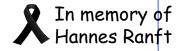
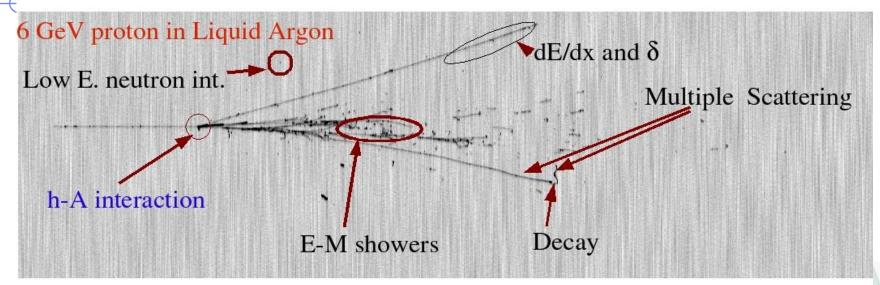


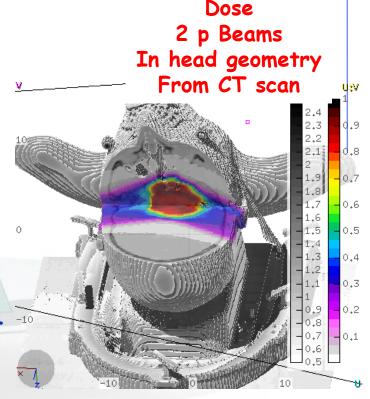
Neutrino interactions in FLUKA: NUNDIS

M. Antonello, G. Battistoni, A. Ferrari, M. Lantz, P. Sala, G. Smirnov

FLUKA: a multi-purpose Monte Carlo code





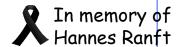


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Web Site: http://www.fluka.org

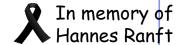
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General framework

- Nuclear models in FLUKA have been developed along the years, initially for hadronnucleus reactions,
- Then extended to treat also
 - Photonuclear reactions
 - Muon-nuclear and electro-nuclear via virtual photon exchange
 - Quasi-eleasic electron scattering
 - Muon capture
 - Neutrino interactions
 - Anti-nucleon reactions
- All sharing the same generalized IntraNuclearCascade + Preequilibrium+Evaporation
 +Gamma deexcitation model
- All DIS share the same fragmentation
- PROs: well tested, valid for all target nuclei, consistent with detector simulations
- CONs: some details of nuclear structure missing, coherent effects have to be inserted ad-hoc



Neutrinos in FLUKA

- Generators of neutrino-nucleon interactions:
 - QuasiElastic
 - Resonance
 - DIS

- Acta Phys.Polon. B40 (2009) 2491-2505 CERN-Proceedings-2010-001 pp.387-394.
- Embedded in FLUKA nuclear models for Initial State and Final State effects
- Only for Argon: absorption of few-MeV (solar) neutrinos on whole nucleus
- Elastic scattering on electrons to be refreshed
- Products of the neutrino interactions can be directly transported in the detector (or other) materials
- Used for all ICARUS simulations/publications



Quasi Elastic

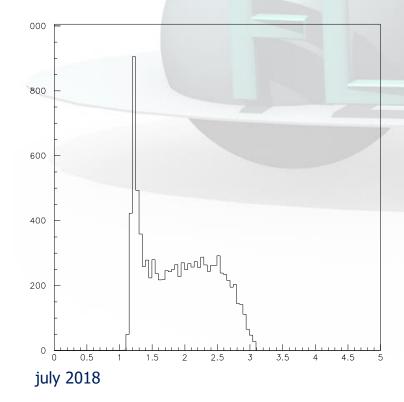
- Following Llewellyn Smith formulation
- $M_A = 1.03$, $M_V = 0.84$
- Lepton masses accounted for
- Polarization of the outgoing lepton is calculated (and used in lepton decay) according to Albright-Jarlskog*

* later applied also to leptons produced in RES/DIS

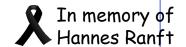


Resonance production

- From Rein-Sehgal formulation
- Keep only Δ production
- No non-resonant background term, assuming that the non-resonant contribution comes from NunDIS
- ullet TRANSITION from RES to DIS: linear decrease of both σ as a function of W

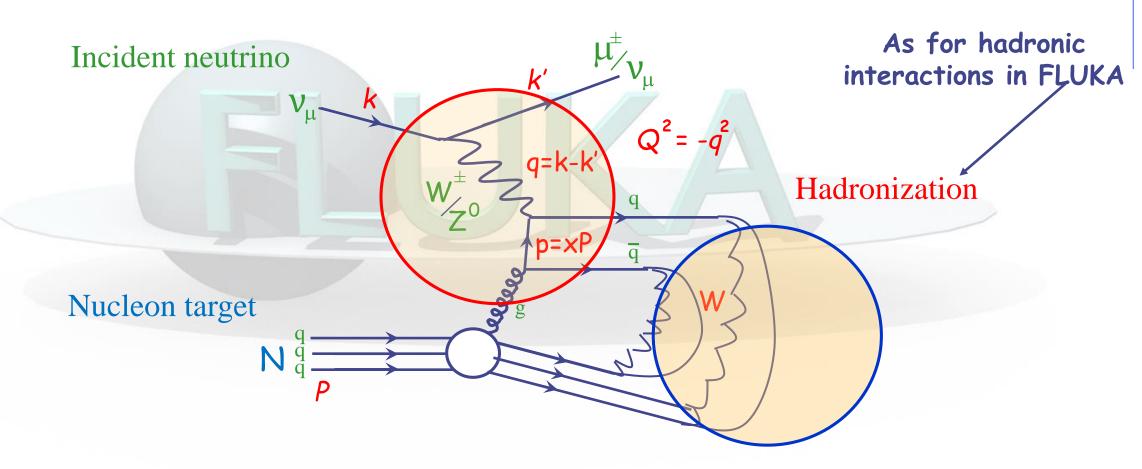


Hadronic mass distribution for ν_{μ} CC on p at 5 GeV



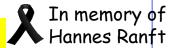
DIS (NUNDIS)

FLUKA hadronization and nuclear interactions work well independently of primary interaction vertex









differential cross sections

$$\frac{d^{2}\sigma}{dxdQ^{2}} = \frac{d^{2}\sigma}{dxdy} \cdot \frac{dy}{dQ^{2}} = \frac{d^{2}\sigma}{dxdy} \cdot \frac{1}{2ME_{\nu}x} ,$$

$$\frac{d^{2}\sigma}{dxdy} = \frac{G_{F}^{2}ME_{\nu}}{\pi(1+Q^{2}/M_{W/Z}^{2})^{2}} \sum_{i=1}^{5} A_{i}(x,y,E_{\nu})F_{i}(Q^{2},x)$$

Structure functions $Fi(\vec{Q},x)$

$$\begin{array}{rcl} F_2^{\nu p}(Q^2,x) & = & 2x[d+\bar{u}+s+\bar{c}] \\ xF_3^{\nu p}(Q^2,x) & = & 2x[d-\bar{u}+s-\bar{c}] \end{array}$$

Callan-Gross relation:
$$F_1 = \frac{F_2}{2x}$$

To be updated to
$$2xF_1(Q^2,x) = F_2(Q^2,x) \frac{1 + 4M^2x^2/Q^2}{1 + R(Q^2,x)}$$

$$F_4 = 0,$$
 $F_5 = \frac{F_2}{x}.$

$$A_1 = y \left(xy + \frac{m_\ell^2}{2ME_\nu}\right)$$

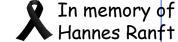
$$A_2 = 1 - y \left(1 + \frac{Mx}{2E_\nu}\right) - \frac{m_\ell^2}{4E_\nu^2}$$

$$A_3 = \pm y \left[x \left(1 - \frac{y}{2}\right) - \frac{m_\ell^2}{4ME_\nu}\right]$$

