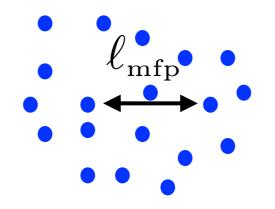
Towards a holographic description of hydro with spin

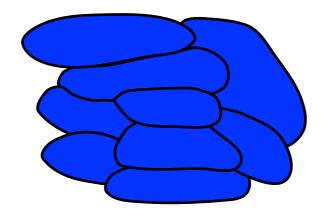
A.Yarom

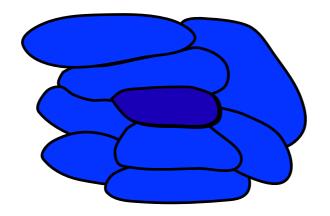
Together with C. Cartwright, A. D. Gallegos, U. Gursoy and R. Klein

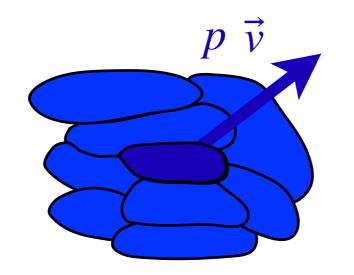
 $\stackrel{\ell_{\rm mfp}}{\longleftrightarrow}$



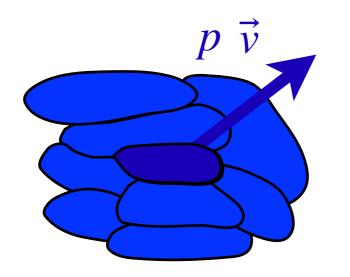




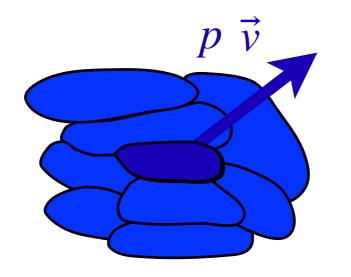




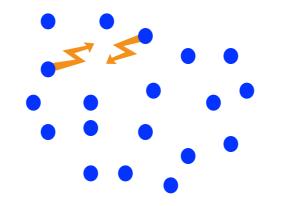
 $\dot{\vec{v}} + \vec{v} \cdot \vec{\nabla} \vec{v} = -\vec{\nabla} p + \eta \nabla^2 \vec{v}$ $\vec{\nabla} \cdot \vec{v} = 0$



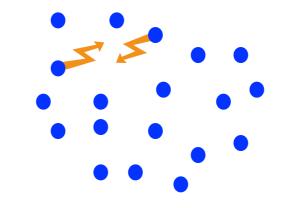
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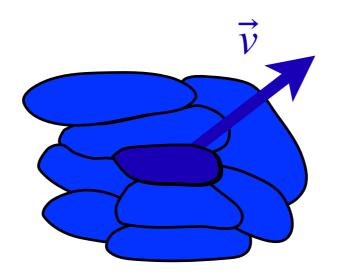


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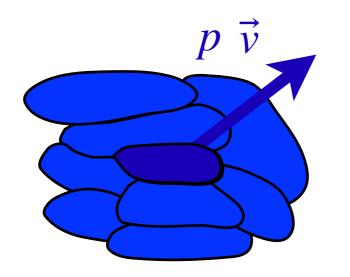


	Viscosity (mPa s)
Water	1
Whole milk	2
Olive oil	56
Pitch	1011

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It works!

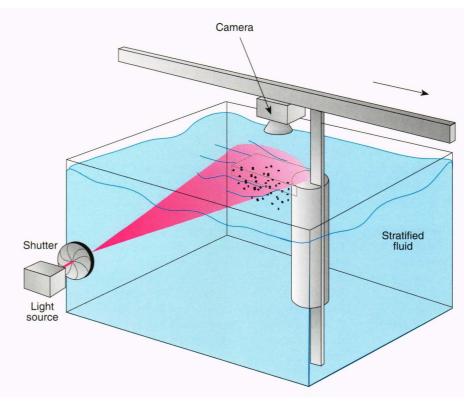


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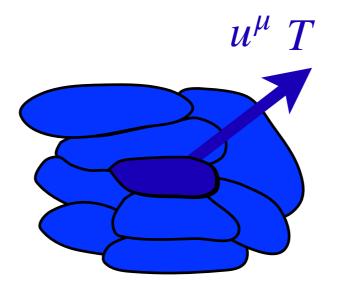
Can be verified!



(S. Diamond, 1994)

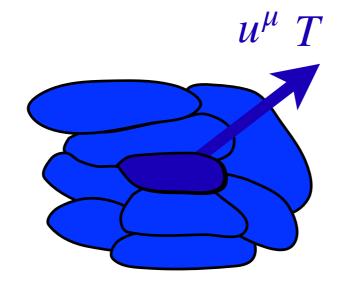
$$\nabla_{\mu}T^{\mu\nu}=0$$

$$T^{\mu\nu} = \epsilon(T)u^{\mu}u^{\nu} + P(T)(\eta^{\mu\nu} + u^{\mu}u^{\nu}) - 2\eta(T)\sigma_{\mu\nu} + \dots$$



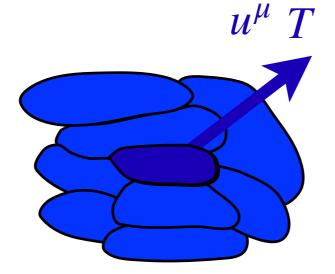
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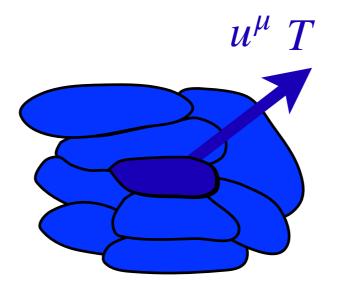
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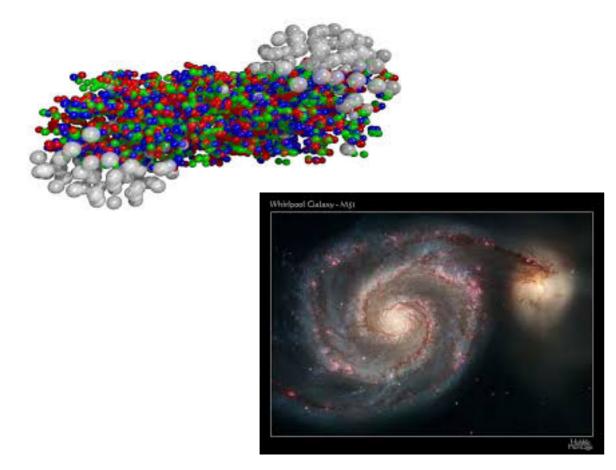


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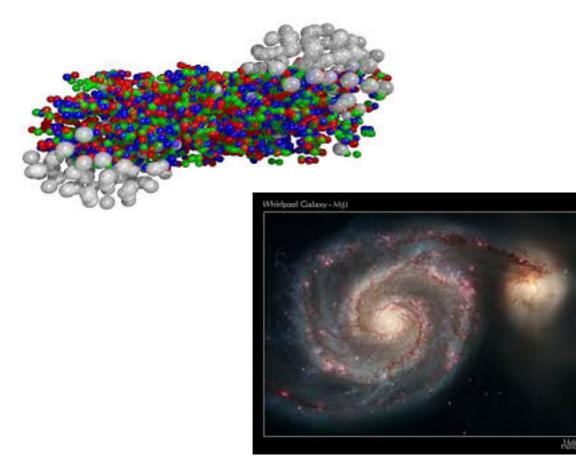
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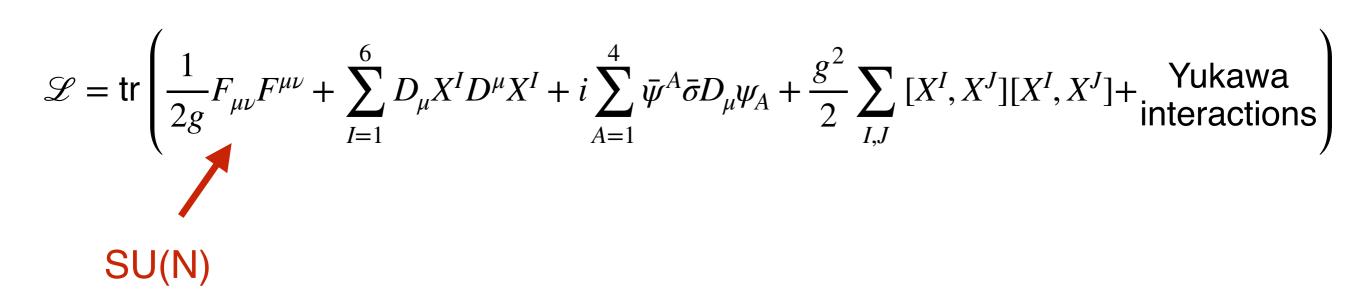
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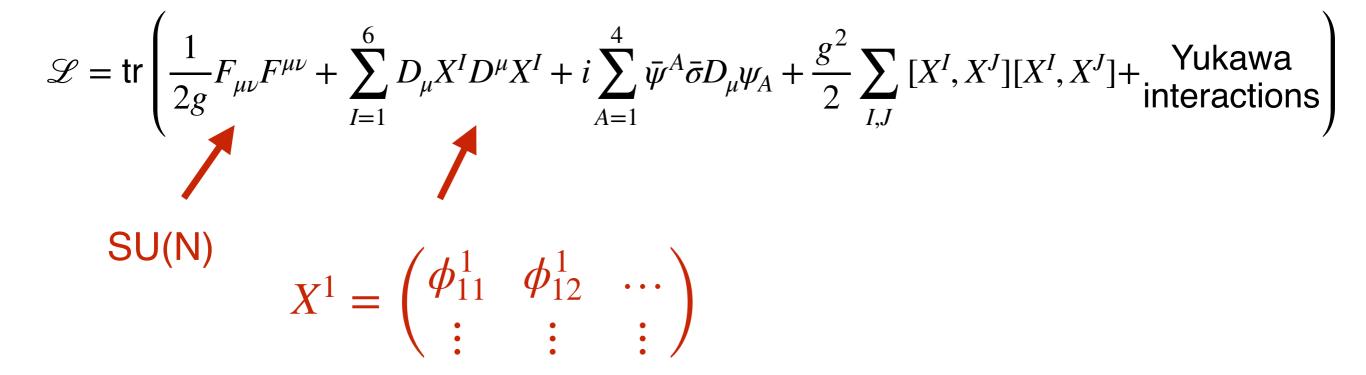
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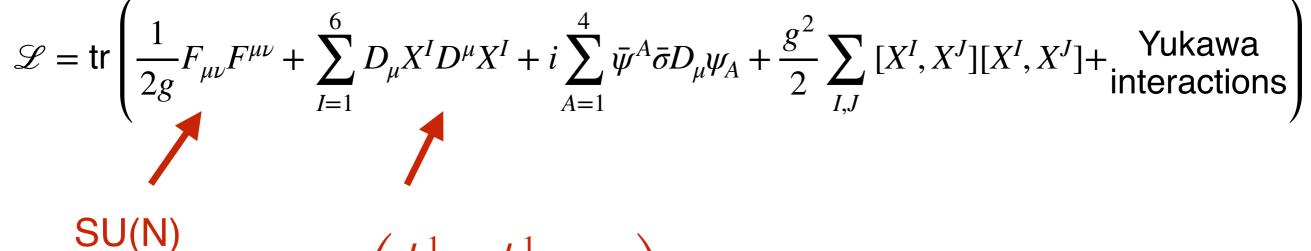
Can be verified.



