

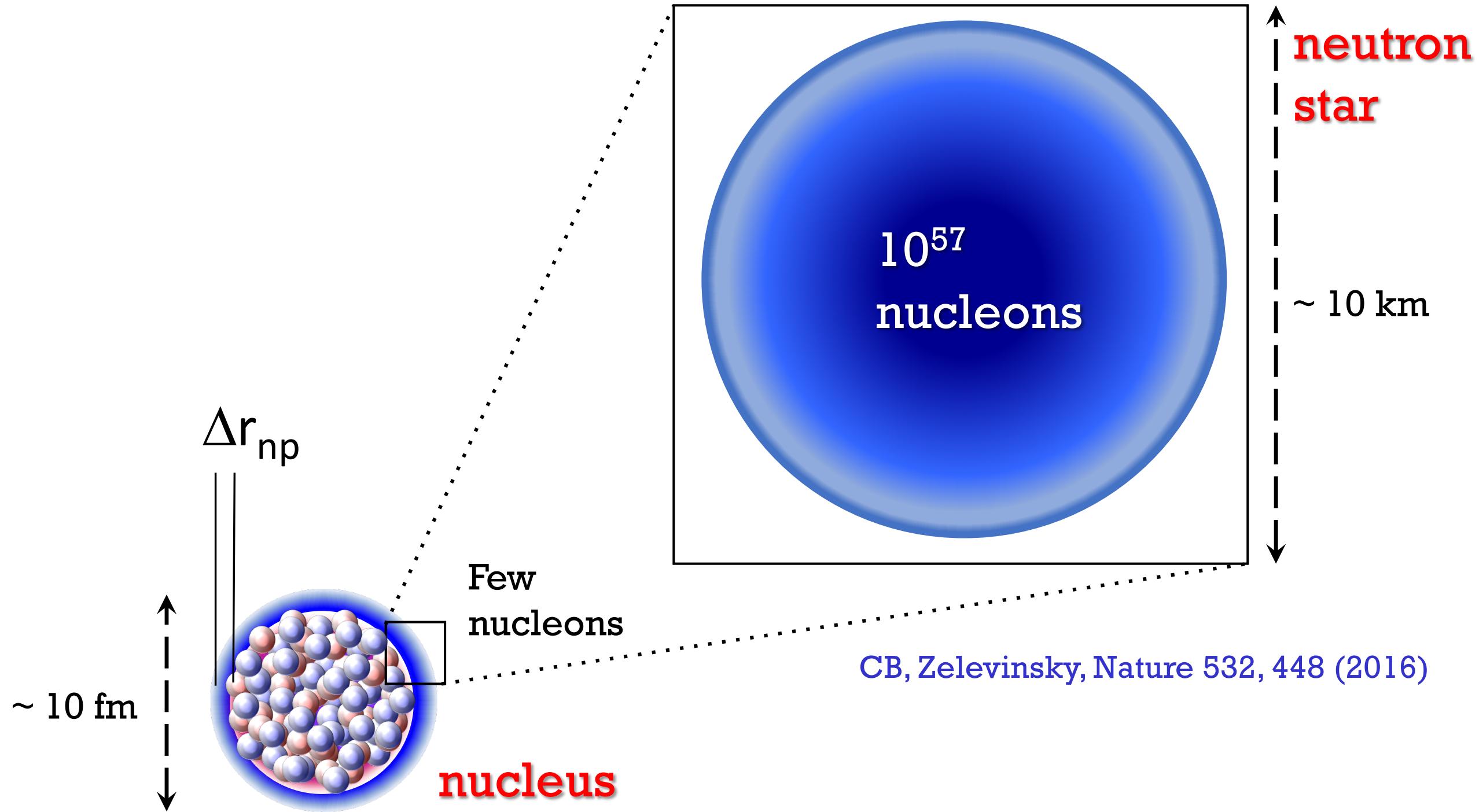
Excitation and decay of light hypernuclei

C.A. Bertulani

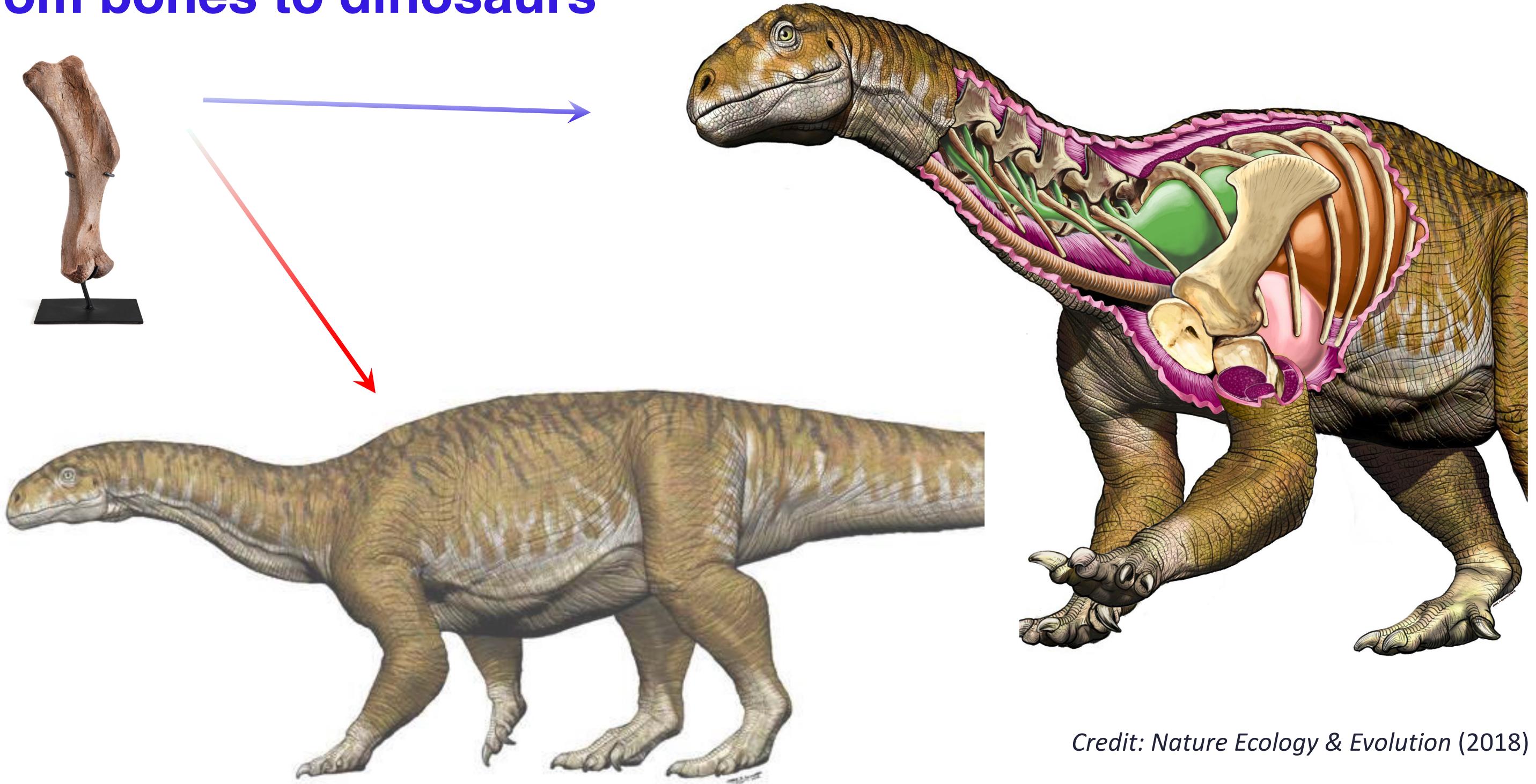


This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 824093

From laboratory experiments to neutron stars



From bones to dinosaurs



Credit: Nature Ecology & Evolution (2018)

Reconstruction of *Ingentia prima* from the Late Triassic (205- 210Ma) of Argentina. Total length 8-10 me...

Neutron stars

EOS

$$p[\rho] = \rho^2 \frac{d}{d\rho} \left(\frac{E}{A} [\rho] \right)$$

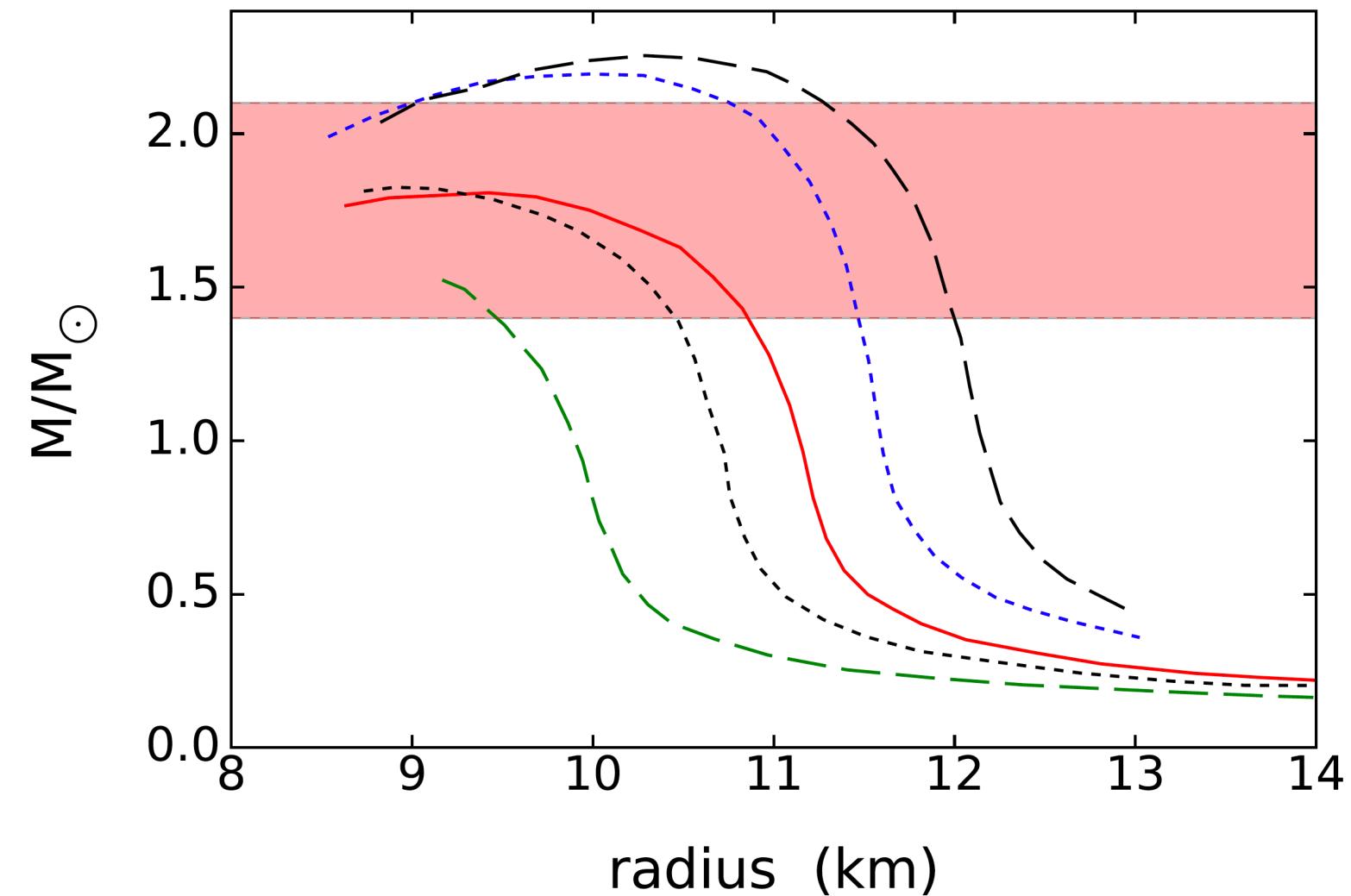
$$\frac{E}{A} [\rho] = \frac{E}{A} [\rho_0] + \frac{1}{18} K_\infty \left(\frac{\rho - \rho_0}{\rho_0} \right)^2 + \dots$$

$$K_\infty = 9\rho^2 \frac{d^2 [E/A]}{d\rho^2} \Big|_{\rho_0}$$

$$\frac{dP}{dr} = -\frac{G\rho(r)M(r)}{r^2} \left[1 + \frac{P(r)}{\rho(r)} \right] \left[1 + \frac{4\pi r^3 P(r)}{M(r)} \right] \left[1 - \frac{2GM(r)}{r} \right]^{-1}$$

$$\frac{dM}{dr} = 4\pi r^2 \rho(r)$$

Tolman-Oppenheimer-Volkoff



EOS + symmetry energy

$$\frac{E}{A}[\rho] = \frac{E}{A}[\rho_0] + \frac{1}{18} K_\infty \left(\frac{\rho - \rho_0}{\rho_0} \right)^2 + S \left(\frac{\rho_n - \rho_p}{\rho} \right)^2 + \dots$$

$$S = \frac{1}{2} \frac{\partial^2 (E/A)}{\partial \delta^2} \Bigg|_{\delta=0} = J + Lx + \frac{1}{2} K_{\text{sym}} x^2 + O(x^3),$$

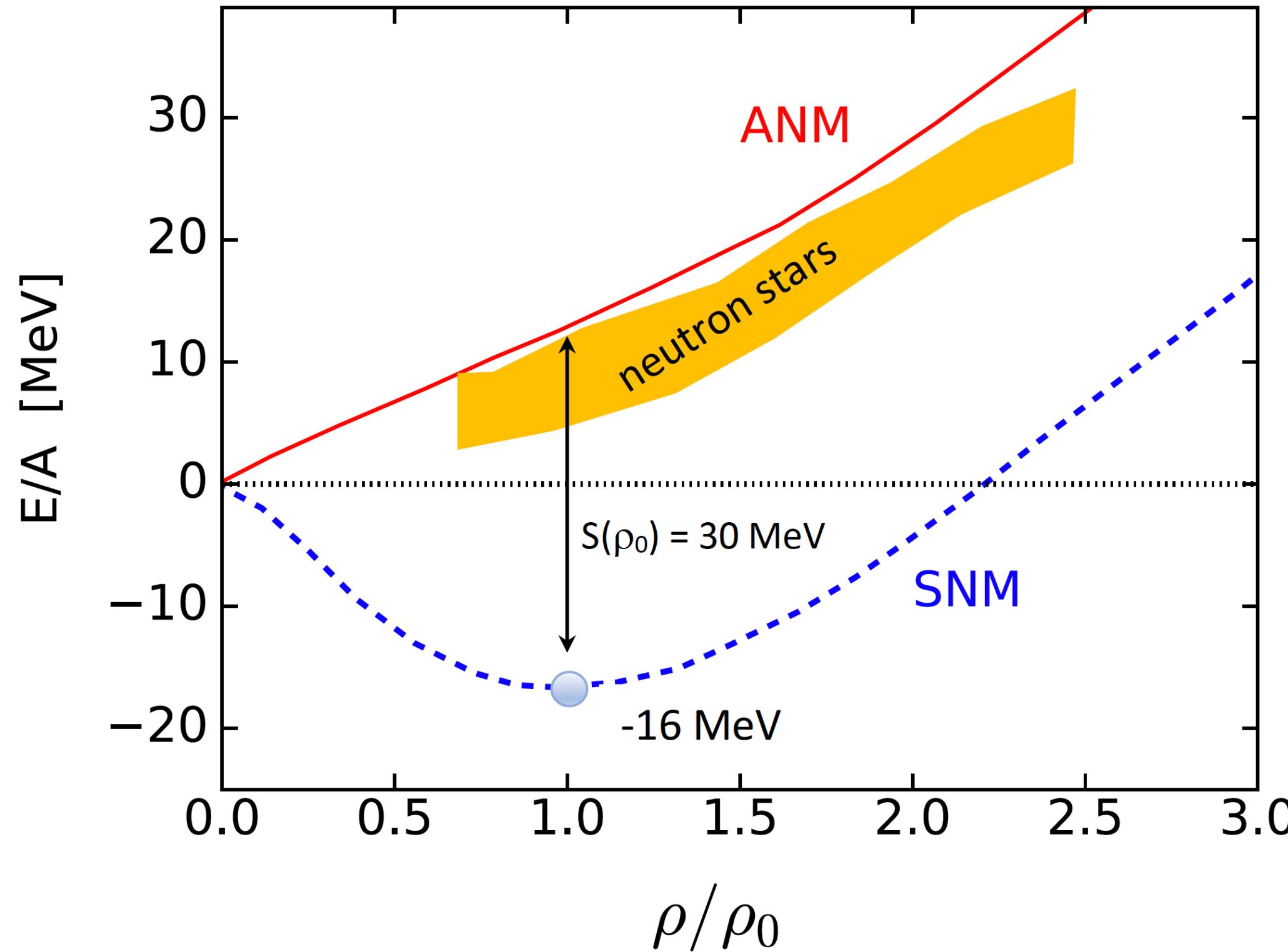
$$L = 3\rho_0 \frac{dS(\rho)}{d\rho} \Bigg|_{\rho_0}, \quad \delta = \frac{\rho_n - \rho_p}{\rho}, \quad x = \frac{(\rho - \rho_0)}{3\rho_0}$$

For $\rho \sim \rho_0$ and $\delta \sim 1 \Rightarrow p = \frac{L\rho_0}{3}$

Skyrme	ρ_0	E_0	K_∞	J	L	K_{sym}
SLy5	0.161	-15.99	229.92	32.01	48.15	-112.76
SkM*	0.160	-15.77	216.61	30.03	45.78	-155.94
Skxs20	0.162	-15.81	201.95	35.50	67.06	-122.31

