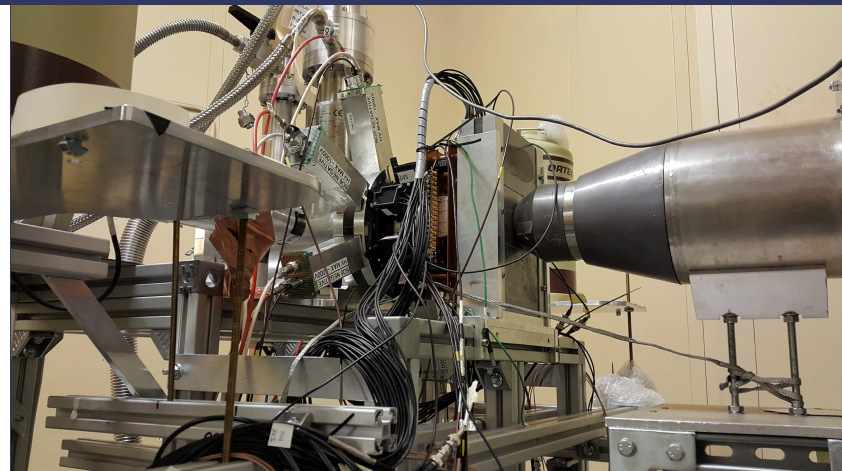
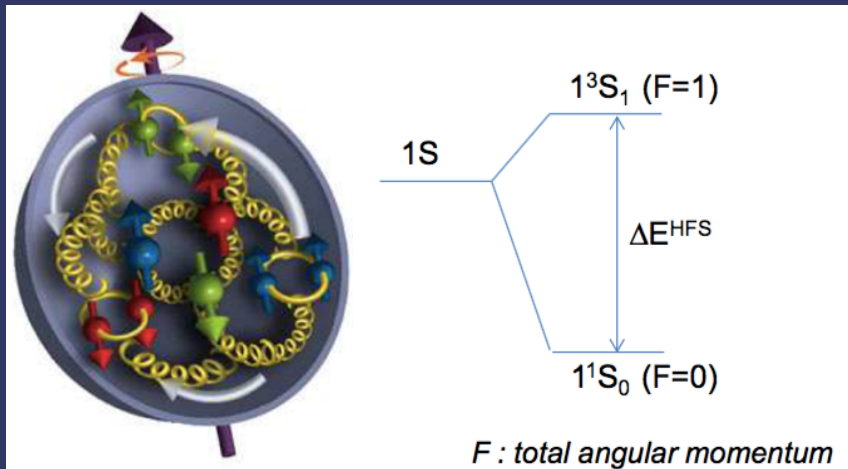


# High precision spectroscopy in muonic hydrogen

## Towards the measurement of the hyperfine transition in muonic hydrogen $\Delta E_{\text{HFS}}(\mu p)_{1S}$

*Andrea Vacchi for the FAMU Collaboration*



Nucleon Spin Structure at Low Q: A Hyperfine View

[ECT\\* - European Centre for Theoretical Studies in Nuclear Physics](#)

July 2-6, 2018

[Trento, Italy](#)

# FAMU Collaboration



**INFN Trieste:** V. Bonvicini, H. Cabrera, E. Furlanetto, E. Mocchiutti, C. Pizzolotto, A. Rachevsky, L. Stoychev, A. Vacchi (also *Università di Udine*), E. Vallazza, G. Zampa, **Elettra-Sincrotrone:** M. Danailov, A. Demidovich, **ICTP:** J. Niemela, K.S. Gadedjisso-Tossou

**INFN Bologna:** L. Andreani, G. Baldazzi, G. Campana, I. D'Antone, M. Furini, F. Fuschino, A. Gabrielli, C. Labanti, A. Margotti, M. Marisaldi, S. Meneghini, G. Morgante, L. P. Rignanese, P. L. Rossi, M. Zuffa

**INFN Milano Bicocca:** A. Baccolo, R. Benocci, R. Bertoni, M. Bonesini, T. Cervi, F. Chignoli, M. Clemenza, A. Curioni, V. Maggi, R. Mazza, M. Moretti, M. Nastasi, E. Previtali, R. Ramponi (also Politecnico Milano CNR)

**INFN Pavia:** A. De Bari, C. De Vecchi, A. Menegolli, M. Rossella, R. Nardò, A. Tomaselli

**INFN Roma3:** L. Colace, M. De Vincenzi, A. Iaciofano, L. Tortora, F. Somma

**INFN Seconda Università di Napoli:** L. Gianfrani, L. Moretti

**INAF-IASF Bologna:** V. Fioretti

**INFN – GSSI:** D. Guffanti,

**RIKEN-RAL:** K. Ishida

**INP, Polish Academy of Sciences:** A. Adamczak

**INRNE, Bulgarian Academy of Sciences:** D. Bakalov, M. Stoilov, P. Danev



# OUTLINE

- Background & motivations
- FAMU's method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

# Muonic hydrogen

Muon ( $e^-$ 's heavier twin) orbiting the proton instead of electron.

$$m_\mu = 207m_e$$

$$r_\mu = \frac{1}{186}r_e$$

0.511 MeV	105.7 MeV
$-1$	$-1$
$\frac{1}{2}$	$\frac{1}{2}$
e	$\mu$
electron	muon

$$m_\mu/m_e \approx 2 \times 10^2$$

- the *radius of the muon orbit* is  $\sim a_0/200$  so that the energy levels of muonic hydrogen are orders of magnitude more “sensitive” to the details of the proton structure than the levels of normal hydrogen.
- the binding energy of the ground state of muonic hydrogen is of the order of 200 Ry,

## *muonic hydrogen*

The muon is tightly bound in hydrogen-like orbits that have very large overlaps with the proton:

- Provide *very high accuracy tests of quantum electrodynamics and the theory of electromagnetic bound states.*
- Moreover, *the values of the fundamental physical constants (particle masses, fine structure constant, proton charge radius, etc.) can be determined more precisely.*  
and how universal is (lepton) universality?

- Very precise spectroscopic measurements of hyperfine splitting offer a way for testing quantum electrodynamics (QED).
- Finding discrepancies between QED and experimental observations
- Could point towards physics beyond the Standard Model of particle physics.
- Verify the theoretical predictions of the nature of quantum mechanics in very strong fields

# why measuring $\Delta E^{\text{hfs}}(\mu^-p)_{1S}$ ?

New independent high precision measurements on  $\mu^-p$  are needed.

The spectroscopic measurement of  $\Delta E^{\text{hfs}}(\mu^-p)_{1S}$ , will :

- provide  $r_Z$ , the **Zemach radius of the proton**, with high precision to disentangle among discordant theoretical values
- **quantify any level of discrepancy between values of  $r_Z$  as extracted from normal and muonic hydrogen atoms** leading to new information on proton structure and muon-nucleon interaction.

The experimental value of  $r_Z$  sets important restrictions on the theoretical models of proton electromagnetic structure and on the parametrization of proton form factors, in whose terms the theoretical values are calculated.

# current status of $(\mu^-p)_{1S} \text{ hfs}$

units fm	rms charge radius $r_{\text{ch}}$	Zemach radius $r_Z$
$e^-p$ scattering & spectroscopy	$r_{\text{ch}} = 0.8751(61)$	$r_Z = 1.037(16)$ Dupays&al' 03 $r_Z = 1.086(12)$ s Friar&Sick' 04 $r_Z = 1.047(16)$ Volotka&al' 05 $r_Z = 1.045(4)$ s Distler&al' 11
$\mu^-p$ Lamb shift spectroscopy	$r_{\text{ch}} = 0.84087(39)$	a 20 years old idea: $r_Z$ from HFS of $(\mu^-p)_{1S}$ Either confirm a $e^-p$ value or admit: $e^-p$ and $\mu^-p$ differ

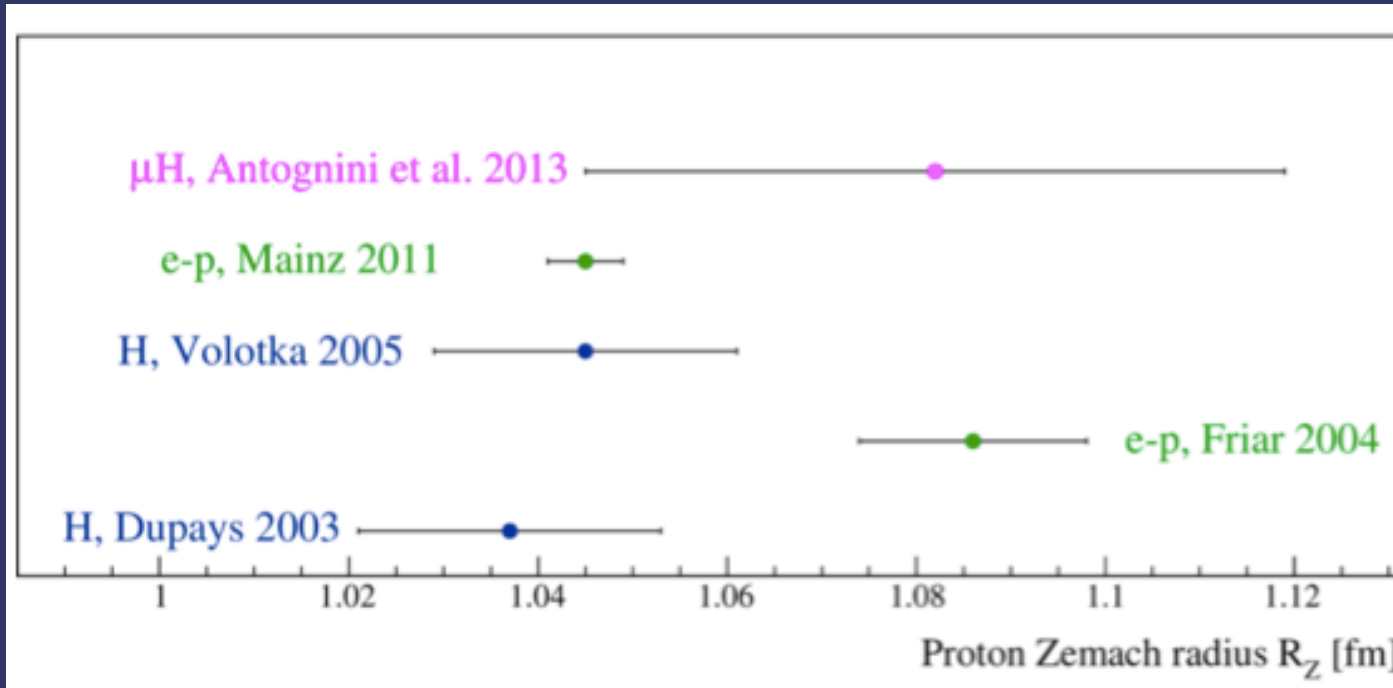
Recently : from hfs of  $(\mu^-p)_{2S}$

$r_Z = 1.082(37)$  [PSI'12]

=> we need new independent measurements

# $r_Z$ current status

## large errors! we need new measurement



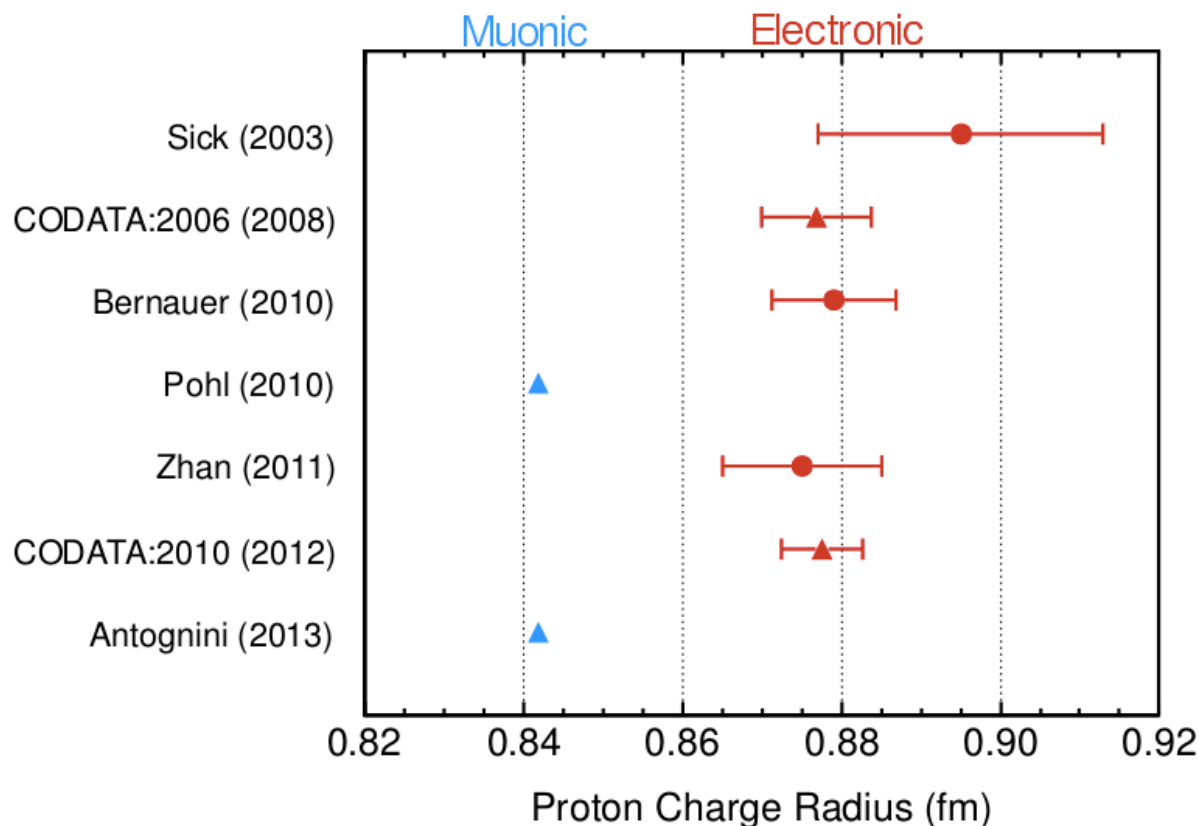
The current theoretical uncertainty of  $r_Z$  significantly exceeds the experimental one.

The experimental results on the proton Zemach radius may be used as a test for the quality of models of the proton in the limit of low transfer momenta.



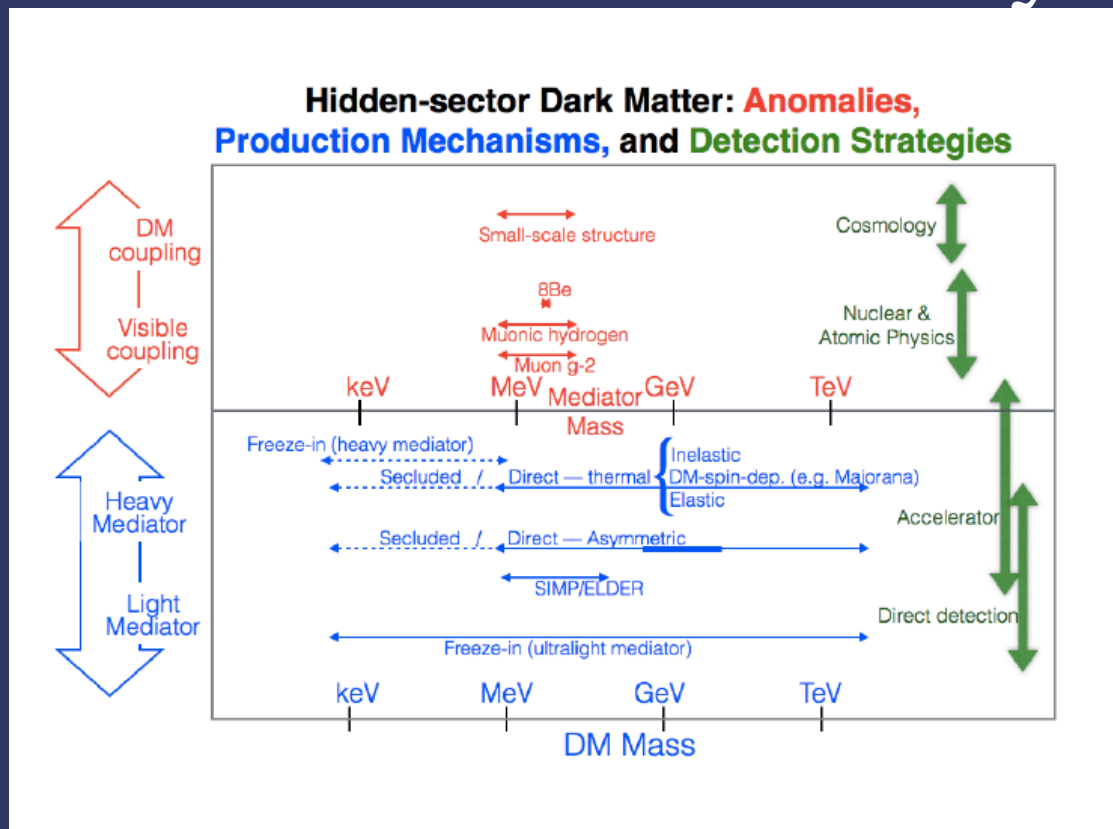


# The proton charge radius can be extracted for each lepton probe from **two** independent methods



The CODATA value of the proton charge radius as obtained from a combination of 24 transition frequency measurements in H and deuterium and several results from elastic electron scattering is **0.88 fm**. However, the **muonic hydrogen Lamb Shift** measurements yield a radius of **0.84 fm**.

# Experimental Anomalies and Hints muon-related anomaly



New models, astrophysical observations, and existing experimental anomalies point to the 1 to 100 MeV mass scale as a high value target region for dark matter and dark mediator searches.

# The FAMU experiment goals

Currently 3 independent experiments plan to measure RZ

- Measure the Hyperfine Splitting (HFS) of  $\mu^-p$  with accuracy  $10^{-5}$
- Extract the Zemach radius of the proton with an accuracy of better than 1%

# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

# a 25 years old idea

Physics Letters A 172 (1993) 277–280  
North-Holland

PHYSICS LETTERS A

## Experimental method to measure the hyperfine splitting of muonic hydrogen $(\mu^-p)_{1S}$

D. Bakalov <sup>1</sup>, E. Milotti, C. Rizzo, A. Vacchi and E. Zavattini

*Dipartimento di Fisica dell'Università di Trieste, via Valerio 2, Trieste 34017, Italy  
and Sezione INFN di Trieste, Area di Ricerca, Padriciano 99, Trieste 34012, Italy*

Received 31 July 1992; revised manuscript received 17 October 1992; accepted for publication 8 November 1992  
Communicated by B. Fricke

We propose an experimental method to measure the hyperfine splitting of the energy level of the muonic hydrogen ground state  $(\mu^-p)_{1S}$  by inducing a laser-stimulated para-to-ortho transition. The method requires an intense low energy pulsed  $\mu^-$  beam and a high power tunable pulsed laser.

### 1. Introduction

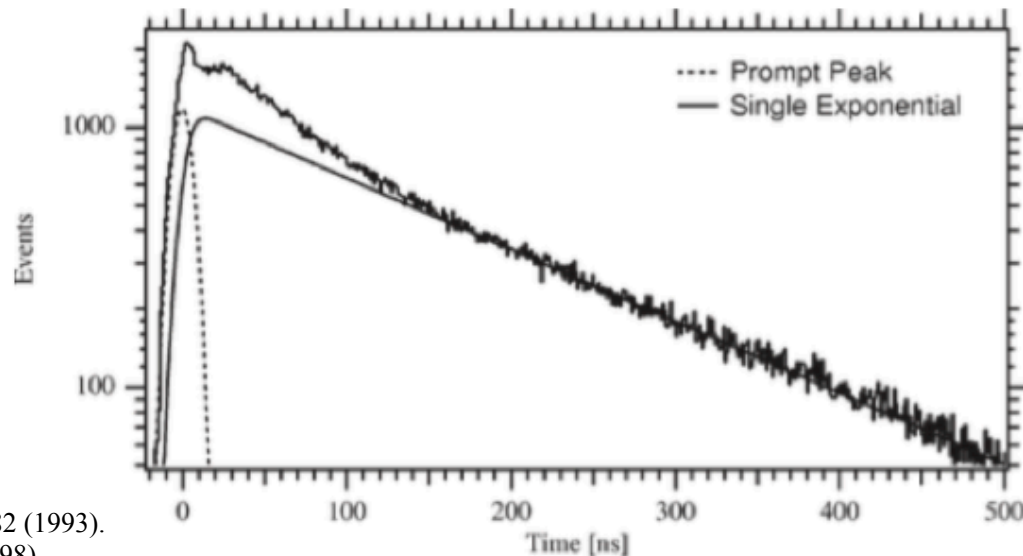
The theoretical expression for the hyperfine splitting

Exploits the *energy dependence of the muon transfer* from muonic hydrogen to higher-Z gas is to detect the spin flip transition in  $\mu p$ .

- For few gases the muon-transfer rate  $\lambda_{pZ}$  is energy dependent  
*Oxygen exhibits a peak in the muon transfer rate  $\lambda_{pZ}^{epith}$  at epithermal energy.*

A. Werthmüller et al. / Muon transfer to oxygen

3

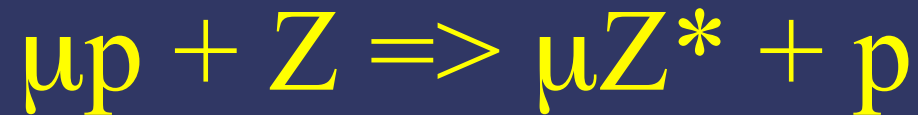


F. Mulhauser, H. Schneuwly, *Hyperfine Interact.* 82 (1993).  
A. Werthmüller, et al., *Hyperfine Interact.* 116 (1998).

Figure 2. Background subtracted time distribution of muonic oxygen  $\mu O(2-1)$  X-rays measured in a gaseous mixture of  $H_2 + 0.4\%O_2$  at 15 bar and room temperature. The prompt peak corresponds essentially to muons directly captured in oxygen whereas the delayed part is due to muon transfer from the ground state of the  $(\mu p)_{1s}$  atom. The solid line represents a pure exponential function to stress the additional structure.

Exploits the *energy dependence of the muon transfer* from muonic hydrogen to higher-Z gas is to detect the spin flip transition in  $\mu p$ .

Adding small quantities of oxygen to hydrogen one can observe the number of hpf transitions which take place from the muon-transfer events this by measuring the time distribution of the oxygen characteristic X-rays of the added gas.



A. Werthmüller et al. / Muon transfer to oxygen

3

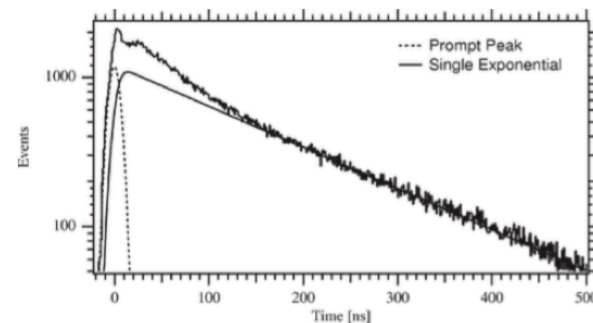


Figure 2. Background subtracted time distribution of muonic oxygen  $\mu O(2-1)$  X-rays measured in a gaseous mixture of  $H_2 + 0.4\%O_2$  at 15 bar and room temperature. The prompt peak corresponds essentially to muons directly captured in oxygen whereas the delayed part is due to muon transfer from the ground state of the  $(\mu p)_{1s}$  atom. The solid line represents a pure exponential function to stress the additional structure.

D. Bakalov, A. Adamczak et al., Phys. Lett. A379 (2014).  
A. Adamczak et al. Hyperfine Interactions 136: 1–7, 2001.

# Laser spectroscopy for $\Delta E_{1S}^{\text{HFS}}$

HOW ?

Method relying on a two-steps process

excited  $\mu p^*$  with  $n > 14$

are formed in a hydrogen gas target,  
in subsequent collisions with H<sub>2</sub> molecules,  
the  $\mu p$  de-excite to the



thermalized  $\mu p$  in the (1S)  $F = 0$  state.



first step

tunable laser pulse

Converts the spin state of the ( $\bar{\mu}$  p) atoms  
from  $^1S_0$  to  $^3S_1$

$\bar{\mu}$ -p( $\uparrow\downarrow$ ) absorbs a photon @ *resonance wavelength*

$$\lambda_0 = hc/\Delta E_{\text{HFS}}^{1S} \sim 6.8 \mu \sim 0.183 \text{ eV}$$



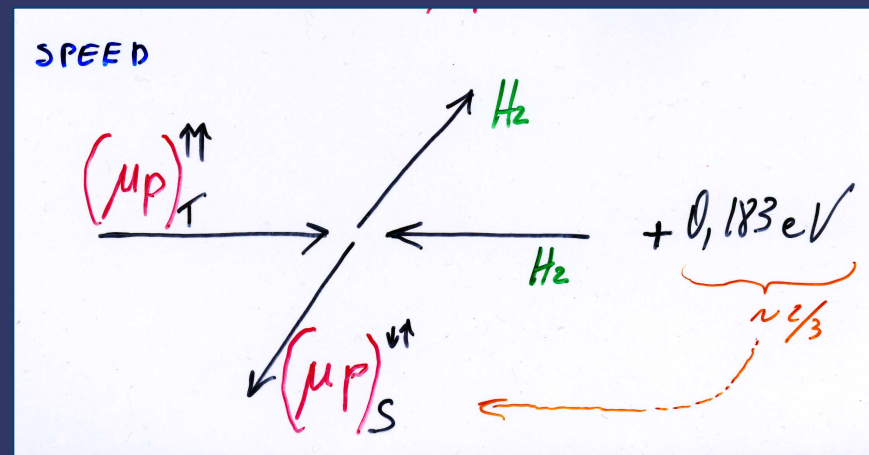
# second-step energy dependent $\mu$ transfer

$\mu^-p(\uparrow\uparrow)$   $^3S_1$  atoms are collisionally de-excited and the transition energy is converted into additional kinetic energy of the  $\mu p$  system

$\mu^-p(\uparrow\downarrow)$   $^1S_0$  and accelerated by  $\sim 0.12 \text{ eV} \sim 2/3 \Delta E_{1S}^{\text{HFS}}$

*Energy-dependent muon transfer rates* change the time distribution of the cascade X-ray events from  $\mu^-Z^{**}$

$\lambda_0$  is recognized by maximal response in the time distribution



D. Bakalov, et al., Phys. Lett. A172 (1993).

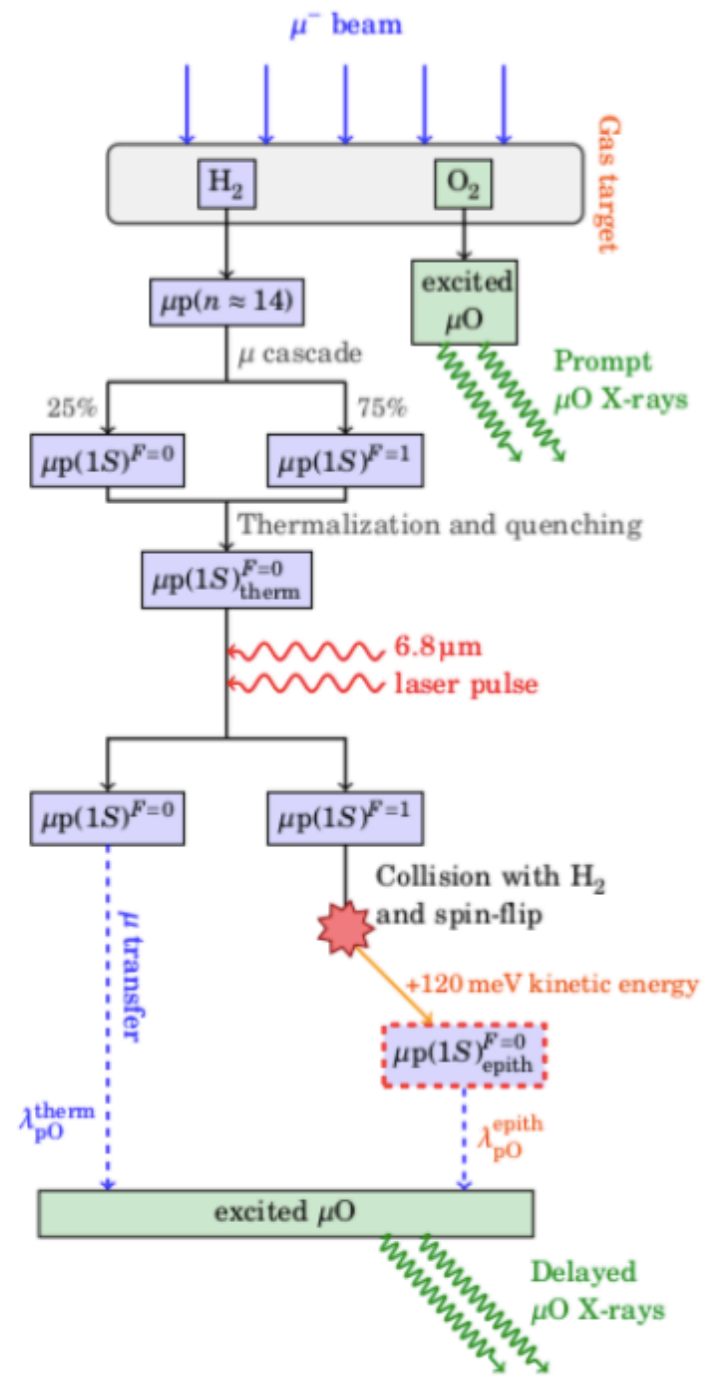
A. Dupays, Phys. Rev. A 68, p. 052503, 2003.

