





INTERACTIONS MODELING IN DUNE

Jarek Nowak for the DUNE Collaboration Modeling neutrino-nucleus interactions, ECT-Villa Tambosi July 12, 2018



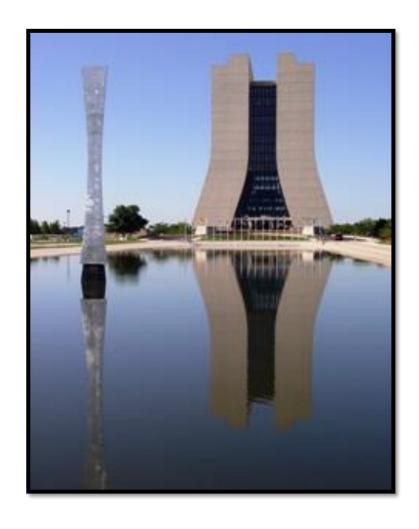
Lancaster Lancaster

University





DEEP UNDERGROUND NEUTRINO EXPERIMENT





- DUNE Experiment: Detectors & Status
- DUNE Physics
- Interaction Monte Carlo generators @DUNE

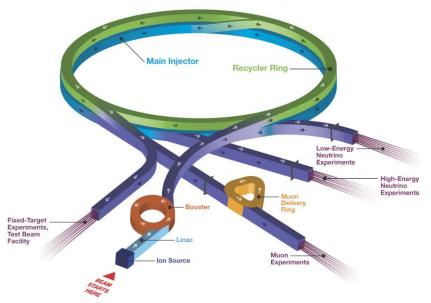
DUNE Experiment

Observe ν_e appearance and ν_μ disappearance at long baseline in wideband beam to measure MH, CPV, and neutrino mixing parameters in a single experiment. Deep underground location reduces cosmogenic background and enables sensitivity to low-energy physics.

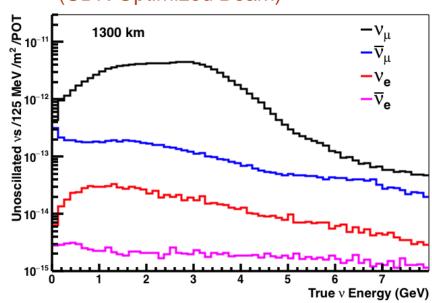


LBNF Beam

Fermilab Accelerator Complex



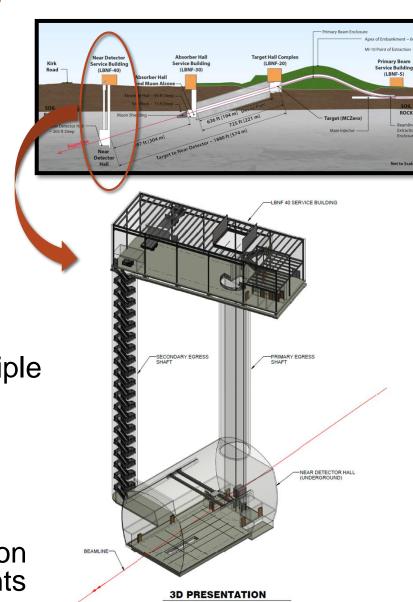
Neutrino Flux at 1300 km (CDR Optimized Beam)



- 60-120 GeV proton beam
- 1.2 MW, upgradeable to 2.4 MW
- Horn-focused neutrino beam line optimized for CP violation sensitivity using genetic algorithm
 - Engineering design of 3-horn focusing system based on optimized parameters in progress
- Neutrino (FHC) and antineutrino (RHC) modes

DUNE Near Detector

- Primary purpose is to constrain systematic uncertainty for longbaseline oscillation analysis
 - Constrain flux, cross-section, and detector uncertainties
- DUNE ND design concept near final
 - Active ND Design Group
 - ND Conceptual Design Report (CDR) planned for 2019
- DUNE ND design concept is an integrated system composed of multiple detectors:
 - Highly segmented LArTPC
 - Magnetized multi-purpose tracker
 - Electromagnetic calorimeter
 - Muon chambers
- Conceptual design will preserve option to move ND for off-axis measurements

