

A global data-driven dispersive optical model

Uncertainty-quantified non-local nucleon-nucleus
scattering and structure description across the chart

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Towards a consistent approach for nuclear structure and reactions:
microscopic optical potentials

ECT*, June 2024

Optical-model potentials

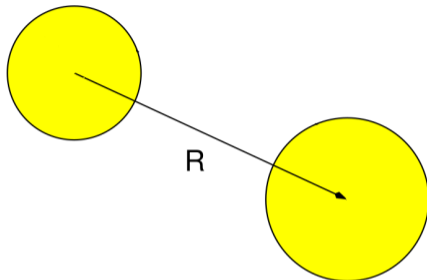
Degrees of freedom: nuclei.

Projection of the Hamiltonian on the elastic channel (Feshbach formalism, e.g. A. Moro, 2019, doi.org/10.3254/978-1-61499-957-7-129)

The projected interaction, U , is the optical-model potential:

$$U = PVP + PVQ \frac{1}{E - QHQ + i\epsilon} QVP \quad P, Q \text{ projections}$$

- A completely consistent U will be complex, non-local, and energy-dependent.
- Imaginary part: flux leaving the elastic channel.
- U cannot be actually computed as above



Goal

Design and train a phenomenological optical model that

- Has all features required for a fully consistent microscopic potential: fully non-local and dispersive.
- Has sound uncertainty quantification, also accounting for model defects. Desirable in more ab-initio models too.
- Provides a good description on a wide area of the chart (global) and can be reliable in extrapolation.

- Introduction
- **Dispersive optical-model**
 - Introduction
 - Link to bound-state properties
- Uncertainty quantification
- Resources and preliminary result

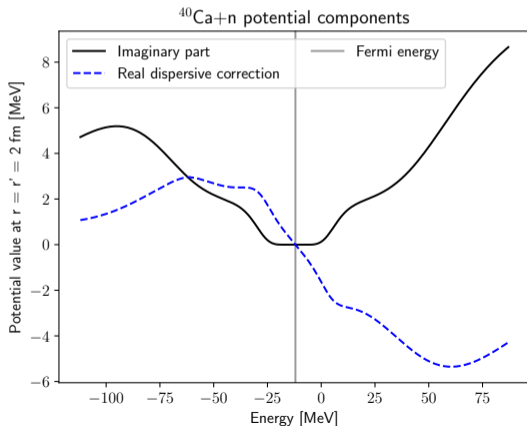
The dispersive optical model

Causality principle requires (J. S. Toll. *Phys. Rev.* 104.6 (1956))

OMP, U , to follow a dispersion (Kramers-Kronig) relation in energy:

$$U(\alpha, \beta, E) = U_{\text{HF}}(\alpha, \beta) + U_D(\alpha, \beta, E), \quad \text{Re}U_D(\alpha, \beta, E) = \frac{1}{\pi} \text{PV} \int_{\mathbb{R}} \frac{\text{Im}U_D(\alpha, \beta, \mathcal{E})}{E - \mathcal{E}} d\mathcal{E}$$

Example of
 $^{40}\text{Ca} + n$ imaginary potential, and
corresponding dispersive correction



Computing the dispersive correction

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In practice, the “usual” (Woods-Saxon, ...) forms are used for

- Energy-independent real part (less parameters)
- Energy-dependent imaginary part, with analytic forms whose integral is known (?).

The energy-dependent real part is then computed as above.

The potential (Hamiltonian) must be defined for all energies, positive and negative, to apply dispersivity.

$\text{Re}U(E)$ is connected to $\text{Im}U$ at all energies (and vice versa).

Consistent H for negative energies \rightarrow structure information

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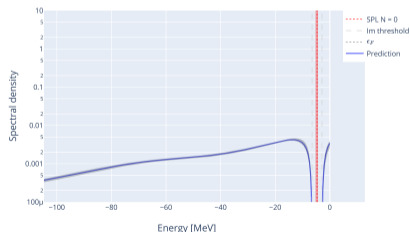
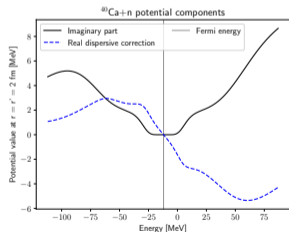
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Consistent H for negative energies \rightarrow structure information

Features of a dispersive optical model

Dispersive optical models:

- Are computationally more expensive (manageable with appropriate analytic forms)
- Are more accurate (a physical condition is being enforced)
- Have less free parameters (the energy dependence of the real part is fixed)
- Describe consistently scattering and bound-state properties



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How to compute spectral functions

- 1 Define the potential in a basis for the nucleus-nucleon motion $\{\alpha\}$, $V_{\alpha\beta}(E)$.
 E nucleon-nucleus rel. energy
- 2 Compute the Hamiltonian, $H_{\alpha\beta}(E)$.
- 3 Compute the propagator, (requires inverting H , i.e. solving the scattering problem)

$$G_{\alpha\beta}(E) = \lim_{\eta \rightarrow 0} \frac{1}{E - H_{\alpha\beta}(E) + i\eta}$$

- 4 The hole spectral functions are

$$S_{\alpha}^h(E) = \frac{1}{\pi} \text{Im} G_{\alpha\alpha}(E)$$

(analogous for particle spectral functions).