D. Bakalov, et al., NIM B281 (2012).

# FAMU Principle of operation

$\mu p$  formation  $\Rightarrow \Rightarrow$

$\mu p$  thermalization  $\Rightarrow \Rightarrow$

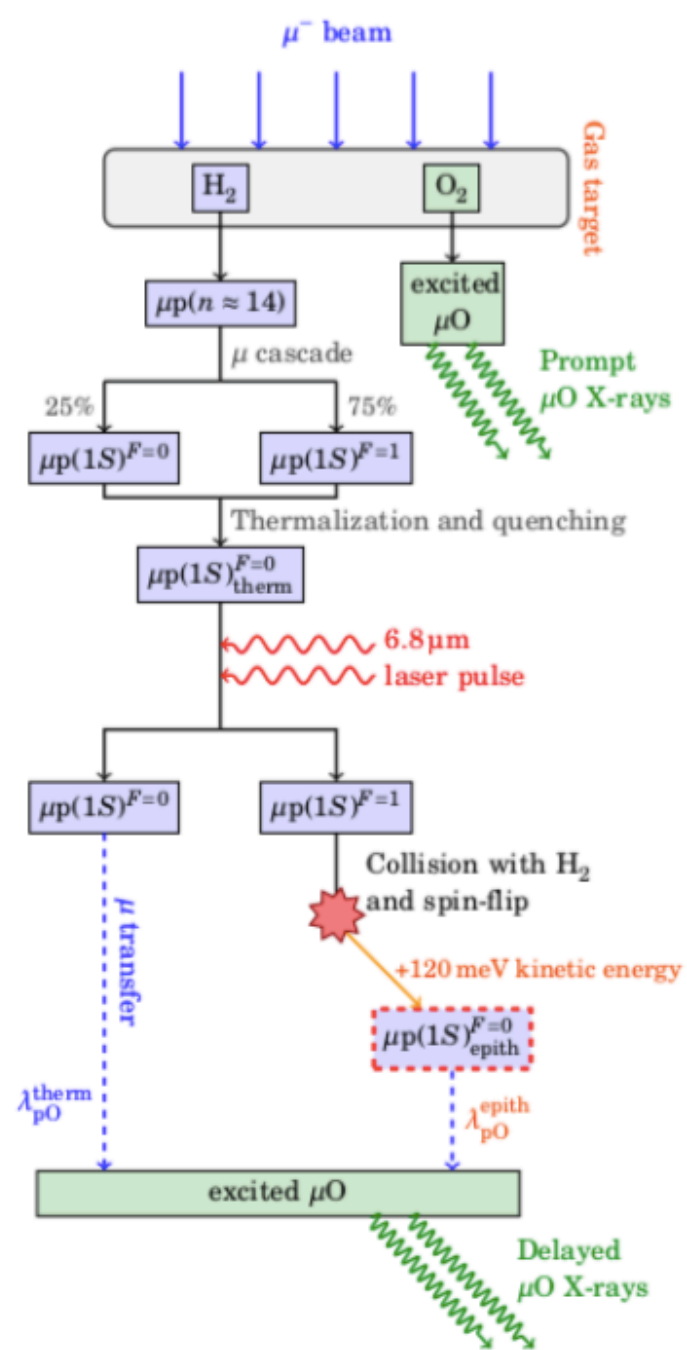


# Lay-Out of the experiment

laser excitation  $F=0 \rightarrow F=1$

$\mu$  enhanced transfer to O

$\mu$ O X-ray time distribution

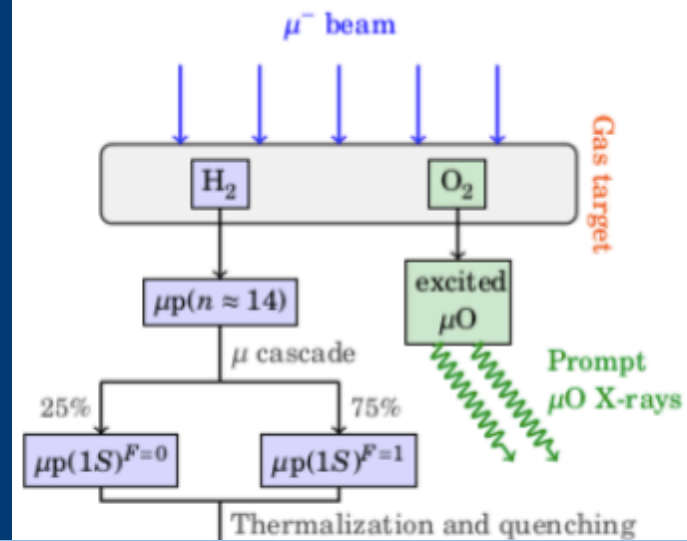


$\Rightarrow \Rightarrow$

$\Rightarrow \Rightarrow$

$\Rightarrow$

# Lay-Out of the experiment

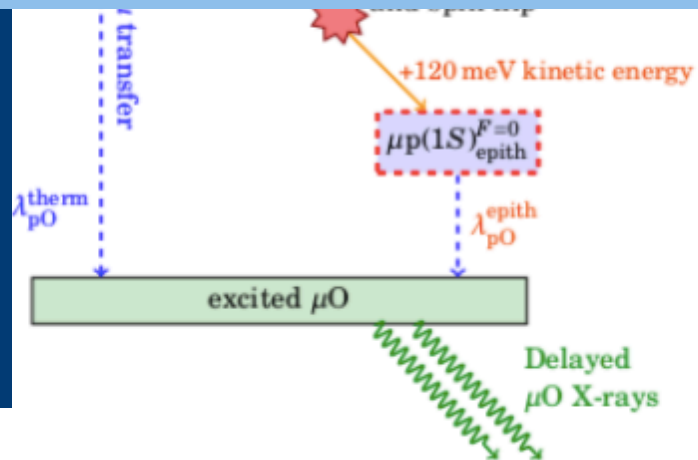


laser excitation

$\lambda_0$  resonance is determined by maximizing the time distribution of  $\mu^-$  transferred events.

$\mu^-$  enhanced transfer to O  $\Rightarrow\Rightarrow$

$\mu\text{O}$  X-ray time distribution  $\Rightarrow$



# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

# FAMU's activity summary

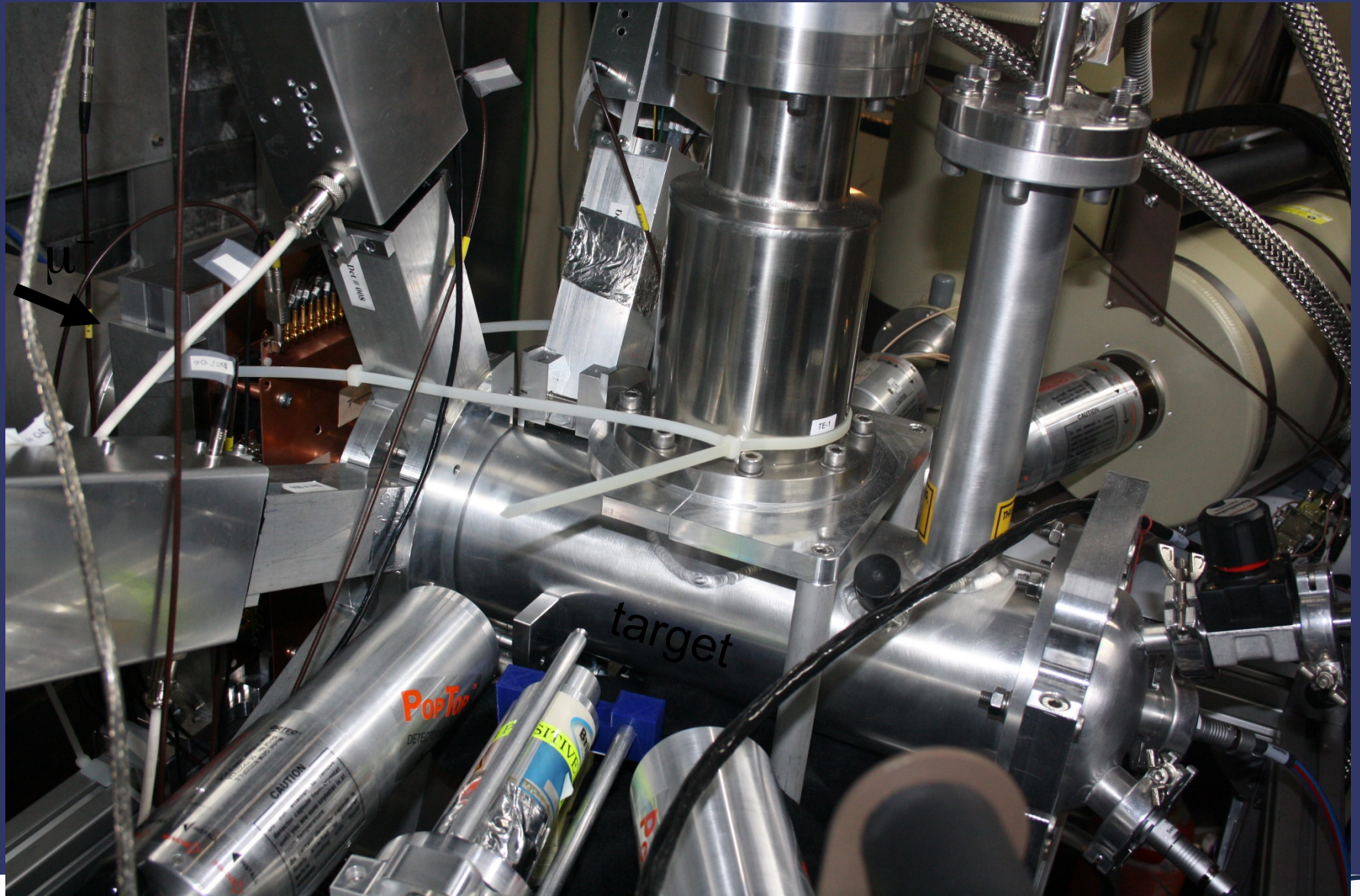
- 2014 characterisation of beam and detector's noise first measurements of transfer rate at room temperature
- 2015-6
  - cryogenic target first measurement of transfer rate between 100 and 300 K
  - laser parts procurement initiated
- 2017-18
  - laser parts delivery completed assembly and characterization on going
  - based on the results of the transfer rate measurements at different temperatures new studies are on going
    - optimal optic cavity design
    - new cryogenic gas target design study and simulation
    - muon beam optimization

# OUTLINE

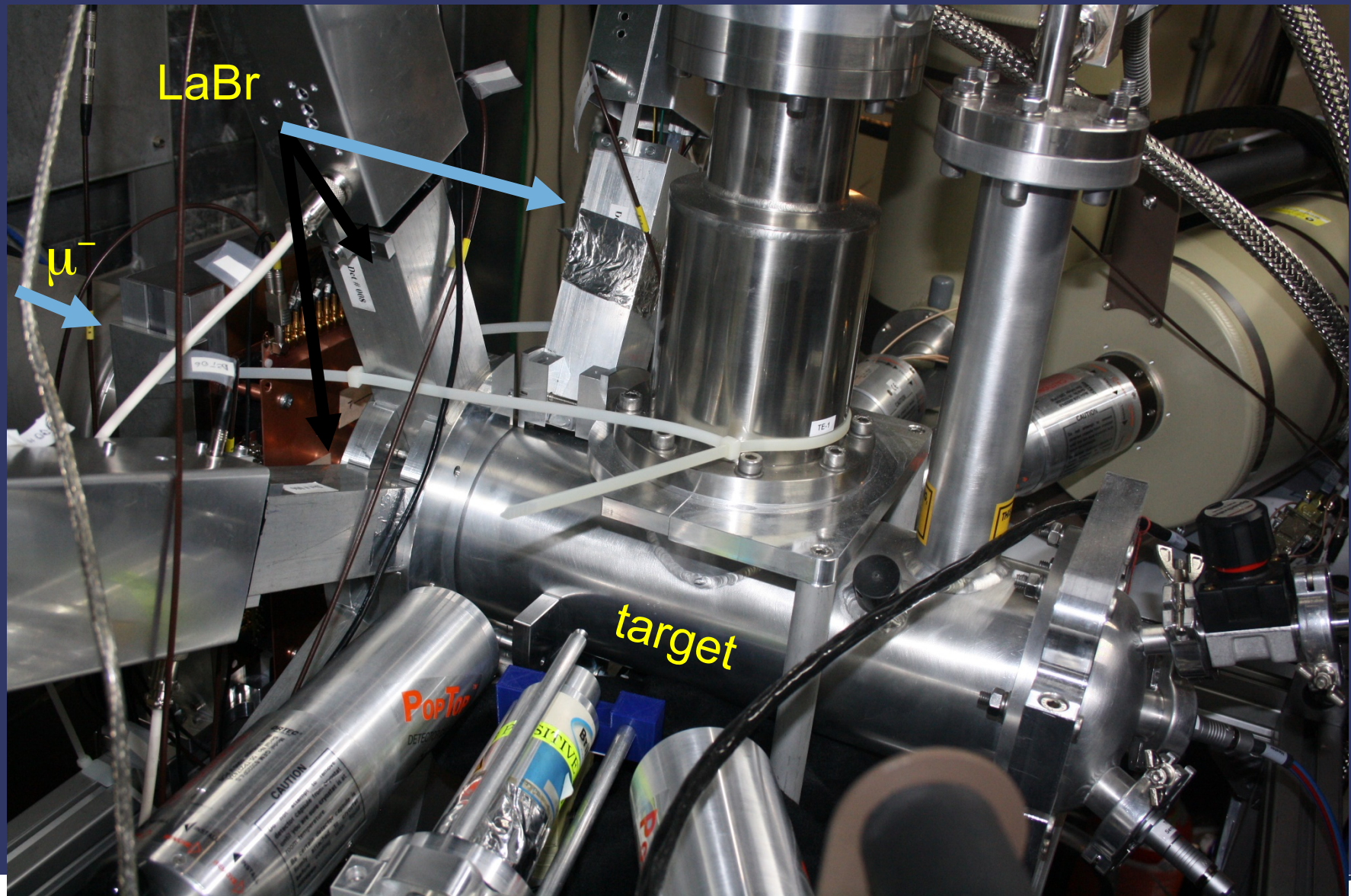
- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions



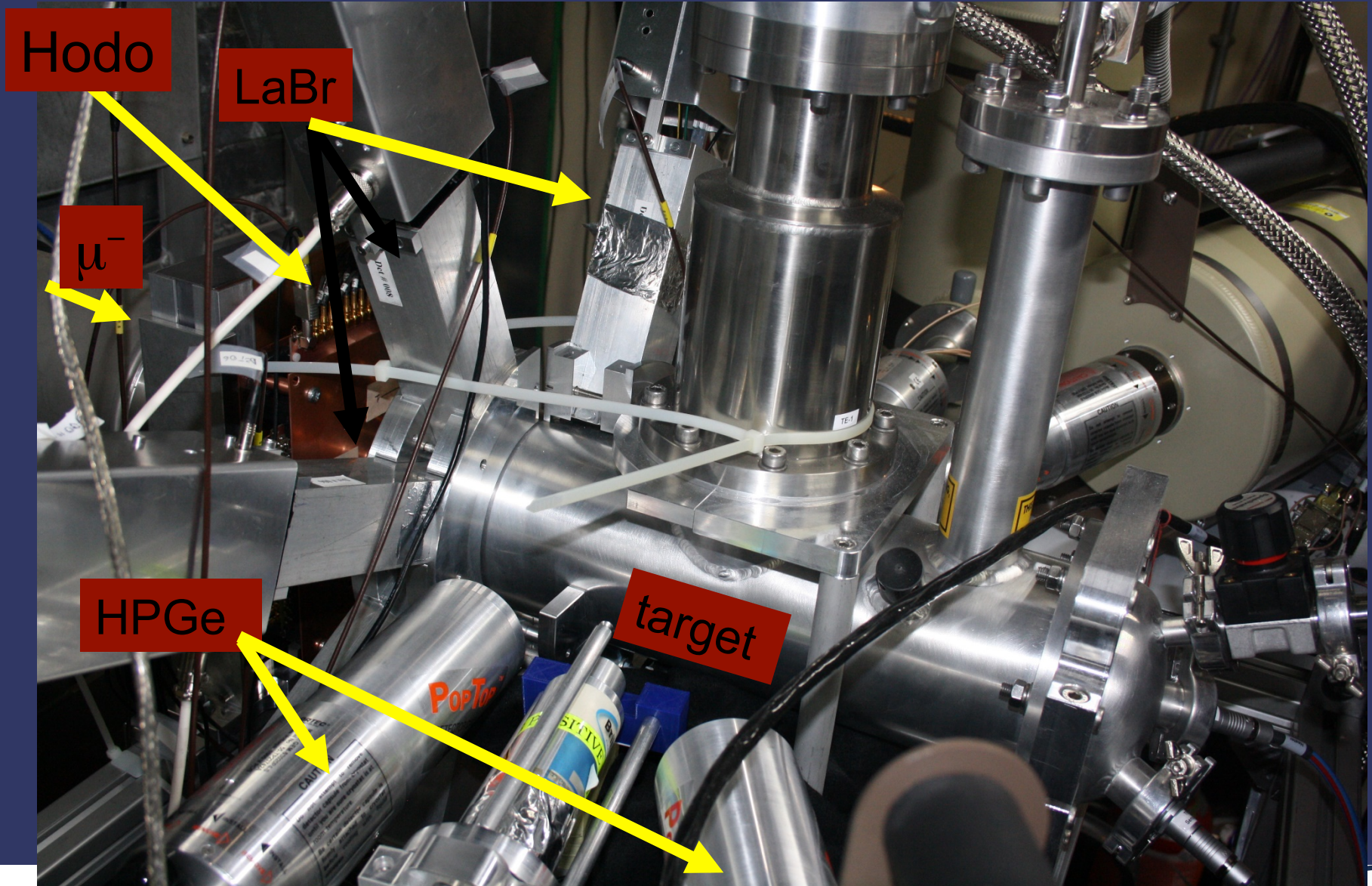
# 2016: experimental setup



# 2016: experimental setup



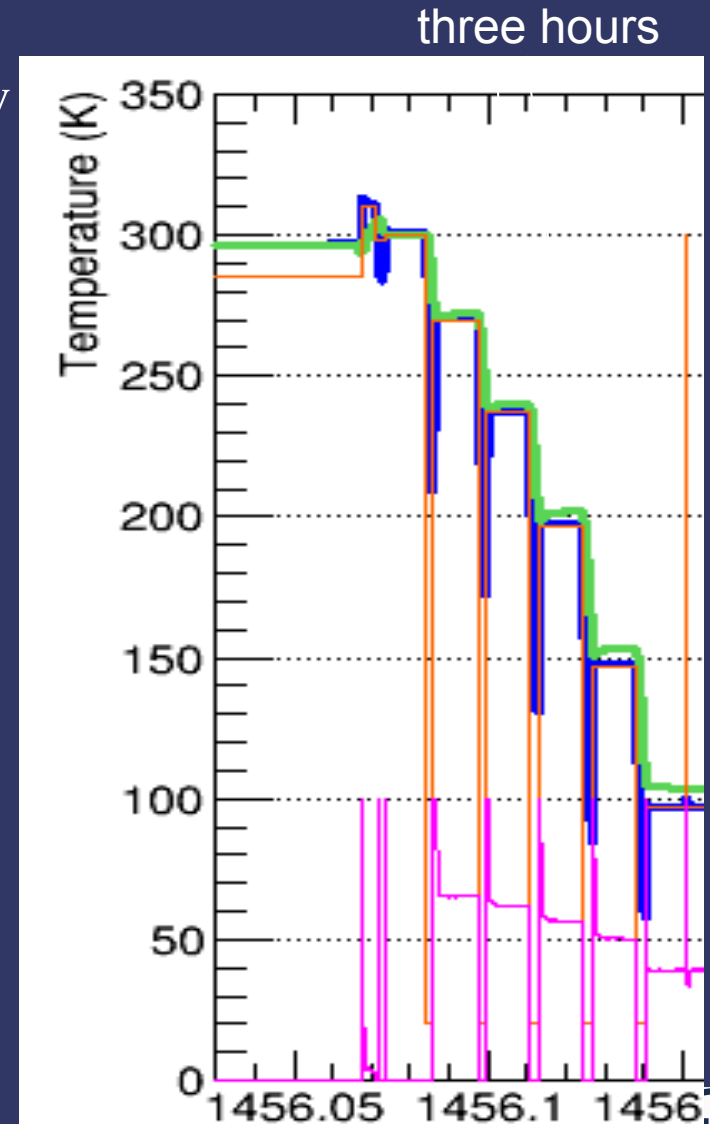
# 2016: experimental setup



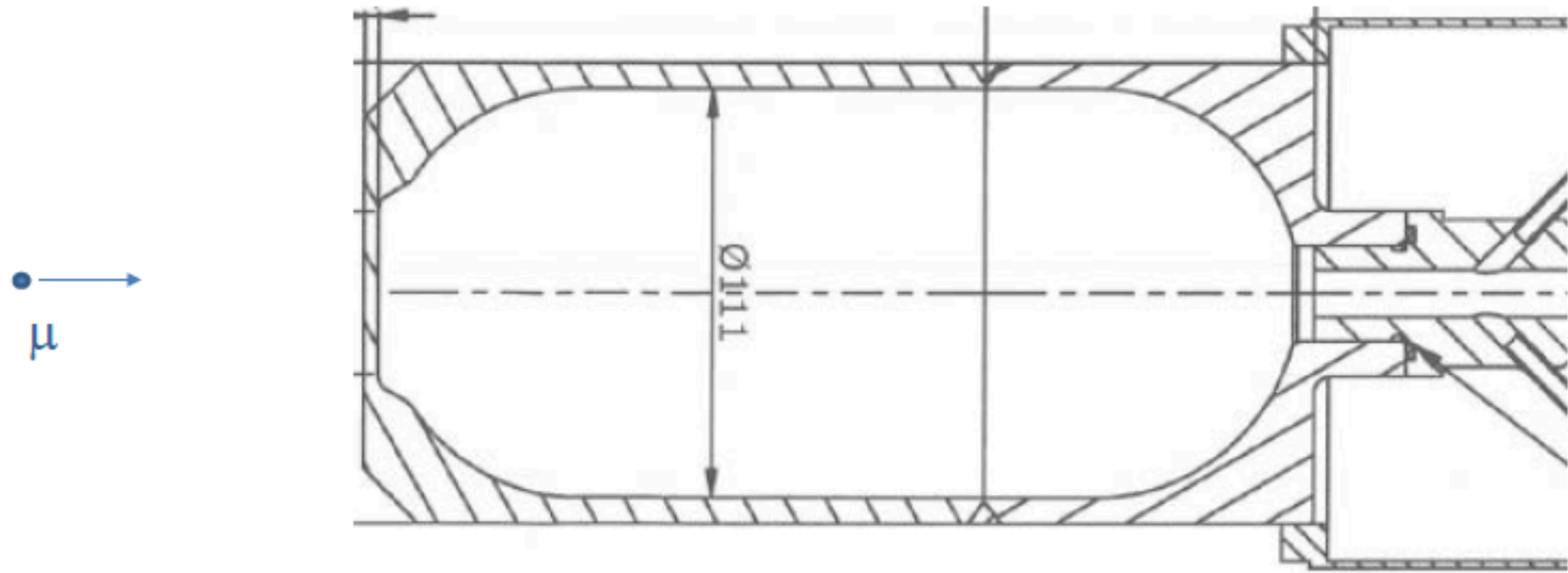
# 2016: transfer rate measurement

Steps:

- 1) fix a target temperature (i.e. mean kinetic energy)
- 2) produce  $\mu p$  and wait for thermalization
- 3) study time evolution of Oxygen X-rays
- 4) repeat with different temperature