$$A_4 = \frac{m_\ell^2}{2ME_\nu} \left(y + \frac{m_\ell^2}{2ME_\nu x}\right)$$

$$A_5 = -\frac{m_\ell^2}{ME_\nu}$$





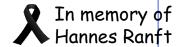
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```
Quark dependence qi(Q,x)
  determined from Parton
Distribution Functions (PDFs)
```

GRV94 Glück et al., Z. Phys. C67 (1995) 433. **G**RV98 Glück et al., Eur. Phys. J. C5 (1998) 461. Bourelly et al., Eur. Phys. J. C23 (2003) 487. BBS J. High Energy Phys. 0207 (2002) 012. CTEQ arXiv:hep-ph/0211080. MRST Phys. Rev. D68 (2003) 014002. Alekhin **DEFAULT OPTION**

NUNDIS WORKS WITH THESE PDFs



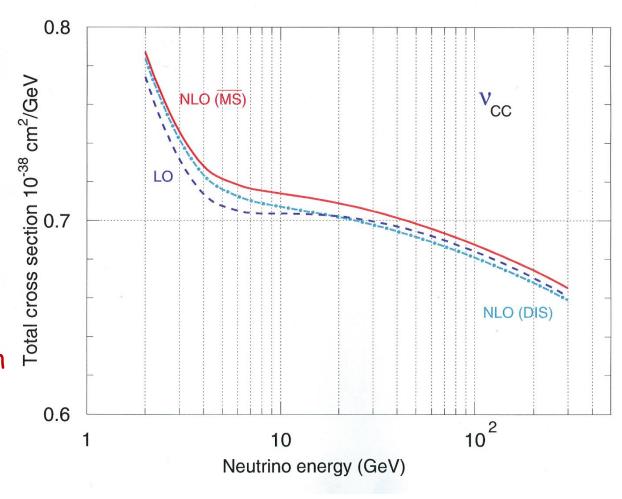


Three versions of pdf from the GRV98 analysis are included as options for evaluating nucleon structure functions

- 1. Leading order analyses (LO)
- 2. Next to leading order analyses (NLO MS-bar)
- 3. Next to leading order analyses (NLO DIS)

An interesting feature of the GRV98 analysis is a low threshold for the transferred , 4-momentum, Q^2 = 0.8 GeV^2

NLO (DIS) is chosen as a default option



M. Gluck, E. Reya and A. Vogt, Eur. Phys. J. C5 (1998) 461

Extrapolation from $Q^2 = 1.0 \text{ GeV}^2$ to $Q^2 = 0$

In memory of Happes Ranft

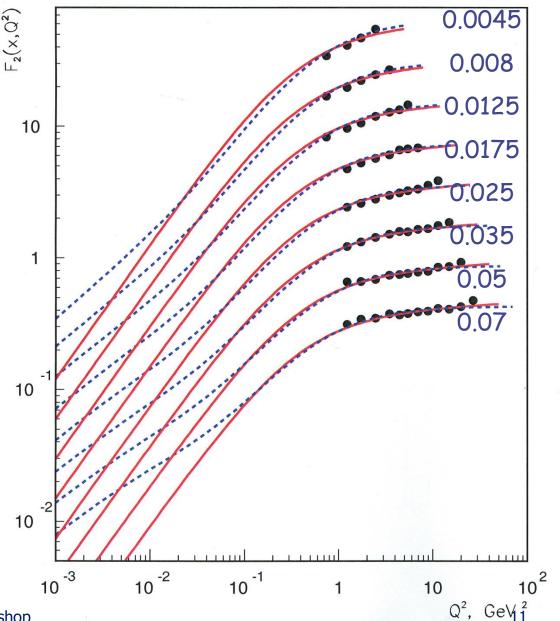
Solid lines: M. Bertini et al. 1996 (Default in NUNDIS)

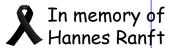
$$F_2(x, Q^2) = A[1 + \epsilon \ln(Q^2(1/x - 1) + M^2)] \ln(1 + Q^2/(Q^2 + a^2))$$
.

Dashed lines: Donnachie-Landshoff 1994

$$F_2(x, Q^2) \sim Ax^{-0.0808} \left(\frac{Q^2}{Q^2 + a}\right)^{1.0808} + Bx^{0.4525} \left(\frac{Q^2}{Q^2 + b}\right)^{0.5475}$$

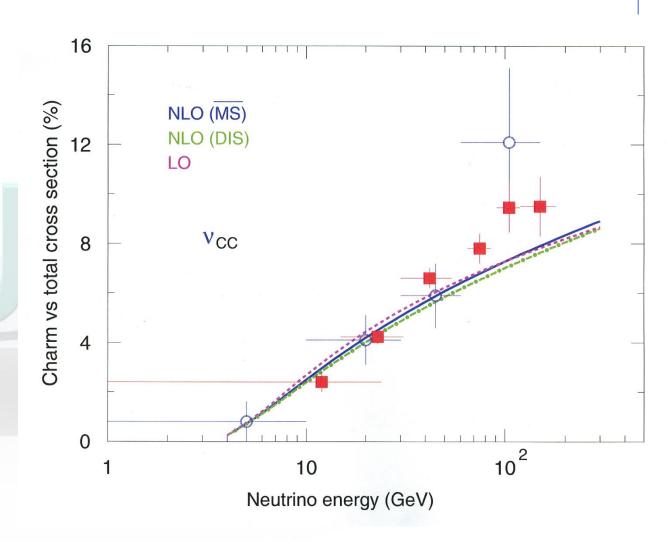
data points from NMC Collab., M. Arneodo et al., Nucl. Phys. B 483 (1997) 3-43 Data/cuves scaled for clarity, factors from 1 to 128

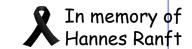




Charm production in neutrino interactions

- Ratio of the charm to total cross sections
- Results of NUNDIS simulation with M_c = 1.35 GeV (curves) and experimental data: E531 (open circles) and CHORUS-2011 (filled squares).



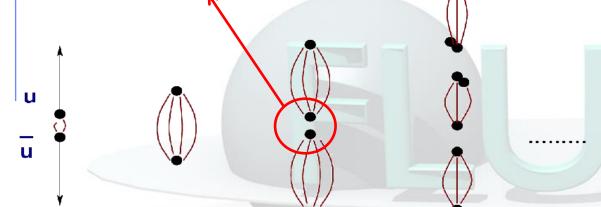


The "hadronization" of chains

In FLUKA:

An example:

Low mass chain: just 2-3 meson/(anti)baryon resonances



Assumes chain universalityFraamentation functions fr

Fragmentation functions from hard processes and e+e-scattering

Transverse momentum from uncertainty considerations

Mass effects at low energies

The same functions and (few)
parameters for all reactions and
energies

ud

ud

 $d\overline{u}$

uud •

udd

us

qq

 $q\overline{q}$

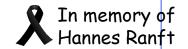
du

Chains from v DIS:

 One quark-diquark chain if interaction of valence quark

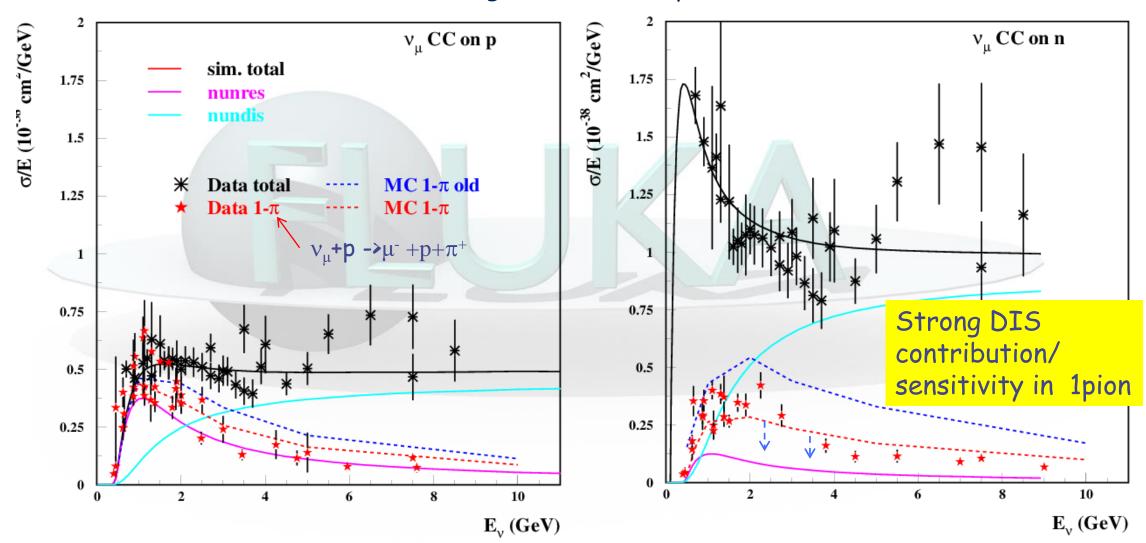
 One quark-diquark plus one qqbar chain if int on sea quark

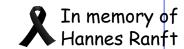
