



UNIVERSITÀ
DI PISA



Istituto Nazionale di Fisica Nucleare

The Roberge-Weiss transition in $N_f = 2+1$ QCD with background magnetic fields

K. Zambello^{1,2} M. D'Elia^{1,2} L. Maio³ G. Zanichelli¹

¹ University of Pisa ² INFN - Sezione di Pisa

³ CPT, Marseille

Trento, 12/09/2024

Introduction

Introduction (I)

- ▶ The phase diagram of QCD in the presence of strong **magnetic fields** has been actively studied during recent years, being relevant for understanding a wide range of physical phenomena, from the physics of the early universe to heavy-ion collision experiments
- ▶ Some **interesting features**:

- chiral symmetry breaking enhanced at zero T , but chiral condensate decreases around T_c (**magnetic catalysis / inverse magnetic catalysis**)

- strengthening of the chiral transition, but chiral restoration temperature T_c **decreases** as a function of eB

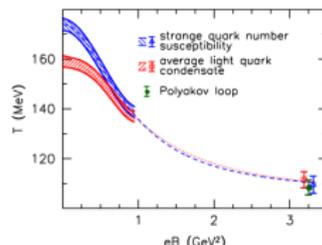
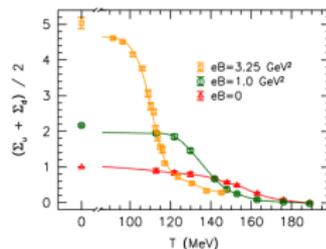


Fig. from JHEP 07, 173 (2015)

Introduction

Introduction (II)

► Some interesting features:

- the transition is a crossover at low eB , turns **first order** somewhere between 4 and 9 GeV^2

JHEP 07, 173 (2015) Phys.Rev.D 105, 034511 (2022)

but see also Phys.Rev.D 102, 054505 (2020)

- the (adimensional) **curvature** coefficient of the chiral crossover temperature $k_2 \sim 0.013(2)$ is weakly dependent on eB ;

the (physical) curvature coefficient $A_2(eB) = k_2(eB)/T_c(eB)$ qualitatively changes behaviour at $eB \sim 0.6 \text{ GeV}^2$

$$\frac{T_c(eB, \mu_B)}{T_c(eB)} = 1 - k_2 \left(\frac{\mu_B}{T_c(eB)} \right)^2$$

Phys.Rev.D 100, 114503 (2019)

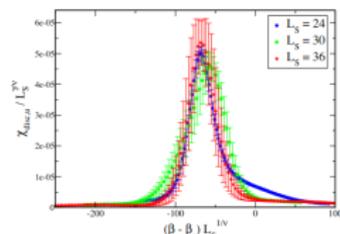
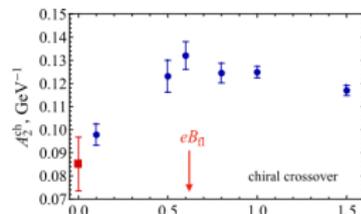
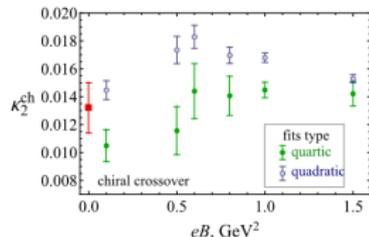


Fig. from Phys.Rev.D 105, 034511 (2022)



Figs. from Phys.Rev.D 100, 114503 (2019)

Introduction

Introduction (III)

- ▶ In this work we investigate the **Roberge-Weiss transition** in the phase diagram at imaginary chemical potentials

- RW line at $\mu_B/T = i\pi$, whose end-point $(i\pi, T_{RW})$ is believed to be a **second order** critical point for physical quark masses

Phys.Rev.D 93, 074504 (2016)

Phys.Rev.D 105, 034513 (2022)