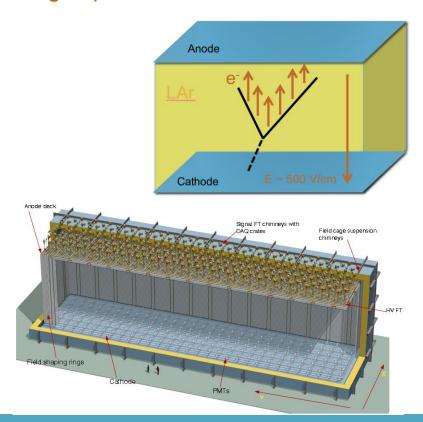


DUNE Far Detector

- 4 10-kt (fiducial) liquid argon TPC modules
- Single- and dual-phase detector designs (1st module will be single phase)
- Integrated photon detection
- Modules will not be identical

Single phase: modular wire-plane readout

Cathode Anode 12 m high 15.5 m wide 58 m long Dual phase: signal extracted; amplified in gas phase



Timeline



2018: protoDUNEs at **CERN**



2019: Technical Design Report



Physics data as soon as 1st module complete **Excavation Begins** Atmospheric vs



2022: First Module **Installation Begins**



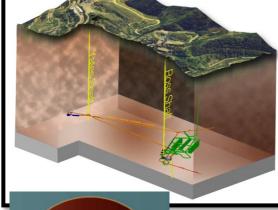
2026: Neutrino Beam Available

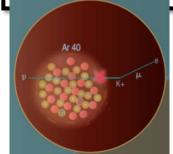
DUNE Far Detector Interim Design Report (2018) Will be made public soon...

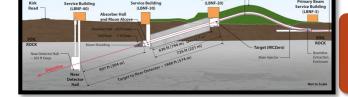












SNB and solar vs

Detector calibration

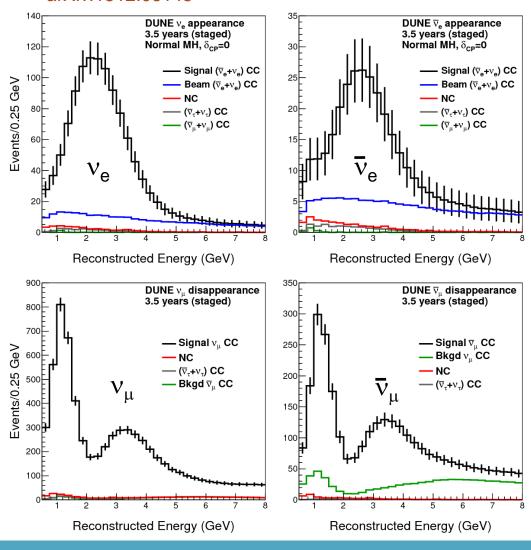
Baryon number

violation

DUNE SCIENCE

Oscillation Sensitivity Calculations

DUNE Conceptual Design Report (CDR) arXiv:1512.06148



- Reconstructed spectra based on GEANT4 beam simulation, GENIE event generator, and Fast MC using detector response parameterized at the single particle level
 - Efficiency tuned using hand scan results
- Order 1000 v_e appearance events in ~7 years of equal running in neutrino and antineutrino mode
- Simultaneous fit to four spectra to extract oscillation parameters

DUNE CDR Systematics

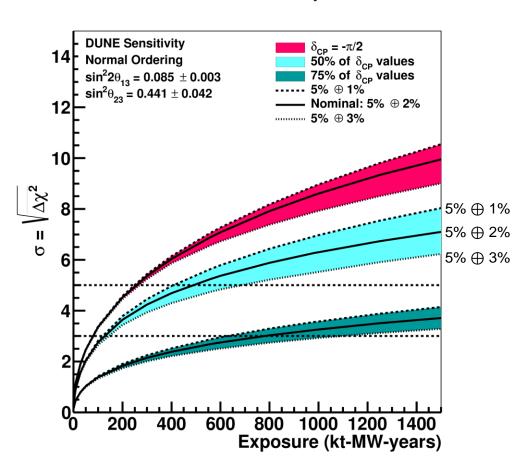
 Sensitivities in DUNE CDR are based on GLoBES calculations in which the effect of systematic uncertainty is approximated using signal and background normalization uncertainties.