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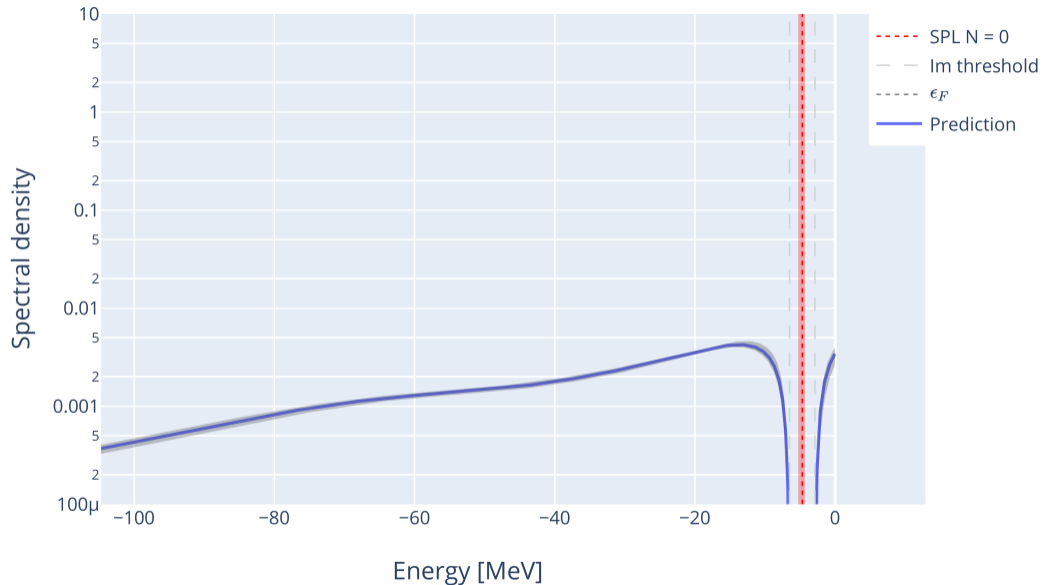
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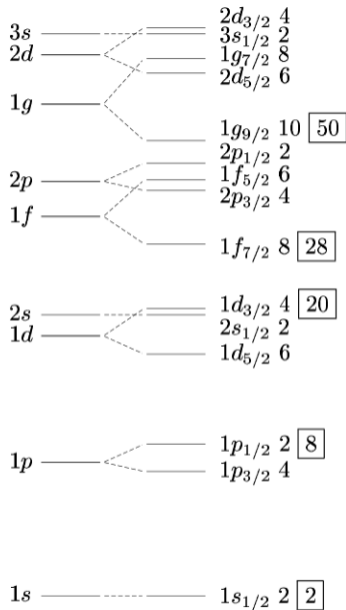
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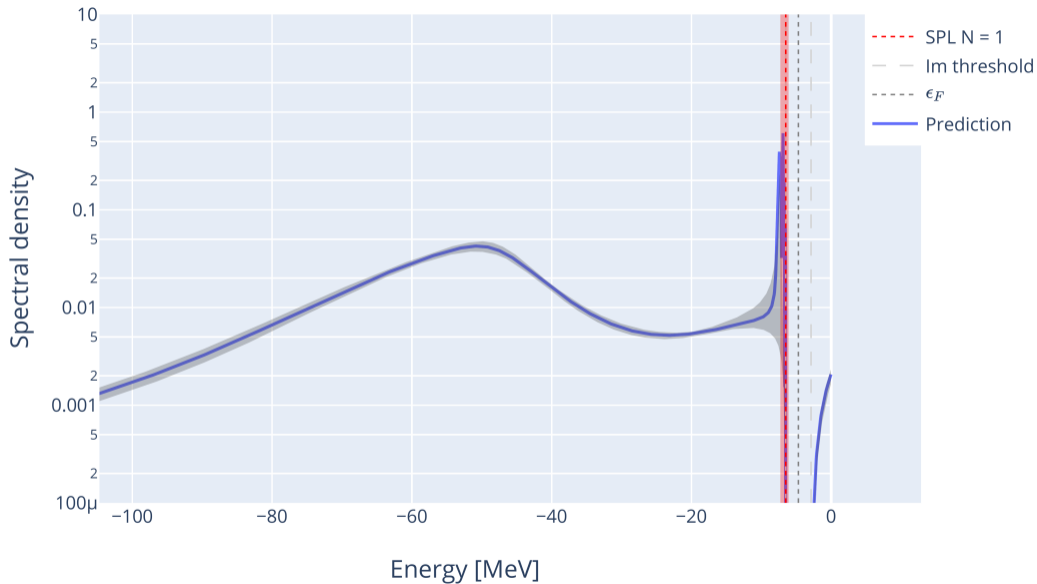
$^{40}\text{Ca} - p$ $L = 2, J = 3/2$ spectral function (fit on ^{90}Zr , $^{112,124}\text{Sn}$, ^{208}Pb)



<https://en.wikipedia.org/wiki/File:Shells.png>



$^{40}\text{Ca} - p$ $L = 0, J = 1/2$ spectral function (fit on ^{90}Zr , $^{112,124}\text{Sn}$, ^{208}Pb)



Given the spectral function $S_\alpha(E)$,

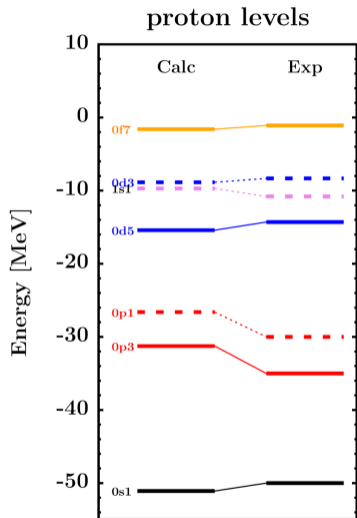
$$\text{one-body density } n_\alpha = \int_{-\infty}^{\mathcal{E}_F} S_\alpha(E) dE \quad , \quad \text{number of particles } n = \sum_{\alpha} n_\alpha$$

If $\alpha = (\vec{r}, a)$, $n_\alpha = \text{density } \rho_a(\vec{r}) \longrightarrow \text{charge radius, electromagnetic moments}$

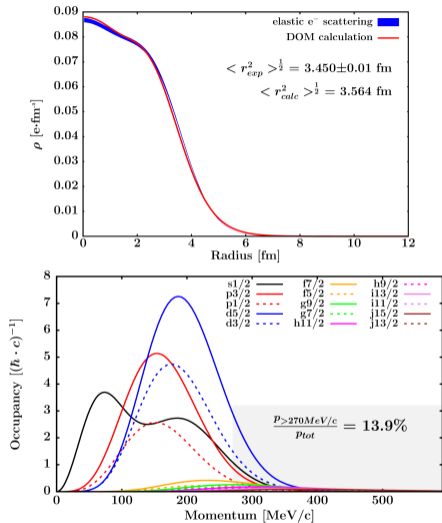
If $\alpha = (r, a)$, “Spectroscopic factors” $S_a(E) = \int S_{a,r}(E) dr$

s.p. energies $E_a = \int_{-\infty}^{\mathcal{E}_F} S_a(E) E dE \longrightarrow \text{binding energy}$

Examples of ^{40}Ca bound-state observable predictions

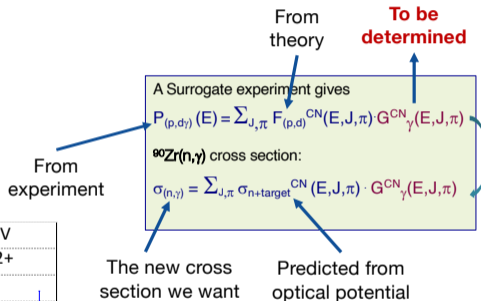
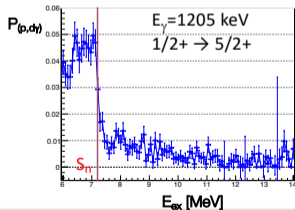
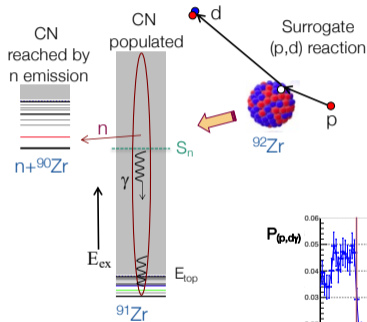


(C. Pruitt, 2019, PhD thesis, 10.7936/hyxx-6e81)



Surrogate reactions method for neutron capture

Escher et al, PRL 121, 052501 (2018)