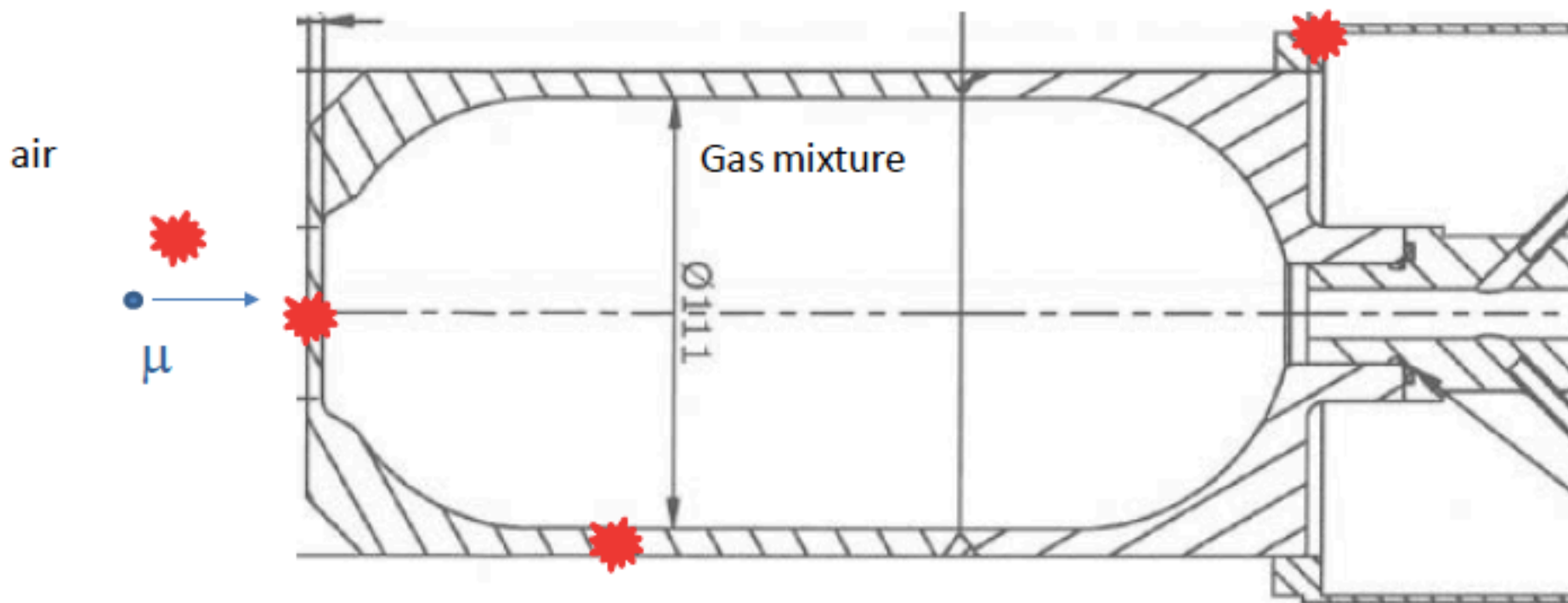


# $\mu$ fate in our setup (300K, 40 atm)



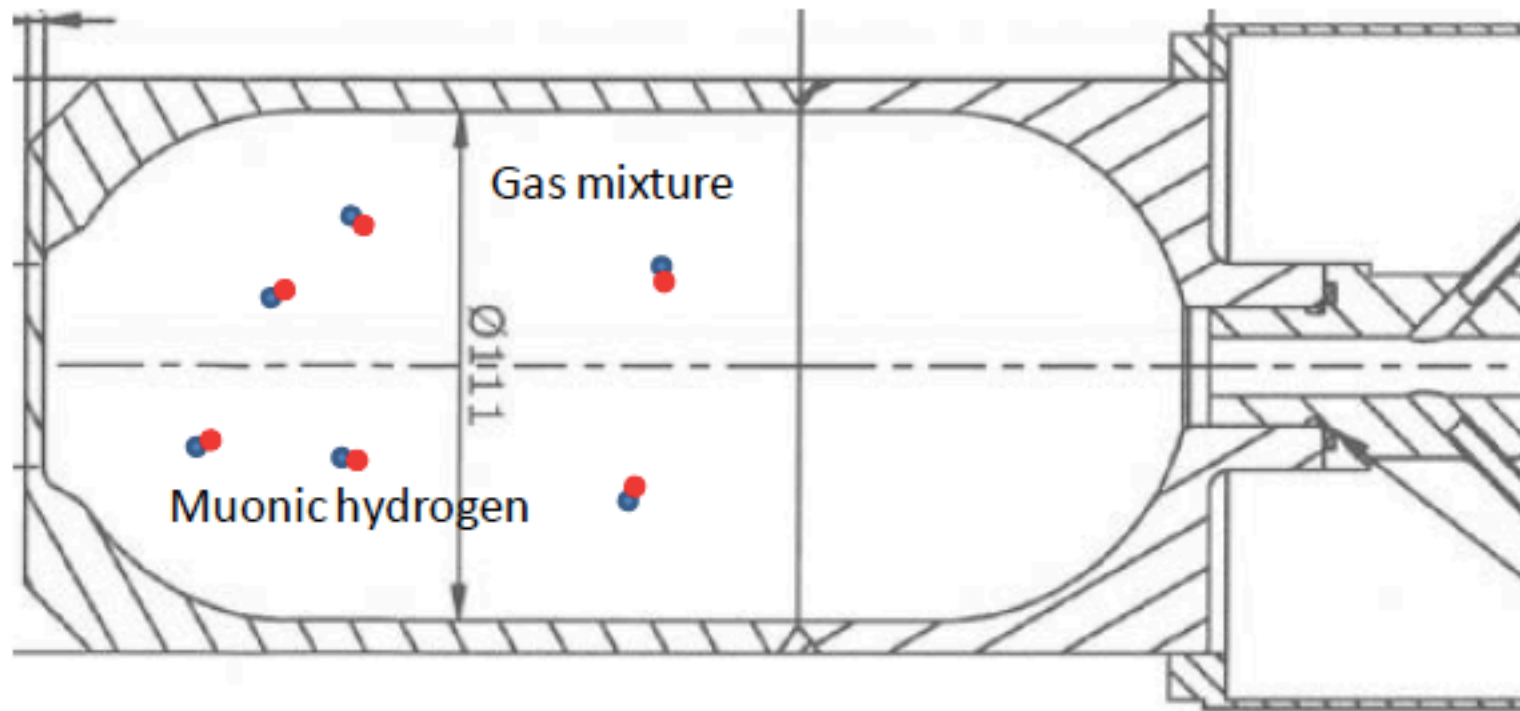
# $\mu$ fate in our setup (300K, 40 atm)

**Muons are capture by target vessel materials: X-ray prompt emission**

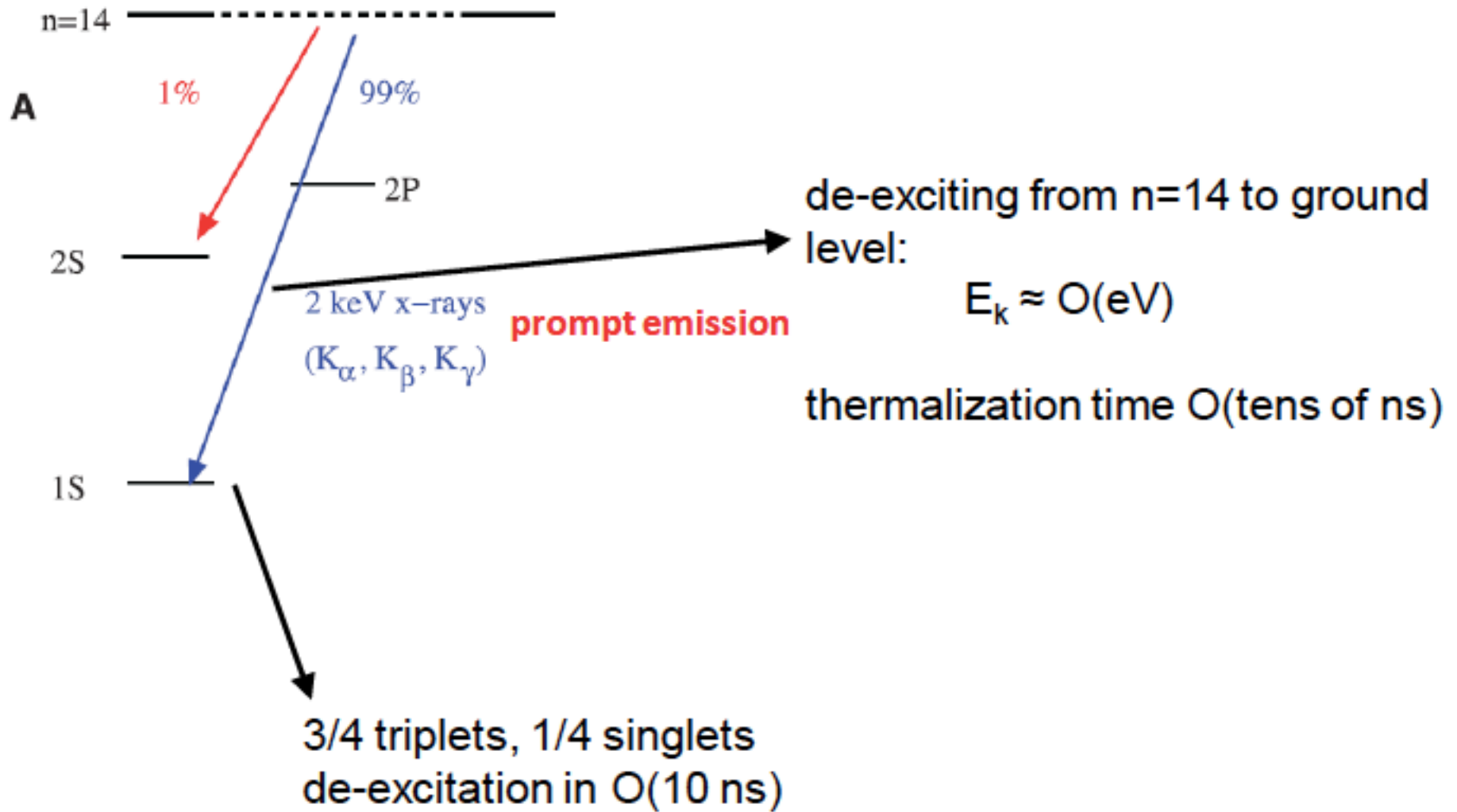


# $\mu$ fate in our setup (300K, 40 atm)

Muons are capture by hydrogen: muonic hydrogen forms inside the target



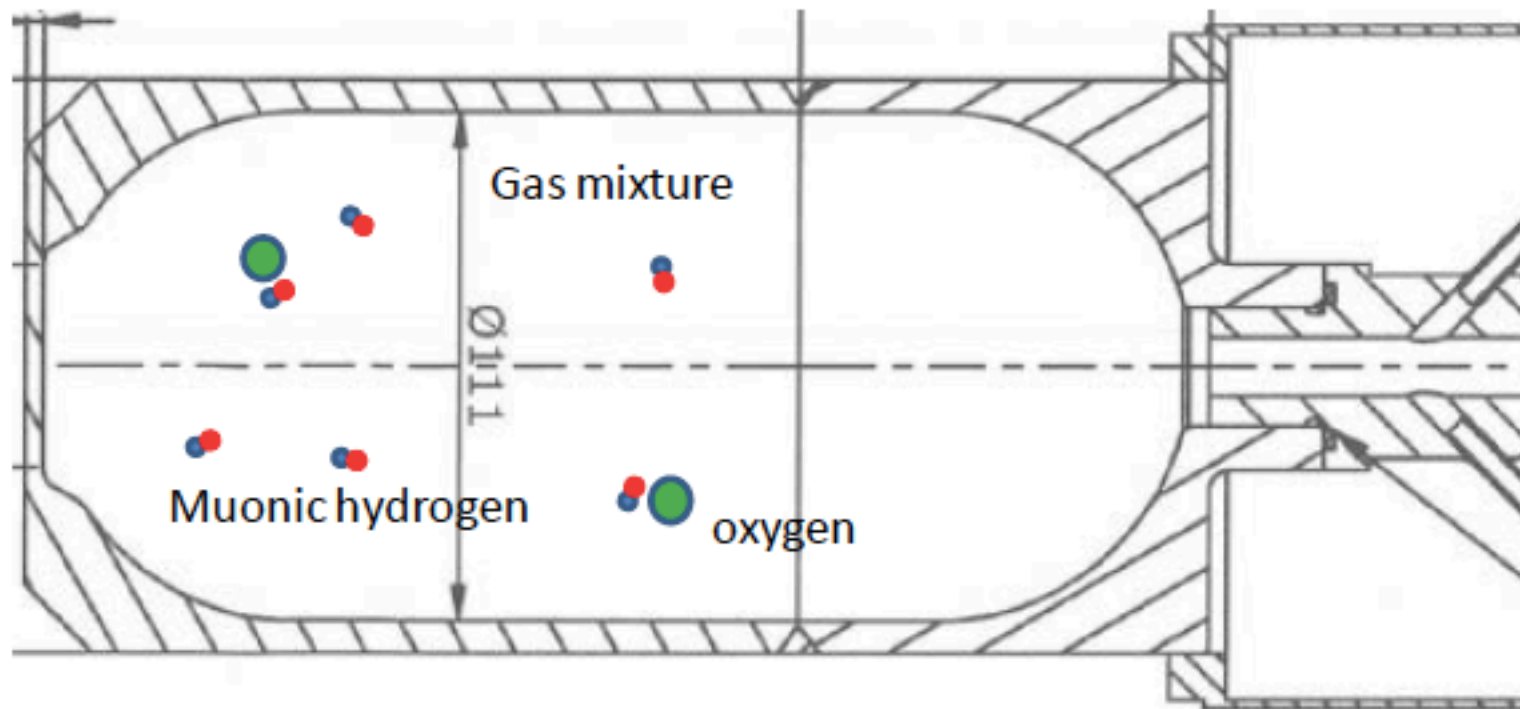
# $\mu$ fate in our setup (300K, 40 atm)





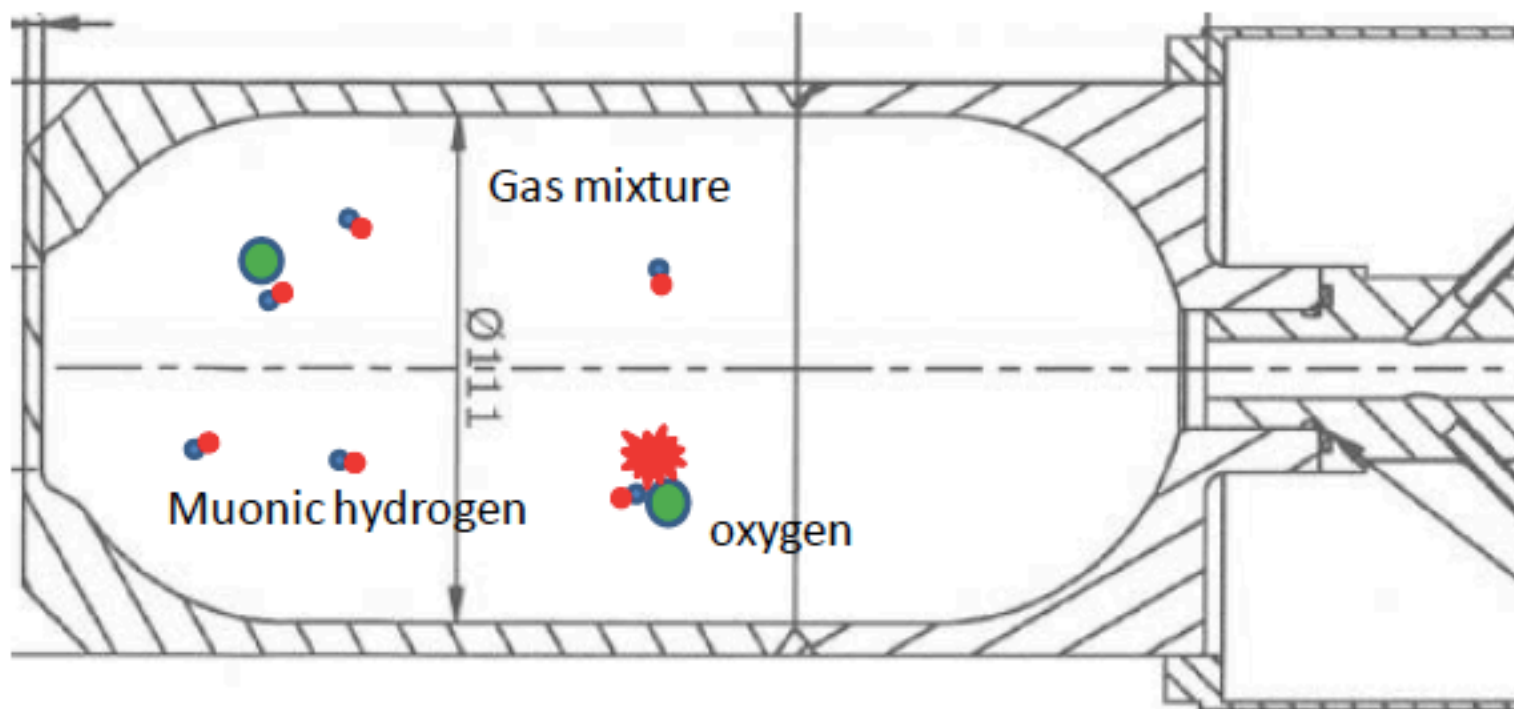
# $\mu$ fate in our setup (300K, 40 atm)

Muons are transferred from muonic hydrogen to other elements



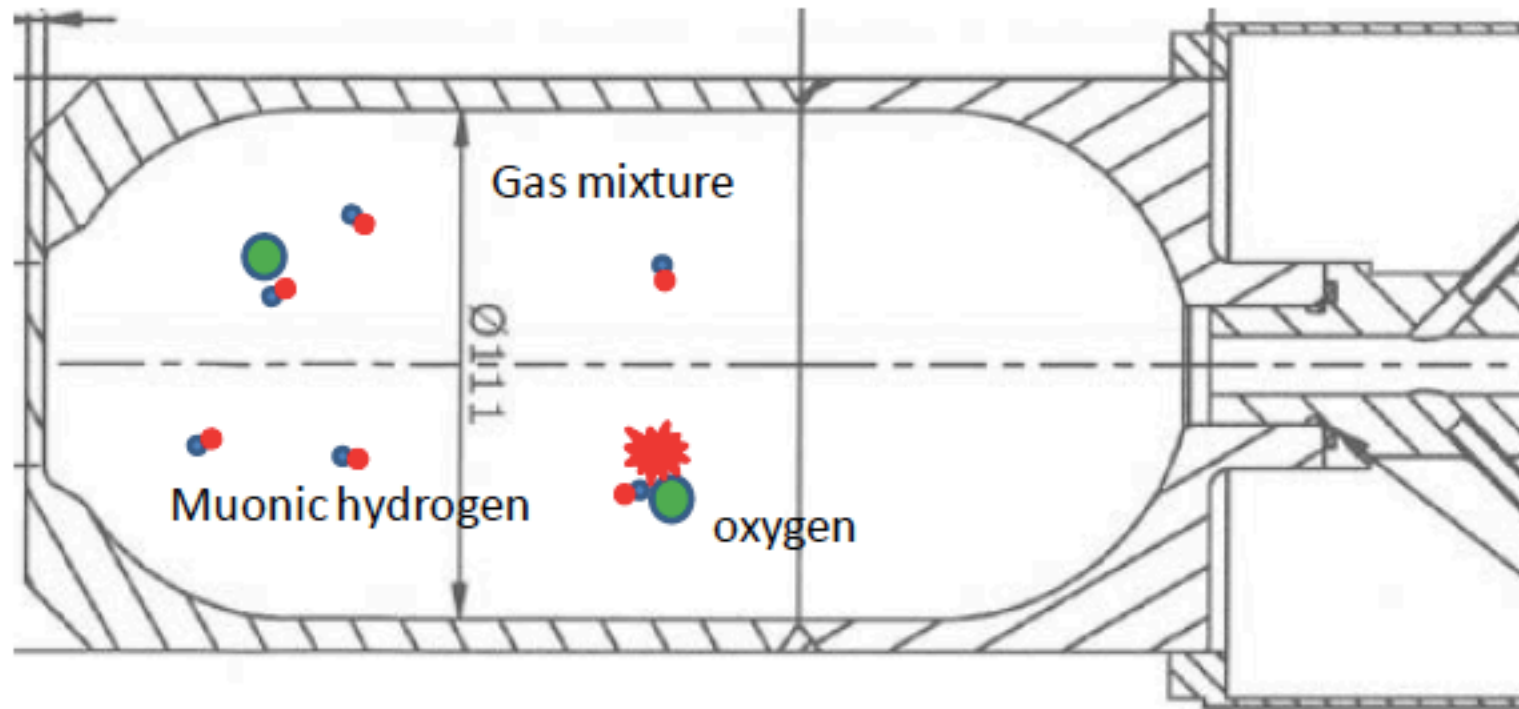
# $\mu$ fate in our setup (300K, 40 atm)

**Muons are transferred from muonic hydrogen to other elements: delayed emission**



# $\mu$ fate in our setup (300K, 40 atm)

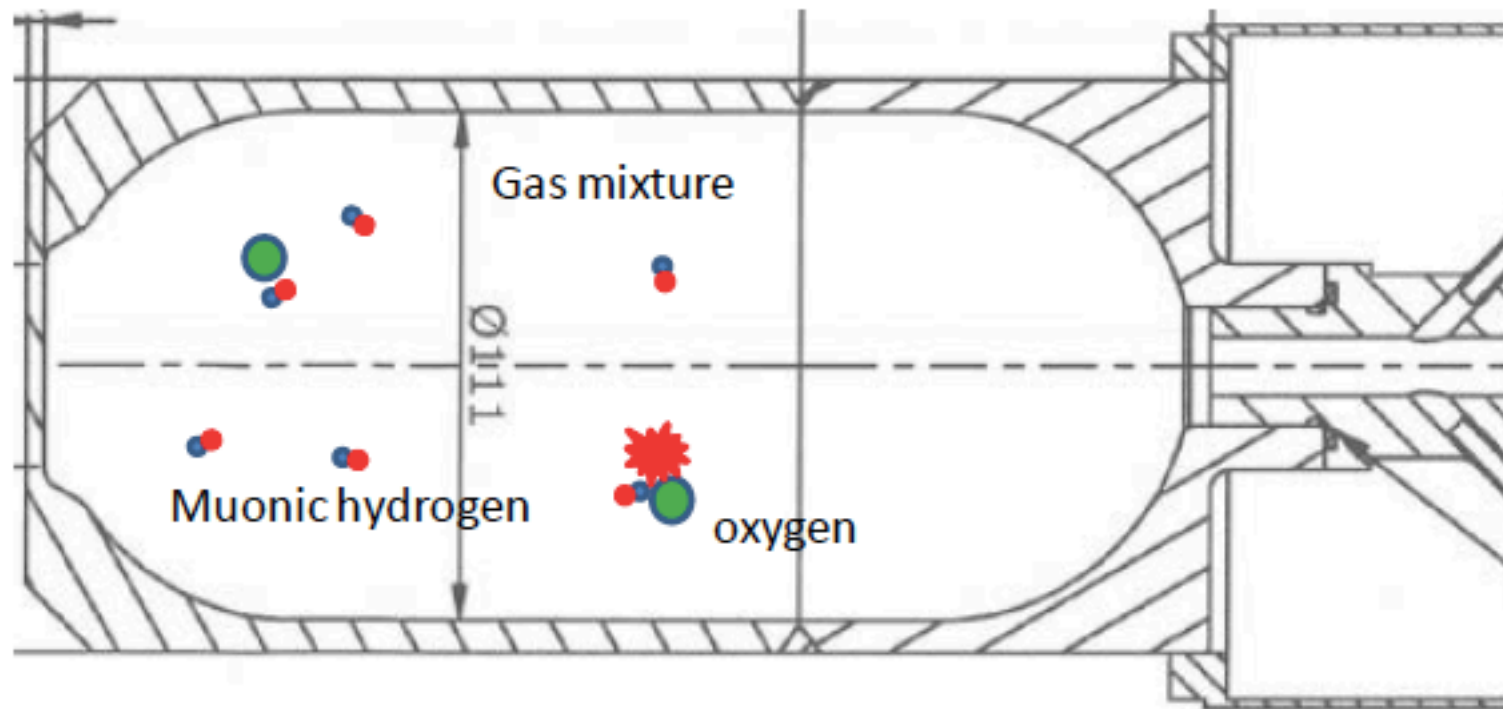
Muons are transferred from muonic hydrogen to other elements: delayed emission



Transfer speed depends on: gas density (i.e. pressure), contaminant concentration, and muonic hydrogen kinetic energy

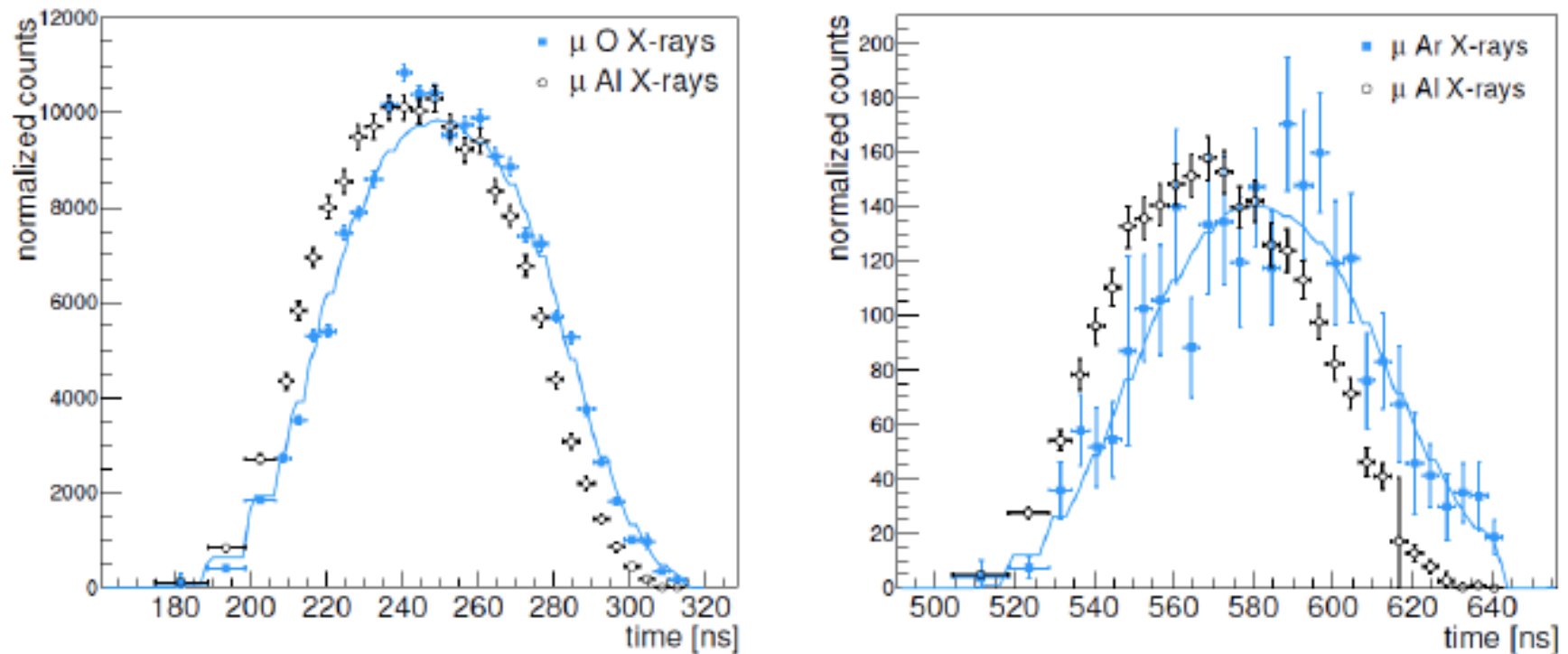
# $\mu$ fate in our setup (300K, 40 atm)

**Muons are transferred from muonic hydrogen to other elements: delayed emission**



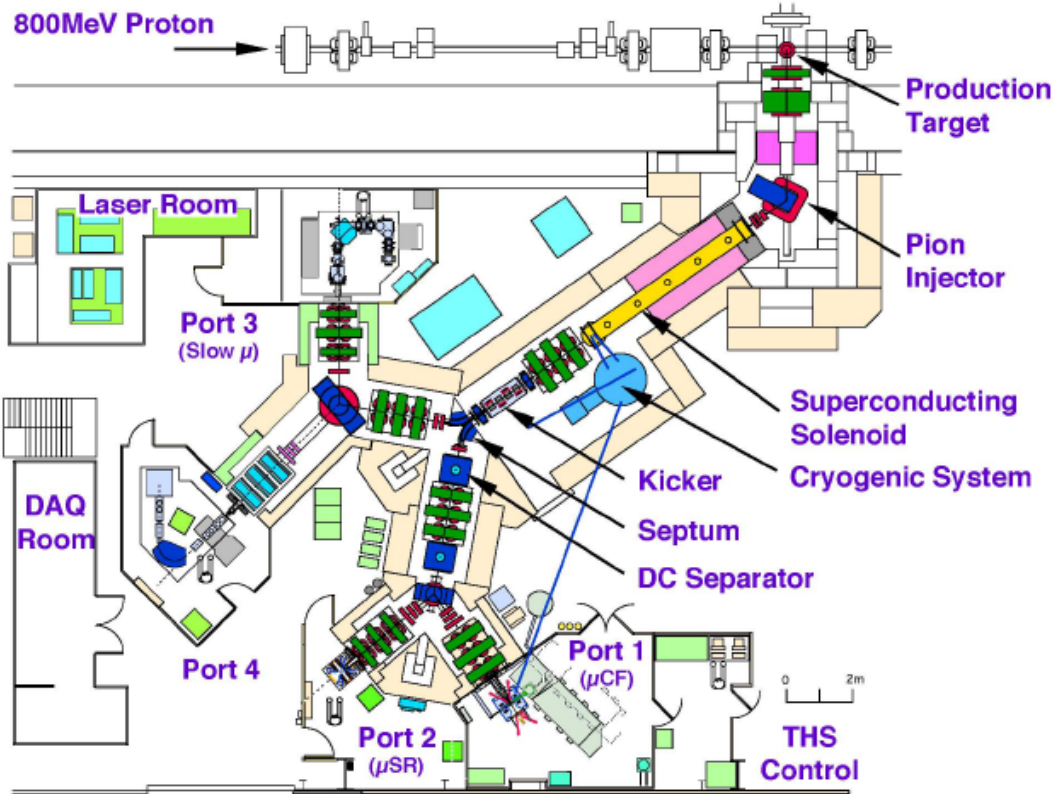
**Transfer speed depends on: gas density (i.e. pressure), contaminant concentration, and muonic hydrogen kinetic energy**  
**hence fast and strong signal until the thermalization of mu-p**

# In 2014 so fast that the delayed emission almost completely overlap the prompt emission

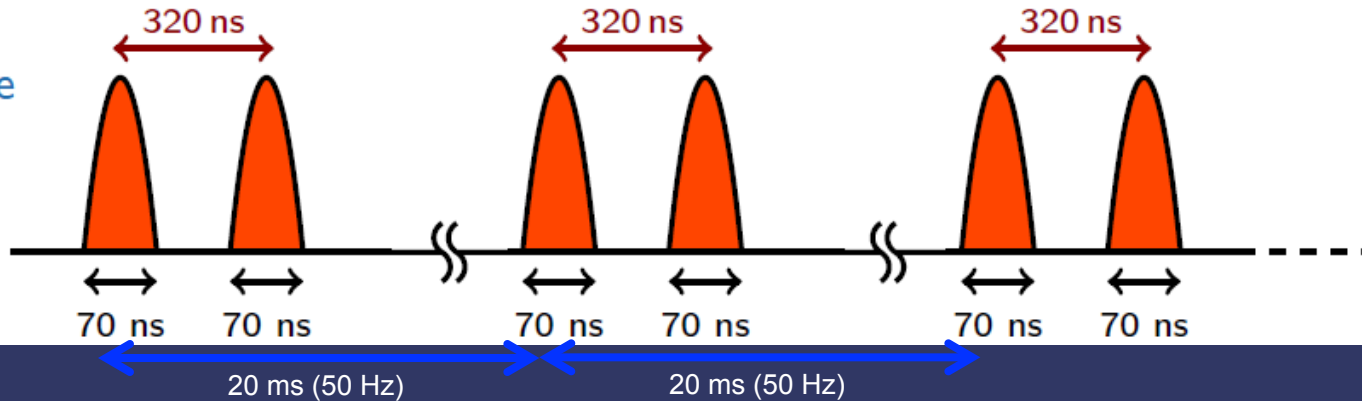


**Figure 1.** Left panel: time evolution of X-rays emission from aluminium and oxygen. Right panel: time evolution of X-rays emission from aluminium and argon. Lines are a fit to the oxygen and argon distributions, see text.