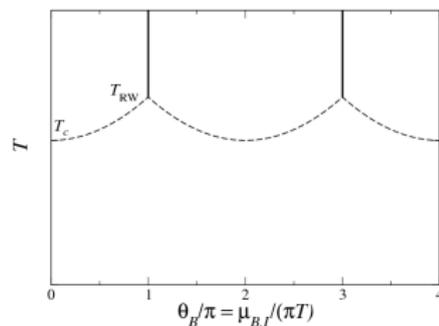
- indications that $T_{RW} \sim T_{Chiral}$ in the chiral limit

Phys.Rev.D 99, 014502 (2019)

Phys.Rev.D 106, 014510 (2022)

- ▶ **Questions:**

- What is the dependence of T_{RW} on eB ?
- What is the fate of the transition at strong magnetic fields?
- Is there any relation between T_{RW} and the chiral restoration temperature?



Introduction

Numerical set-up

Numerical set-up:

- ▶ $N_f = 2 + 1$, stout-staggered fermions with physical masses, tree-level Symanzik improved action
- ▶ $N_t = 6, 8$ lattices with different volumes
- ▶ We stay at constant chemical potential $\mu_f/T = i\pi$
- ▶ We use the (imaginary part of the) Polyakov loop $L = \langle |Im L| \rangle$ as the order parameter of the transition.

For different eB we estimate T_{RW} as the inflection point of $L(T)$ and the peak of its susceptibility $\chi_L(T)$,

$$L = \langle |Im L| \rangle$$

$$\chi_L = N_t N_s^3 (\langle (Im L)^2 \rangle - \langle |Im L| \rangle^2)$$

At fixed N_t and b_z , the temperature T is tuned by changing a . The magnetic field is $eB = \frac{6\pi b_z}{(aN_s)^2} = \frac{6\pi b_z N_t^2}{N_s^2} T^2$.

- ▶ Finite-size scaling analysis to determine the order of the transition

$$\chi_L = N_s^{\frac{\gamma}{\nu}} \phi(t N_s^{\frac{1}{\nu}}), \quad t = \frac{T - T_{RW}}{T_{RW}}$$

Transition at finite eB

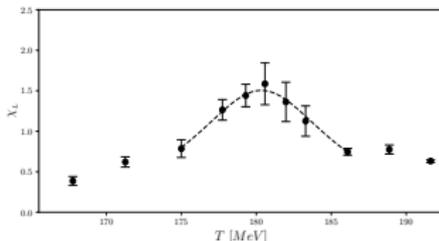
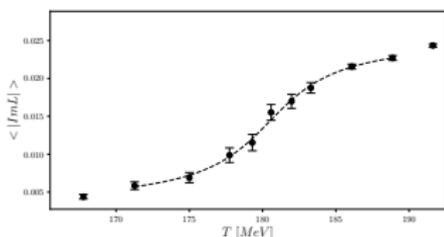
Transition at 0.2, 0.4, 0.6 GeV^2

$N_t = 6$ runs at $eB = 0.2, 0.4, 0.6 \text{ GeV}^2$ (and 0 GeV^2)

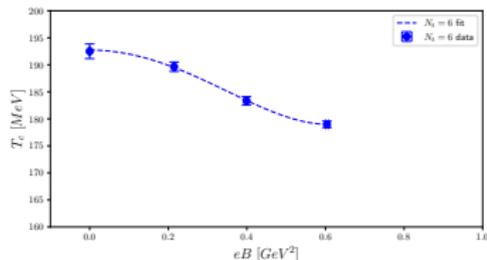
- ▶ $N_s = 18, 24$ give similar results for T_{RW} , finite-size effects are tiny

