Bayes of My Life: Bayesian analysis of nucleon-nucleon scattering data in pionless EFT

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The Nuclear Interaction: Post-Modern Developments @ ECT\* 20 August 2024

Washington University in St. Louis



### Outline

- Motivation/Bayesian EFT Model Calibration
- BUQEYE Formalism
- Interaction Choice
- Results
- Questions



## Outline

- Motivation/Bayesian EFT Model Calibration





### arXiv:2408.02480







nucleon interactions.



### We are interested in calibrating the next generation of EFT nucleon-

### arXiv:2408.02480









nucleon interactions.

These models should have robust uncertainty quantification:











nucleon interactions.

These models should have robust uncertainty quantification: • Parametric uncertainty











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- Parametric uncertainty
- Truncation uncertainty











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This must be accomplished in the model calibration.















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How do we find them?





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Compute on the lattice?





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Let's cast the problem this way:



### $pr(\vec{a} | \vec{y}, I) \propto pr(\vec{y} | \vec{a})$



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 Compute on the lattice? So how do we calibrate the LECs?

Let's cast the problem this way:

additional external information.



## $pr(\vec{a} | \vec{y}, I) \propto pr(\vec{y} | \vec{a})$ where $\vec{a}$ is the vector of parameters, $\vec{y}$ is a vector of data, and I is







### We have





### We have

Posterior



## $pr(\vec{a} | \vec{y}, I) \propto pr(\vec{y} | \vec{a})$ Likelihood



### We have

Posterior

How do we calculate the likelihood?



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Posterior How do we calculate the likelihood? data, we assume this is given by



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# Let's say we have some experimental data, $\vec{y}_{exp}$ . For each piece of



## We have $pr(\vec{a} | \vec{y}, I) \propto pr(\vec{y} | \vec{a})$ Likelihood Posterior How do we calculate the likelihood? Let's say we have some experimental data, $\vec{y}_{exp}$ . For each piece of data, we assume this is given by $y_{\text{exp}} = y_{\text{th}}(\vec{a}) + \delta y_{\text{exp}}, \quad \delta y_{\text{exp}} \sim \mathcal{N}(0,\sigma^2)$

Likelihood Modeling









can easily arrive at the result





Since we made a Gaussian assumption for the experimental discrepancy, we can easily arrive at the result

 $pr(y | \vec{a}) \sim e$ 



$$y_{exp} - \left(y_{exp} - y_{th}(\vec{a})\right)^2 / 2\sigma^2$$



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$$y_{exp} - \left(y_{exp} - y_{th}(\vec{a})\right)^2 / 2\sigma^2$$

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$$-\left(y_{\exp}-y_{th}(\vec{a})\right)^2/2\sigma^2$$

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### Minimize → Find best model

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## Bayes' Theorem





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Posterior What we found for a Least-squares optimization essentially had  $pr(\vec{a} \mid I) = C.$ 



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• Ex: LECs are natural, i.e., order



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$$1 \rightarrow \operatorname{pr}(\vec{a} | \mathbf{I}) \sim \mathcal{N}\left(\vec{0}, \Sigma_{\mathrm{pr}}\right)$$







different from least-squares



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### Least-squares Parameters are fixed

different from least-squares

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- Robust handling of data outliers



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- Rigorous uncertainties



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- Approximate uncertainties

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# **Bayesian calibration**

- Parameters are random variables
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- Use of prior information
- Robust handling of data outliers
- Rigorous uncertainties
- Easy extension to full model uncertainties



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- Approximate uncertainties
- Convoluted extension to model uncertainties







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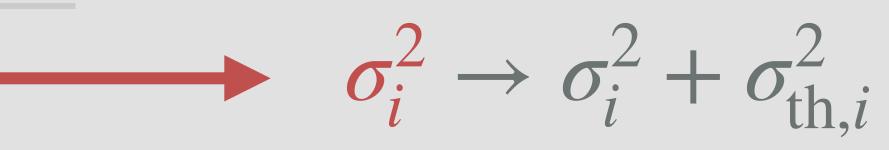


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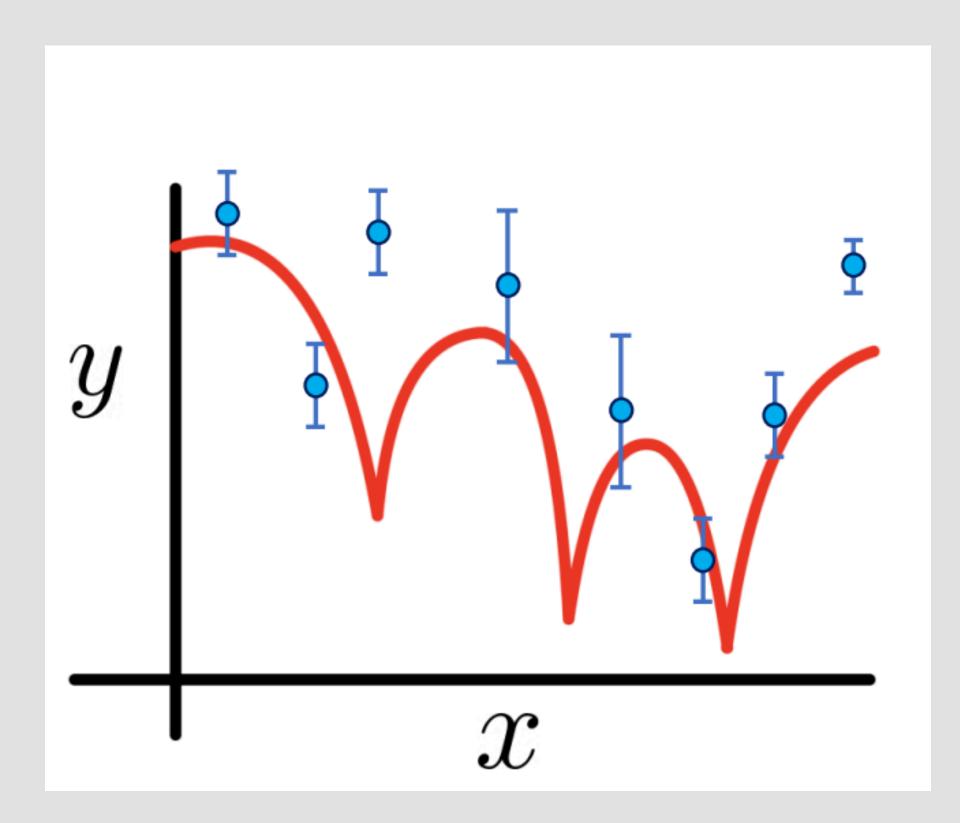






Why are theory errors necessary in calibration?

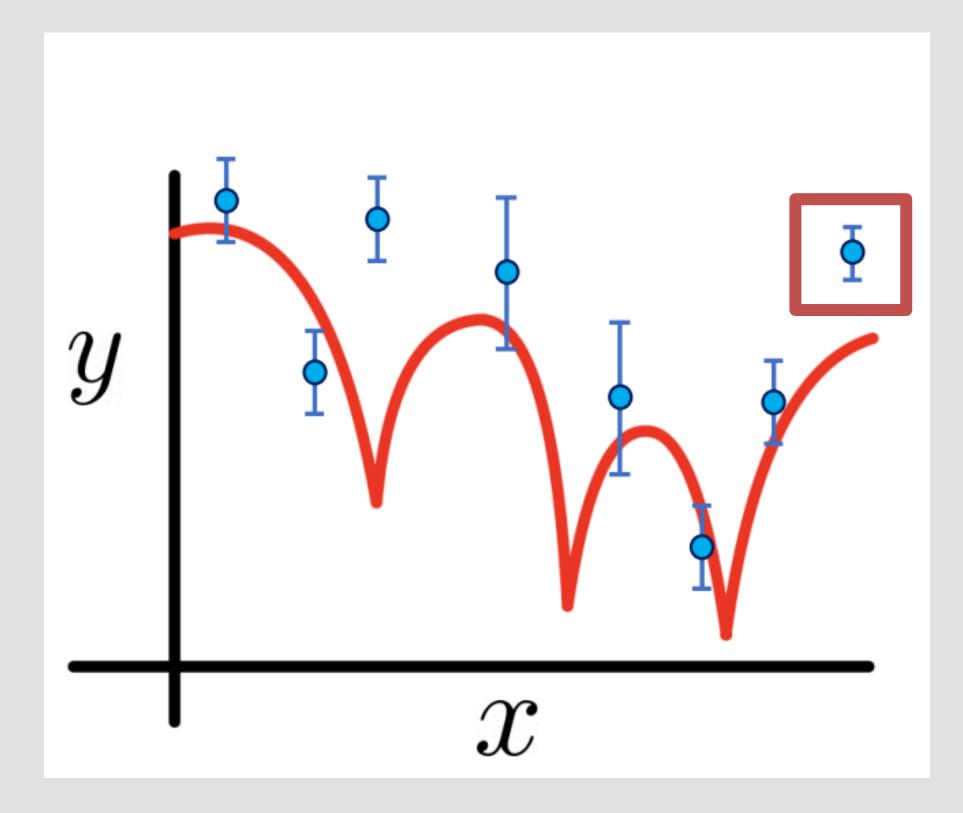






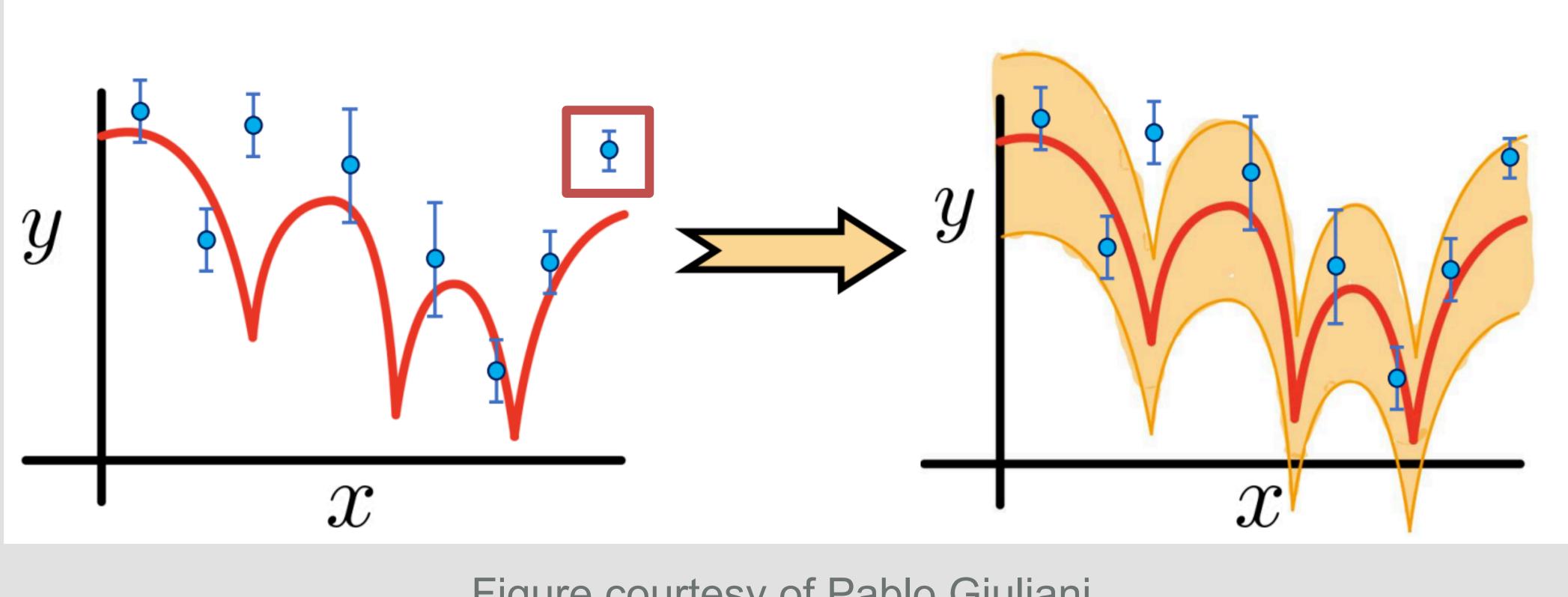
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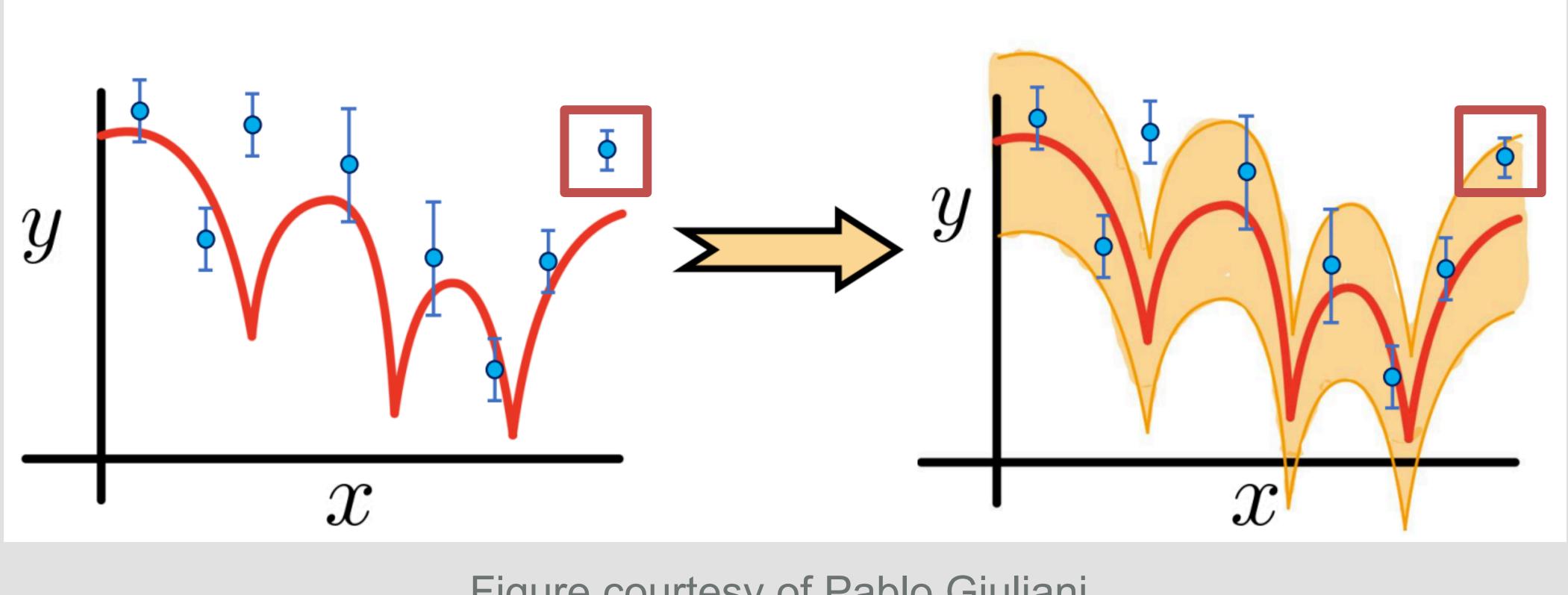
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#### Figure courtesy of Pablo Giuliani

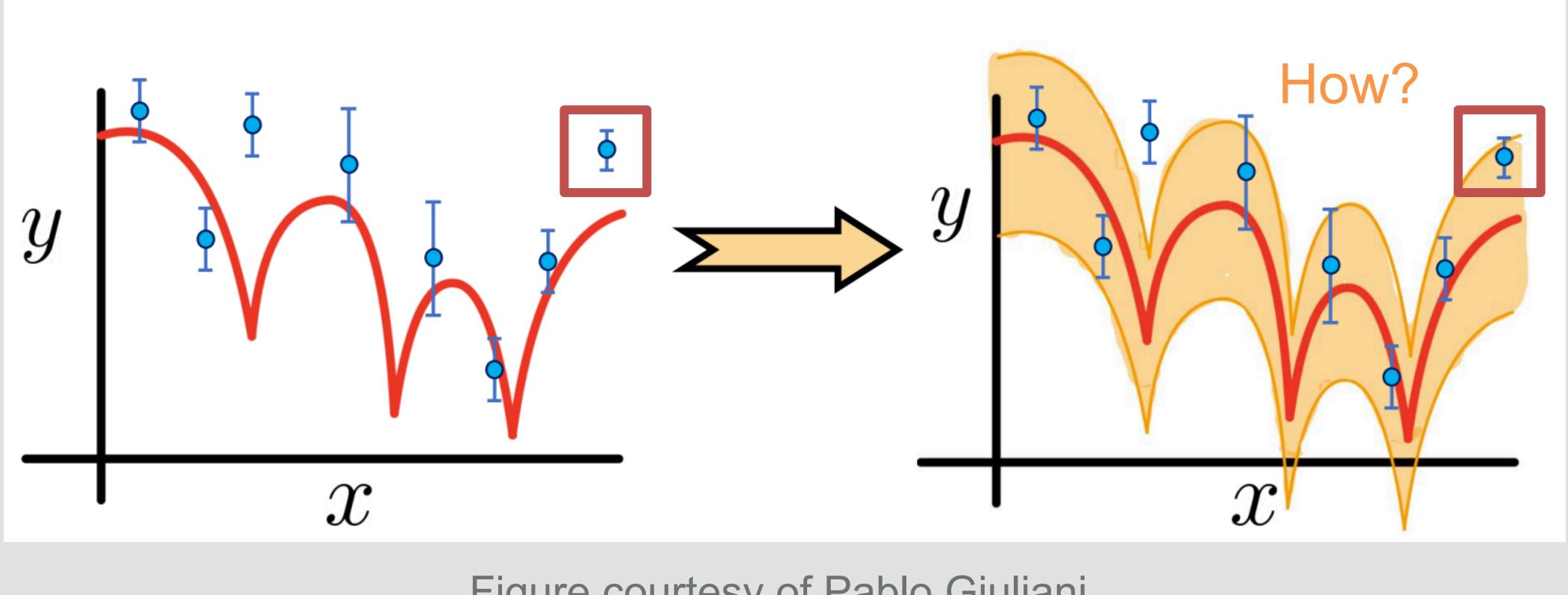
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# Outline

- Motivation/Bayesian
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# **-T Model Calibration**







Since our model is a perturbative series, we can model it as such\*: n=0



# $y_{\text{th}}(x) = y_{\text{ref}}(x) \sum_{n=1}^{\infty} c_n(x) Q^n(x), \quad Q \equiv \frac{\max[p_{soft}, p]}{\Delta x},$

\*R. J. Furnstahl et. al. Phys. Rev. C 92, 024005





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$$y_{\text{th}}(x) = y_{\text{ref}}(x) \sum_{n=0}^{\infty} c_n(x)$$

where  $y_{ref}(x)$  sets a reference scale for the observable  $y_{th}$  and  $\Lambda_h$  is the EFT breakdown scale.