gradual transition of low energies chains to "phase space explosion" constrained in p_{T} , including baryons, mesons, resonances.



Single pion production

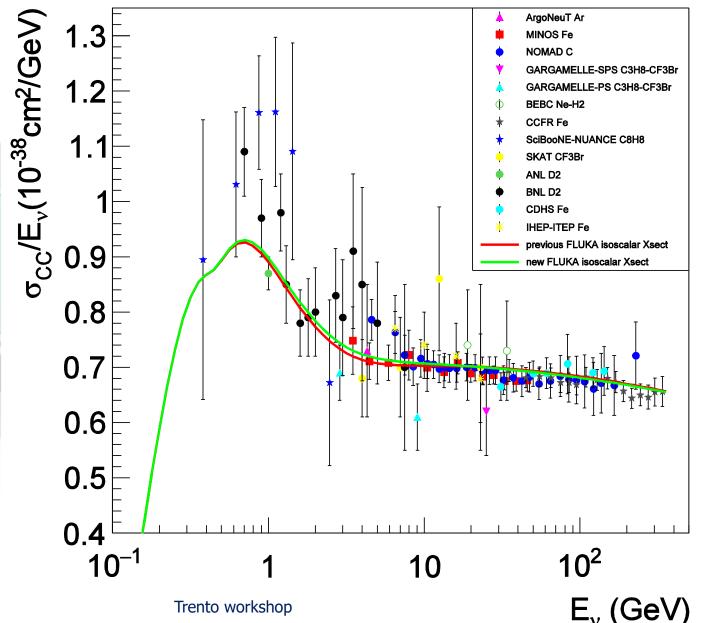
New low-mass chain treatment of fragmentation -> improvements in the RES-DIS transition





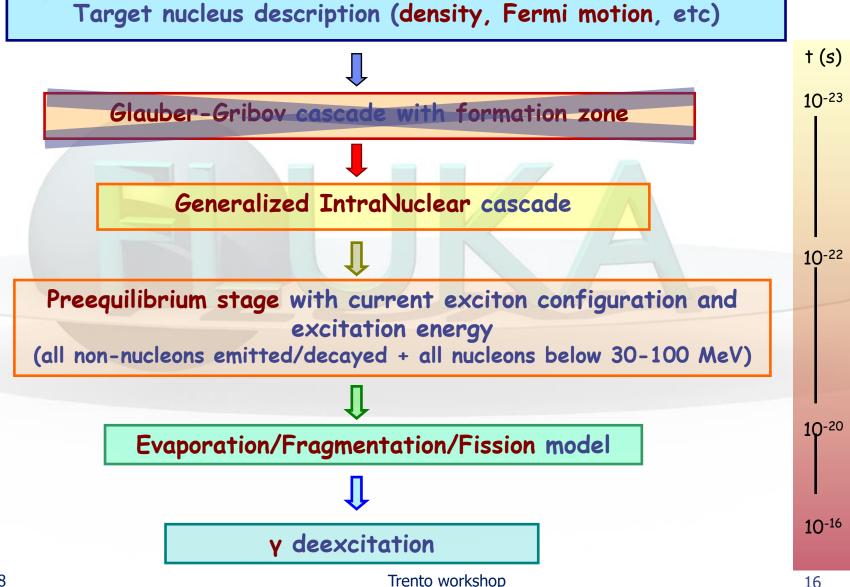
Comparison with data on total cross section

Isoscalar v_{μ} - Nucleon total CC cross section Fluka (lines) with two pdf options Vs Experimental data



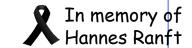
Nuclear interactions in FLUKA: the PEANUT mode





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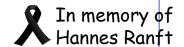
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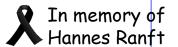
(Generalized) IntraNuclear Cascade Primary and secondary particles moving in the nuclear medium

- Target nucleons motion and nuclear potential well according to local Fermi gas model
- Interaction probability σ_{free} includes inelastic σ_{free} + Fermi motion × $\rho(r)$ + exceptions (ex. π)
- Glauber cascade at higher energies
- Classical trajectories (+) nuclear mean potential (resonant for π)
- Curvature from nuclear potential \rightarrow refraction and reflection
- Interactions are incoherent and uncorrelated
- Interactions in projectile-target nucleon CMS
- Fully relativistic
- Multibody absorption for π , μ
- Special for K^- , antinucleon, π (phase shifts, annihilation)
- Quantum effects (Pauli, formation zone, correlations...)
- Exact conservation of energy, momenta and all additive quantum numbers, including nuclear recoil
- First excited nuclear levels accounted for (more levels in evaporation/gamma deexc)

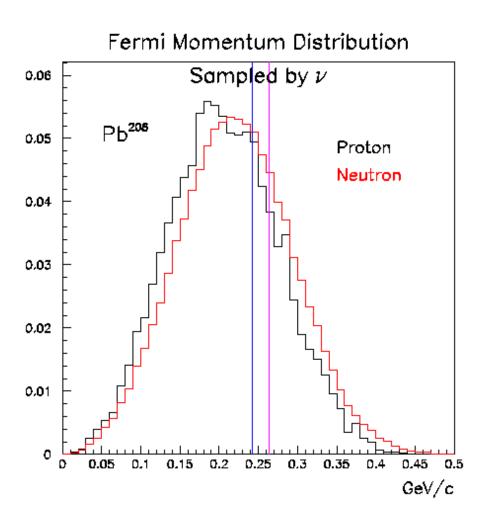


Nucleon Fermi Motion in FLUKA

- Momentum smearing according to uncertainty principle assuming a position uncertainty = $\sqrt{2}$ fm
- Nuclear density given by symmetrized Woods-Saxon for A>16 and by a harmonic oscillator shell model for light isotopes
- Proton and neutron densities are different

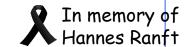


Example of Fermi distribution

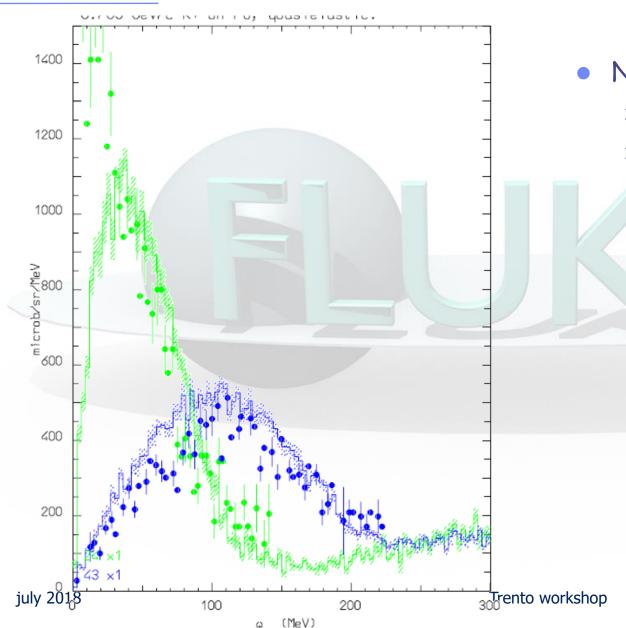


Fermi momentum distribution as "seen" by interacting neutrinos on lead.

Vertical lines: maximum Fermi momentum according to un-smeared distribution



Positive kaons as a probe of Fermi motion



K⁺ and K⁰

- No low mass S=1 baryons
 - > weak K+N interaction
 - > Only elastic and charge exchange up to ≈ 800 MeV/c

 $K^+ Pb \rightarrow K^+ Pb 705 MeV/c$

Residual excitation spectrum

With K^{+'} at 24^o (green)

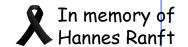
at 43° (blue)

Histogram: FLUKA

Dots: data (Phys. Rev. C51,669 (1995))

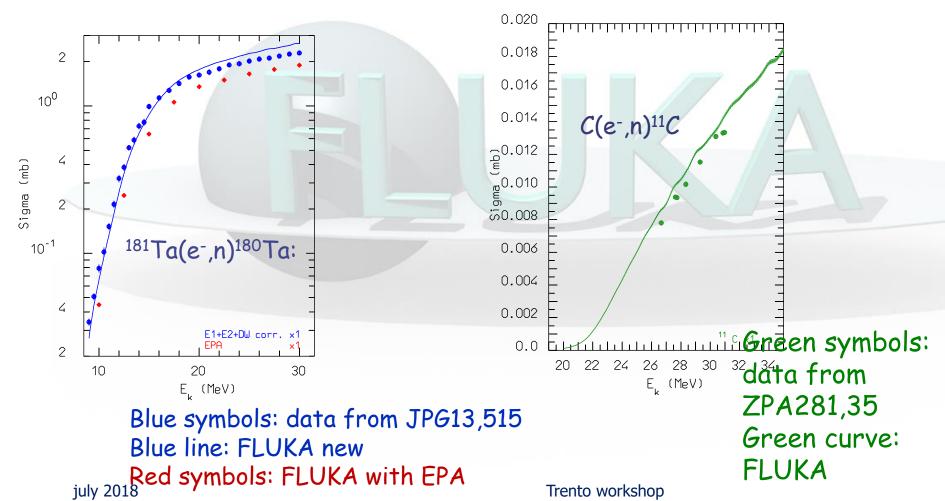
On free nucleon: recoil at 43 MeV or 117 MeV

O-deg tail is elastic on nucleus, not included in sim



Electron scattering