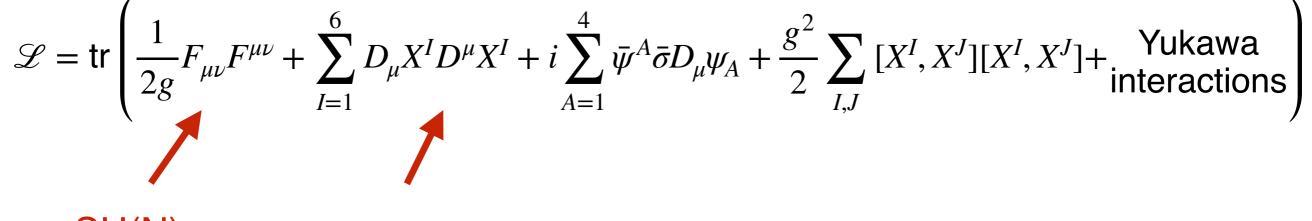
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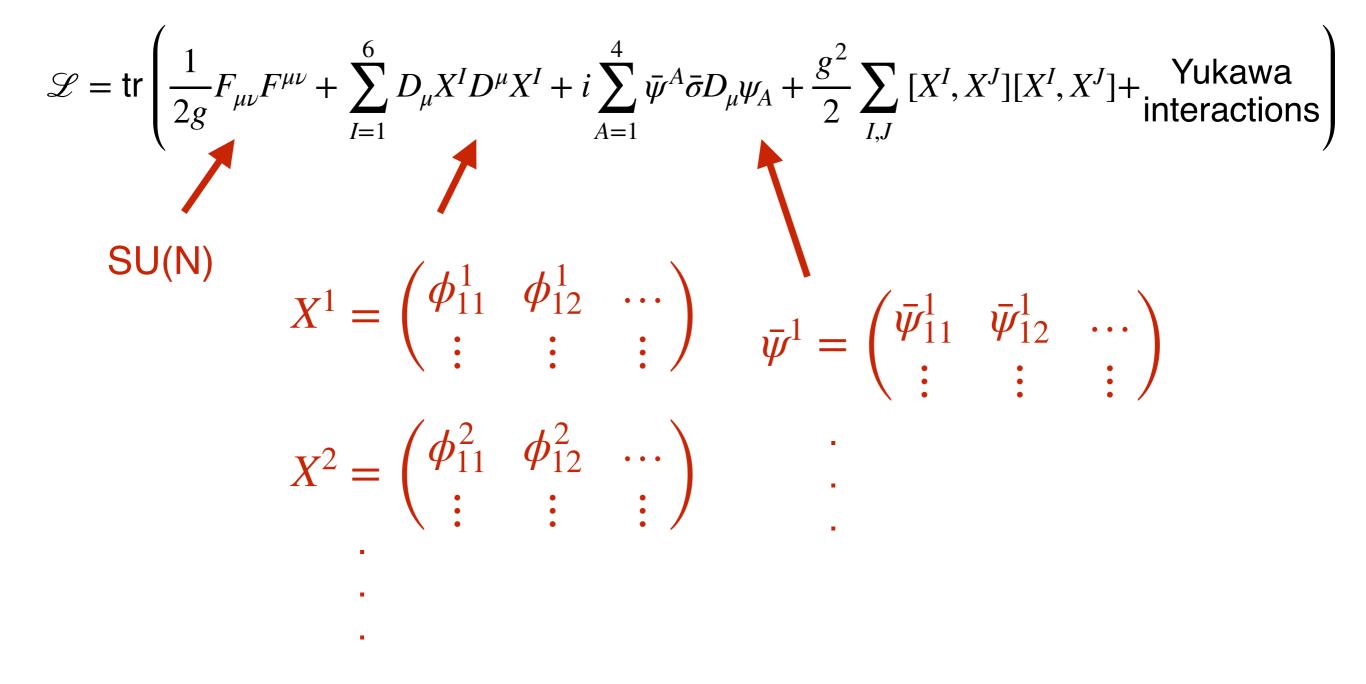


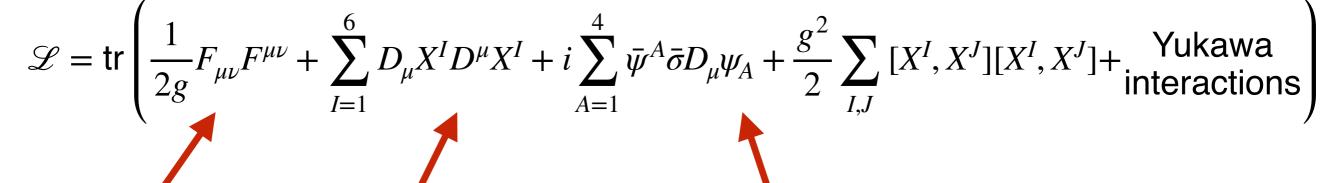
$$X^{1} = \begin{pmatrix} \phi_{11}^{1} & \phi_{12}^{1} & \cdots \\ \vdots & \vdots & \vdots \end{pmatrix}$$
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SU(N)

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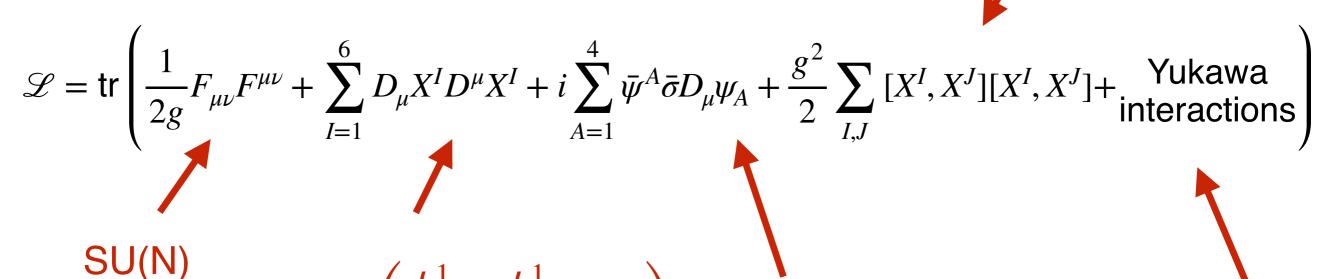




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•



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$$X^{2} = \begin{pmatrix} r & 11 & r & 12 \\ \vdots & \vdots & \vdots \end{pmatrix}$$

 ψ_{12}

I U 1 1

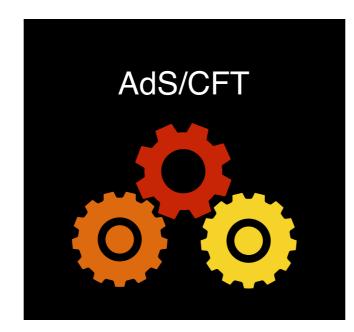
x 7

Holographic hydrodynamics
$$\mathscr{L} = \operatorname{tr} \left(\frac{1}{2g} F_{\mu\nu} F^{\mu\nu} + \sum_{I=1}^{6} D_{\mu} X^{I} D^{\mu} X^{I} + i \sum_{A=1}^{4} \bar{\psi}^{A} \bar{\sigma} D_{\mu} \psi_{A} + \frac{g^{2}}{2} \sum_{I,J} [X^{I}, X^{J}] [X^{I}, X^{J}] + \operatorname{Yukawa}_{\text{interactions}} \right)$$

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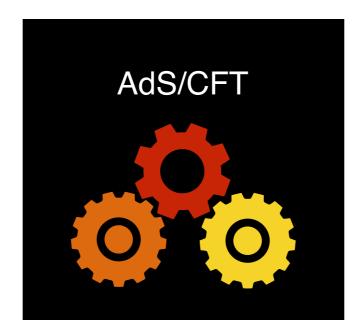


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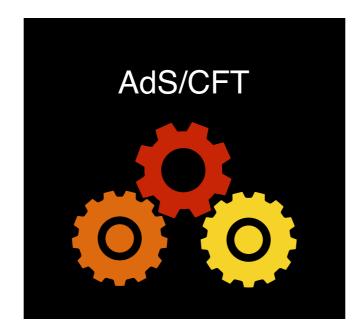
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$$\epsilon = \frac{3\pi^2 N^2}{8} T^4$$

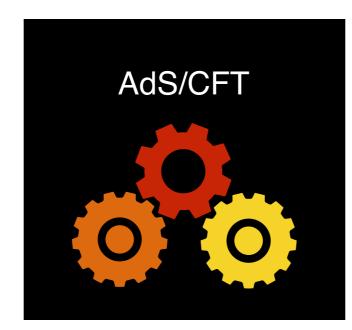
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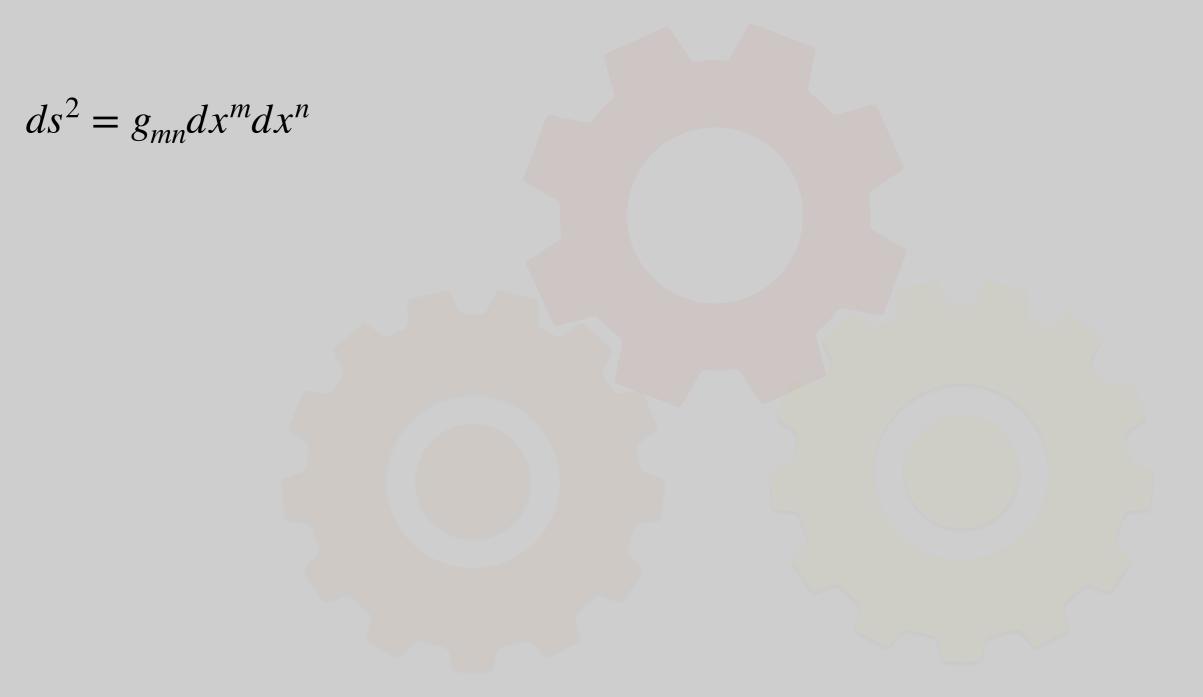