Spectral uncertainty not included in this treatment.

- Signal normalization uncertainties are treated as *uncorrelated* among the modes $(v_e, \bar{v}_e, v_\mu, \bar{v}_\mu)$ and represent the **residual uncertainty** expected after constraints from the near detector and the four-sample fit are applied.
 - $v_{\mu} = \overline{v}_{\mu} = 5\%$ Flux uncertainty after ND constraint • $v_{e} = \overline{v}_{e} = 2\%$ Residual uncertainty after v_{μ} and v/\overline{v} constraint
- Oscillation parameter central values and uncertainties are taken from NuFit 2016 (arXiv:1611.01514). Parameters are allowed to vary constrained by 1/6 of the ±3σ range in the global fit.

Effect of Normalization Uncertainty

CP Violation Sensitivity



Statistically limited for ~100 kt-MW-years.

Uncertainty in v_e
appearance sample
normalization must be
~5% ⊕ 2% to discover
CPV in a timely manner.

STRATEGY FOR CONSTRAINING SYSTEMATIC UNCERTAINTIES

- The uncertainty could be categorised into three groups.
 - Flux will be constrained using near detector and external data

Detector response - will be constrained using prototype detectors,
 SBN detectors, test beam experiments, near detector)

Interaction modelling (discussed here)

Strategy for Interaction Model

- Prospects for improved interaction models:
 - Improved models becoming available
 - Intermediate neutrino program measurements in LAr TPCs
- ND constraint:
 - High precision near detector designed to constrain cross-section and hadronization uncertainties, resolving many individual particles produced by resonance and DIS interactions
 - Argon nuclear targets in ND allows significant cancellation of cross section uncertainties common to near and far detectors
- FD constraint:
 - Four FD samples allow cancellation of uncertainties that are correlated between v_e/v_u or v/\overline{v}

Improving Interaction Models

- Worldwide effort that will benefit DUNE!
- Alternative models being implemented in GENIE include:
 - Long- and short-range correlations among nucleons
 - Effect of random phase approximations
 - Meson exchange currents
 - Effective spectral functions
 - Coherent pion production
 - Updated PDFs for DIS interactions
- Variation of tunable parameters within existing models
- Comparisons among generators
- Neutrino interaction data available or coming soon from:
 - ArgoNeuT, MINERvA, NOvA-ND, T2K-ND280, μBooNE, SBND, ICARUS, ...
- Electron-argon scattering data coming from JLab

DUNE Monte Carlo Generator

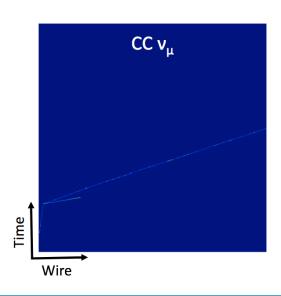
- In the DUNE CDR neutrino interactions were simulated using GENIE 2.8.4.
- The newest Monte Carlo production uses GENIE 2.12.8
 - configuration DefaultPlusMECWithNC (https://cdcvs.fnal.gov/redmine/projects/nutools/wiki/GENIE_Configuration_Files)
 - This is rather "old" configuration with Rein-Sehgal models for pion production interactions and experimental MEC model (by S. Dytman)
- Dune collaboration is now in the process of defining modern GENIE configuration/tune.
 - I will report opinions from this workshop to the working group.
 - The new configuration will be used for the Technical Design Report (2019).

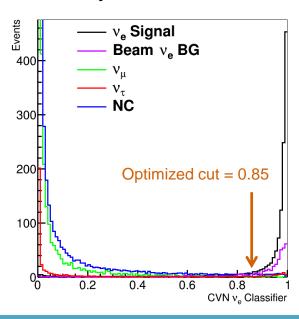
Monte Carlo Analysis (New!)