# High intensity muon beam



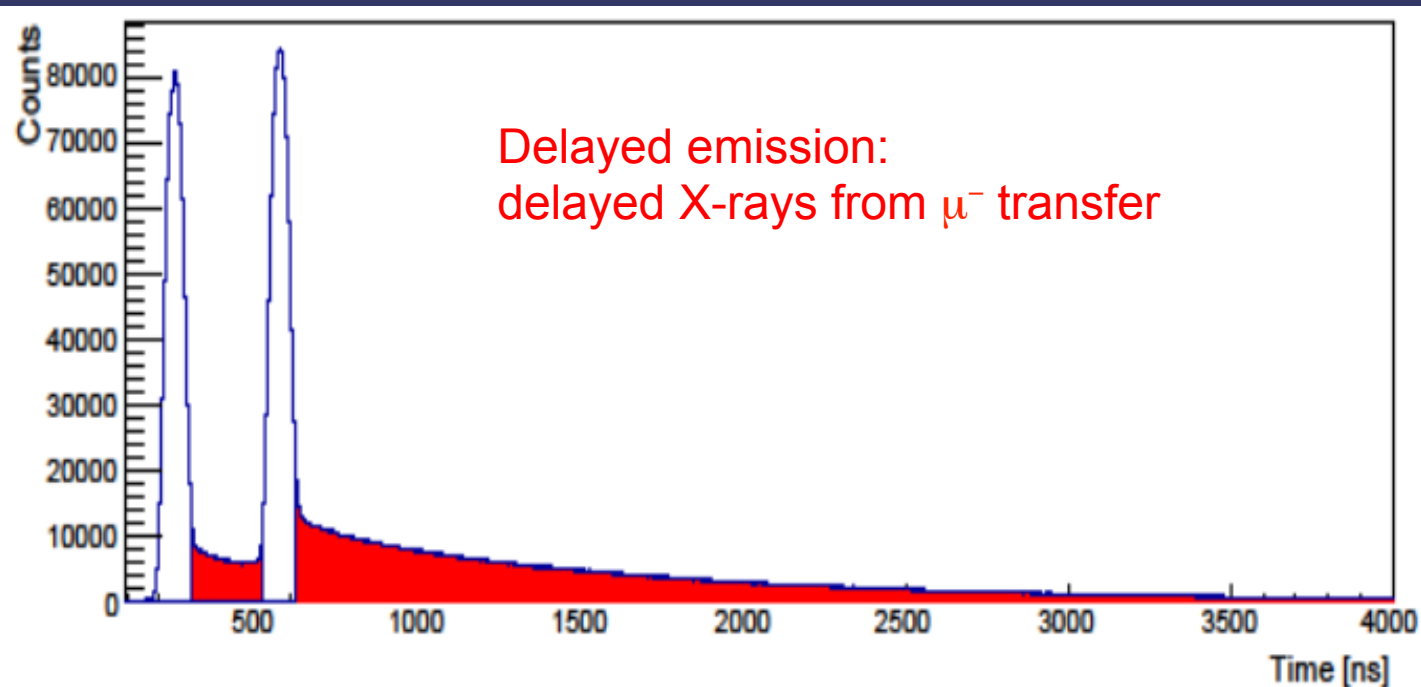
Beam time structure

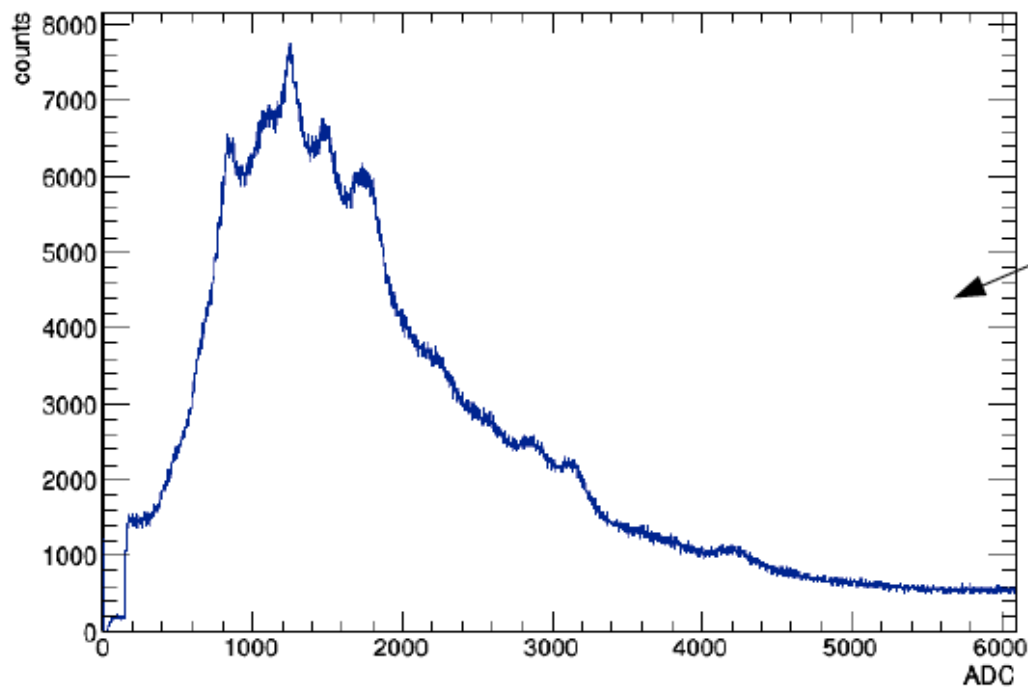


# 2016: transfer rate measurement

Steps:

- 1) fix a target temperature (i.e. mean kinetic energy of gas constant)
- 2) produce  $\mu p$  and wait for thermalization
- 3) study time evolution of Oxygen X-rays
- 4) repeat with different temperature



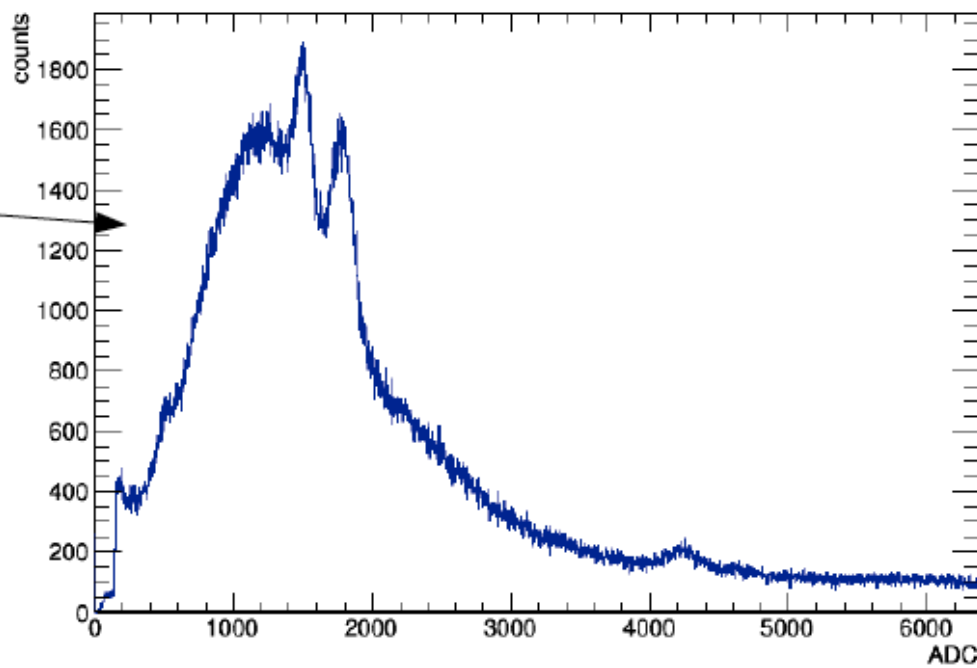


PROMPT SPECTRUM  
( $t < 1200$  ns)

One detector (LaBr 3)  
0.3% oxygen  
concentration

DELAYED SPECTRUM  
( $t > 1200$  ns)

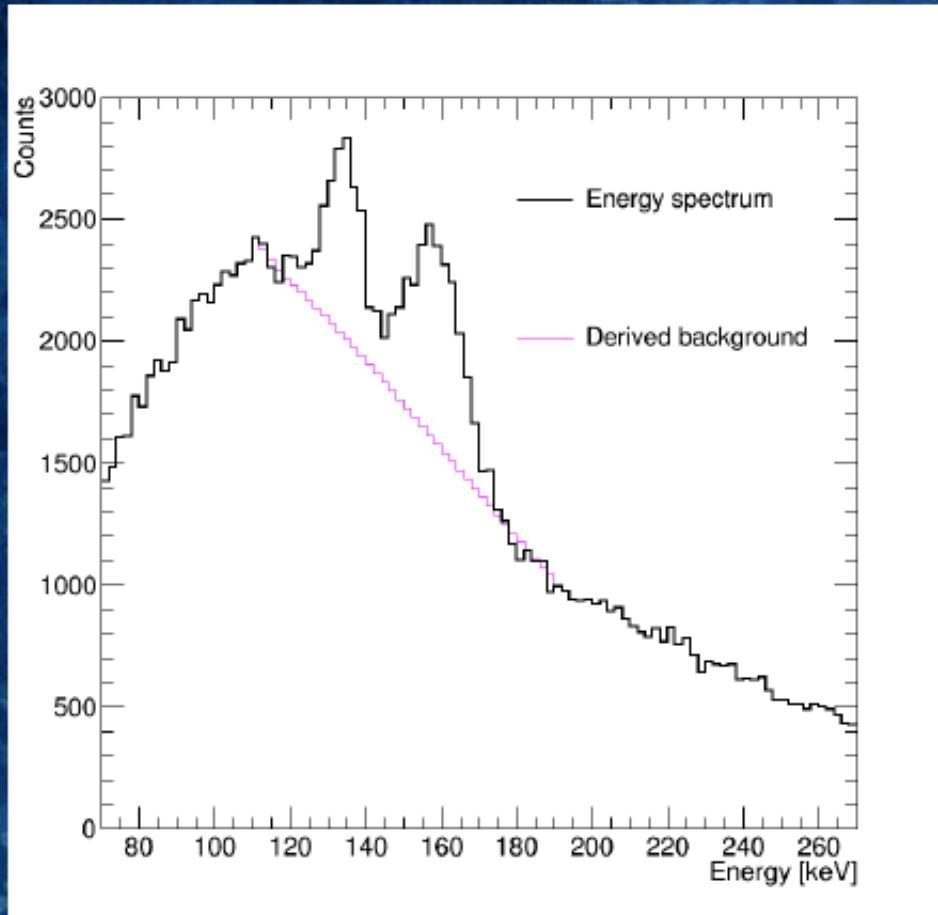
One detector (LaBr 3)  
0.3% oxygen  
concentration





# BACKGROUND EVALUATION

Simplest solution: “straight line”



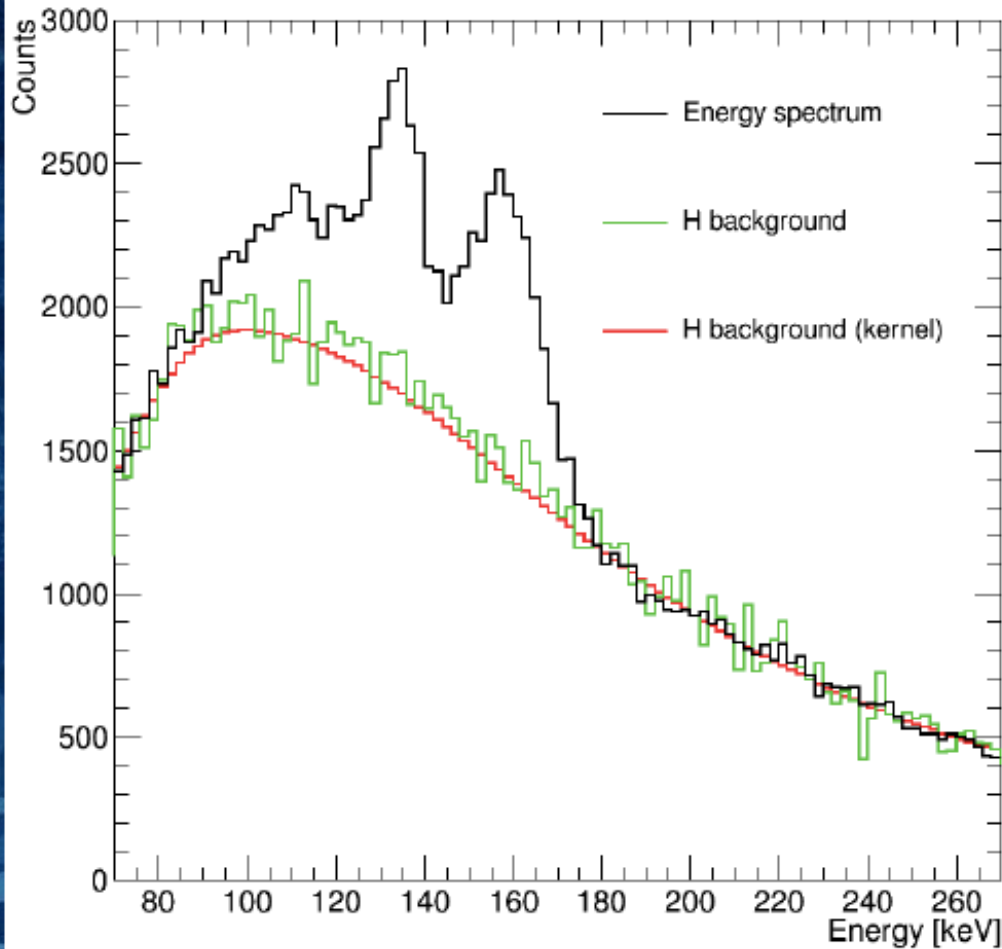
$T = 300 \text{ K}$

Time bin = [1450, 1650] ns

Using ROOT/“TSpectrum” class –  
spectroscopic algorithms

Problem: unstable results...

# Best solution: pure H smoothing

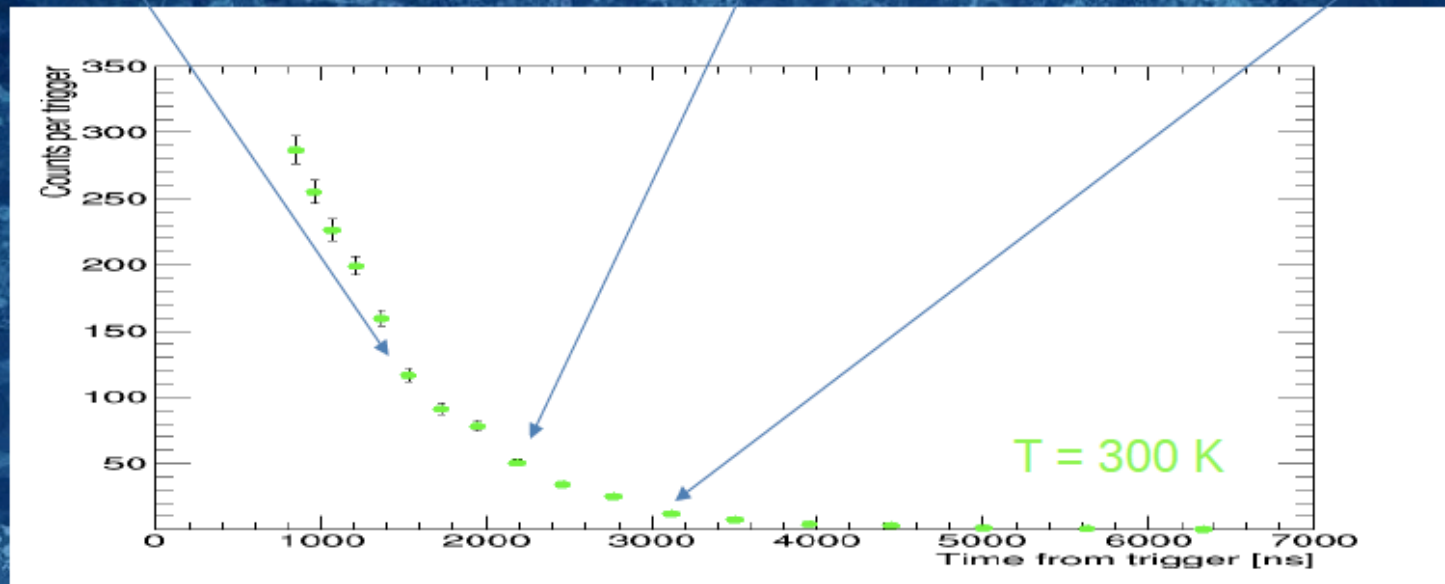
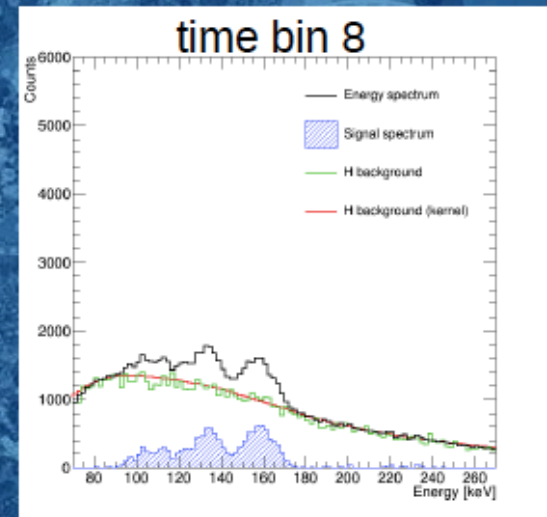
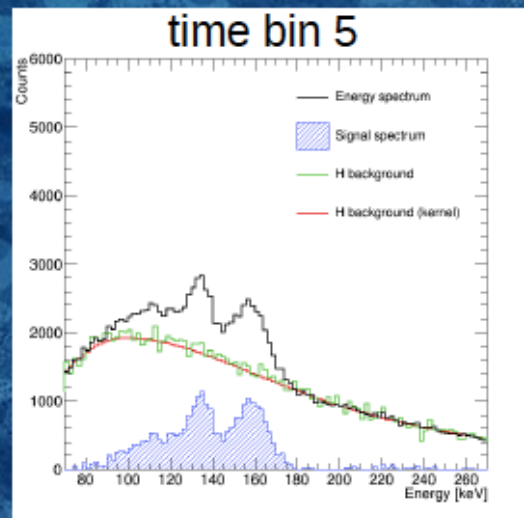
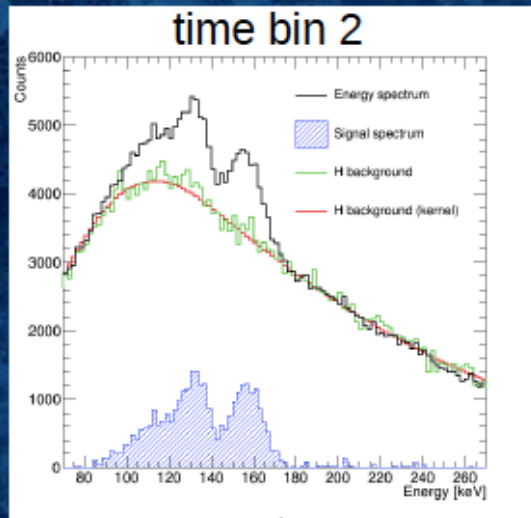


$T = 300 \text{ K}$

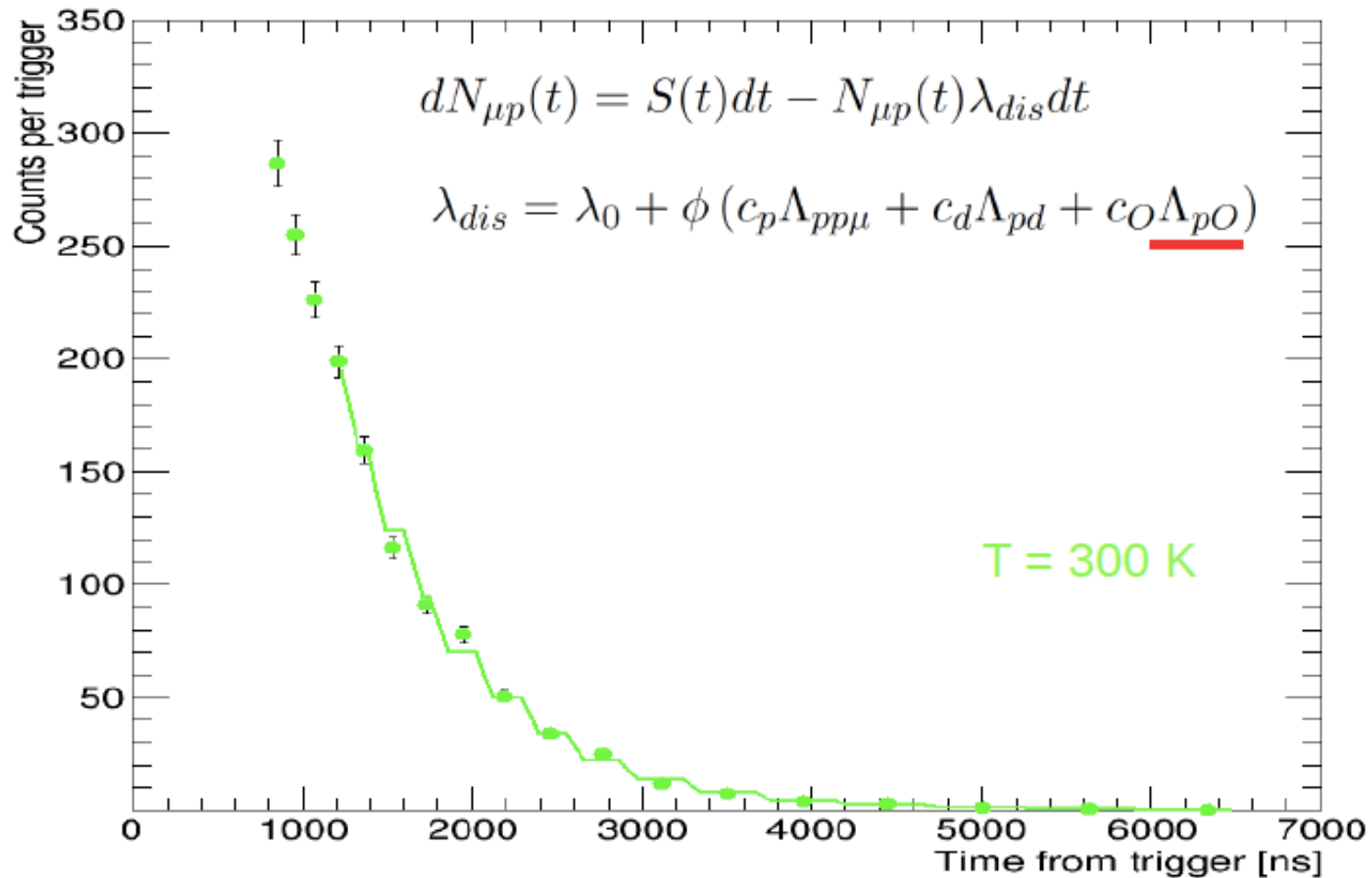
Time bin = [1450,1650] ns

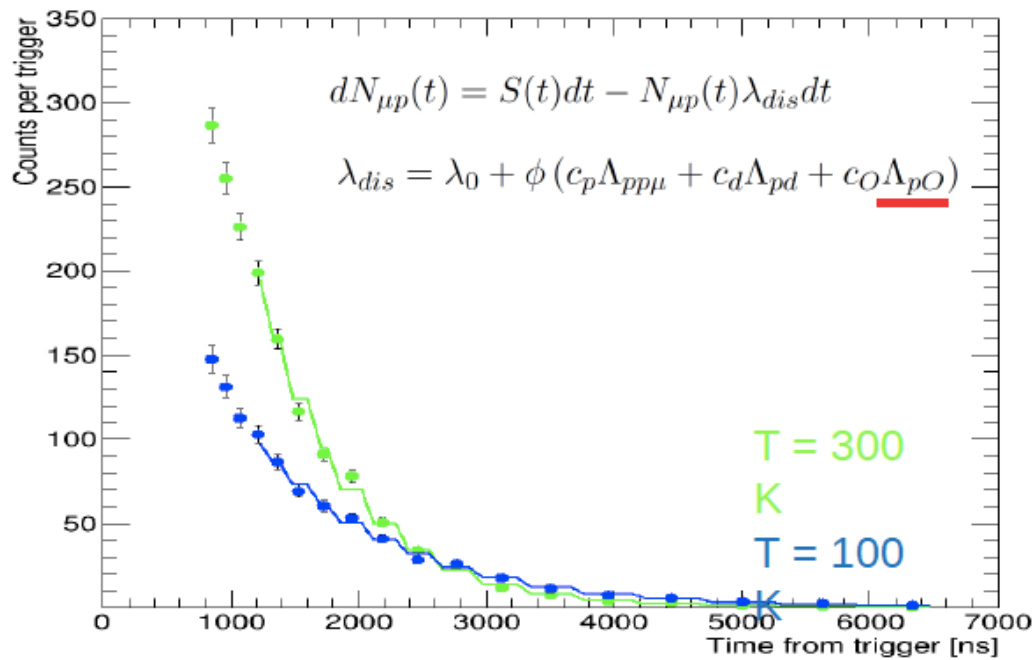
Pure hydrogen data taking within the same beam time and with the same pressure and temperatures.