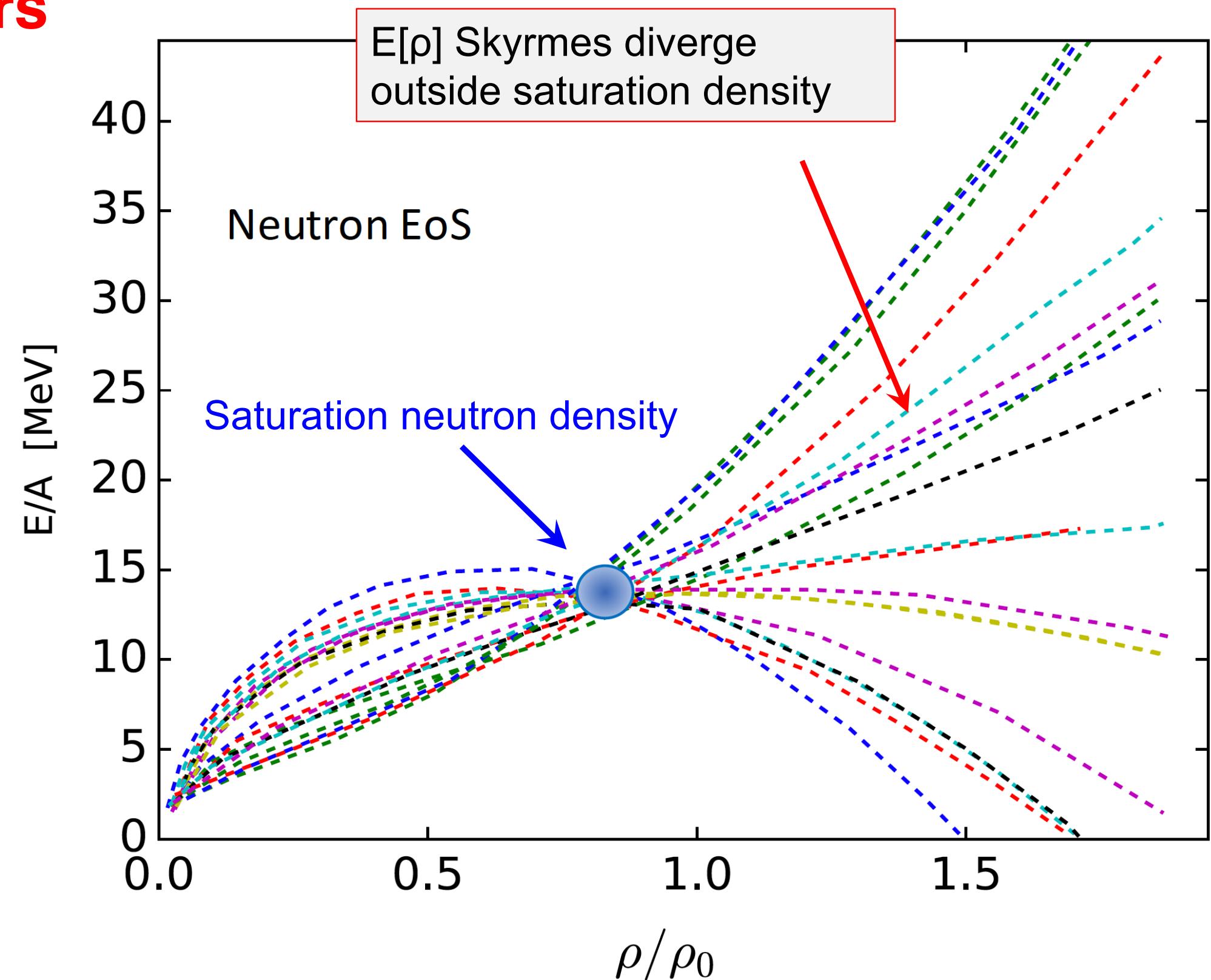
L crucial for neutron matter

EOS of neutron stars



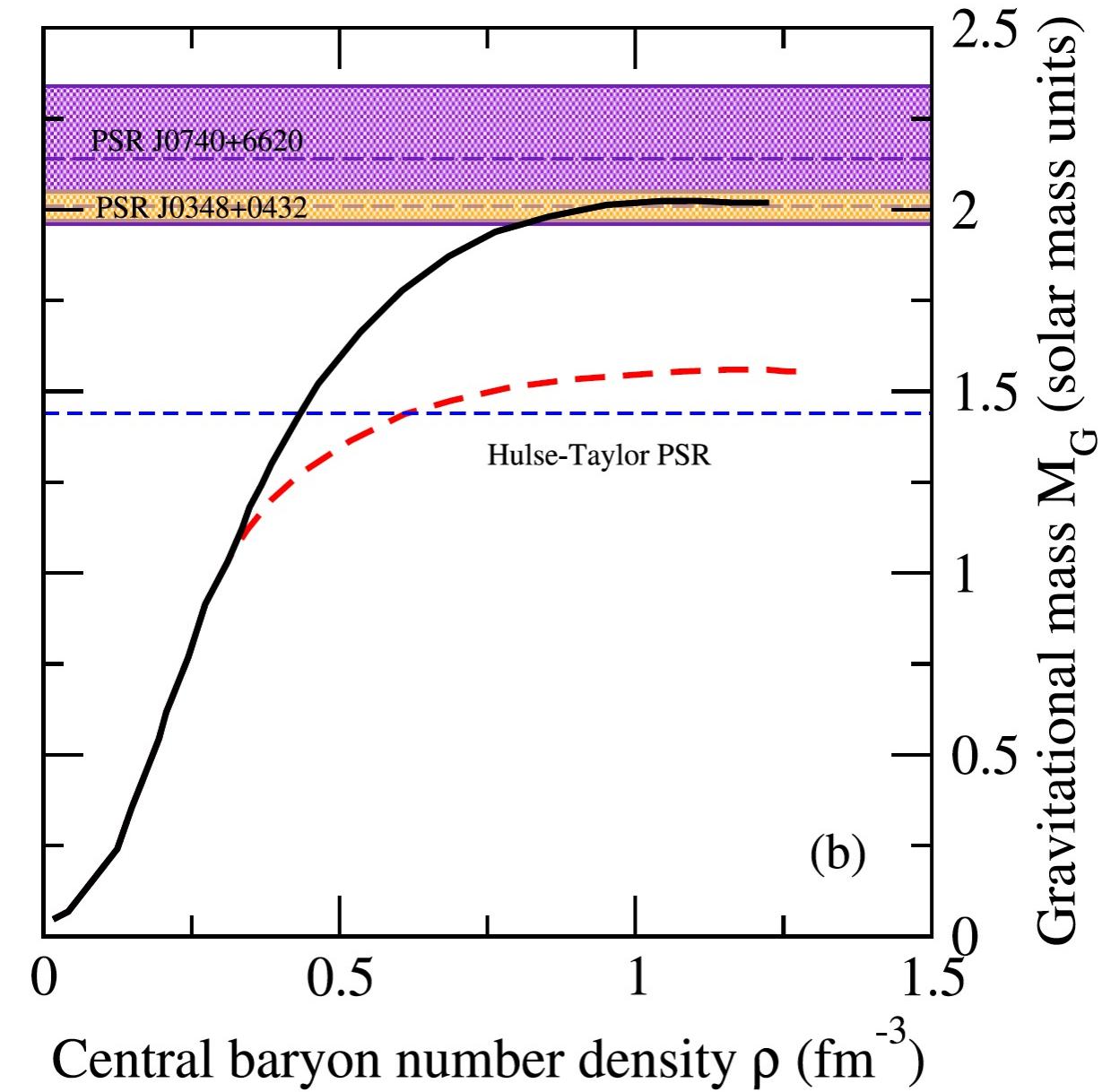
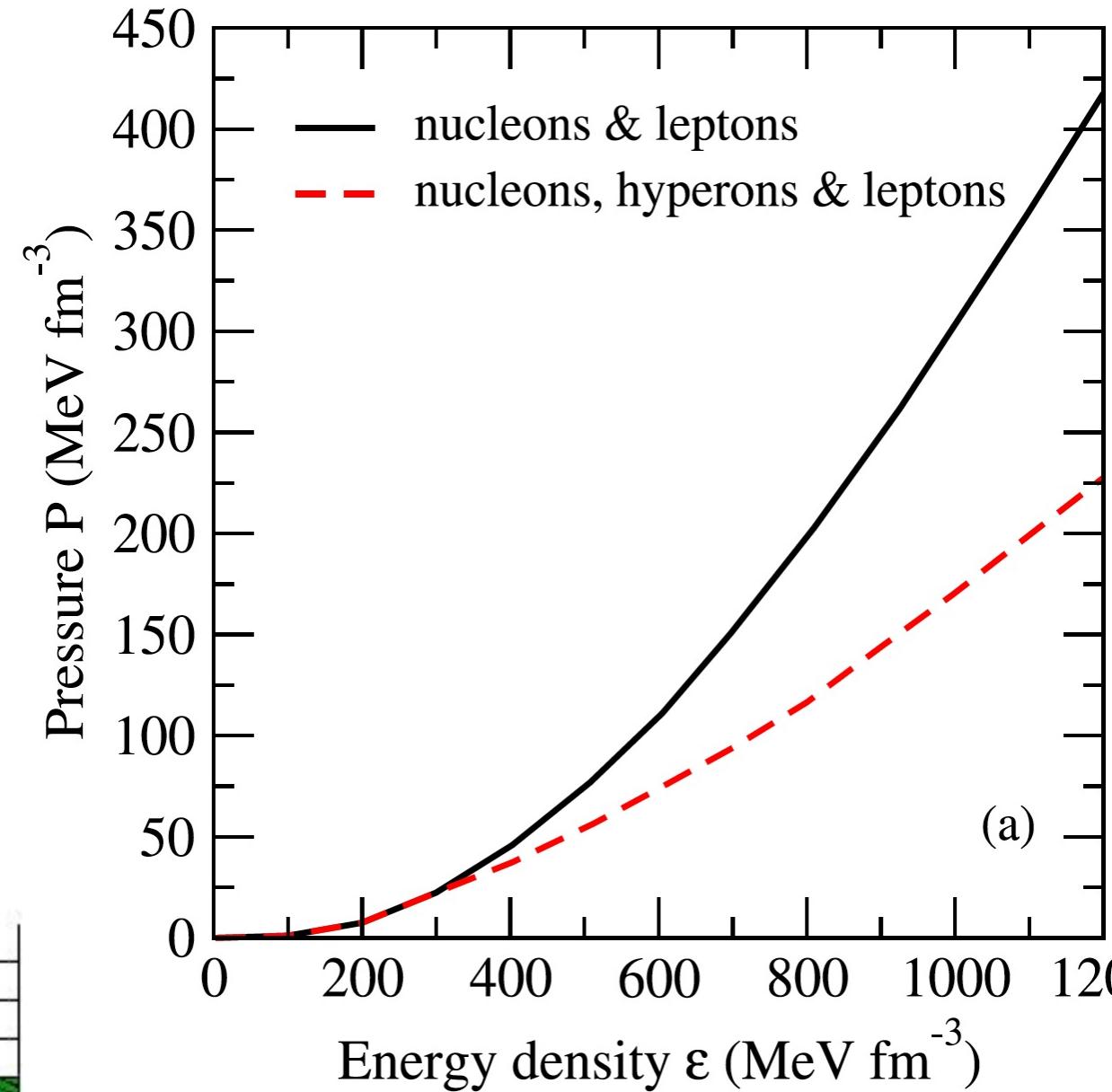
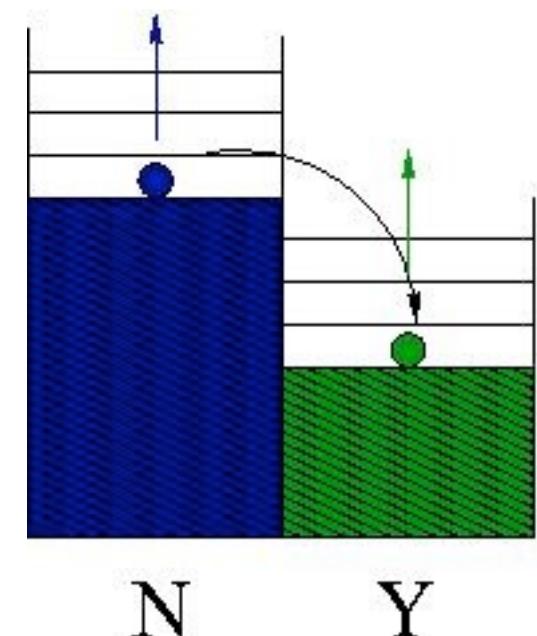
EOS & Neutron stars

Pethick, Ravenhall,
ARNPS 45 (1995) 429



Hyperons and neutron stars

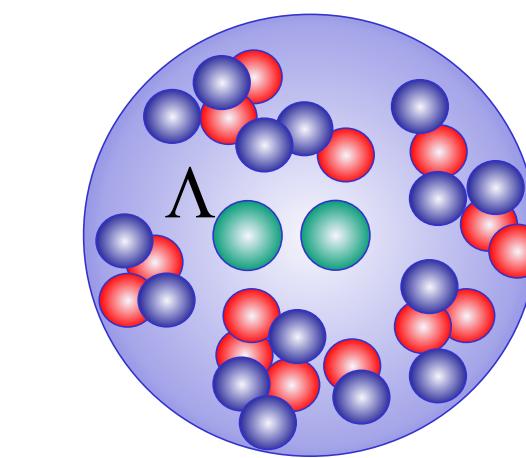
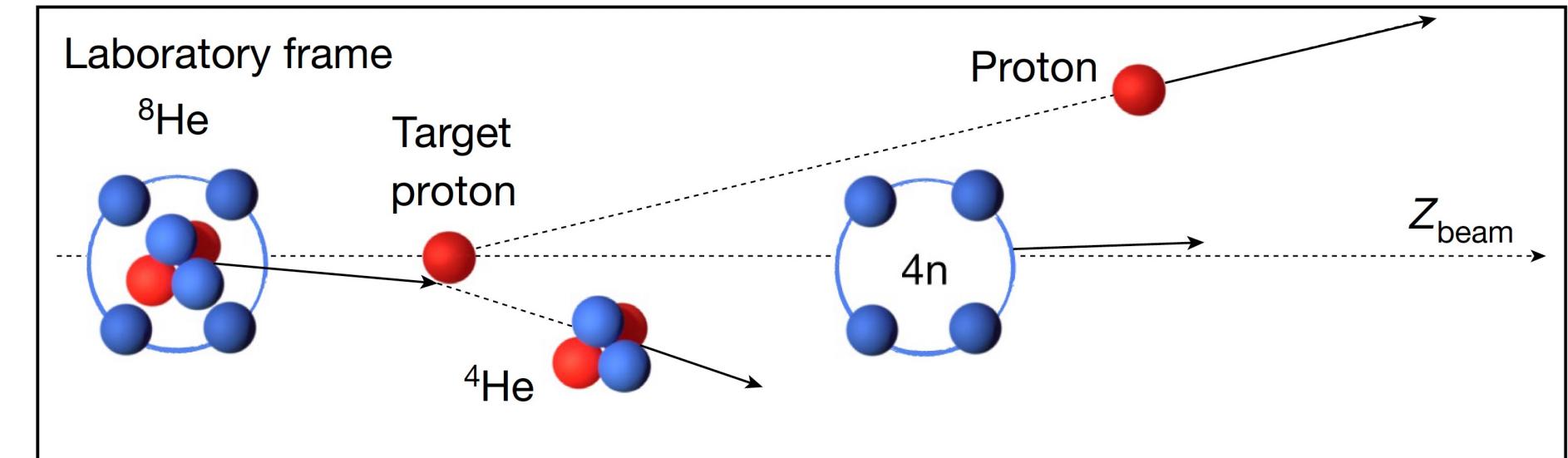
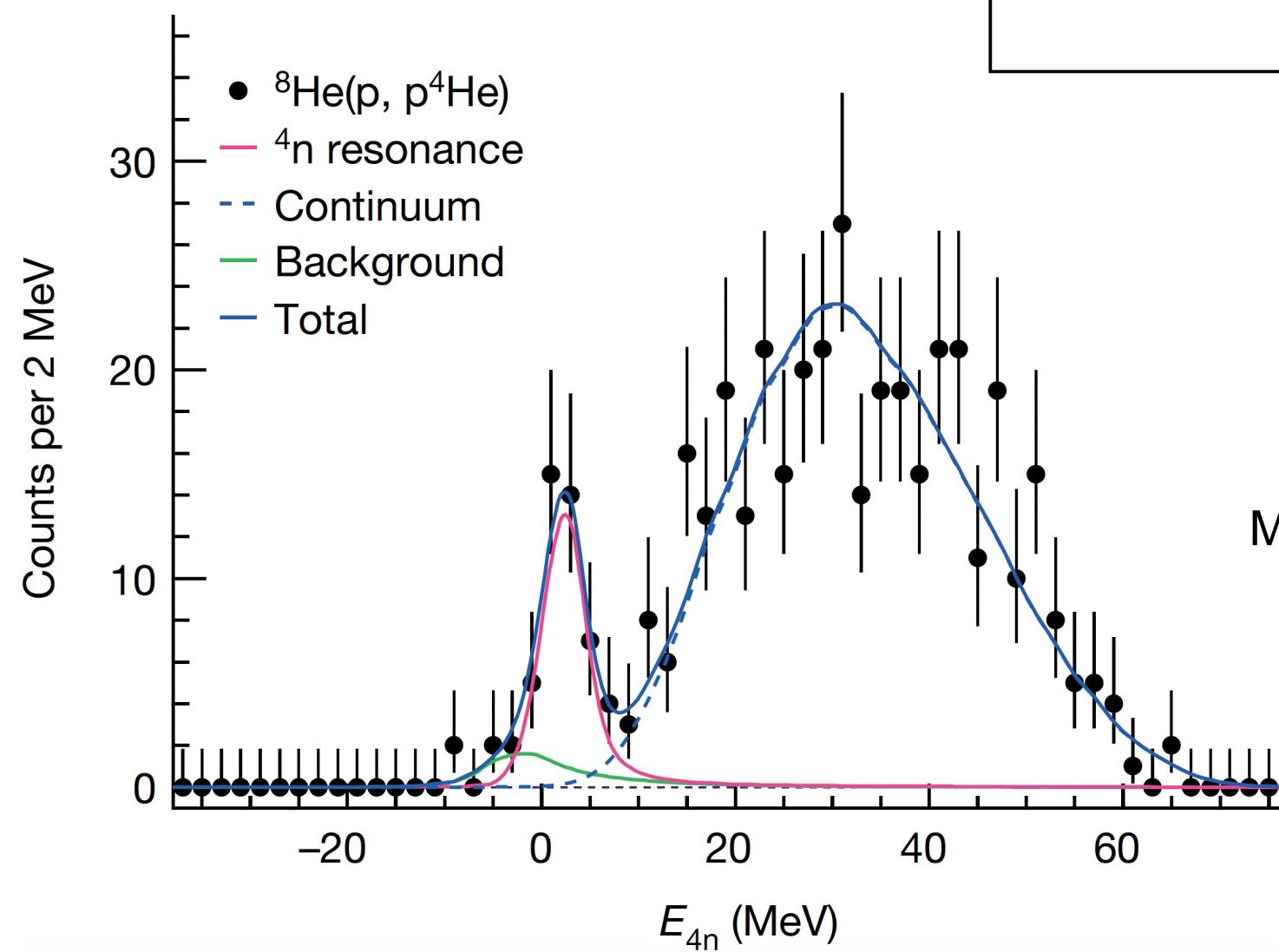
Hyperons make the EoS softer → reduction of the mass



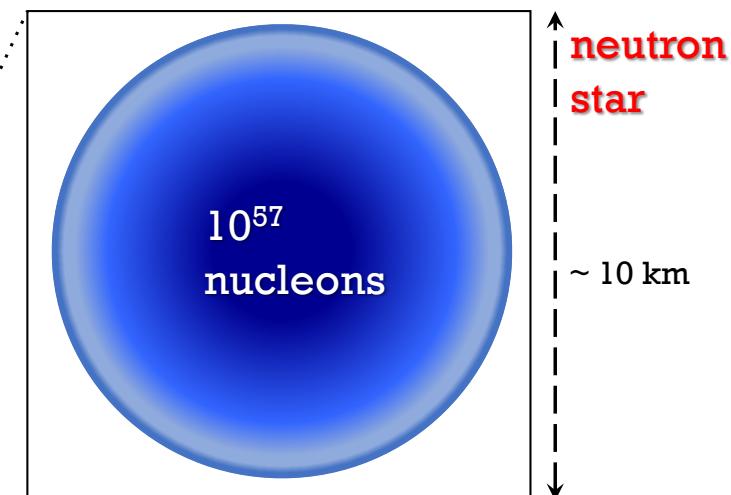
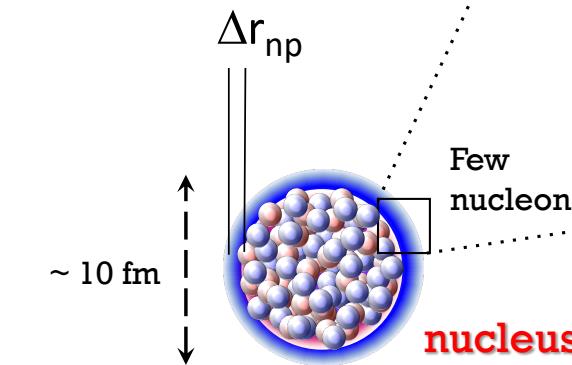
Hyperon “puzzle”

NN interactions and NΛ interactions

Duer, et al.,
Nature 606, 678 (2022)



Modified clusterization



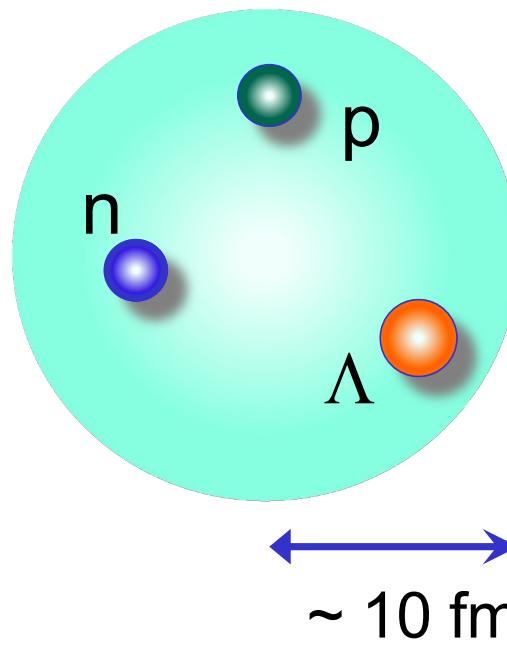
Hypertriton and ΛN interaction

$$\tau_\Lambda = 263 \text{ ps}$$

$$\tau_{\Lambda^3\text{H}} = 253 \pm 11(\text{stat.}) \pm 6(\text{sys.}) \text{ ps}$$