- Quasi-Elastic on nucleon (+ all nuclear)
- Inelastic via virtual photon exchange, recently improved (E1+E2)

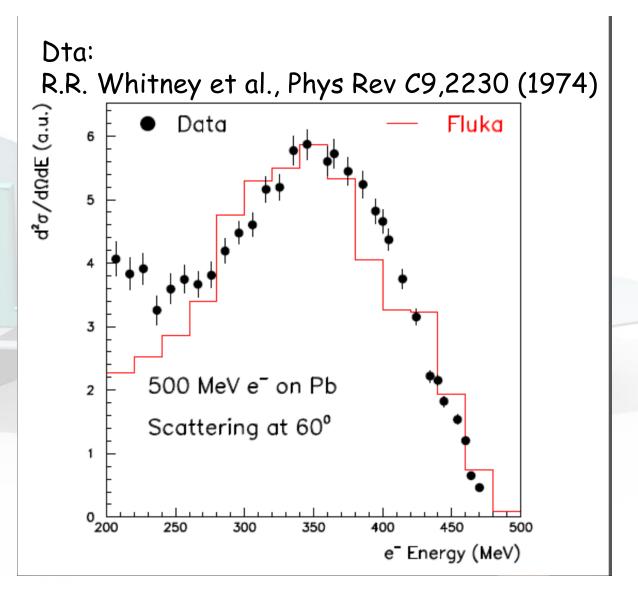


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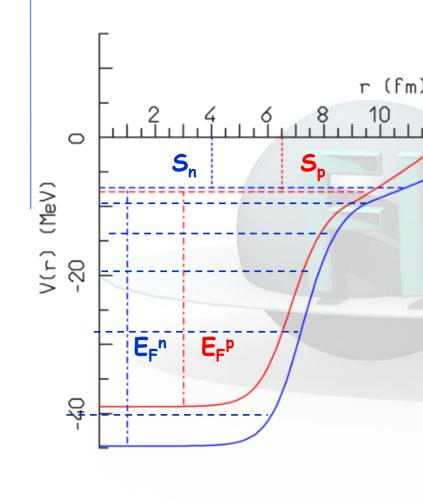
First checks with electrons

- Quasi-Elastic scattering of electrons on Lead, outgoing electron spectrum at 60°
- Inelastic tail not included in simulation
- To be improved wit the inclusion of energydependent nuclear well, as already there for nucleon-induced reactions
- Much more tests needed



Nucleon levels inside the nuclear potential: schematic drawing

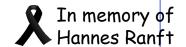




> Blue: neutron

> Red: proton

- p/n are grouped on energy levels space according to the shell model (blue lines, shown for neutrons)
- Depending on the level, the maximum radius for the corresponding nucleon is less or equal to the nucleus radius (equal for the nucleons on the Fermi level)
- ☐ The potential well depth depends on the nucleon energy (not yet implemented for neutrino and electron interactions)
- ☐ Hit nucleon must go above Fermi level, can stay below separation energy.



Nucleon correlation function:

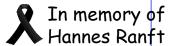
Correlation function: it can be computed within the Fermi-gas model

Due to the anti-symmetrization of the fermion's wave function, given a nucleon in a position \vec{r} in a nucleus with density ρ_0 , the probability of finding another like nucleon in a position \vec{r} is decreased for small values of the distance $d=|\vec{r}-\vec{r}'|$ by a factor

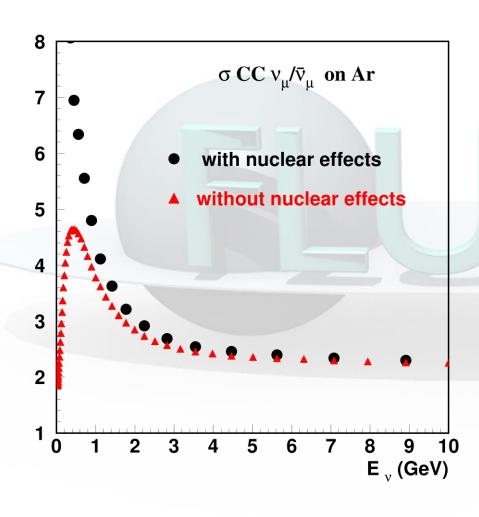
$$g(x) = 1 - \frac{1}{2} \left[\frac{3}{x^2} \left(\frac{\sin x}{x} - \cos x \right) \right]^2$$

where $x = K_F d$, and the factor 1/2 in front of the parenthesis accounts for the two possible spin orientations.

Nucleon hard core effects are also taken into account, forbidding to "find" a nucleon of the same or different type at less than 1-1.5 fm distance. This check is applied at every possible re-interaction, checking against all nucleons already involved in previous interactions

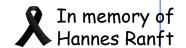


Effect of Pauli Blocking: example

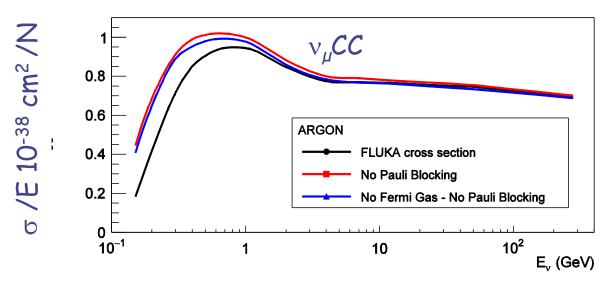


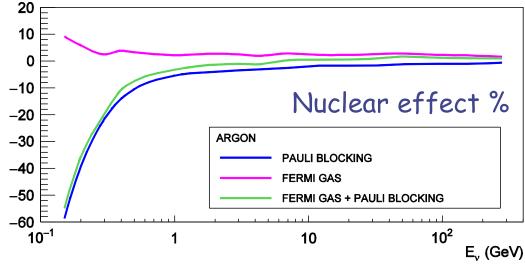
Ratio of Neutrino/antineutrino σ CC vs (a)neutrino energy For interactions in Ar nuclei, v_{μ} As calculated with FLUKA Black: full calculation Red: simple sum of v-N cross section

Smaller q² in anti-neutrino results in higher Pauli-blocking probability



Total cross section: nuclear effects in Ar



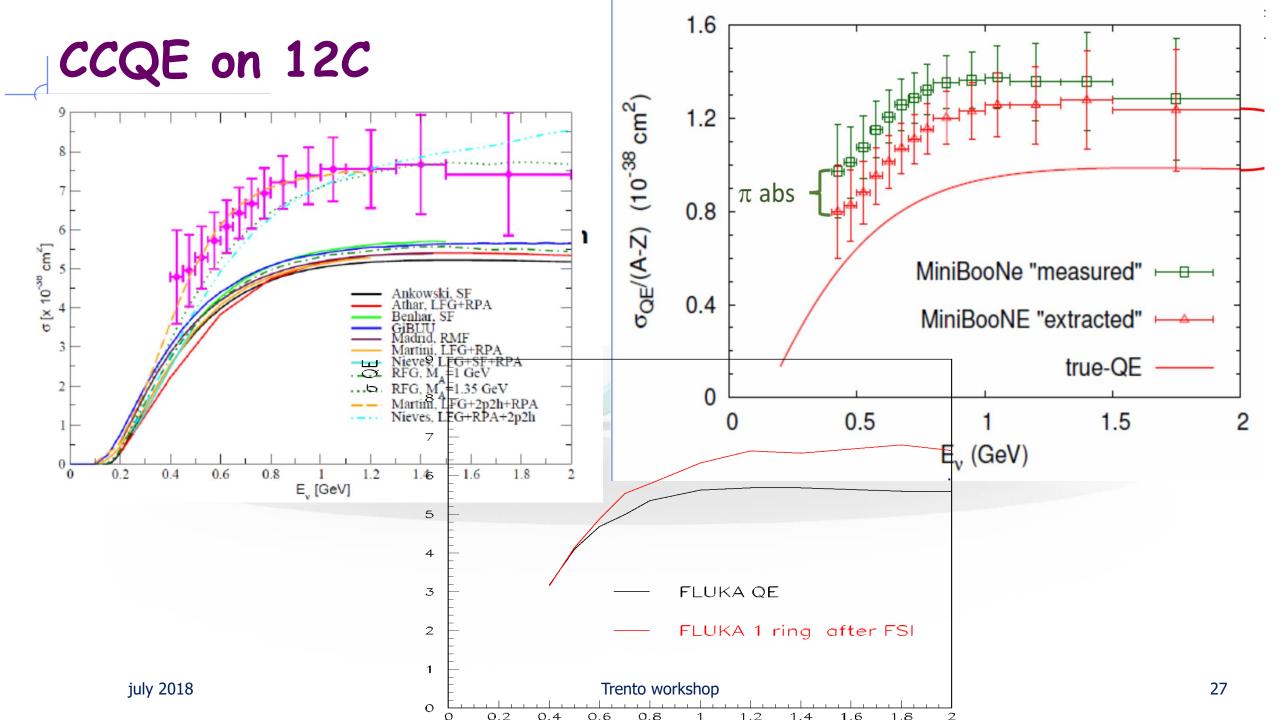


5 GeV < Ev < 50 GeV
Pauli Blocking effect
and Fermi Gas effect
separately have an
impact of ~ 2-3%
Globally Nuclear
effects stay within
±1%

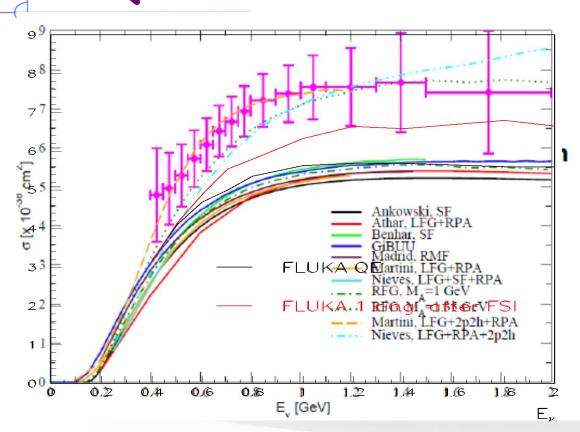
Ev < 5 GeV

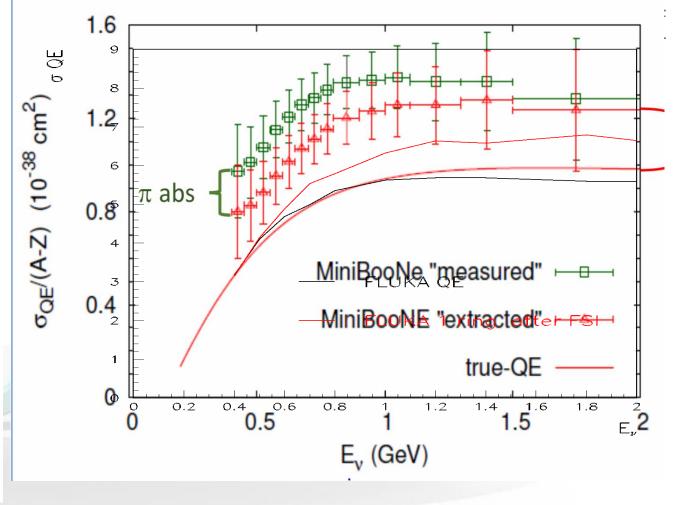