# $(p)Q^n(x), \quad Q \equiv \frac{\max[p_{soft}, p]}{\Lambda},$

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This series follows the truncation scheme of the EFT:

$$y_{\text{th}}(x) = y_{\text{ref}}(x) \sum_{n=0}^{k} c_n Q^n + y_{\text{ref}}(x)$$



# $(p)Q^n(x), \quad Q \equiv \frac{\max[p_{soft}, p]}{\Lambda},$

# $_{\rm f}(x) \sum c_n Q^n = y_{\rm th}^{(k)}(x) + \delta y_{\rm th}^{(k)}(x).$ n=k+1\*R. J. Furnstahl et. al. Phys. Rev. C 92, 024005









From the neglected terms, we have



# $\infty$ $\delta y_{\text{th}}^{(k)}(x) = y_{\text{ref}}(x) \sum c_n(x)Q^n(x).$ n=k+1



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This is a geometric series in Q, so we can find\*



# $\delta y_{\rm th}^{(k)}(x) = y_{\rm ref}(x) \sum_{n=1}^{\infty} c_n(x)Q^n(x).$ n=k+1 $\delta y_{\text{th}}^{(k)}(x) = \frac{y_{\text{ref}} \,\bar{c} \, Q^{(k+1)}}{1 - O},$

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Where we assume that  $c_n | \bar{c} \sim \mathcal{N}(0, \bar{c}^2)$ .



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#### Theoretical Covariance





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From the truncation uncertainty, we can construct a covariance matrix, assuming  $\delta y_{th}$  is normally distributed,



 $\Sigma_{ij}^{\text{th}} = \frac{\left(y_{\text{ref},i}\,\overline{c}\,Q_i^{(k+1)}\right)\left(y_{\text{ref},j}\,\overline{c}\,Q_j^{(k+1)}\right)}{1 - Q_i Q_j}r(x_i, x_j; \,\overline{l}),$ 



#### Theoretical Covariance

From the truncation uncertainty, we can construct a covariance matrix, assuming  $\delta y_{th}$  is normally distributed,  $\Sigma_{ij}^{\text{th}} = \frac{\left(y_{\text{ref},i}\,\overline{c}\,Q_i^{(k+1)}\right)\left(y_{\text{ref},j}\,\overline{c}\,Q_j^{(k+1)}\right)}{1 - Q_i Q_j}r(x_i, x_j;\,\overline{l}),$ were we introduce a kernel  $r(x_i, x_j; \vec{l})$  to smooth and handle

correlations.









We can build a total covariance,



### $\Sigma_{ij} = \Sigma_{ij}^{\exp} \delta_{ij} + \Sigma_{ij}^{\text{th}}.$



We can build a total covariance,

 $\Sigma_{ij} = \Sigma_{ij}$ And our correlated likelihood is n

 $pr(\vec{y} | \vec{a}, I) \propto e^{-(\vec{y}_{exp} - \vec{y}_{exp})}$ 



$$\sum_{ij}^{exp} \delta_{ij} + \sum_{ij}^{tn}$$
now
$$\vec{y}_{th} \int^{T} \Sigma^{-1} (\vec{y}_{exp} - \vec{y}_{th}) = e^{-d_{M}(\vec{a})}$$

1



We can build a total covariance,

 $\Sigma_{ij} = \Sigma$ 

And our correlated likelihood is now

 $pr(\vec{y} \mid \vec{a}, I) \propto e^{-\left(\vec{y}_{exp} - \vec{y}_{th}\right)^{T} \Sigma^{-1} \left(\vec{y}_{exp} + \vec{y}_{exp}\right)^{T} \Sigma^{-1} \left(\vec{y}_{exp} + \vec{y}_{exp}\right)^{T}$ 

efine the Mahalanobis distance  $d_{M}(\vec{a}) = \left(\vec{y}_{exp} - \vec{y}_{th}\right)^{T} \Sigma^{-1} \left(\vec{y}_{exp} - \vec{y}_{th}\right).$ 



$$\Sigma_{ij}^{\exp}\delta_{ij} + \Sigma_{ij}^{\text{th}}$$

$$\vec{y}_{th} \right)^{T} \Sigma^{-1} \left( \vec{y}_{exp} - \vec{y}_{th} \right) = e^{-d_{M}(\vec{a})}$$



We can build a total covariance,

And our correlated likelihood is now

 $pr(\vec{y} | \vec{a}, I) \propto e^{-(\vec{y}_{exp} - \vec{x}_{exp})}$ where we define the Mahalanobis distance



## $\Sigma_{ij} = \Sigma_{ij}^{\exp} \delta_{ij} + \Sigma_{ij}^{\text{th}}$

$$\vec{y}_{th}$$
)<sup>T</sup> $\Sigma^{-1}$  $(\vec{y}_{exp} - \vec{y}_{th}) = e^{-d_M(\vec{a})}$ 

# $d_{M}(\vec{a}) = \left(\vec{y}_{exp} - \vec{y}_{th}\right)^{T} \Sigma^{-1} \left(\vec{y}_{exp} - \vec{y}_{th}\right).$

Correlated version of  $\chi^2$ 









#### In this process, we have introduced two new parameters: $\bar{c}$ and $\Lambda_{h}$ .



This changes the posterior we need to find:



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This changes the posterior we need to find:  $pr(\vec{a}, \vec{c}^2, \Lambda_b | \vec{y}_{exp}, I) \propto pr(\vec{y}_{exp} | \vec{a}, \Sigma, I) pr(\vec{a} | I)$ 

Total posterior



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We can find a closed form of  $pr(\bar{c}^2 | \Lambda_b, \bar{a}, I)$  and  $pr(\Lambda_b | \bar{a}, I)$ .









standard choice of prior for an unknown variance:  $\bar{c}^2 \sim \chi$ 



Since we had  $c_n | \bar{c} \sim \mathcal{N}(0, \bar{c}^2)$ , where  $\bar{c}^2$  is a population variance, we make the

$$\chi^{-2}\left(\nu_0,\tau_0^2\right)$$





Since we had  $c_n | \bar{c} \sim \mathcal{N}(0, \bar{c}^2)$ , where  $\bar{c}^2$  is a population variance, we make the standard choice of prior for an unknown variance:  $\bar{c}^2 \sim \chi$ 

This yields a conjugate posterior



$$\chi^{-2}\left(\nu_0,\tau_0^2\right)$$

 $\operatorname{pr}(\mathbf{\bar{c}}^2 | \mathbf{I}) \sim \chi^{-2}(\nu_0, \tau_0^2) \iff \operatorname{pr}(\mathbf{\bar{c}}^2 | \mathbf{\bar{a}}, \Lambda_{\mathbf{b}}, \mathbf{I}) \sim \chi^{-2}(\nu, \tau^2(\mathbf{\bar{a}}, \Lambda_{\mathbf{b}})).$ 





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Where we have hyperparameters:



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Where we have hyperparameters:

$$\nu = \nu_0 + N_{\rm obs} n_{\rm ord}$$



$$\chi^{-2}\left(\nu_0,\tau_0^2\right)$$

$$r(\mathbf{\bar{c}}^2 | \mathbf{\bar{a}}, \Lambda_{\mathbf{b}}, \mathbf{I}) \sim \chi^{-2} \left( \nu, \tau^2(\mathbf{\bar{a}}, \Lambda_{\mathbf{b}}) \right).$$

ders, degrees of freedom





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Where we have hyperparameters:

$$\nu = \nu_0 + N_{\text{obs}} n_{\text{orders}}, \text{ degrees of freedom}$$
$$\tau^2 \left(\vec{a}, \Lambda_b\right) = \frac{1}{\nu} \left( \nu_0 \tau_0 + \sum_{i,n} c_{n,i}^2(\vec{a}, \Lambda_b) \right), \text{ scale}$$



$$\chi^{-2}\left(\nu_0,\tau_0^2\right)$$

$$r(\mathbf{\bar{c}}^2 | \mathbf{\bar{a}}, \Lambda_{\mathbf{b}}, \mathbf{I}) \sim \chi^{-2} \left( \nu, \tau^2(\mathbf{\bar{a}}, \Lambda_{\mathbf{b}}) \right).$$





Since we had  $c_n | \bar{c} \sim \mathcal{N}(0, \bar{c}^2)$ , where  $\bar{c}^2$  is a population variance, we make the standard choice of prior for an unknown variance:  $\bar{c}^2 \sim c$ 

This yields a conjugate posterior  $pr(\bar{c}^2 | I) \sim$ 

Where we have

$$\begin{aligned} & \chi^{-2}(\nu_0, \tau_0^2) \iff \operatorname{pr}(\bar{\mathbf{c}}^2 \,|\, \bar{\mathbf{a}}, \Lambda_b, \mathbf{I}) \sim \chi^{-2}\left(\nu, \tau^2(\bar{\mathbf{a}}, \Lambda_b)\right). \end{aligned}$$
hyperparameters:
$$\nu = \nu_0 + N_{\mathrm{obs}} n_{\mathrm{orders}}, \text{ degrees of freedom} \qquad c_{n,i} = \frac{y_i^n - y_i^{(n-1)}}{y_{\mathrm{ref},i} \, Q_i^n} \\ \tau^2\left(\vec{a}, \Lambda_b\right) = \frac{1}{\nu} \left(\nu_0 \tau_0 + \sum_{i,n} c_{n,i}^2(\vec{a}, \Lambda_b)\right), \text{ scale} \\ \overset{*}{}_{\mathrm{J}} \text{ A. Melendez et. al. Phys. Rev. C 100, 044} \end{aligned}$$



$$\chi^{-2}\left(\nu_0,\tau_0^2\right)$$





### Posterior for $\Lambda_b$





### Posterior for $\Lambda_b$

Our posterior for the breakdown scale also uses these hyperparameters:





### Posterior for $\Lambda_h$

#### Our posterior for the breakdown scale also uses these hyperparameters:



 $pr(\Lambda_{b} | \vec{a}, I) \propto \frac{pr(\Lambda_{b} | I)}{\tau^{\nu} \prod_{n,i} \left(\frac{p_{i}}{\Lambda_{b}}\right)^{n}}$ 



### Posterior for $\Lambda_h$

Our posterior for the breakdown scale also uses these hyperparameters:

 $pr(\Lambda_b | \vec{a}, I)$ 

This posterior needs to be numerically normalized as the normalization constant is dependent on  $\vec{a}$ .



$$\propto \frac{\text{pr}(\Lambda_{b} | \mathbf{I})}{\tau^{\nu} \prod_{n,i} \left(\frac{\mathbf{p}_{i}}{\Lambda_{b}}\right)^{n}}$$



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Our posterior for the breakdown scale also uses these hyperparameters:

 $pr(\Lambda_{\rm b} | \vec{a}, I)$ 

This posterior needs to be numerically normalized as the normalization constant is dependent on  $\vec{a}$ .

With all our components, we can estimate our parameters.



$$\propto \frac{\operatorname{pr}(\Lambda_{b} | \mathbf{I})}{\tau^{\nu} \prod_{n,i} \left(\frac{\mathbf{p}_{i}}{\Lambda_{b}}\right)^{n}}$$



### Outline

- Motivation/Bayesian
  BUQEYE Formalism
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### **-T Model Calibration**







#### Hidden in the *I* in our distributions is the choice of interaction.





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Hidden in the I in our distributions is the choice of interaction.

This includes: Degrees of freedom





Hidden in the I in our distributions is the choice of interaction.

- Degrees of freedom
- Power counting





Hidden in the I in our distributions is the choice of interaction.

- Degrees of freedom
- Power counting
- Representation





Hidden in the I in our distributions is the choice of interaction.

- Degrees of freedom
- Power counting
- Representation
- Regularization scheme





#### Pionless EFT





#### Pionless EFT

#### We are working in a "Weinberg-ized" pionless EFT.





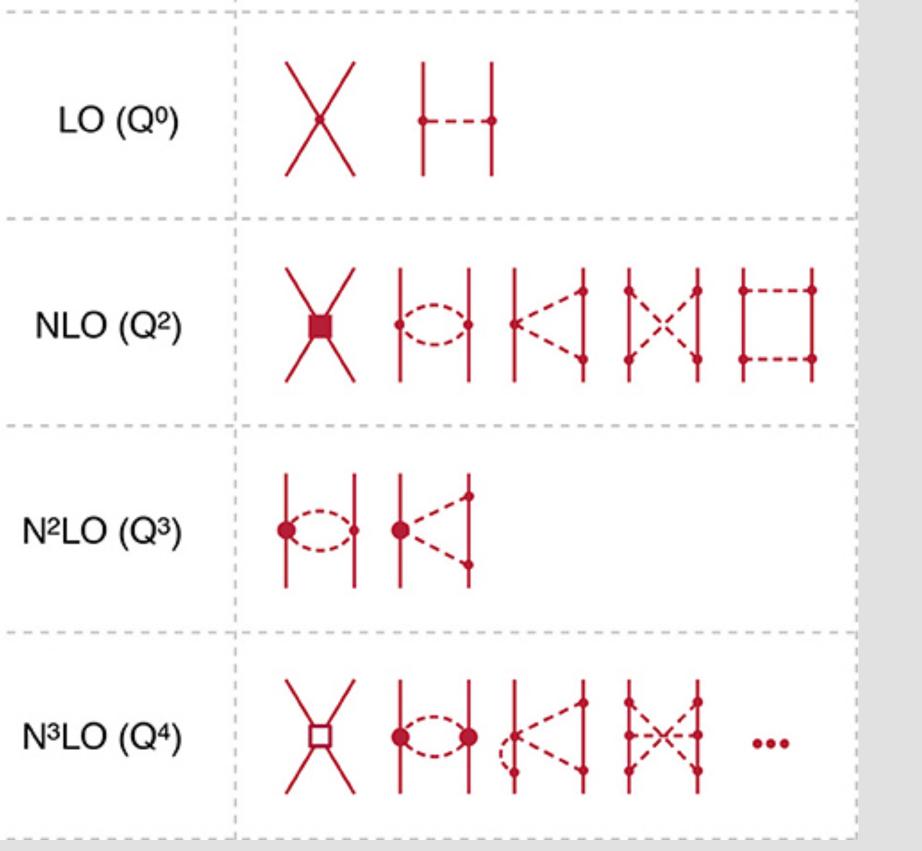
#### We are working in a "Weinberg-ized" pionless EFT.