Philipp Eckert, et al., EPJ Web Conf. 271 (2022) 01006

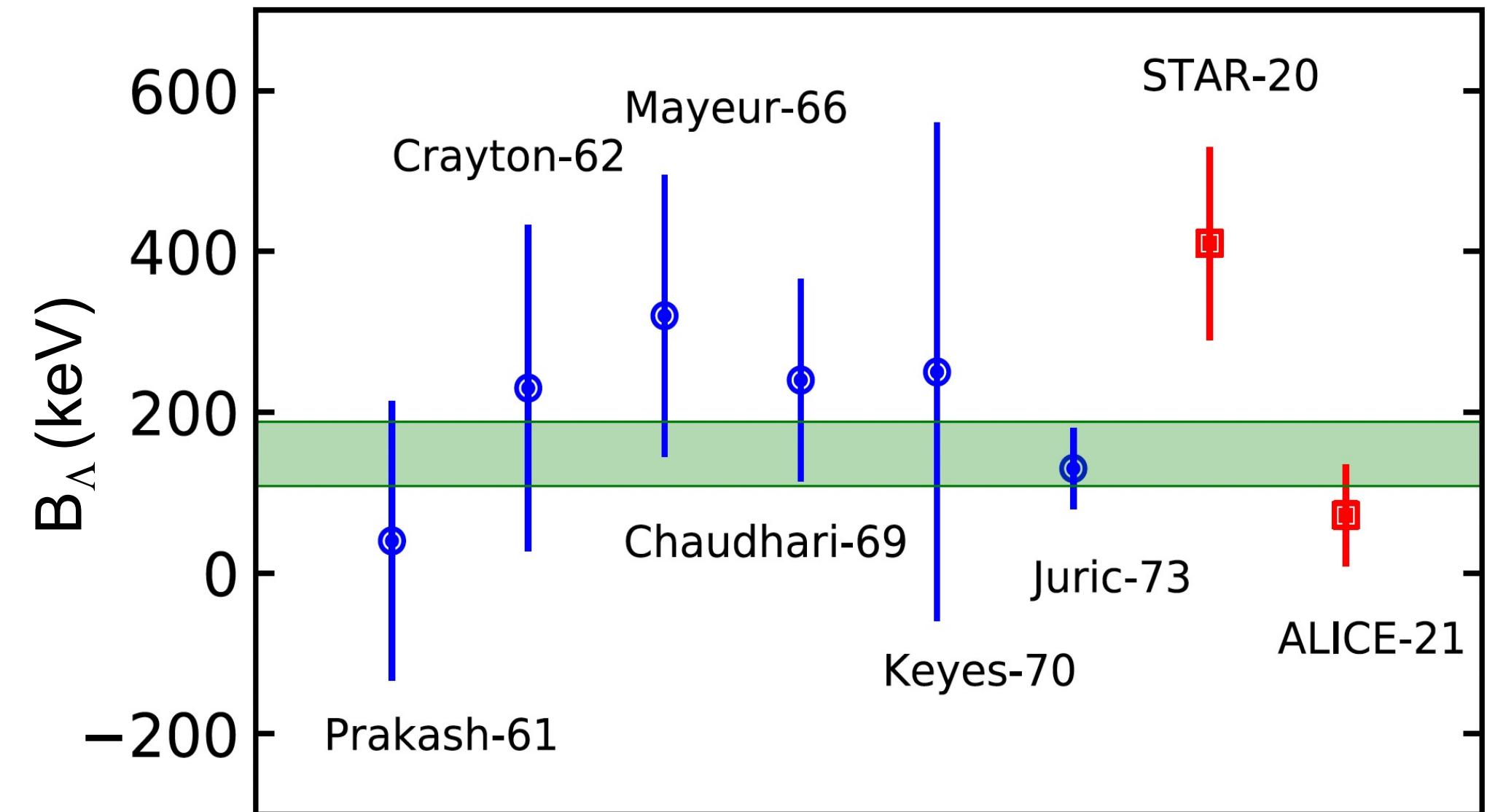
<https://hypernuclei.kph.uni-mainz.de>



Halo

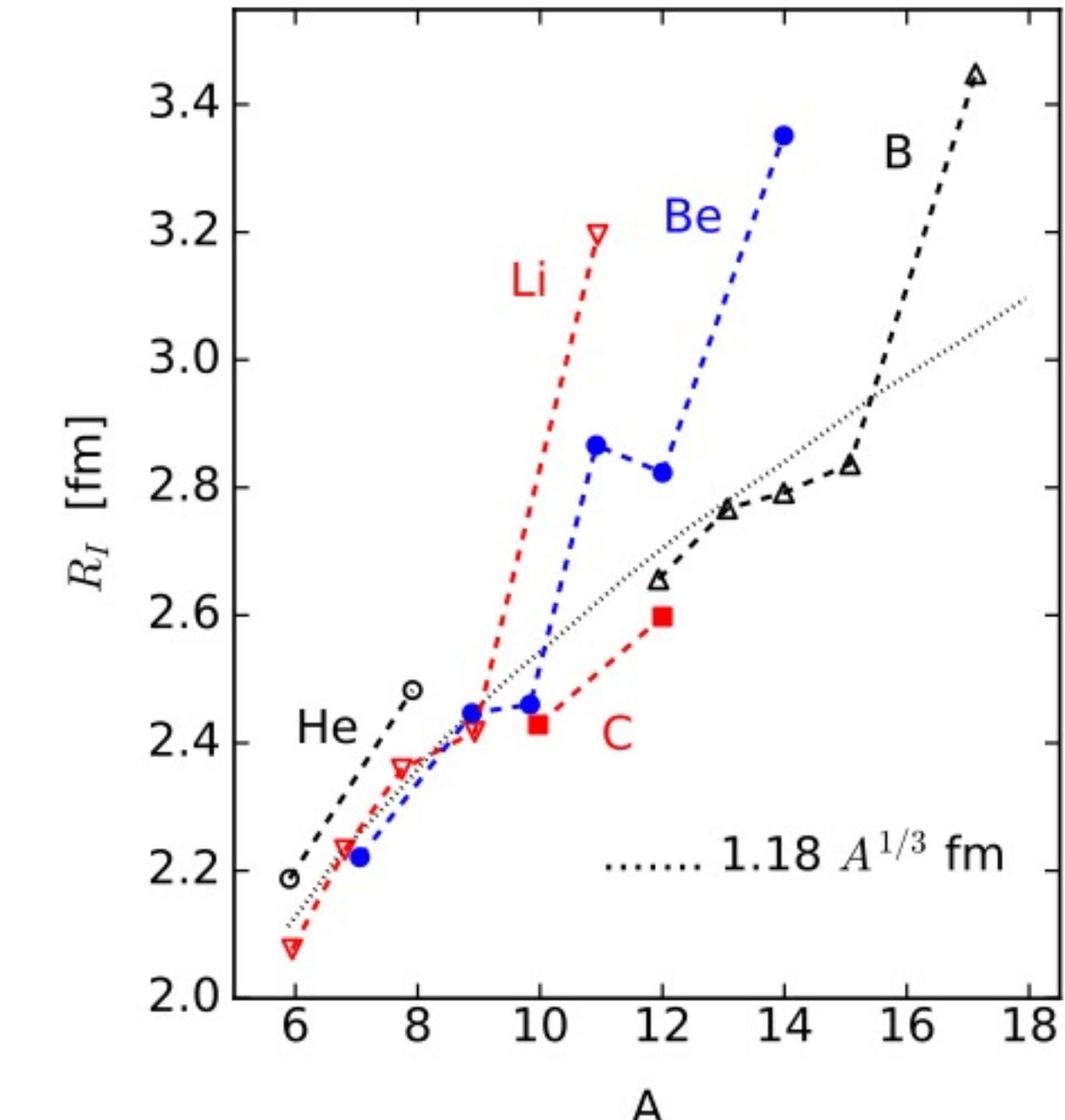
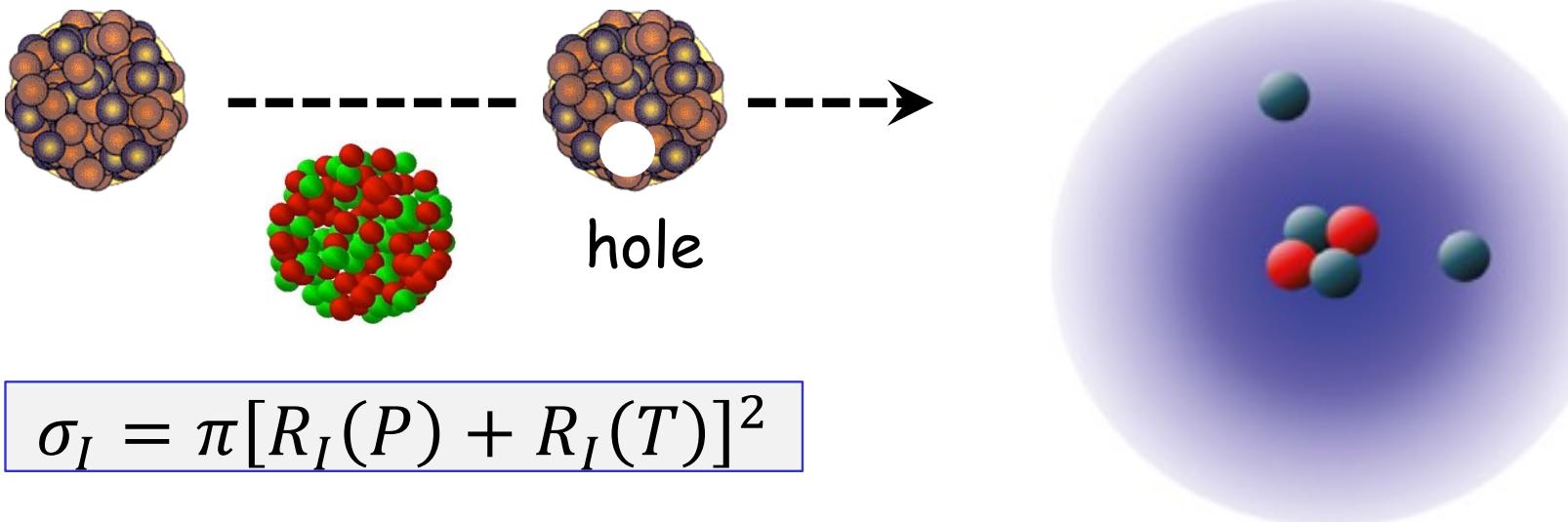
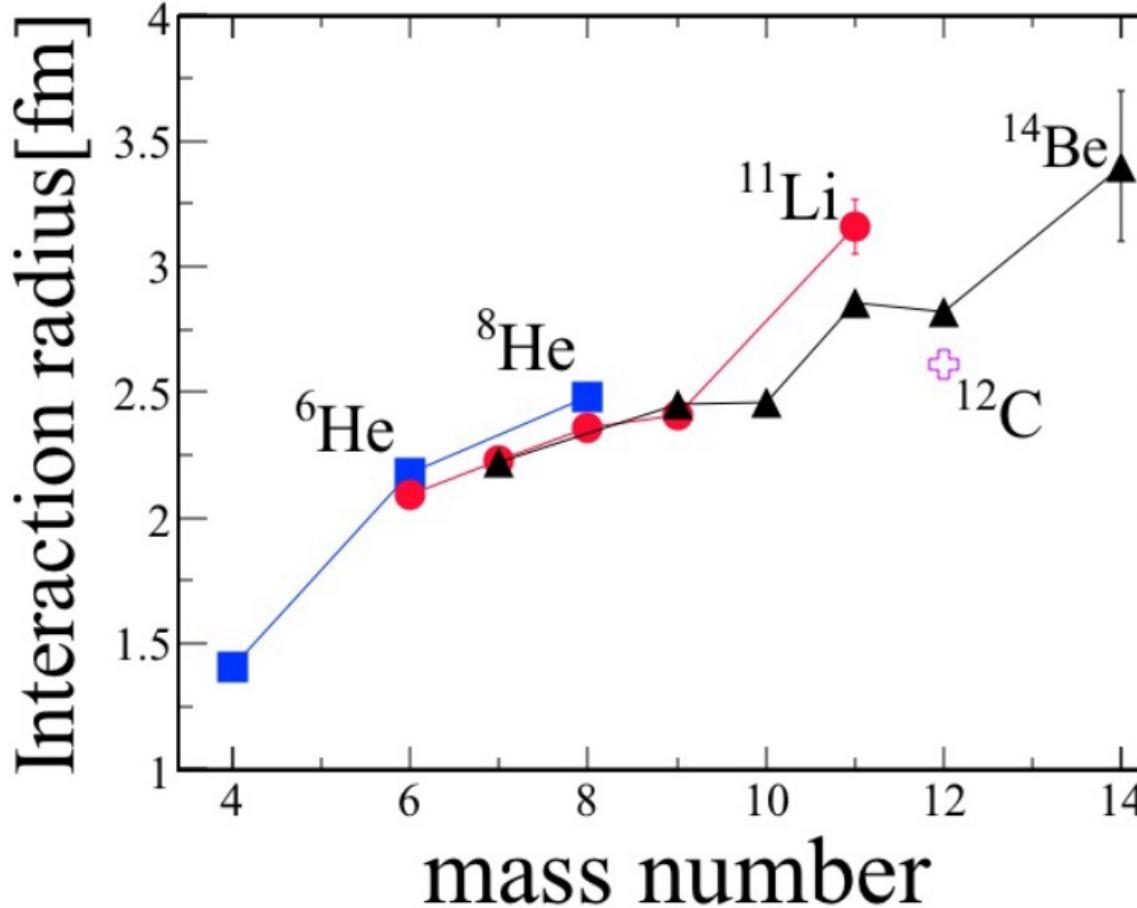
${}^6_\Lambda\text{He}$ ($S_n = 0.17 \text{ MeV}$)

${}^7_\Lambda\text{Be}$ ($S_{2p} = 0.67 \text{ MeV}$)

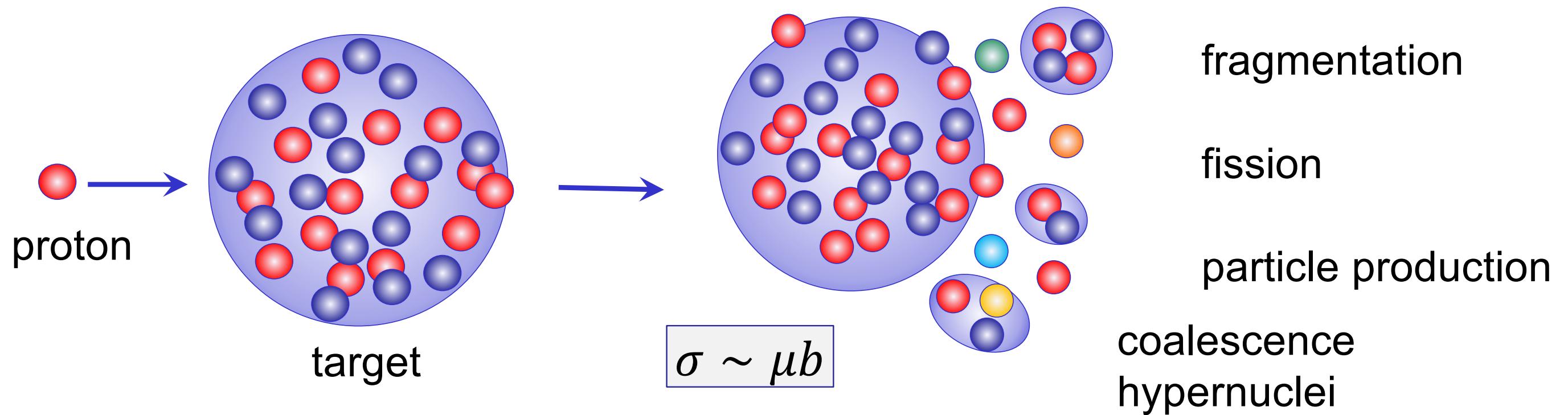


Nuclear halos and interaction radius

Tanihata et al.,
PRL 55, 2676 (1985)



Production of hypernuclei



$$C_A = \frac{2J_A + 1}{2^A} \frac{1}{\sqrt{A} m_T^{A-1}} \left(\frac{2\pi}{R^2 + \left(\frac{r_A}{2}\right)^2} \right)^{\frac{3}{2}(A-1)}$$

Particles produced coalesce into nuclei if they are close in space and momentum.

R = source size, r_A = nuclear size
 m_T = transverse mass of coalesc. part.

Production & Decay of Hypertriton

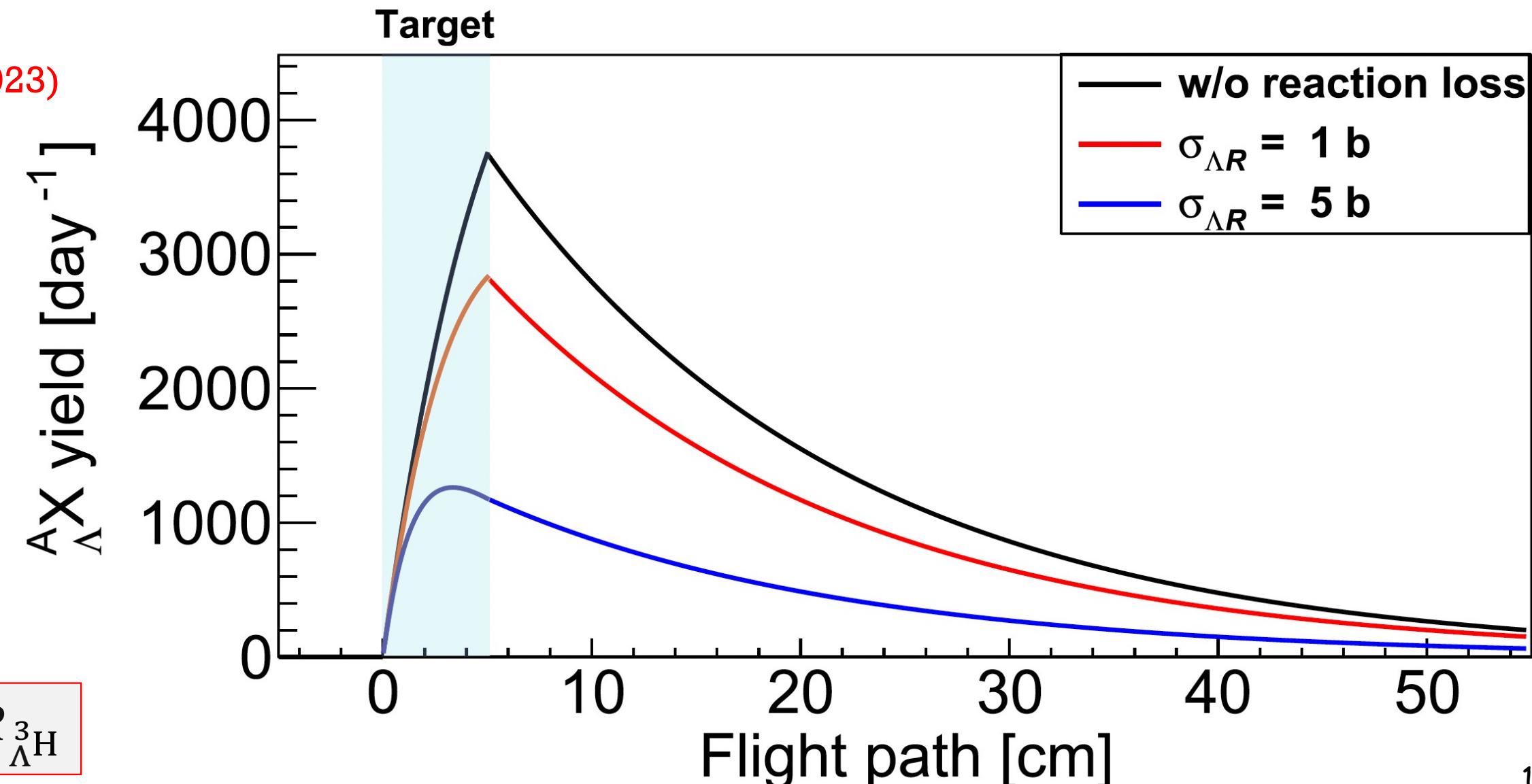
Use active target to:

- (a) Produce hypertriton
- (b) Reconstruct hypertriton by measuring weak decay (${}^3_{\Lambda}\text{H} \rightarrow \pi^- + {}^3\text{He}$)
- (c) Interaction cross section through meson decay vertex distribution

Velardita et al.,

Eur. Phys. J. A 59, 139 (2023)

$$\sigma_I = \sigma_{\Lambda R} \sim 1.8 \text{ b}$$
$$\tau \sim 200 \text{ ps}$$

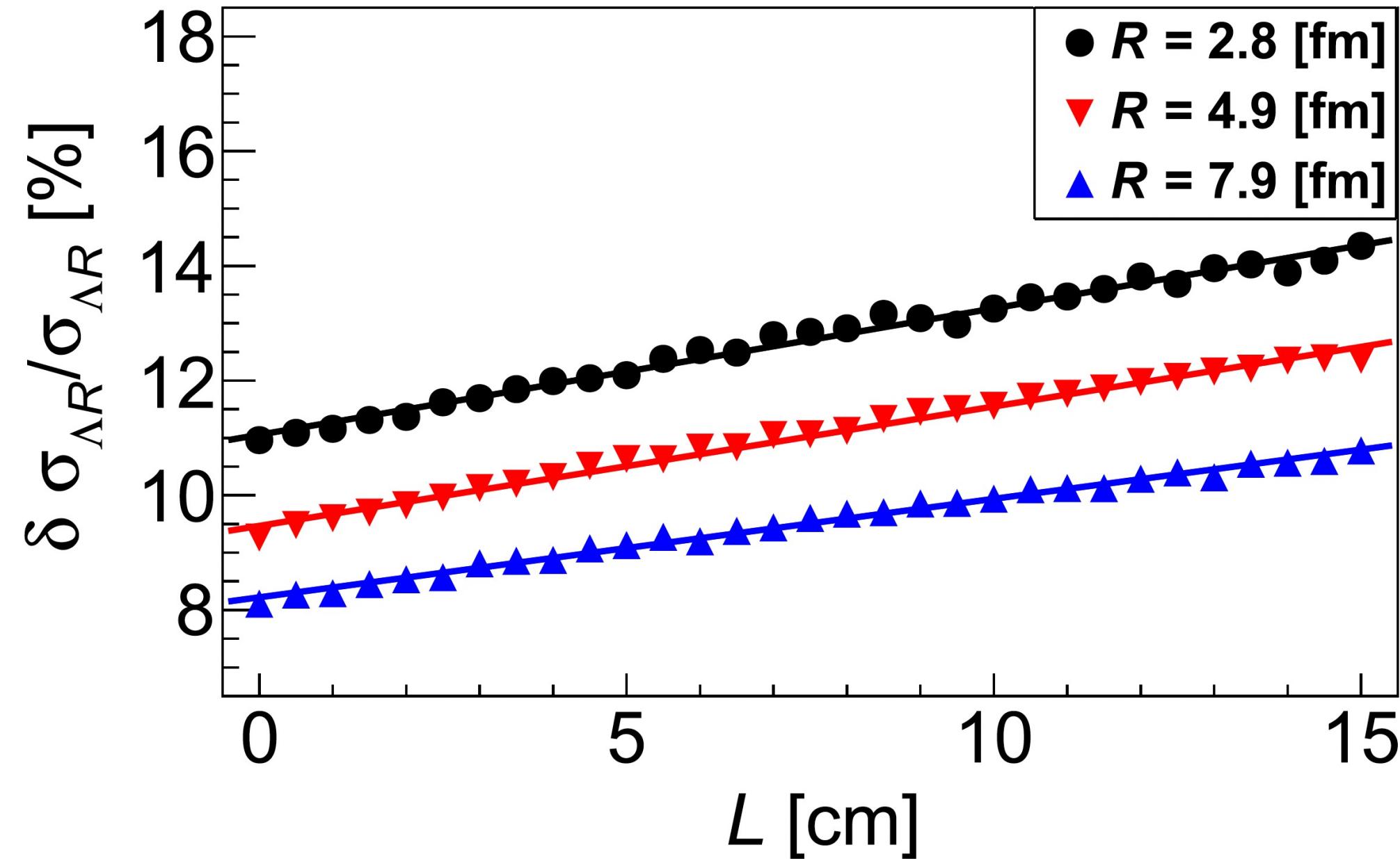


Experiment $\rightarrow \sigma_I \rightarrow R_{{}^3\text{H}}$

Extraction of Hypertriton radius

Velardita et al.,
Eur. Phys. J. A 59, 139 (2023)

$^{12}\text{C} + ^{12}\text{C}$ collisions
at 1.9 GeV/nucleon
 L = active target width

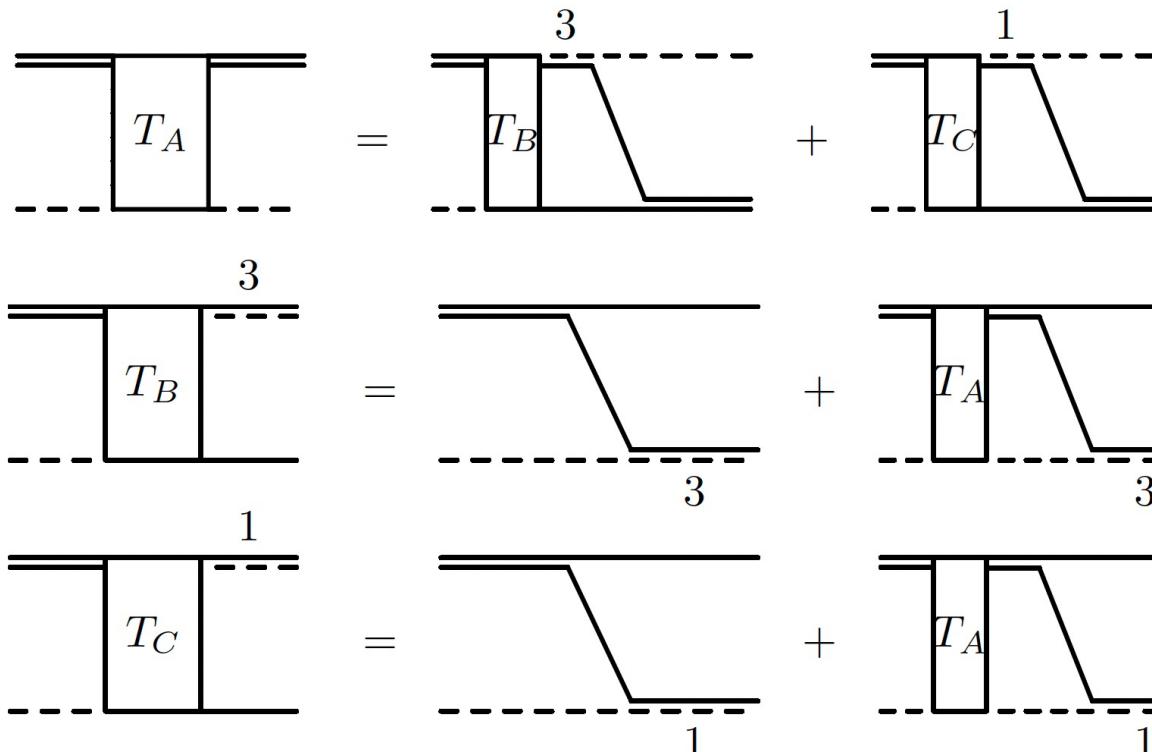


Hypertriton wavefunction

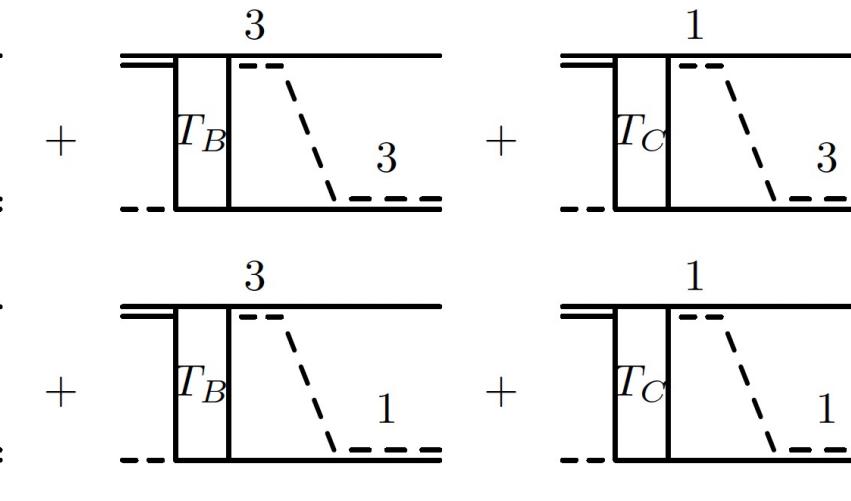
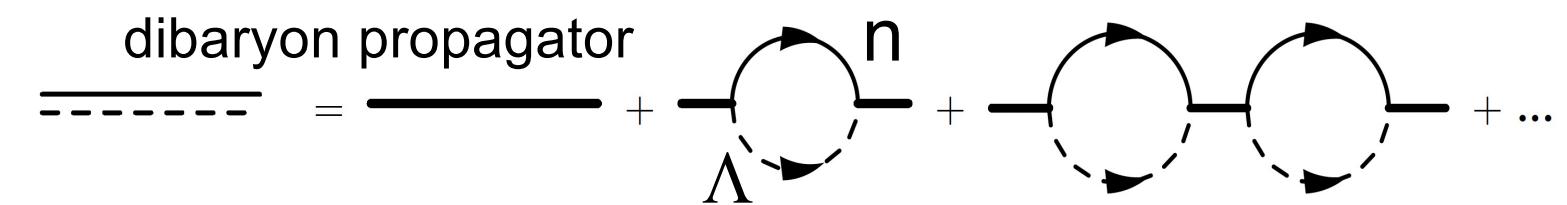
Pionless EFT

- (a) Momentum scales < pion mass
- (b) Dibaryon field Δ_d
- (c) Cutoff parameter Λ_c
- (d) Simplified 3-body force $H(s_0, \Lambda_c)$
- (e) Asymptotic analysis

Hildenbrand, Hammer,
Phys. Rev. C, 100, 034002 (2019)