nuclear effects are dominated by the Pauli Blocking and rapidly increase to the order of 10% and above

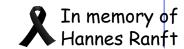
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CCQE on 12C







Nuclear effects in Minerva

Beam: $v\mu$ NuMi Low Energy (average 4 GeV) Main Target : CH

Measured also with C, Fe, Pb targets PRL 112, 231801 (2014)

Here: ratio of cross sections / the one in CH

Left: total CC vs neutrino Energy:

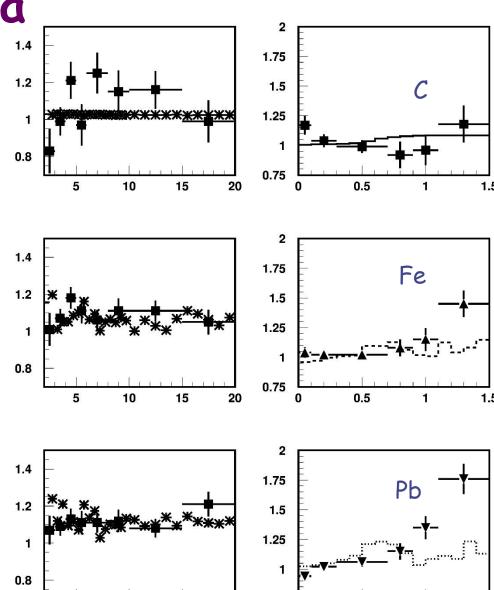
squares: data

crosses: FLUKA

Right: do/dx

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symbols: data histos: Fluka expt: reduction at low x and enhancement at high x with incr. A Fluka: fails the highest x (same for Genie)



0.75

Trento workshop



Pions: nuclear medium effects

july 2018

Non resonant channel Free π N interactions \Rightarrow

P-wave resonant Δ production

 Δ in nuclear \Rightarrow decay \Rightarrow elastic scattering, charge exchange $\stackrel{\text{medium}}{\Longrightarrow} \text{reinteraction} \stackrel{\textstyle >}{\Longrightarrow} \text{Multibody pion absorption}$

Assuming for the free resonant σ a Breit-Wigner form with width $\sigma_{res}^{Free} = \frac{8\pi}{p_{cms}^2} \frac{M_{\Delta}^2 \Gamma_F^2(p_{cms})}{\left(s - M_{\Delta}^2\right)^2 + M_{\Delta}^2 \Gamma_F^2(p_{cms})}$

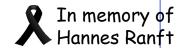
An ``in medium'' resonant σ (σ^A_{res}) can be obtained adding to Γ_F the imaginary part of the (extra) width arising from nuclear medium

 $\frac{1}{2}\Gamma_T = \frac{1}{2}\Gamma_F - \text{Im}\Sigma_{\Delta} \quad \Sigma_{\Delta} = \Sigma_{qe} + \Sigma_2 + \Sigma_3 \quad \text{(Oset et al., NPA 468, 631)}$ quasielastic scattering, two and three body absorption

The in-nucleus σ_t^A takes also into account a two-body s-wave absorption σ_s^A derived from the optical model

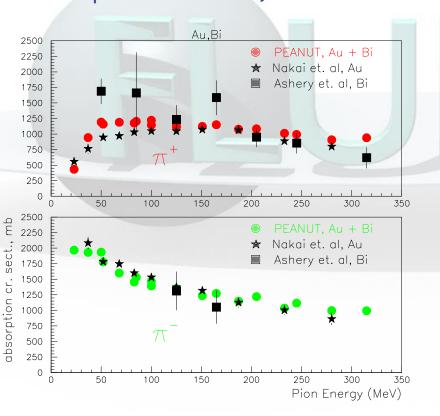
$$\sigma_t^A = \sigma_{res}^A + \sigma_t^{Free} - \sigma_{res}^{Free} + \sigma_s^A \quad \sigma_s^A(\omega) = \frac{4\pi}{p} \left(1 + \frac{\omega}{2m}\right) \operatorname{Im} B_0(\omega) \rho$$
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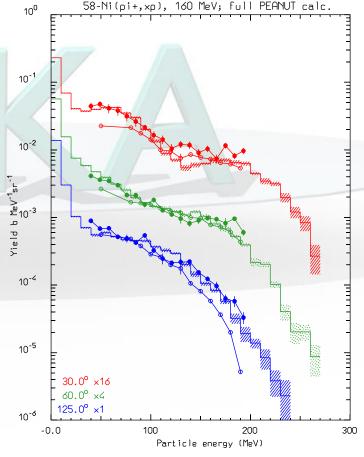


Pion absorption examples

Pion absorption cross section on Gold and Bismuth in the Δ resonance region (multibody absorption in PEANUT)

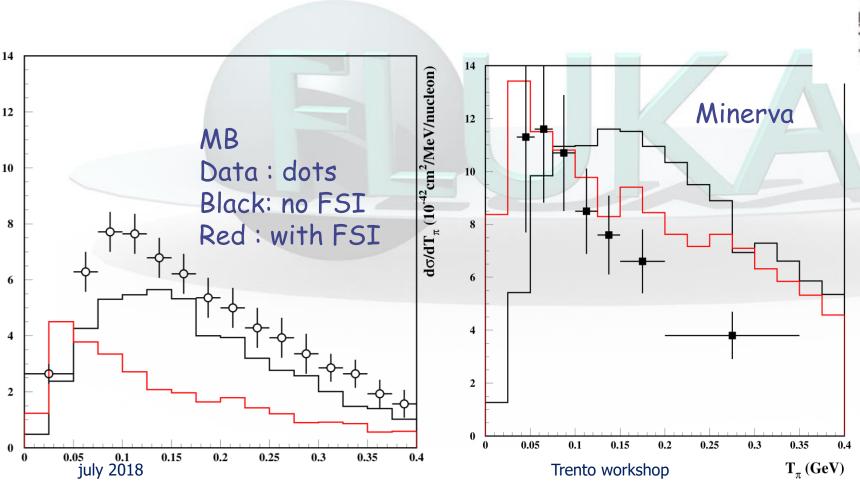


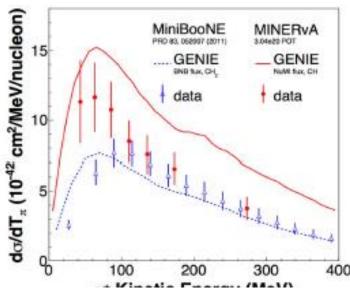
Emitted proton spectra at different angles , 160 MeV π^+ on 58 Ni Phys. Rev. C41,2215 (1990) Phys. Rev. C24,211 (1981) Proton spectra extend up to 300 MeV



MB and Minerva



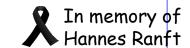




MiniBoone: CH₂, Ev≈0.8 GeV, cut on single pion, PHYS. REV.D 83, 052007 (2011)

Minerva : CH, Ev \approx 4 GeV, cut on W<1.4 arXiv:1406.6415v3 (2015)

Tension betw the two data sets vs models/ extent of FSI