# Pionless EFT We are working in a "Weinberg-ized" pionless EFT. Two-nucleon force LO (Qº) N²LO (Q³)

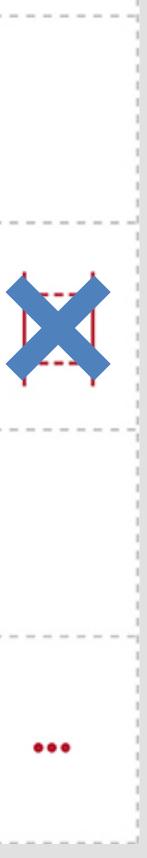






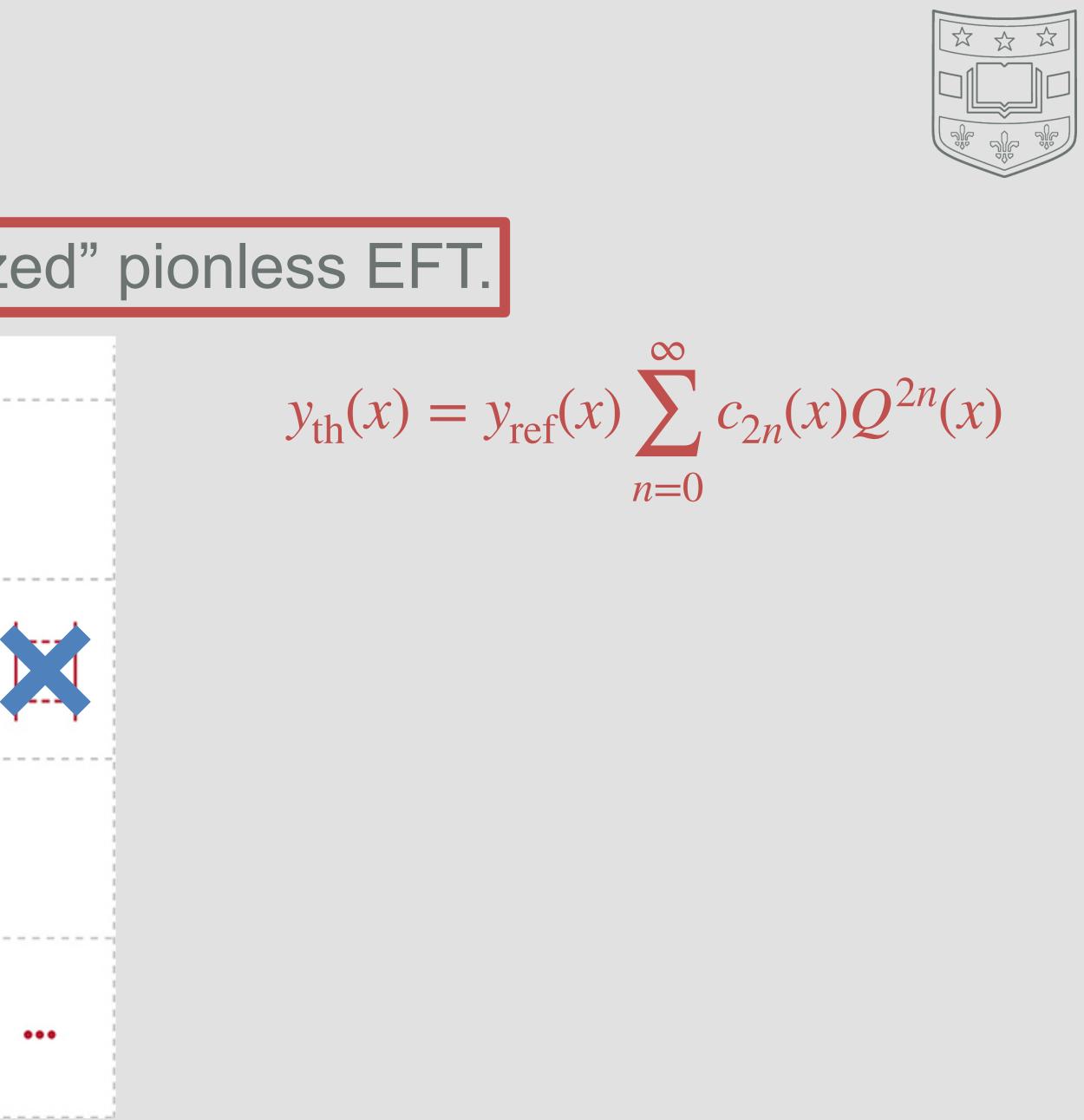
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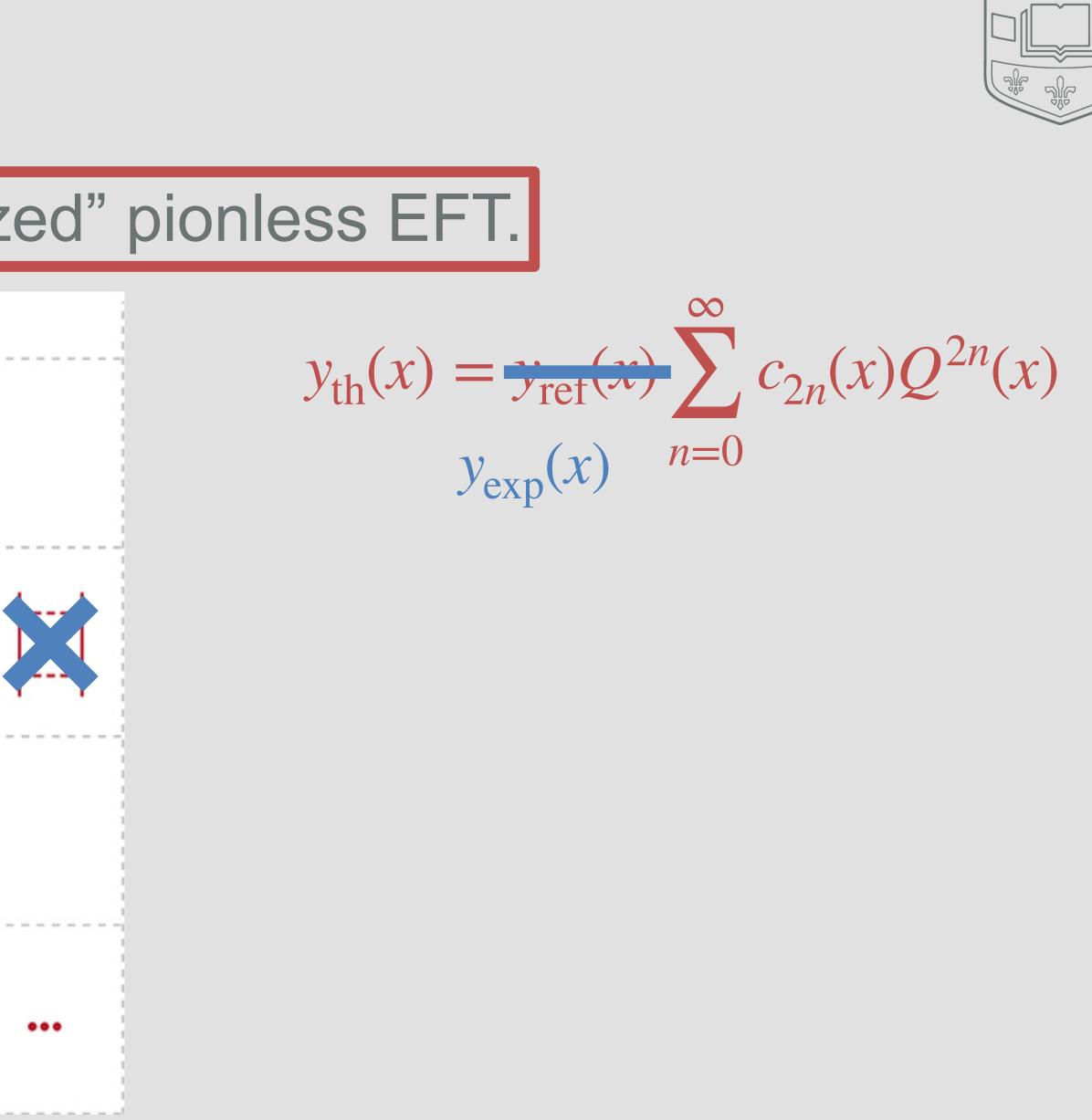




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# Pionless EFT We are working in a "Weinberg-ized" pionless EFT. Two-nucleon force N<sup>2</sup>LO (Q<sup>3</sup>)





We are working in a "Weinberg-ized" pionless EFT.

Our interaction takes the form:



# $y_{\text{th}}(x) = \underbrace{\sum_{\text{ref}(x)} \sum_{n=0}^{\infty} c_{2n}(x) Q^{2n}(x)}_{y_{\text{exp}}(x) n=0}$



We are working in a "Weinberg-iz

Our interaction takes the form:

zed" pionless EFT.  

$$y_{th}(x) = \frac{y_{ref}(x)}{y_{ref}(x)} \sum_{n=0}^{\infty} c_{2n}(x)Q^{2n}(x)$$

$$y_{exp}(x) = \frac{y_{ref}(x)}{y_{exp}(x)} \sum_{n=0}^{\infty} c_{2n}(x)Q^{2n}(x)$$

 $v_{\rm LO} = C_S + C_T \sigma_1 \cdot \sigma_2$ 



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 $v_{\rm LO} = C_{\rm S} +$ 

 $v_{\text{NLO}}^{\text{CI}}(\vec{k}, \vec{K}) = C_1 k^2 + C_2 k^2 \sigma_1 \cdot \sigma_2 + C_3 S_{12}(k) + C_4 k^2 \tau_1 \cdot \tau_2$ 

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$$y_{\text{th}}(x) = \underbrace{y_{\text{ref}}(x)}_{y_{\text{exp}}(x)} \sum_{n=0}^{\infty} c_{2n}(x) Q^{2n}(x)$$

$$+ C_T \sigma_1 \cdot \sigma_2$$

 $+iC_5\vec{S}\cdot(\vec{K}\times\vec{k})+C_6k^2\tau_1\cdot\tau_2\sigma_1\cdot\sigma_2+C_7S_{12}(k)\tau_2\cdot\tau_2$ 



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ed" pionless EFT.  

$$y_{th}(x) = \frac{y_{ref}(x)}{y_{ref}(x)} \sum_{n=0}^{\infty} c_{2n}(x) Q^{2n}(x)$$

$$y_{exp}(x)$$

 $v_{\rm LO} = C_{\rm S} + C_T \sigma_1 \cdot \sigma_2$ 

 $+iC_5\vec{S}\cdot(\vec{K}\times\vec{k})+C_6k^2\tau_1\cdot\tau_2\sigma_1\cdot\sigma_2+C_7S_{12}(k)\tau_2\cdot\tau_2$  $v_{\text{NLO}}^{\text{CD}} = C_0^{\text{IT}} T_{12} + C_0^{\text{IV}} (\tau_{1z} + \tau_{2z})$ 







be local in coordinate space (for QMC).



## To use these interactions, they must be regularized in some fashion and must



To use these interactions, they must be regularized in some fashion and must be local in coordinate space (for QMC).

We employ a Gaussian cutoff in coordinate space, which smears  $\delta$ -functions upon Fourier transformation



 $f(r) = \frac{1}{\pi^{3/2} R_s^3} e^{-\left(\frac{r}{R_s}\right)^2}$ 



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 $f(r) = \frac{1}{\pi^{3/2} R_s^3} e^{-\left(\frac{r}{R_s}\right)^2}$ We choose  $R_s \in [1.5, 2.0, 2.5]$  fm which are  $\sim \frac{400}{R_s}$  MeV in momentum

space.







#### To estimate all of these parameters, we need data to calibrate to:



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Our choice of data is the pp and np Granada database (251 differential cross sections, 133 total cross sections, 4 polarized cross sections) up to 5 MeV + deuteron binding energy + nn scattering length.



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Our choice of data is the pp and np Granada database (251 differential cross sections, 133 total cross sections, 4 polarized cross sections) up to 5 MeV + deuteron binding energy + nn scattering length.

We then use Markov Chain Monte Carlo (MCMC) to sample the to estimate  $\bar{c}$  and  $\Lambda_{h}$ .



posteriors at LO  $(Q^0)$ , NLO  $(Q^2)$ , and N3LO  $(Q^4)$ , allowing for the order-by-order convergence analysis for LO $\rightarrow$ NLO and NLO $\rightarrow$ N3LO

# **Prior Choices** • $pr(\vec{a} | I) \sim \mathcal{N}\left(\vec{a}_{p.s.}^{MAP}, \vec{10}^2\right)$ • $pr(\Lambda_b | I) \sim \mathcal{N}(500 \text{ MeV}, 1000^2 \text{ MeV}^2)$ • $\operatorname{pr}(\bar{c}^2 | \mathbf{I}) \sim \chi^{-2}(\nu_0 = 1.5, \tau_0^2 = 1.5^2)$ • $r(x_i, x_j; \vec{l}) = e^{|p_i - p_j|/2l_p} e^{|\theta_i - \theta_j|/2l_\theta} \delta_{\text{type}_i, \text{type}_j}, \quad l_p = 0.3 \text{ MeV}, \ l_\theta = 20^\circ$ $p_{\rm soft} = \begin{cases} p_d \sim 45 \; {\rm MeV}/c, & {\rm for} \; np \; {\rm scattering} \\ 1/{}^1 a_{\rm pp} \sim 25 \; {\rm MeV}, & {\rm for} \; pp \; {\rm scattering} \; . \end{cases}$





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### **-T Model Calibration**