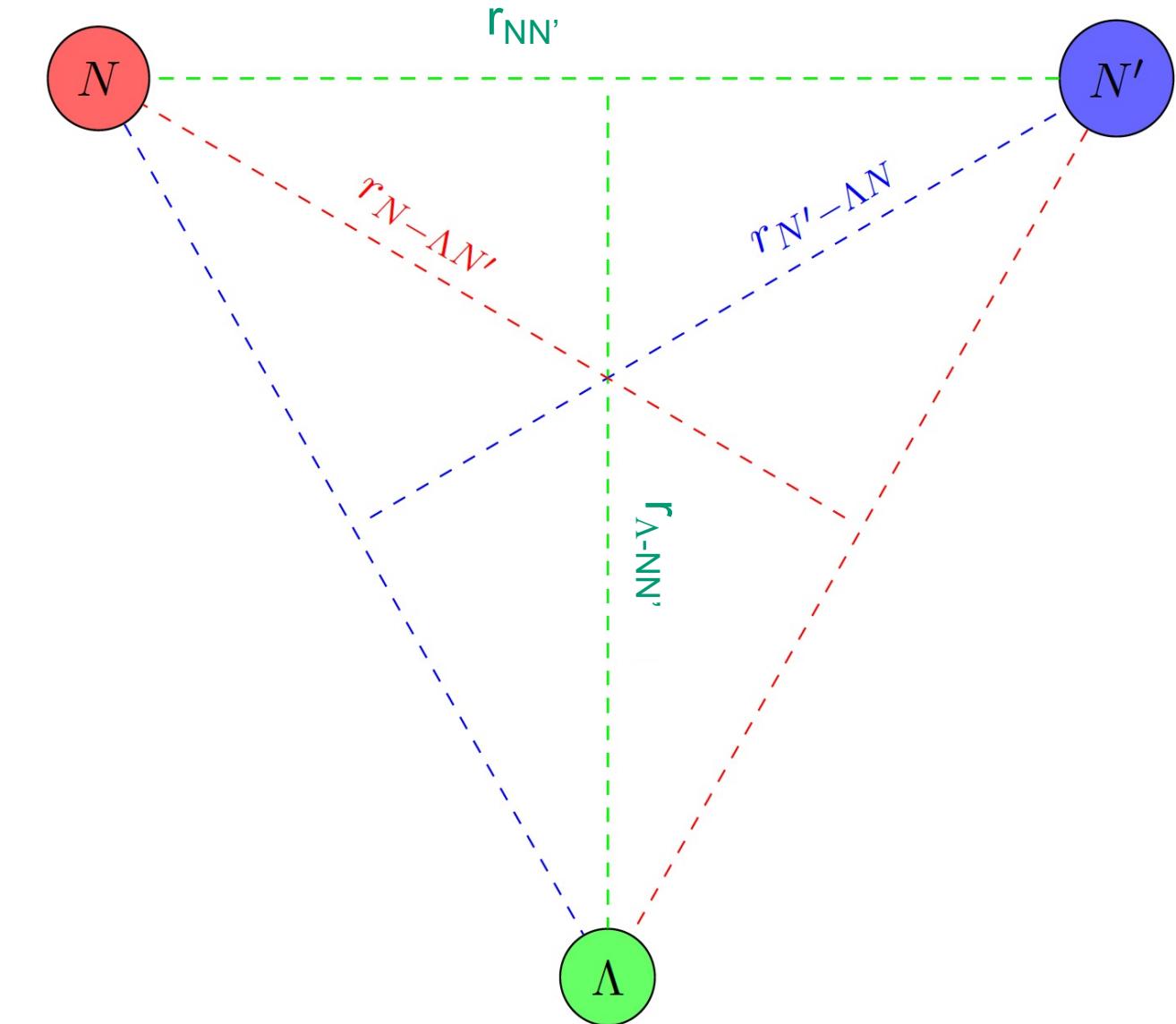
$$\begin{aligned} \mathcal{L} = & N^\dagger \left(i\partial_t + \frac{\nabla^2}{2M} \right) N + \Lambda^\dagger \left(i\partial_t + \frac{\nabla^2}{2M_\Lambda} \right) \Lambda \\ & + \Delta_d d_l^\dagger d_l - \frac{g_d}{2} \left[d_l^\dagger N^T (i\tau_2) (i\sigma_l \sigma_2) N + \text{H.c.} \right] \\ & + \Delta_s s_j^\dagger s_j - \frac{g_s}{2} \left[s_j^\dagger N^T (i\tau_j \tau_2) (i\sigma_2) N + \text{H.c.} \right] \\ & + \Delta_3 (u_l^3)^\dagger u_l^3 - g_3 \left[i (u_l^3)^\dagger \Lambda^T (i\sigma_l \sigma_2) N + \text{H.c.} \right] \\ & + \Delta_1 (u^1)^\dagger u^1 - g_1 \left[i (u^1)^\dagger \Lambda^T (i\sigma_2) N + \text{H.c.} \right] + \dots \end{aligned}$$



Hypertriton wavefunction (EFT)

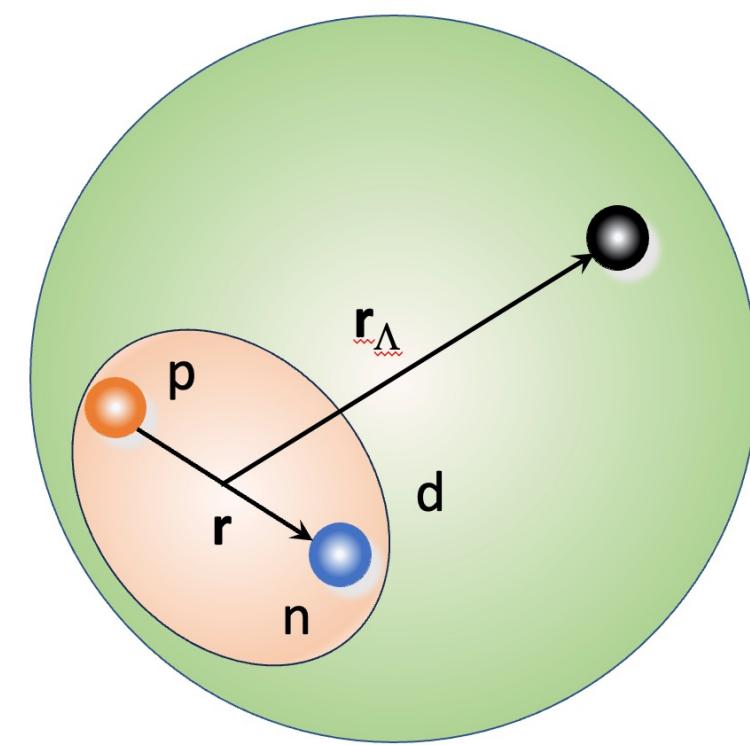
$\sqrt{\langle r_{\Lambda-NN'}^2 \rangle} [\text{fm}]$	$\sqrt{\langle r_{N'-\Lambda N}^2 \rangle} [\text{fm}]$	$\sqrt{\langle r_{NN'}^2 \rangle} [\text{fm}]$
10.79	3.96	2.96
+3.04/-1.53	+0.40/-0.25	+0.06/-0.05
+0.03/-0.02	+0.03/-0.03	+0.03/-0.04

Confirms that the "picture" as a **two body system** consisting of a deuteron and a Λ is a good approximation.



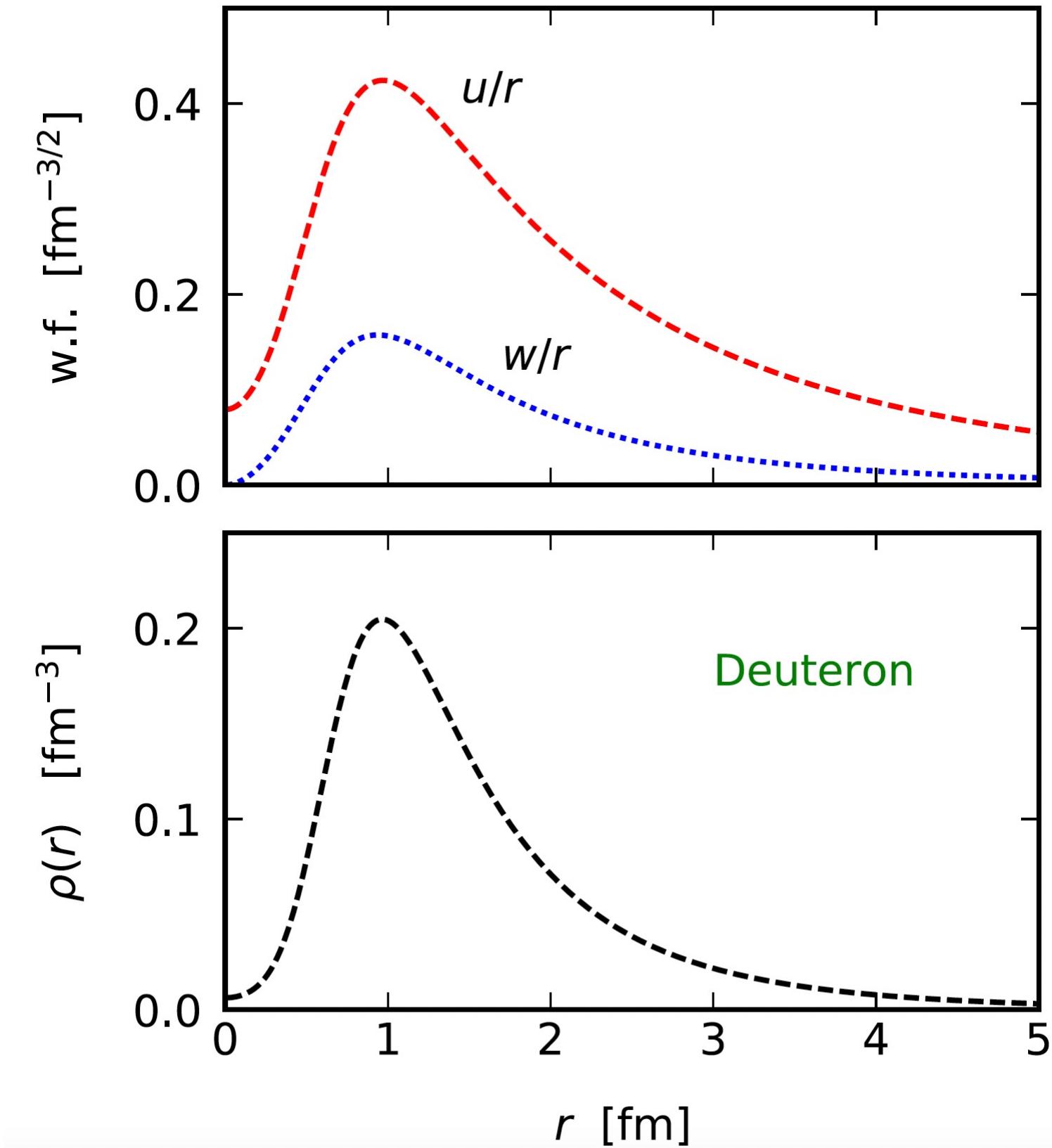
Congleton, JPG 18, 339 (1992)
+ many others

Hypertriton wavefunction



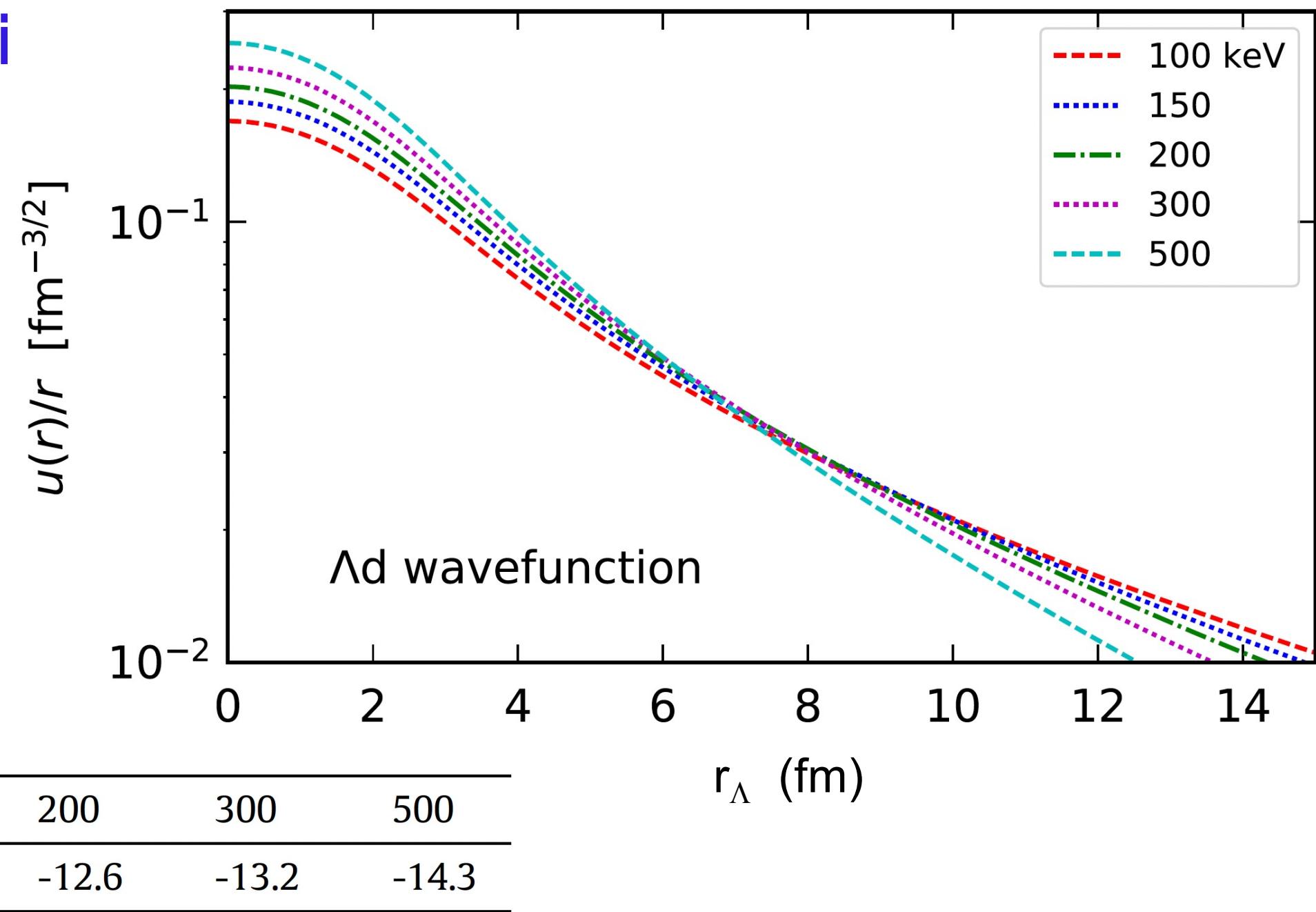
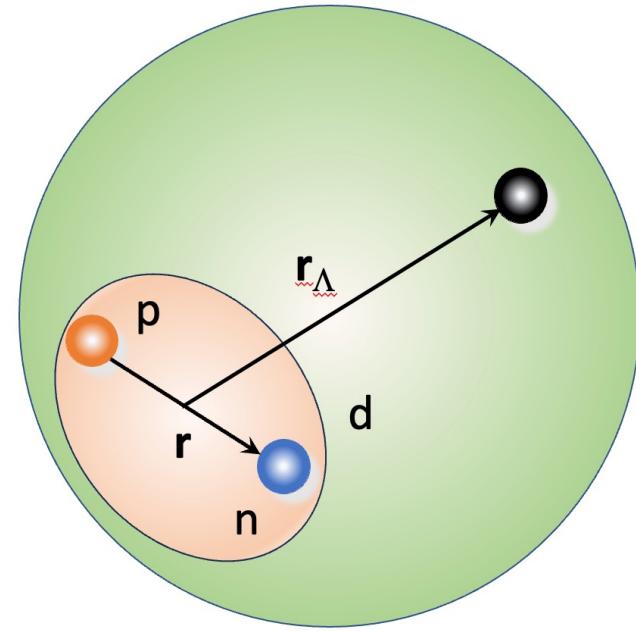
Deuteron radial s-wave, $u(r)/r$, and d-wave, $w(r)/r$ as a function of the proton-neutron distance r using Av18 interaction.

Wiringa, Stoks, Schiavilla
Phys. Rev. C 51, 38 (1995)



Hypertriton wavefunction

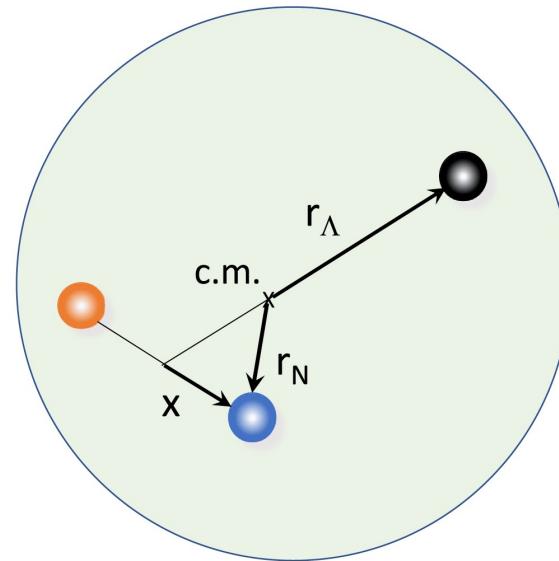
CB, PLB 837, 137639 (2023)



Λ -deuteron wavefunction - Woods-Saxon, radius $R = 2.5\text{fm}$, diffuseness $a = 0.65 \text{ fm}$.

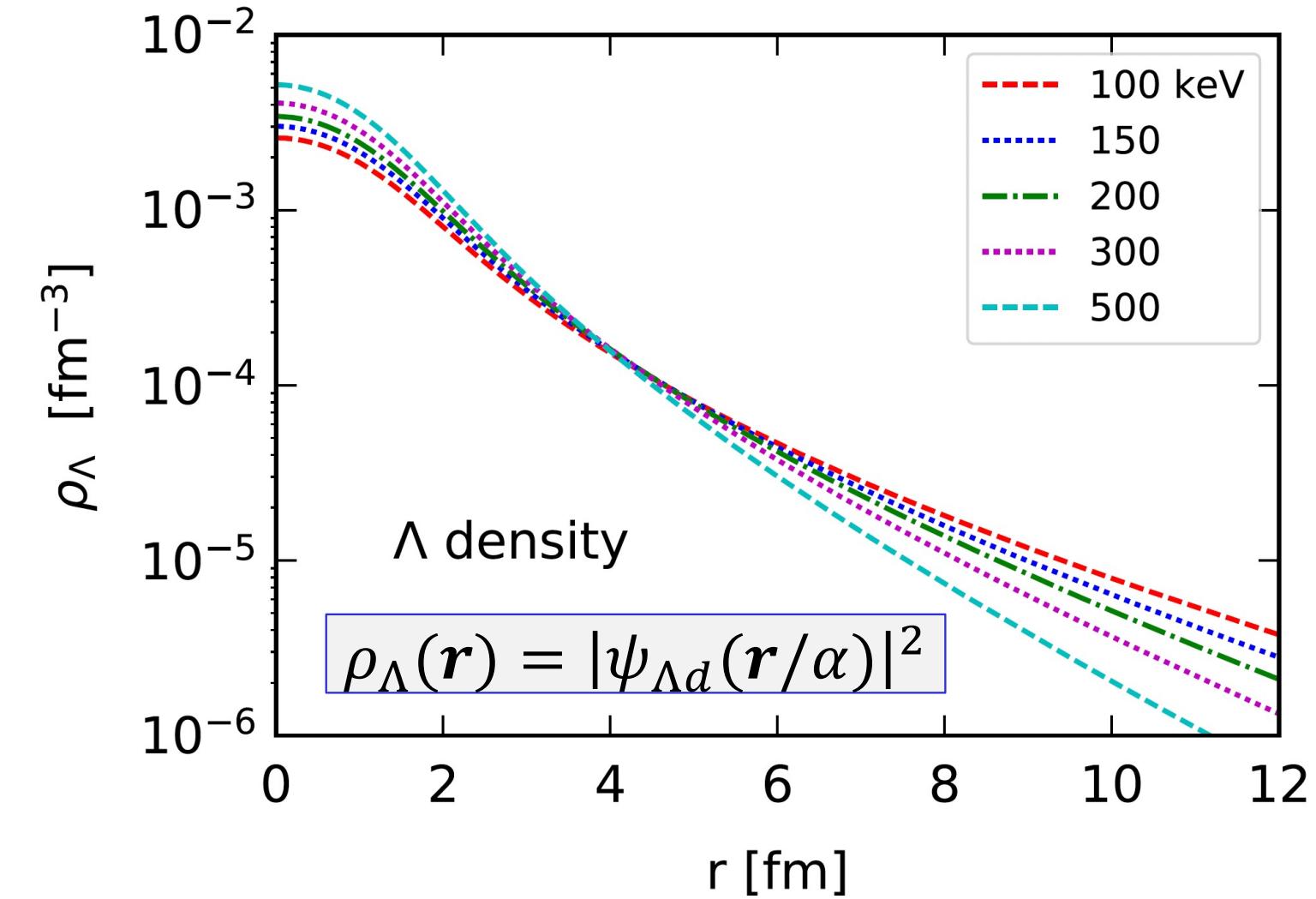
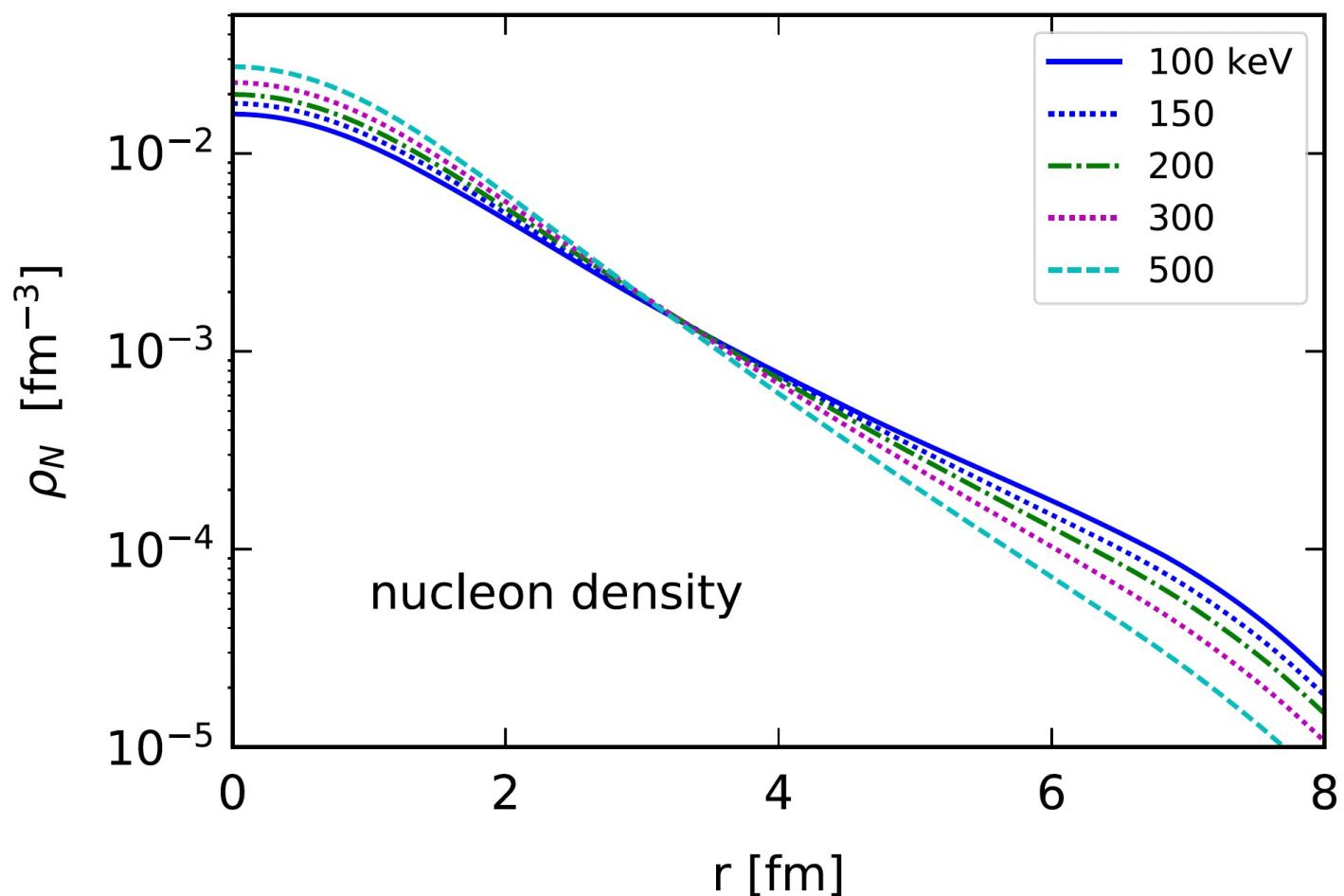
$u(r)/r$ for the s-wave and binding energies $B = 100, 150, 200, 300$ and 500 keV .