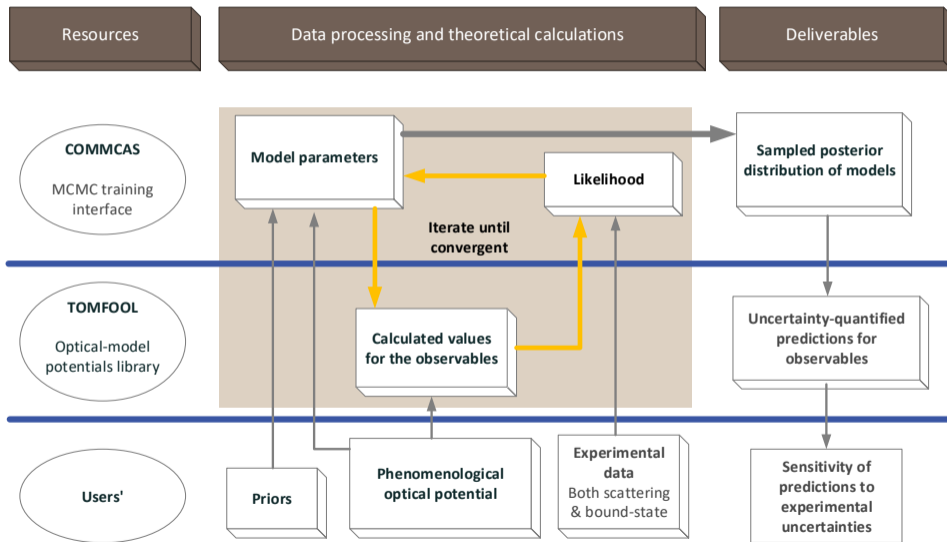
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- **Dispersive optical-model**
- **Uncertainty quantification**
 - Markov-chain Monte Carlo training
 - Likelihood function
- **Resources and preliminary result**

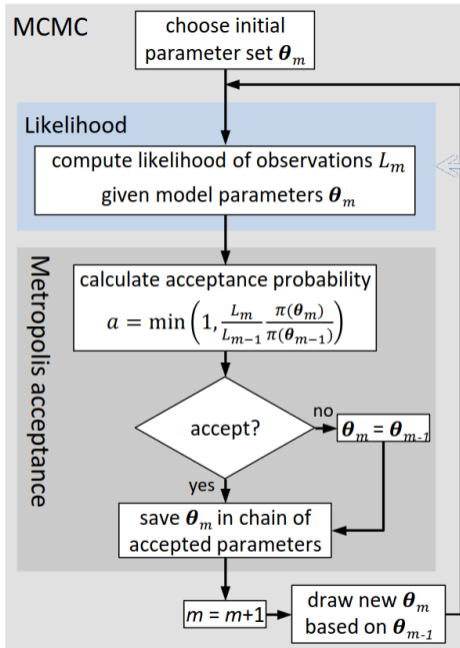
- Highly non-linear model.
- High-dimensional parameter space (e.g. Koning-Delaroche has 46 parameters).
- Need good uncertainty quantification.

→ Fit performed through Markov-chain Monte Carlo (Bayesian):

- Computationally demanding.
- Statistical assumptions more explicit (priors).
- Straightforward and reliable error estimation (“fit” yields a sample of posterior distribution of potentials):
 - No truncation to 2nd moments.
 - No Gaussian distribution assumption.

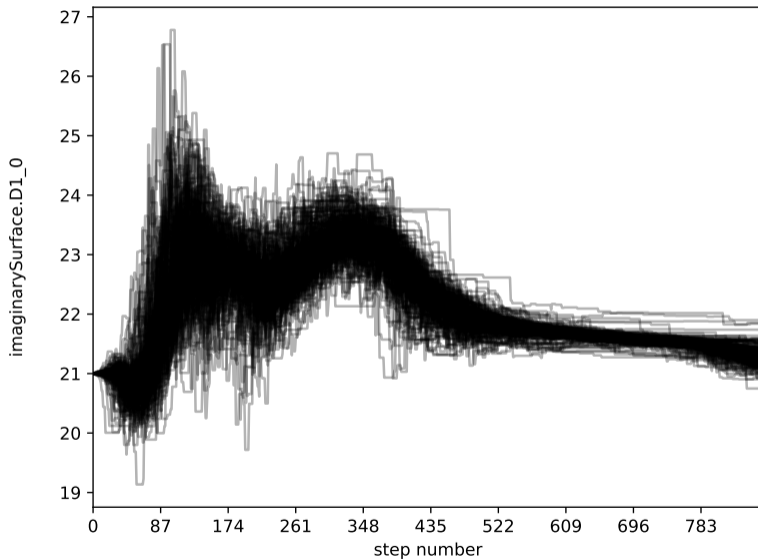
MCMC training of data-driven optical models





ebooks.iospress.nl/
publication/48604

A parameter trace for a Markov-Chain-Monte-Carlo OMP inference



The MCMC “fit” yields a sample of the posterior distribution of potentials:

Potential 1: $r_{C0} = 1.21$ fm, $V_{00} = 51$ MeV, ...

Potential 2: $r_{C0} = 1.25$ fm, $V_{00} = 46$ MeV, ...

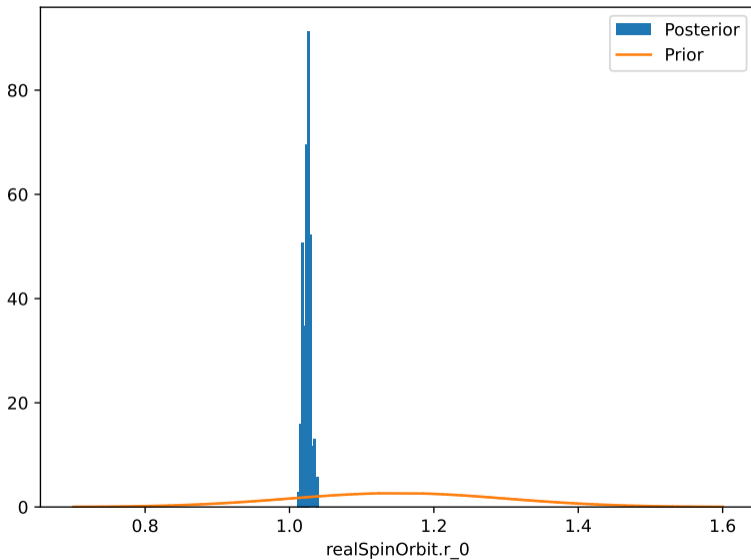
...

Potential 485: ...