i.e. $T_{RW}(N_s = 18, eB = 0.6 \text{ GeV}^2) = 180.38(69) \text{ MeV}$

$T_{RW}(N_s = 24, eB = 0.6 \text{ GeV}^2) = 178.99(59) \text{ MeV}$



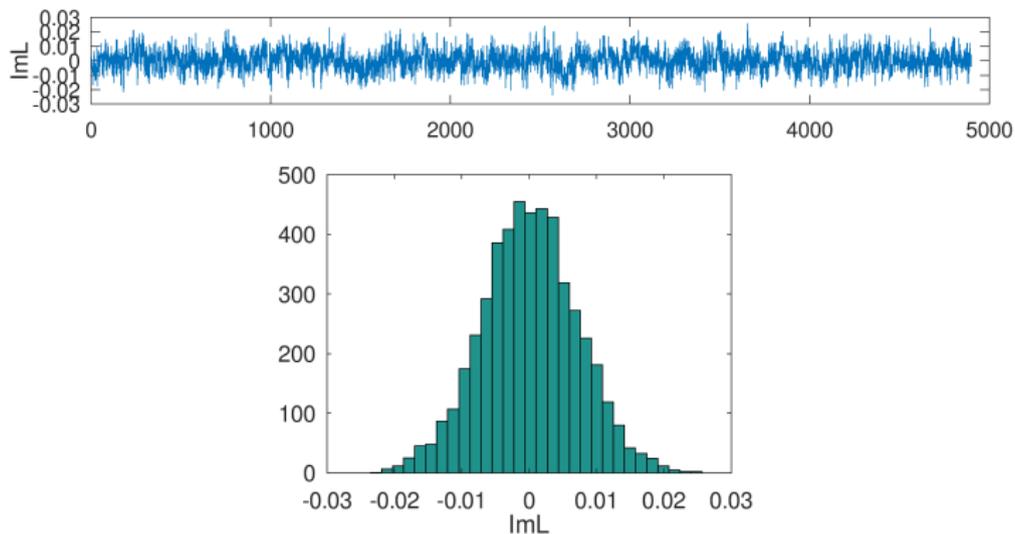
- ▶ We observe that T_{RW} decreases as a function of eB and the data fit well to a rational function $T_{RW}(eB) = T_{RW}^0 \frac{1+a(eB)^2}{1+b(eB)^2}$



Transition at finite eB

Crosscheck for the RW line (low T)

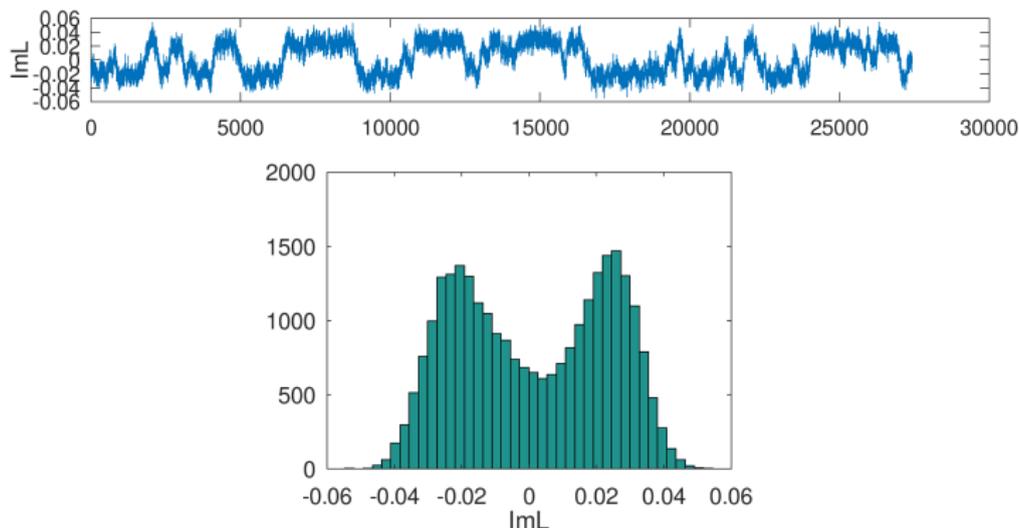
Monte Carlo history and histogram of $\text{Im}L$ at $\frac{\mu_f}{T} = i\pi$, $T < T_{RW}$



Transition at finite eB

Crosscheck for the RW line (high T)

Monte Carlo history and histogram of $\text{Im}L$ at $\frac{\mu_f}{T} = i\pi$, $T \gtrsim T_{RW}$