Hypertriton + nucleon density



$$\alpha = \frac{m_d}{m_\Lambda + m_d}$$

$$\beta = \frac{m_\Lambda}{m_\Lambda + m_d}$$



$$\rho_1(\mathbf{r}_\Lambda) = |\psi_{\Lambda d}(\mathbf{r}/\beta)|^2 \quad \text{Lambda dist. within } {}^3_\Lambda\text{H}$$

$$\rho_2(\mathbf{r}_p) = |\psi_{deut}^{free}(2x)|^2 \quad \text{p dist. within deut.}$$

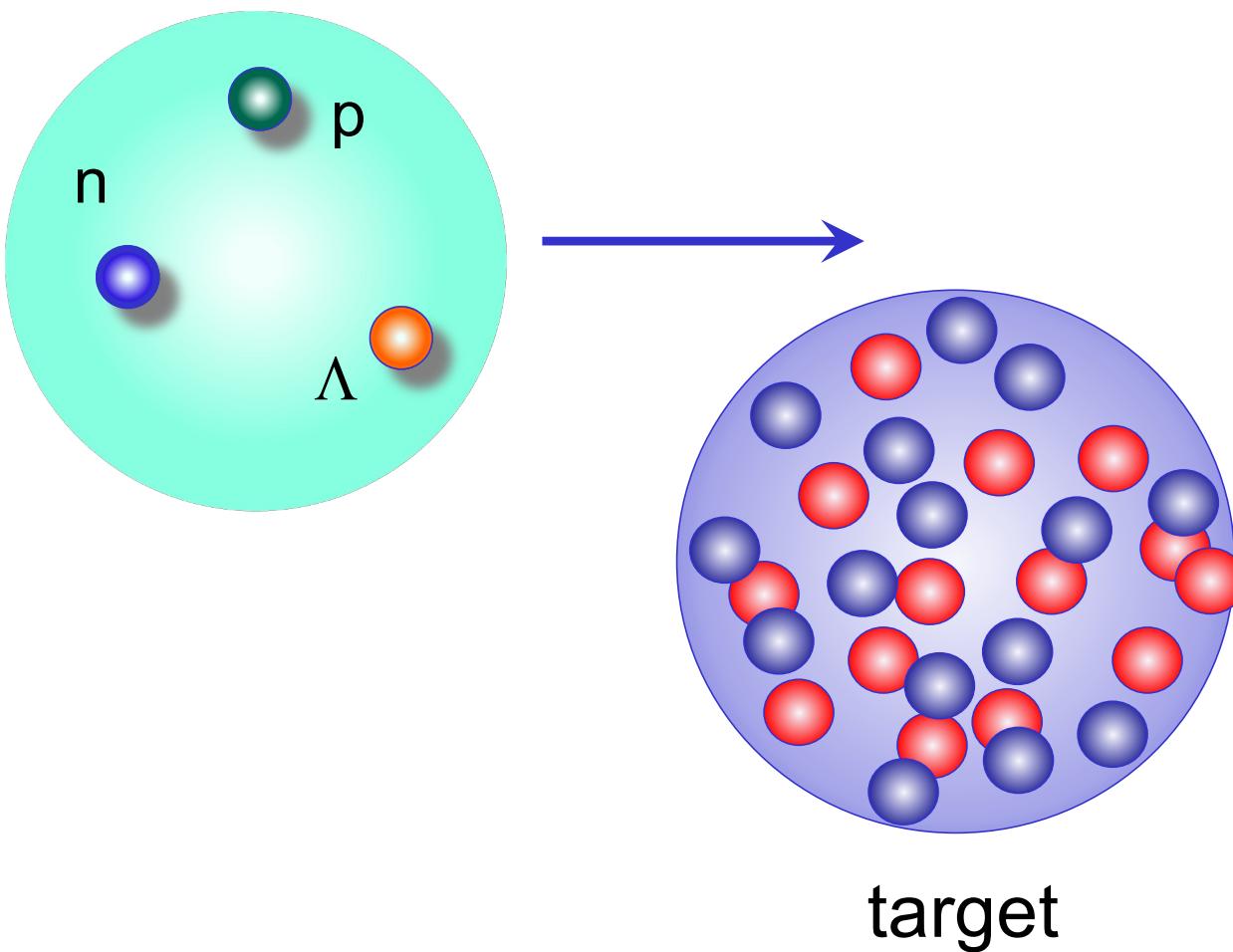
$$\rho_N(\mathbf{r}_N) = \int \rho_1(|\mathbf{r} - \mathbf{x}|) \rho_2(\mathbf{x})$$

nucleon distribution within ${}^3_\Lambda\text{H}$

Hypertriton destruction

$$T(b) = T_{\Lambda}(b) \ T_p(b) \ T_n(b)$$

Transmission probability



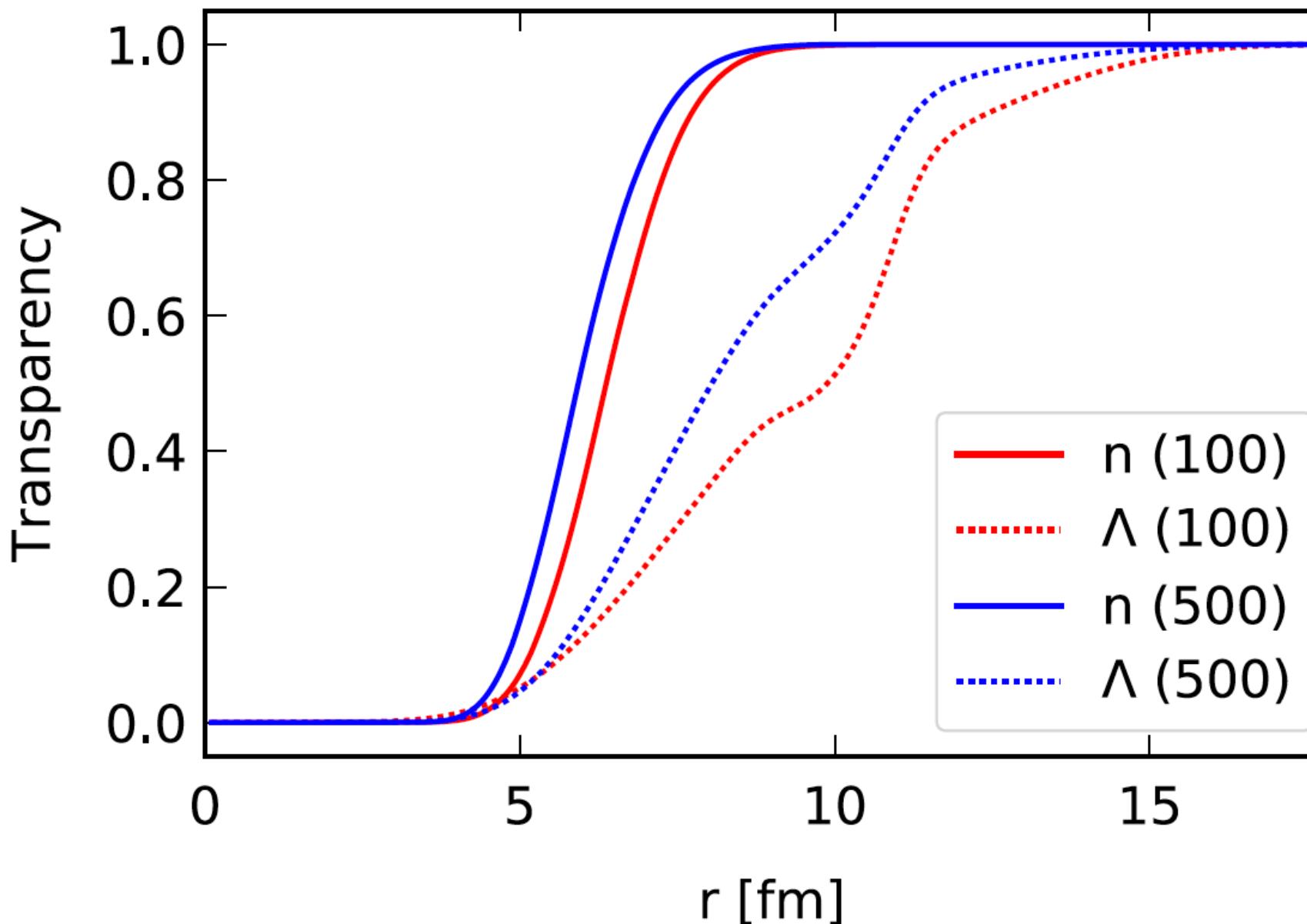
$$\begin{aligned} T_i(b) &= \int d^2 s_i dz_i(z_i, \mathbf{s}_i - \mathbf{b}) \\ &= \exp \left[-\sigma_{pi} Z_T \int dz' \rho_p^T(z', \mathbf{s}) \right] \\ &= \exp \left[-\sigma_{ni} N_T \int dz' \rho_n^T(z', \mathbf{s}) \right] \end{aligned}$$

$i = \Lambda, p, \text{ or } n$

$$\begin{aligned} &1.5 \text{ GeV/nucleon } {}^3\text{H incident on } {}^{12}\text{C}, {}^{120}\text{Sn}, {}^{208}\text{Pb} \\ &\sigma_{np} = 45.8 \text{ mb} \\ &\sigma_{pp} = 40 \text{ mb} \\ &\sigma_{\Lambda p} = 35 \text{ mb} \end{aligned}$$

Hypertriton destruction

Transmission probability



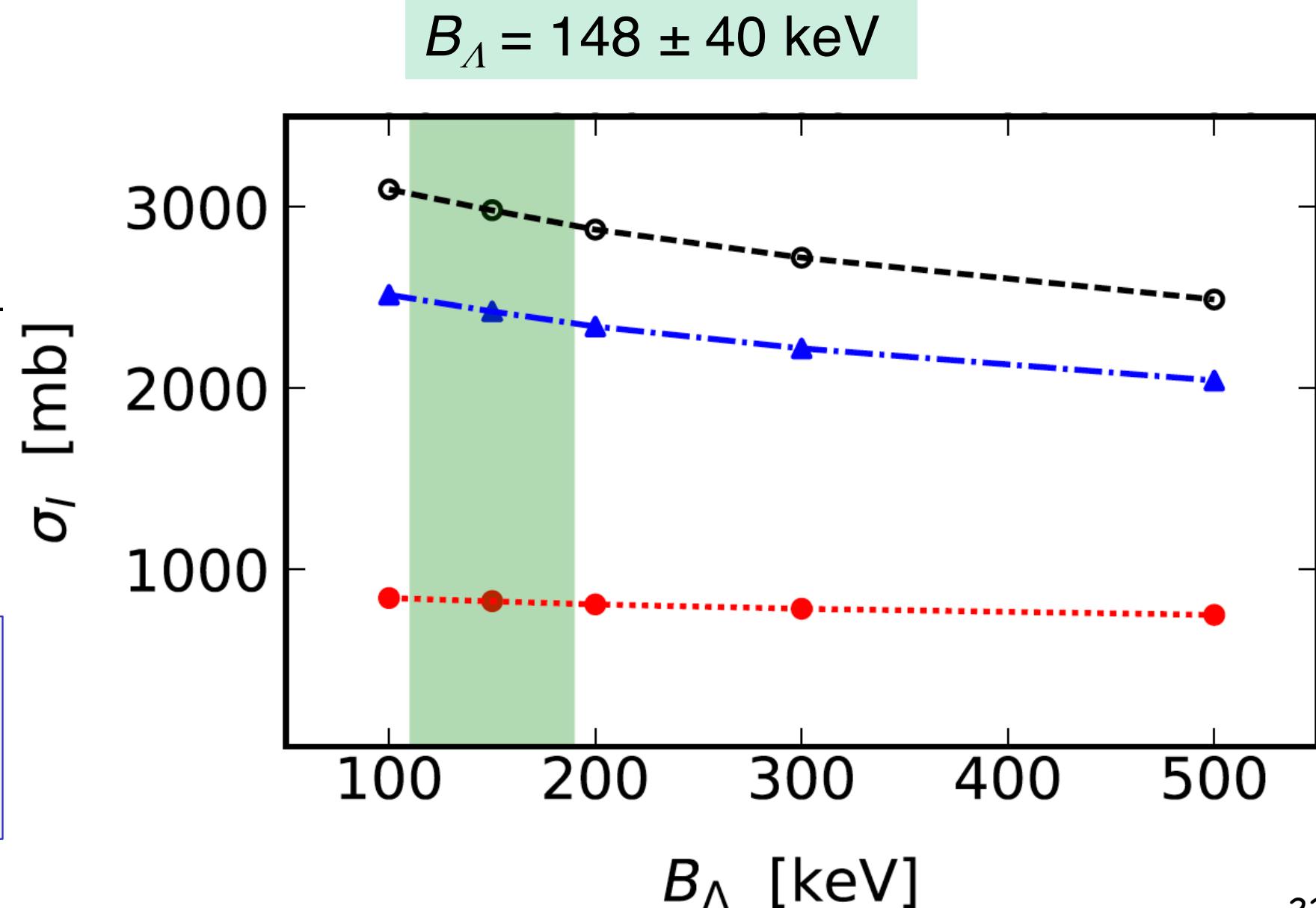
Transition from full opaque-ness to full transparency displays changes in the slope due to structure of the deuteron within ${}^3\Lambda\text{H}$.

Sensitivity to hypertriton interaction cross sections

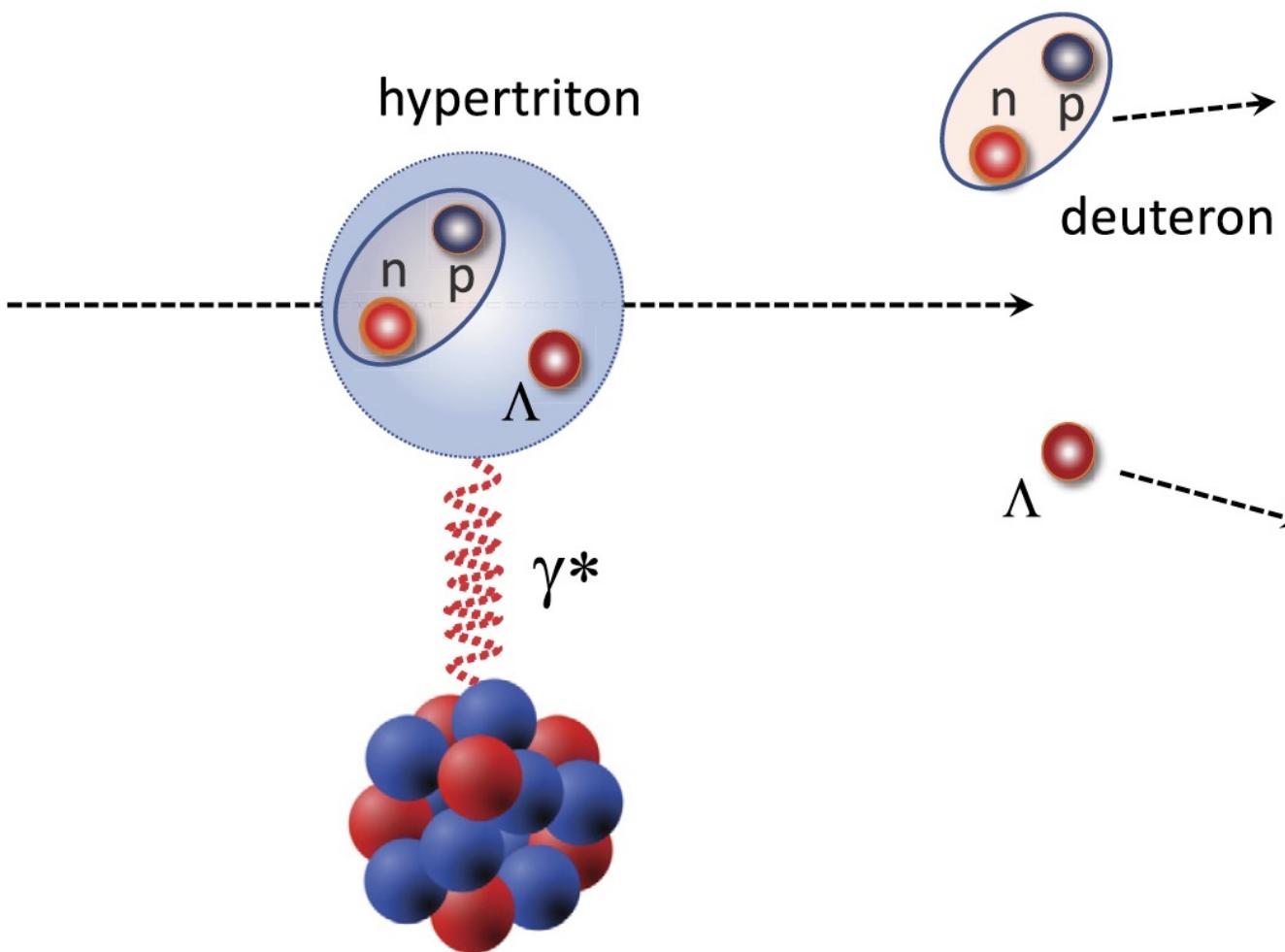
1.5 GeV/nucleon $^3\Lambda$ H incident on ^{12}C , ^{120}Sn , ^{208}Pb

B_Λ (keV)	$\sigma_I(\text{C})$	$\sigma_I(\text{Sn})$	$\sigma_I(\text{Pb})$
100	842.	2516.	3098.
150	824.	2424.	2982.
200	807.	2341.	2876.
300	783.	2220.	2721.
500	749.	2043.	2490.

For a C target: a 12% reduction of the cross section $\sigma_{\Lambda I}$ from $B = 100$ MeV to $B = 500$ MeV.



Electromagnetic response of the hypertriton



$$\frac{d\sigma_C}{dE} = \frac{16\pi^3}{9\hbar c} n(E) \frac{dB(E)}{dE}$$

$$n(E) = \frac{2Z_T^2\alpha}{\pi} \left(\frac{Ec}{\gamma\hbar\nu^2} \right)^2 \int_0^\infty dbb \left[K_1^2 + \frac{1}{\gamma^2} K_0^2 \right] T(b)$$

$$x = Eb/\gamma\nu$$

First-order perturbation theory

$$\frac{dB(E)}{dE} = \frac{1}{\hbar} \sqrt{\frac{\mu}{2E}} |\langle g.s. || \mathcal{O}_{E1} || E, l \rangle|^2$$

$$\langle g.s. || \mathcal{O}_{E1} || E, l \rangle = (-1)^l \frac{e_{eff}}{\sqrt{4\pi}} \int_0^\infty dr r u_{g.s.}(r) u_{E,l}(r)$$

$$u_{E,l}(r) \rightarrow \sqrt{2\mu_{\Lambda d}/\pi\hbar^2 k} e^{i\delta_l} \sin(kr + \delta_l)$$

Electromagnetic response of the hypertriton

$$\frac{dB(E)}{dE} = \frac{1}{\hbar} \sqrt{\frac{\mu}{2E}} |\langle g.s. || \mathcal{O}_{E1} || E, l \rangle|^2$$

Analytical model:

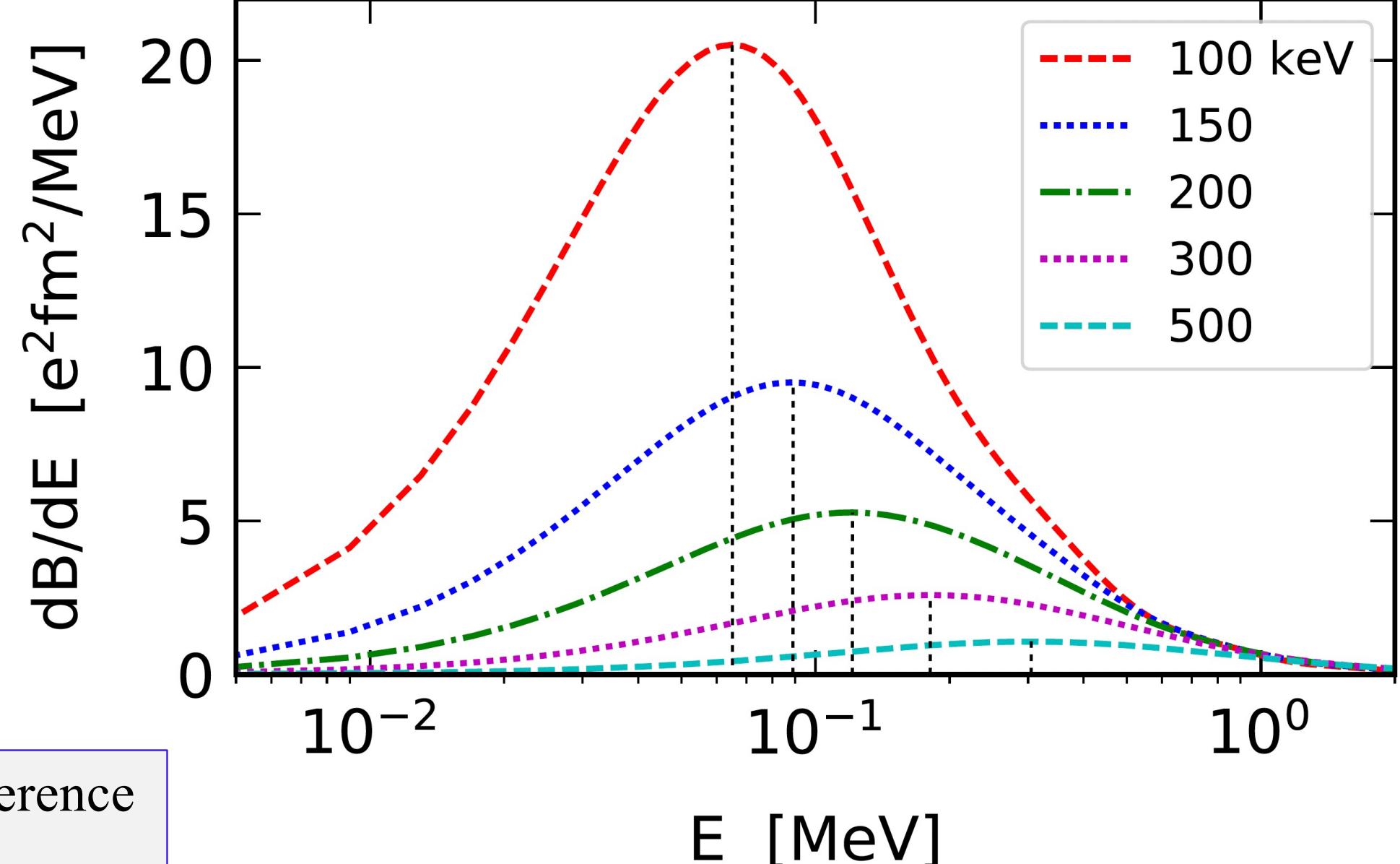
CB, Sustich

Phys. Rev. C 46 (1992) 2340

$$\frac{dB(E)}{dE} = C \sqrt{B_\Lambda} \frac{E^{3/2}}{(E + B_\Lambda)^4}$$

$$E_{max} = \frac{3}{5} B_\Lambda$$

- Excellent agreement < 4% difference
- FSI small

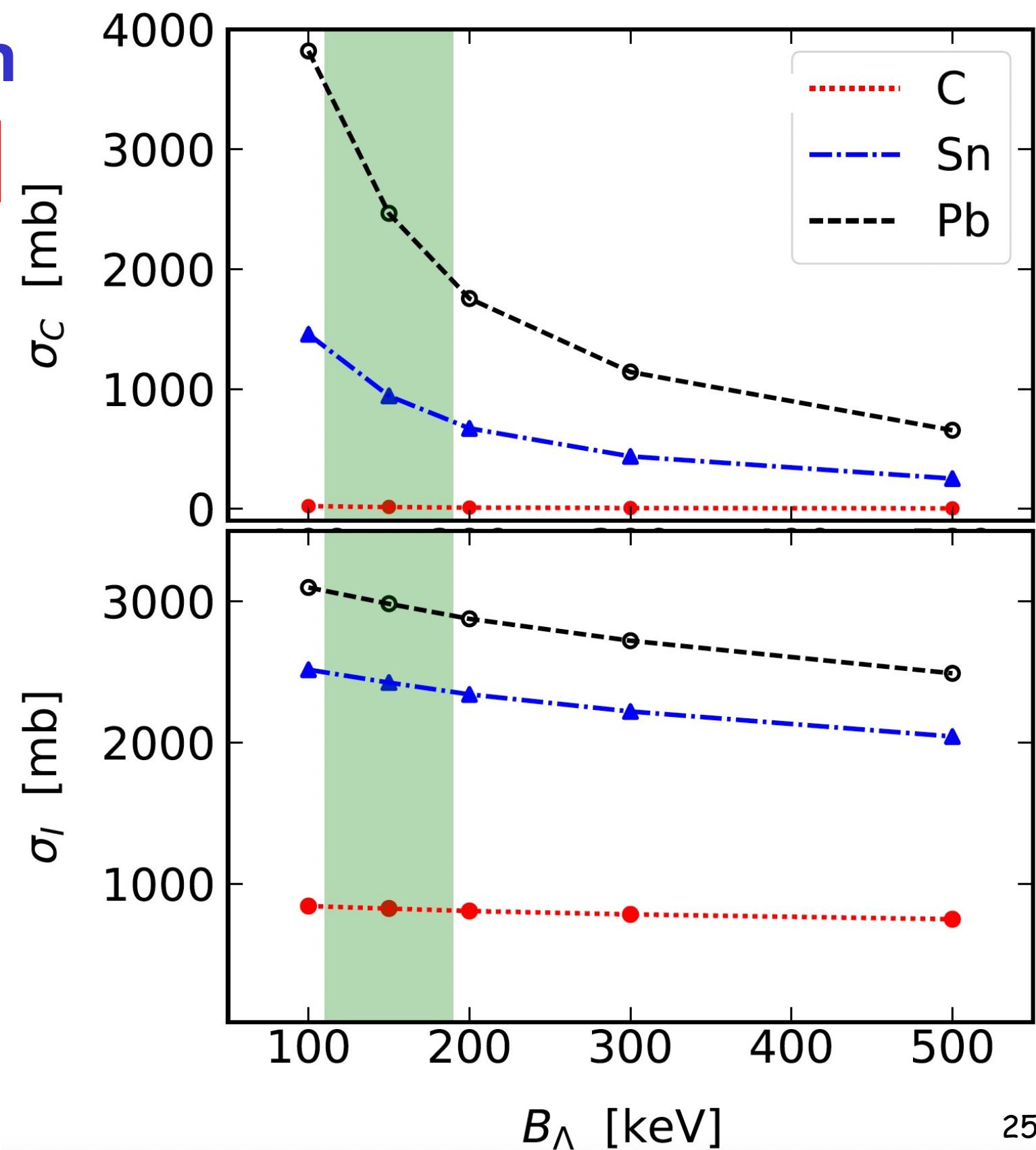


EM response of the hypertriton

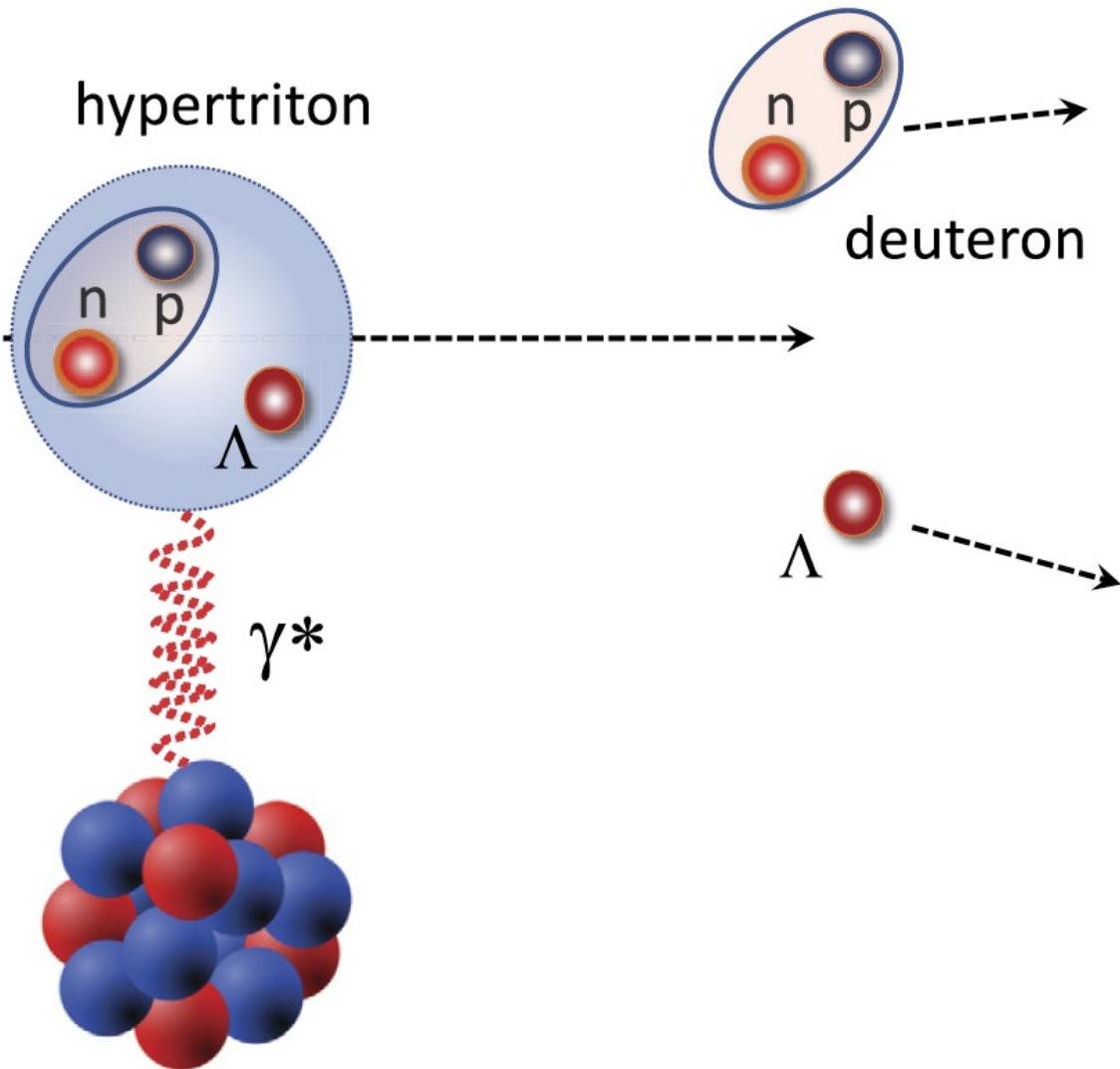
1.5 GeV/nuc. ${}^3\Lambda$ H incident on ${}^{12}\text{C}$, ${}^{120}\text{Sn}$, ${}^{208}\text{Pb}$

B_Λ (keV)	$\sigma_C(\text{C})$	$\sigma_C(\text{Sn})$	$\sigma_C(\text{Pb})$
100	22.9	1457.	3820.
150	14.9	942.	2464.
200	10.7	672.	1755.
300	7.1	438.	1142.
500	4.1	253.	656.

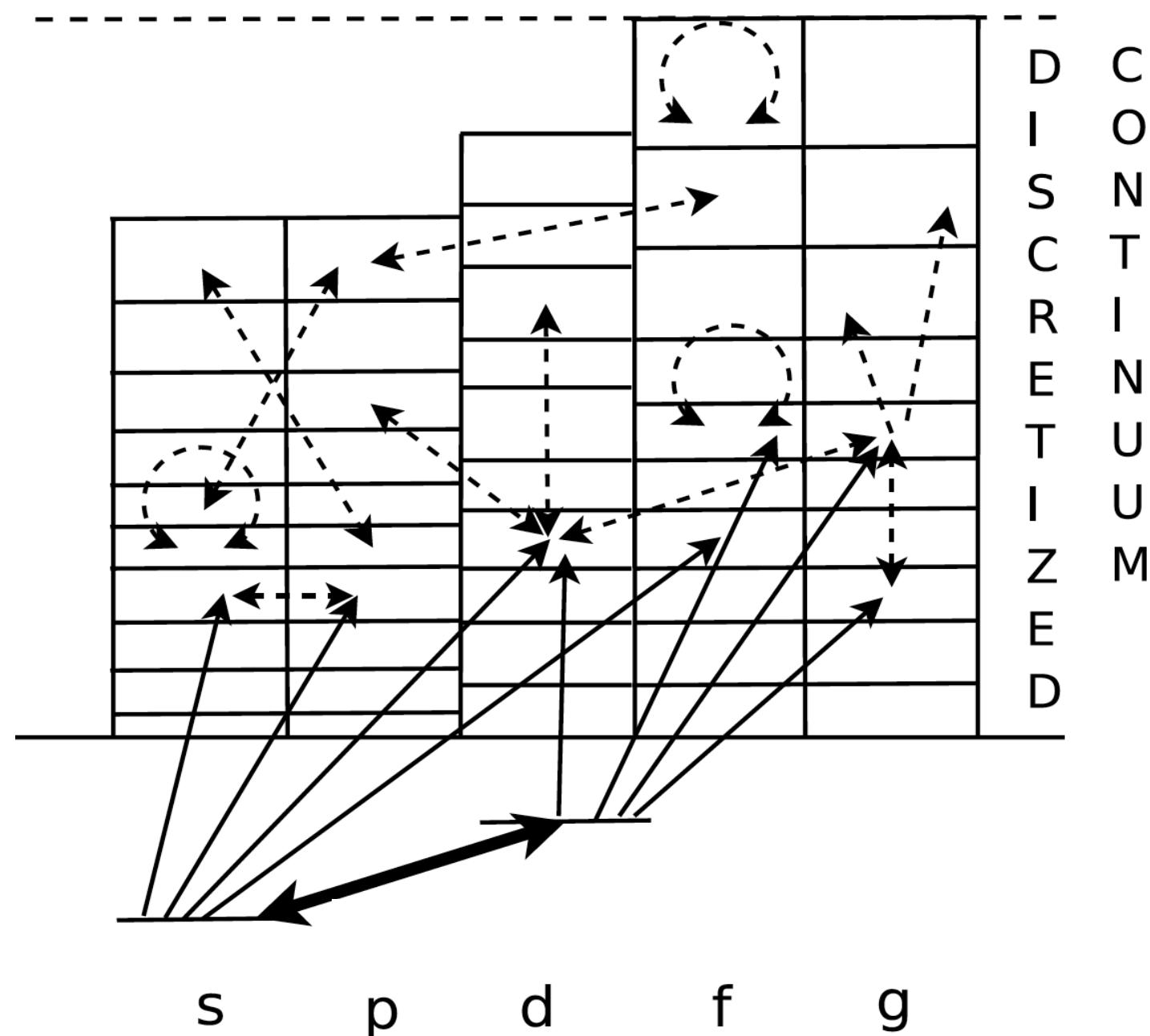
B_Λ (keV)	$\sigma_I(\text{C})$	$\sigma_I(\text{Sn})$	$\sigma_I(\text{Pb})$
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200	807.	2341.	2876.
300	783.	2220.	2721.
500	749.	2043.	2490.



EM response of the hypertriton



- Basis expansion for intrinsic ϕ
- Eikonal scattering waves
- Nuclear + EM potentials
- Relativity



- Continuum discretization
- Coupled-channels
(relativistic CDCC)

Electromagnetic response of loosely bound nuclei

$$i\hbar\nu \frac{d}{dz} S_c(\mathbf{b}, z) = \sum_{c'} \langle \Phi_c | H_{int}(\mathbf{b}, z) | \Phi_{c'} \rangle S_{c'}(\mathbf{b}, z) \exp \left[i \frac{E_{cc'} z}{\hbar c} \right]$$

$$f_c(\mathbf{q}) = -\frac{ik}{2\pi} \int db \exp(i\mathbf{q} \cdot \mathbf{b}) [S_c(\mathbf{b}, z) - \delta_{0c}]$$

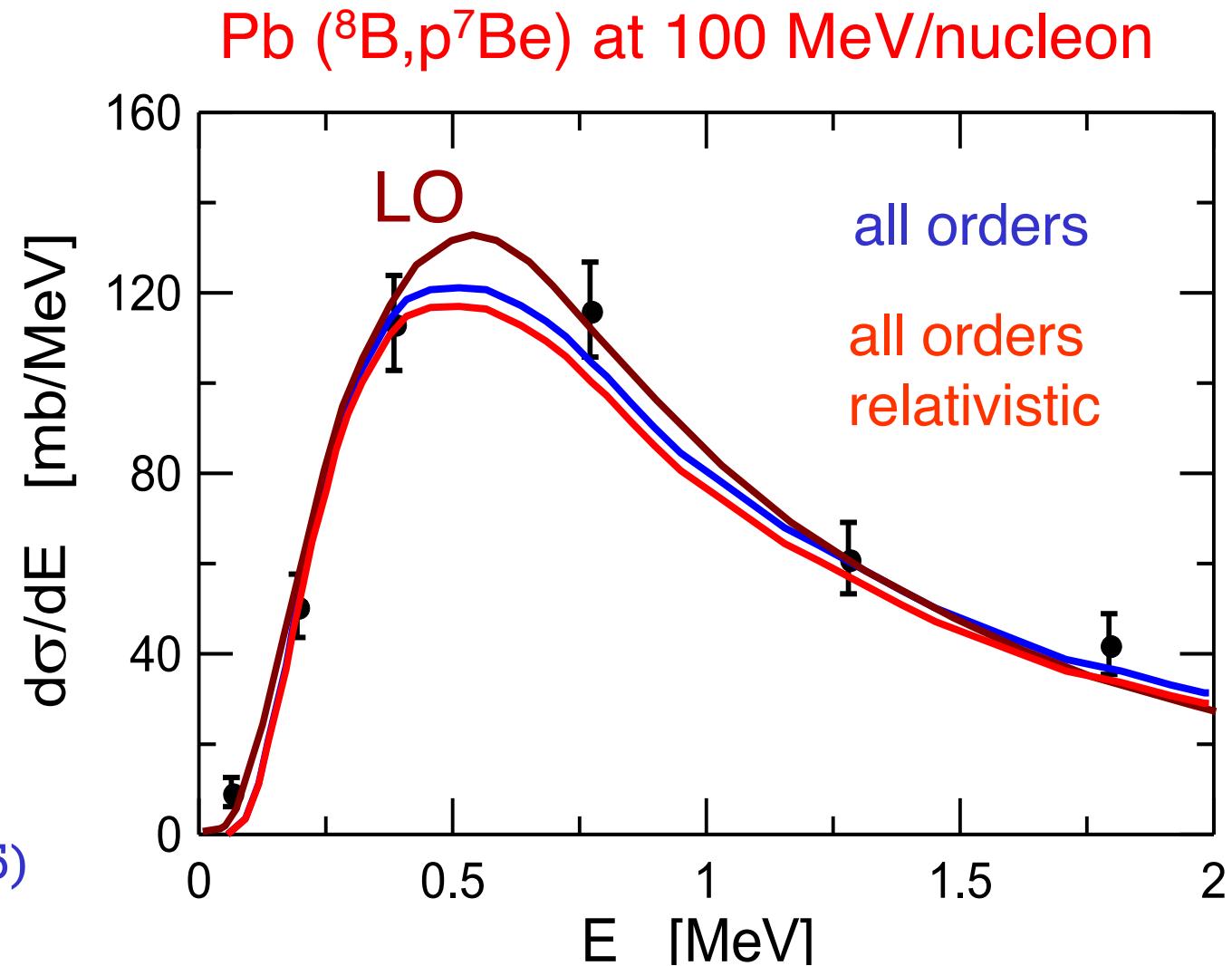
$$\frac{d\sigma_c}{d\Omega} = \sum_{M_0, M_c} \left| f_c^{(M_c - M_0)}(\theta, E_c) \right|^2$$

$$|E_c\rangle = \int dE'_c \Gamma(E'_c) |E'_c\rangle$$

$$\Gamma(E_j) = \begin{cases} \frac{1}{\Delta E} & \text{if } (j-1)\Delta E < E_c < j\Delta E \\ 0, & \text{otherwise} \end{cases}$$

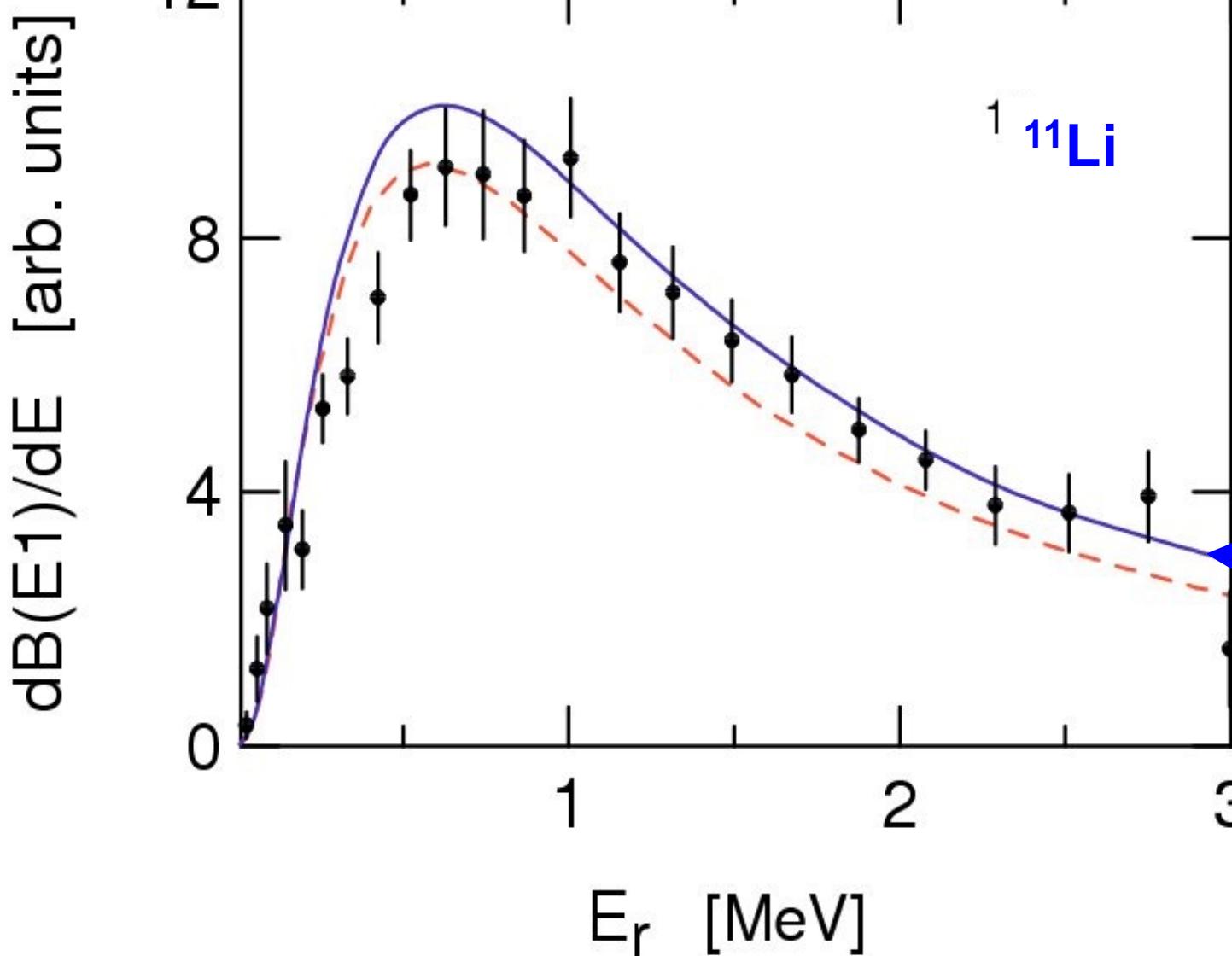
CB, PRL 94, 072701 (2005)

$H_{int} = H_{EM} = \text{easy}$
 $H_{int} = H_{\text{nucleus-nucleus}} = \text{very complicated}$