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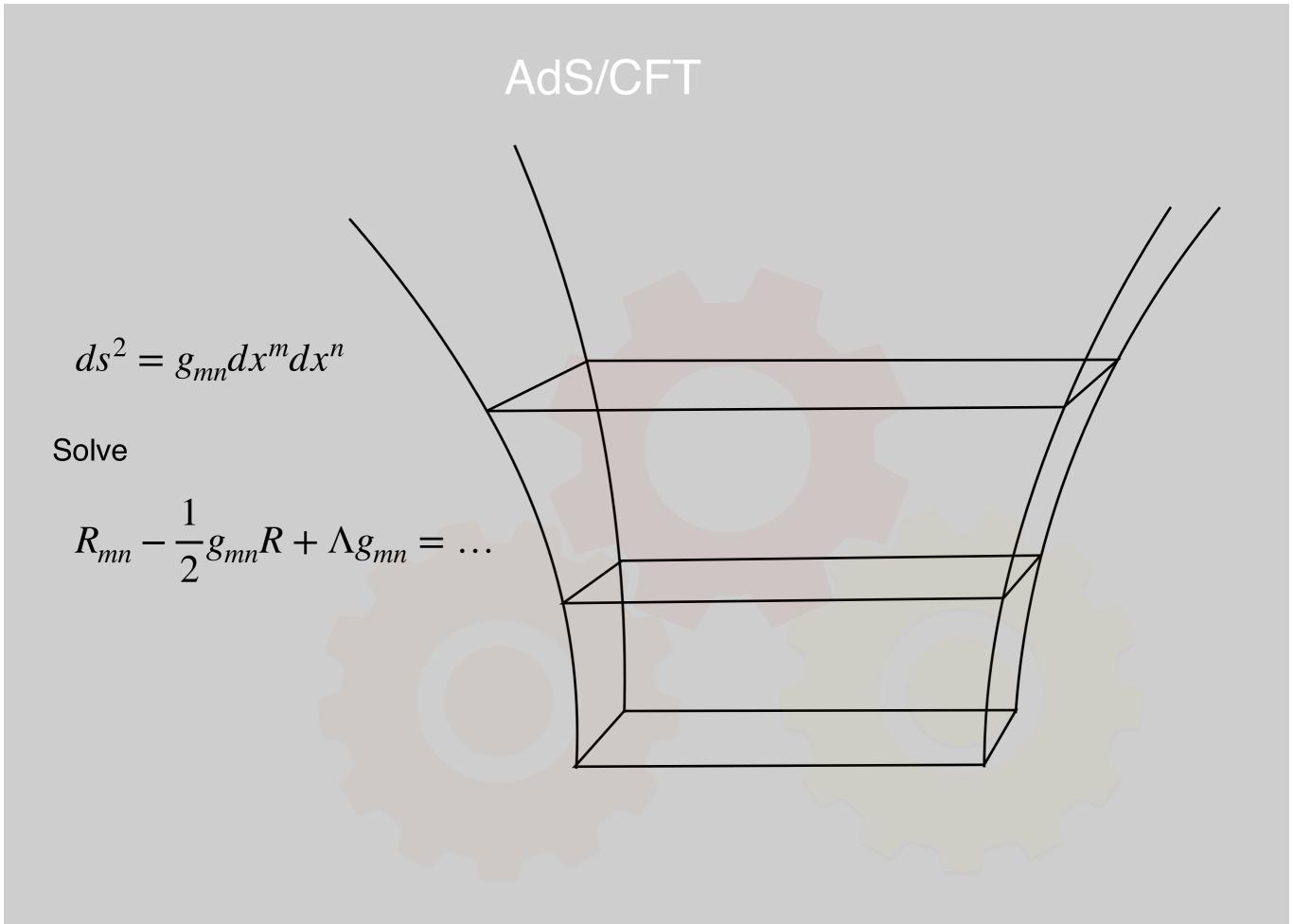


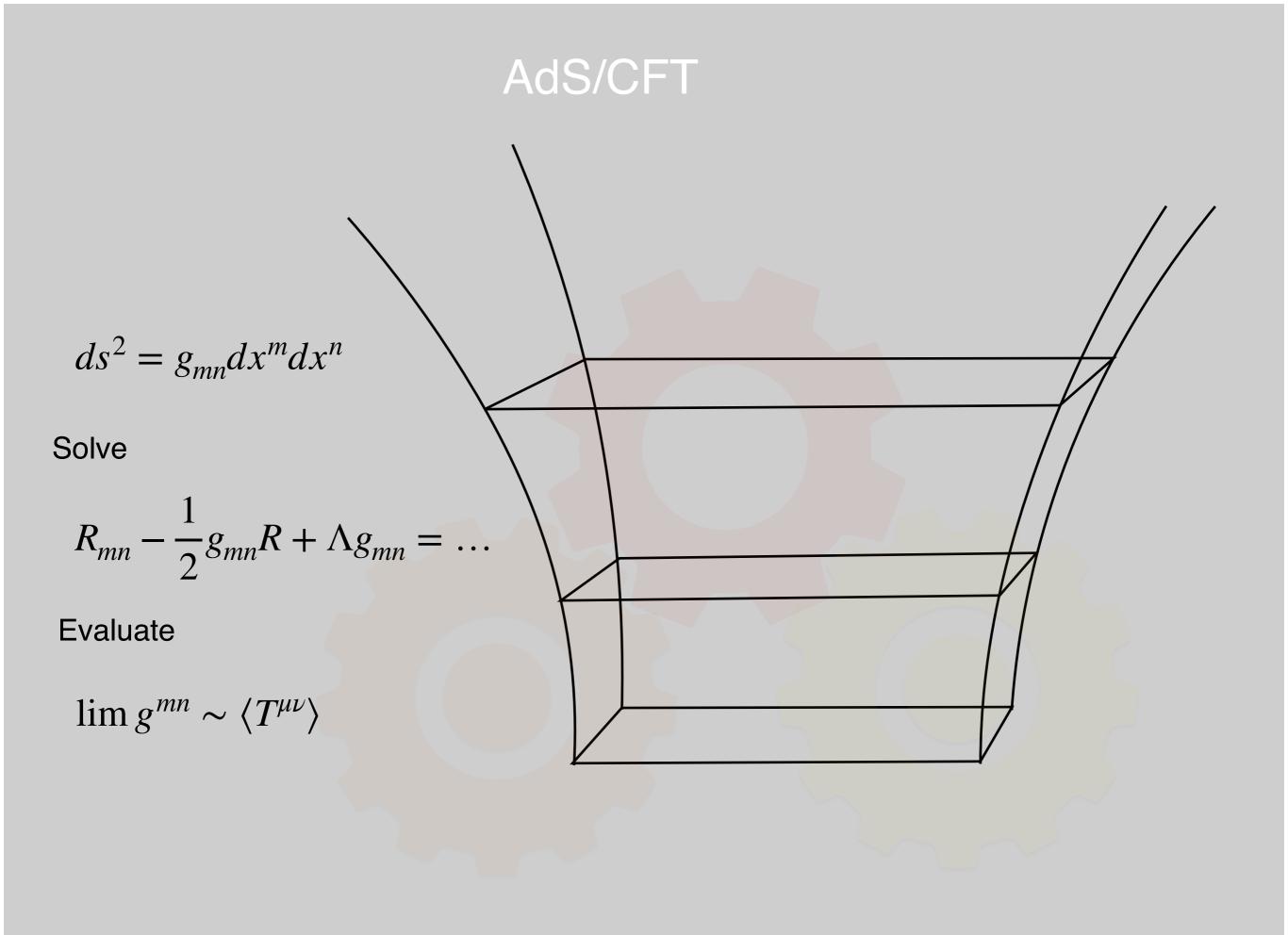


$$ds^2 = g_{mn}dx^m dx^n$$

Solve

$$R_{mn} - \frac{1}{2}g_{mn}R + \Lambda g_{mn} = \dots$$





Relativistic hydrodynamics

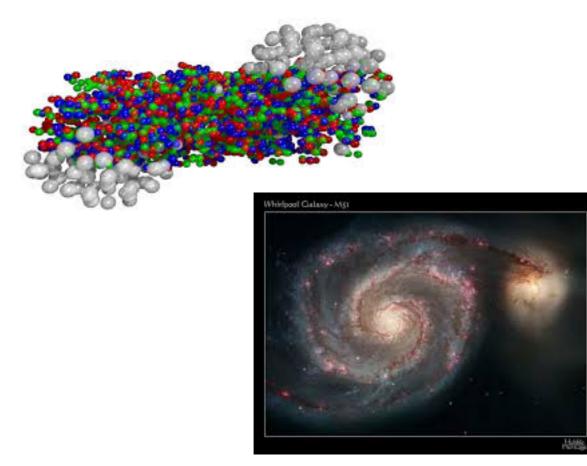
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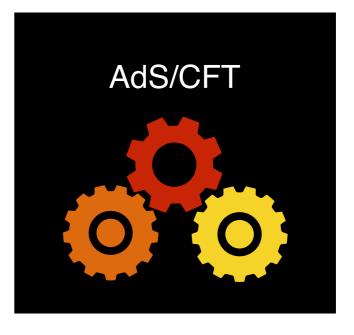
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It works.

Can be verified.