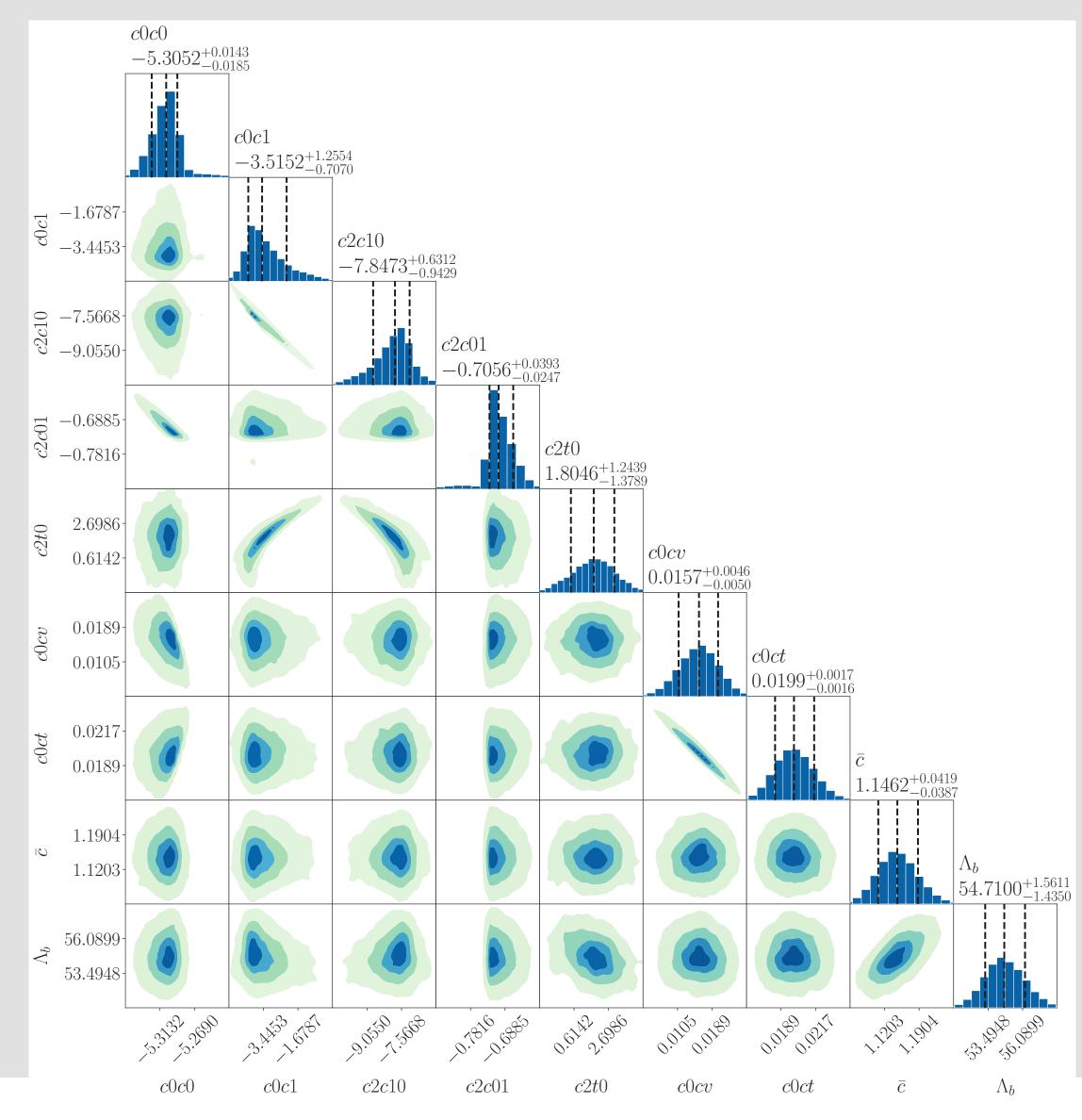








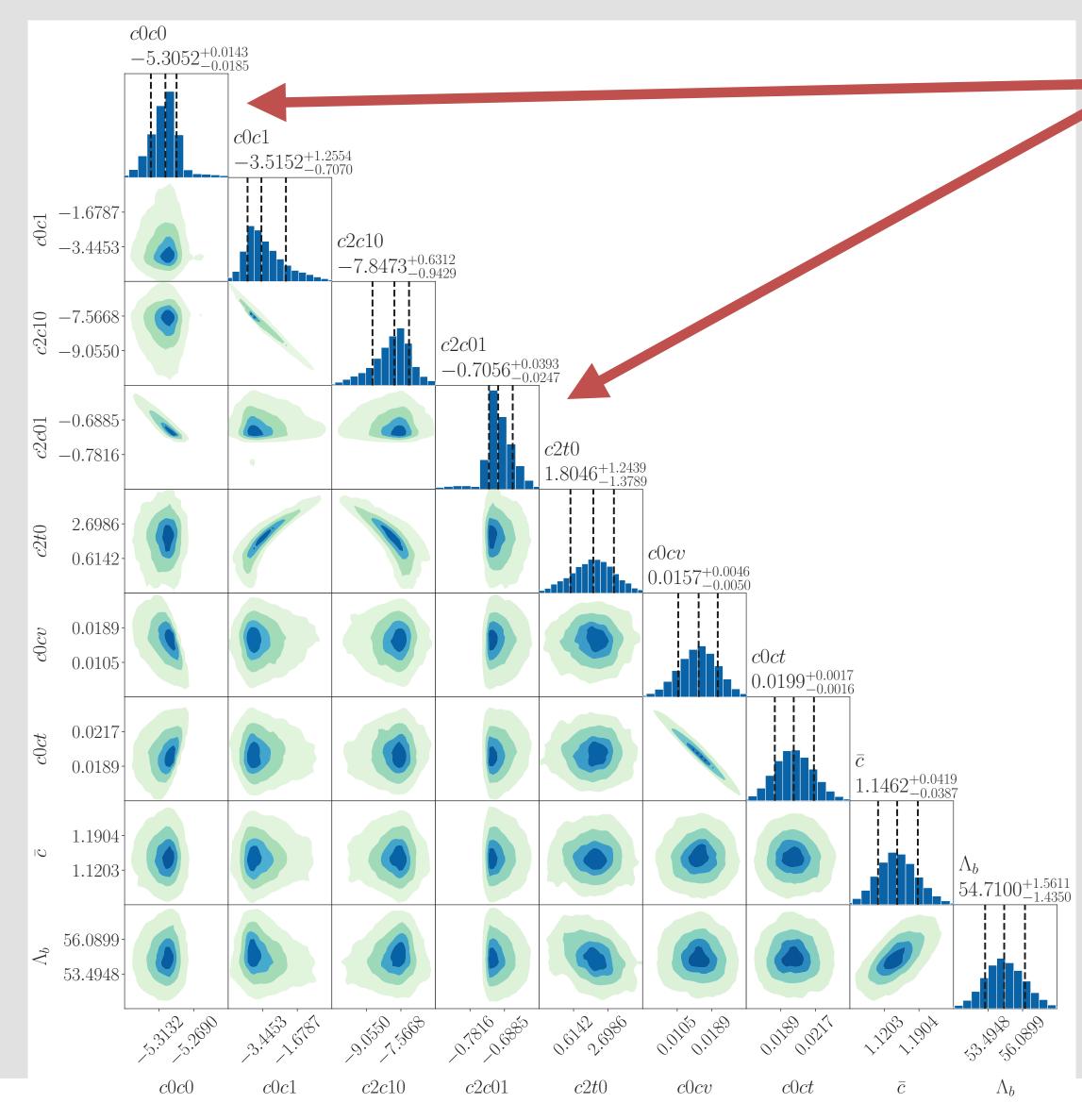




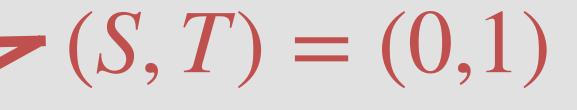








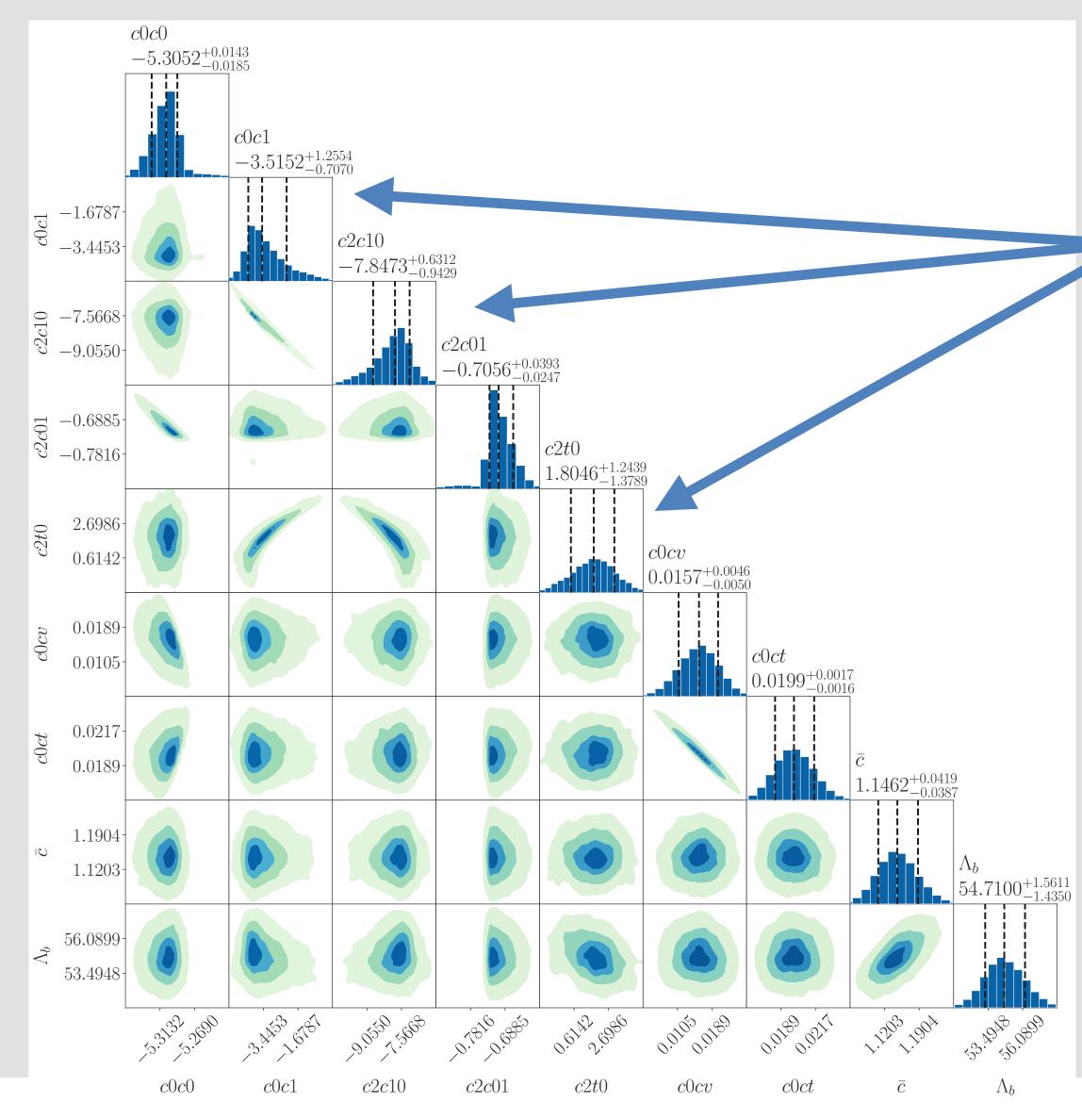




#### Figures:







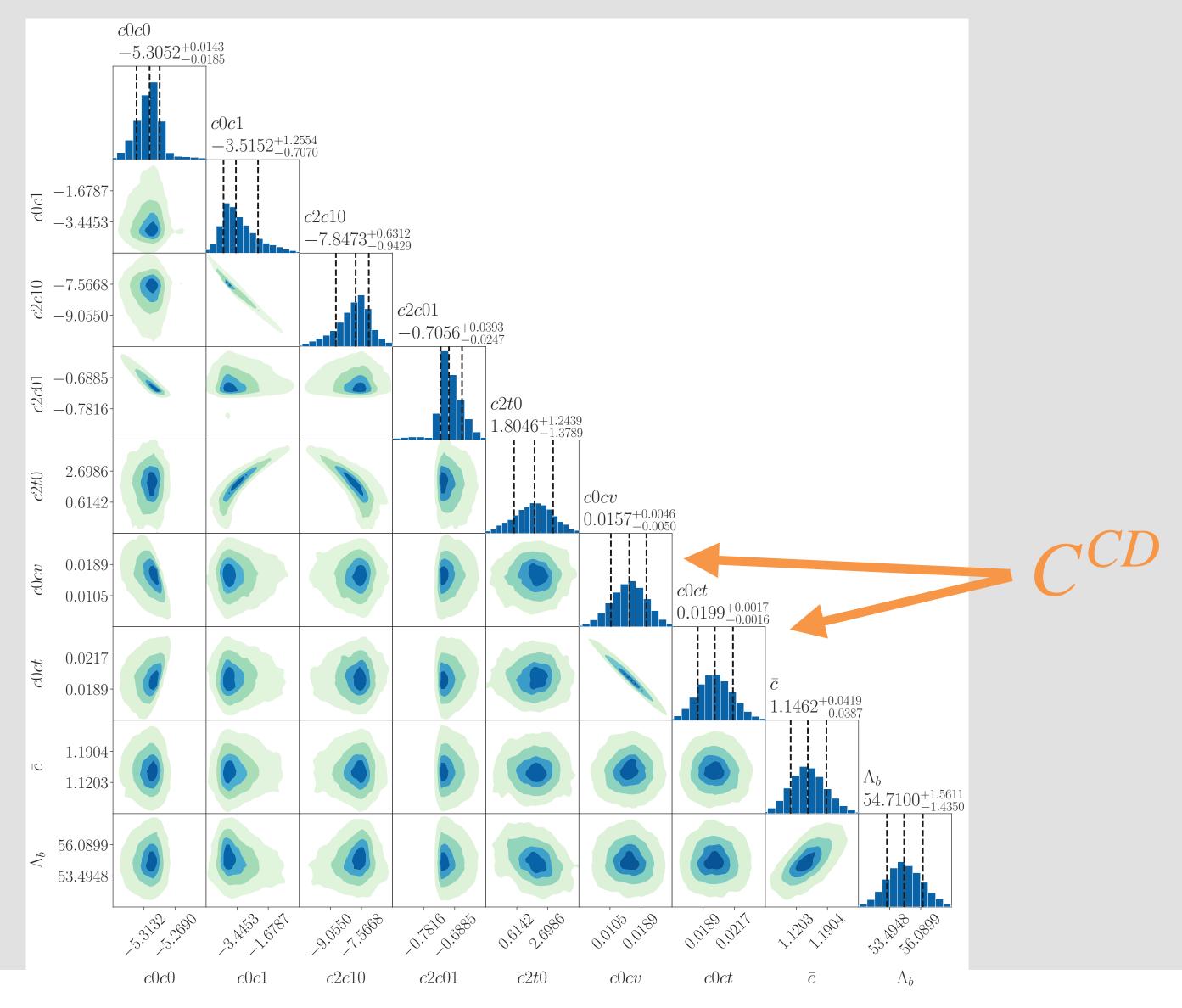


#### (S, T) = (1,0)









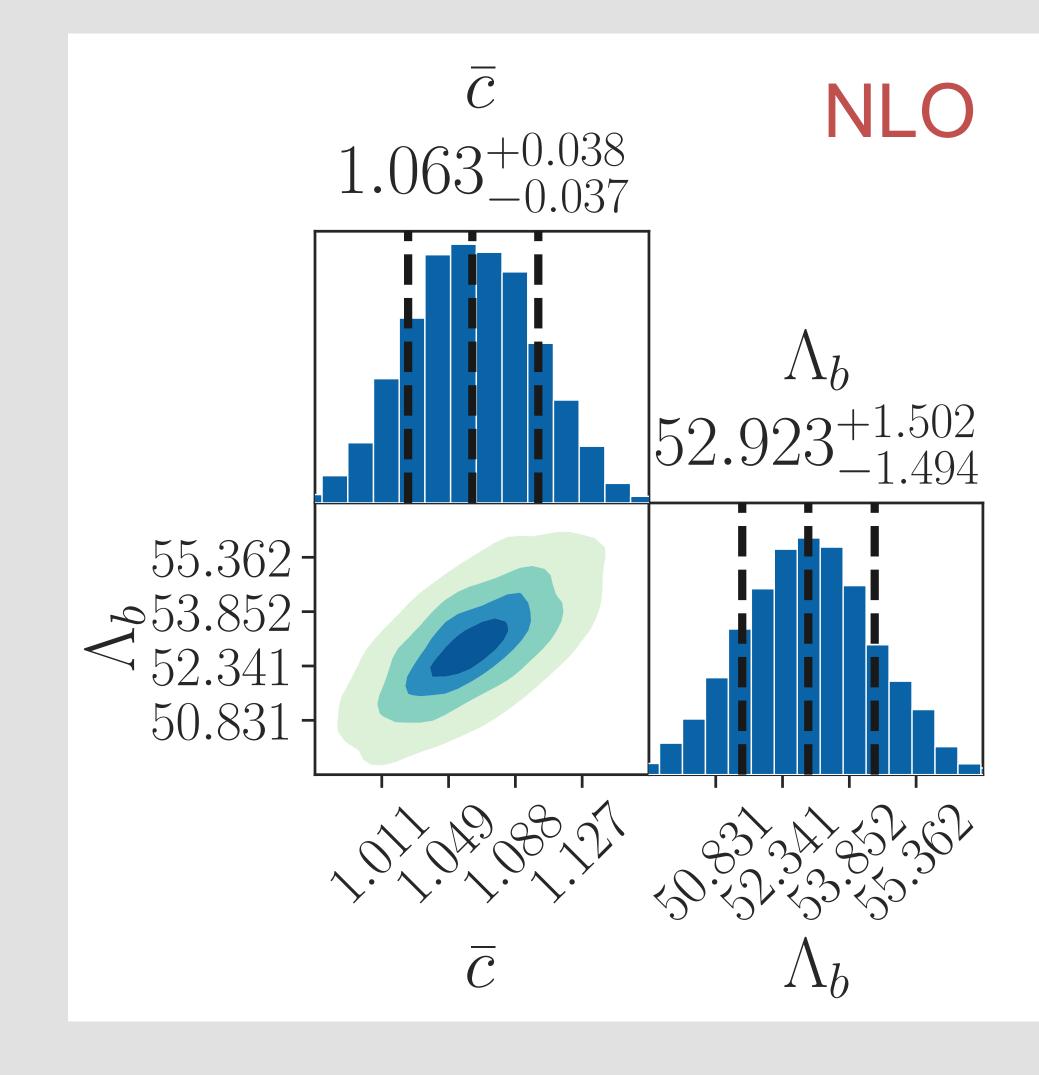




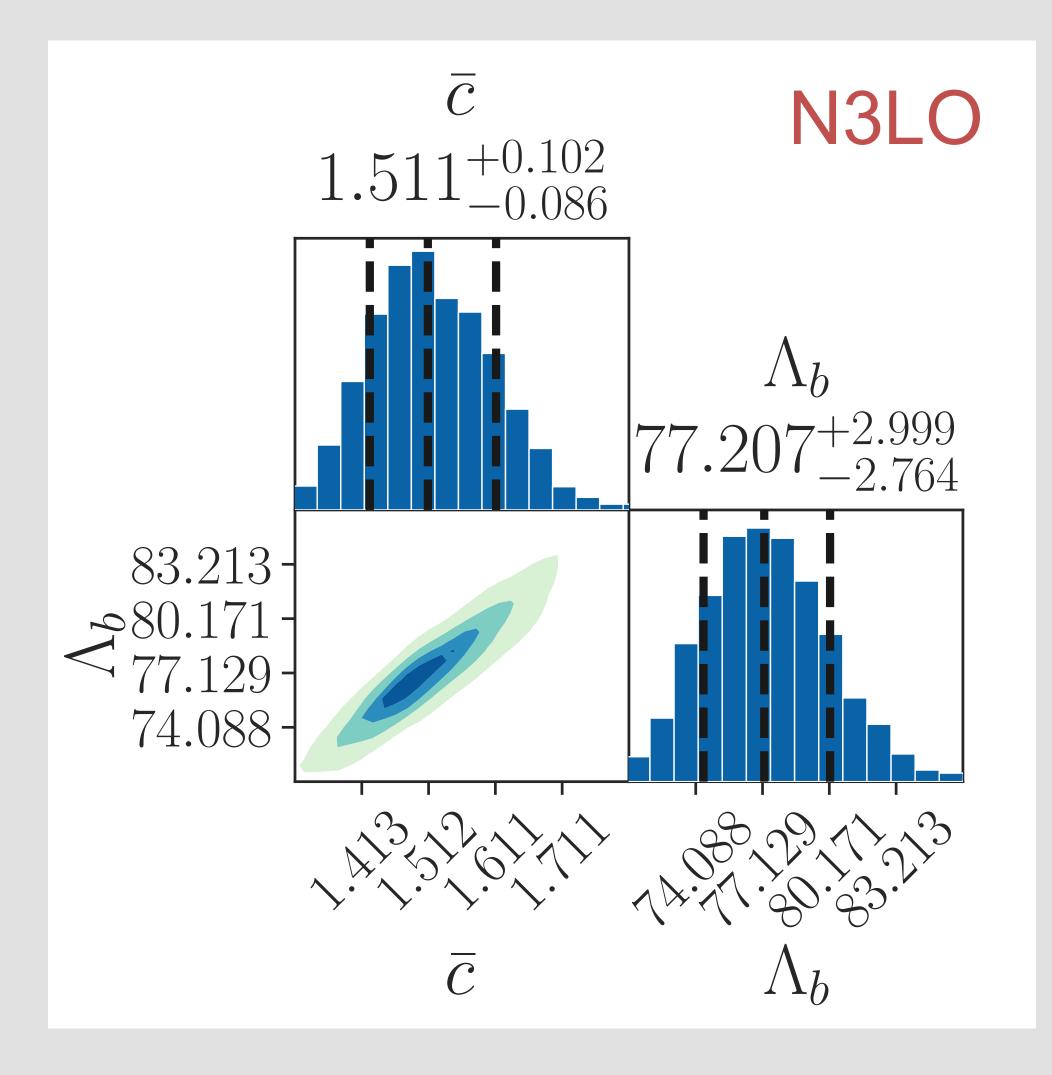