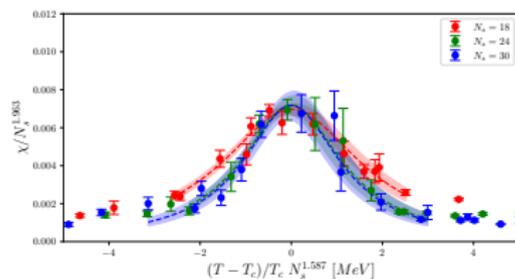
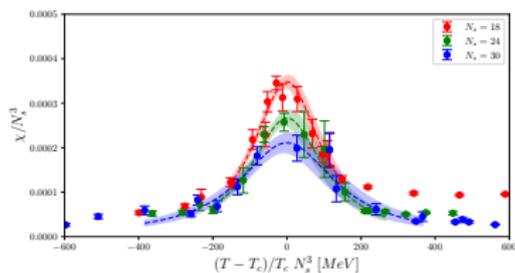
→ first order phase transition between left and right sectors

Transition at finite eB

FSS analysis at 1 GeV^2

$N_t = 6$ runs at $eB = 1 \text{ GeV}^2$

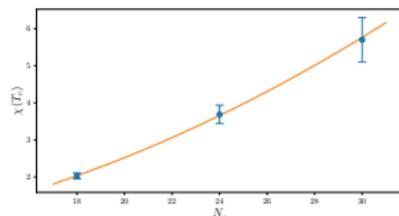
- ▶ Finite-size scaling analysis for $N_s = 18, 24, 30$, collapse plots:



→ compatible with a **second order** transition of the $Z(2)$ universality class

- ▶ Fitting the peaks of χ_L :

$$\chi_{max}(N_s) = \alpha N_s^{\frac{\gamma}{\nu}} \rightarrow \frac{\gamma}{\nu} = 2.04(19)$$

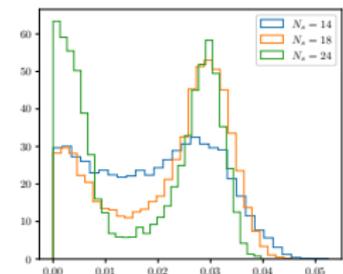
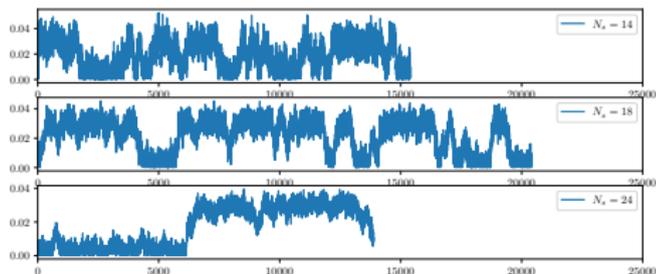


Transition at finite eB

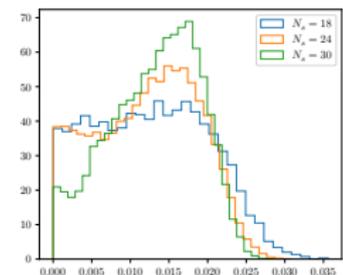
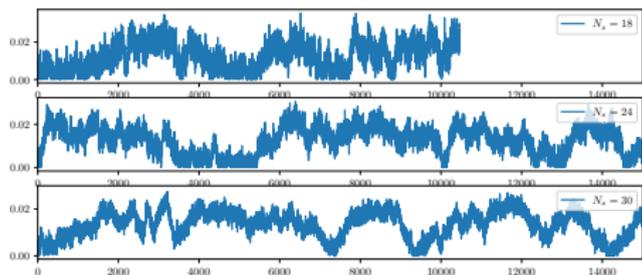
Transition at 2.5 GeV^2 (I)

$N_t = 6$ runs at $eB = 2.5 \text{ GeV}^2$

Histograms show a **double peaked** distribution, suggesting the presence of metastable states typical of a **first order** transition.