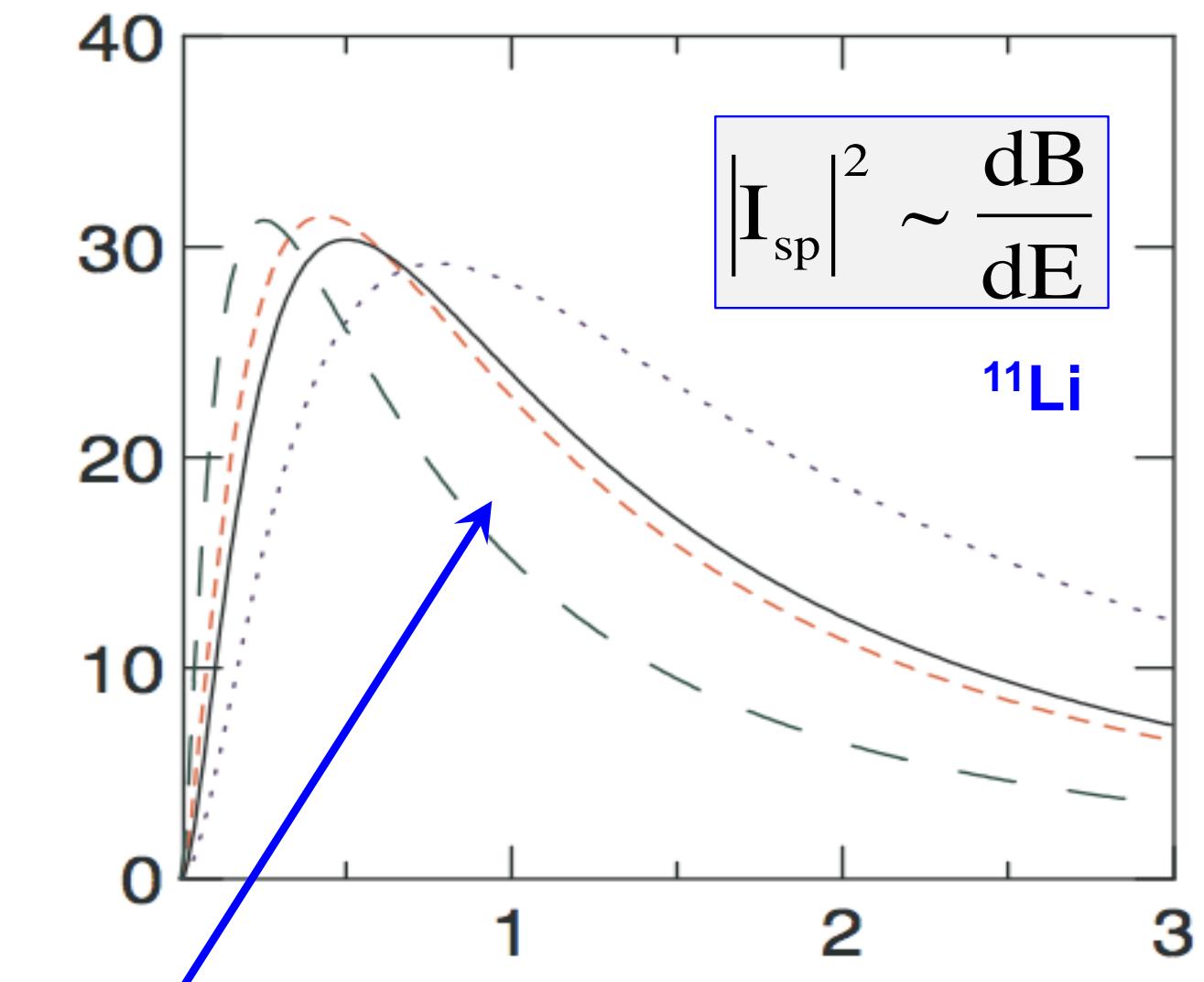


FSI in 3-body models

3-body configurations and FSI widen
 $\frac{dB(E1)}{dE}$



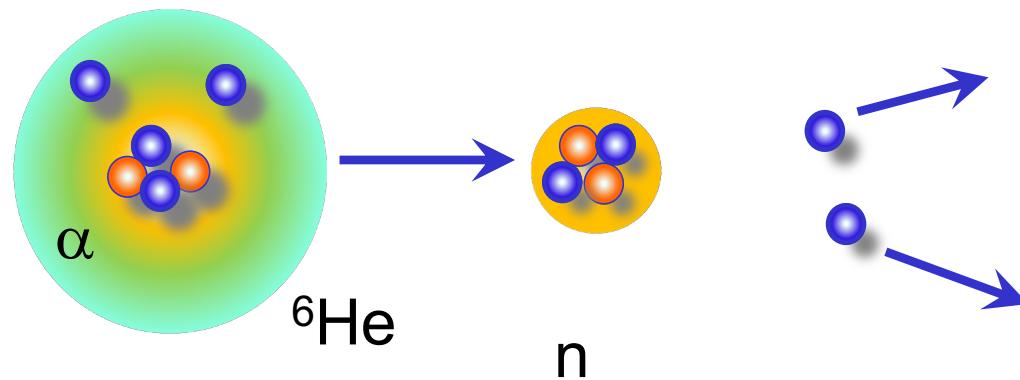
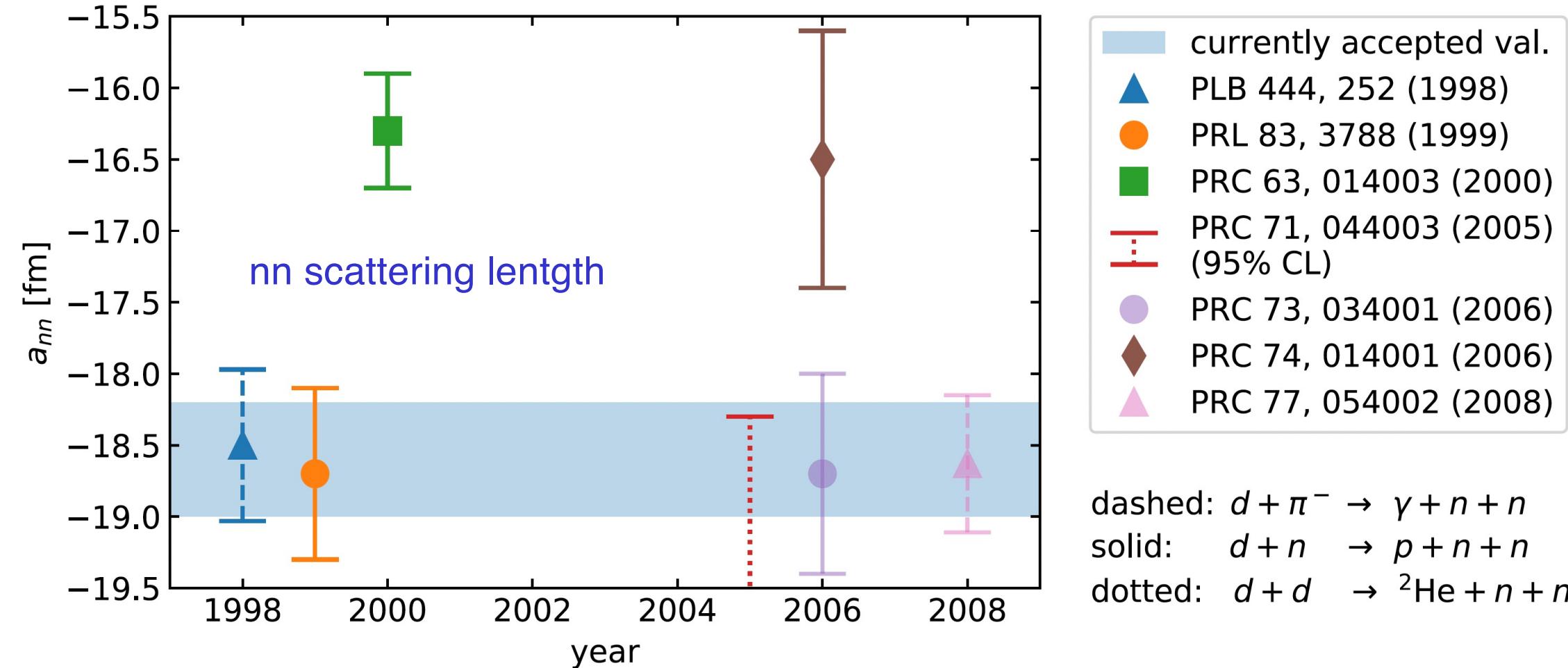
$|\mathcal{I}_{sp}|^2 \text{ [MeV fm}^5]$



FSI: Different
scattering lengths,
effective ranges

CB, PRC 75, 024606 (2007)

nn scattering length



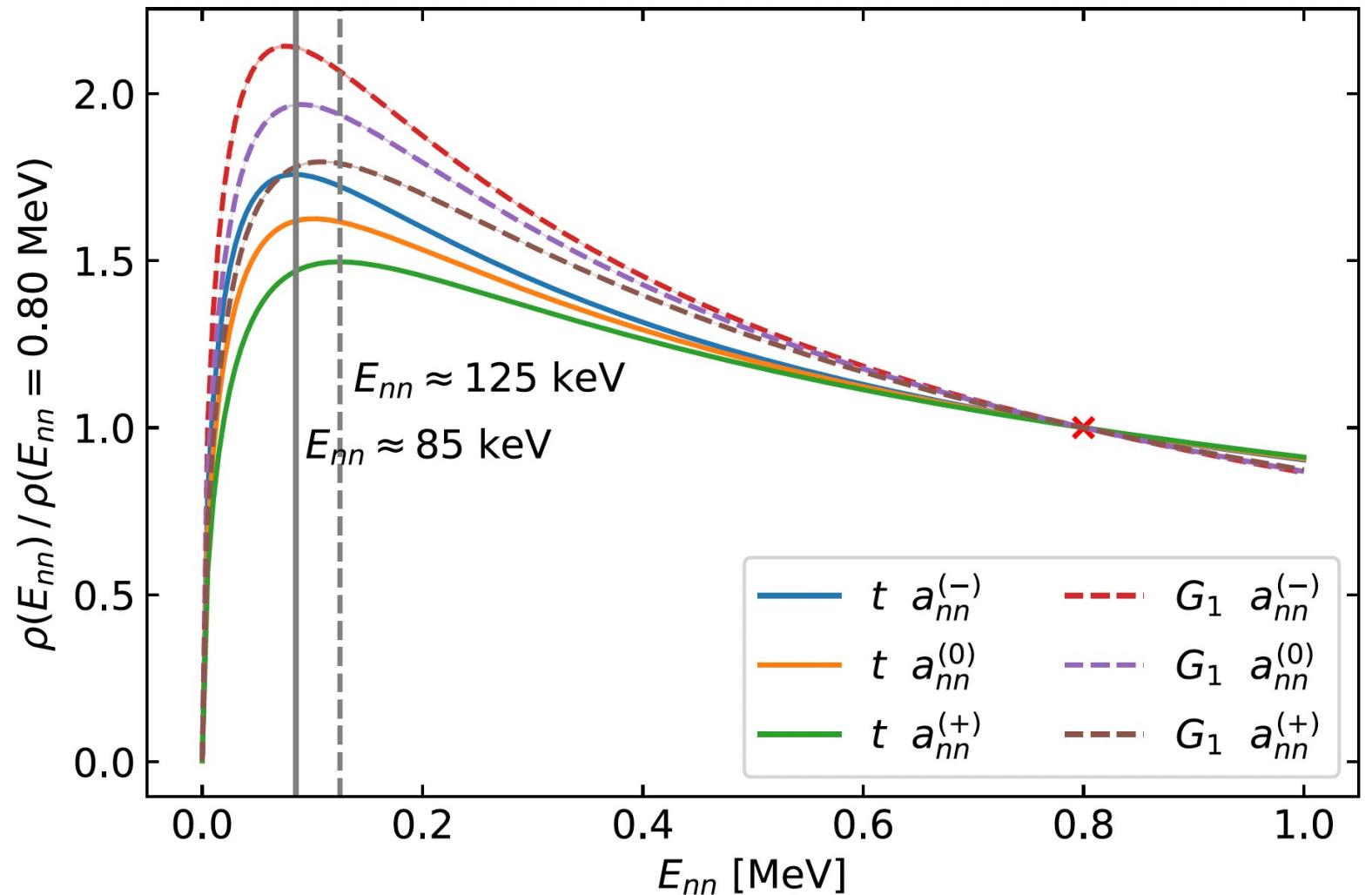
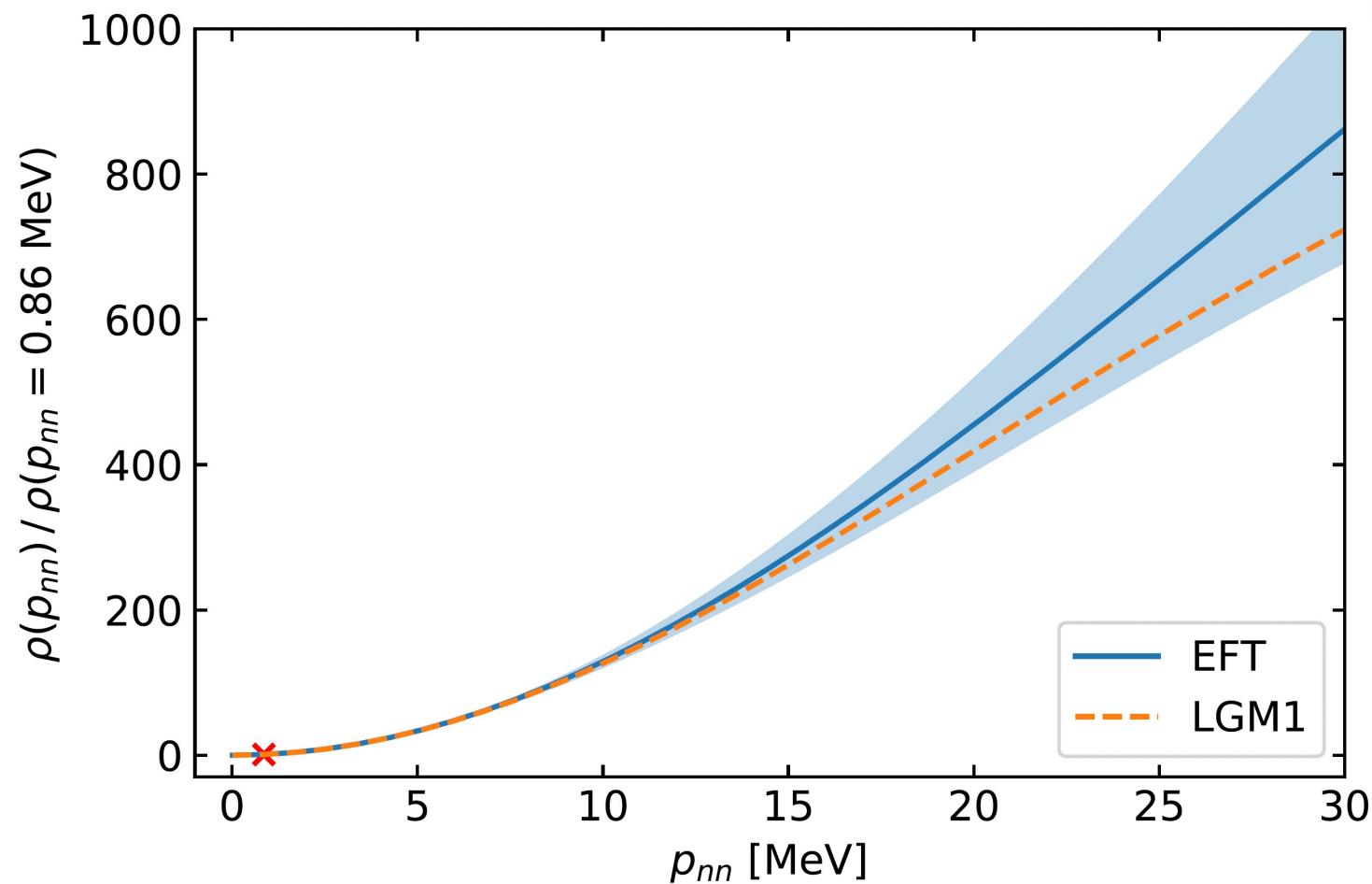
${}^6\text{He}(p,p\alpha)nn$ reaction in inverse kinematics at high energies

Göbel, et al, PRC 104, 024001 (2021)

nn scattering length

Göbel, et al, PRC 104, 024001 (2021)

Halo EFT \rightarrow nn configuration
in ${}^6\text{He} \rightarrow E_{nn}$ density



LGM1 = Faddeev calculations with local Gaussian potentials

pp scattering length

v_{NN} charge independence violation: $m_{\pi^\pm} \neq m_{\pi^0}$

v_{NN} charge symmetry violation: $m_{down} \neq m_{up}$

$p + d \rightarrow p + p + n \rightarrow p + p$ scattering

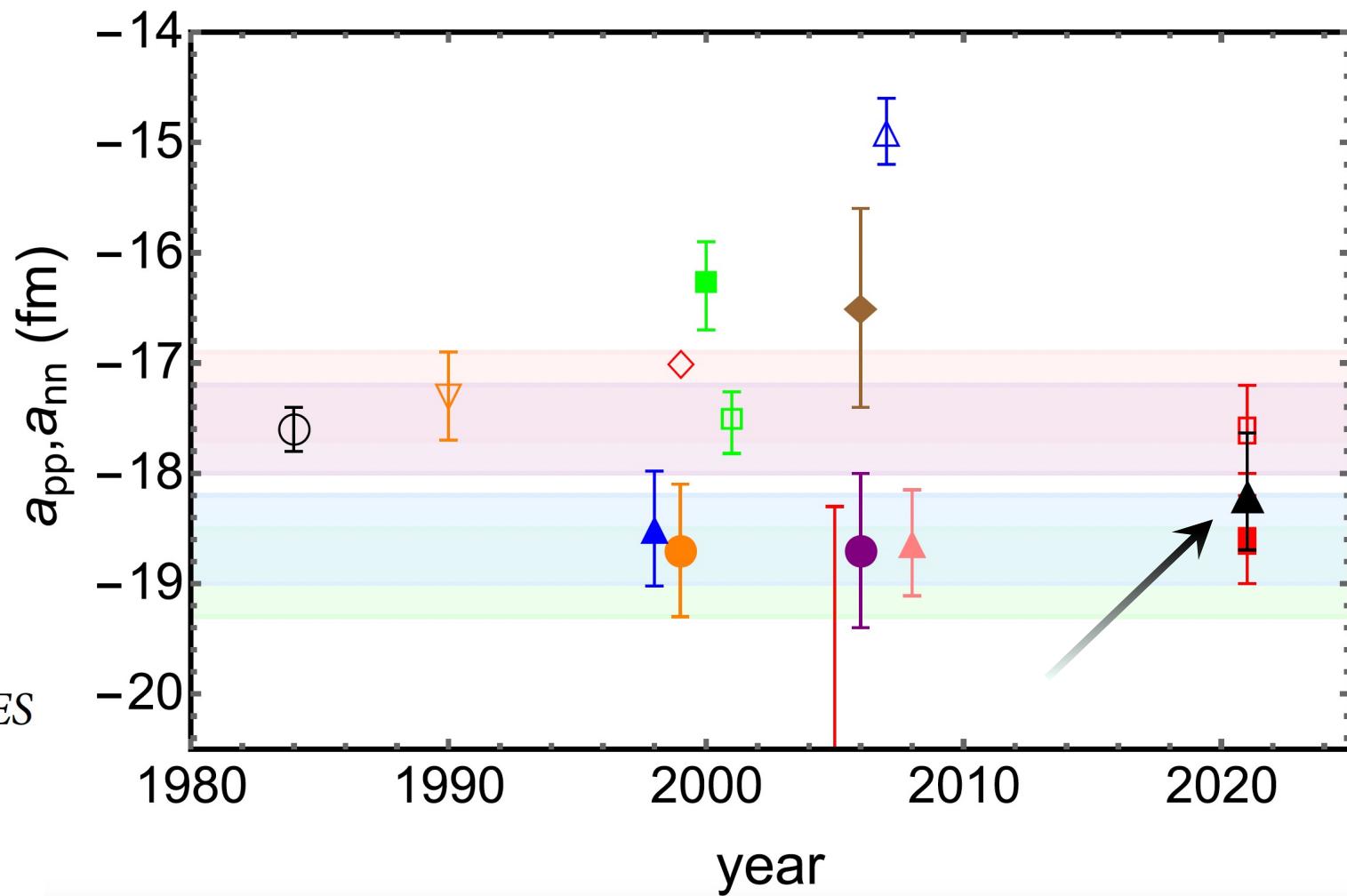
Tumino, et al,
Nature Comm. Phys. 6, 106 (2023)

$$\frac{d^3\sigma}{d\Omega_B d\Omega_b dE_B} = (KF) \cdot |\phi(p_{xs})|^2 \cdot \left[\frac{d^2\sigma_{xA \rightarrow bB}}{dE_{c.m.} d\Omega} \right]^{HOES}$$

$$\left(\frac{d\sigma}{d\Omega_{c.m.}} \right)^{HOES} = \frac{1}{4k^2} \left(|F(\mathbf{p}, \mathbf{k}) - 2T_{CN}(p, k)|^2 + 3|F(\mathbf{p}, \mathbf{k})|^2 \right)$$

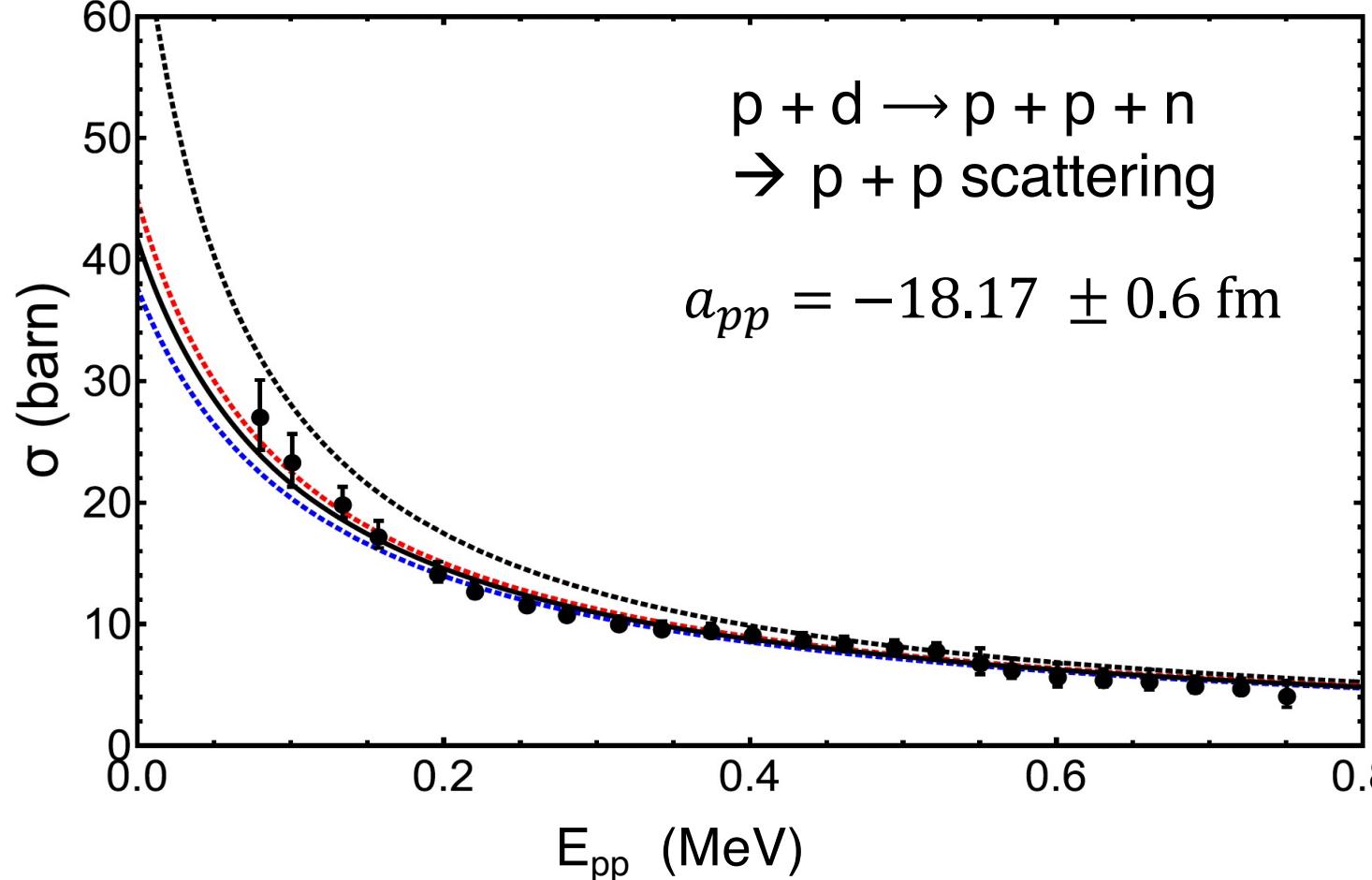
$$F(\mathbf{p}, \mathbf{k}) = m_p e^2 e^{-\pi\eta} \Gamma(1 + i\eta) (p^2 - k^2)^{i\eta} g(\mathbf{p}, \mathbf{k})$$

$$g(\mathbf{p}, \mathbf{k}) = (\mathbf{p} - \mathbf{k})^{-2(1+i\eta)} \pm (\mathbf{p} + \mathbf{k})^{-2(1+i\eta)}$$



Energy differences
between mirror nuclei
small as compared to
experiment (Nolen-
Schiffer anomaly)

pp scattering length



NN s-wave phase shift

$$k \cot \delta = -\frac{1}{a} + \frac{1}{2} r_0 k^2$$

$$\frac{4\pi}{4\pi}$$

$$\sigma = \frac{\left(\frac{1}{a} - \frac{1}{2} r_0 k^2\right)^2 + k^2}{\left(\frac{1}{a} - \frac{1}{2} r_0 k^2\right)^2 + k^2}$$

$$V_{NN}(R) = V_0 e^{-\frac{r^2}{r_G^2}} + \frac{e_{NN}^2}{r}$$

disable EM \rightarrow d from short-range universal window

Table 2 The effect of electromagnetic terms. Low-energy parameters with and without the inclusion of the electromagnetic terms for the two potential models indicated.