# Fixed temperature: time evolution

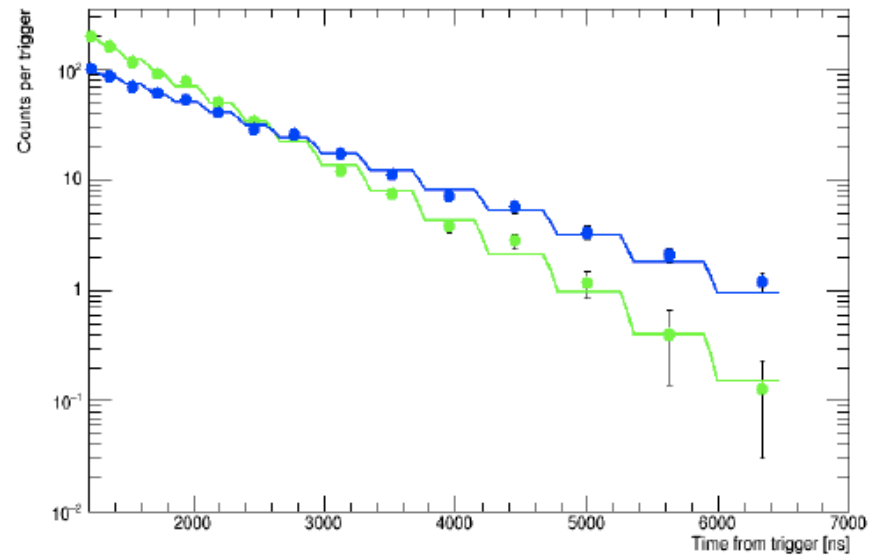


# Temperature and time evolution



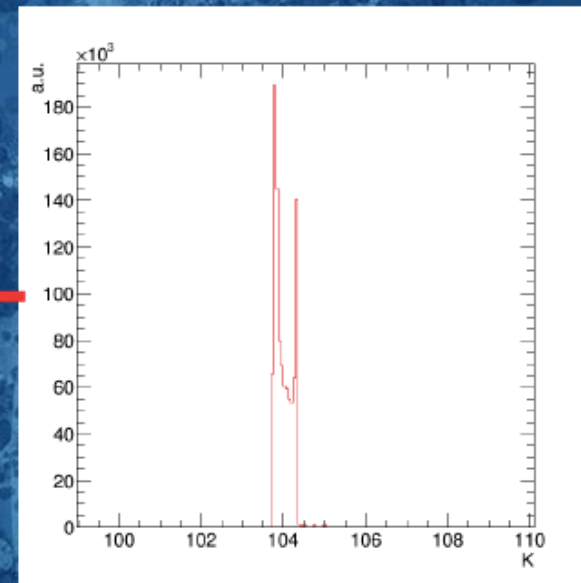
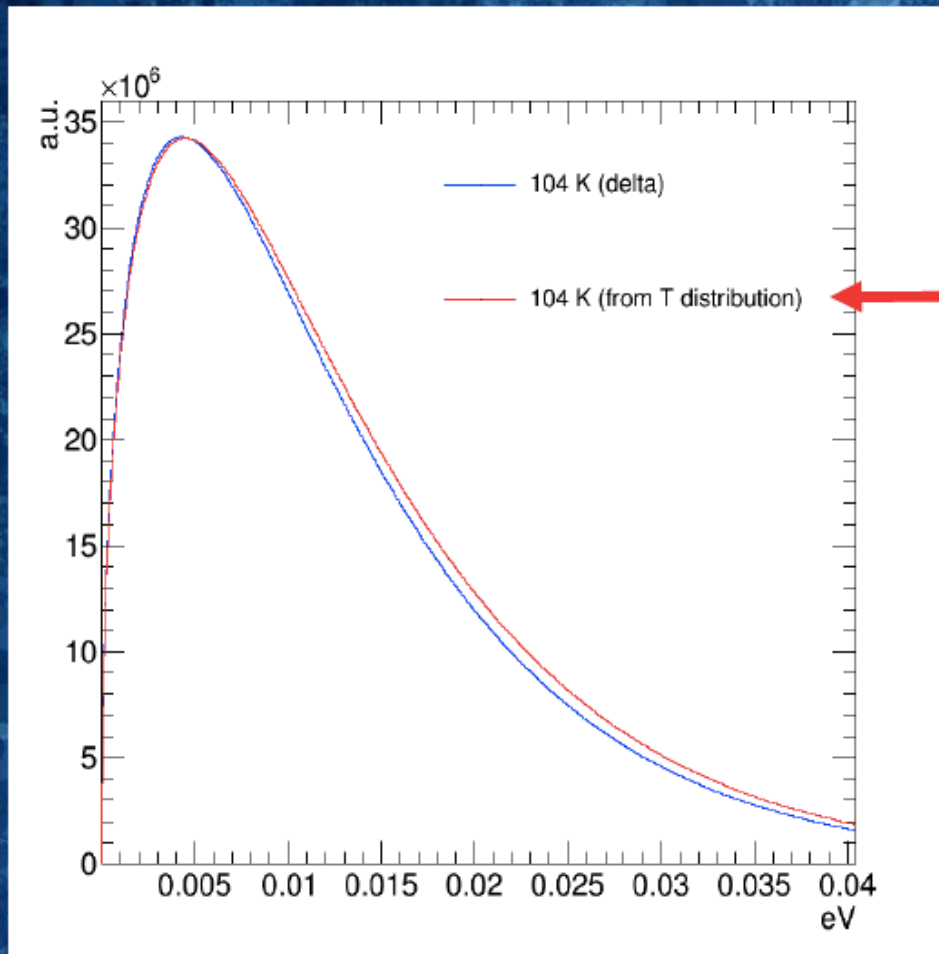


Transfer rate dependence  
on temperature



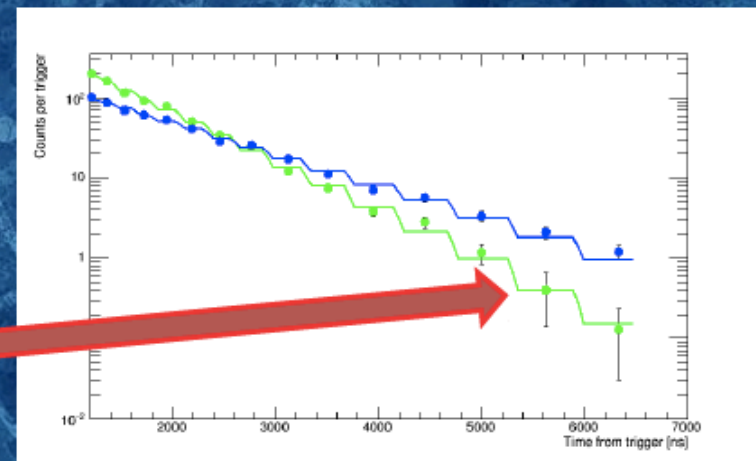
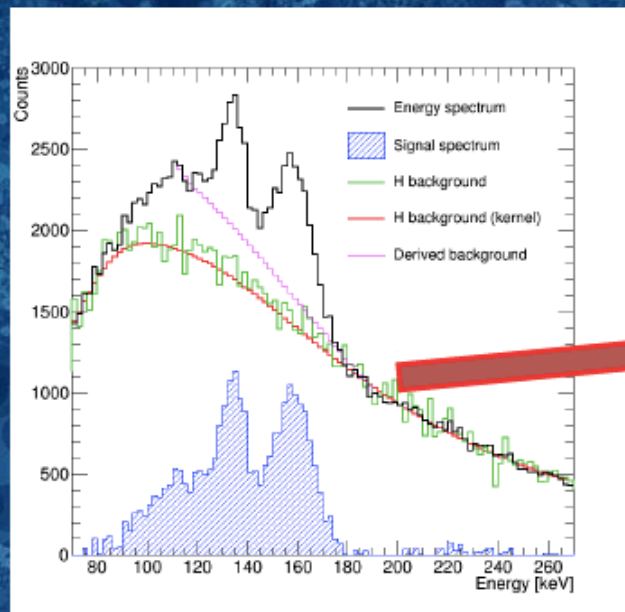
# Temperature $\leftrightarrow$ energy conversion

- Definition of temperature interval
- Generation of a Maxwell-Boltzmann distribution for each temperature in that interval



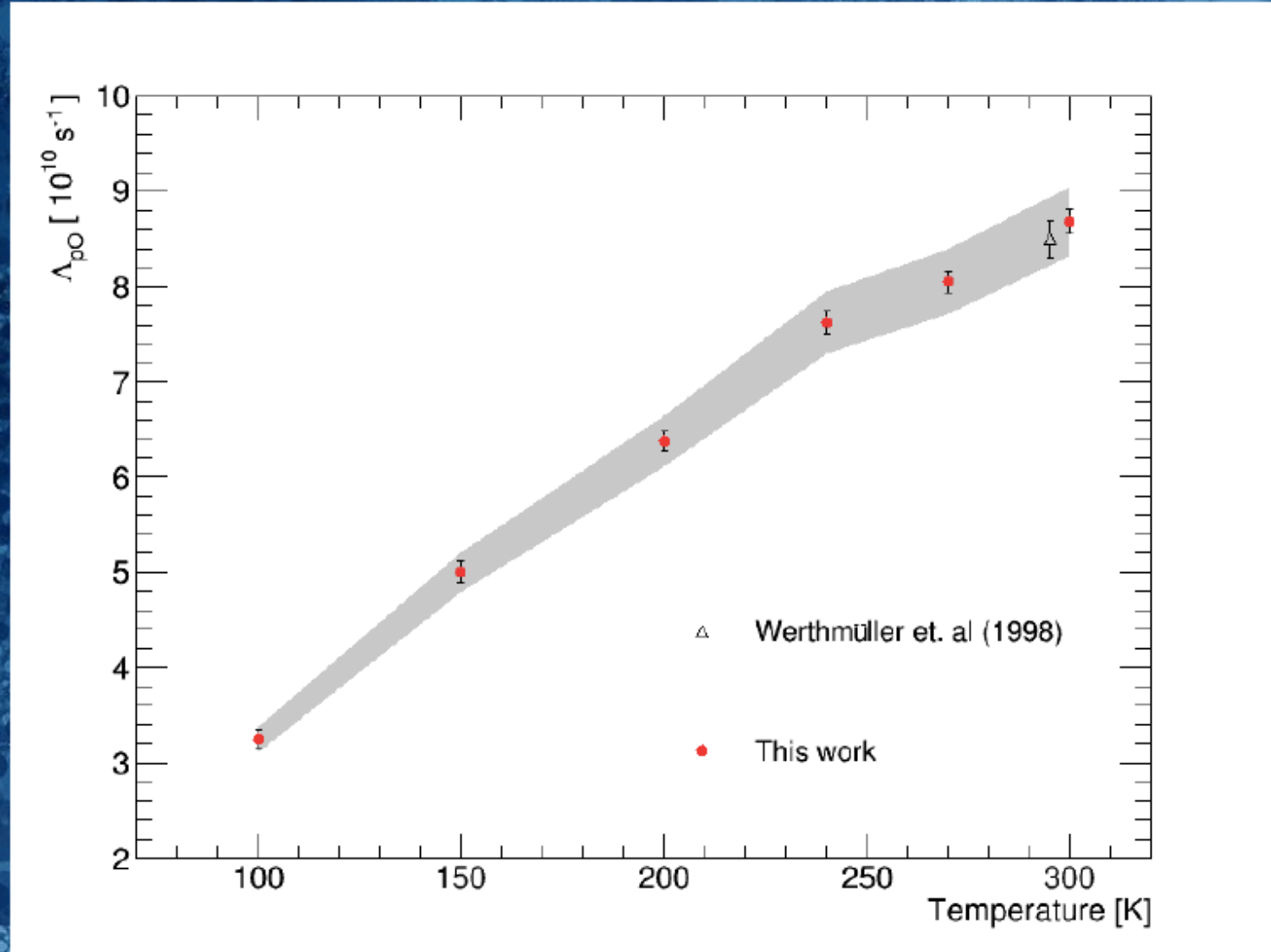
# Sistematic errors contributions

- 3% given by the O concentration calculation
- 3% given by the density calculation
- About 5% due to the procedure of the background subtraction



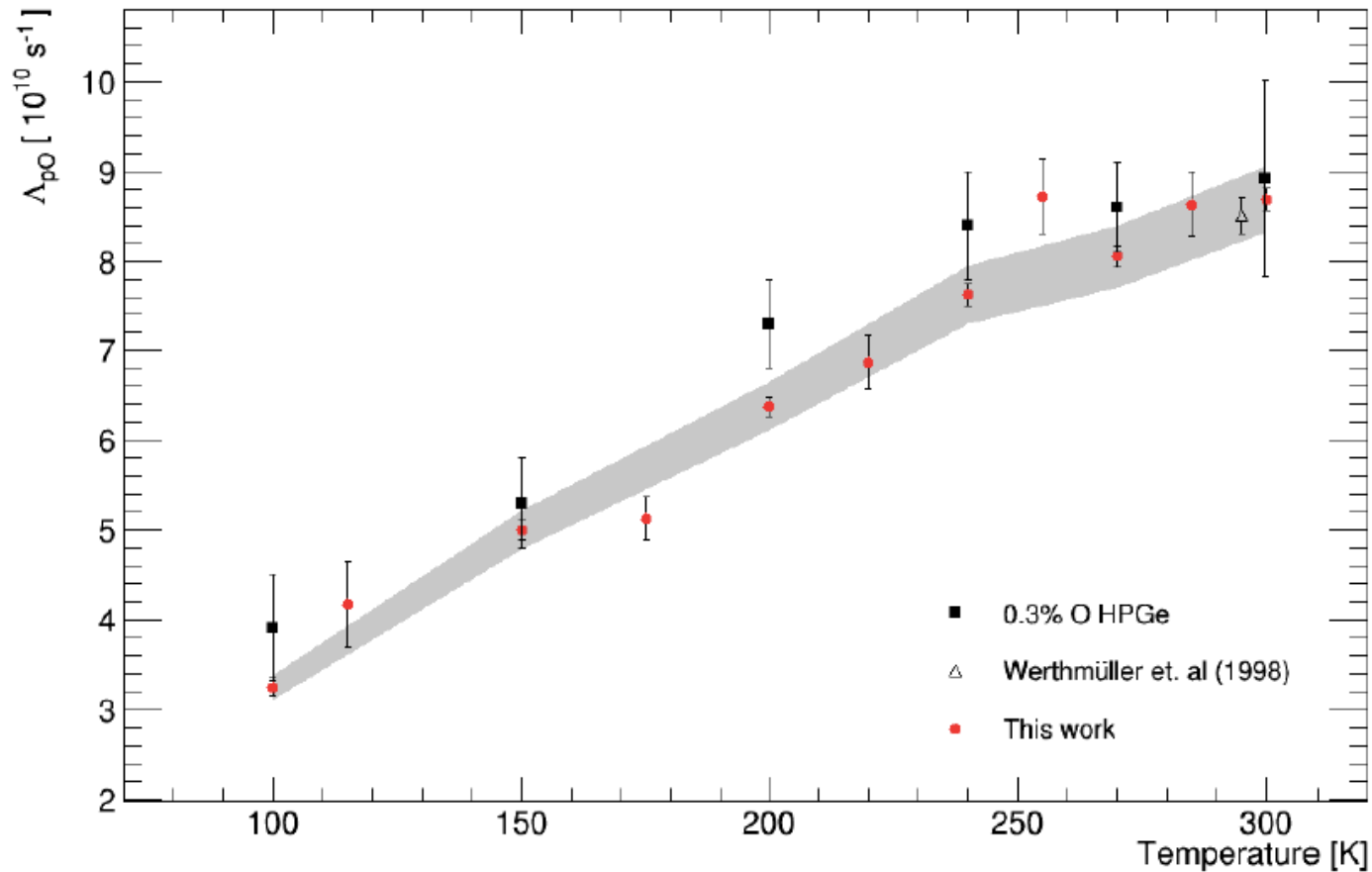
Other uncertainties, negligible ( $\ll$  statistical error)

# Final result for oxygen concentration at 0.3%

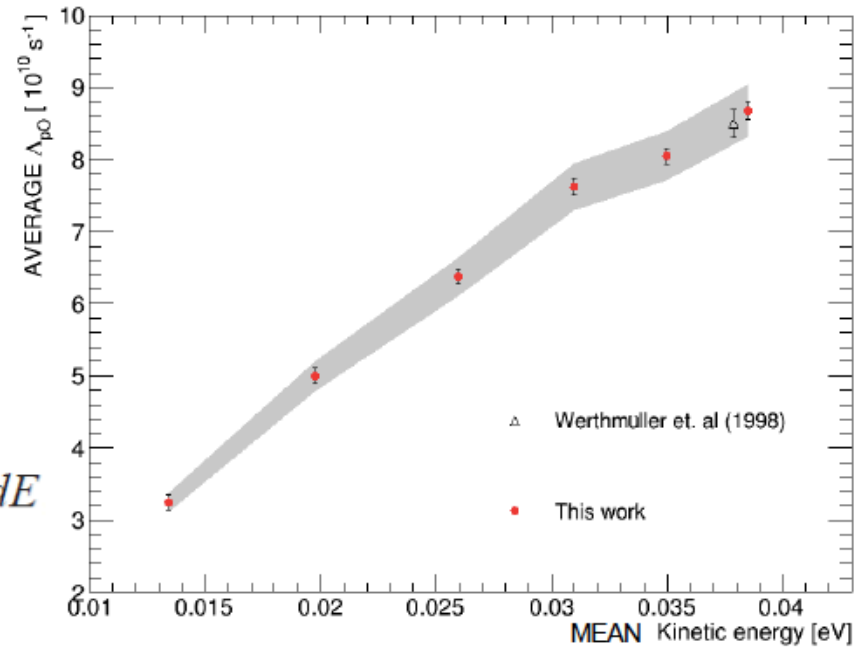
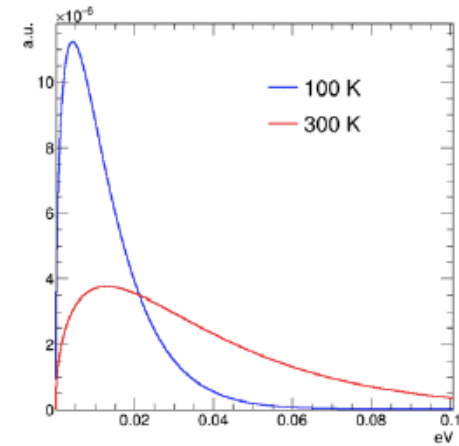
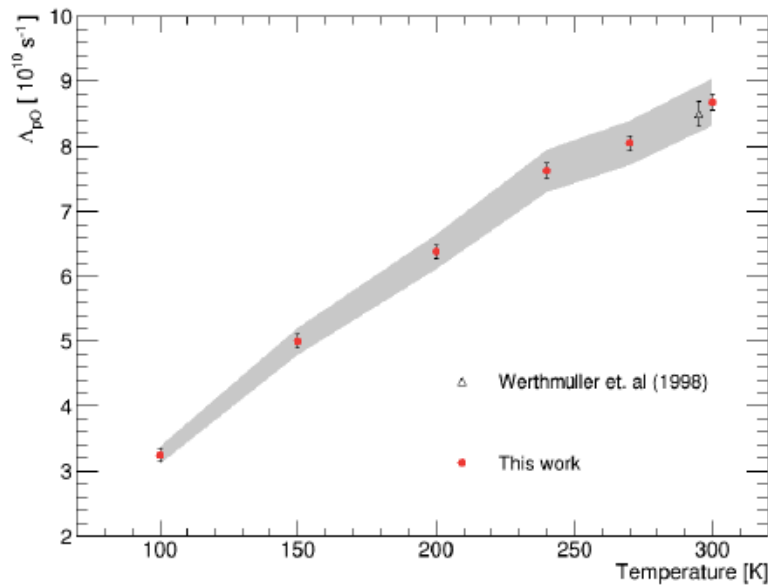




# Comparison between the transfer rate result obtained analysing the data acquired with HPGe for the 0.3% oxygen concentration



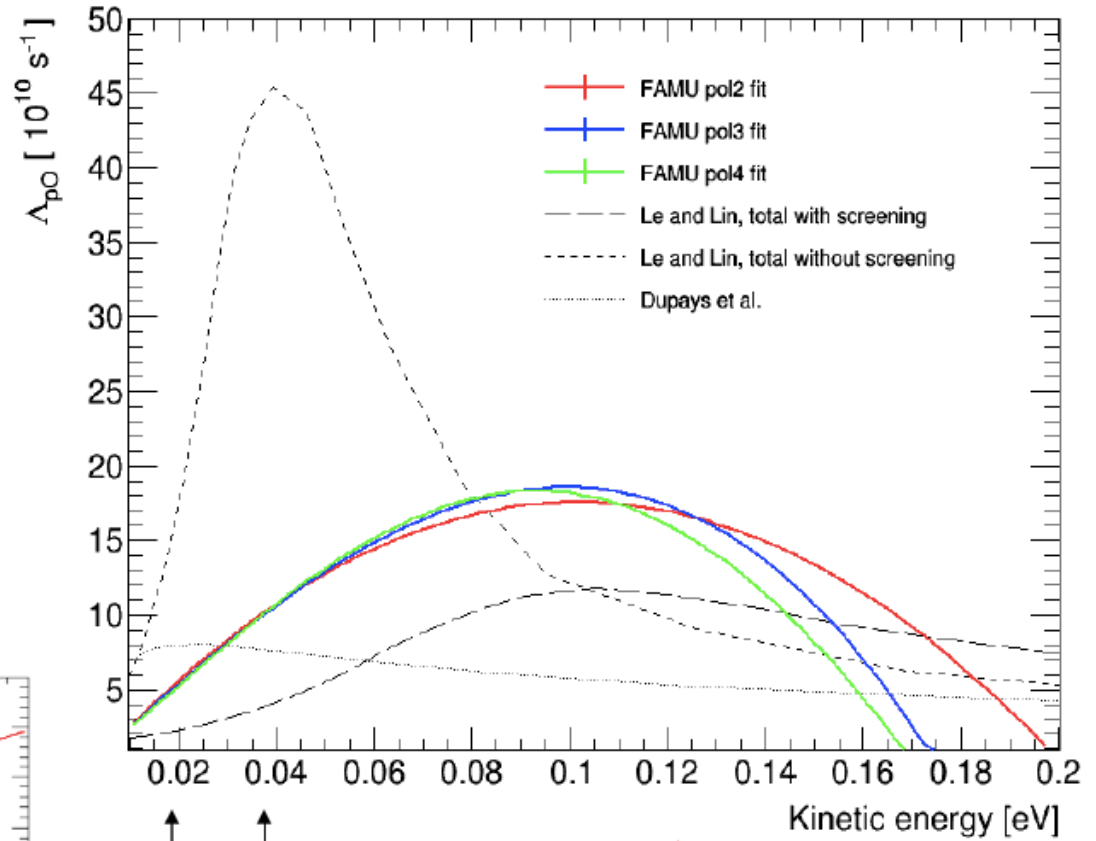
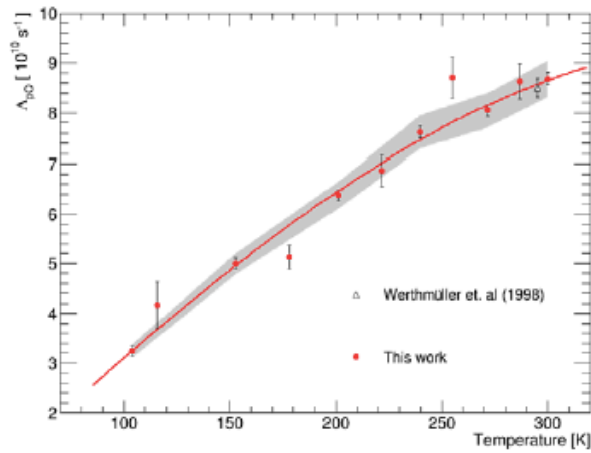
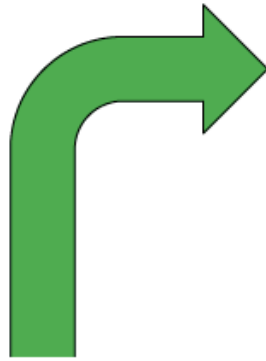
# Conversion to kinetic energy



$$\lambda(T) = \int_0^{\infty} \lambda(E) \sqrt{\frac{4E}{\pi(kT)^3}} \exp(-E/kT) dE$$

# Phenomenological model

$$\lambda(T) = \int_0^{\infty} \lambda(E) \sqrt{\frac{4E}{\pi(kT)^3}} \exp(-E/kT) dE$$



$\uparrow$  ~100 K  
 $\uparrow$  ~300 K

$\uparrow$   
 $\mu\text{p}$  energy after de-excitation

# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

## two competing processes

- the strength of x-ray signal, from mu-p's accelerated via the laser shot, is proportional to
  - the ratio of the muon transfer rate to oxygen and
  - the thermalization rate.
- both of them are proportional to the target density,