This can be compared with $eB = 1.0 \text{ GeV}^2$,

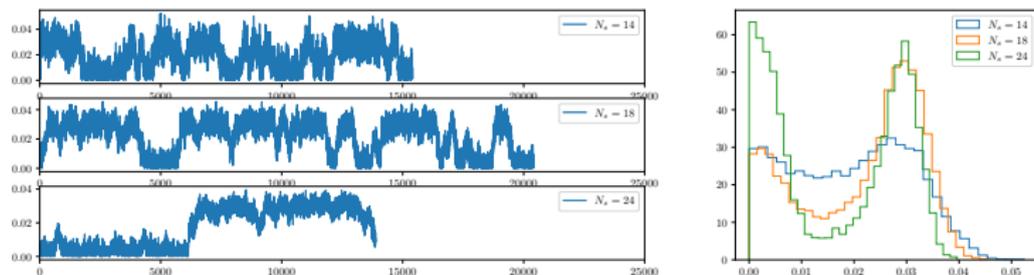


Transition at finite eB

Transition at 2.5 GeV^2 (II)

$N_t = 6$ runs at $eB = 2.5 \text{ GeV}^2$

Histograms show a **double peaked** distribution, suggesting the presence of metastable states typical of a **first order** transition.



But we expect large discretization effects.

$$eB = \frac{6\pi b_z N_t^2}{N_s^2} T^2, \text{ we want } \frac{b_z}{N_s^2} \ll 1$$

$$\text{At } 1 \text{ GeV}^2, b_z = 30 \text{ and } N_s = 24 \rightarrow \frac{b_z}{N_s^2} \approx 0.05$$

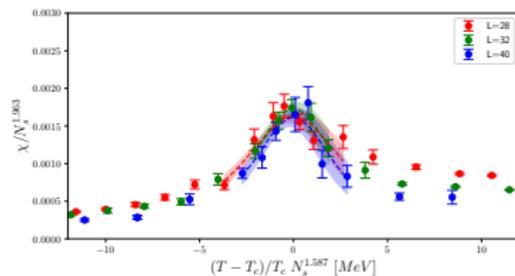
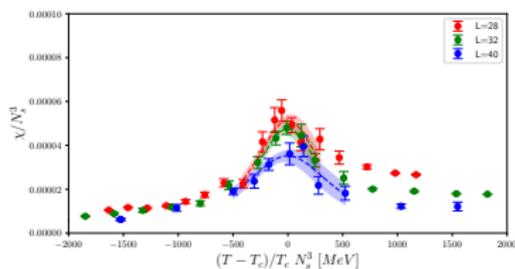
$$\text{At } 2.5 \text{ GeV}^2, b_z = 89 \text{ and } N_s = 24 \rightarrow \frac{b_z}{N_s^2} \approx 0.15$$

Towards the continuum limit

FSS analysis at 1 GeV^2

$N_t = 8$ runs at $eB = 1 \text{ GeV}^2$

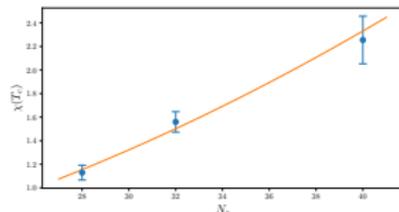
- ▶ Finite-size scaling analysis for $N_s = 28, 32, 40$, collapse plots:



→ compatible with a **second order** transition of the $Z(2)$ universality class

- ▶ Fitting the peaks of χ_L :

$$\chi_{max}(N_s) = \alpha N_s^{\frac{\gamma}{\nu}} \rightarrow \frac{\gamma}{\nu} = 1.97(28)$$