another FSI: Formation zone

Naively: "materialization" time (originally proposed by Stodolski).

Qualitative estimate:

In the frame where p/l = 0

$$\bar{t} = \Delta t \approx \frac{\hbar}{E_T} = \frac{\hbar}{\sqrt{p_T^2 + M^2}}$$

Particle proper time

$$\tau = \frac{M}{E_T}\bar{t} = \frac{\hbar M}{p_T^2 + M^2}$$

Going to the nucleus system

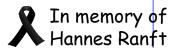
$$\Delta x_{for} \equiv \beta \ c \cdot t_{lab} \approx \frac{p_{lab}}{E_T} \bar{t} \approx \frac{p_{lab}}{M} \tau = k_{for} \frac{\hbar p_{lab}}{p_T^2 + M^2}$$

Condition for possible reinteraction inside a nucleus:

$$\Delta x_{for} \le R_A \approx r_0 A^{\frac{1}{3}}$$

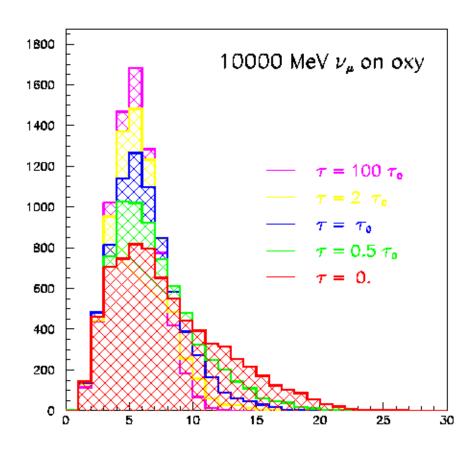
Decrease of the reinteraction probability

Applied also to DIS neutrino interactions and, in an analogue way, to QE neutrino interactions

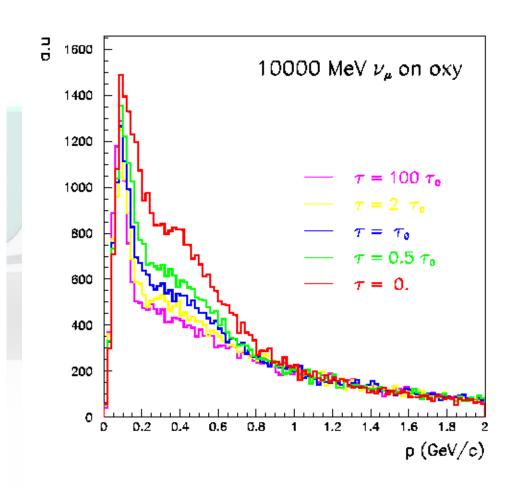


Effect of formation zone

Total hadron multiplicity

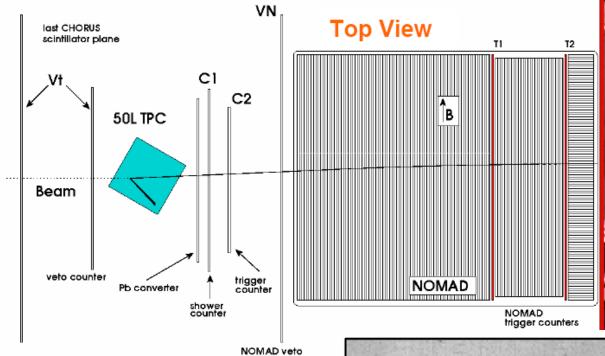


Charged hadron spectra



The 501 LAr TPC in the WANF

neutrino beam (1997)





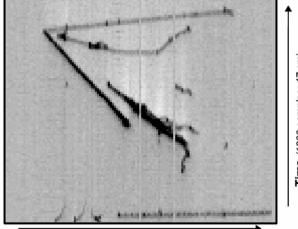
Trigger and μ reconstruction: NOMAD

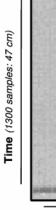
Event selection: "GOLDEN sample"

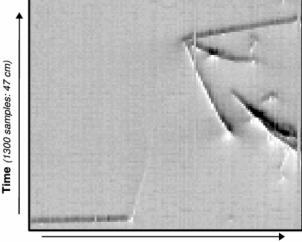
july 2018

= 1 μ and 1 proton >40MeV fully contained

Phys.Rev. D74 (2006) 112001



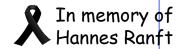




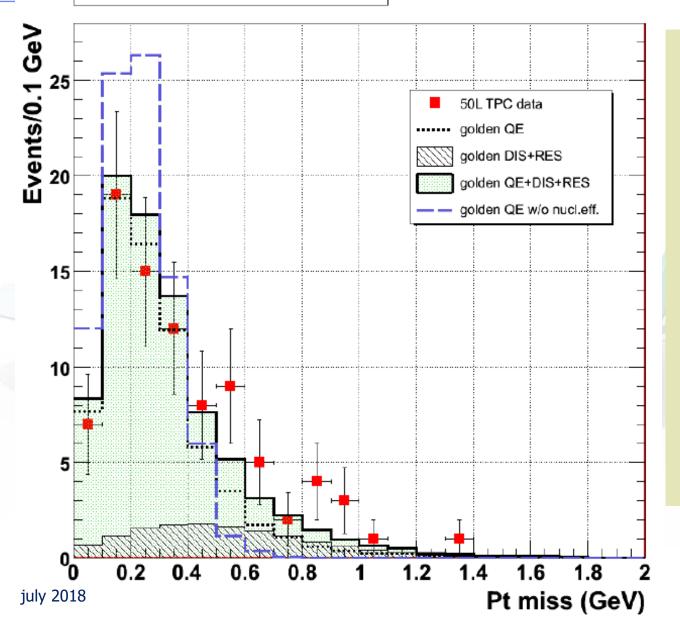
Collection wires. (128 wires: 32 cm.)

Induction wires. (128 wires: 32 cm.)

Ranft







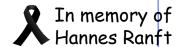
- from 400 QE golden fraction
 16%
 - background additional 20%
 finally expected

80±9(stat.)±13(syst.→ mainly QE fraction and beam simul)

to be compared with 86 events observed

Very good consistency with expectations

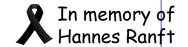
Note: here DIS and RES from old coupling with the NUX code (A. Rubbia)



Last steps of the reaction

- The INC in FLUKA is carried on until all involved nucleons drop below 30-50 MeV kinetic (smooth threshold, depends on number of excited nucleons)
- A PRE-EQUILIBRIUM steps comes in: nucleon-hole pairs sharing statistically the residual excitation energy. Exciton number increased by "collisions", particle emission still possible.
- When the exciton number reaches equilibrium, EVAPORATION / FISSION comes in.
 Statistical, includes nucleons and heavy fragments, includes sub-barrier emission,
 takes into account single excited levels.
- At excitation energies < separation energy→ emission of gamma rays (actually also in competition with evaporation). Uses atlas of excited levels/transitions whenever available.

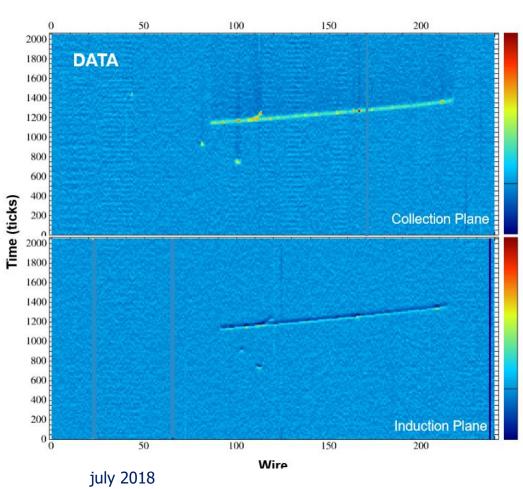
First Demonstration of LArTPC MeV-Scale Physics in ArgoNeuT





Ivan Lepetic APS_April2018

ILLINOIS INSTITUTE OF TECHNOLOGY For the ArgoNeuT Collaboration