From the sample, extract:

- Statistics on parameters
- Sample of distribution of predictions for any observable

Example of a posterior parameter distribution



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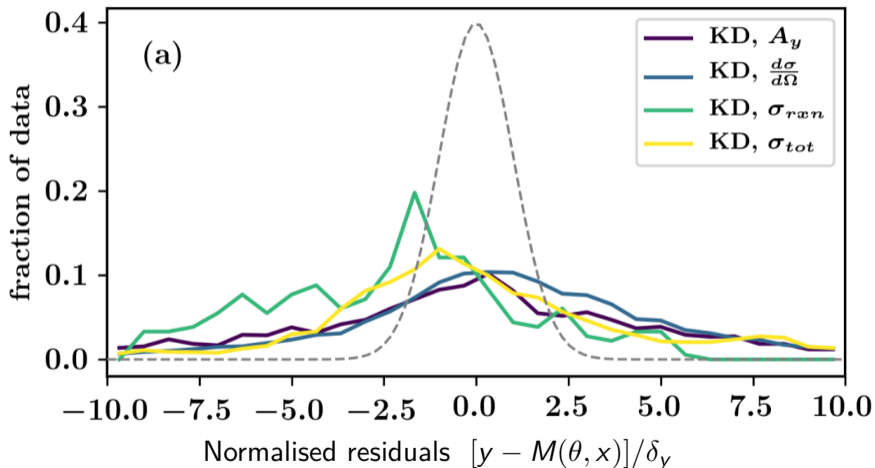
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Limitations of OMP fitting: past literature shows that...

Koning-Delaroche training dataset, by observable type

(C. D. Pruitt et al. *Phys. Rev. C* 107.1 (2023))



...good
uncertainty
quantification
is hard for
this problem.

(C. D. Pruitt et al. *Phys. Rev. C* 107.1 (2023)

and e.g. A. J. Koning et al. *Nuclear Physics A* 713.3 (2003))

- There are outliers.
- Uncertainties (statistical and systematic) are underestimated.
- Errors due to experiment and model defects are not disentangled.
- Observations are not independent.
- $\chi^2 \gg 1$ and error estimation based on it is not meaningful

Unaccounted-for-uncertainties estimation

Covariance matrix in C. D. Pruitt et al. *Phys. Rev. C* 107.1 (2023)

(the expression was simplified for illustrative purposes):

$$\tilde{\Sigma} = \frac{k}{N} \text{diag}(\vec{\Delta}) \quad , \quad \vec{\Delta} = \left\{ \delta_y^2 + \left(y \delta_{\hat{t}(y)} \right)^2 \right\}$$

δ_y reported data error, y observation, $M(x)$ model prediction,
 $\hat{t}(y)$ “type” (proton elastic σ , neutron analyzing power, ...) of y ,
 $\delta_{\hat{t}}$ fitted parameters.

Likelihood (being $r = y - M(\vec{\theta}, x)$):

$$L = \left[(2\pi)^k |\tilde{\Sigma}| \right]^{-1/2} \exp \left[-\frac{1}{2} \frac{r^2}{|\tilde{\Sigma}|} \right]$$

“Sigma clipping”:

- 1 Converge fit on chosen data set (initially, all data).
- 2 Chose data set as points such that

$$y - M(\vec{\theta}, \mathbf{x}) \leq 3\sqrt{\delta_y^2 + \text{var}(M(\vec{\theta}, \mathbf{x}))}$$

(discarded data can be recovered at a later step).

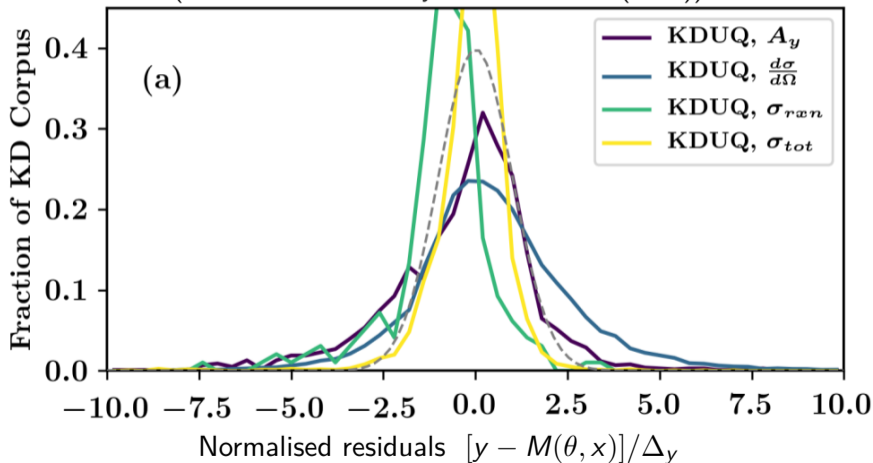
- 3 Repeat until convergent.

Past literature shows that...

Koning-Delaroche training dataset, by observable type

Uncertainty-quantified potential

(C. D. Pruitt et al. *Phys. Rev. C* 107.1 (2023))



...missing data- and model-uncertainties can be effectively estimated.

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How much computing power do we need?

- Currently, 0.7 s for one Green's function, 0.04 s for one scattering energy.
(Could be a great use case for emulators)
- 2×10^4 MCMC steps (at least, if fitted UAU and outlier rejection desired)
- 150 target nuclei (for a “global” fit).
- 2 projectiles (proton and neutron) per target.
- 400 MCMC walkers (~ 40 fit parameters)

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- Half walkers can run in parallel
(stretch move): 200 cores, ~ 416 days.
- Each system can run in parallel:
 6×10^4 cores, ~ 33 hours.

Quartz (108 648 cores)

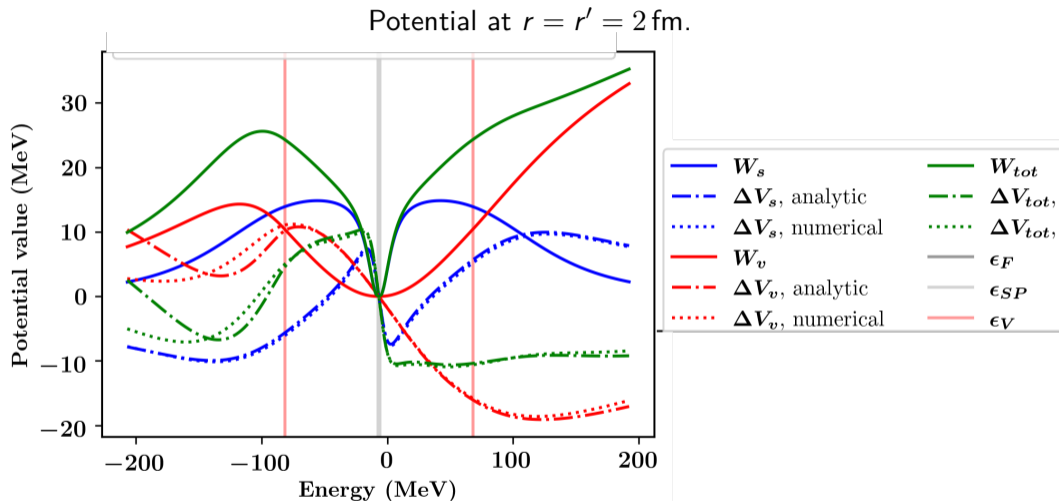


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The currently-adopted potential