Relativistic hydrodynamics

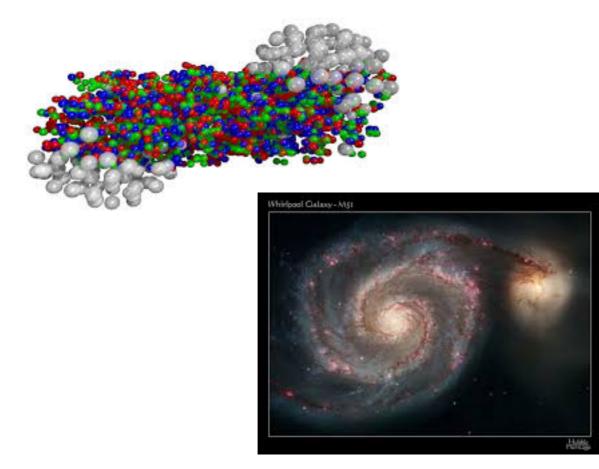
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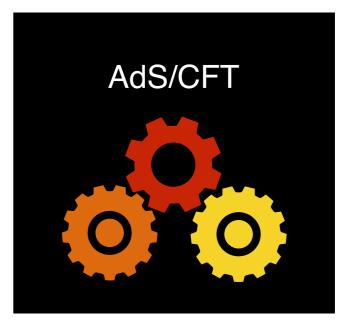
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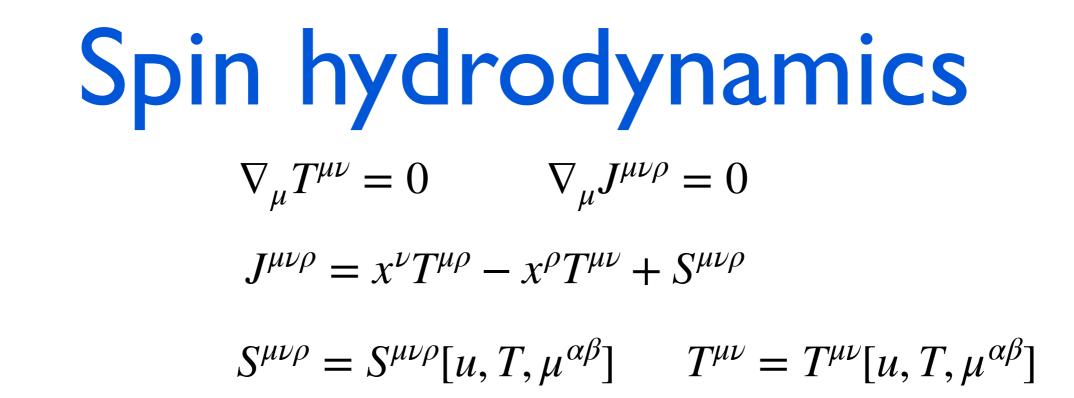
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Can be verified(?)

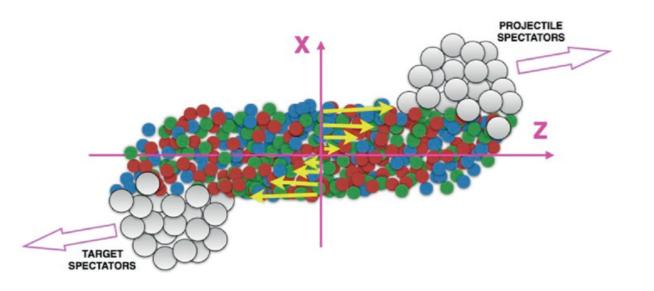


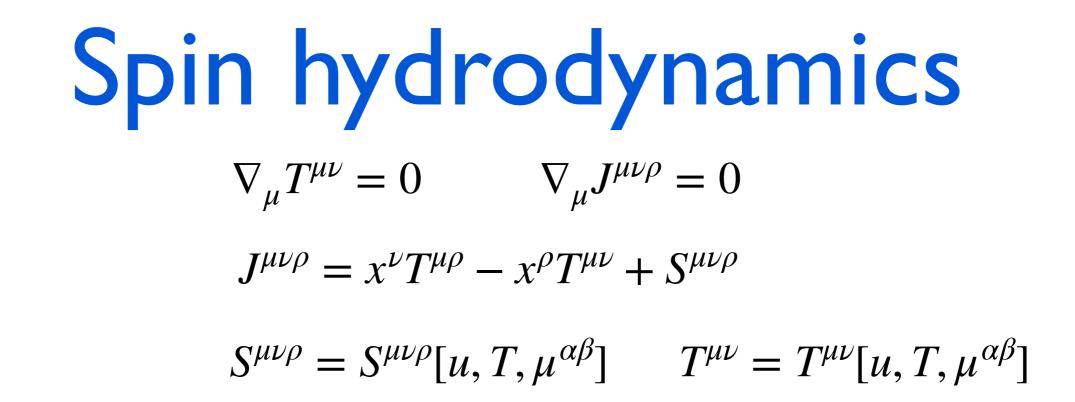


Spin hydrodynamics $\nabla_{\mu}T^{\mu\nu} = 0 \qquad \nabla_{\mu}J^{\mu\nu\rho} = 0$ $J^{\mu\nu\rho} = x^{\nu}T^{\mu\rho} - x^{\rho}T^{\mu\nu} + S^{\mu\nu\rho}$ $S^{\mu\nu\rho} = S^{\mu\nu\rho}[u, T, \mu^{\alpha\beta}] \qquad T^{\mu\nu} = T^{\mu\nu}[u, T, \mu^{\alpha\beta}]$ $u^{\mu} T \mu^{ab}$



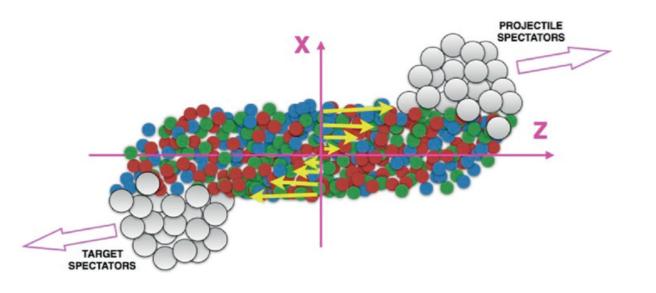






It works?

Can be verified?

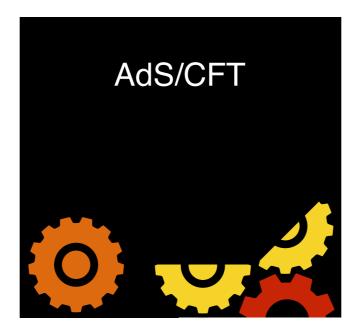


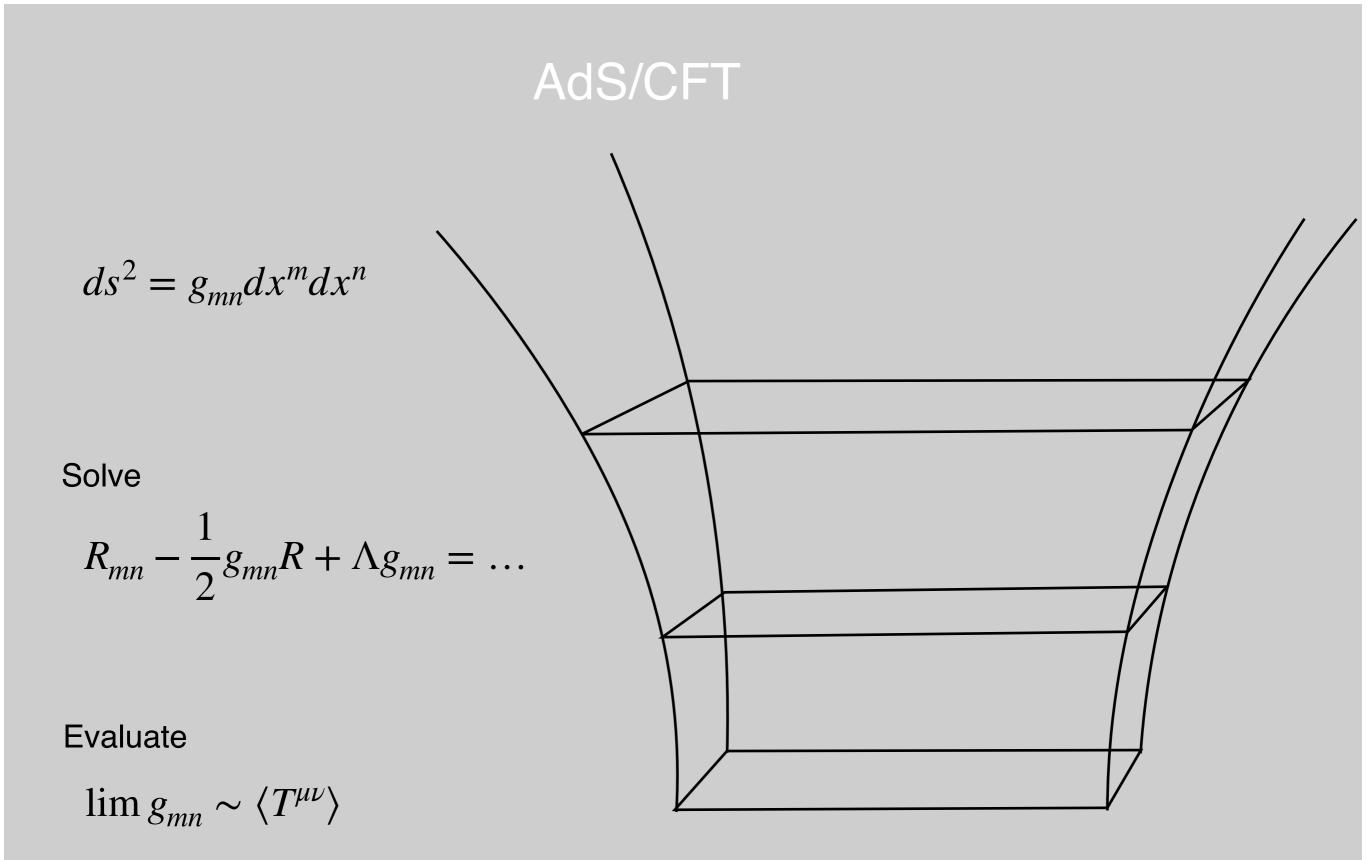
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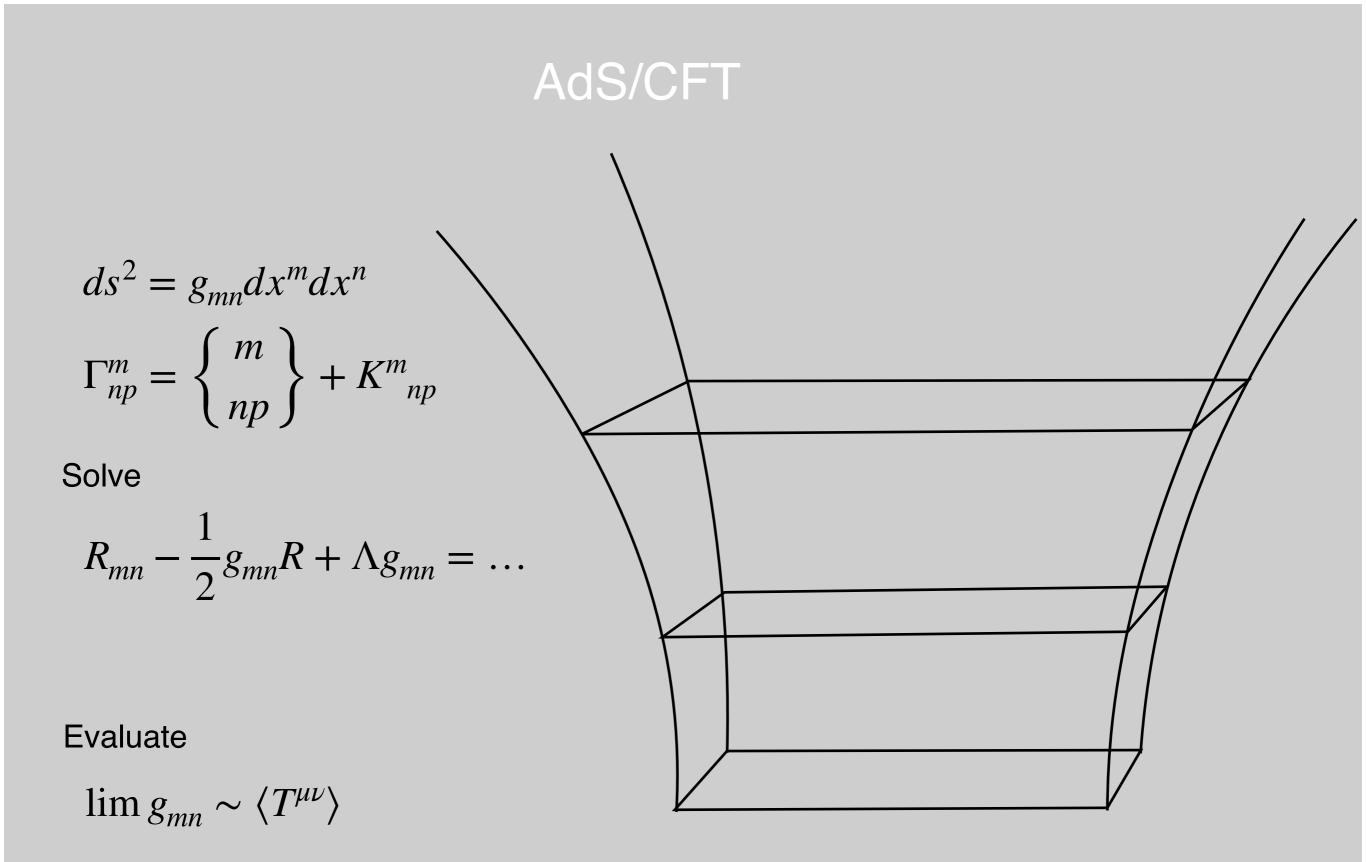
What is $\langle S^{\mu\nu\rho} \rangle$ in a hydrodynamic configuration? (For large N and large g)

