Shape coexistence in gold, mercury and bismuth isotopes studied by insource laser spectroscopy at RILIS-ISOLDE

Andrei Andreyev University of York(UK), JAEA(Tokai, Japan), CERN-ISOLDE on behalf of RILIS-Windmill-ISOLTRAP collaboration

- Reminder on shape coexistence in the lead region
- Isotope Shift(IS) and Hyperfine Splitting (HFS) measurements
- ISOLDE and our detection tools
- Hg chain is there life after staggering?
- Au chain back to sphericity
- Bi chain following the steps of Hg's?



Windmill-ISOLTRAP-IDS-RILIS Collaboration '2018

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MR-TOF@ISOLTRAP team

S. Kreim, V. Manea, F. Wienholtz +...

- Highly collaborative
- 15 institutions, >40 atomic and nuclear physicists
- Many PhD students, both atomic and nuclear physics

Shape coexistence around closed proton and/or neutron shells (and subshells)

- spherical and deformed structures co-exist in the nucleus at low energy
- its study can contribute in finding a unified description for atomic nuclei
- supplies information about the mixing between these configurations



Nilsson Diagrams around Z~82 & 82<N<126 (WS)



Around Z=82 and neutron mid-shell N=102-108, protons and neutrons coherently produce low-lying coexisting oblate and prolate shapes





) Oblate



Prolate

(An example) Modern 'state-of-the art' beyond mean-field calculations (SLy6+GCM) J.M.Yao, M. Bender, P.-H. Heenen, PRC87,034322(2013)







FIG. 4. (Color online) Quadrupole deformation energy surface for ¹⁸⁴Hg, normalized to the absolute minimum and projected on particle numbers. Each contour line is separated by 0.2 MeV. The inset shows the energy as a function of γ deformation along the path joining the two axial minima.

FIG. 13. (Color online) (Upper panel) Variation of the charge radii $\delta \langle r_{ch}^2 \rangle$ for the lowest 0⁺ states in Hg [normalized to the ground state (g.s.) of ¹⁹⁴Hg], Pb (normalized to the g.s. of ¹⁹⁴Pb), Po (normalized to the g.s. of ²¹⁰Po), and Rn (normalized to the g.s. of ²⁰⁴Rn) isotopes, compared to the the experimental data for ground states taken from Refs. [18,49].



Mass Number A R. Julin, K. Helariutta, M. Muikku, J. Phys. G 27 (2001)

Studies in the Lead Region

- Particle decay studies (alpha/beta,E0's)
- In-beam γ -ray studies (RDT/IDT), isomeric states
- Plunger/lifetimes
- Coulex
- Moments

• Optical/Laser spectroscopy: Hyperfine splitting, isotope shift (HFS/IS)

Pre-2003: Charge Radii in the Lead Region



- Shape coexistence around N~104
- Sphericity around N=126, kink in radii, high-spin isomers
- Octupole effects around N~132, inverse odd-even radii staggering

2003-2011: Charge Radii in the Lead Region



The **ISOLDE** facility at **CERN HRS** Target **1.4 GeV Protons From PSB** RIL WISARD NICOLE VITO collections **MR-To** MINIBALL Windmil COLLAPS IDS CRIS

Resonance Laser Spectroscopy of an Atom



•Measuring α/γ or ToF spectra while scanning the frequency

Resonance Laser Spectroscopy of an Atom



- More complex in odd-A (odd-odd-A) cases
- Needs to consider HFS splitting (see examples further in the talk)



Our tools '2018 WindMill (WM), FC, MR-TOF MS



Hg with VADLIS@ISOLDE from Molten Lead Target

Nuclear Instruments and Methods in Physics Research B 376 (2016) 39-45



Blurring the boundaries between ion sources: The application of the RILIS inside a FEBIAD type ion source at ISOLDE



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Windmill Detection System

A. Andreyev et al., PRL 105, 252502 (2010)



- 34% geometrical efficiency at implantation site.
- Alpha-gamma coincidences
- Digital electronics

MR-ToF Mass Spectrometer



MR-ToF MS counts ions, thus is not limited by decay scheme or long half-lives
MR-ToF MS offers a way to separate background for direct single-ion detection using MCP (time scale: tens of ms).

April 2015: IS598@ISOLDE - Hg isotopes (or, what happens after shape staggering in ¹⁸¹⁻¹⁸⁵Hg?)