#### 2.5 fm $\bar{c}$ and $\Lambda_b$ Posteriors

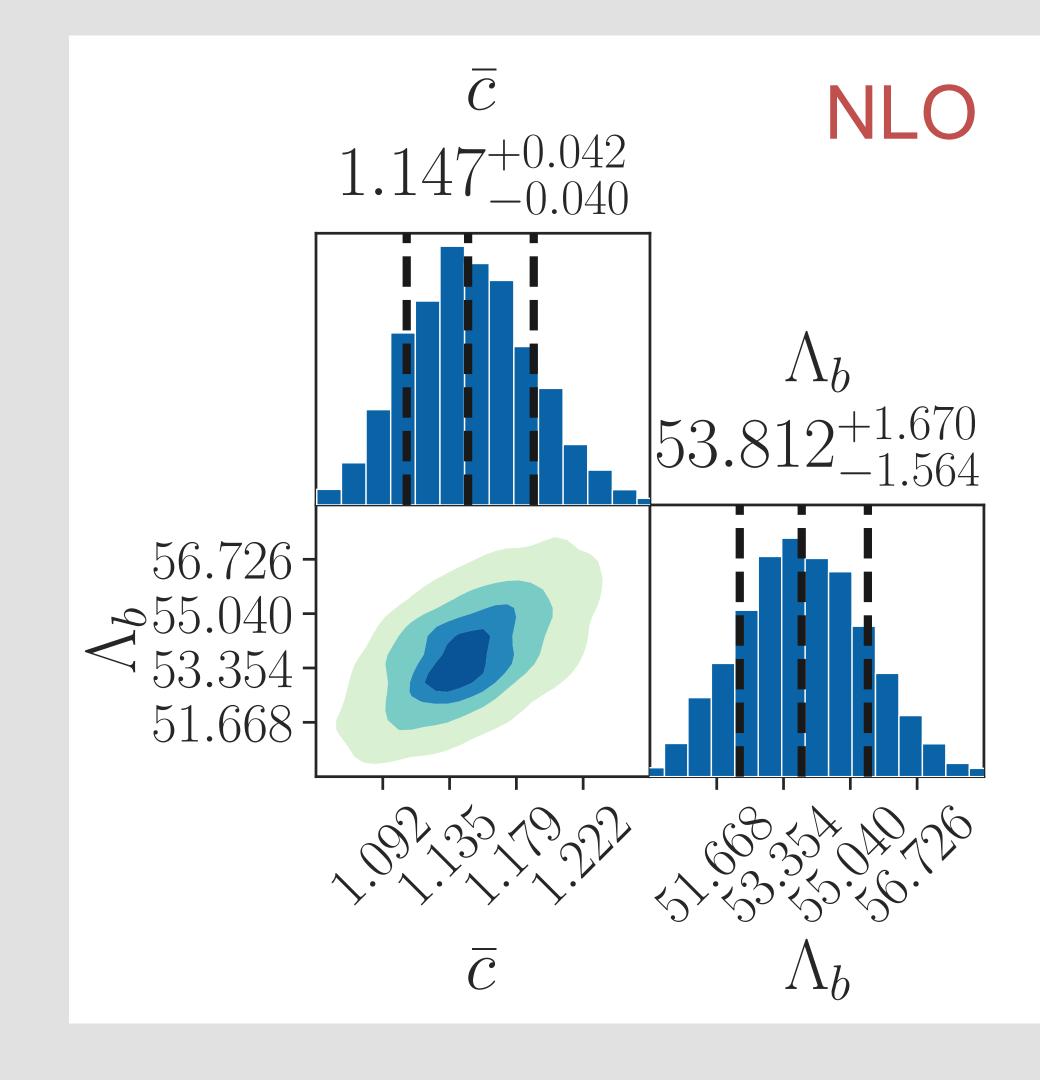




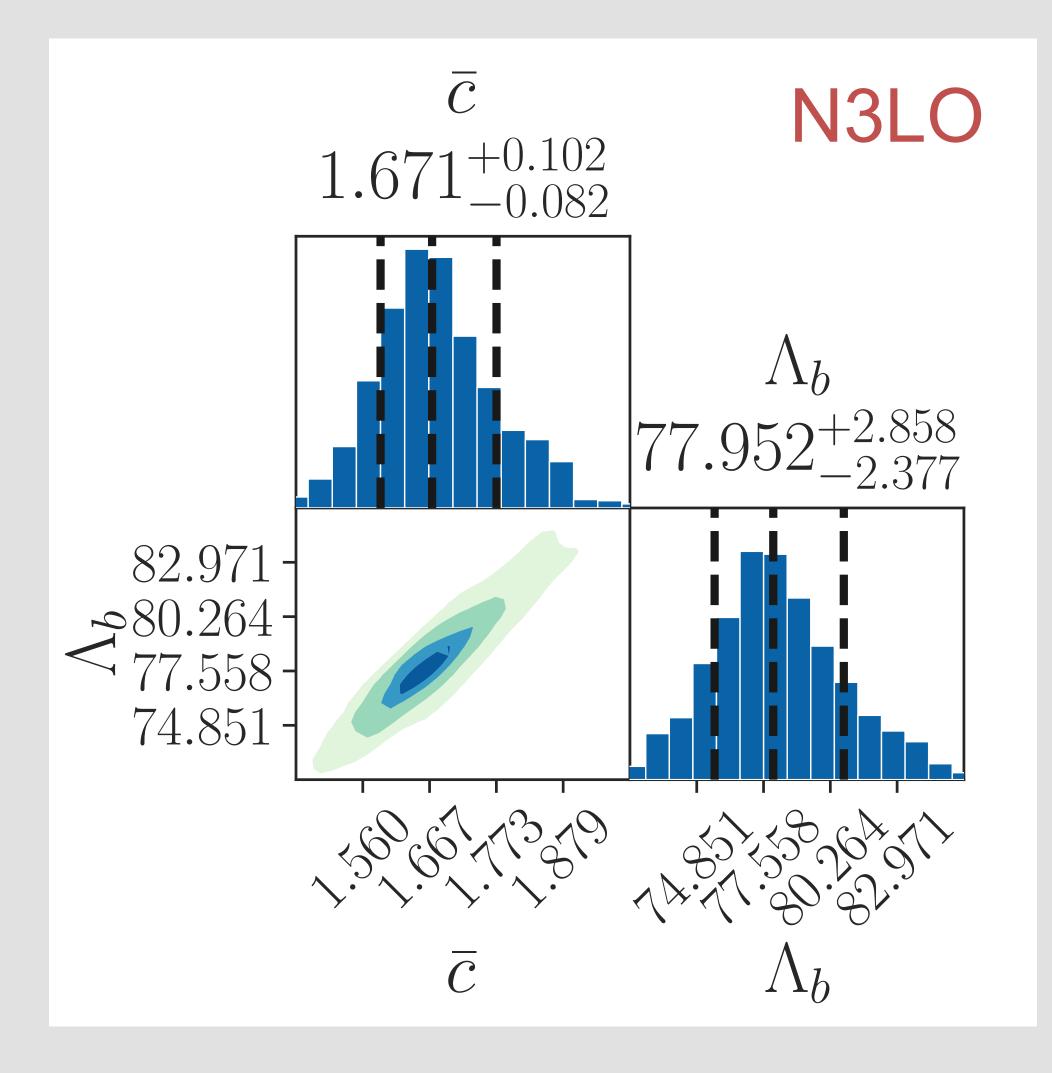




### 2.0 fm $\bar{c}$ and $\Lambda_b$ Posteriors

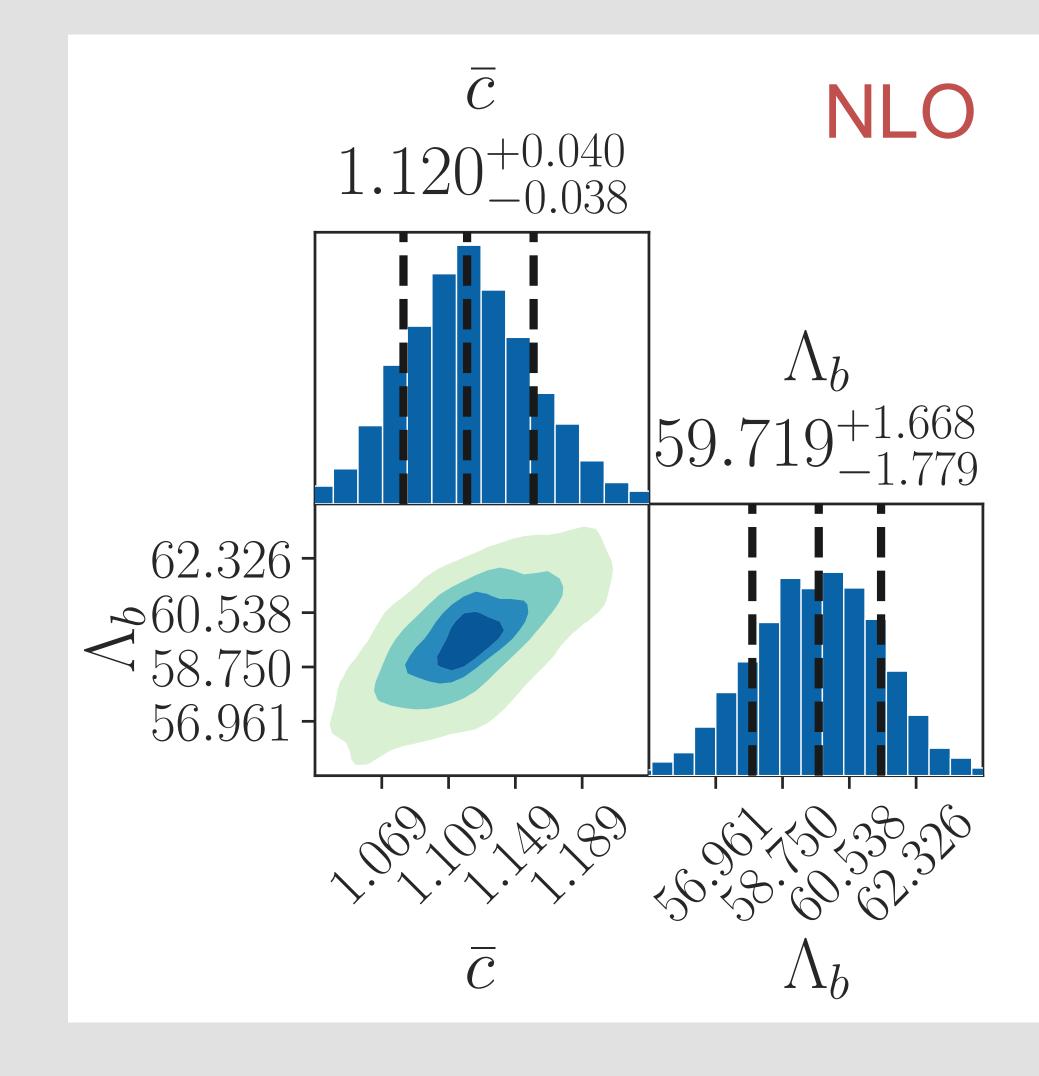




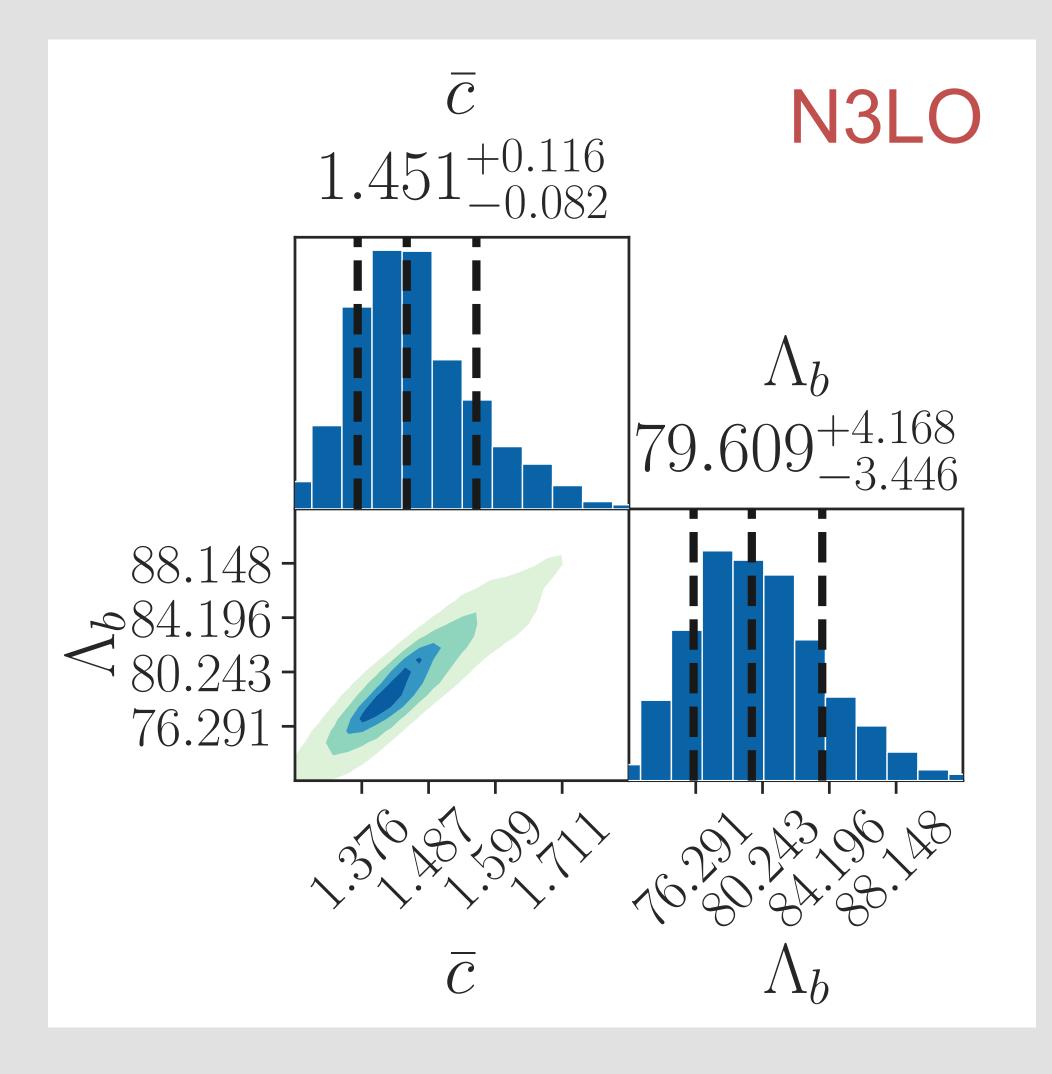




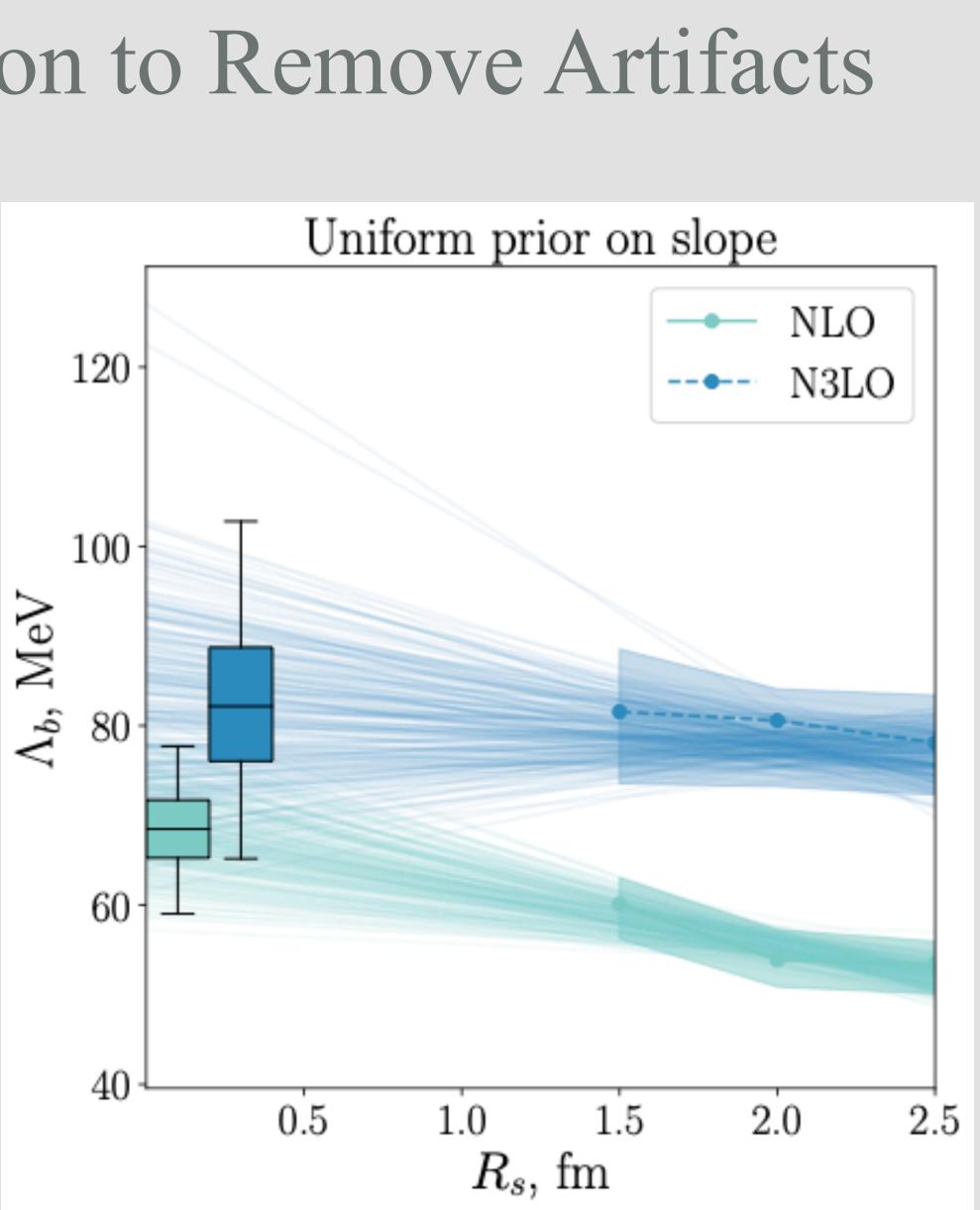
### 1.5 fm $\bar{c}$ and $\Lambda_b$ Posteriors