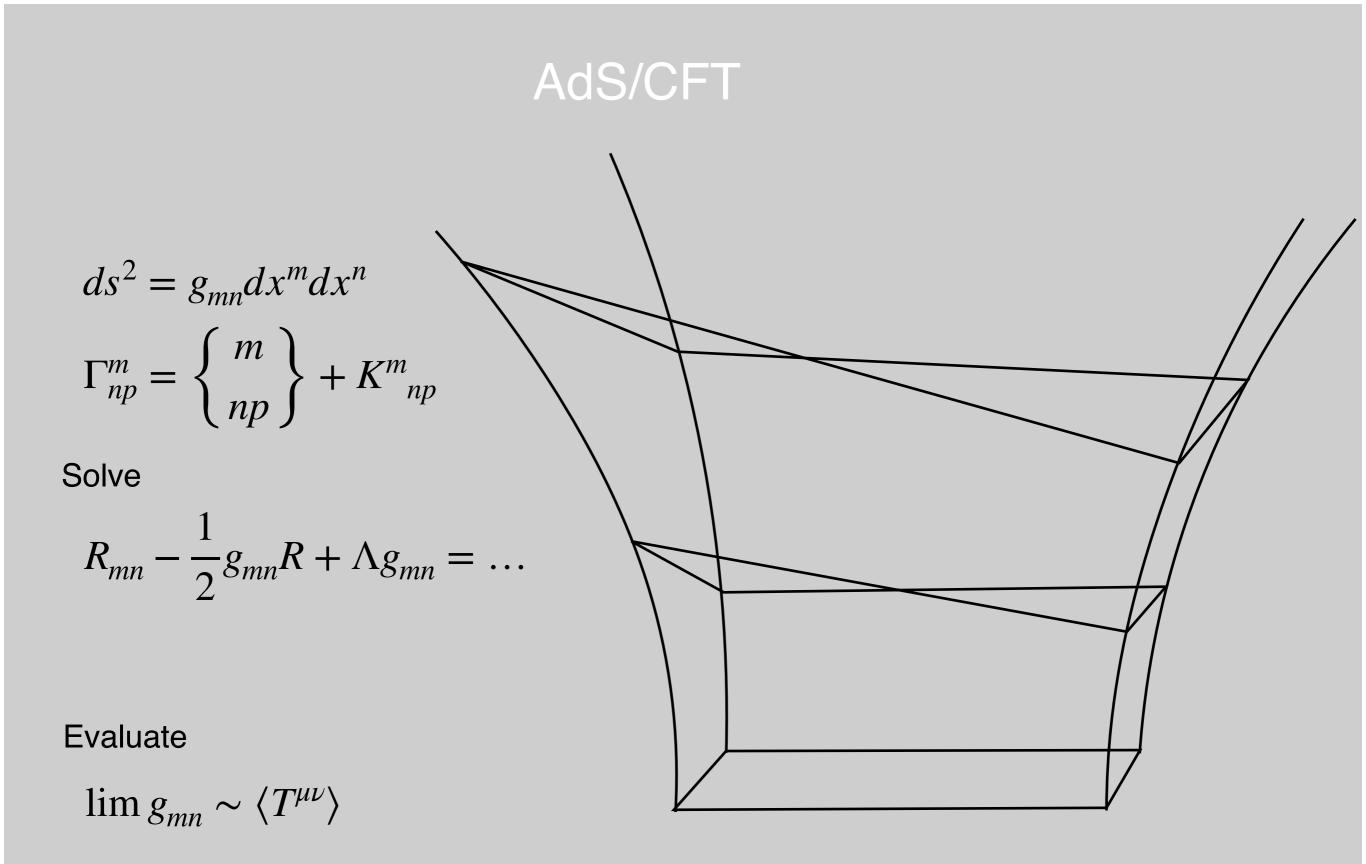
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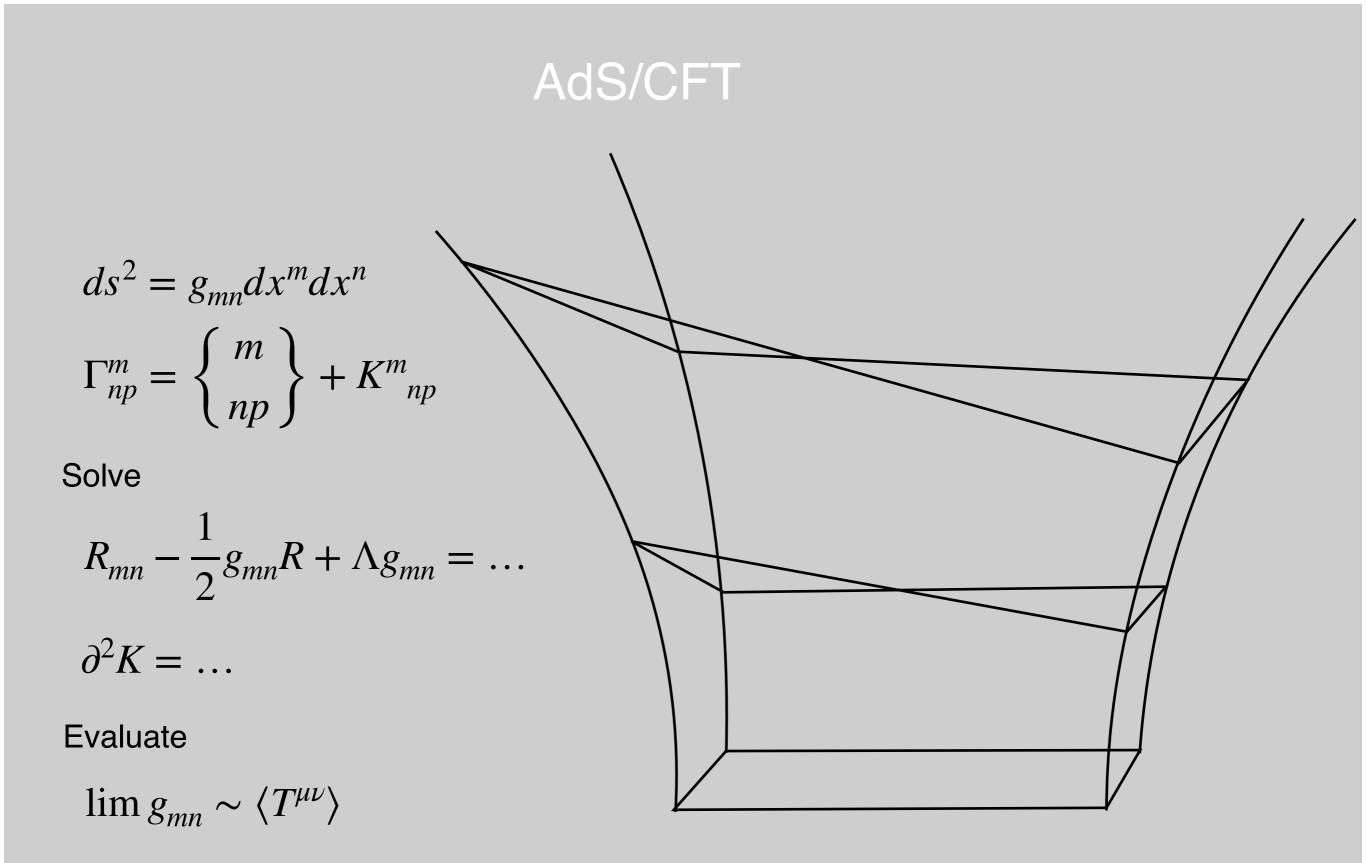
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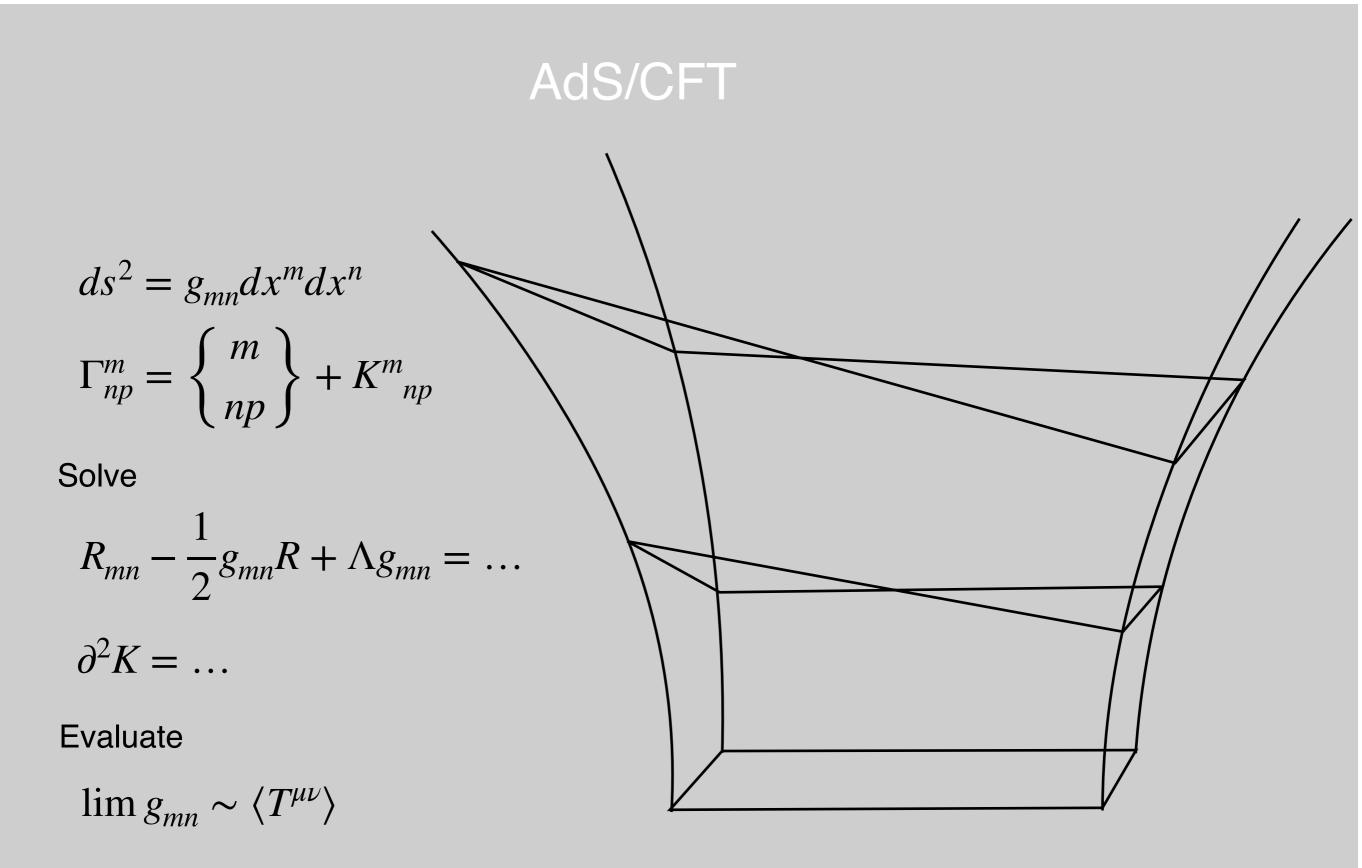