- GEANT4 beam simulation of updated beam design
- Full LArSoft Monte Carlo simulation
 - Shared framework among many LArTPC experiments
 - GENIE event generator
 - GEANT4 particle propagation
 - Detector readout simulation including realistic waveforms and white noise
- Automated signal processing and hit finding
- Automated energy reconstruction
 - Muon momentum from range (contained) or multiple Coulomb scattering (exiting)
 - Electron and hadron energy from calorimetry
- Event selection using convolutional visual network (CVN)
- Oscillation analysis using CAFAna fitting framework
 - Shared framework with NOvA
- CDR-style systematics analysis (update coming in 2019)
- Results shown here are for single phase; dual phase analysis in progress

CVN Event Selection

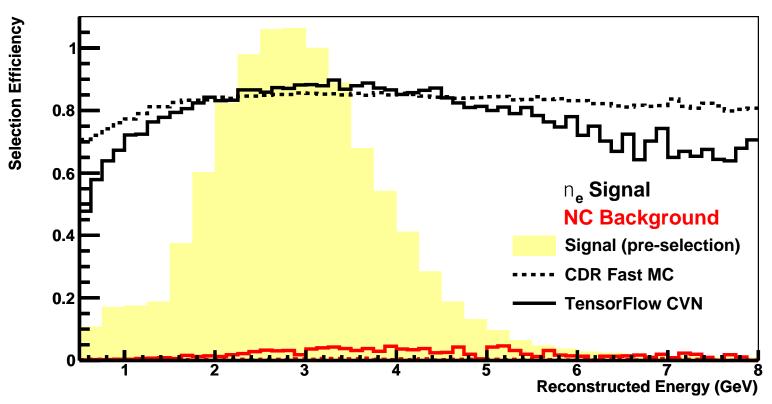
- ResNet architecture implemented in TensorFlow (arXiv:1605.07678)
- Training performed on sets of 500 x 500 DUNE MC images
- DUNE MC images classified into categories
 - v_e CC, v_u CC, v_τ CC, NC
- Event selection performed by applying cuts on ν_e CC-like and ν_μ CC-like CVN classifiers
 - v_eCC-like cut chosen by optimizing CPV sensitivity





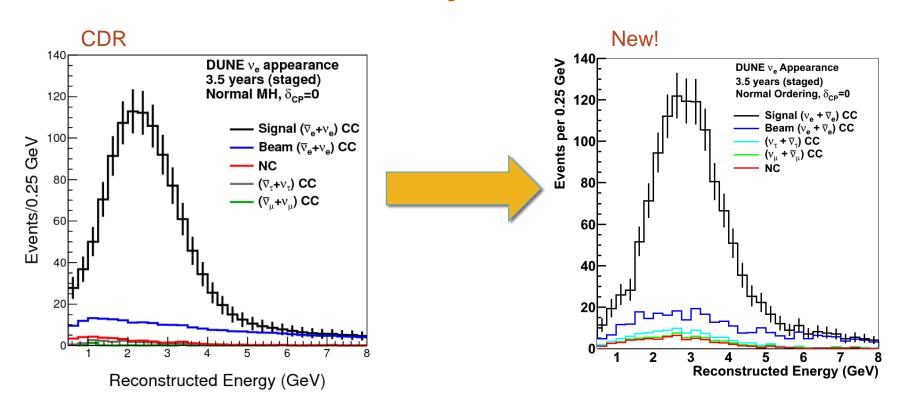
Selection Efficiency





CVN v_e event selection efficiency similar to that from CDR Fast MC

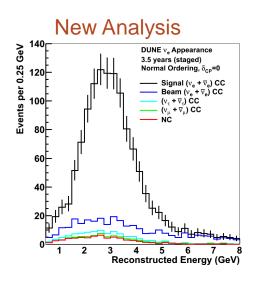
Monte Carlo Analysis Results



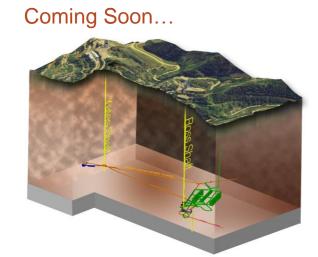
Good agreement between CDR and New MC-based analysis

Full update of sensitivity plots with detailed systematics planned for TDR in 2019

Summary



protoDUNE



- LBNF and DUNE making rapid progress on facility construction, detector design, and physics analysis
- New MC-based analysis is ready for sensitivity studies.
- The neutrino interactions modeling in DUNE MC will be updated for the Technical Design Report in 2019!
- Expect first DUNE FD data in ~2024...