	Argonne v_{18}	w/o v^{EM}		Argonne v_{18}	w/o v^{EM}
$^1a_{nn}(\text{fm})$	-18.487	-18.818	$^1a_{pp}(\text{fm})$	-7.806	-17.164
$^1r_{nn}(\text{fm})$	2.840	2.834	$^1r_{pp}(\text{fm})$	2.788	2.865
	Gaussian	w/o v^{EM}		Gaussian	w/o v^{EM}
$^1a_{nn}(\text{fm})$	-18.487	-18.89 ± 0.02	$^1a_{pp}(\text{fm})$	-7.806	-17.19 ± 0.03
$^1r_{nn}(\text{fm})$	2.85 ± 0.07	2.83 ± 0.07	$^1r_{pp}(\text{fm})$	2.77 ± 0.08	2.86 ± 0.08

FSI in light hypernuclei decay

$$m_\Lambda = 1115.7 \text{ MeV}$$

$$\Gamma = 2.5 \times 10^{-6} \text{ eV}$$

$$\tau = 263 \pm 2 \text{ ps}$$

$$\Lambda \rightarrow p\pi^- (64.1\%)$$

$$\Lambda \rightarrow n\pi^0 (35.7\%)$$

Nonmesonic decays of light hypernuclei is the only tractable way to study $\Delta S = 1$ weak baryon-baryon interaction.

Parker, PRC 76, 035501 (2007)

$$\begin{aligned} \Gamma_{\text{total}} &= \frac{\Gamma_{\text{mesonic}}}{\Gamma_{\pi^-} + \Gamma_{\pi^0} + \Gamma_{\pi^+}} + \frac{\Gamma_{\text{nonmesonic}}}{\Gamma_p + \Gamma_n + \Gamma_{mb}} \end{aligned}$$

Two-body nonmesonic decay modes, $\Lambda N \rightarrow NN$, distinguishable from $\Lambda \rightarrow \pi N$ because of large decay energy $M_\Lambda - M_N = 176 \text{ MeV}$. Sensitive to weak interaction couplings (such as $g_{\Lambda N \rho}$ or g_{NNK}) not available to the free Λ decays.

At the quark level:

$$H_{\text{weak}} = \frac{G_F}{2} \sin\theta_c \cos\theta_c [\bar{u}\gamma_\mu(1-\gamma_5)s\bar{d}\gamma^\mu(1-\gamma_5)u]$$

Effective Lagrangian for $\Delta T = 1/2$:

$$H_{\text{weak}} = -G_F m_\pi^2 \bar{\psi}_N (A_\pi + B_\pi \gamma_5) \phi_\pi \cdot \tau \psi_\Lambda \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

Isgur et al, PRL 64, 161 (1990)

$$A_\pi = 1.05, \quad B_\pi = -7.15$$

$$\Gamma_{\pi^-}: \Lambda \rightarrow p\pi^-$$

$$\Gamma_{\pi^0}: \Lambda \rightarrow n\pi^0$$

$$\Gamma_{\pi^+}: p\Lambda \rightarrow nn\pi^+$$

$$\Gamma_p: \Lambda p \rightarrow np$$

$$\Gamma_n: \Lambda n \rightarrow nn$$

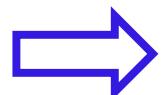
$$\Gamma_{mb}: \Lambda NN \rightarrow NNN$$

FSI in light hypernuclei decay

CB, Lobato, EPJ A 57, 67 (2021)

Fadeev + approximations

$$\psi_f^{(-)} = \left[1 + (E^{(-)} - H_0)^{-1} t_{pn}^\dagger \right] \chi^{(-)}(\mathbf{p}_p) \chi^{(-)}(\mathbf{p}_n) \psi_{A-2}$$



$$\begin{aligned} \Gamma = & \frac{2\pi}{\hbar} \sum \int \left| \langle \chi^{(-)}(\mathbf{p}_p) \chi^{(-)}(\mathbf{p}_n), \Psi_{A-2}(S M_S J_F M_F) \rangle \right| \\ & \times \left[1 + t_{pn} \left(\Delta_n - \frac{\hbar^2 \mathbf{P}_{pn,cm}^2}{4m_N} \right) (\Delta_n - H_0 + i\varepsilon)^{-1} \right] \\ & \times V_{weak} \left| \Psi_A(J_I M_I) \right|^2 \delta(E_{pn,cm} + E_{pn,rel} - \Delta_n) \\ & \times \frac{d\mathbf{p}_p}{(2\pi)^3} \frac{d\mathbf{p}_n}{(2\pi)^3}. \end{aligned}$$

and similarly for nn

H_0 = kinetic energies of p and n

Δ_n = total energy of three-body system

FSI in light hypernuclei decay

Watson, PR 88, 1163 (1952)

Migdal, JETP 1, 2 (1954)

$$F(E_{nN}) = \left| \frac{\psi(k_{nN}, r_{nN})}{\psi^{(0)}(k_{nN}, r_{nN})} \right|^2$$

$$= \frac{(1/r_{nN} - 1/a_{nN} + k_{nN}^2 r_{nN}/2)^2}{(-1/a_{nN} + k_{nN}^2 r_{nN}/2)^2 + k_{nN}^2}$$

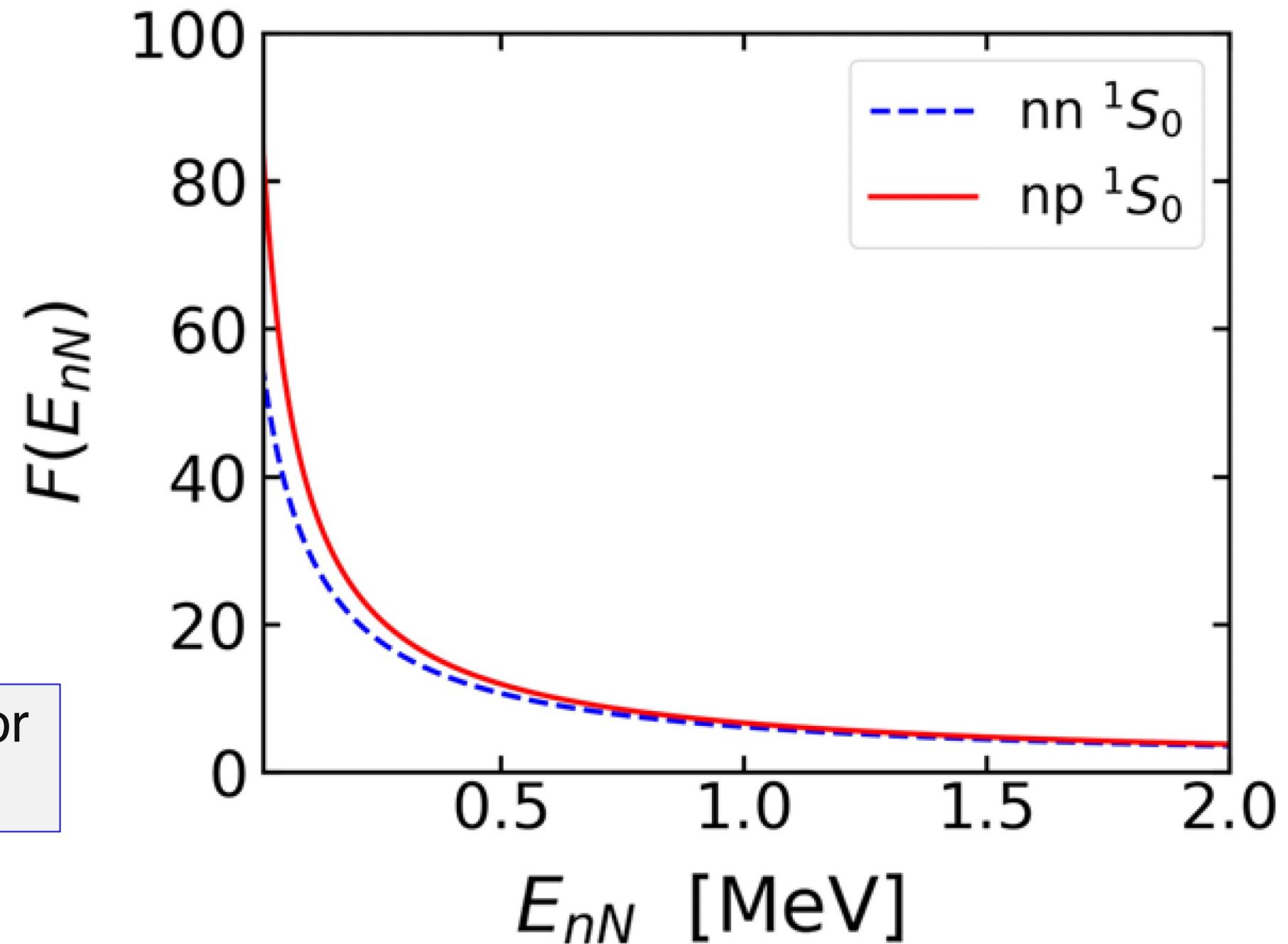
$$\Gamma = \frac{2\pi}{\hbar} \sum \int \left| \left\langle \chi^{(-)}(\mathbf{p}_p) \chi^{(-)}(\mathbf{p}_n), \Psi_{A-2}(SM_S J_F M_F) \right| \right.$$

$$\times \left[1 + t_{pn} \left(\Delta_n - \frac{\hbar^2 \mathbf{P}_{pn,cm}^2}{4m_N} \right) (\Delta_n - H_0 + i\varepsilon)^{-1} \right]$$

$$\times V_{weak} \left| \Psi_A(J_I M_I) \right\rangle \left| \delta(E_{pn,cm} + E_{pn,rel} - \Delta_n) \right|^2$$

$$\times \frac{d\mathbf{p}_p}{(2\pi)^3} \frac{d\mathbf{p}_n}{(2\pi)^3} .$$

FSI in light hypernuclei decay



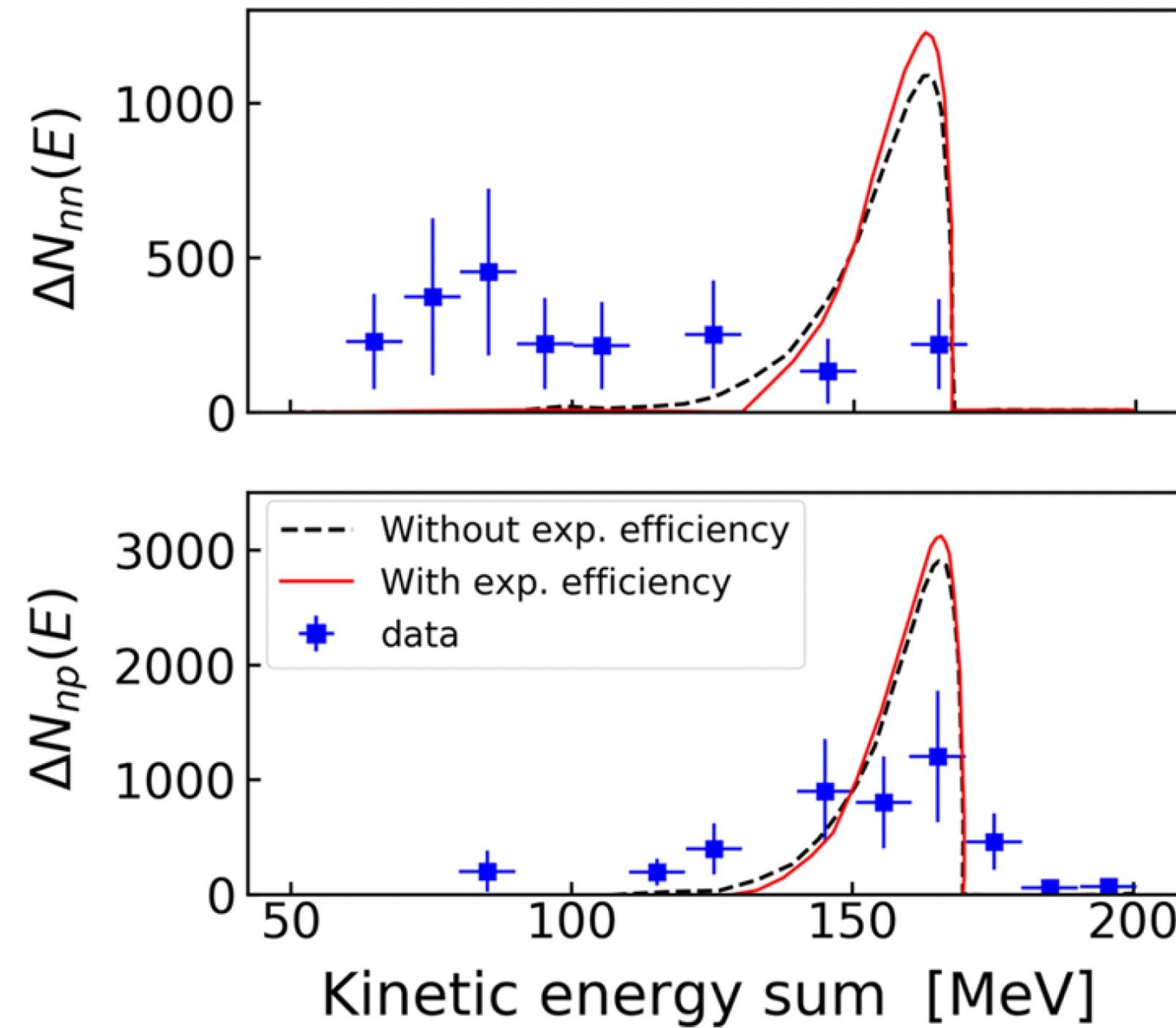
Migdal-Watson enhancement for
 1S_0 singlet nn and np systems

FSI in light hypernuclei decay

Low energy peak reasonable for the pn channel, while it is quite bad for the nn channel

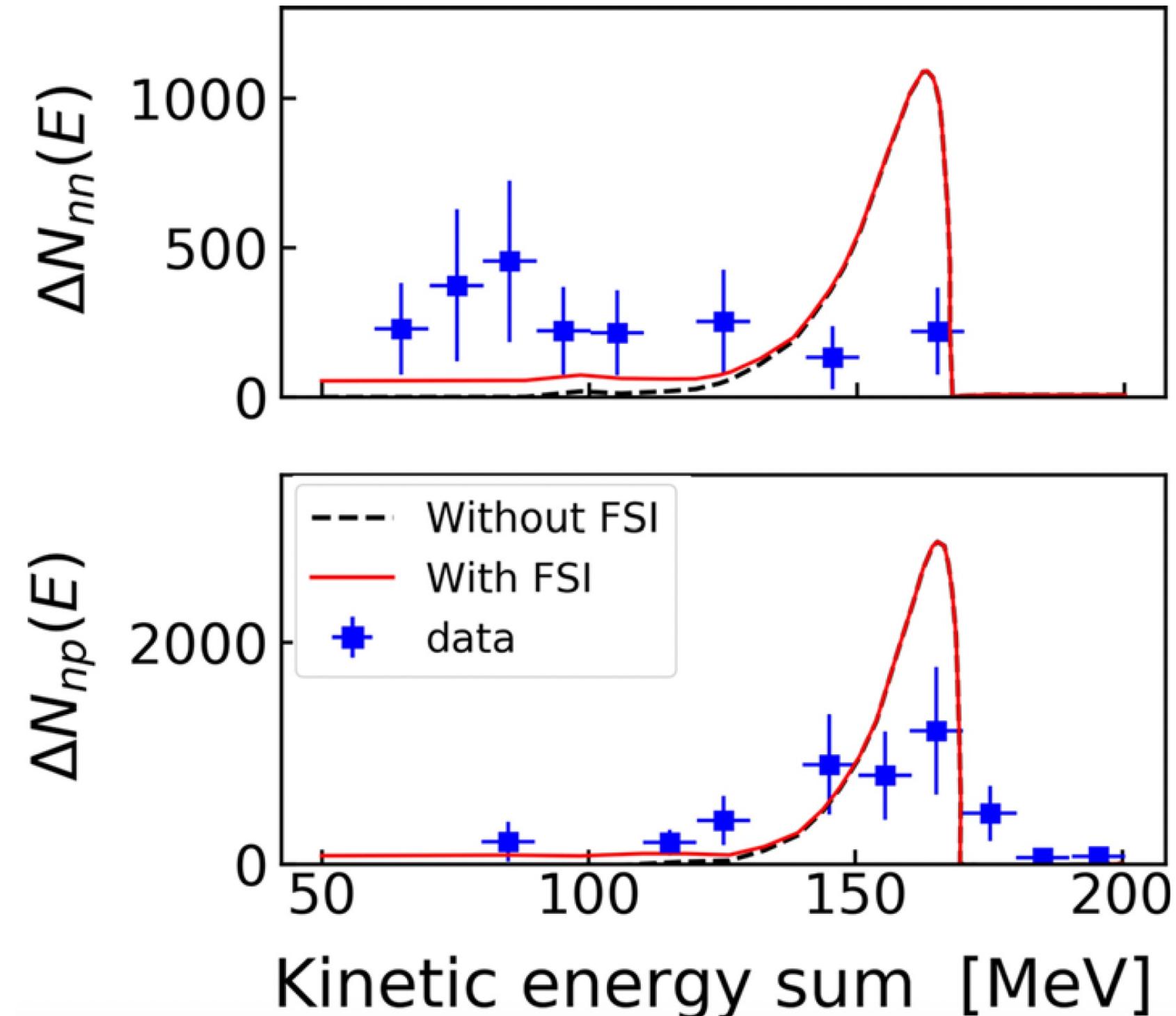
Data:

Parker, PRC 76, 035501 (2007)



FSI in NMWD

Effect of FSI the 1S_0 singlet neutron-neutron (nn) and neutron-proton (np) Λ A non-mesonic decay



Spin correlations

$$\mathcal{A}(\mathbf{p}_1, \mathbf{r}_1) = \int d^3 r \chi^{(+)}(\mathbf{p}_1, \mathbf{r}) K(\mathbf{r}, \mathbf{r}_1) \psi_0(\mathbf{r}_1).$$

$K(\mathbf{r}, \mathbf{r}_1)\psi_0(\mathbf{r}_1)$ is the propagator for wavefunction evolution from the source to detector

Λ^\pm amplitude for singlet (triplet)

$$\Lambda^{(\pm)}(\mathbf{p}_1, \mathbf{p}_2, \mathbf{r}_1, \mathbf{r}_2) = \frac{1}{\sqrt{2}} \left[\mathcal{A}(\mathbf{p}_1, \mathbf{r}_1) \mathcal{A}(\mathbf{p}_2, \mathbf{r}_2) \right. \\ \left. \pm \mathcal{A}(\mathbf{p}_2, \mathbf{r}_1) \mathcal{A}(\mathbf{p}_1, \mathbf{r}_2) \right],$$

$$P(\mathbf{p}_1, \mathbf{p}_2) = \int d^3 r_1 d^3 r_2 \left| \left| \Lambda^{(+)}(\mathbf{p}_1, \mathbf{p}_2, \mathbf{r}_1, \mathbf{r}_2) \right|^2 \right.$$

$$\left. \pm \mathcal{M} \Lambda^{(-)}(\mathbf{p}_1, \mathbf{p}_2, \mathbf{r}_1, \mathbf{r}_2) \right|^2 \Bigg],$$

Correlation function: relative contribution of the singlet and the triplet states in the initial configuration

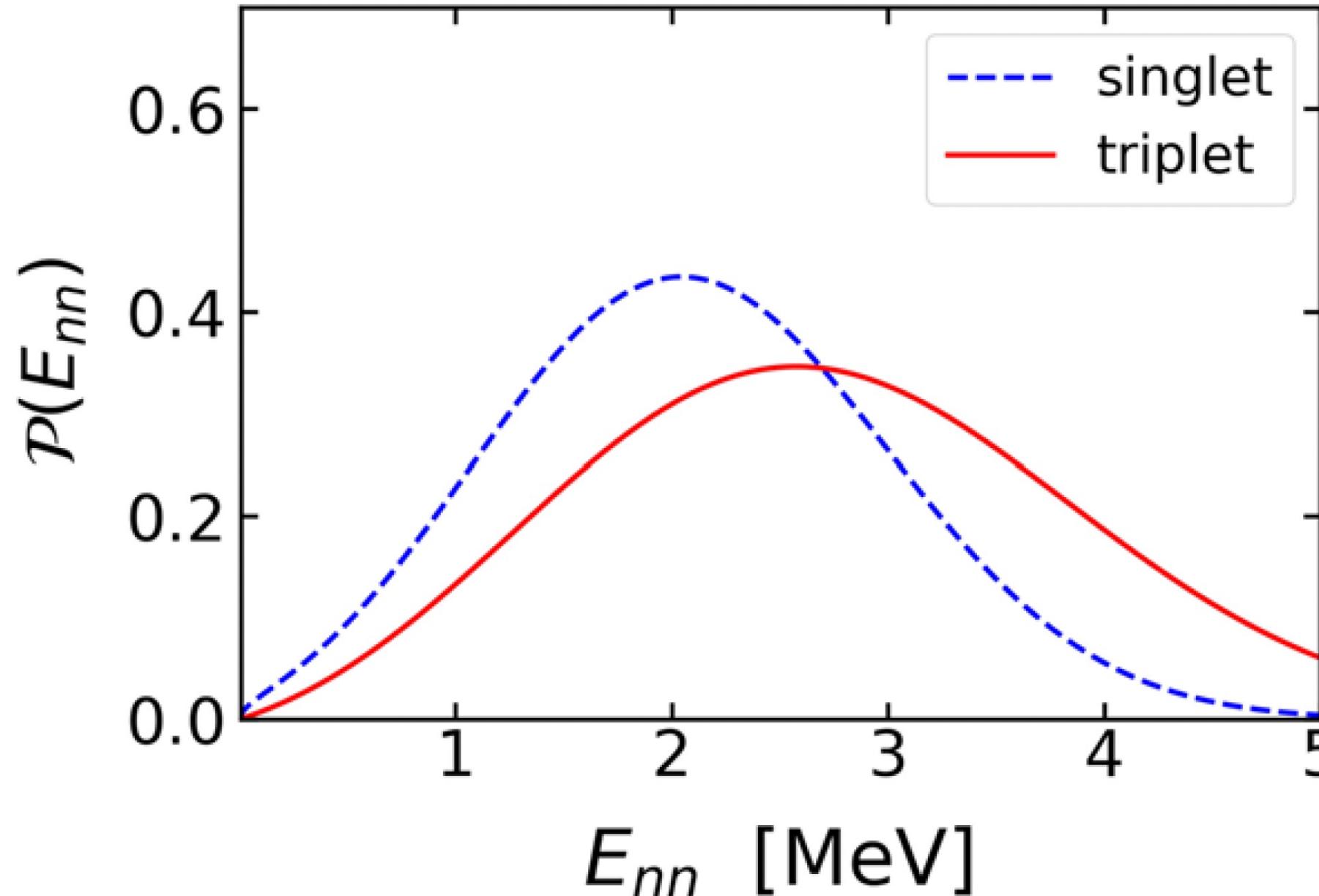
Probability admixture for singlet and triplet states

$$C(\mathbf{p}_1, \mathbf{p}_2) = \frac{P(\mathbf{p}_1, \mathbf{p}_2)}{P(\mathbf{p}_1) P(\mathbf{p}_2)}$$

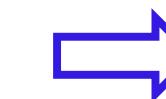
Spin correlations

Two-body probability density $\mathcal{P}(E_{nn})$

$$P(E_{nn}) = \mathcal{P}(E_{nn}) dE_{nn}$$



Assuming $r_0 = 4$ fm for the average initial distance of the nn pair right after the NMWD



Spin correlations more important than FSI in NMWD

Summary

- Production and fragmentation of hypertriton sensitive to its radius.
- Electromagnetic response of the hypertriton also useful. Maybe will become state-of the art probe in the future.
- Nonmesonic decays of light hypernuclei is the only tractable way to study $\Delta S = 1$ weak baryon-baryon interaction.
- FSI important and can be used to our benefit (scattering lengths, probe of spin correlations, etc).