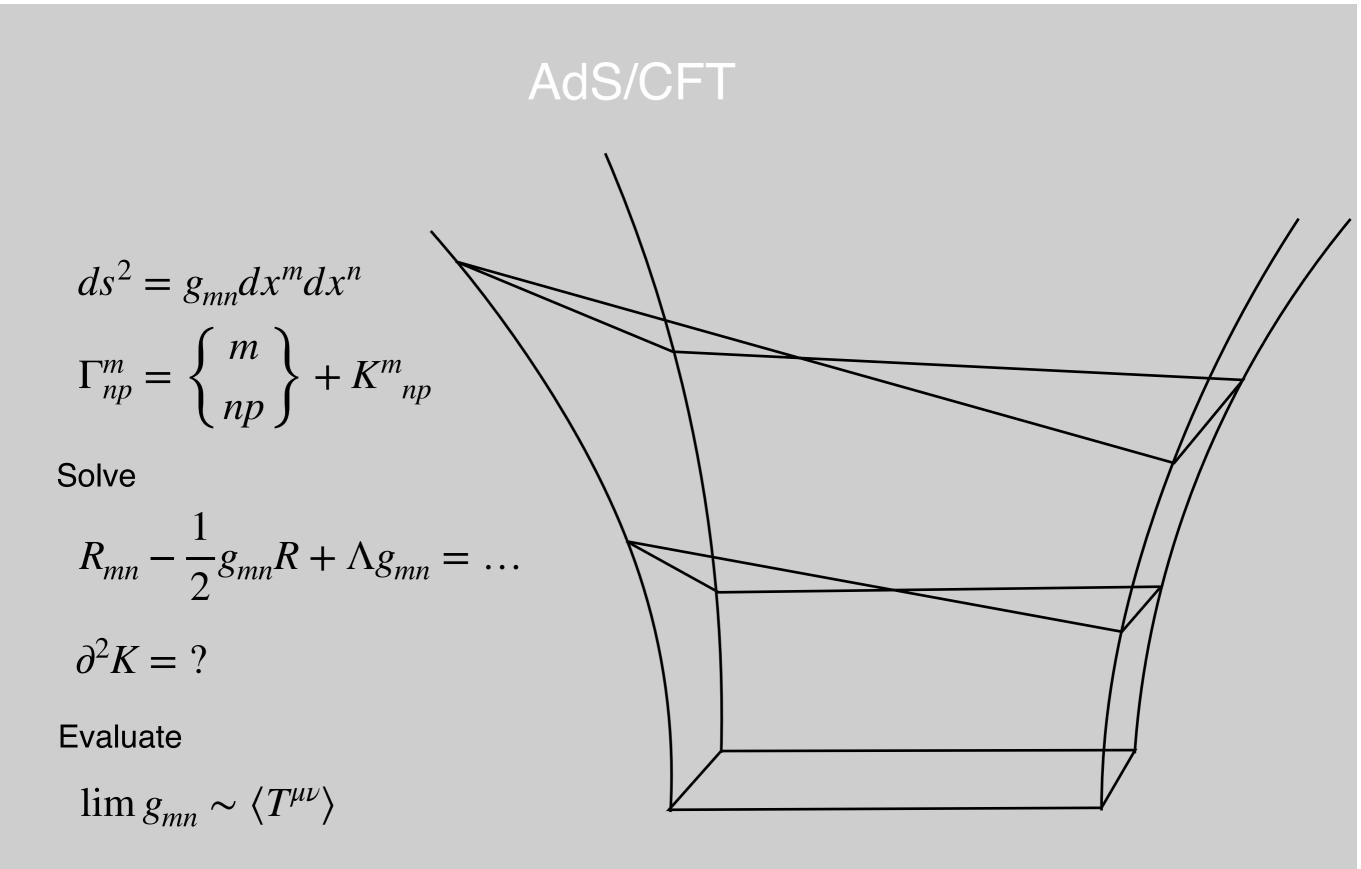








 $\lim K^m_{\ np} \sim \langle S^\mu_{\ \nu\rho} \rangle$



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$$V^1 \to V^1 \cos \theta + V^2 \sin \theta$$

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Likewise,

$$T^{\mu\nu} = \left(\begin{array}{cc} \epsilon & \dots \\ \vdots & \ddots \end{array}\right)$$

SUSY is a symmetry that mixes bosons and fermions, schematically,

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Our main result is that

$$S^{\mu\nu\rho} = \epsilon^{\mu\nu\rho\sigma} R_{\sigma}$$

for a large (infinite) class of supersymmetric theories.

SUSY is a symmetry that mixes bosons and fermions, schematically,

$$\phi \to \phi + \bar{\theta} \psi$$

Composite operators will also transform, schematically,

$$T^{\mu\nu} \to T^{\mu\nu} + \bar{\theta}^{\mu} J^{\nu}$$

Thus,

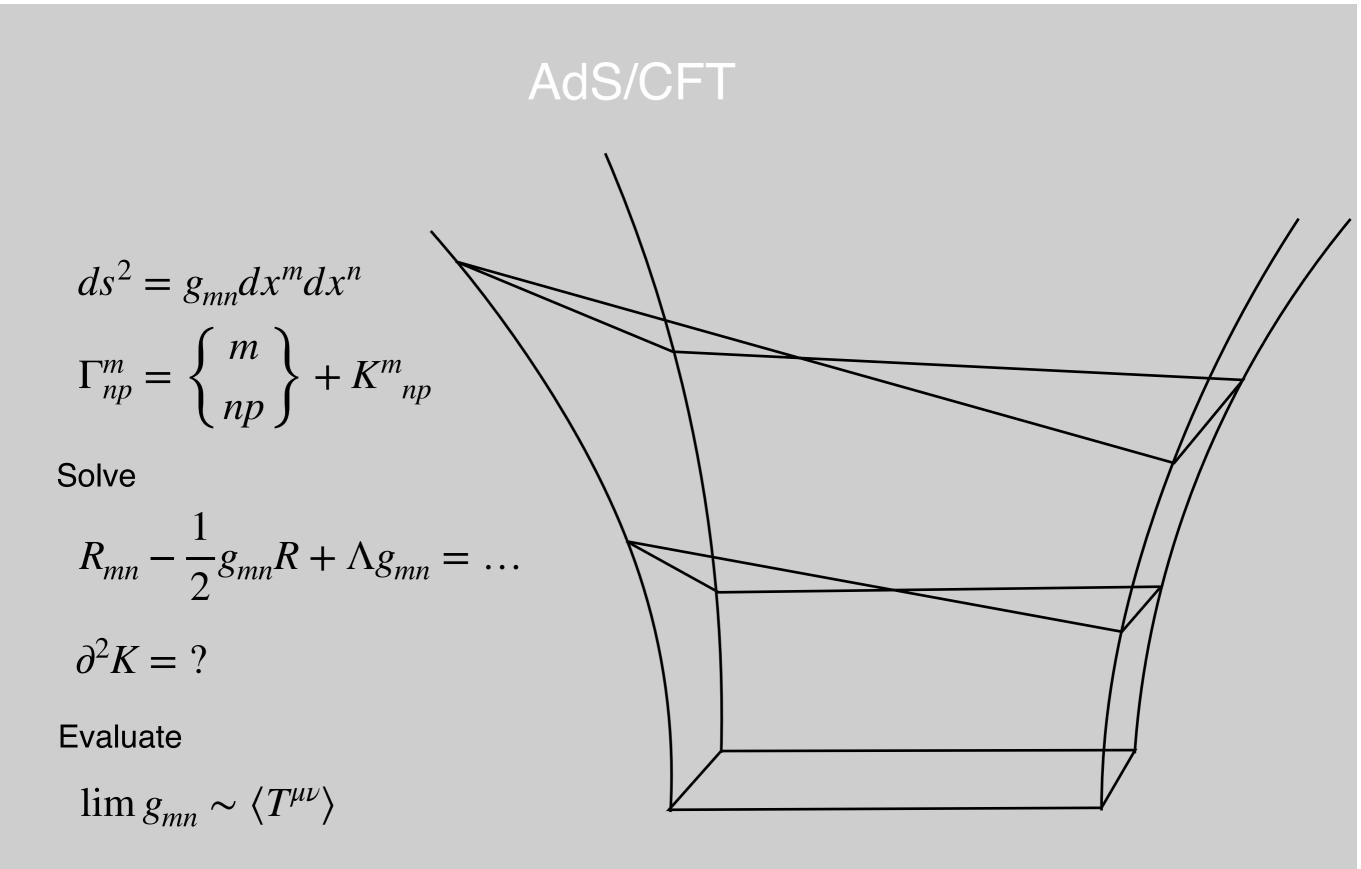
$$\mathbf{T}^{\mu} = \left(T^{\mu\nu} \quad J^{\mu}_{\alpha} \quad R^{\mu}\right)$$

$$\mathscr{L} = \operatorname{tr}\left(\frac{1}{2g}F_{\mu\nu}F^{\mu\nu} + \sum_{I=1}^{6}\partial_{\mu}X^{I}\partial^{\mu}X^{I} + i\sum_{A=1}^{4}\bar{\psi}^{A}\bar{\sigma}D_{\mu}\psi_{A} + \frac{g^{2}}{2}\sum_{I,J}[X^{I}, X^{J}][X^{I}, X^{J}] + \operatorname{Yukawa}_{\text{interactions}}\right)$$