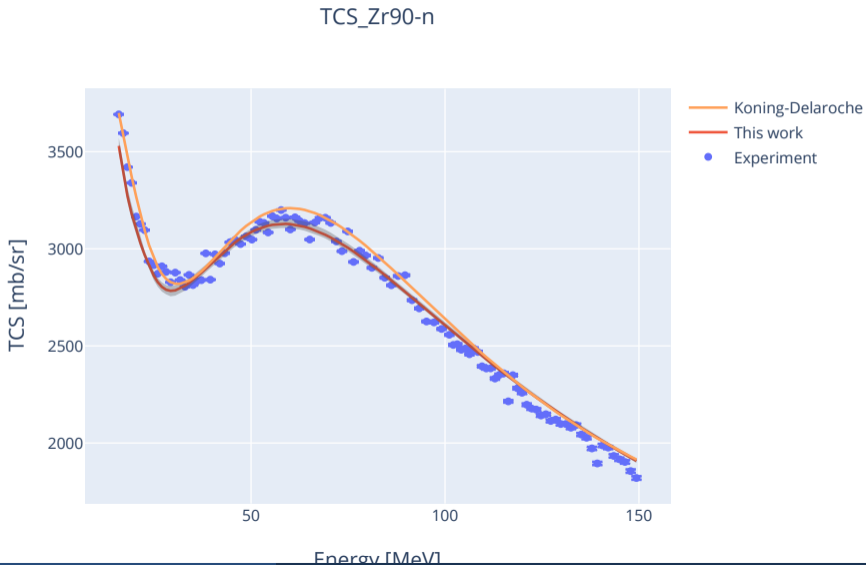
$$\begin{aligned} & Z_1 Z_2 \text{Coulomb} \left(R_1, r_{C0} A^{1/3} + r_{C1} A^{-1/3} + r_{C2} A^{-4/3} \right) + \text{NLP} \cdot U \\ U = & - \left(V_{v0} - V_{vA} A^{-1/3} \pm V_{vs} \frac{Z - N}{A} \right) \text{WS}(R, r_{v0} A^{1/3} - r_{v1}) + \\ & + V_{SO} \vec{L} \cdot \vec{S} \frac{1}{r} \frac{d}{dr} \text{WS} - iW(E, W_1, W_2, E_F, \dots) \text{WS} + \\ & - \text{asymptotic imaginary volume correction}(E_a, \dots) + \\ & - \text{imaginary surface symmetric around } E_F + \text{imaginary spin-orbit} + \\ & + \text{dispersive correction}(E, W_1, \dots) \\ & R = (R_1 + R_2)/2 \\ \text{NLP} = & \frac{2\sqrt{R_1 R_2}}{\beta^2} \exp \left(-\frac{R_1^2 R_2^2}{\beta^2} \right) J(L + 0.5, 2R_1 R_2 / \beta^2) \end{aligned}$$

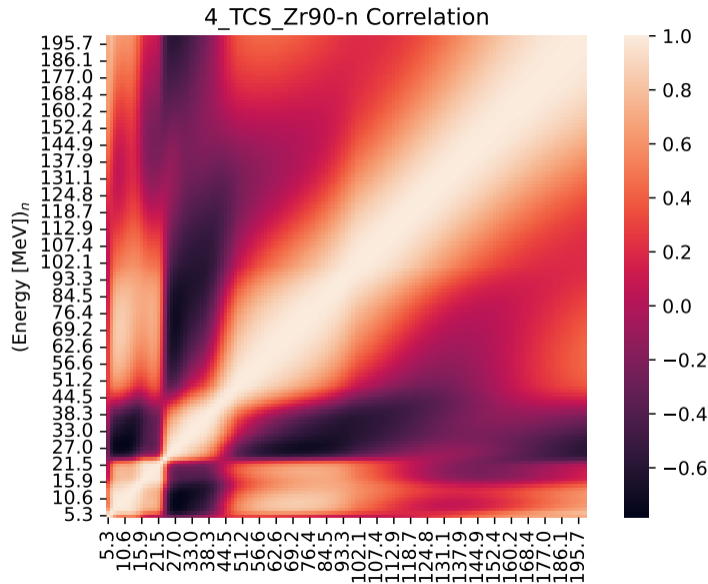
Plot with current imaginary forms



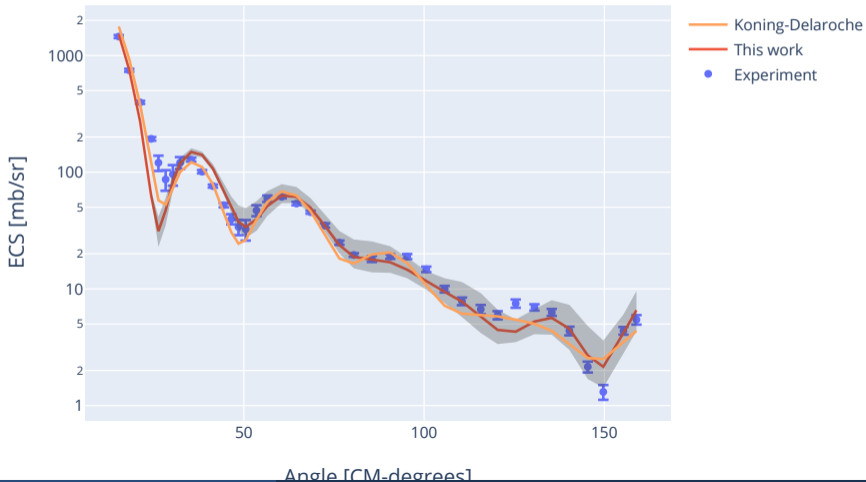
Analytic dispersive correction at $E < E_F$ valid at all E only when E_a is big.

- Fit on ^{90}Zr , $^{112,124}\text{Sn}$, ^{208}Pb : RCS, TCS, n_{nucleon} ; fixed UAU, no rejections
- Fit on $^{40,48}\text{Ca}$, $^{58,64}\text{Ni}$, ^{90}Zr bound properties only; fit UAU, no rejections