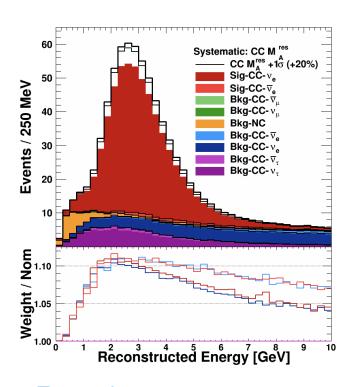
Extra Slides

Fast MC/MGT Details

- Fast MC = Flux simulations + GENIE + parameterized detector response
- Detector response parameterization based on inputs from LArSoft simulations, GEANT4, and ICARUS
- Reconstructed quantities and selection criteria based on realistic kinematics
- MGT fitter includes statistical limitations on constraints from four-sample fit and uncertainty in sample-sample correlation

Single particle detector response. Many values quite conservative.

Particle type	Detection Threshold (KE)	Energy/Momentum Resolution	Angular Resolution
μ^\pm	30 MeV	Contained track: track length Exiting track: 30%	1°
π^\pm	100 MeV	μ -like contained track: track length π -like contained track: 5% Showering or exiting: 30%	1°
${\sf e}^{\pm}/\gamma$	30 MeV	$2\% \oplus 15\%/\sqrt{E}$ [GeV]	1°
р	50 MeV	p<400 MeV/c: 10% p>400 MeV/c: 5% \oplus 30%/ \sqrt{E} [GeV]	5°
n	50 MeV	$40\%/\sqrt{E}[{\sf GeV}]$	5°
other	50 MeV	$5\% \oplus 30\%/\sqrt{E}$ [GeV]	5°



Example: v_e appearance spectrum showing variation that is induced by changing the value of M_A^{RES} by +1 σ in the simulation.

DUNE CDR

Normalization uncertainties

- Estimate uncertainties after ND and external data constraints
- Understand advantages of LArTPC, and cancellations in FD 4-sample fits
- Consider experience from T2K and MINOS
 - MINOS similarities
 - Flux shape, v energies
 - Longer baseline
 - Similar cross sections
 - T2K similarities
 - Different near and far detector technologies
 - Similar analysis strategies
- Strategies to address required increase in precision

Source of	MINOS	T2K	DUNE
Uncertainty	$ u_e$	$ u_e$	$ u_e$
Beam Flux	0.3%	3.2%	2%
after N/F			
extrapolation			
Interaction	2.7%	5.3%	$\sim 2\%$
Model			
Energy scale	3.5%	included	(2%)
(u_{μ})		above	
Energy scale	2.7%	2.5%	2%
(u_e)		includes	
		all FD	
		effects	
Fiducial	2.4%	1%	1%
volume			
Total	5.7%	6.8%	3.6 %
Used in DUNE			$5\% \oplus 2\%$
Sensitivity			
Calculations			

Background Uncertainties

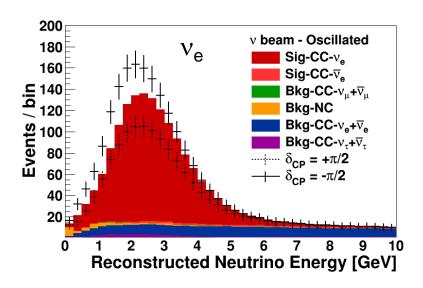
Table 3.9: Normalization uncertainties and correlations for background to the ν_e , $\bar{\nu}_e$, ν_{μ} , and $\bar{\nu}_{\mu}$ data samples