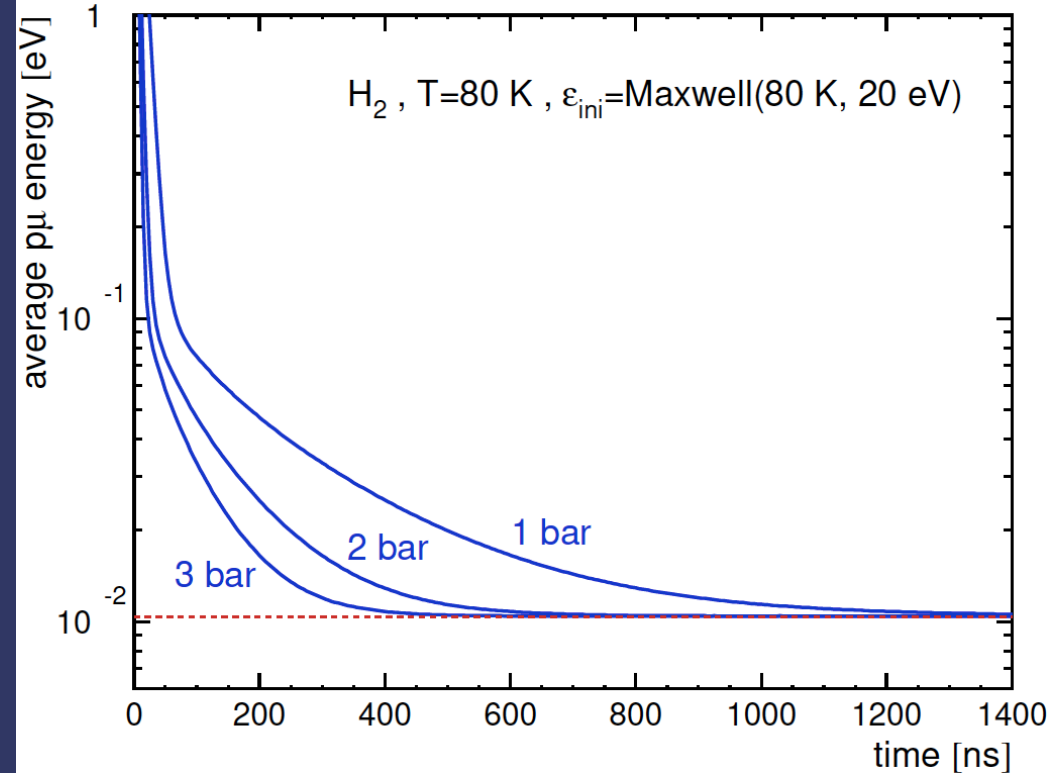
# need to optimize the relevant parameters

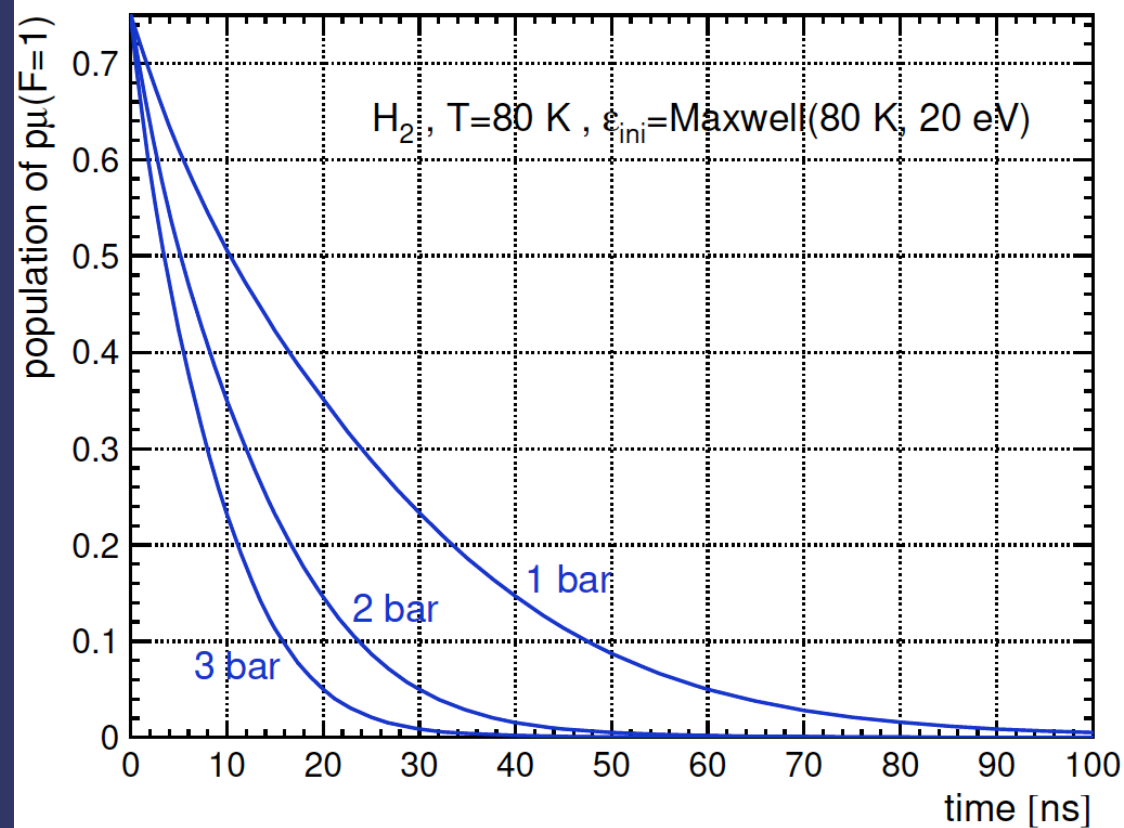
- The only parameter we can use to enhance the signal is the oxygen concentration - the x-ray signal is directly proportional to this concentration.
- The target pressure cannot be too small. We need a reasonable amount of the muon stops within the volume of laser field.
- The overall **optimal condition** for the HFS measurement is thus a convolution of these two optimization functions **is going to be determined by the HFS-measurement simulations, which is underway.**

the time evolution of mean kinetic energy, which also illustrates the thermalization time. This is an answer for the moment of laser ignition - mu-p atoms should be thermalized

Also, this picture shows a mean time of deceleration from about 0.11 eV

(meanenergy of mu-p's after the laser excitation and downwards spin flip) to about 0.04 eV (lower energy of a relatively high muon transfer rate to oxygen,). Within this time, the most of muon-transfer events should take place, in order to have a strong signal. We can increase this signal only by increasing the oxygen concentration, within certain limits.

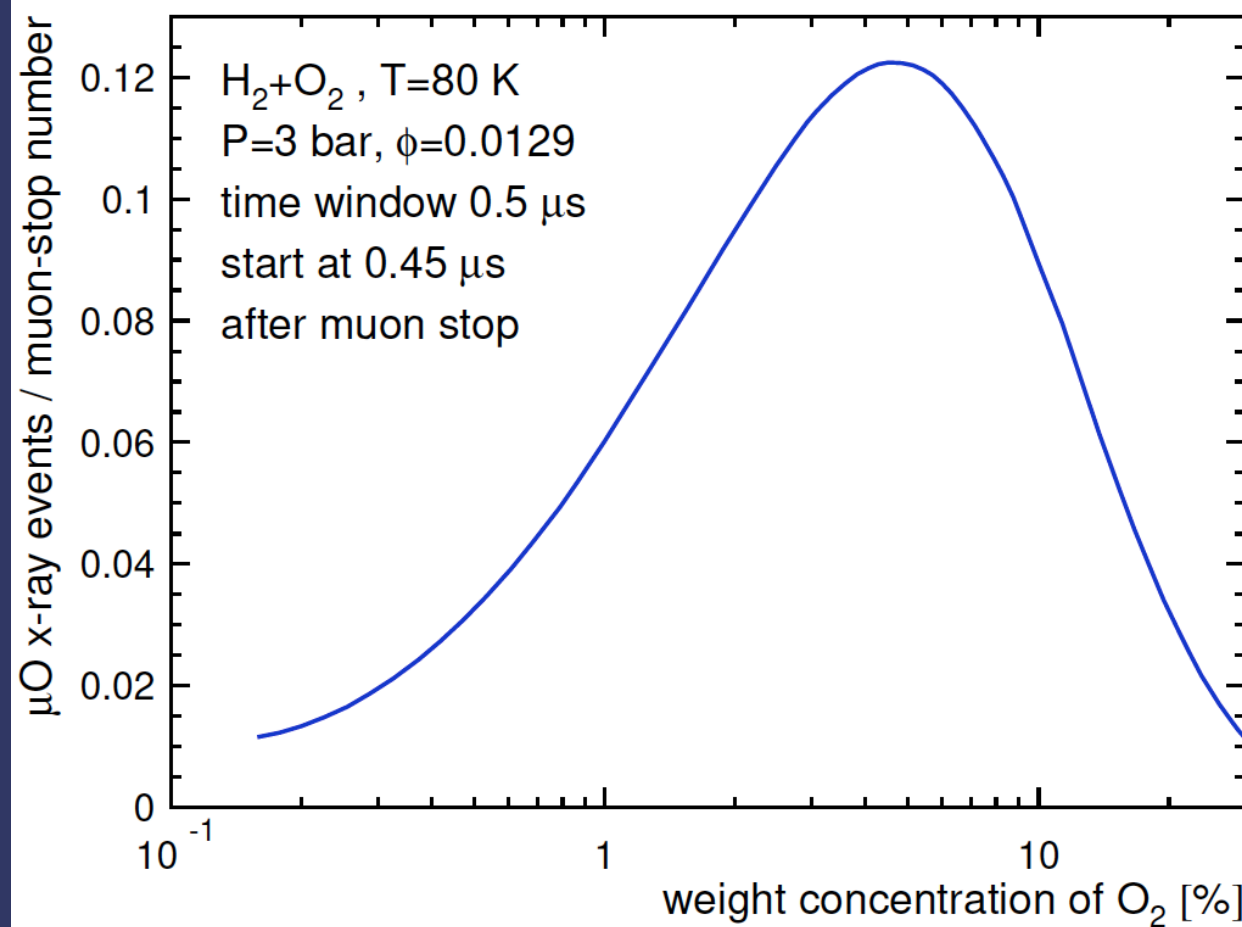




mu-p spin de-excitation versus time.

The de-excitation time informs us about how long we should wait for the acceleration of mu-p atom which was excited by a laser photon.

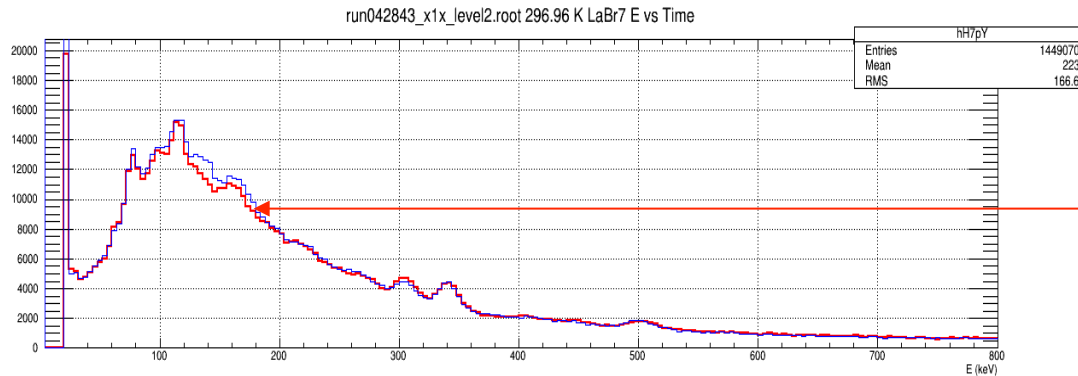




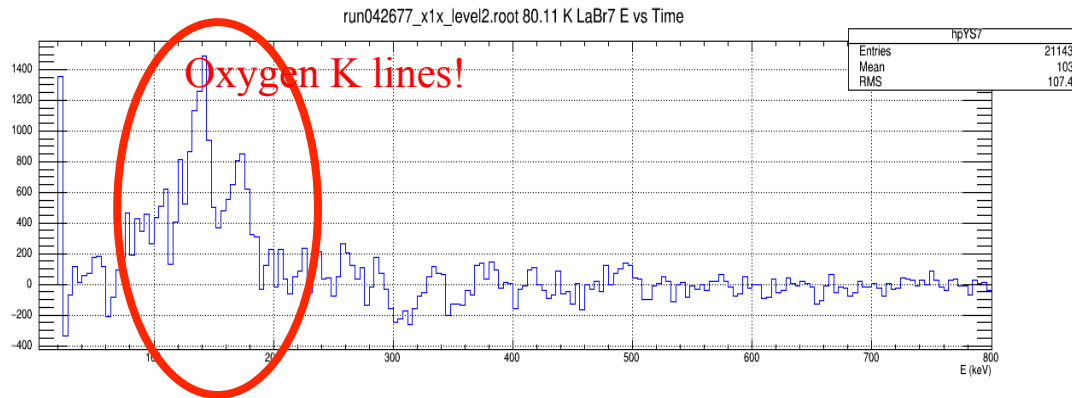
The plotted functions show a number of existing mu-p's in the time window of 500 ns, divided by the number of muon stops. The beginning of the time window corresponds to the moment of full thermalization. The time windows is approximately equal to the time of laser-field presence in the multipass cavity.

# from 2018 data old cryo target low pressure

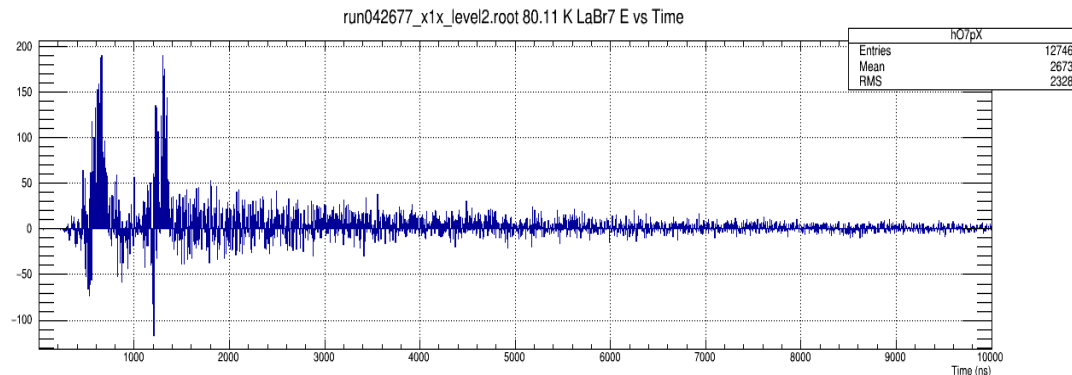
Energy spectra  
Blue O(3%) mixture  
Red pure hydrogen  
(3 bar 80 K)  
(normalized according  
to acq number of  
trigger, i.e. time)



LaBr re-calibrated

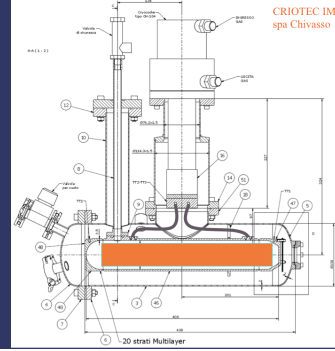


Oxygen mixture - hydrogen

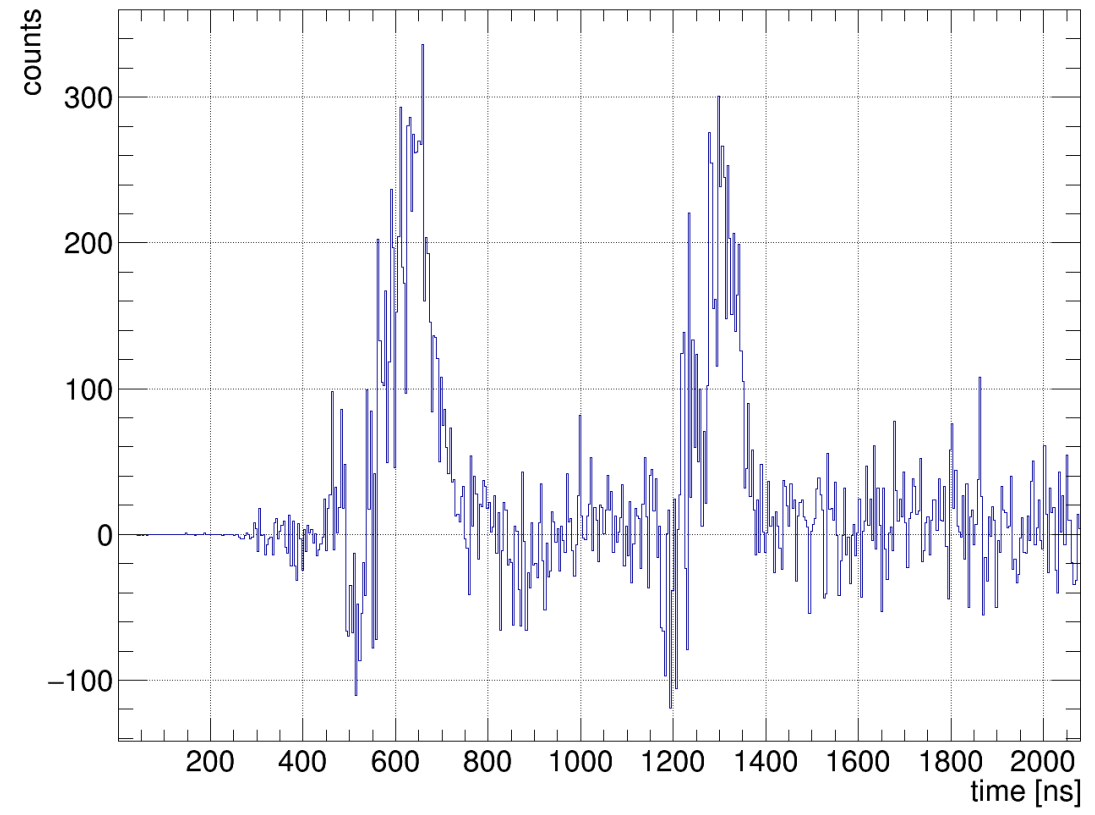


Time distribution of  
oxygen mixture –  
hydrogen in the  
energy range 110 –  
200 keV

x-ray time spectrum,  
for a low pressure and a fixed temperature around 80 K  
at 3 bar at 80 K with a mixture of hydrogen  
and oxygen 3% (by mass).

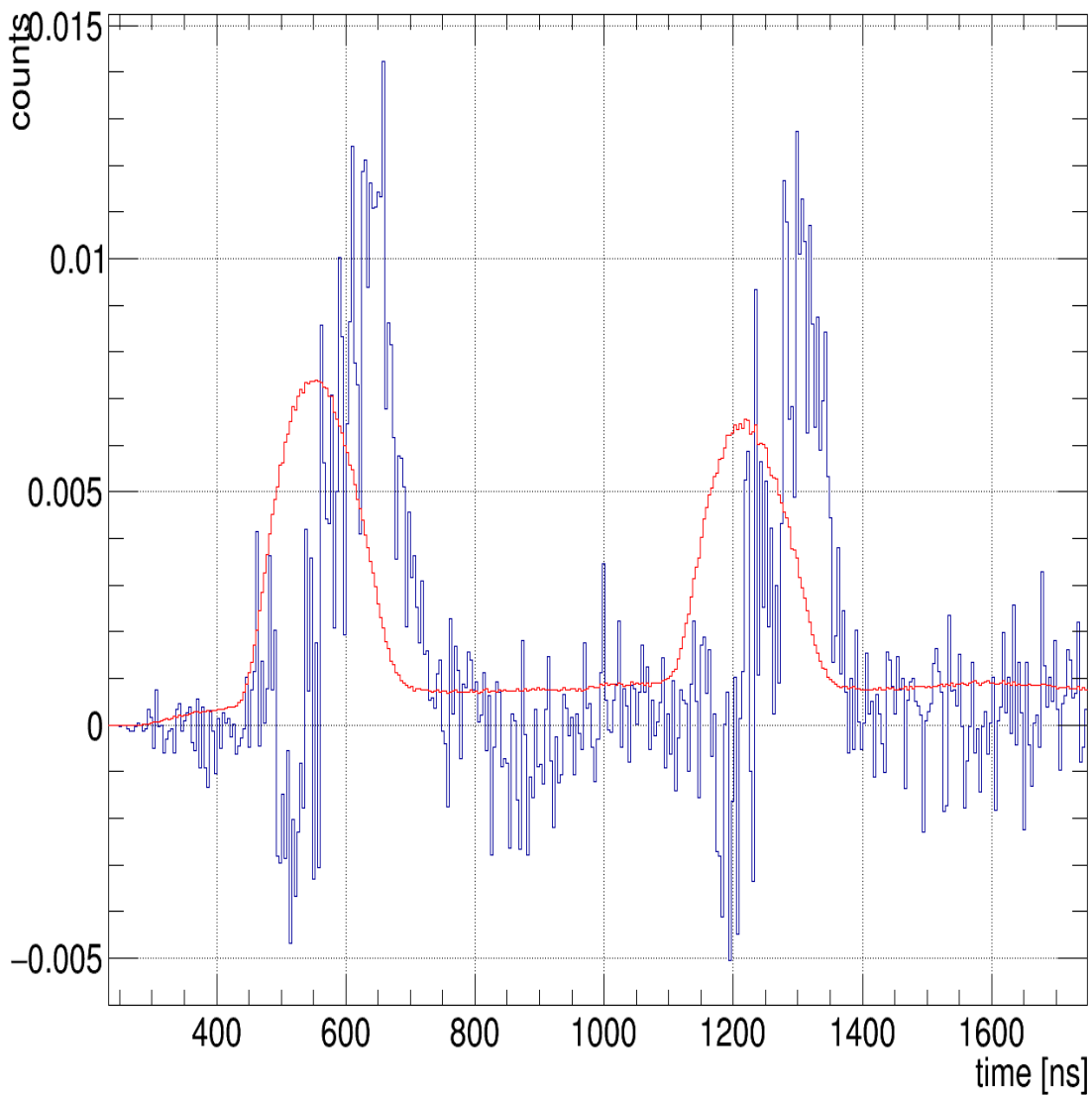


time evolution of oxygen  
lines signal obtained with the  
target used for the transfer  
rate Oxygen lines time  
evolution  
(very fast due to high oxygen  
concentration).



the y axis represents the number of counts with a data taking of 8 hours  
for two detectors. Delayed events disappear very quickly.



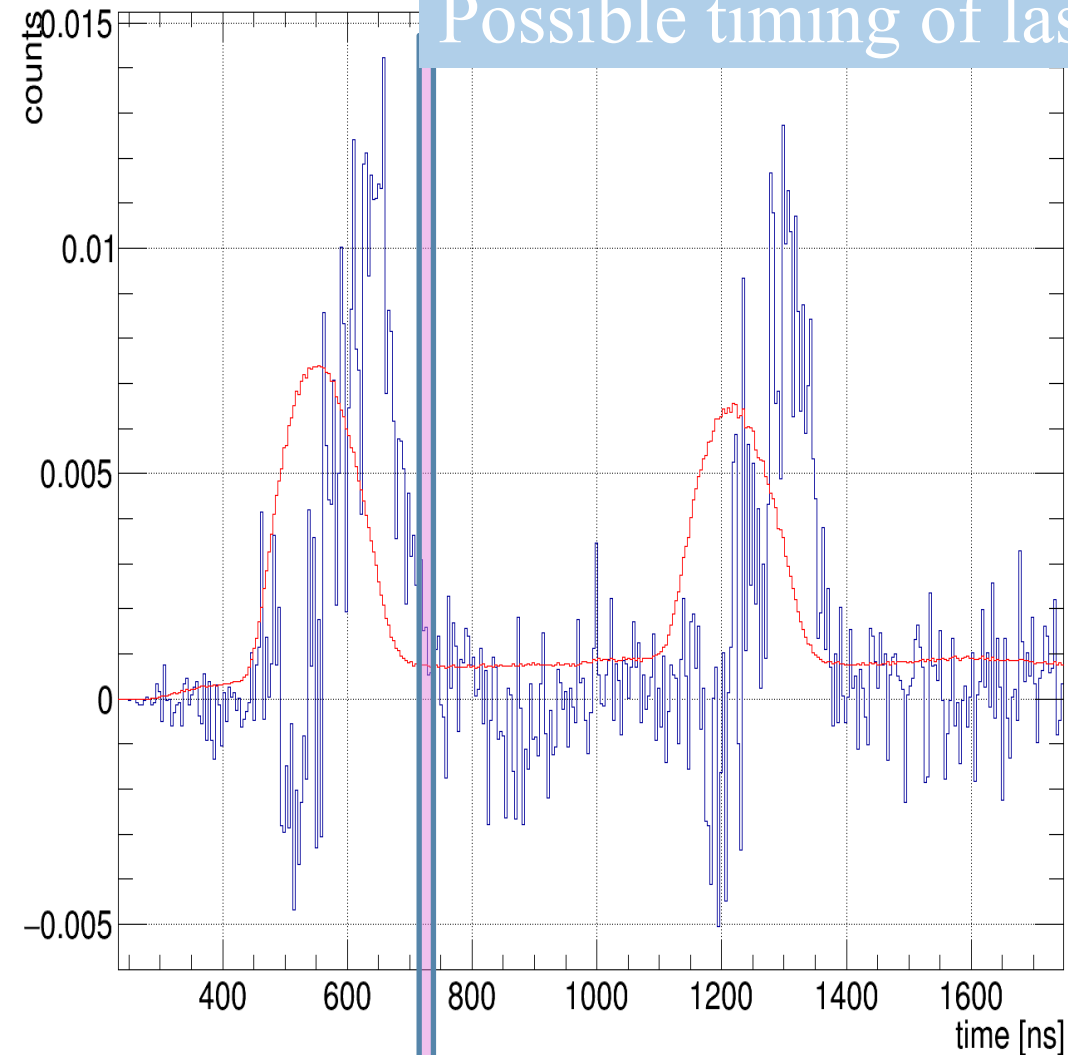


Oxygen lines time evolution  
(very fast due to high oxygen concentration).

Comparison with muon beam arrival time  
(actually prompt X-ray signal).

Oxygen signal is delayed but still overlapping the prompt signal.

## Possible timing of laser shot



Oxygen lines time evolution  
(very fast due to high oxygen concentration).

Comparison with muon beam arrival time  
(actually prompt X-ray signal).

Oxygen signal is delayed but still overlapping the prompt signal.

In order to have a signal/fluctuation  $> \sim 1$   
a factor 10 is needed in the acquisition time, in fact:

For  $2.2 \times 10^4$  counts:

Statistical fluctuations:  $\sqrt{2.2 \times 10^4} = 148$

Expected signal:  $2.2 \times 10^4 \times 0.008 = 176$

80 hours = 3.3 days (one frequency measurement).