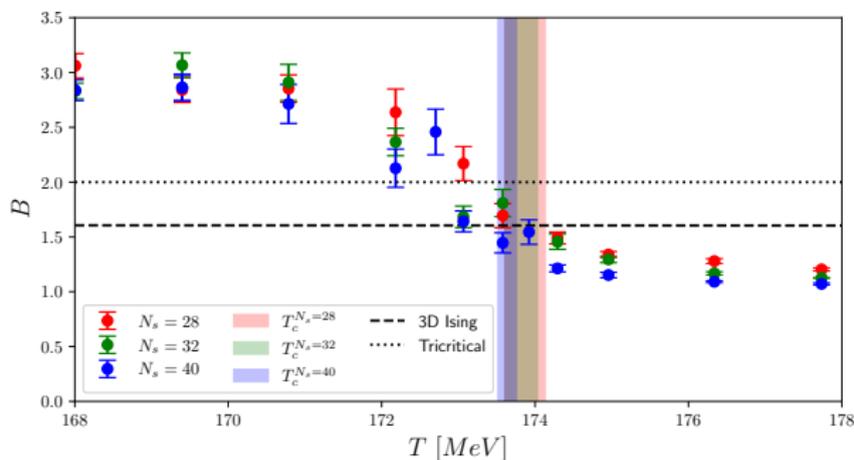


Towards the continuum limit

FSS analysis at 1 GeV^2

Binder cumulant:

Results are compatible with a critical point belonging to the $Z(2)$ universality class, but a tricritical point cannot be ruled out.

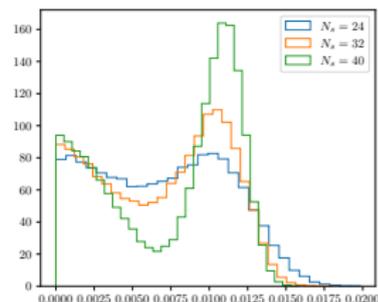
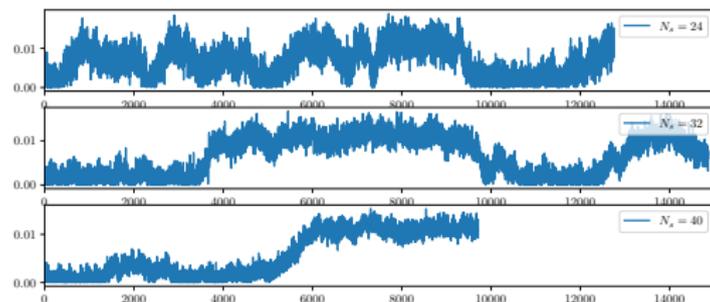


Towards the continuum limit

FSS analysis at 2.5 GeV^2

$N_t = 8$ runs at $eB = 2.5 \text{ GeV}^2$

Histograms still show a **double peaked** distribution, suggesting the presence of metastable states typical of a **first order** transition.



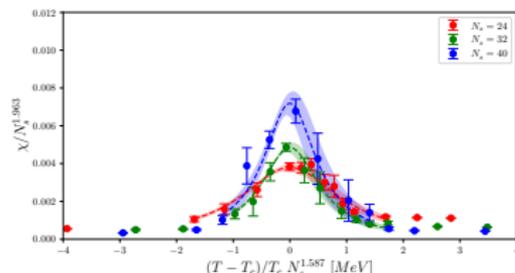
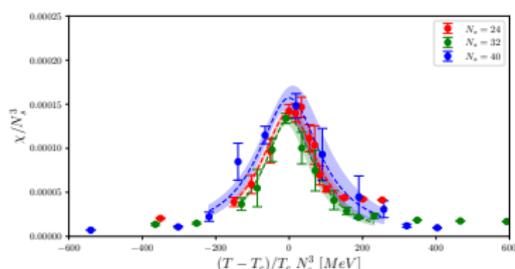
In this case $\frac{b_Z}{N_s^2} \approx 0.08$.

Towards the continuum limit

FSS analysis at 2.5 GeV^2

$N_t = 8$ runs at $eB = 2.5 \text{ GeV}^2$

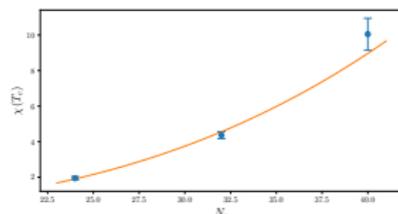
- ▶ Finite-size scaling analysis for $N_s = 24, 32, 40$, collapse plots:



→ compatible with a **first order** transition

- ▶ Fitting the peaks of χ_L :

$$\chi_{max}(N_s) = \alpha N_s^{\frac{\gamma}{\nu}} \rightarrow \frac{\gamma}{\nu} = 3.03(18)$$