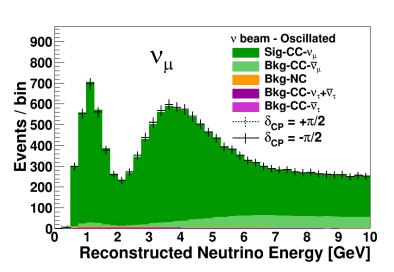
Background	Normalization Uncertainty	Correlations		
For $\nu_e/\bar{\nu}_e$ appe	For $ u_e/\bar{ u}_e$ appearance:			
Beam ν_e	5%	Uncorrelated in $ u_e$ and $ar u_e$ samples		
NC	5%	Correlated in $ u_e$ and $ar u_e$ samples		
ν_{μ} CC	5%	Correlated to NC		
$\nu_{ au}$ CC	20%	Correlated in $ u_e$ and $ar u_e$ samples		
For $ u_{\mu}/\bar{\nu}_{\mu}$ disappearance:				
NC	5%	Uncorrelated to $ u_e/ar u_e$ NC background		
$\overline{ u_{ au}}$	20%	Correlated to $ u_e/ar{ u}_e$ $ u_ au$ background		

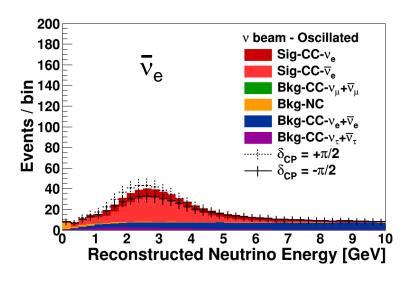
The DUNE Experimental Setup

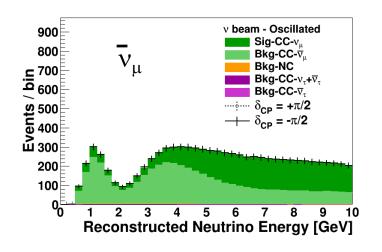
	CDR Reference Design	Optimized Design	
ν mode (150 kt · MW · year)			
$ u_e$ Signal NH (IH)	861 (495)	945 (521)	
$ar{ u}_e$ Signal NH (IH)	13 (26)	10 (22)	
Total Signal NH (IH)	874 (521)	955 (543)	
Beam $ u_e + \bar{\nu}_e$ CC Bkgd	159	204	neutrinos
NC Bkgd	22	17	
$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	42	19	
$ u_{\mu} + ar{ u}_{\mu} \; CC \; Bkgd$	3	3	
Total Bkgd	226	243	
$\bar{\nu}$ mode (150 kt · MW · year)			
ν_e Signal NH (IH)	61 (37)	47 (28)	
$ar{ u}_e$ Signal NH (IH)	167 (378)	168 (436)	
Total Signal NH (IH)	228 (415)	215 (464)	
Beam $ u_e + ar{ u}_e$ CC Bkgd	89	105	antineutrinos
NC Bkgd	12	9	
$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	23	11	
$ u_{\mu} + ar{ u}_{\mu} \; CC \; Bkgd$	2	2	
Total Bkgd	126	127	