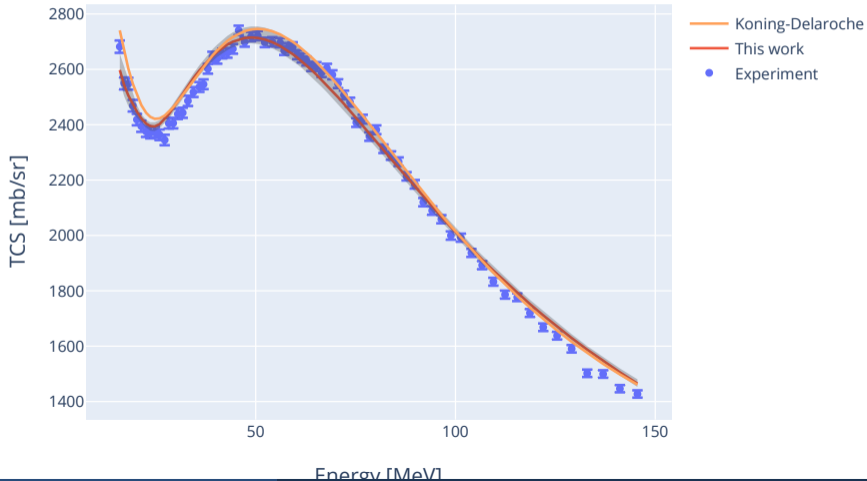




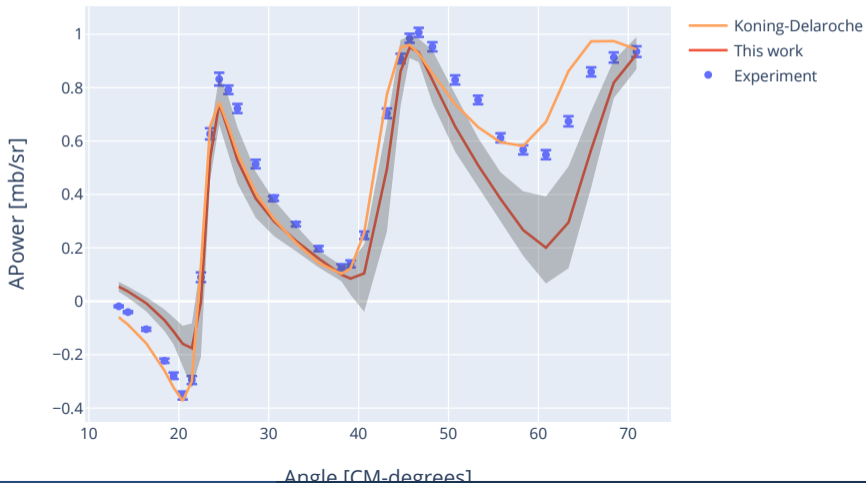
ECS_Zr90-n_24.0MeV



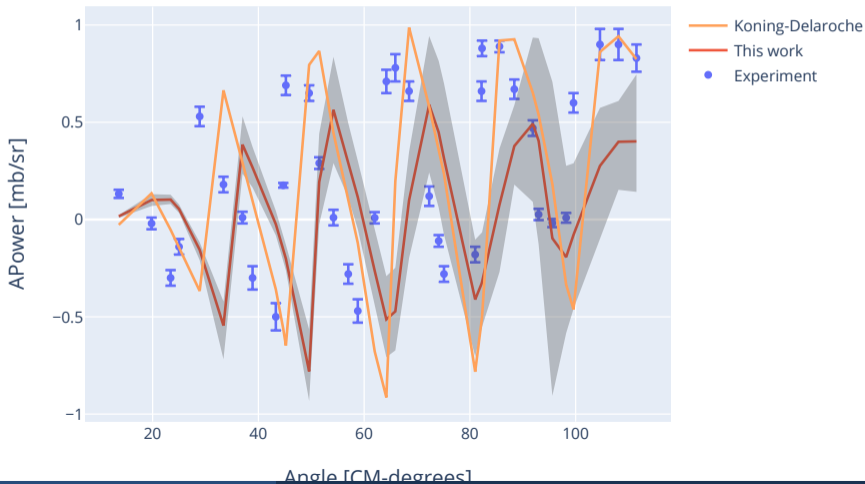
TCS_Ni64-n



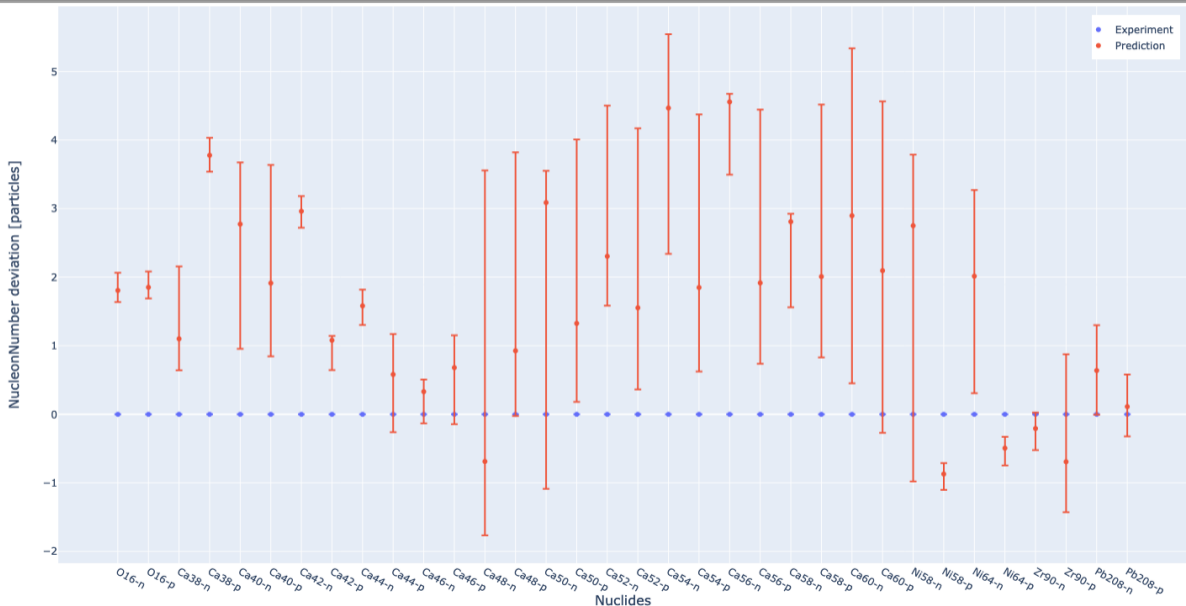
APower_Ni64-p_65.0MeV

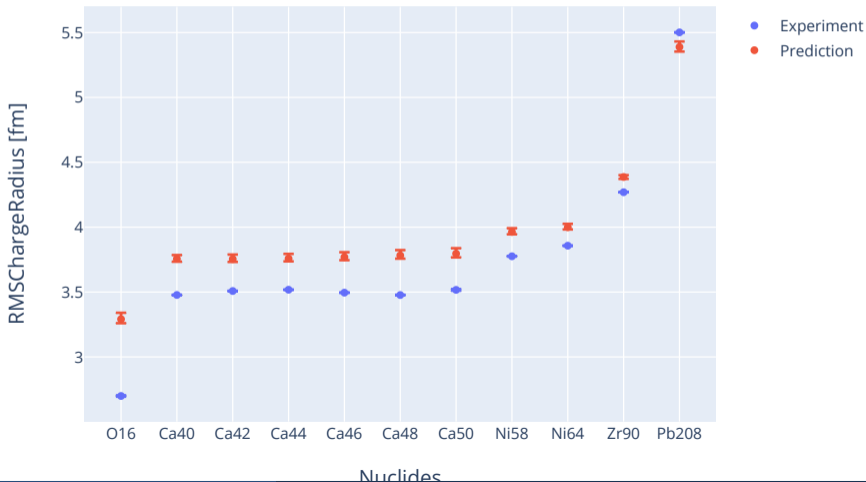


APower_Pb208-p_49.35MeV

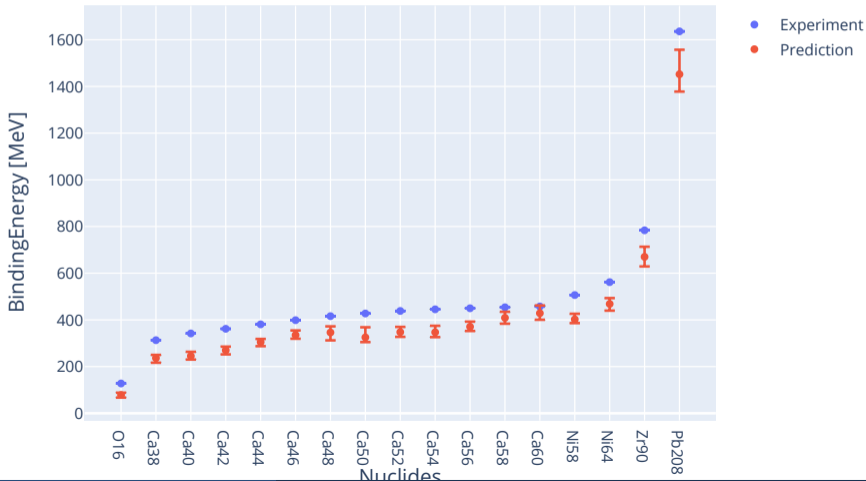