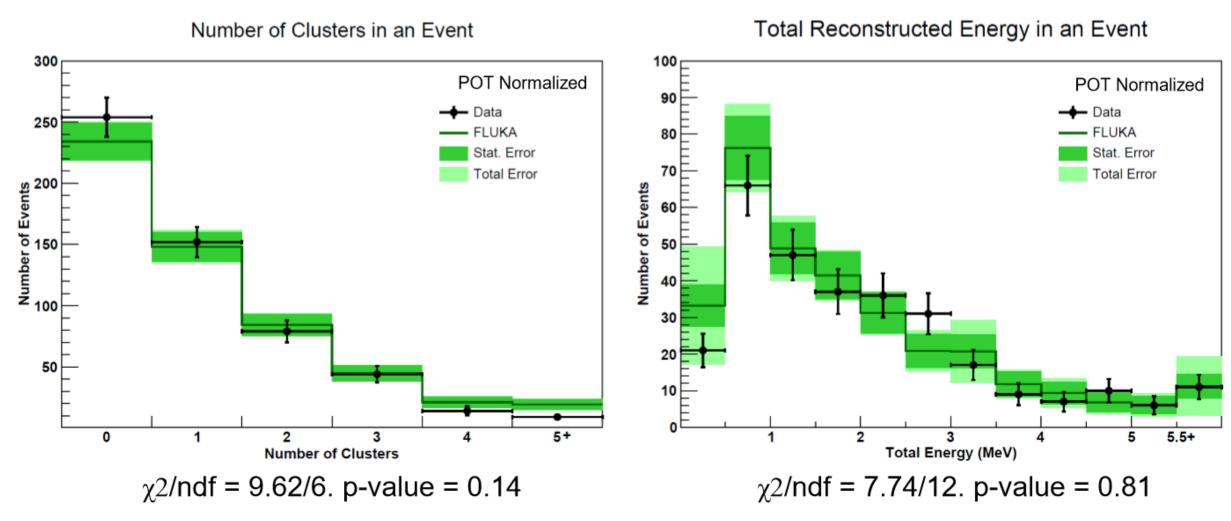


Low-energy gammas produced by neutrino-nucleus interactions in ArgoNeuT

Photons from neutrino-produced nuclear de-excitation and inelastic neutron scattering -

e.g.
$$v_{\mu}$$
 + 40Ar $\rightarrow \mu$ - + p + 39Ar*
 \rightarrow 39Ar + γ

Data and FLUKA

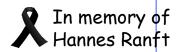


Agreement is far worse if either de-excitation or neutron produced gammas are removed.



Conclusions and perspectives

- A neutrino event generator (NUNDIS) is implemented in FLUKA
- QE, RES, DIS interactions
- Hadronization as for hadronic interactions in FLUKA
- Nuclear effects from the FLUKA nuclear models
- Encouraging comparisons with expt data
- More has to be done:
- Coherent pion production
- \triangleright Coherent effects (see high x in Minerva and proton pairs in Argoneut)
- > More coherent / nuclear structure effects for low energy QE
- \triangleright Meson exchange in QE (high x in Minerva)
- Radiative corrections in DIS (ongoing)
- Comparisons against data







The FLUKA International Collaboration



In memory of Hannes Ranft

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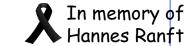
A. Fassò, M.V. Garzelli, E. Gadioli, J. Ranft

M. Chin, Malaysia









Nuclear potential for pions

For pions, a complex nuclear potential can be defined out of the π -nucleon scattering amplitude to be used in conjunction with the Klein-Gordon equation

$$\left[\left(\omega - V_c \right)^2 - 2\omega U_{opt} - K^2 \right] \Psi = m_\pi^2 \Psi$$

In coordinate space (the upper/lower signs refer to π^+/π^-):

$$2\omega U_{opt}(\omega,r) = -\beta(\omega,r) + \frac{\omega}{2M} \nabla^2 \alpha(\omega,r) - \nabla \frac{\alpha}{1 + g\alpha(\omega,r)} \nabla$$

$$\beta = 4\pi \left[\left(1 + \frac{\omega}{M} \right) \left(b_0(\omega) \mp b_1(\omega) \frac{N - Z}{A} \right) \rho(r) + \left(1 + \frac{\omega}{2M} \right) B_0(\omega) \rho^2(r) \right]$$

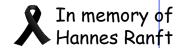
$$\alpha = 4\pi \left[\frac{1}{\left(1 + \frac{\omega}{M}\right)} \left(c_0(\omega) \mp c_1(\omega) \frac{N - Z}{A} \right) \rho(r) + \frac{1}{\left(1 + \frac{\omega}{M}\right)} C_0(\omega) \rho^2(r) \right]$$

Using standard methods to get rid of the non-locality, in momentum space

$$2\omega U_{opt}(\omega, r) = -\beta - K^2 \frac{\alpha}{1 + g\alpha} + \frac{\omega}{2M} \nabla^2 \alpha$$

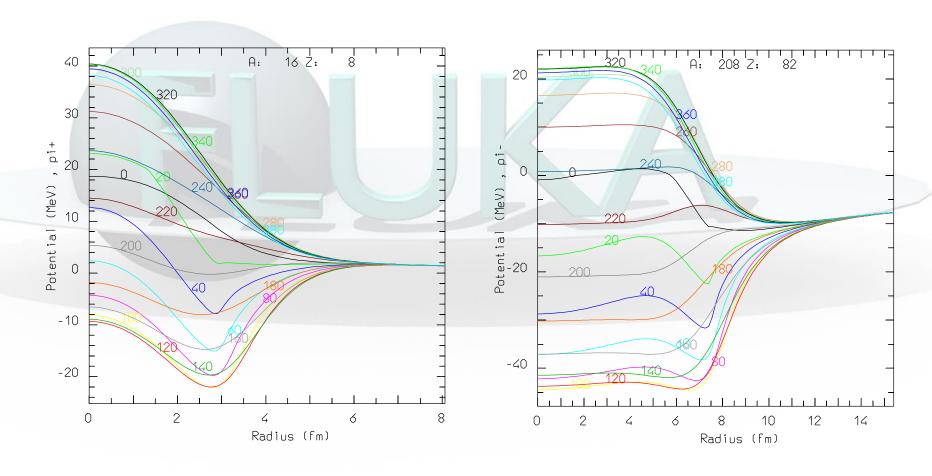
$$K^{2} = k_{0}^{2} + V_{c}^{2} - 2\omega V_{c}^{2} - 2\omega U_{opt}(\omega, r) = \frac{k_{0}^{2} + V_{c}^{2} - 2\omega V_{c}^{2} + \beta - \frac{\omega}{2M} \nabla^{2} \alpha}{1 - \overline{\alpha}}$$

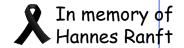
$$\overline{\alpha} = \frac{\alpha}{1 + g\alpha}$$



Nuclear potential for pions: examples