Expected FD Spectra

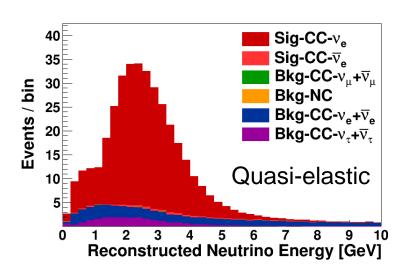


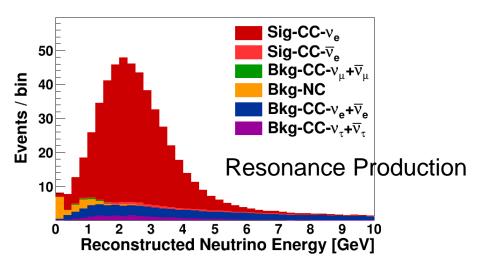


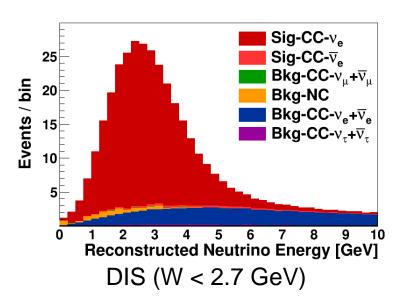


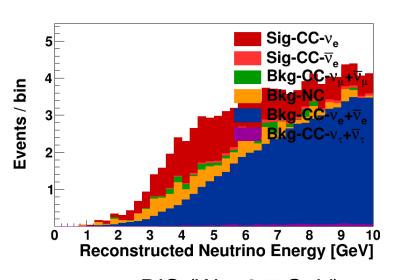


Spectra By Cross Section Model







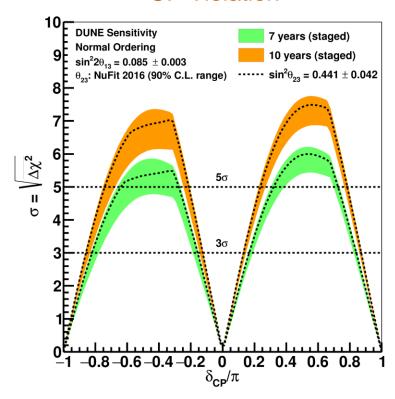


DIS (W > 2.7 GeV)

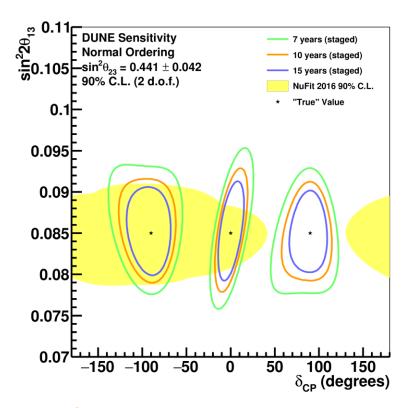
CP Violation Sensitivity

DUNE CDR:

CP Violation



Width of band indicates variation in possible central values of θ_{23}

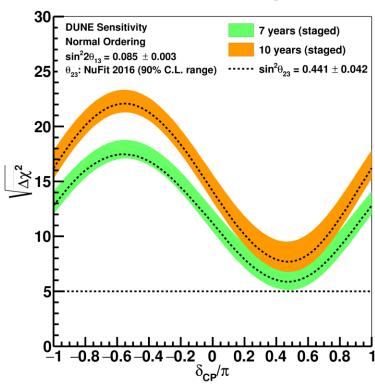


Simultaneous measurement of neutrino mixing angles and δ_{CP}

Other Oscillation Physics

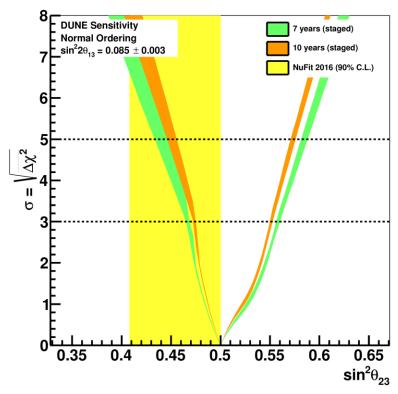
DUNE CDR:

Mass Ordering



Width of band indicates variation in possible central values of θ_{23}

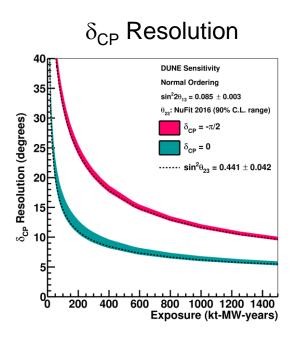
Octant

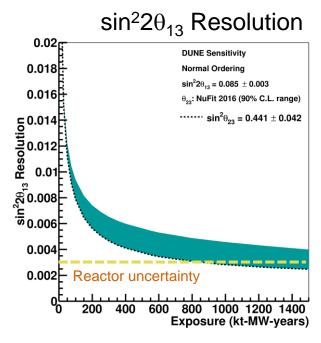


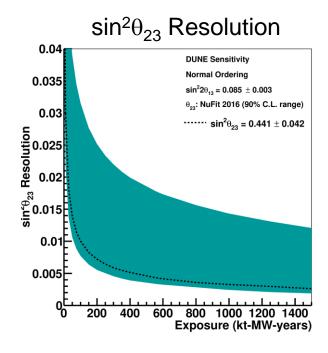
Width of band indicates variation in possible true value of δ_{CP}