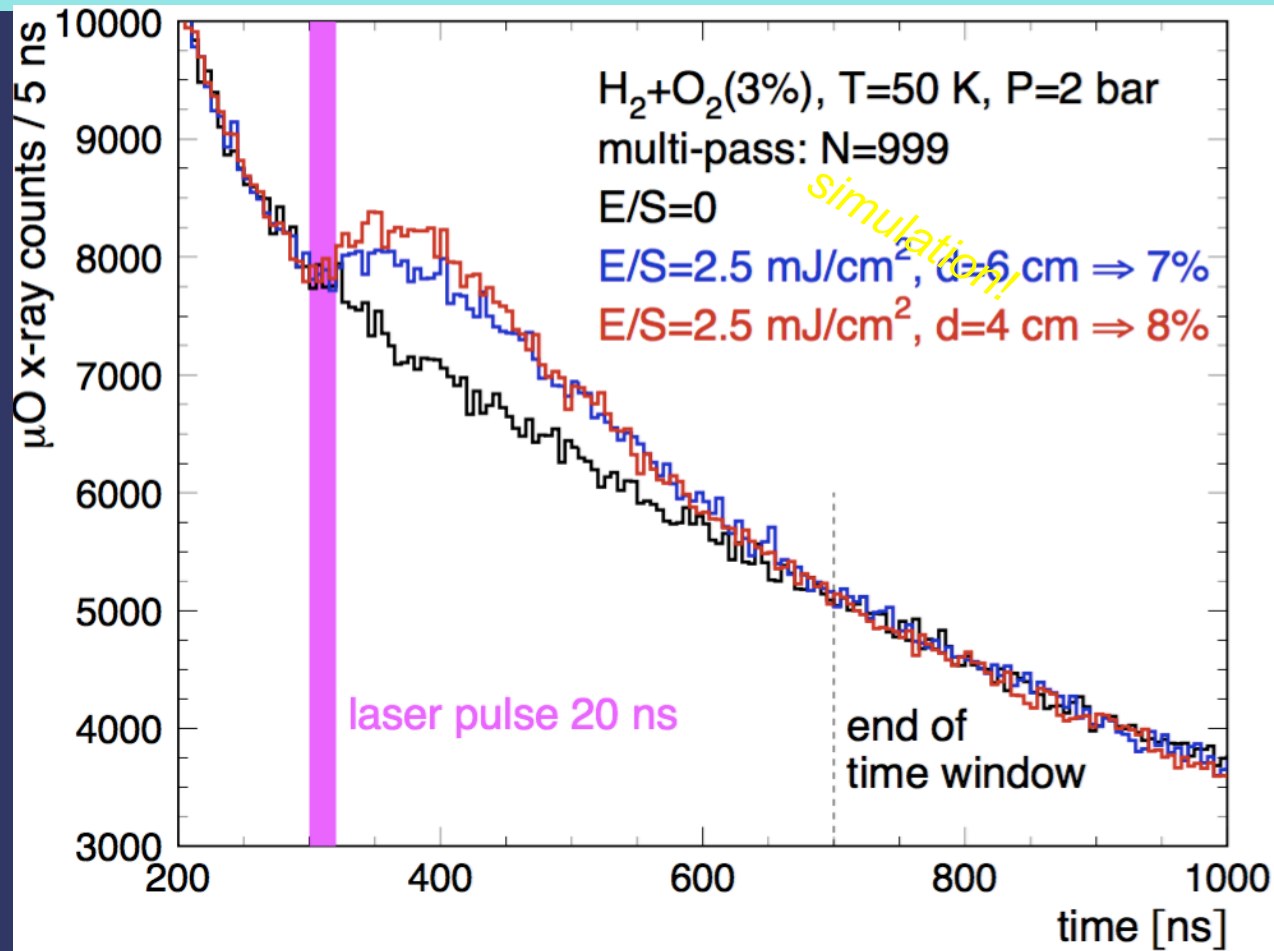
NB: possible optimization and (time) reduction factors:

- 1) Number of detectors (factor 2, 16 LaBr instead of 8)
- 2) Muon focalization (possible factor 2)
- 3) Software reconstruction (probably a factor 2, results presented with quick and dirty “quicklook” analysis)
- 4) laser can be at 8mJ  $\Rightarrow$  4% transition prob (1,6% Signal )
- 5) optimized target.
- 6) gas pressure and concentration

NB2: no systematics taken into account, no background measurement  
(working at 30/50 Hz but one of the pulses could be used to study the background),  
no new target materials and momentum.

# Study of best setup to maximize signal

- Shape and orientation of the optical cavity
- Characteristics of the cryo-target
- Pressure and oxygen concentration

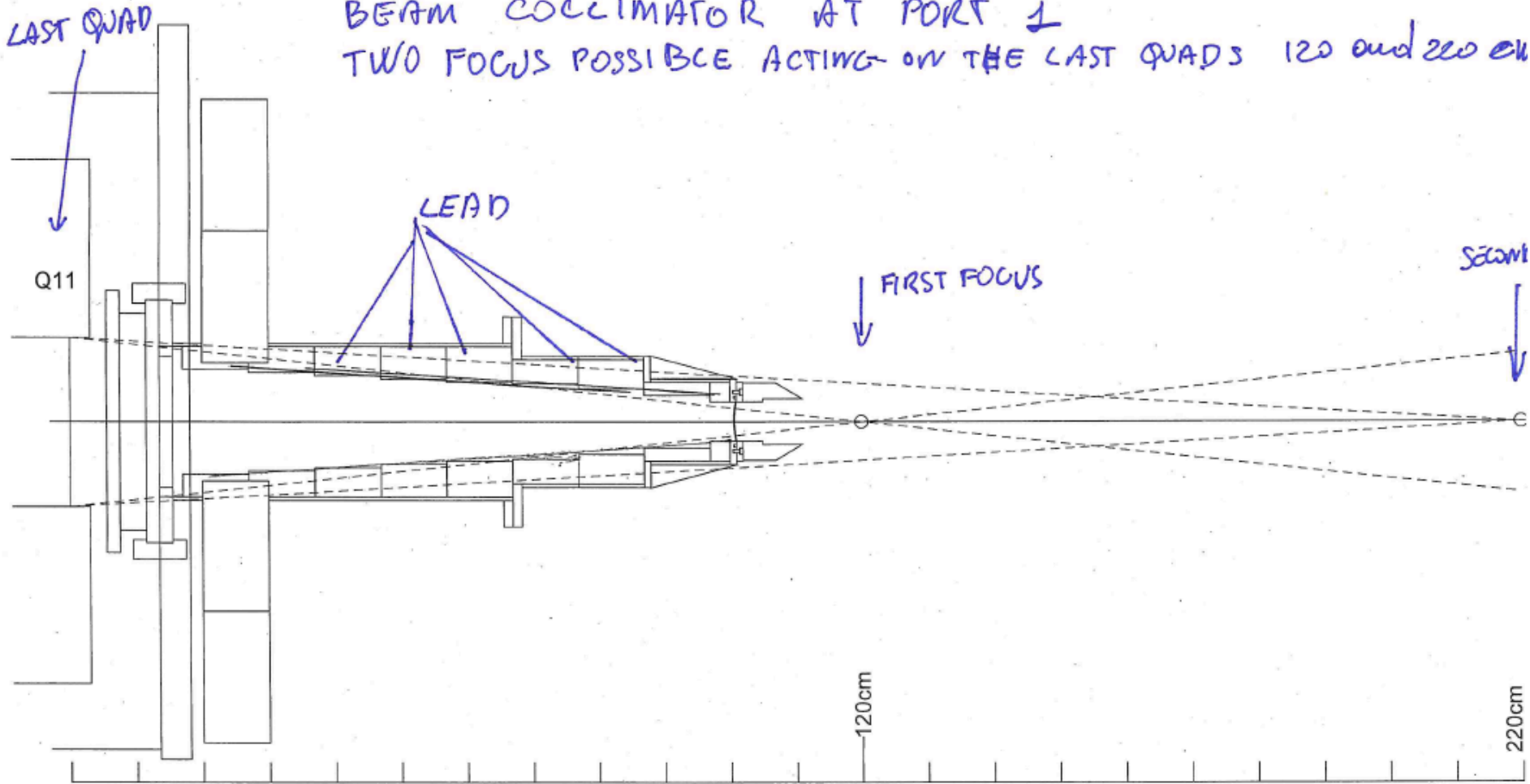


# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - pulsed high intensity muon beam
  - high energy MIR fine-tunable laser
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions



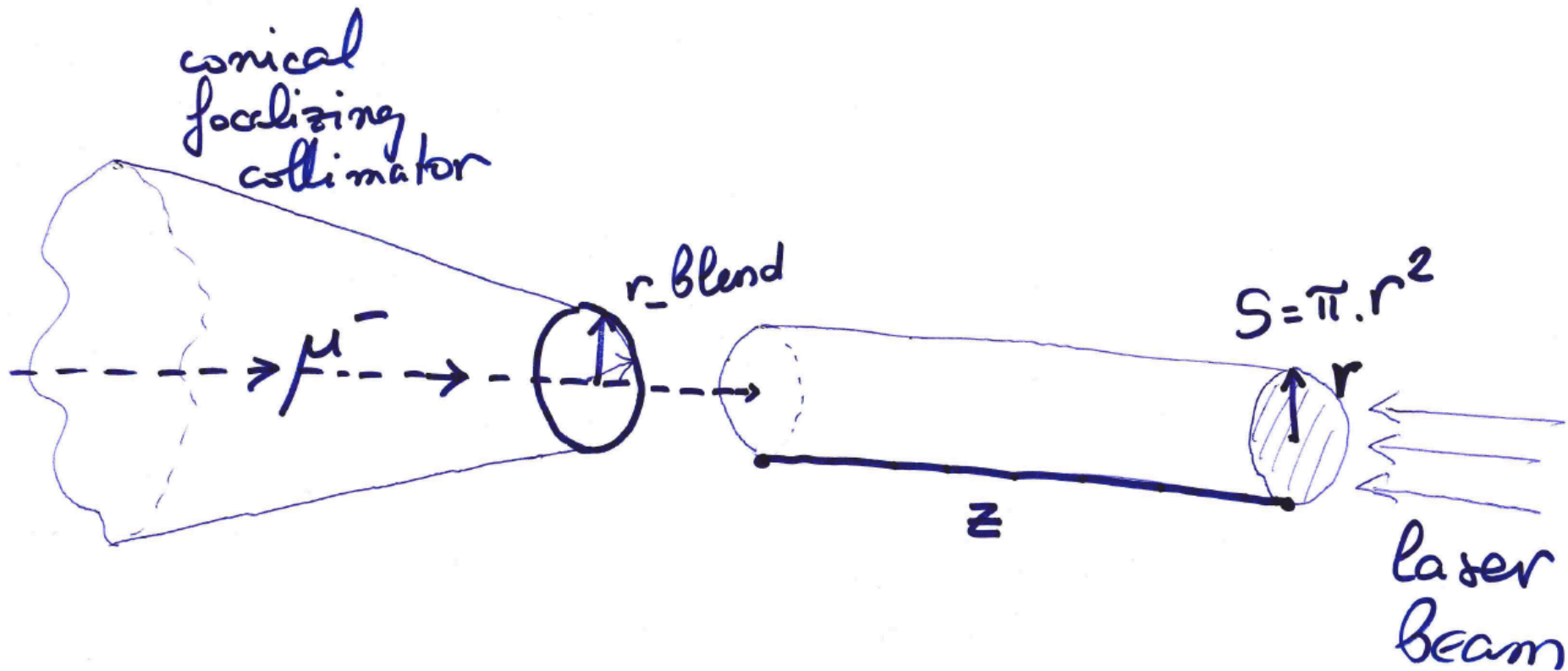
BEAM COLLIMATOR AT PORT 1  
 TWO FOCUS POSSIBLE ACTING ON THE LAST QUADS 120 and 220 cm



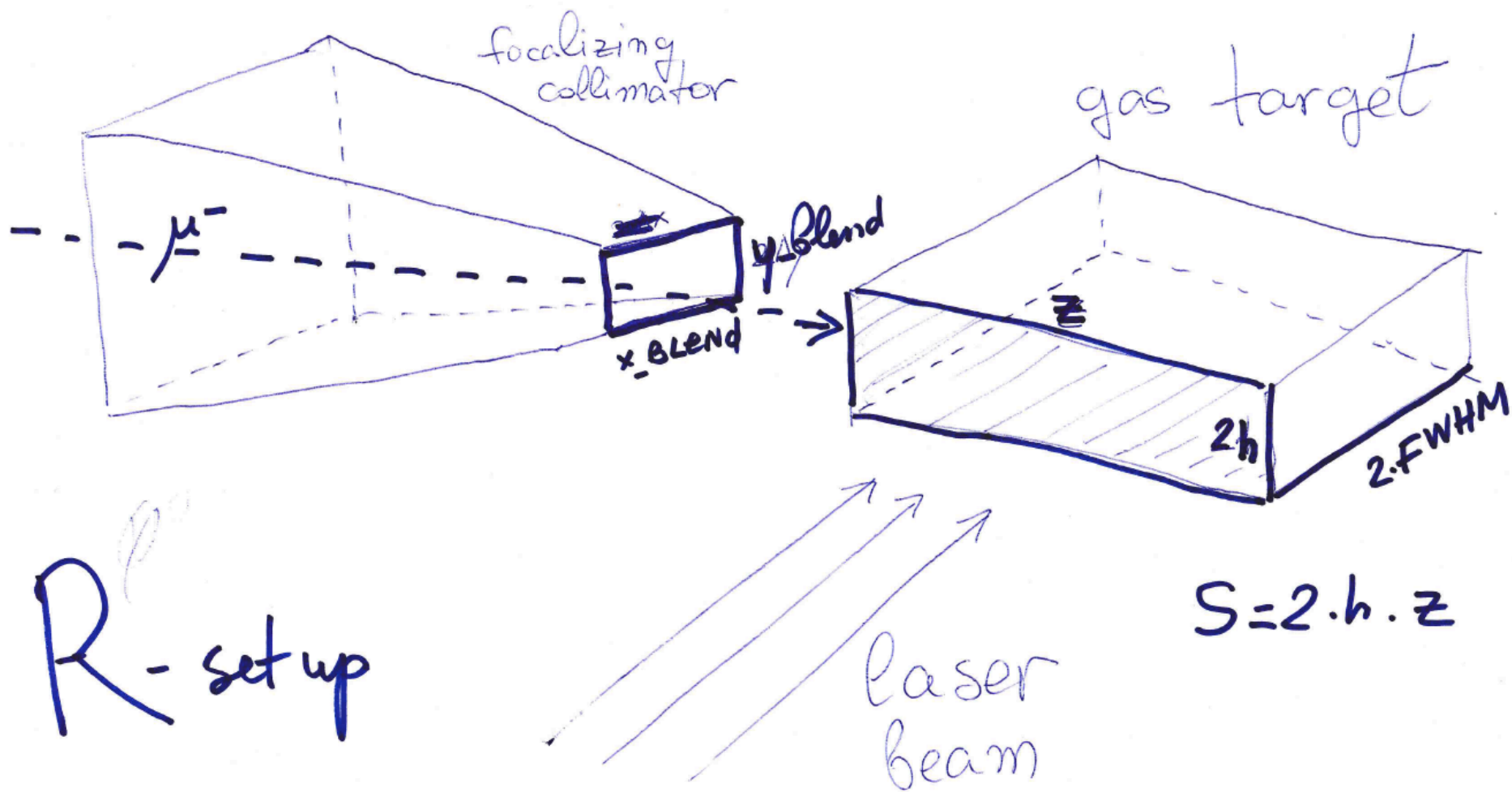
FROM THE  
 PART TO THE  
 RIGHT ONE CAN  
 EXTRACT THE  
 LEAD

COLLIMATOR  
 THE PART TO THE  
 LEFT CAN BE  
 DISMOUNTED

THE IDEA IS TO  
 SUBSTITUTE THE COLLIMATOR  
 WITH A FOCUSING STRUCTURE



C - setup

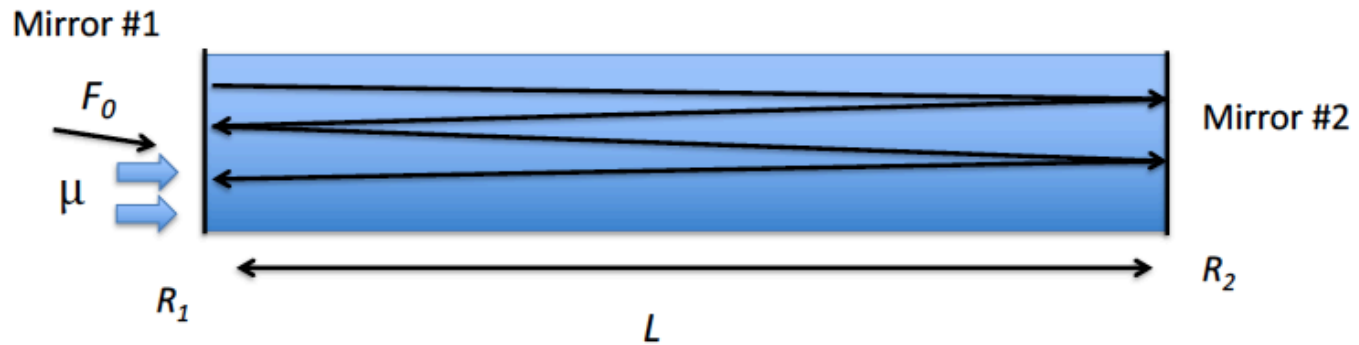


# Multipass Optical Cavity

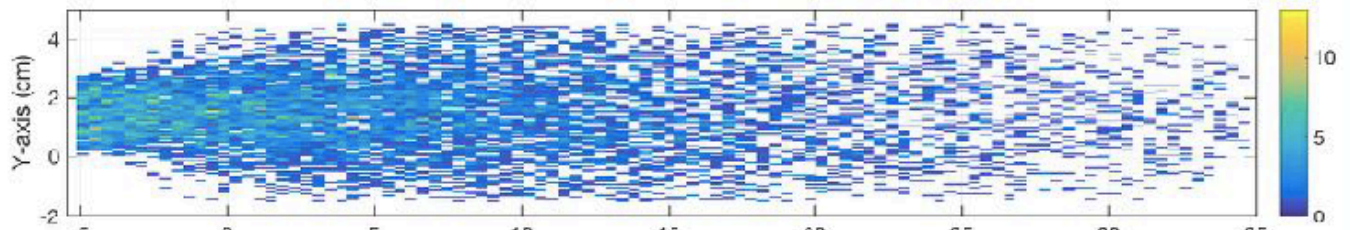
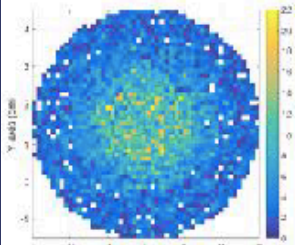
Luigi Moretti, Livio Gianfrani

9497 events

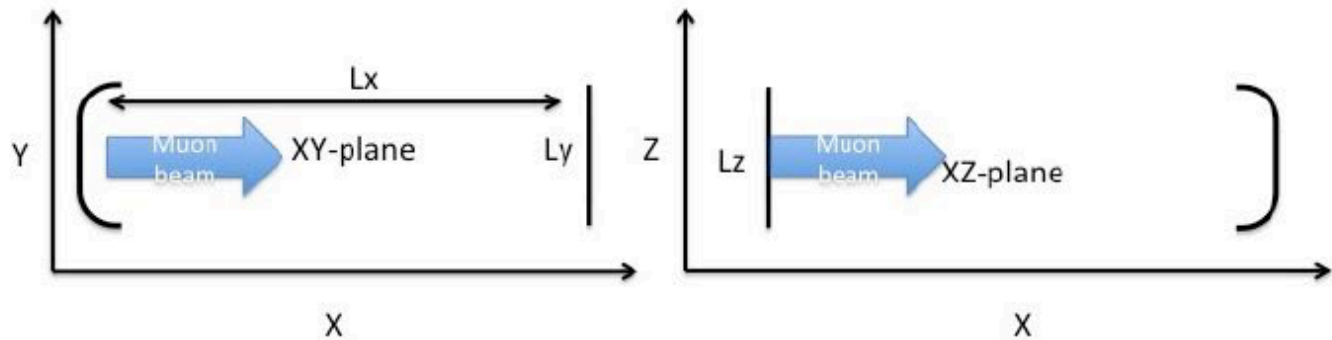
$$V_{\text{est}} = \pi (2 \text{ cm})^2 (15 \text{ cm}) = 94 \text{ cm}^3$$



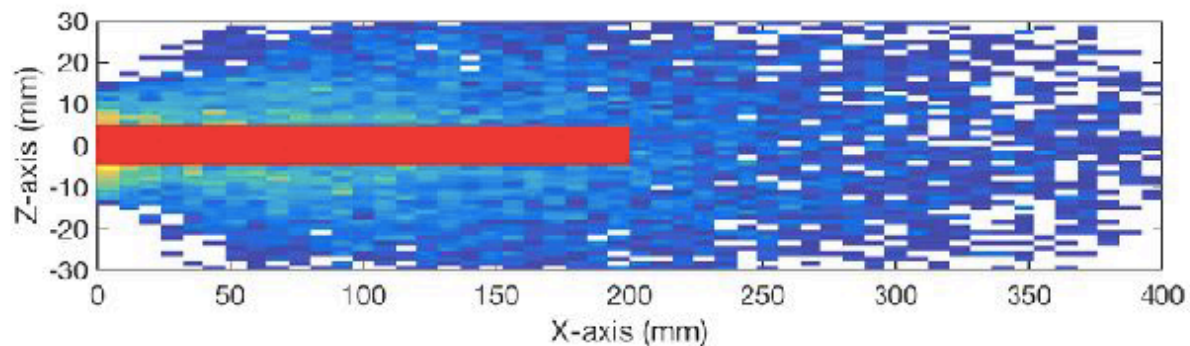
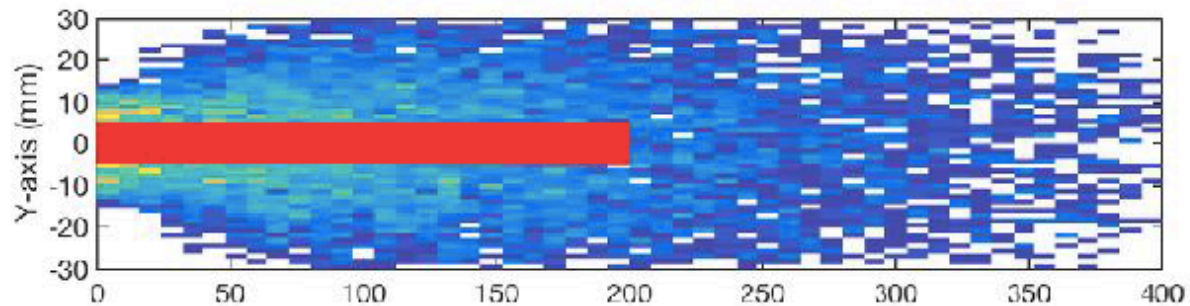
$$L_{\text{eff}} = \frac{L}{\alpha} = \frac{20 \text{ cm}}{12 \times 10^{-4}} \approx 166 \text{ m}$$



# Optical design of cavity



$$L_y = L_z = 2.5 \text{ cm}$$
$$L_x = 20 \text{ cm}$$



# Cavity enhancement effect at glance

$$E_l = 2.5 \text{ mJ}$$

$$N_R = 700$$

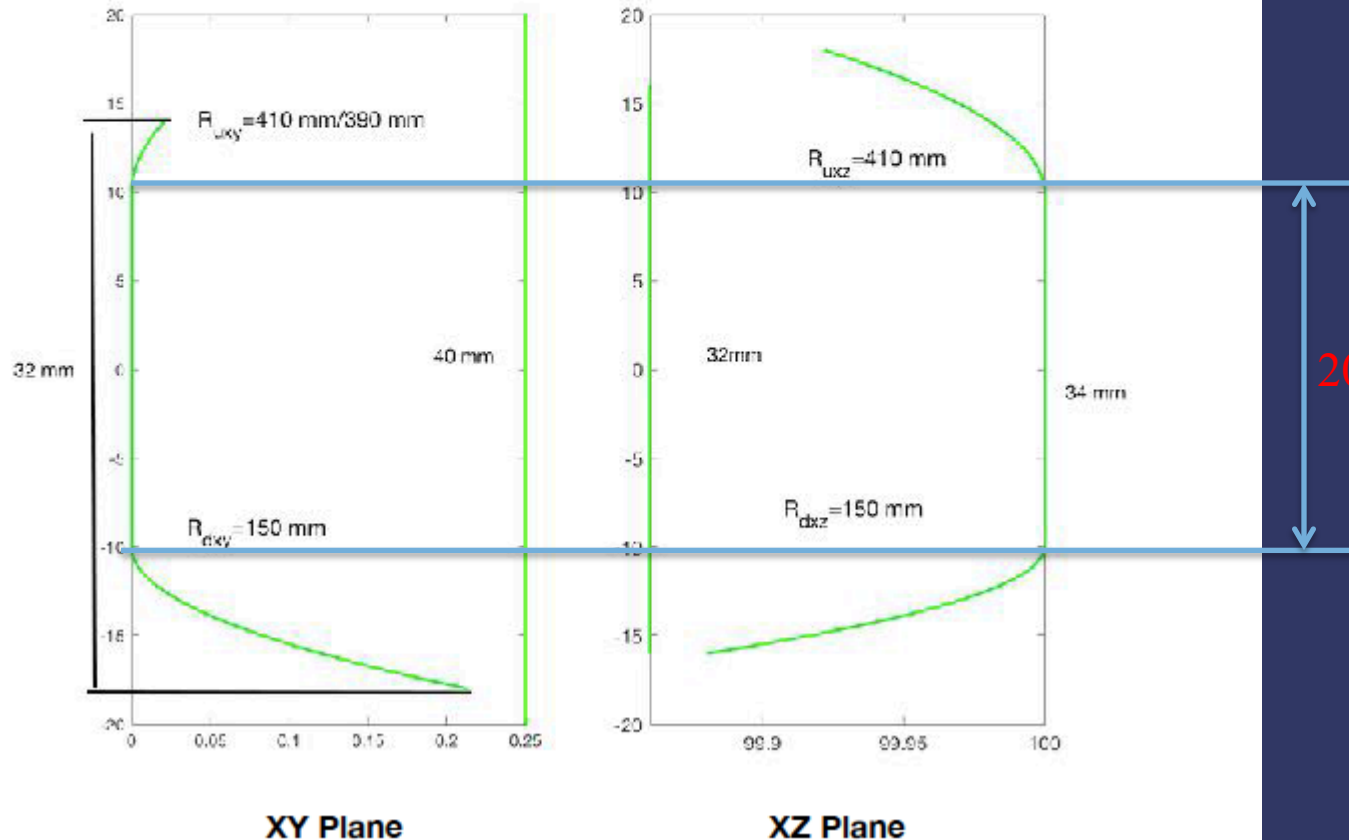
$$R_1 = R_2 = 0.9989$$

$$\text{New design } S_{ill} \approx (2 \cdot 2) \text{ cm}^2 \quad (\alpha = 12 \times 10^{-4}) \quad \left. \vphantom{S_{ill}} \right\} \rightarrow D_{cav} = \frac{N_R E_l}{S_{ill}} = 438 \frac{\text{mJ}}{\text{cm}^2}$$

$$\bar{P} = \frac{D_{cav}}{D_{sat}} \approx 0.01$$

# New geometry of cavity-2

HR mirrors



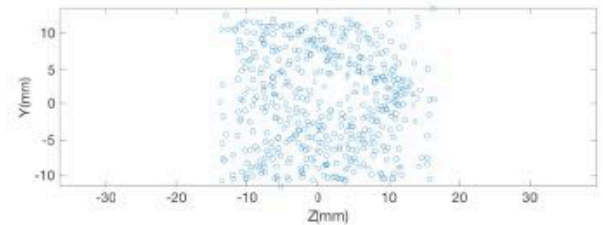
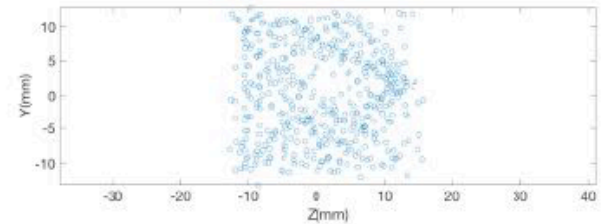
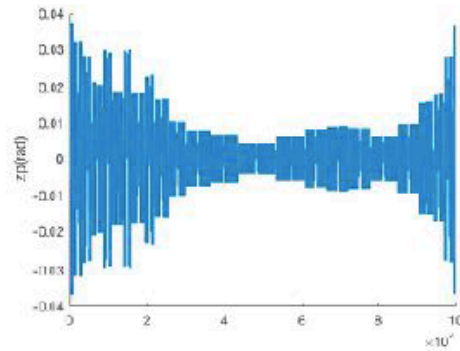
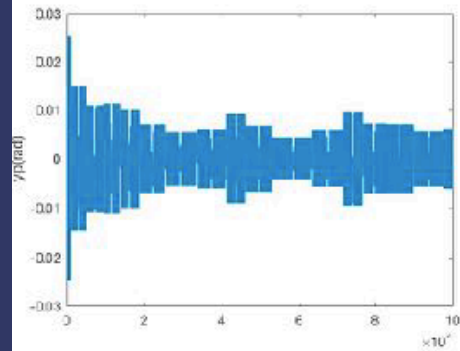
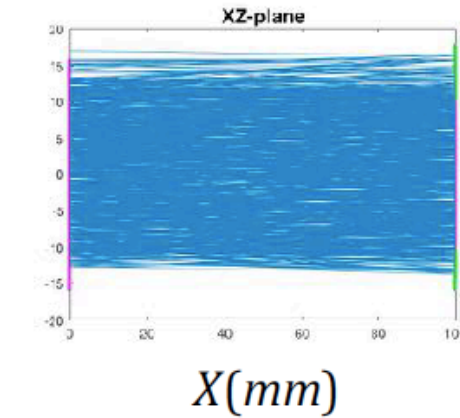
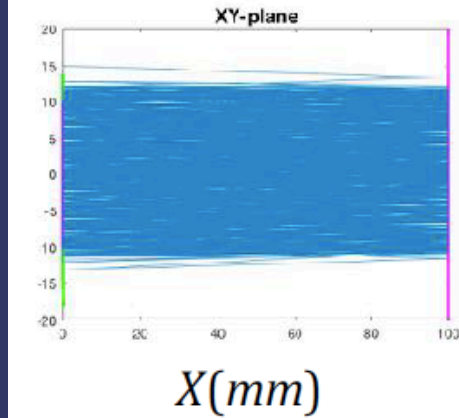
$A = 3,14 \text{ cm}^2$

Substrate Fused Silica;

# New geometry of cavity-2

$$N_{rif} = 1000$$

$$\alpha_{xy} = 15\text{mrad} \quad \alpha_{xz} = 6\text{mrad}$$



$$S_{ill} \approx (2.3 \times 2.5)\text{cm}^2$$

After about 1000 reflection the ray light escape from the cavity



# Mirror Fabrication

- **Substrate** : 1) Fused Silica; 2)Stainless steel
- **HR coating**: 1)Semiconductor multilyer Ge/ZnS; 2) Metallic

Laser Components Company with Optoprim as Italian reference

- Test for the Laser damage;
- Test for the low temperature.