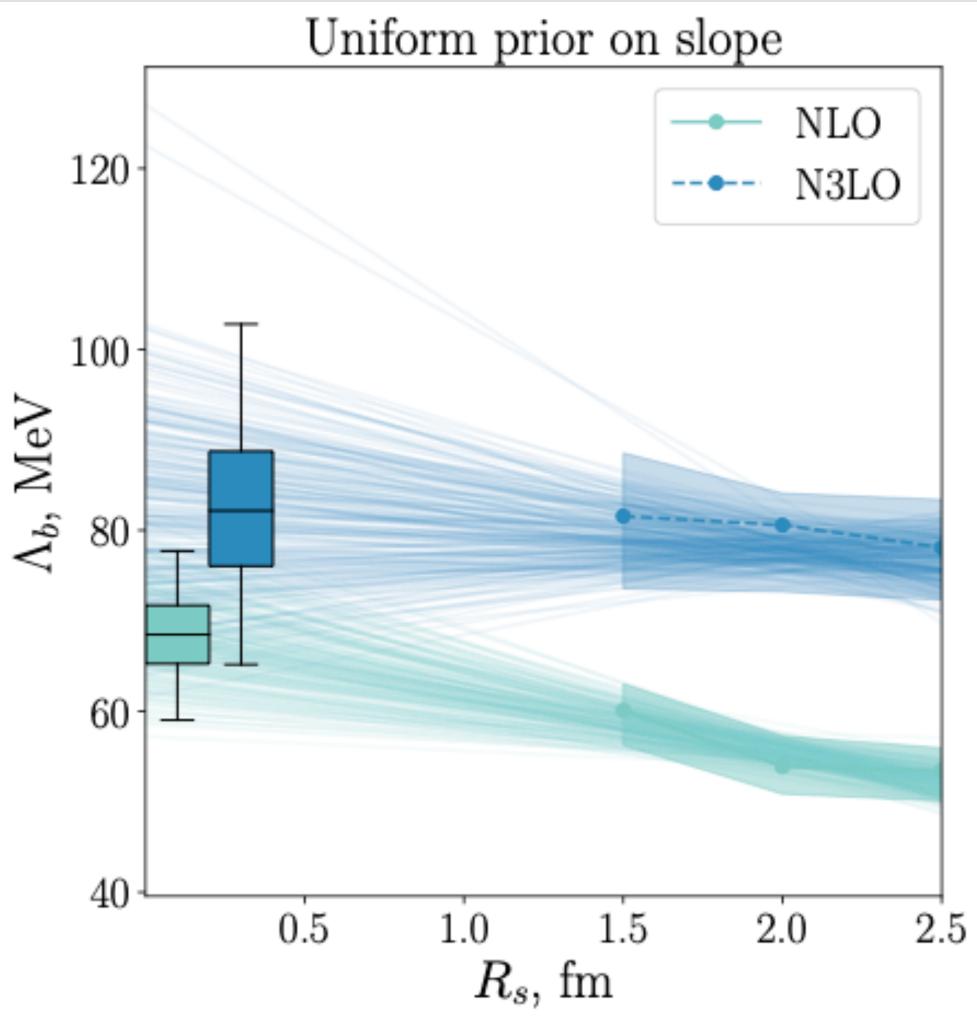






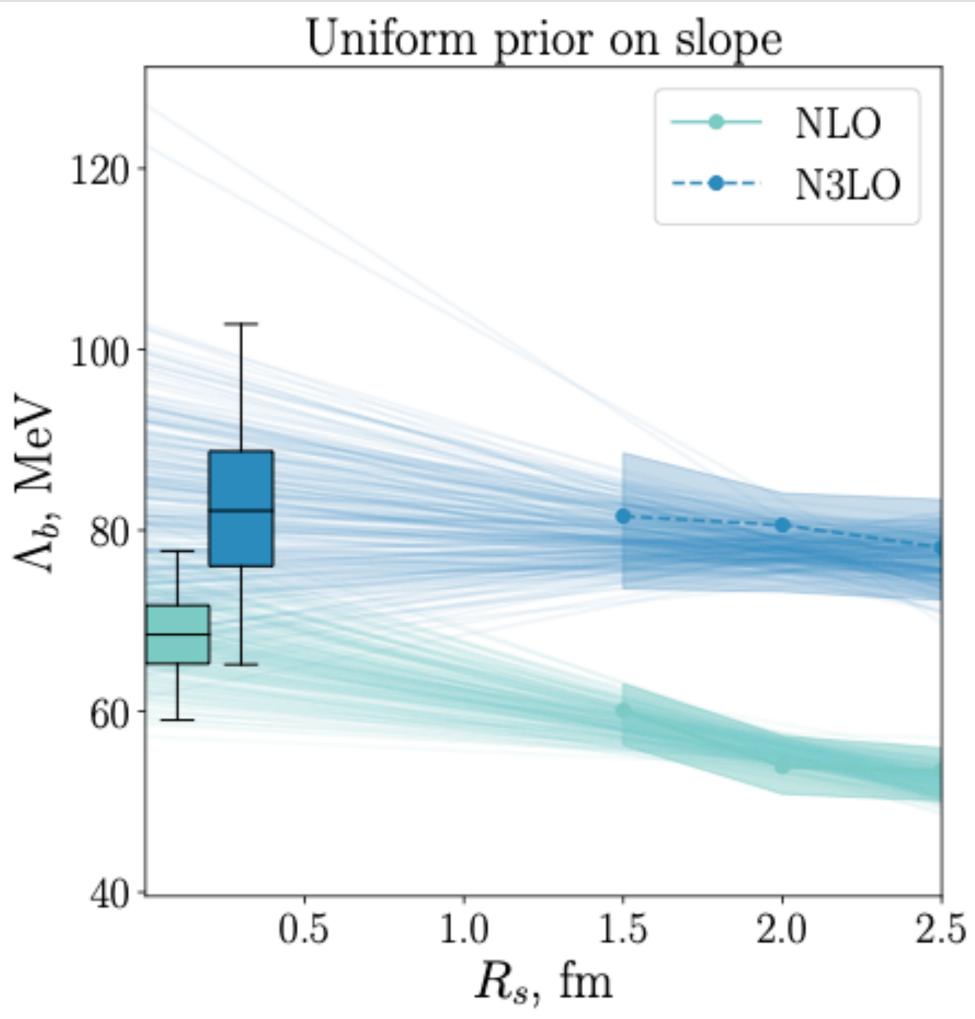








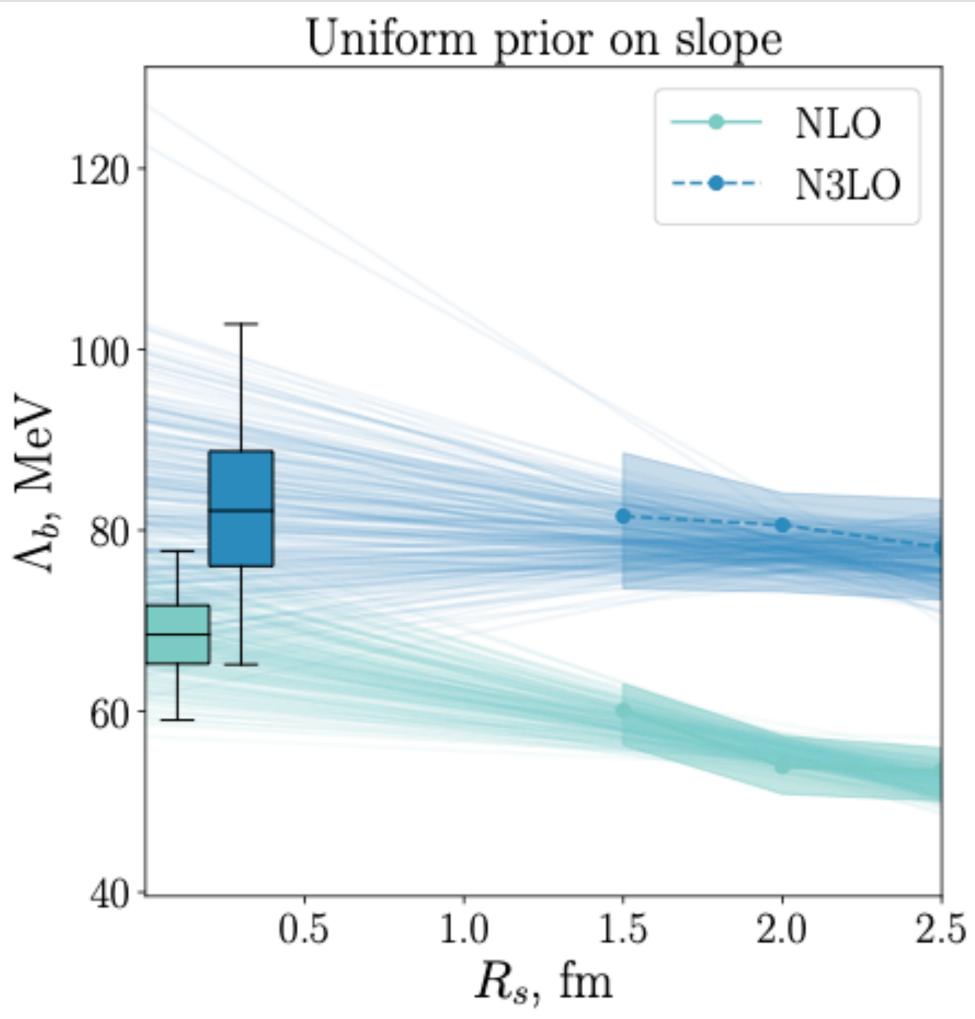
#### Why is there dependence on the order?





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• Power counting?

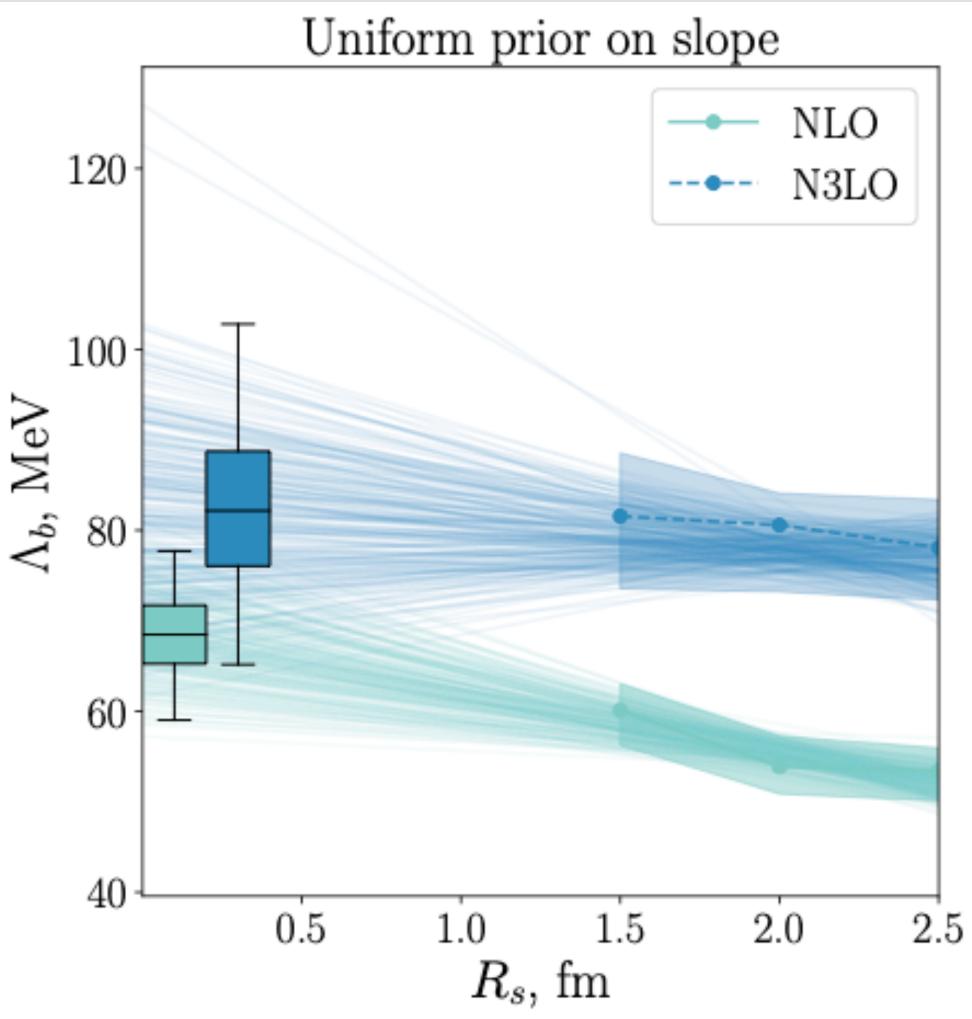




Why is there dependence on the order?

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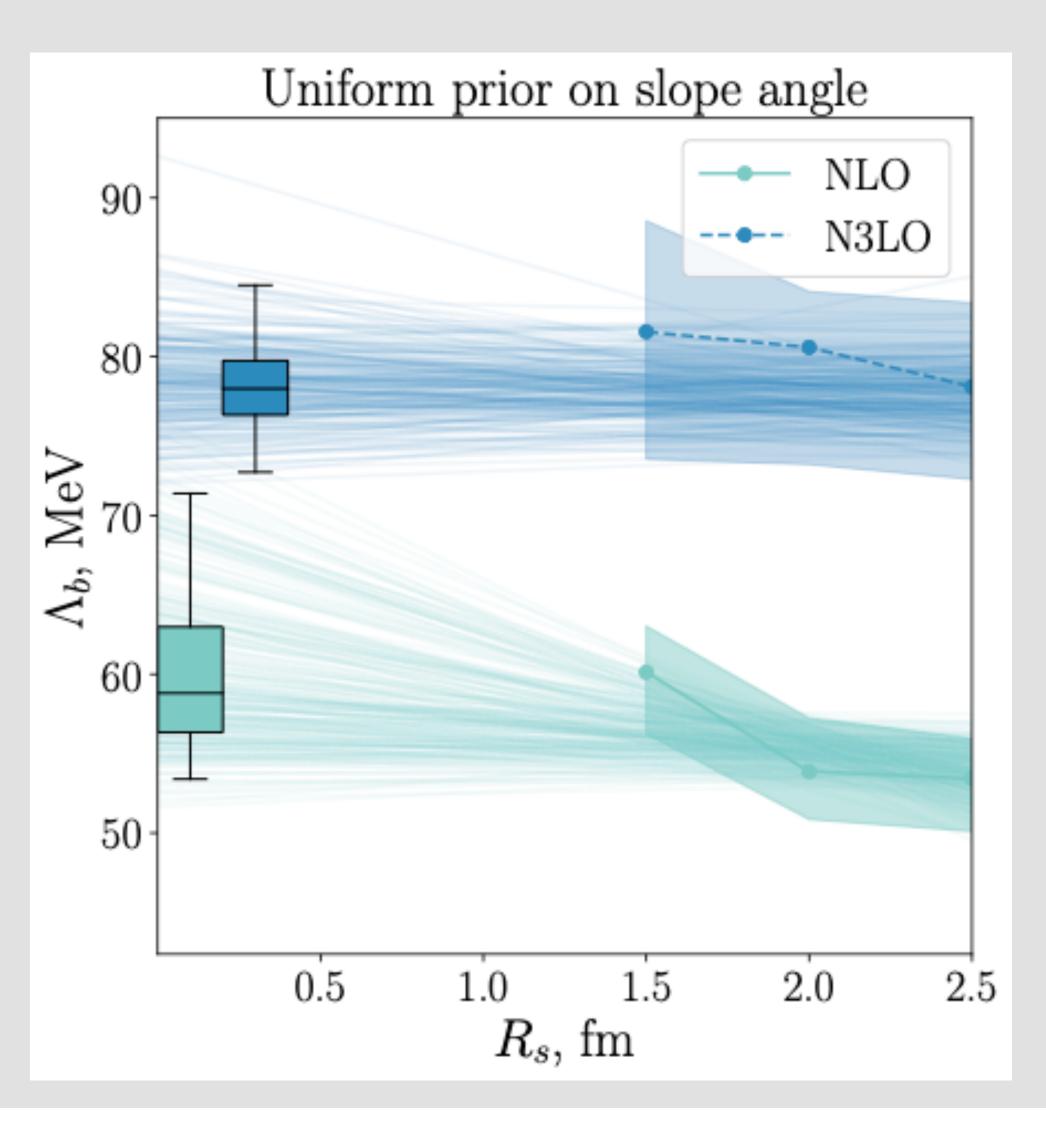


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- Power counting?
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### Proper choice of prior!!!





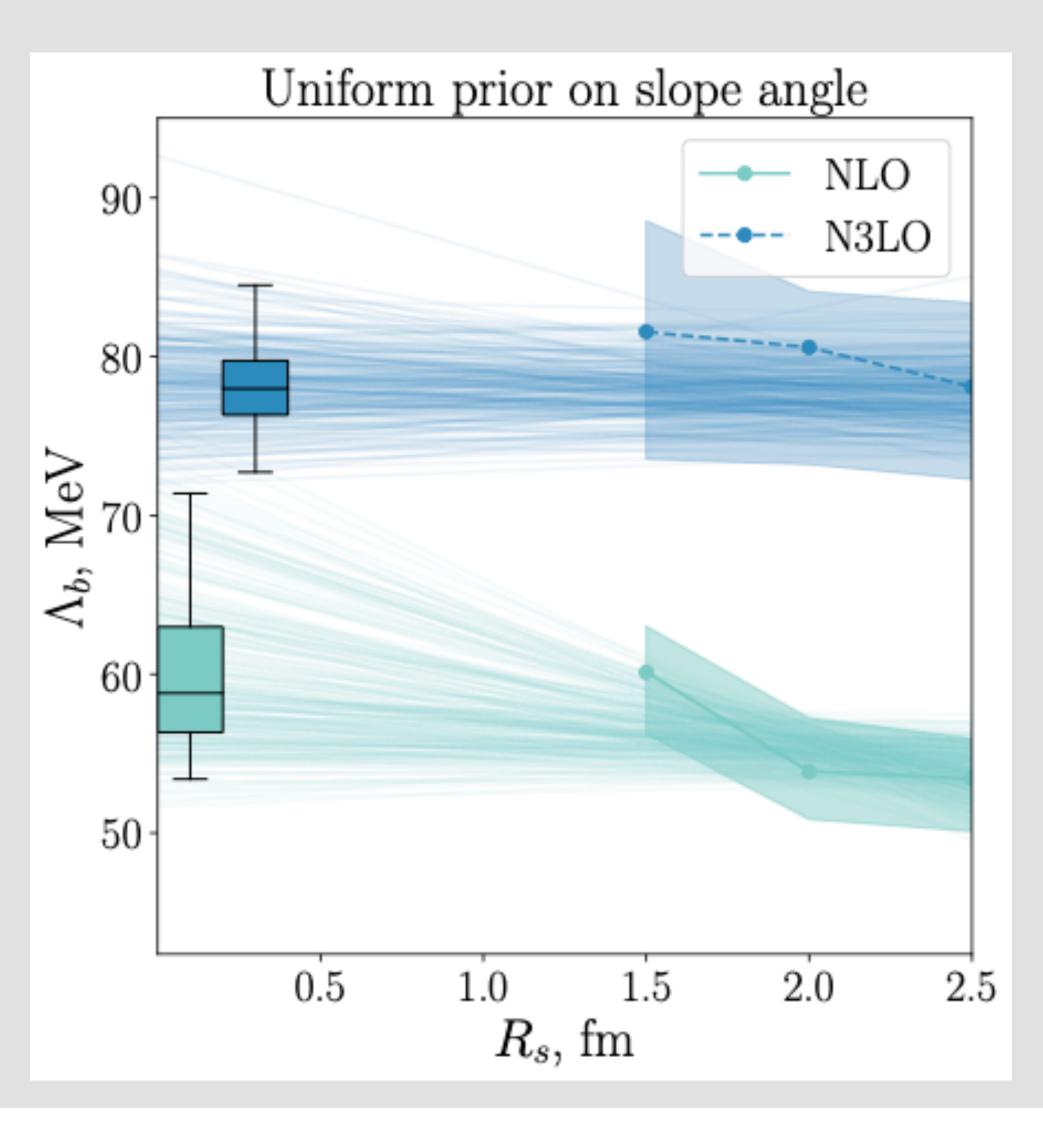
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### Unconstrained *p*-waves





#### Unconstrained *p*-waves

With  $\Lambda_h \sim 50$  MeV, the max lab energy is given by





## Unconstrained *p*-waves

With  $\Lambda_b \sim 50$  MeV, the max lab energy is given by

 $p_{c.m.}^{(max)} = - \Lambda_b$ 

F(max)

 $\Lambda_b$ 

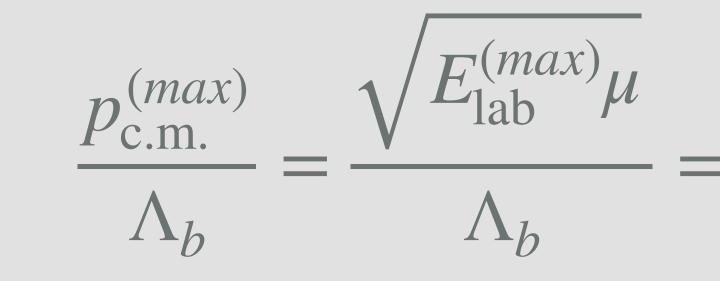


$$= 1 \Rightarrow E_{\text{lab}}^{(max)} = \frac{\Lambda_b^2}{\mu} \sim 5 \text{ MeV}$$



## Unconstrained *p*-waves

With  $\Lambda_h \sim 50$  MeV, the max lab energy is given by



MeV that constrains  ${}^{1}P_{1}$  and  ${}^{3}P_{0}$  channels which have poor experimental and theoretical constraints.