Previous Hg radii data (ISOLDE) ^{183,185,187,199,201}Hg: J. Bonn et al, PLB38, 308 (1972) ¹⁸¹⁻¹⁹¹Hg: J. Bonn et al., ZPhys A 276, 203 (1976) ¹⁸²⁻¹⁹⁸Hg: G. Ulm Z. Phys. A 325, 247 259 (1986) **Note: gs of** ^{182,184,186}Hg – weakly oblate

In-beam Hg data - examples (RDT, plunger, Coulex) ^{184,186}Hg; L.P. Gaffney LP *et al.* PRC89 024307(2014) ¹⁷²Hg: M. Sandzelius et al, PRC79, 064315 (2009) ¹⁷⁵Hg: D. O'Donnell et al., PRC79, 051304(2009) ^{176,178}Hg: M. P. Carpenter et al, PRL78, 3650(1997) ¹⁷⁶Hg: M. Muikku et al, PRC58, 3033 (1998) ^{180,182}Hg: T. Grahn et al, PRC80, 014324 (2009)

¹⁸⁰Hg@Windmill



¹⁸¹Hg@Windmill



179,185,207,208**Hg**



Isotopes with N>126 ²⁰⁷Hg HFS spectra@MR-ToF, I=9/2 also ²⁰⁸Hg! I=0



HFS spectra and Charge radii for Hg isotopes



Comparison Hg vs Pb chains



S.Sels et al, to be submitted to PRC, July 2018

Density Functional Theory (DTF) Calculations (York-Lyon-Brussels Collaboration)

- Extensive Density Functional Theory blocked calculations performed by the York-Lyon-Brussels Collaboration (14 parametrizations of Skyrme functionals were probed)
- UNEDF0, UNDEF1, UNDED1^{SO}, SLy4, SkM*, SGII
- 8 parametrizations of SLy5sX (different surface tension)
- Variation of pairing strength or surface tension
- Full account of calculations: S.Sels et al, to be submitted to PRC, July 2018

Some details of DFT calculations (changing pairing strenght or surface tension)



FIG. 11. Potential energy surfaces calculated using the UNEDF1^{SO} functional for even-mass mercury isotopes between A = 174 and A=196 as a function of the axial

FIG. 12. Differences between energies of weakly and strongly prolate configurations and those corresponding to the oblate configurations, as calculated with UNEDF1^{SO} (left) and the SLy5sX parameterizations (right). For UNEDF1^{SO} calculations with standard pairing (top), pairing amplified by 8% (middle) and pairing amplified by 20% (bottom) are presented. For the SLy5sX family, results for SLy5s1 (top), SLy5s4 (middle) and SLy5s8 (bottom) are presented.



Some details of DFT calculations (changing pairing strenght or surface tension)





Mean-field theory summary:



- The gs radii staggering in Hg is a subtle effect of an interplay between closelylying weakly oblate/prolate and strongly prolate shapes
- •The effect is extremely sensitive to fine details of interactions, e.g. a slight change of pairing strength in UNEDF1⁵⁰ or change of surface tension in Sly5sX results in similar final pictures.
- This balance also depends on other properties of interactions, e.g. sp energies.
- •Overall, no preference can be given between different approaches

Monte-Carlo Shell Model (MCSM) Calculations for Hg isotopes (T.Otsuka et al)

- Largest calculation of its kind, avoids diagonalization of >2x10⁴²-dimensional H matrix
- ^{132}Sn inert core, with up to 30 valence protons in $g_{7/2}\text{-}i_{13/2}$ and up to $\ 24$ valence neutrons in $h_{9/2}\text{-}j_{15/2}$
- Calculated magnetic (and quadrupole) moments
- Radii staggering due to increased proton occupancy of the $h_{9/2}$ intruder state coupled with larger neutron $i_{13/2}$ occupancy (driven by favourable monopole interaction between them)



IS534@ISOLDE: Charge radii of Au isotopes



Are the light gold isotopes deformed, A(Au)<183?
What are the spins of ground and isomeric states?

Previous radii data (ISOLDE) ¹⁸⁵⁻¹⁹⁰Au : K. Wallmeroth et al, NPA493,224 (1989) ^{183,184}Au: U. Kronert et al, Z.Phys. A331, 521 (1988) ^{184mg}Au: F. Le Blanc et al. PRL79, 2213 (1997)

IS534: Nuclear spins for ^{177,179}Au extracted from HFS spectra



Based on the number of HFS components and their intensity ratio, the gs spins of ^{177,179}Au are experimentally determined as 1/2

(In)Famous Kink at N=126?

PRL **110,** 032503 (2013)

PHYSICAL REVIEW LETTERS

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Charge Radius Isotope Shift Across the N = 126 Shell Gap P. M. Goddard, P. D. Stevenson, and A. Rios



FIG. 1 (color online). Isotope shifts are given by the difference in the mean square charge radius between a series of even isotopes, denoted by their mass number A, and that of ²⁰⁸Pb. Across the N = 126 shell closure, a strong increase in the slope of the experimental data (diamonds) is observed. Theoretical predictions, obtained with different Skyrme parametrizations, are also presented. Only a handful of Skyrme sets are able to reproduce the increase of slope above N = 126.

Occupation of the neutron $1h_{11/2}$ orbital provides a better overlap with majority of n=1 proton orbitals (via symmetry energy), thus driving them to a larger radius?



FIG. 3 (color online). Neutron single-particle energies around the Fermi surface in ²¹⁰Pb for five sets of Skyrme forces. States with a significant BCS occupation, >3%, above N = 126 are emboldened. Whenever the $1i_{11/2}$ state is substantially populated, the kink in isotope shift can be reproduced.