The real part of the pion optical potential for π^- on ^{16}O (left) and π^+ on ^{208}Pb (right) as a function of radius for various pion energies (MeV)



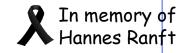


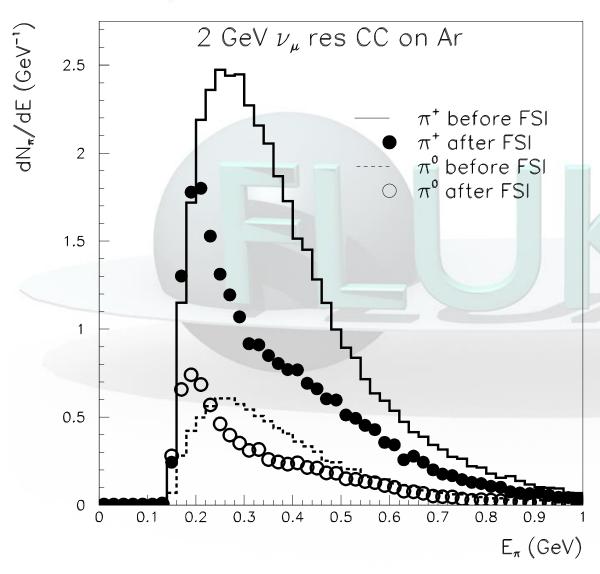
NUNDIS 2015: kinematics

 Considered kinematical limits for the *PDF* available from GRV94, GRV98, and BBS analyses.

	Required	GRV94		GRV98		BBS	
Variable		Default	Tested	Default	Tested	Default	Tested
$E_{min} ({ m GeV})$	_	0.050					
$E_{max} { m (GeV)}$	\geq 10 4	$70 \cdot 10^{3}$			${\bf 10}^5$		
$Q^2_{min} \; ({ m GeV^2})$	\leq 5.5·10 ⁻¹²	0.4	0.4	0.8	0.8	2	0.8
$Q^2_{max}~({ m GeV^2})$	\geq 1.9 \cdot 10 4	10^{6}	10^{9}	10^{6}	10^{9}	10^4	$2\cdot 10^4$
$oldsymbol{x}_{min}$	$\leq 1.4 \cdot 10^{-11}$	10^{-5}	10^{-30}	10^{-9}	10^{-30}	10^{-4}	10^{-30}
x_{max}	1	0.99999	0.99999	1	1	1	1

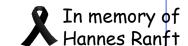




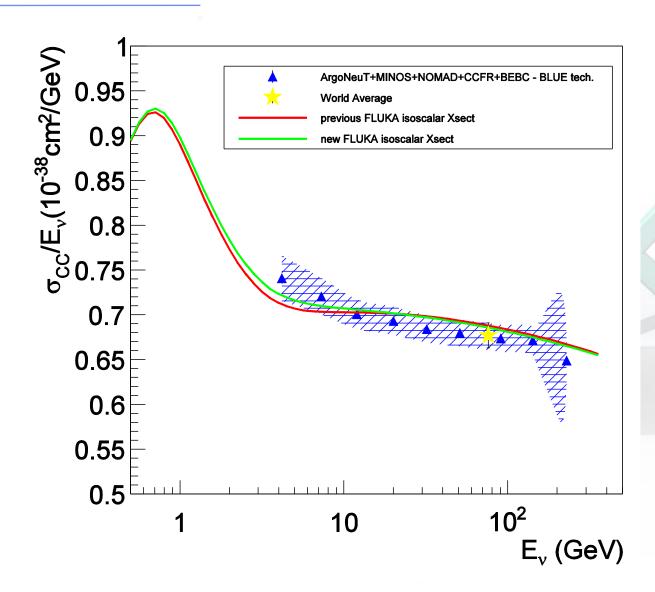


Example of expected effect: $2 \text{ GeV } v_{\mu} \text{ CC RES}$ interaction in Ar: Pion production vs pion total E Lines: before FSI Symbols: after FSI

Solid and filled symbols: positive pions Dashed and open symbols: pizero



Same, with evaluation of data systematics

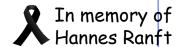


Work in progress: Attempt to compare with a combined estimate from available data and relative systematic error, properly accounting for correlations

Focus on the CNGS energy range (5-30 GeV)

Recent experiments (like MINOS, NOMAD, CCFR 1997): measure the shape of neutrino flux and get the Absolute normalization from Old measurements at high energy, performed using Narrow Band Beams (CCFR-E701 / CCFRR-E616 / CDHS) or Wide Band Beams (GARGAMELLE / BEBC)

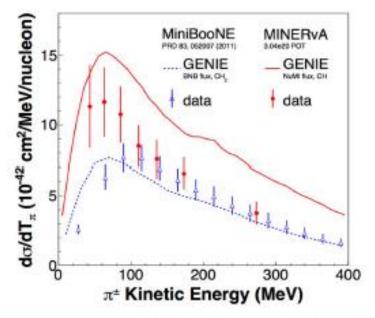
→ Common systematic errors



Data on pion production

Thoughts on MINERVA vs. MiniBooNE

- Shapes very similar, no significant dip in either!
- Small difference in slope (Kinematics, FF, nonres differences).
- Biggest difference is at low energy.

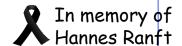


MiniBoone: $CH_{2,}$ Ev ≈ 0.8 GeV, cut on single pion, PHYS. REV.D 83, 052007 (2011)

Minerva: CH, Ev ≈ 4 GeV, cut on W<1.4 arXiv:1406.6415v3 (2015)

Tension betw the two data sets vs models/ extent of FSI

july 2018



Pion production in p-p collisions:

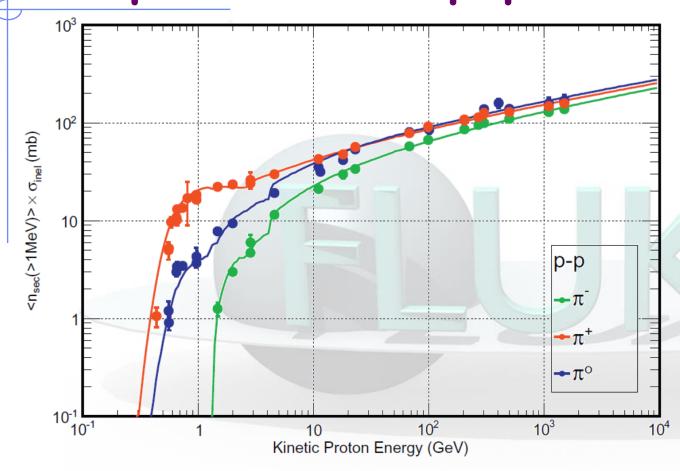
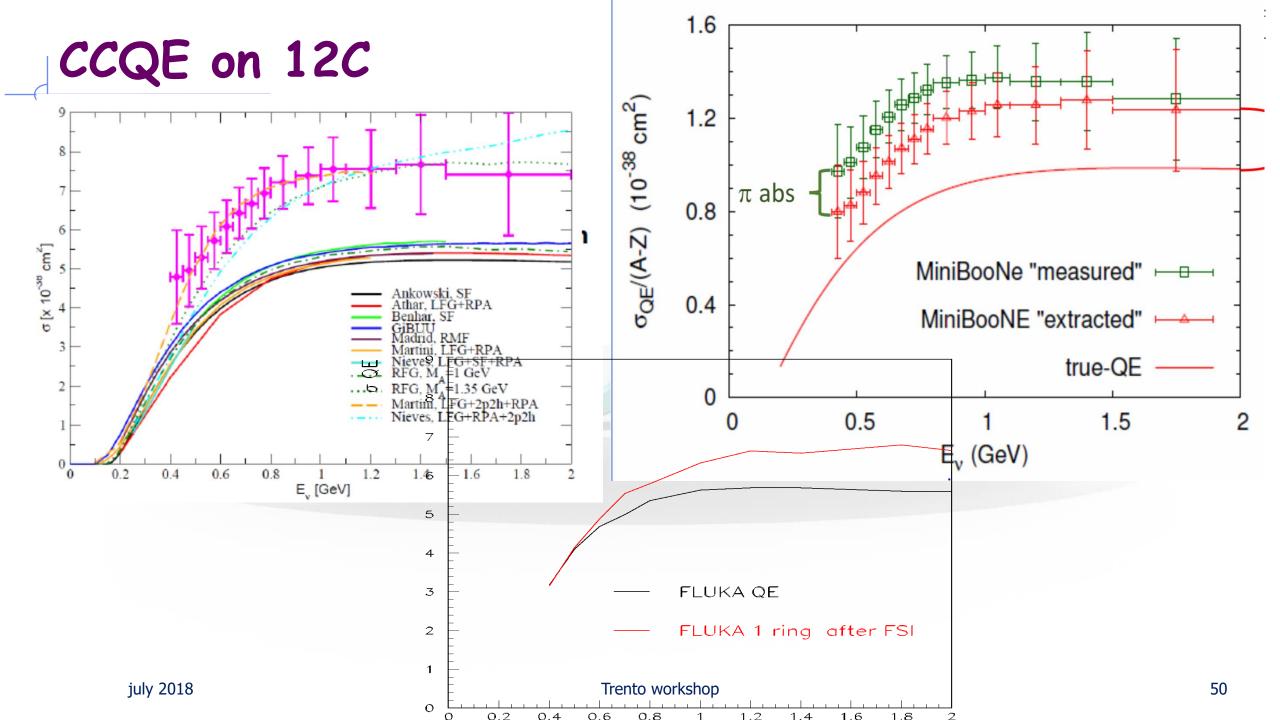


Fig. 2. Inclusive cross sections for the production of π^0 (blue), π^+ (red) and π^- (green) in p–p collision as function of the incoming proton kinetic energy. Lines: FLUKA simulation; points: data from Ref. [28]. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

Inclusive cross section for the production of π^0 (blue), π^+ (red), and π^- (green) in p-p collisions as a function of the proton kinetic energy. Lines: simulations, symbols exp. Data. (figure from AstrPhys81, 21 (2016))



CCQE on 12C