Curvature of the critical line

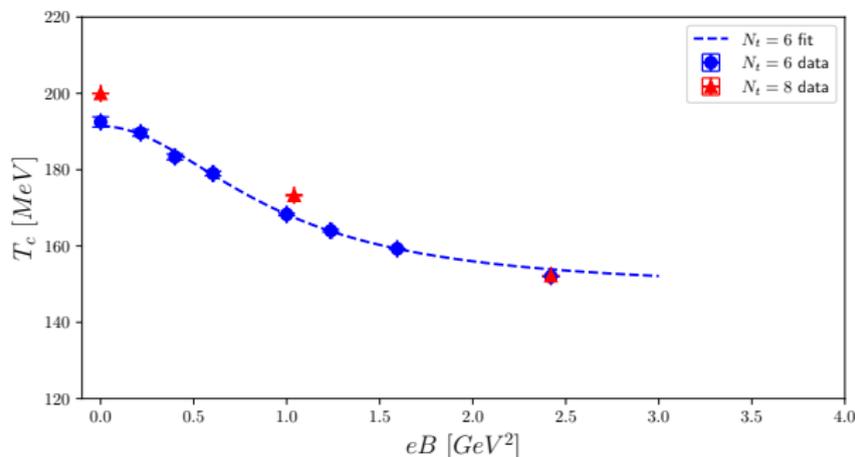
Curvature of the critical line (I)

All in all, data fit well to a rational function

$$T_{RW}(eB) = T_{RW}^0 \frac{1+a(eB)^2}{1+b(eB)^2} \text{ up to } 1.6 \text{ GeV}^2.$$

Full data set (including 2.5 GeV^2) well parametrized by

$$T_{RW}(eB) = T_{RW}^0 \frac{1+a(eB)^2+c(eB)^4}{1+b(eB)^2+d(eB)^4}.$$



Curvature of the critical line

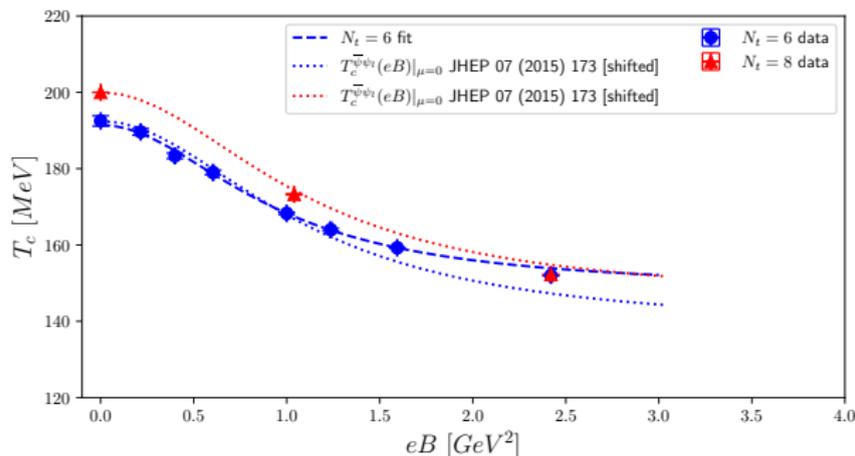
Curvature of the critical line (II)

We can Taylor expand the rational function ansatz around $eB = 0$:

$$T_{RW}(eB) = T_{RW}^0 + k(eB)^2$$

Curvature close to the curvature of the chiral critical line at $\mu = 0$ found by ref. JHEP 07, 173 (2015) from $\langle \bar{\psi}_I \psi_I \rangle$:

$$k \sim -44.8 \leftrightarrow k_{(eB)^2} = -50.0(3.5), \quad k_{(eB)^4} = -56(10)$$



To summarize:

- ▶ We have studied the RW end-point in the presence of background magnetic fields
- ▶ The RW temperature decreases as a function of the strength of the magnetic field
- ▶ We have found indications that the RW transition becomes first order between 1 and 2.5 GeV^2
- ▶ The curvature of the critical line is close to the curvature of the chiral critical line (from the light quark condensate) at $\mu = 0$

Thank you for listening!