Fit on ^{90}Zr , $^{112,124}\text{Sn}$, ^{208}Pb : RCS, TCS, n_{nucleon} ; fixed UAU, no rejections



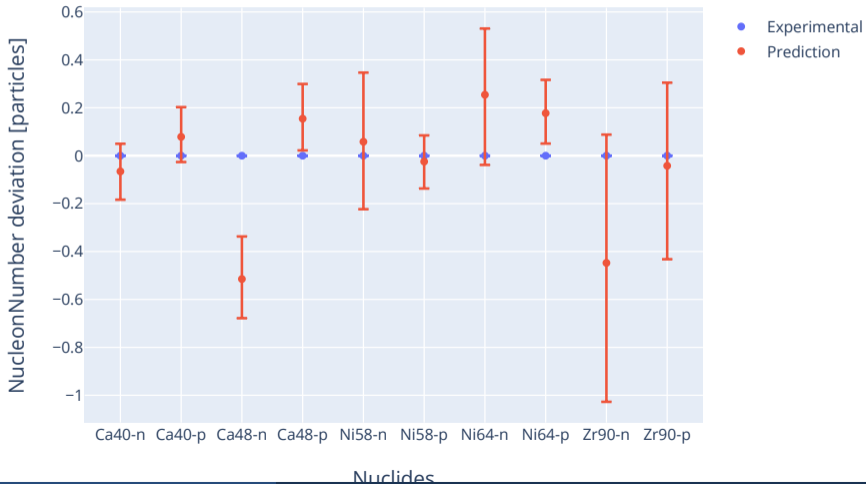


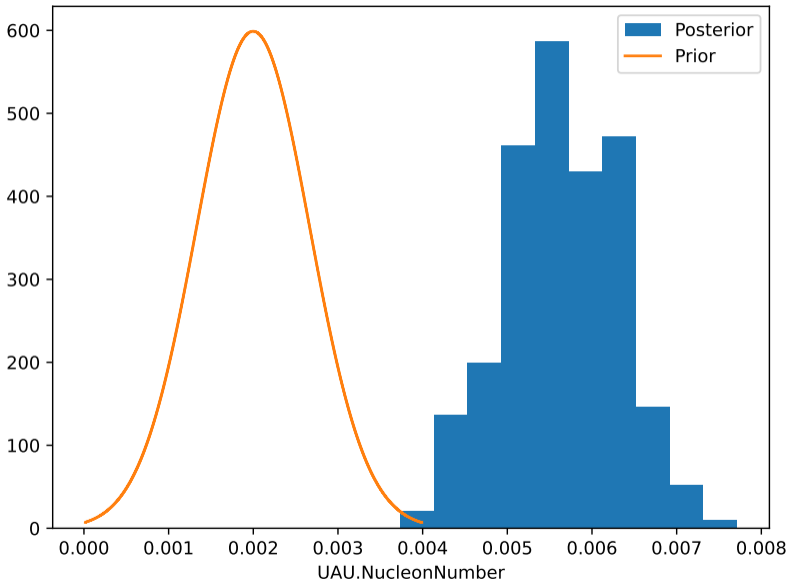
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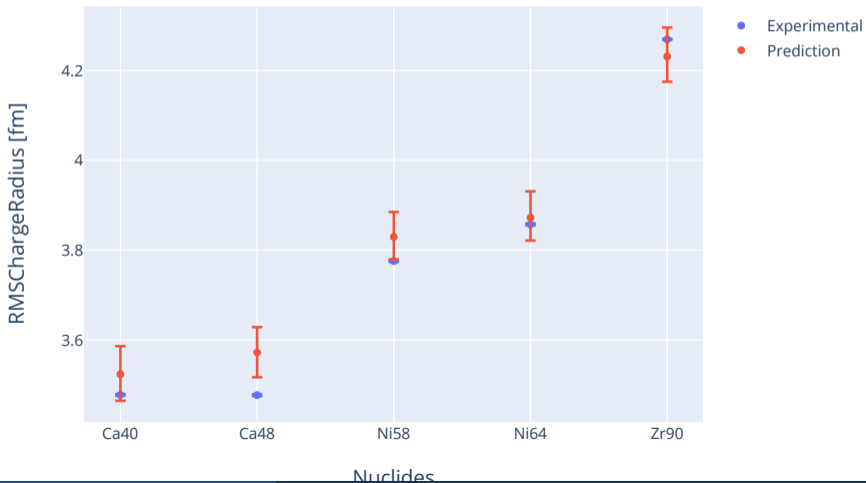
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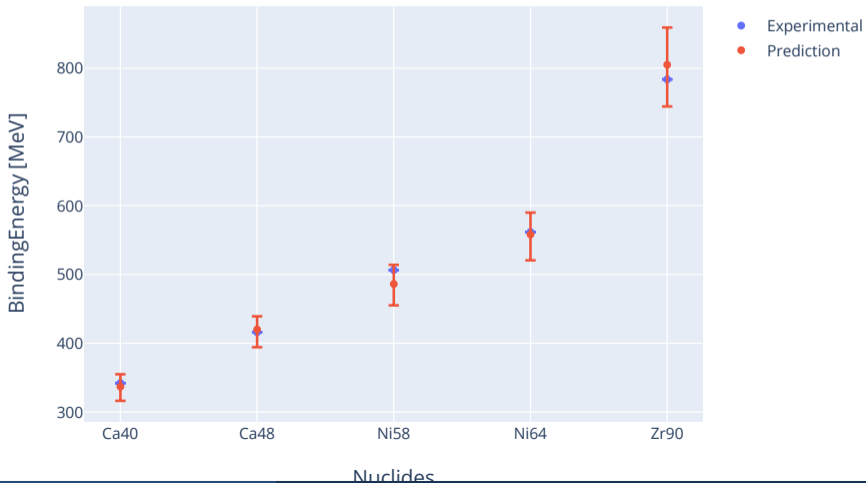