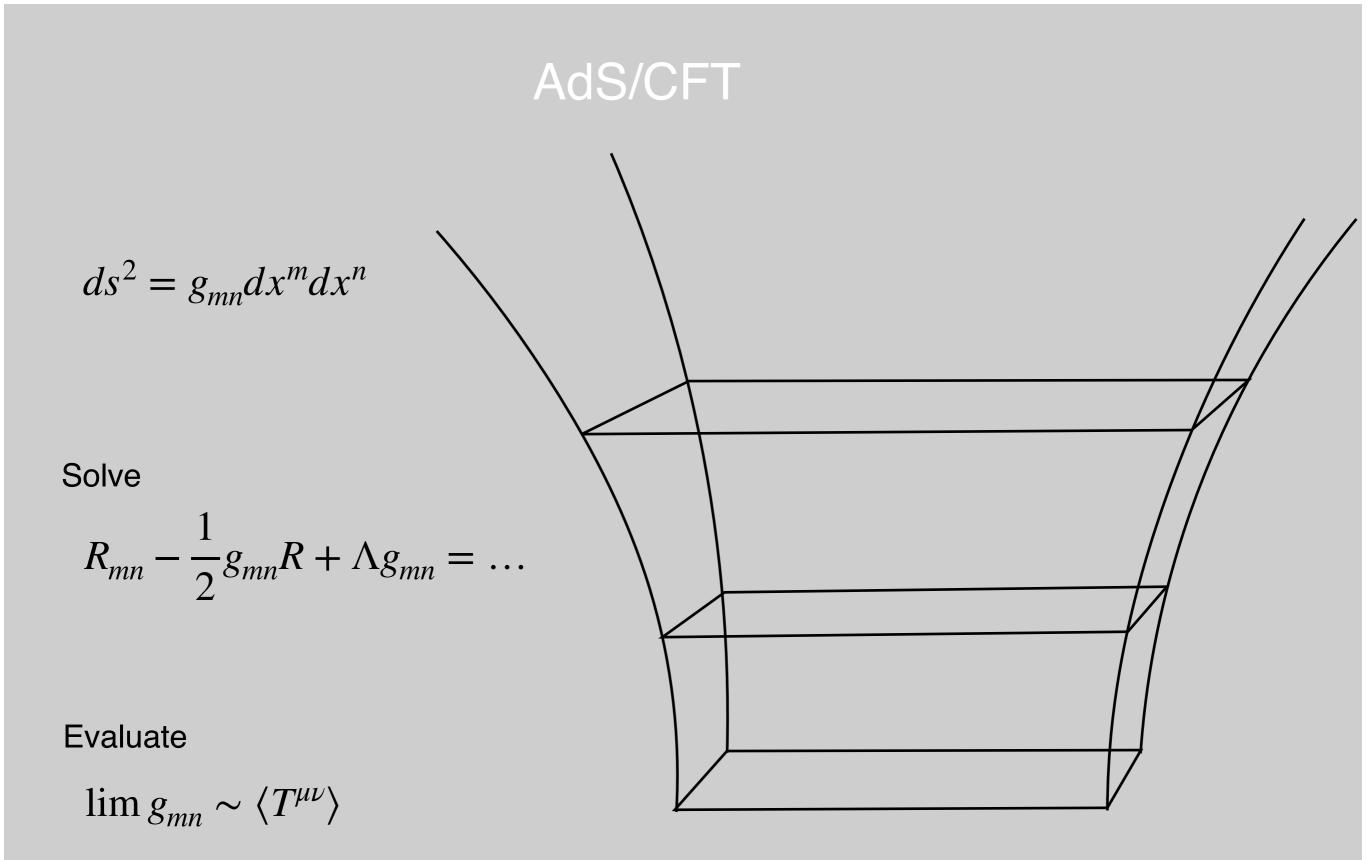
$$S^{\mu\nu\rho} = \epsilon^{\mu\nu\rho\sigma}R_{\sigma}$$

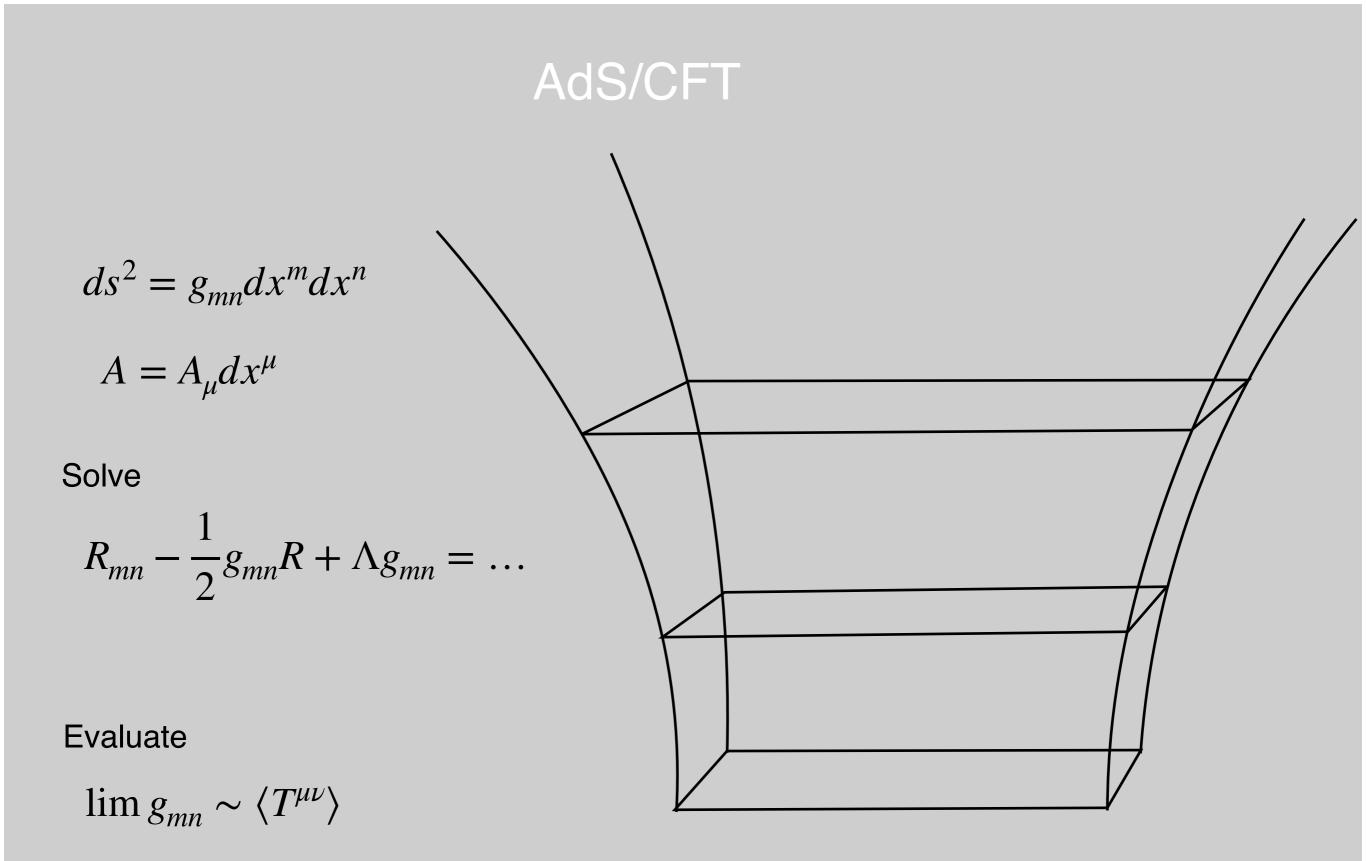
Our main result is that

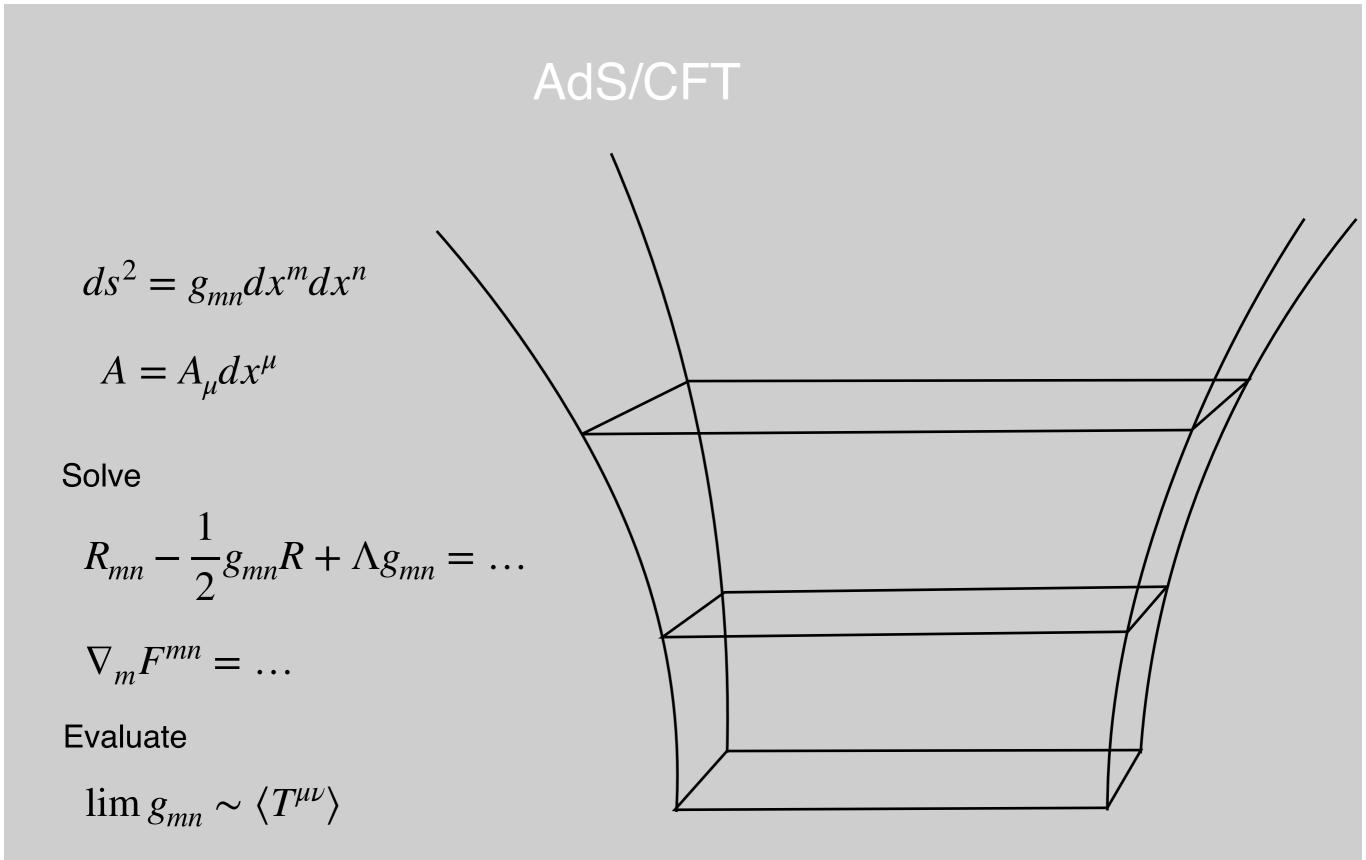
for a large (infinite) class of supersymmetric theories.