Oscillation Parameter Sensitivity

DUNE CDR:







Staging Assumptions

- Year 1 (2026): 20-kt FD with 1.07 MW (80-GeV) beam and initial ND constraints
- Year 2 (2027): 30-kt FD
- Year 4 (2029): 40-kt FD and improved ND constraints
- Year 7 (2032): upgrade to 2.14 MW (80-GeV) beam (technically limited schedule)

Exposure (kt-MW-years)	Exposure (Years)
171	5
300	7
556	10
984	15

NuFit 2016

NuFIT 3.0 (2016)

	Normal Ord	dering (best fit)	Inverted Orde	$ring (\Delta \chi^2 = 0.83)$	Any Ordering
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	3σ range
$\sin^2 heta_{12}$	$0.306^{+0.012}_{-0.012}$	$0.271 \to 0.345$	$0.306^{+0.012}_{-0.012}$	$0.271 \rightarrow 0.345$	$0.271 \to 0.345$
$ heta_{12}/^\circ$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$31.38 \rightarrow 35.99$
$\sin^2 \theta_{23}$	$0.441^{+0.027}_{-0.021}$	$0.385 \to 0.635$	$0.587^{+0.020}_{-0.024}$	$0.393 \to 0.640$	$0.385 \to 0.638$
$ heta_{23}/^\circ$	$41.6_{-1.2}^{+1.5}$	$38.4 \rightarrow 52.8$	$50.0^{+1.1}_{-1.4}$	$38.8 \rightarrow 53.1$	$38.4 \rightarrow 53.0$
$\sin^2 \theta_{13}$	$0.02166^{+0.00075}_{-0.00075}$	$0.01934 \rightarrow 0.02392$	$0.02179^{+0.00076}_{-0.00076}$	$0.01953 \rightarrow 0.02408$	$0.01934 \rightarrow 0.02397$
$ heta_{13}/^\circ$	$8.46^{+0.15}_{-0.15}$	$7.99 \rightarrow 8.90$	$8.49^{+0.15}_{-0.15}$	$8.03 \rightarrow 8.93$	$7.99 \rightarrow 8.91$
$\delta_{ m CP}/^\circ$	261^{+51}_{-59}	$0 \rightarrow 360$	277_{-46}^{+40}	$145 \rightarrow 391$	$0 \rightarrow 360$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.50_{-0.17}^{+0.19}$	$7.03 \rightarrow 8.09$	$7.50_{-0.17}^{+0.19}$	$7.03 \rightarrow 8.09$	$7.03 \rightarrow 8.09$
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.524^{+0.039}_{-0.040}$	$+2.407 \rightarrow +2.643$	$-2.514_{-0.041}^{+0.038}$	$-2.635 \to -2.399$	$ \begin{bmatrix} +2.407 \to +2.643 \\ -2.629 \to -2.405 \end{bmatrix} $

For 1σ uncertainty in DUNE sensitivity calculations, we take 1/6 of the $\pm 3\sigma$ range, to account for non-Gaussian PDFs in NuFit.

DUNE Systematics: TDR

- Systematics analysis building on expertise developed in MINERvA, T2K, and NOvA
 - "DUNEResponse" ← "T2KReweight"
- CAFAna fitting framework facilitates more sophisticated treatment of systematic uncertainty than was possible for CDR
 - Systematic uncertainties in TDR will be based on detailed evaluation of flux, neutrino interaction, and detector uncertainties
 - Sensitivity calculations will be based on fits combining information from near and far detectors
- Flux and interaction systematics evaluated using reweighting technique (including GENIE and non-GENIE reweights)
 - Impact of systematic variations propagated through full analysis chain
 - Ability to consider systematics impacting kinematic distributions as well as normalization
- Detector systematics evaluated within the fit
 - Detector calibration task force evaluating magnitude and sources of detector uncertainty