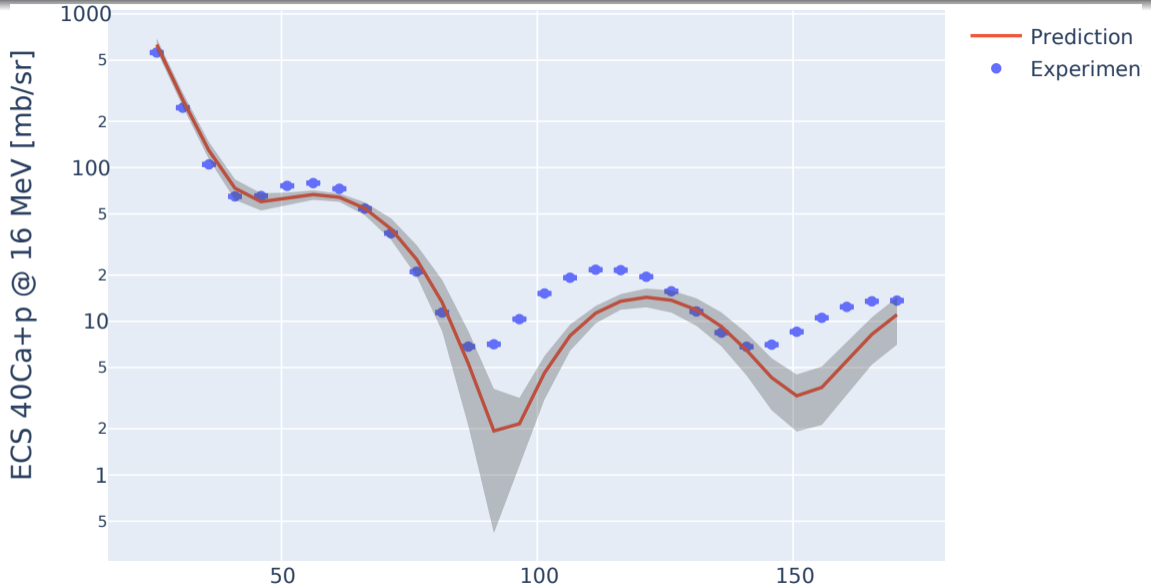
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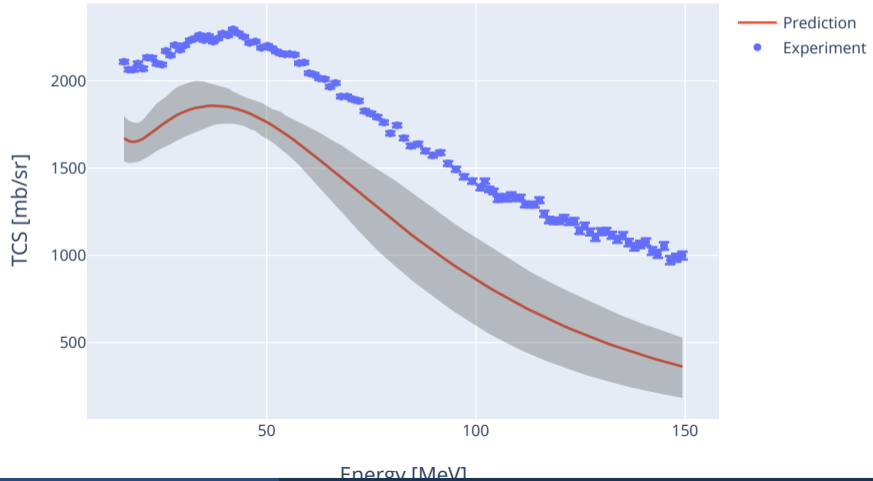
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Fit on $^{40,48}\text{Ca}$, $^{58,64}\text{Ni}$, ^{90}Zr bound properties only; fit UAU, no rejections



TCS_Ca40-n



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New potential form (under testing)

$$\text{Coulomb}(R_1, \dots) + \text{NLP} \cdot U$$

$$U = -\text{Real energy-independent volume}(R, \dots) + \text{spin-orbit}$$

–only for $E > E_F$: Imaginary volume with asymptotic correction

–imaginary surface non-symmetric around E_F

–at $E < E_F$: im. volume with surface-like E -dependence (no asymptotic-corrected volume)

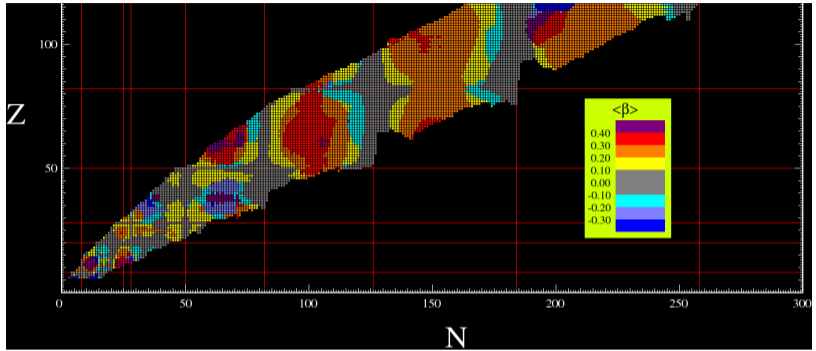
+dispersive correction(E, W_1, \dots)

$$R = (R_1 + R_2)/2$$

$$\text{NLP} = \frac{2\sqrt{R_1 R_2}}{\beta^2} \exp\left(-\frac{R_1^2 R_2^2}{\beta^2}\right) J(L + 0.5, 2R_1 R_2/\beta^2)$$

- Preliminary testing look promising.

`www-phynu.cea.fr/science_en_ligne/carte_potentiels_microscopiques/carte_potentiel_nucleaire_eng.htm`



“Spherical” potentials perform worse for deformed nuclei
[e.g. A. J. Koning et al. *Nuclear Physics A* 713.3 (2003)].

- 1 Take into account effects of deformation through models (e.g. coupled channels).
- 2 Extract equivalent OMP, minimal added complexity for e.g. elastic studies.

- Reliable global optical model, fully dispersive and non-local, trained on scattering and bound-state data, with sound uncertainty quantification, is within reach. (Nucleon numbers, binding energies, etc., available for very unstable systems).
- Changes to “traditional” potential form required for full dispersivity and good data reproduction (under testing now).
- User-friendly library handling such potentials (TOMFOOL) will be released.

To do:

- Test new form.
- Include more quantities (single-particle energies, charge exchange, skins, ...).
- Improve computational efficiency (more parallel, maybe emulators).

Thank you for your attention



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