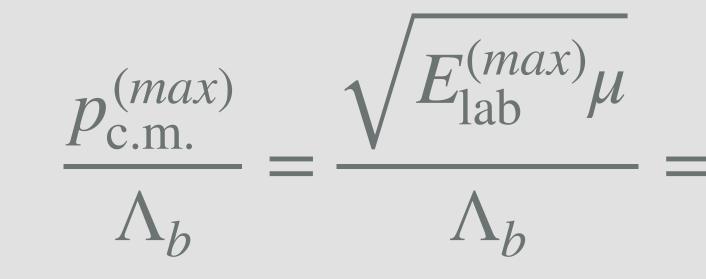
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The Granada database has 4 data (polarized cross sections) up to 5



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MeV that constrains  ${}^{1}P_{1}$  and  ${}^{3}P_{0}$  channels which have poor experimental and theoretical constraints.

The data is telling us that our "Weinberg-ized" pionless EFT is predominantly *s*-wave physics!



$$1 \Rightarrow E_{\text{lab}}^{(max)} = \frac{\Lambda_b^2}{\mu} \sim 5 \text{ MeV}$$

The Granada database has 4 data (polarized cross sections) up to 5







calculations.





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We calculate a posterior predictive distribution (p.p.d.) for the observables





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 $\operatorname{pr}(\vec{y}_{\mathrm{th}} | \vec{y}, \vec{x}, I) = d\vec{a} \, d\vec{c}^2 \, d\Lambda_k$ 



$$_{b} \mathcal{N}\left(\vec{y}_{th}, \Sigma_{th}\right) \operatorname{pr}(\vec{a}, \vec{c}^{2}, \Lambda_{b} | \vec{y}_{exp}, I)$$



calculations.

We calculate a posterior predictive distribution (p.p.d.) for the observables

which is done via sampling the posterior.



### We can now easily and rigorously propagate uncertainty to observable

# $\operatorname{pr}(\vec{y}_{\text{th}} | \vec{y}, \vec{x}, I) = \left[ d\vec{a} \, d\bar{c}^2 \, d\Lambda_b \, \mathcal{N}\left(\vec{y}_{\text{th}}, \Sigma_{\text{th}}\right) \operatorname{pr}(\vec{a}, \bar{c}^2, \Lambda_b | \vec{y}_{\text{exp}}, I) \right]$



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 $\operatorname{pr}(\vec{y}_{\mathrm{th}} | \vec{y}, \vec{x}, I) = \int d\vec{a} \, d\vec{c}^2 \, d\Lambda_k$ 

which is done via sampling the posterior.

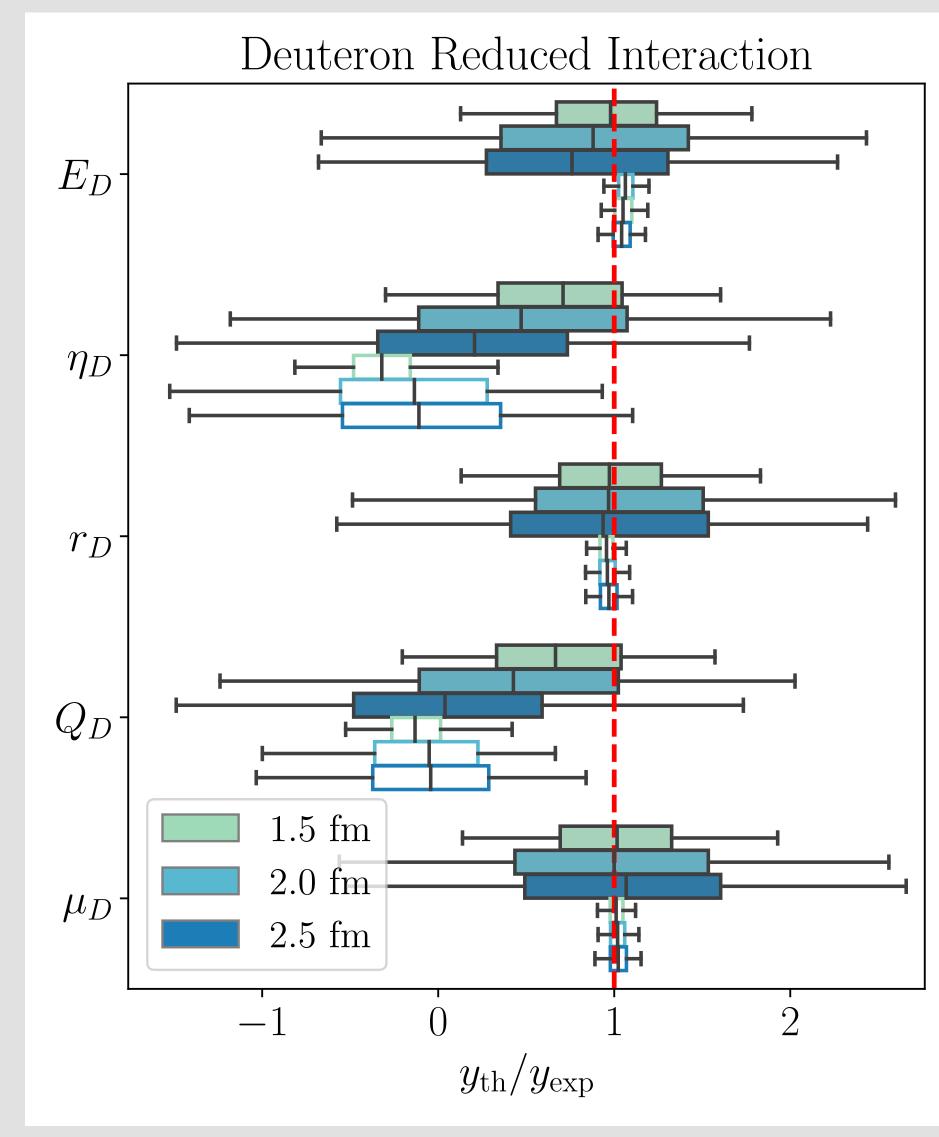
This can be done for any calculation of nuclear observables.



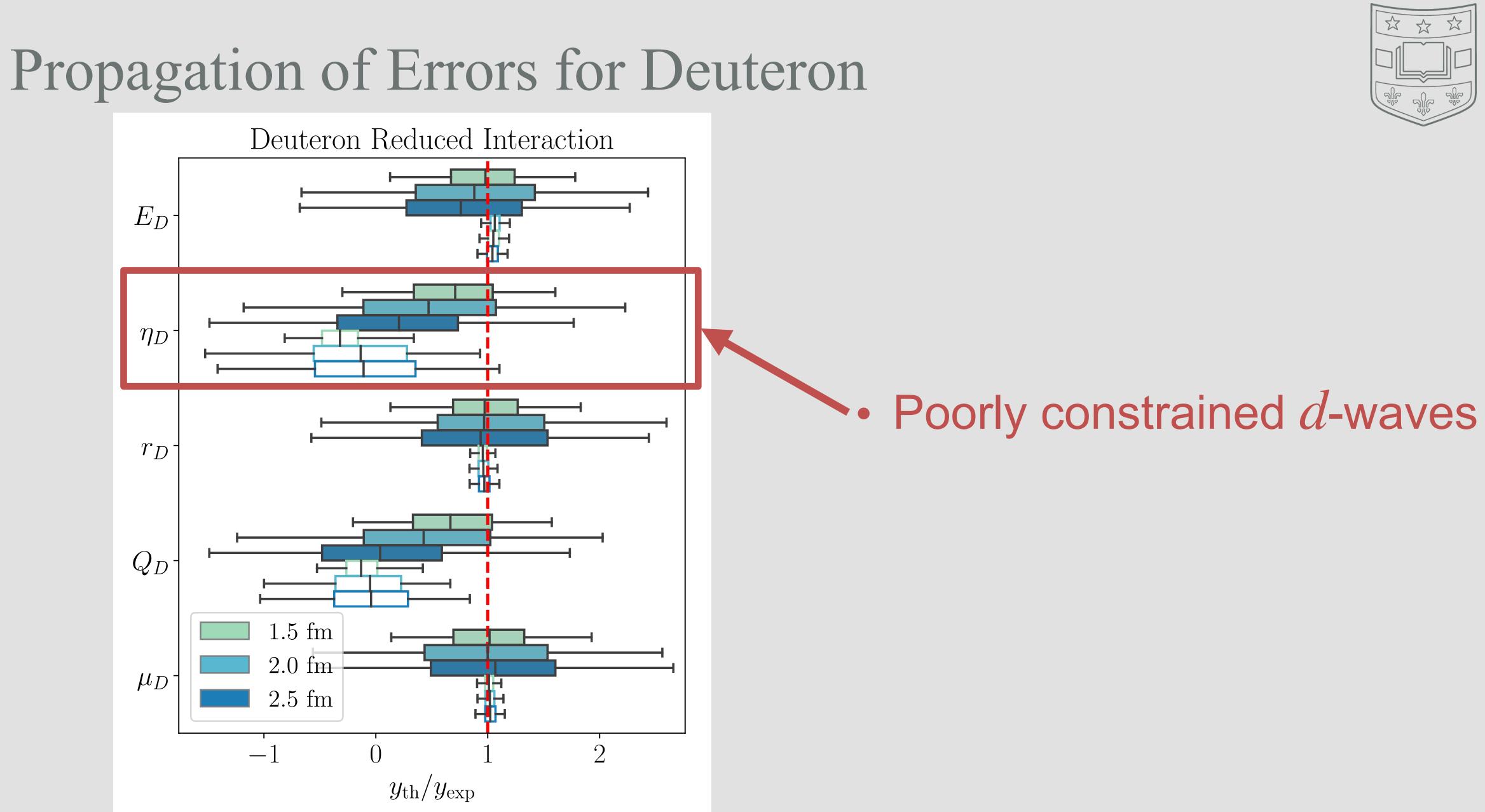
$$_{b} \mathcal{N}\left(\vec{y}_{\text{th}}, \Sigma_{\text{th}}\right) \text{pr}(\vec{a}, \vec{c}^{2}, \Lambda_{b} | \vec{y}_{\text{exp}}, I)$$



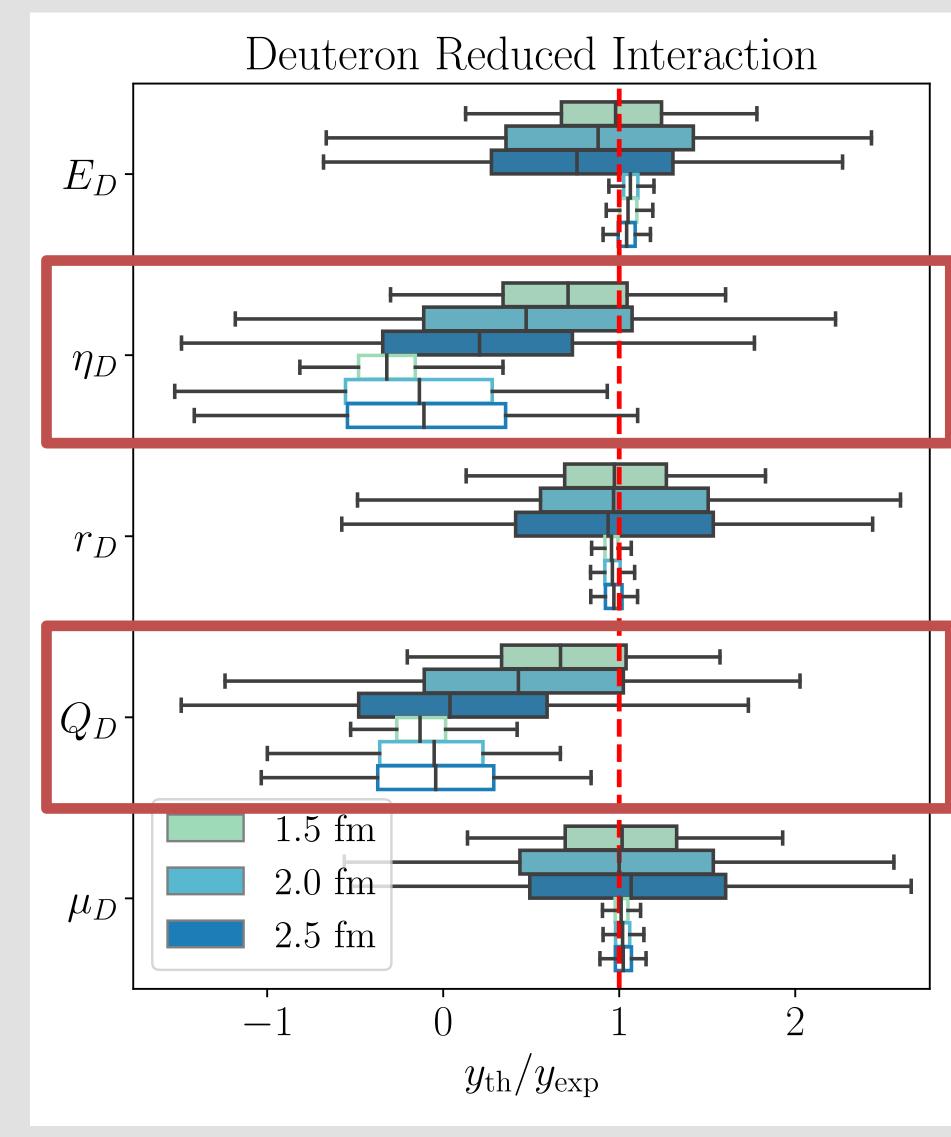
## Propagation of Errors for Deuteron







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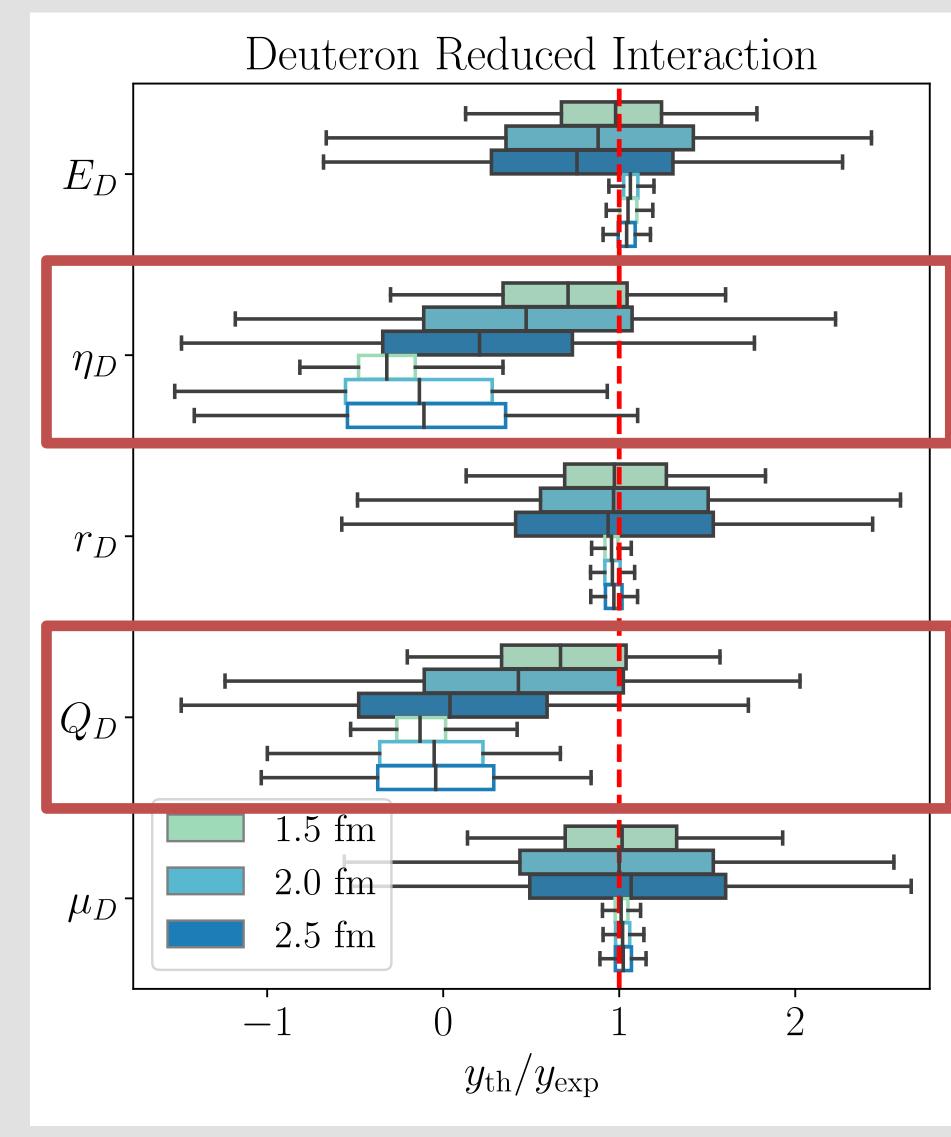




