\frac{\theta}{3} \left[ (r+2)\psi_r^{\langle\mu\nu\rangle} - (r-1)m^2\psi_{r-2}^{\langle\mu\nu\rangle} \right] \theta + r\psi_{r-1}^{\langle\mu\nu\rangle,\alpha} \dot{u}_\alpha \\ &\quad - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\gamma \psi_{r-1}^{\alpha\beta,\gamma} + (r-1)\psi_{r-2}^{\langle\mu\nu\rangle,\alpha\beta} \sigma_{\alpha\beta}\end{aligned}$$

- No dependence on **equilibrium** quantities appears because moments of spin-rank 2 do not appear in any conserved current
- Nonetheless, they are important to e.g. describe **tensor polarization** of spin-1 particles

DW, NW, ES, 2207.01111 (2022)

- ▶ Main argument:  $\mathfrak{C}_{r-1}^{\langle\mu\nu\rangle} \simeq A_r \psi_r^{\mu\nu} + B_r \pi^{\mu\nu}$  leads to  $\psi_r^{\mu\nu} \sim \pi^{\mu\nu}$  in the Navier-Stokes limit

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$$\Delta_{\alpha\beta}^{\mu\nu} := (\Delta_\alpha^{(\mu} \Delta_\beta^{\nu)})/2 - (1/3)\Delta^{\mu\nu}\Delta_{\alpha\beta}$$

# Alignment of $\phi$ -mesons

- Spin-1 particles feature tensor polarization ( $\hat{\wedge}$  alignment)
- Intuition: “Polarization counts the difference between spin-projections +1 and -1, alignment counts spin-projection 0”

- ▶ Larger than expected
- ▶ Many theoretical developments

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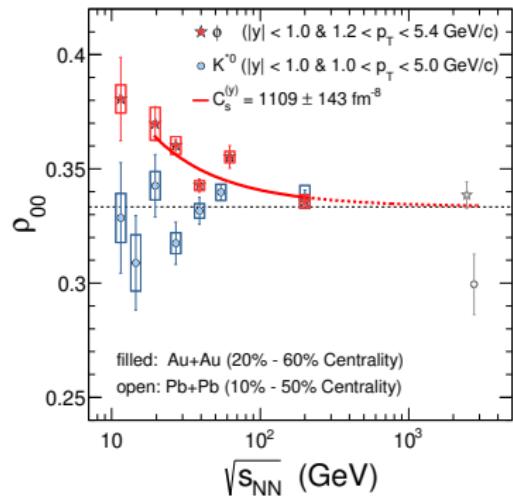
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# Alignment in QKT

- Spin-1 Wigner function has 16 components, 9 of which are independent

$$W^{\mu\nu} = E^{\mu\nu} f_E + \frac{k^{(\mu}}{2k} F_S^{\nu)} + F_K^{\mu\nu} + K^{\mu\nu} f_K + i \frac{k^{[\mu}}{2k} F_A^{\nu]} + i \epsilon^{\mu\nu\alpha\beta} \frac{k_\alpha}{m} G_\beta$$

- ▶  $f_K$  → particle-number density
- ▶  $G^\mu$  → polarization
- ▶  $F_K^{\mu\nu}$  → tensor polarization/ alignment
- Several approaches on the market (incomplete list)
  - ▶ Coalescence  
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  - ▶ Strong force fields  
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