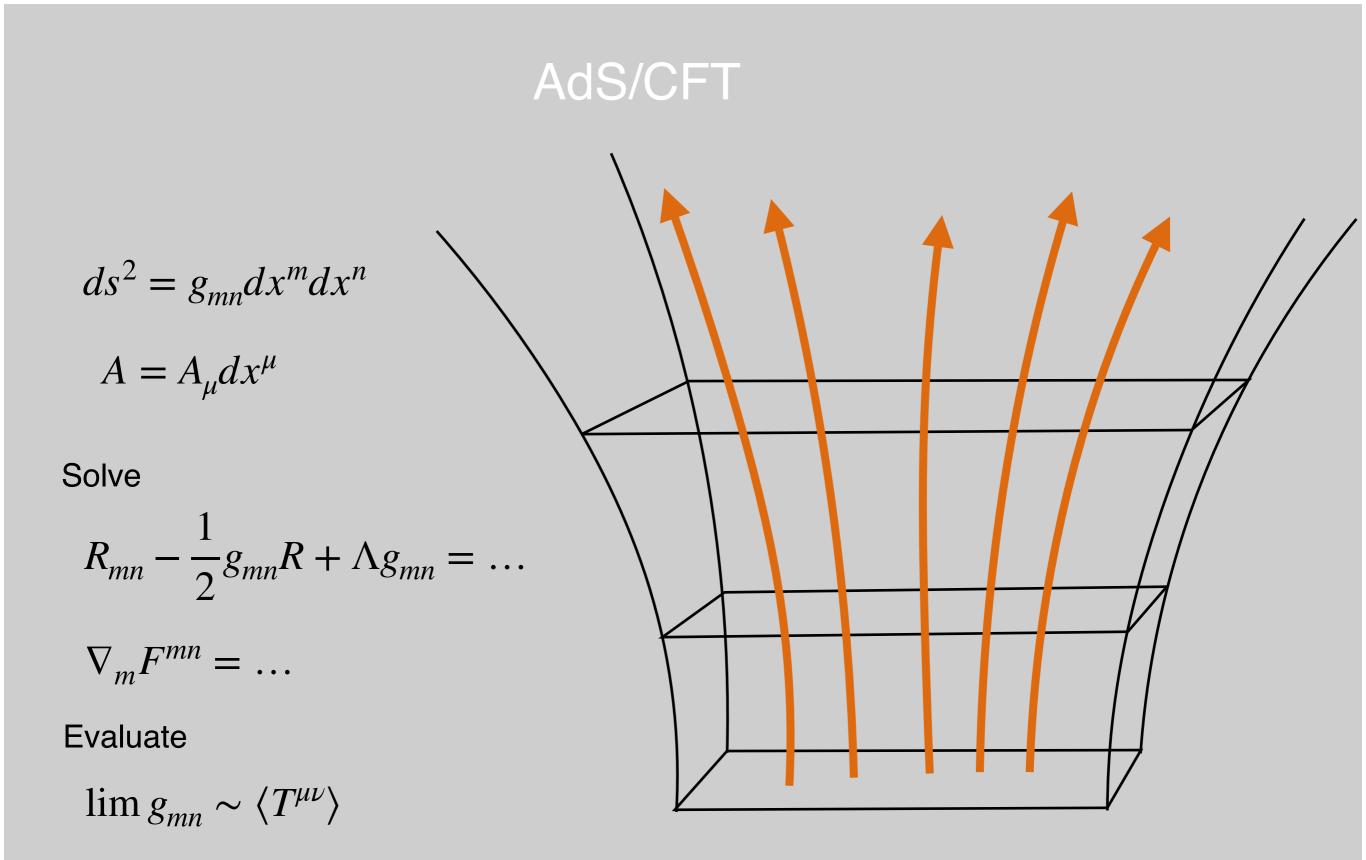


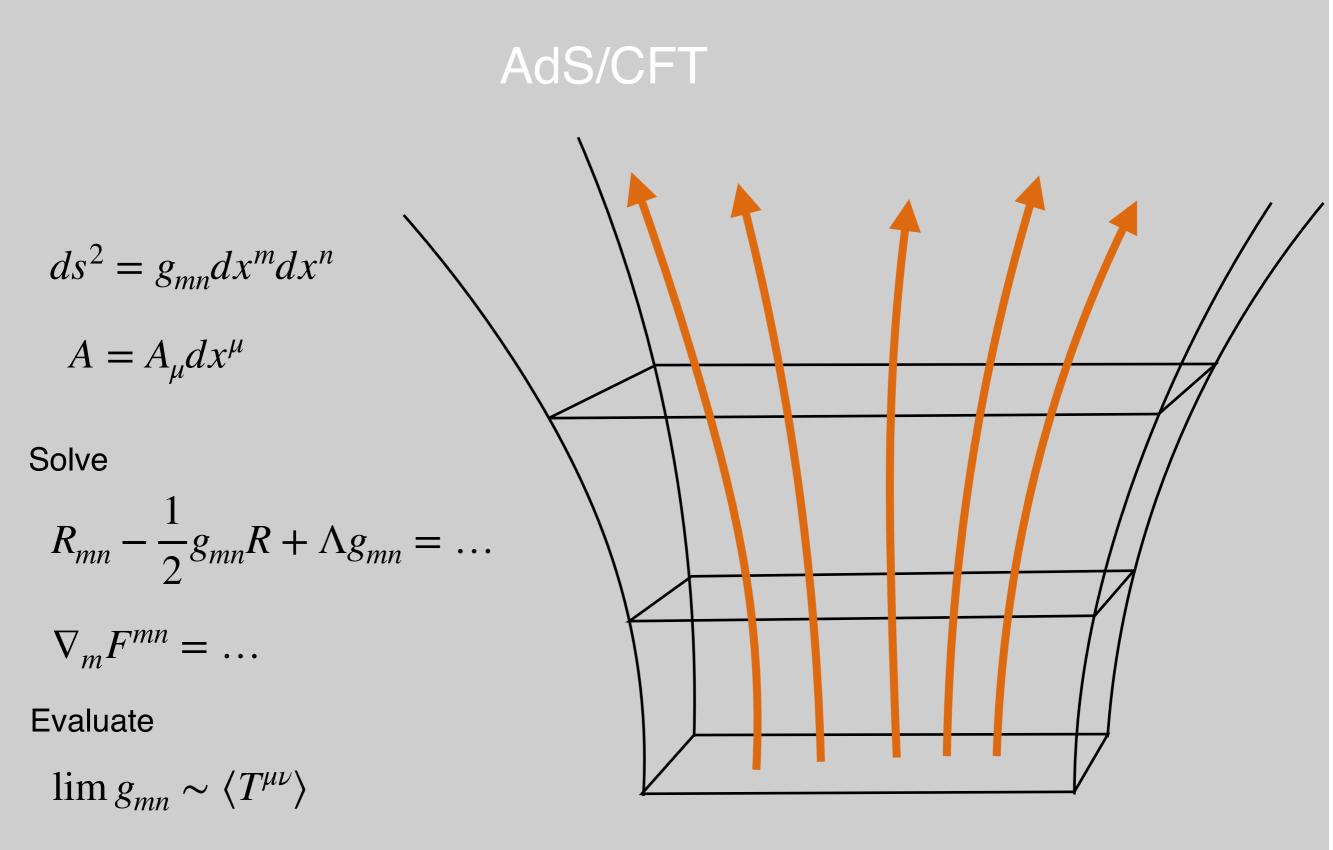
 $\lim K^m_{\ np} \sim \langle S^\mu_{\ \nu\rho} \rangle$







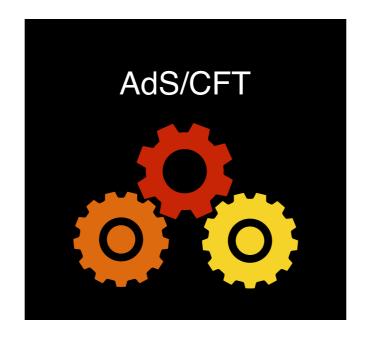


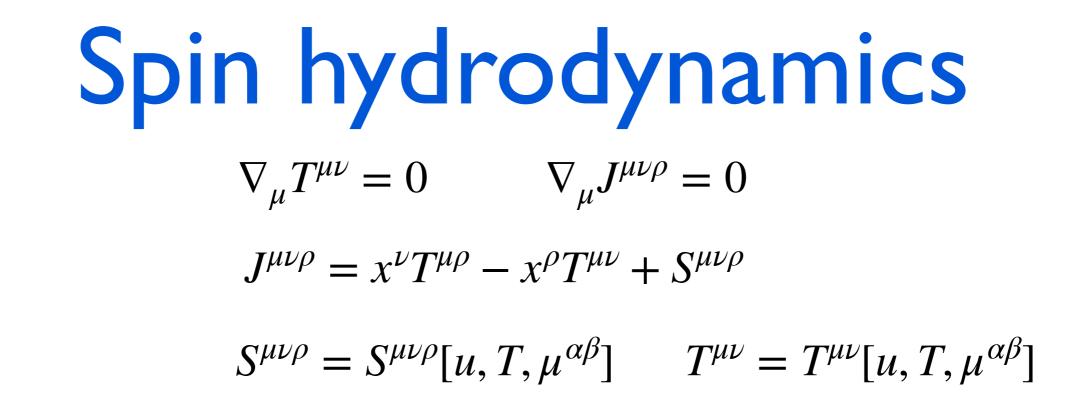


 $\lim A^m \sim \langle \epsilon^{\mu\nu\rho\sigma} S_{\nu\rho\sigma} \rangle$

Holographic hydrodynamics
$$\mathscr{L} = \operatorname{tr}\left(\frac{1}{2g}F_{\mu\nu}F^{\mu\nu} + \sum_{I=1}^{6}\partial_{\mu}X^{I}\partial^{\mu}X^{I} + i\sum_{A=1}^{4}\bar{\psi}^{A}\bar{\sigma}D_{\mu}\psi_{A} + \frac{g^{2}}{2}\sum_{I,J}[X^{I},X^{J}][X^{I},X^{J}] + \operatorname{Yukawa}_{\text{interactions}}\right)$$

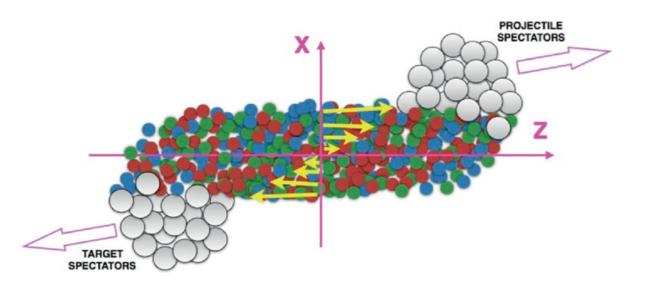
What is $\langle S^{\mu\nu\rho} \rangle$ in a hydrodynamic configuration? (For large N and large g)





It works?

It can be verified





Thank you