## Poorly constrained *d*-waves • 2b corrections at $O(Q^5)$



## Propagation of Errors for Deuteron

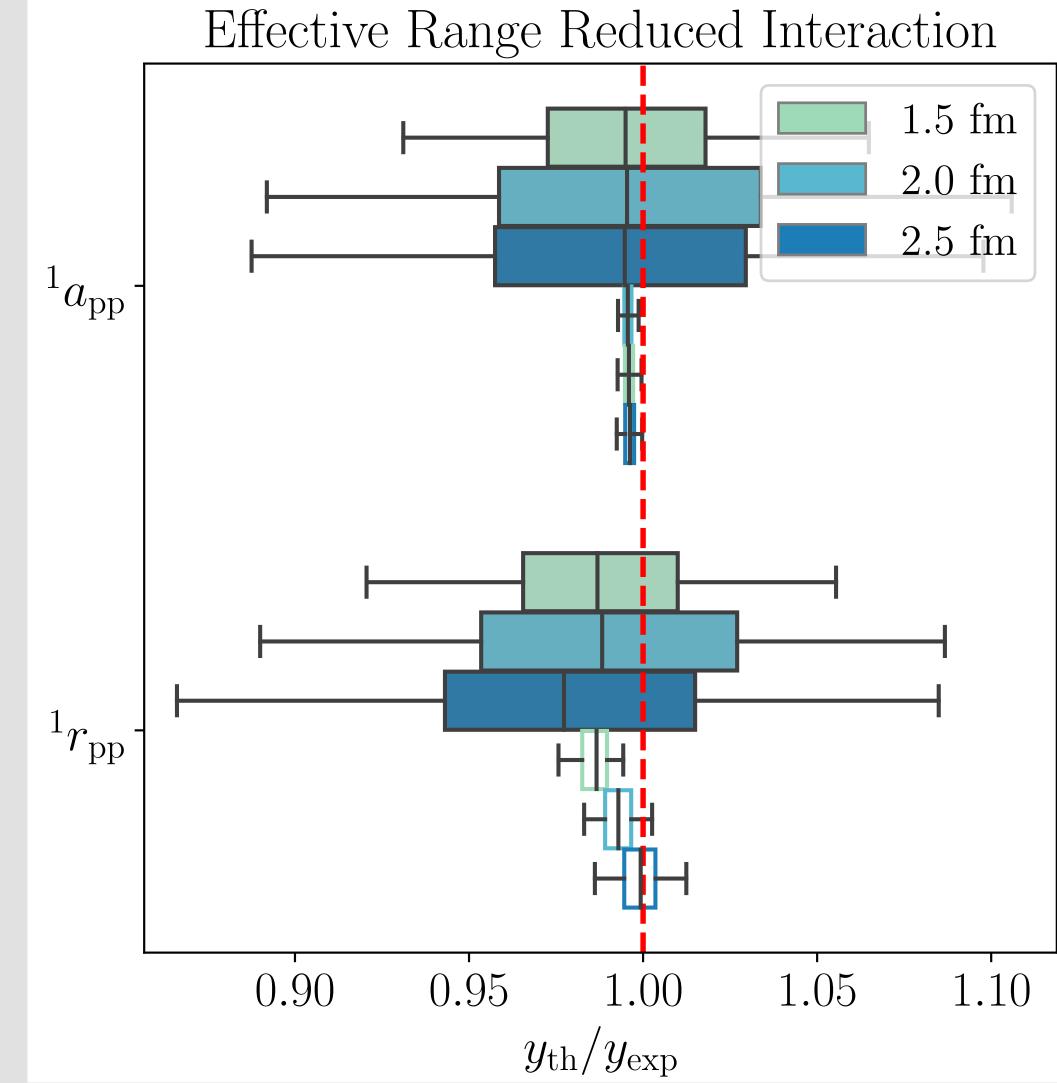




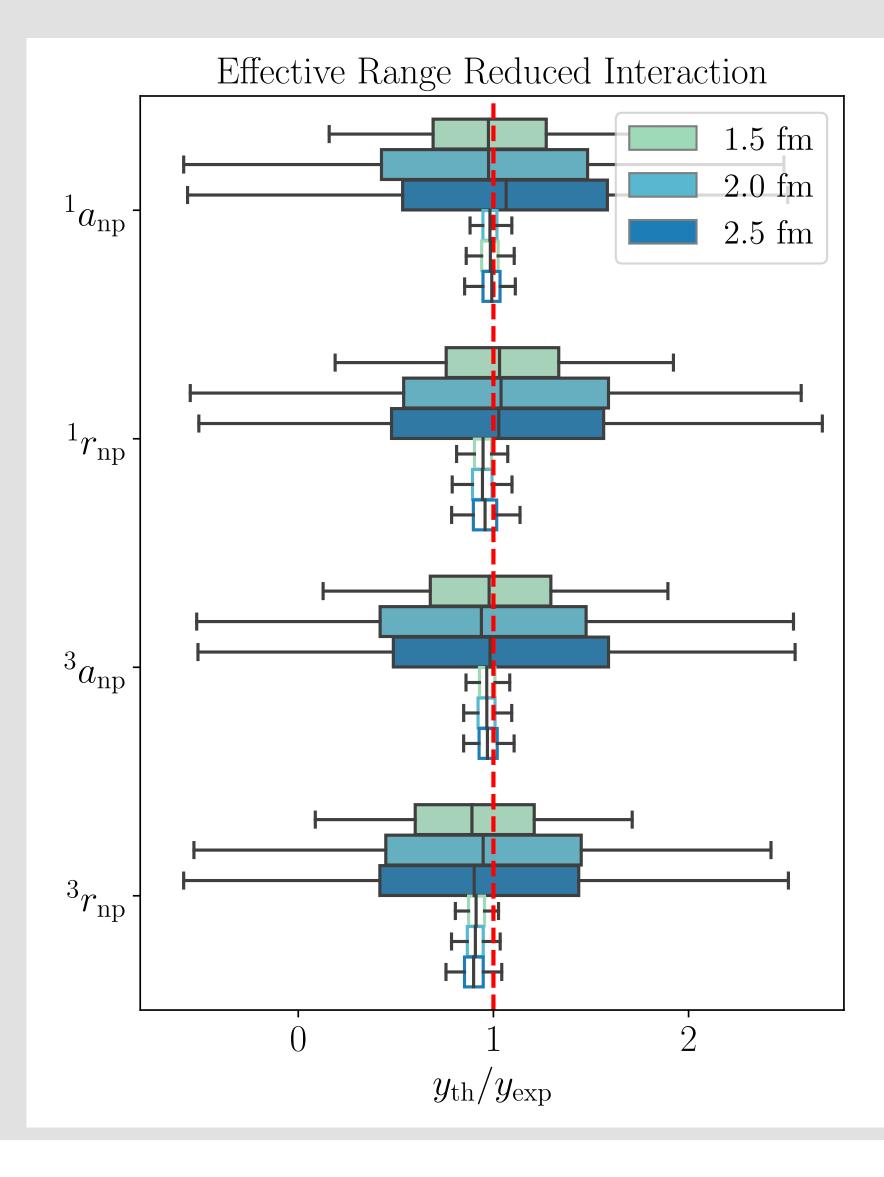
## Poorly constrained *d*-waves • 2b corrections at $O(Q^5)$ $\rightarrow$ SHOULD BE **CONSISTENT WITH 0**



## Propagation of Errors for ERPs















### • For pion- and $\Delta$ -full interactions, we must look at higher energy data (~200 MeV)



- Emulation for calculation of scattering observables



• For pion- and  $\Delta$ -full interactions, we must look at higher energy data (~200 MeV)



- Emulation for calculation of scattering observables
- $\sim 10^5$  parallel sample steps at  $\sim 4$  min. per step  $\rightarrow$  280 days of wall time on an HPC

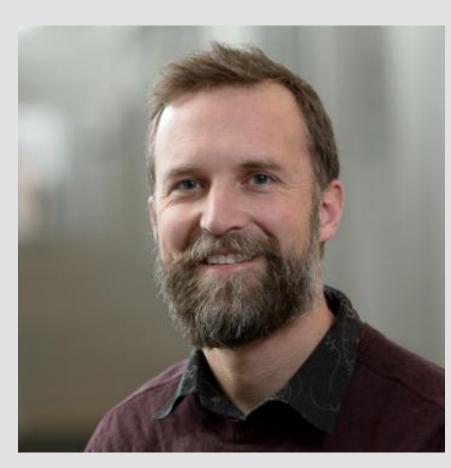


• For pion- and  $\Delta$ -full interactions, we must look at higher energy data (~200 MeV)



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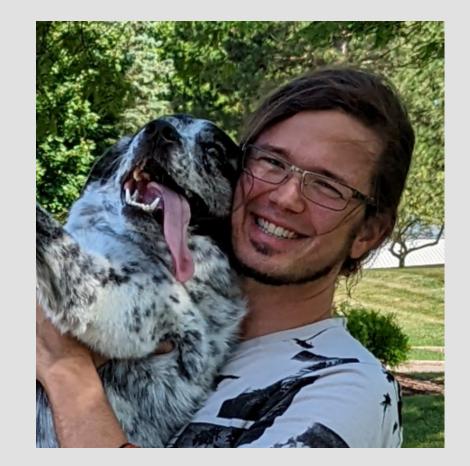


### Ozge Surer Stefan Wild LBNL Miami University



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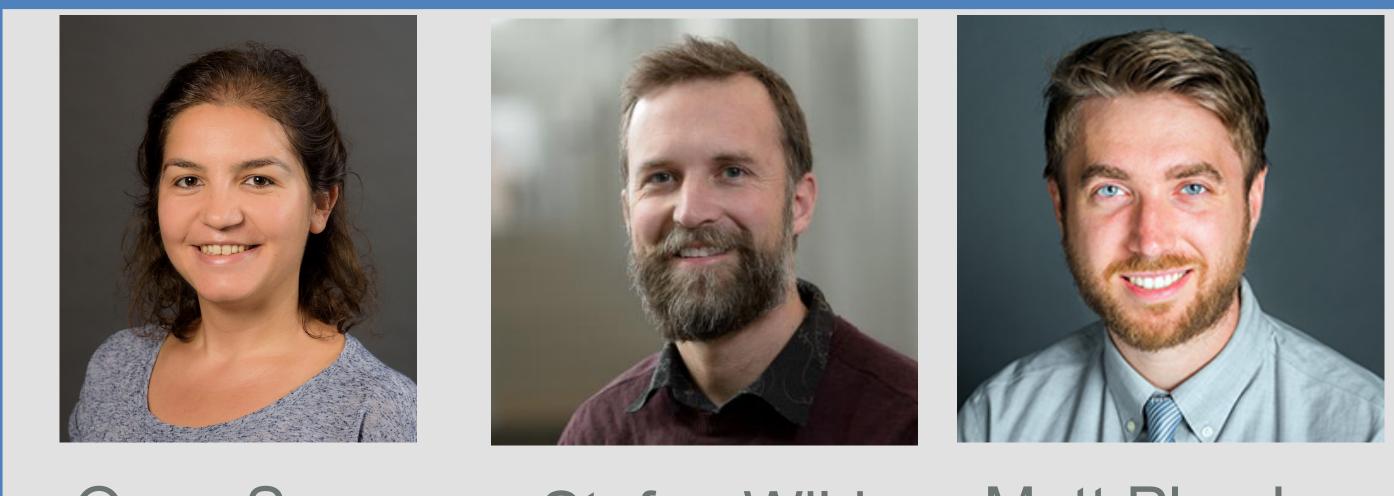


Pablo Giuliani MSU/FRIB

**Daniel Odell** SRNL



- Emulation for calculation of scattering observables

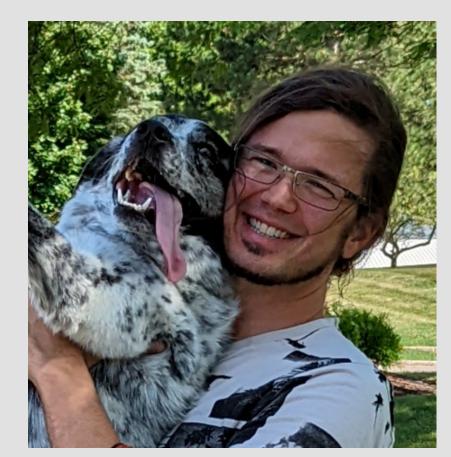


### Ozge Surer Matt Plumlee Stefan Wild Miami University LBNL Amazon

**Gaussian Process Emulation** 



## • For pion- and $\Delta$ -full interactions, we must look at higher energy data (~200 MeV)



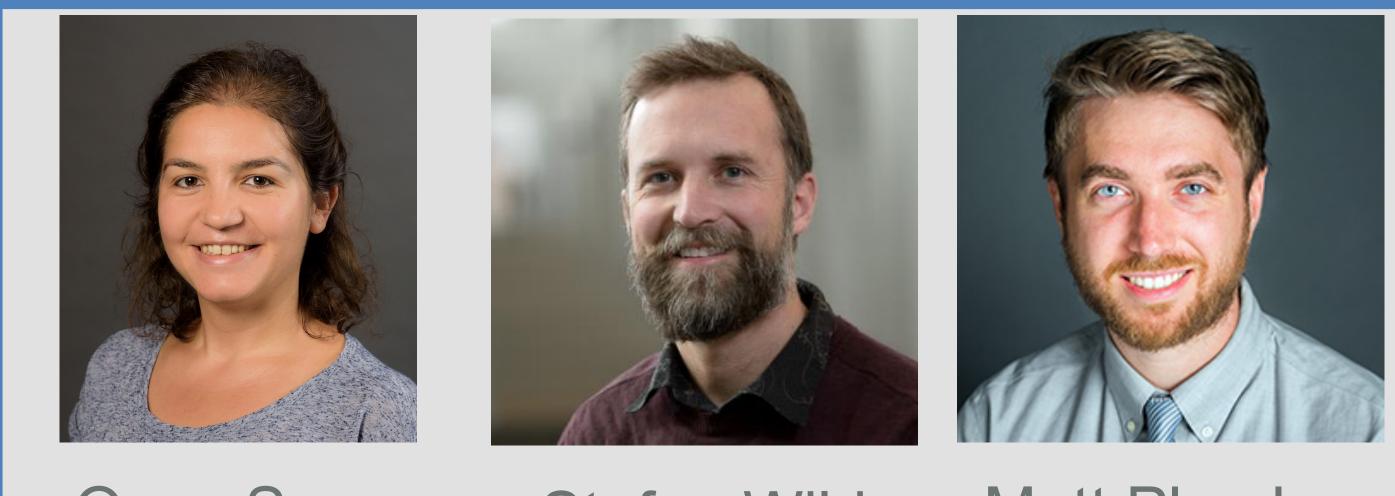
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**Gaussian Process Emulation** 



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Matt Plumlee Amazon





### Pablo Giuliani **Daniel Odell MSU/FRIB** SRNL

**Reduced Basis Methods via Galerkin Projection** 







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## **-T Model Calibration**









### Does proper inclusion of theoretical uncertainties in the calibration of EFT-inspired potentials bring all such models in agreement?



- Can Bayesian machinery help us identify consistent power counting? E.x. breakdown discrepancies across orders?



 Does proper inclusion of theoretical uncertainties in the calibration of EFT-inspired potentials bring all such models in agreement?



- Can Bayesian machinery help us identify consistent power counting? E.x. breakdown discrepancies across orders?
- body forces and currents?



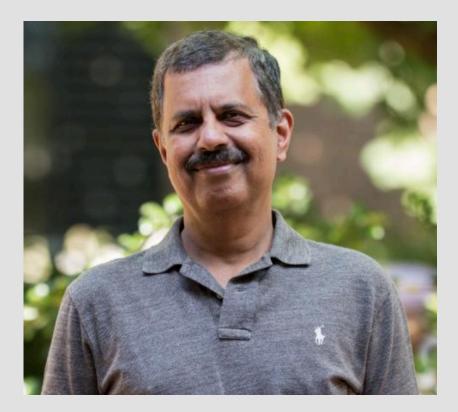
 Does proper inclusion of theoretical uncertainties in the calibration of EFT-inspired potentials bring all such models in agreement?

 Data can inform us about EFTs if we work consistently. How do we move towards consistent order-by-order inclusion of many-



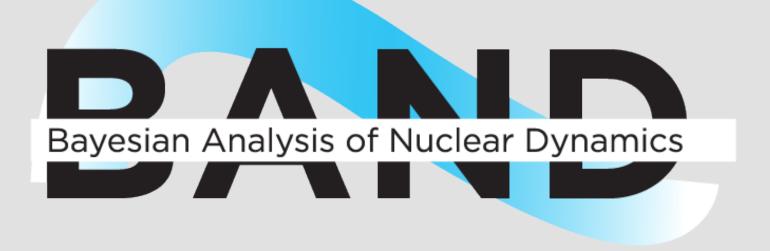
## Acknowledgements

QMC@WashU Piarulli (PI), Pastore (PI), Novario (SS), Weiss(PD\*), Flores (PD), Chambers-Wall (GS)



Sai lyer WashU

### Fellowship/Travel

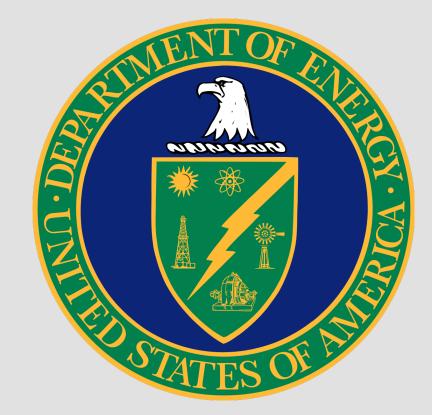




### Computational Resources



### Funding



### Collaborations

NTNP Nuclear Computational Low-Energy Initiative

A SciDAC-5 Project