# New target simulations:

TARGET 2016

vacuum window: 0.8 mm Al  
pressure vessel window: 2.84 mm Al  
with hodoscope (1mm fibers)  
gas: ~cylinder, 6 cm  $\varnothing$  40 cm length  
with Ni (100 microns) + Au (10 microns) coating  
with multi-layer insulators in front, on sides, on the bottom  
lead collimator: wall with hole 3 cm  $\varnothing$

**40 BAR @ 300 K**

**2 BAR @ 80 K**

TARGET 2018

vacuum window: 1 mm Kapton  
pressure vessel window: 1.5 mm fused Silica  
no hodoscope  
gas: cylinder 2 cm  $\varnothing$  15 cm length  
no coating  
with multi-layer insulators in front (same of 2016)  
lead collimator: wall with hole 2 cm  $\varnothing$

**2 BAR @ 80 K**

# New target simulations:

@ 300 K

TARGET 2016

vacuum window  
pressure vent  
with hodoscop  
gas: ~cylinder  
with Ni  
with multi-layer  
lead collimator

low pressure :

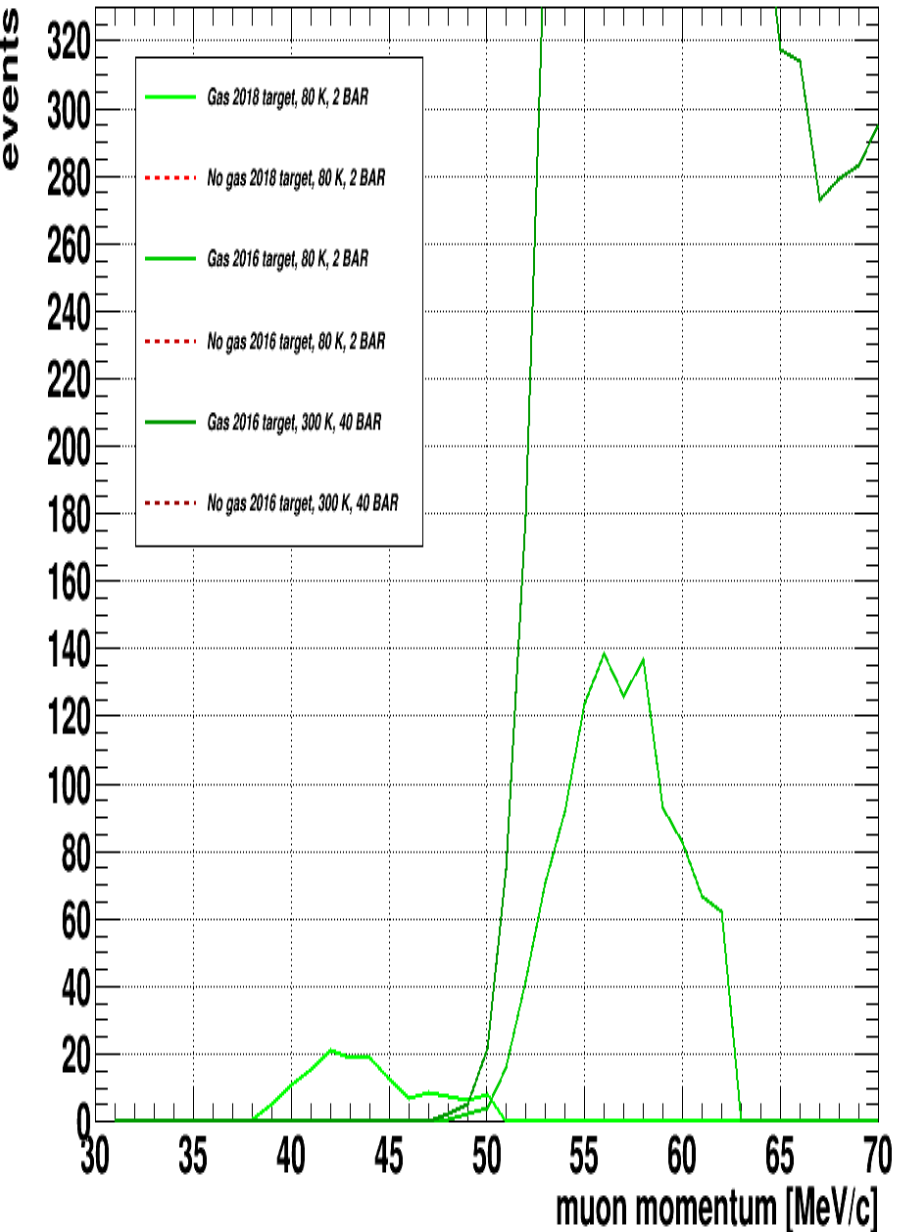
- less stopped muons
- lower momentum = less muons in beam
- & small optical cavity

Need careful optimization  
of all elements

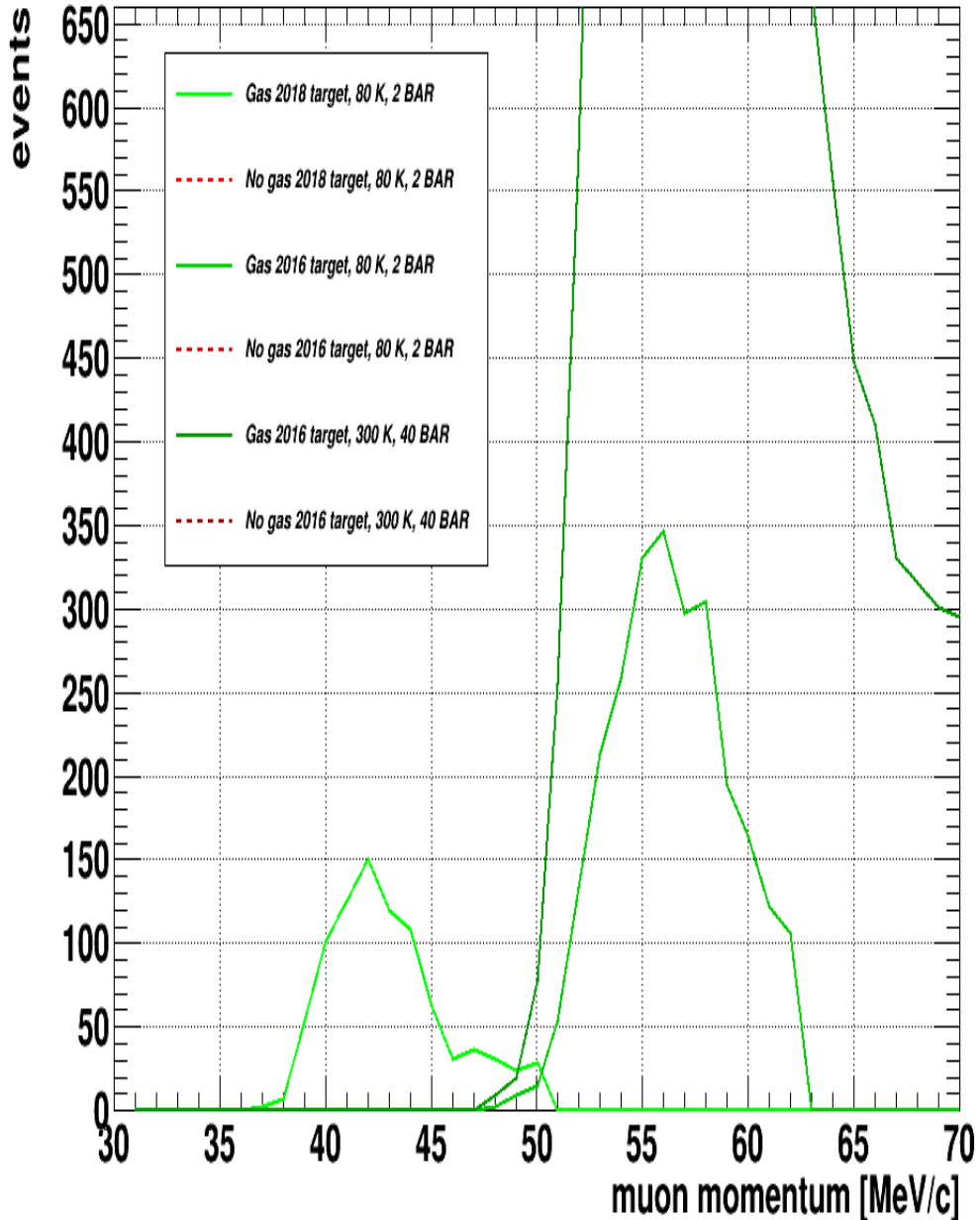
TARGET 2018

vacuum window  
pressure vent  
no hodoscop  
gas: cylinder 2  
no coating  
with multi-layer insulator  
lead collimator: wall with

### Stop in gas weighted with muons in the beam as function of p



### Stop in gas assuming constant number of muons as function of p



# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - high energy MIR fine-tunable laser
  - pulsed high intensity muon beam
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

## Tunable pulsed IR laser at $\lambda=6.8\mu$

Direct difference frequency generation  
in non-oxide non linear crystals using  
single-mode Nd:YAG laser and tunable Cr:forsterite laser

Wavelength:  $\lambda = 6785$  nm                      44.22 THz

Line width:  $\Delta\lambda = 0.07$  nm                      450 MHz

Tunability range: 6785 +/- 10 nm                      130 GHz

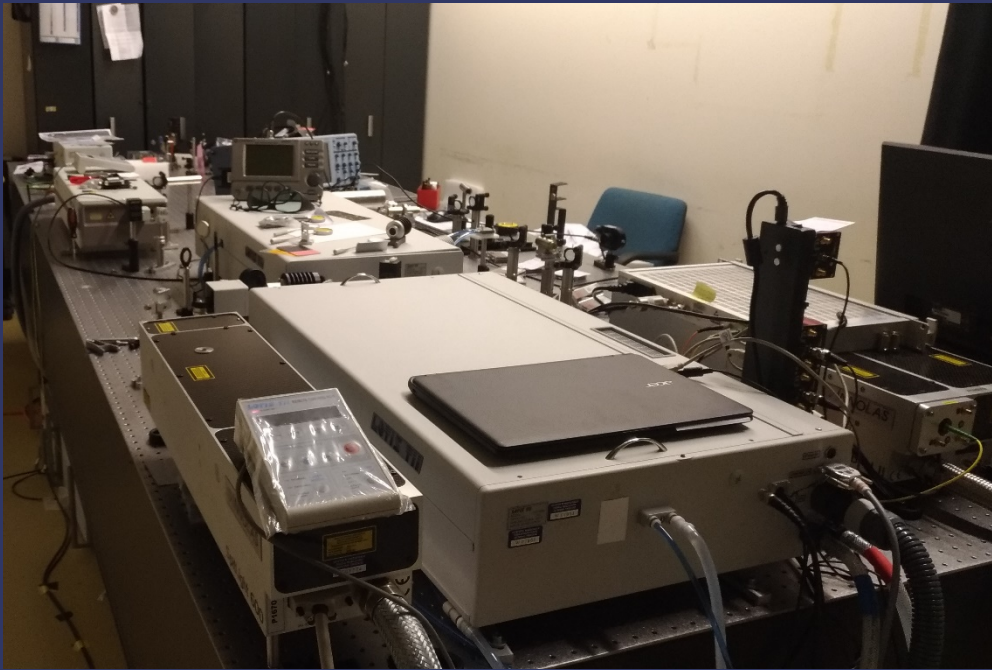
Tunability step                      = 0.007nm                      45 MHz

Repetition rate: 25 Hz

Pulse Energy at 6780 nm: > 1 mJ

# The lab at the moment

- Available - All lasers
- Available - Most optics and electronics
- Available - Most test and measurement equipment

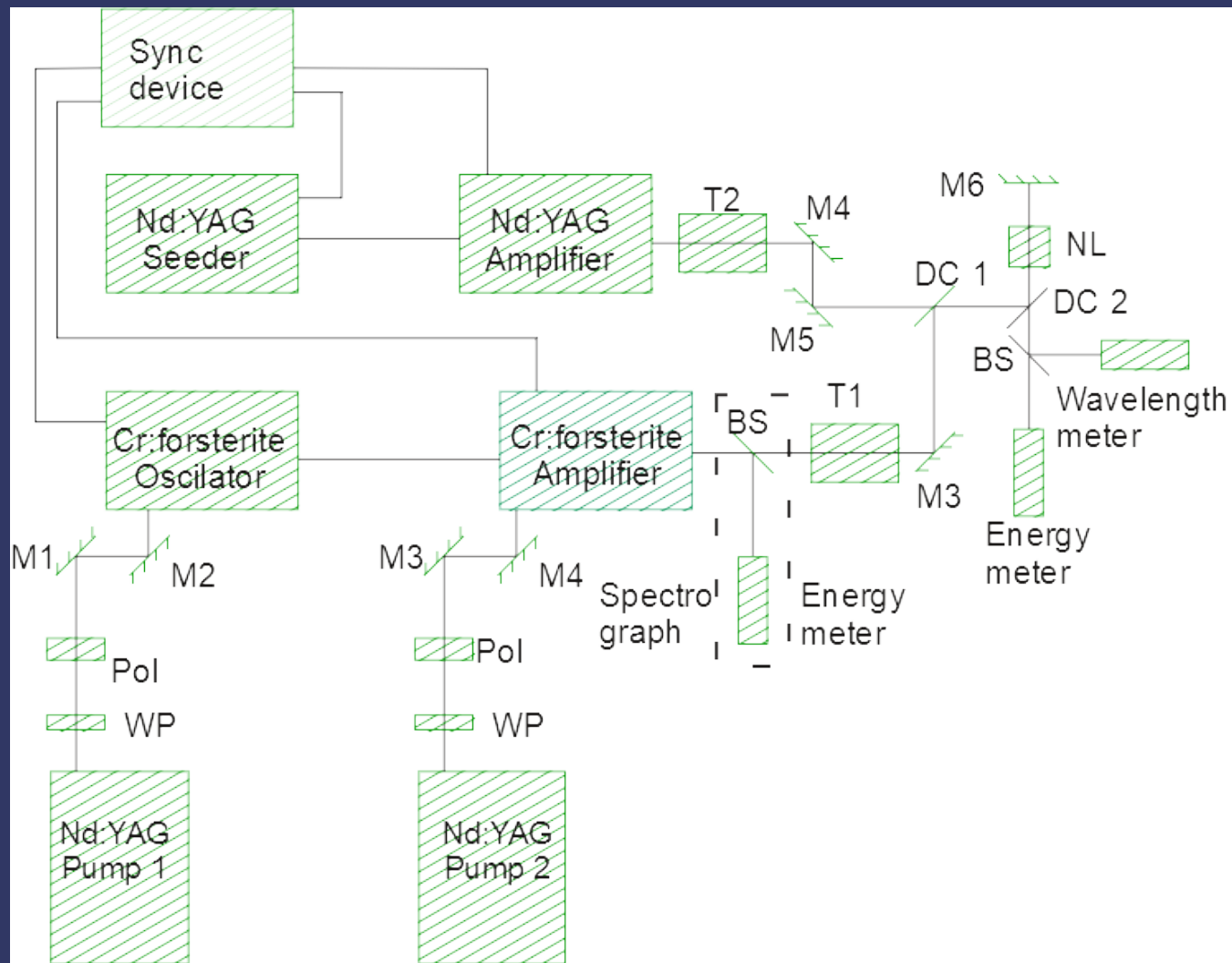


# Final scheme of the DFG based laser system

The Nd:YAG will be at "fixed" wavelength 1064.14nm with linewidth max - 0.34pm (90MHz) and min - 0.11pm (30MHz).

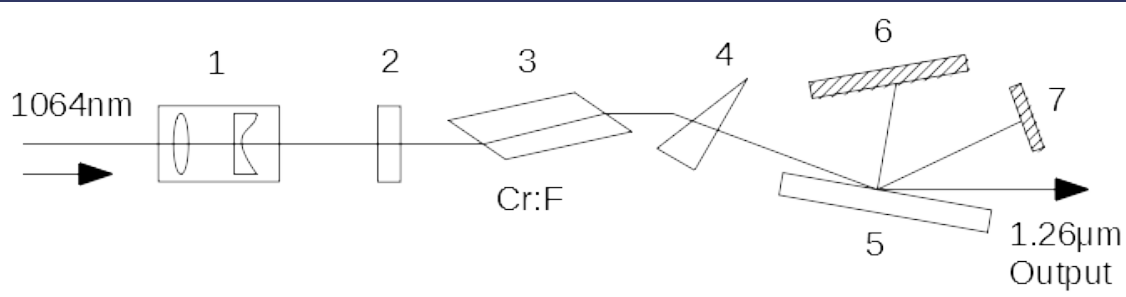
The Cr:forsterite will have linewidth max - 1pm (188MHz) and min - 0.5pm (90MHz).

The Cr:forsterite will be tunable from 1252nm to 1272 nm which corresponds to tunability from 6500nm to 7090nm, which is 3765GHz. The required tunability 6760nm  $\pm$ 3nm corresponds to tunability range  $\sim$  39GHz.

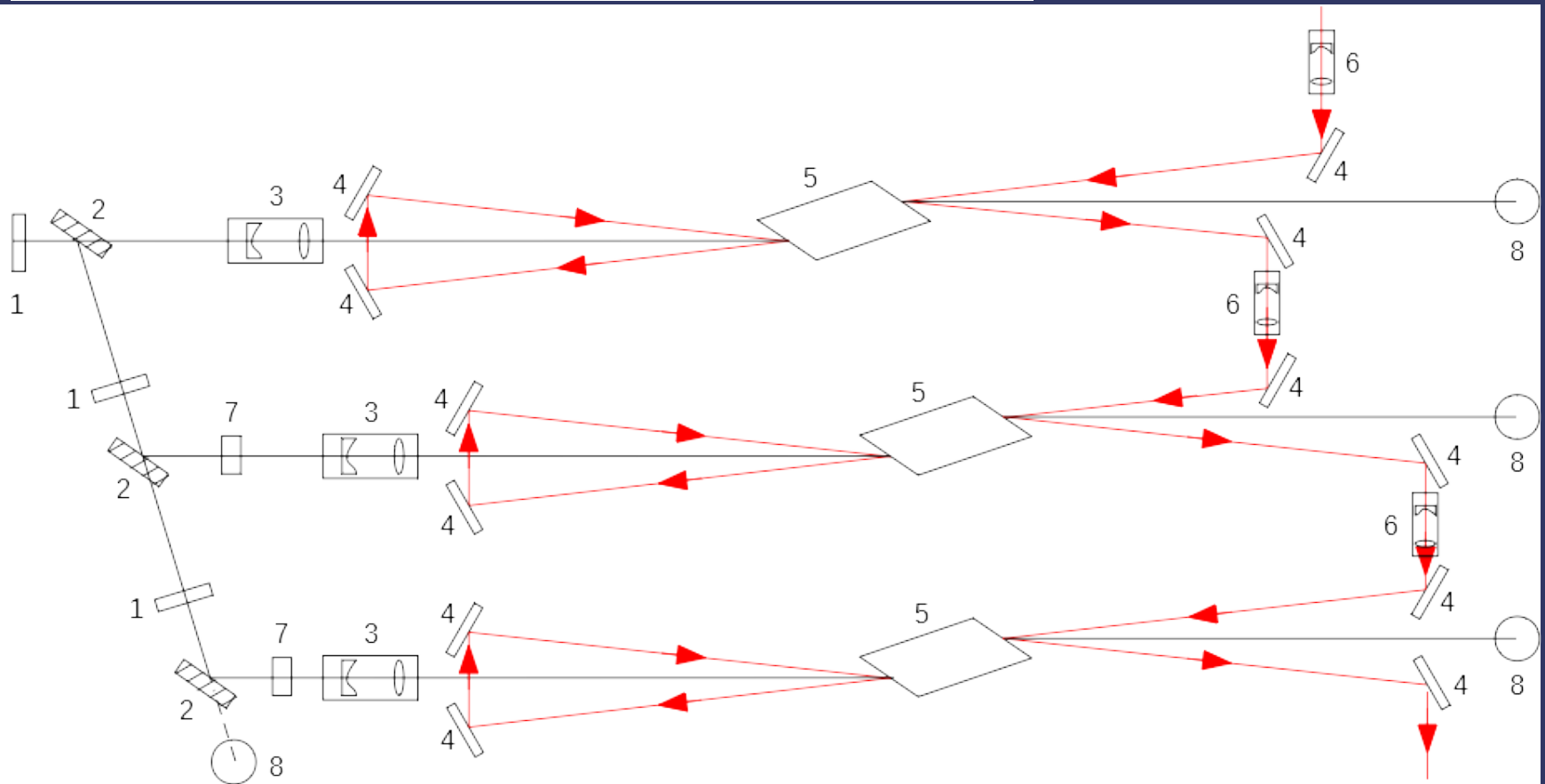


WP - waveplate, Po - polarizer, M1-M5 - mirrors, T1 and T2 - telescopes, BS - beamsplitters, DC1 - dichroic mirror (reflecting 1.26 $\mu$ m, transmitting 1.06 $\mu$ m), DC2 - dichroic mirror (reflecting 1.06 and 1.26  $\mu$ m, transmitting 6.76 $\mu$ m)





# Cr:Forsterite oscillator & amplifier stages setups



1 – half-wave plates ( $\lambda/2$ , 1064 nm); 2 – polarizers (1064 nm); 3 – decreasing telescopes (1064 nm); 4 – turning mirrors (1262 nm); 5 – Cr:Forsterite crystals; 6 – increasing telescopes (1262 nm); 7 - rotators (90°, 1064 nm); 8 – beam stops

# Available NL crystals & Expected output energies at 6760 – 6780 nm

## Nonlinear crystals

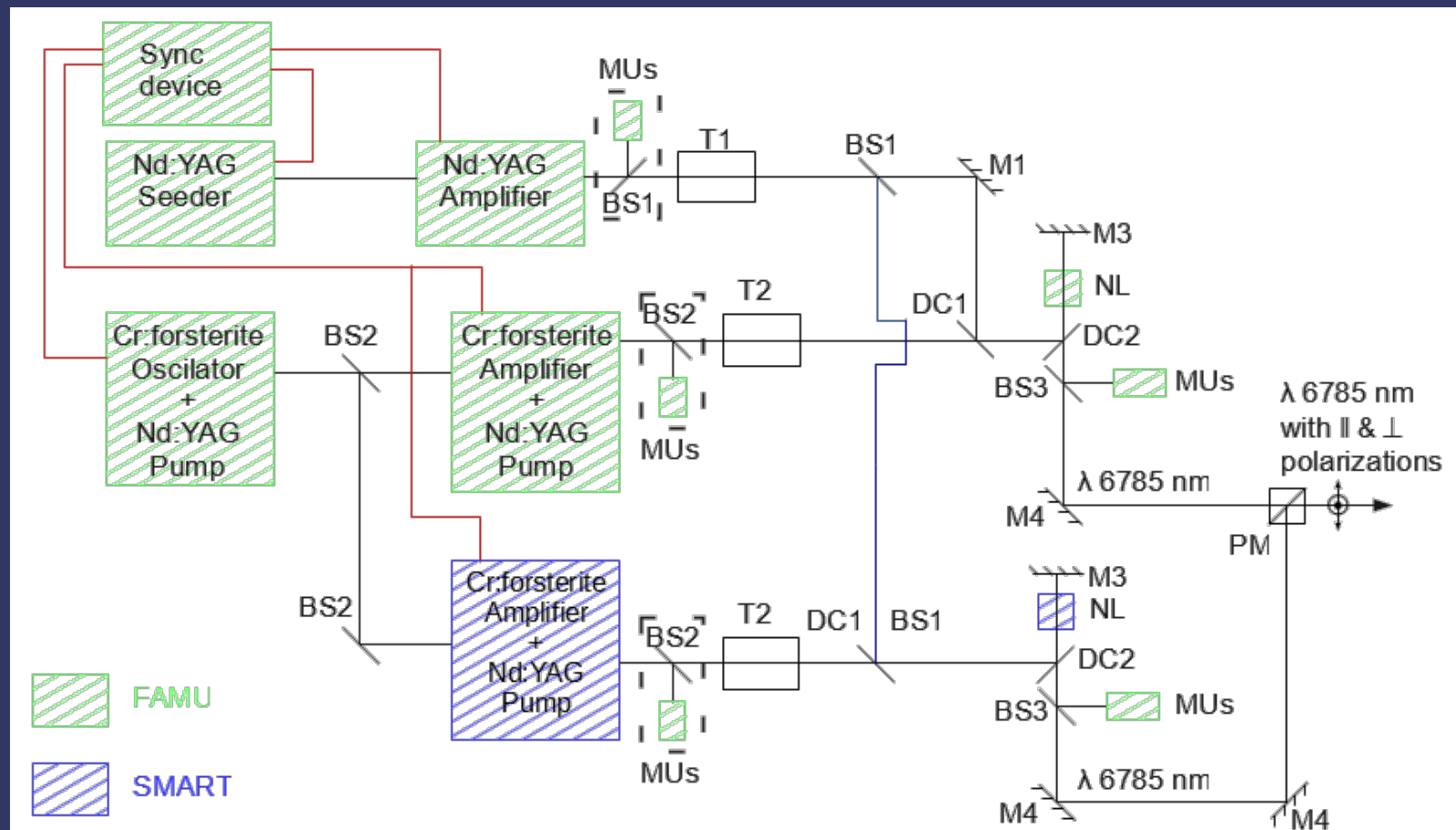
Available	Ordered
LiInS <sub>2</sub> - 7x7x20 mm	LiInSe <sub>2</sub> - 7x7x15 mm
LiInS <sub>2</sub> - 8x8x18	LGS – 5x5x4 mm
	BaGa <sub>4</sub> S <sub>7</sub> – in progress

Expected energies:

LiInS<sub>2</sub> & LiInSe<sub>2</sub>: 1 – 1.5 mJ

LiGaS<sub>2</sub> & BaGa<sub>4</sub>S<sub>7</sub> ~ 2mJ

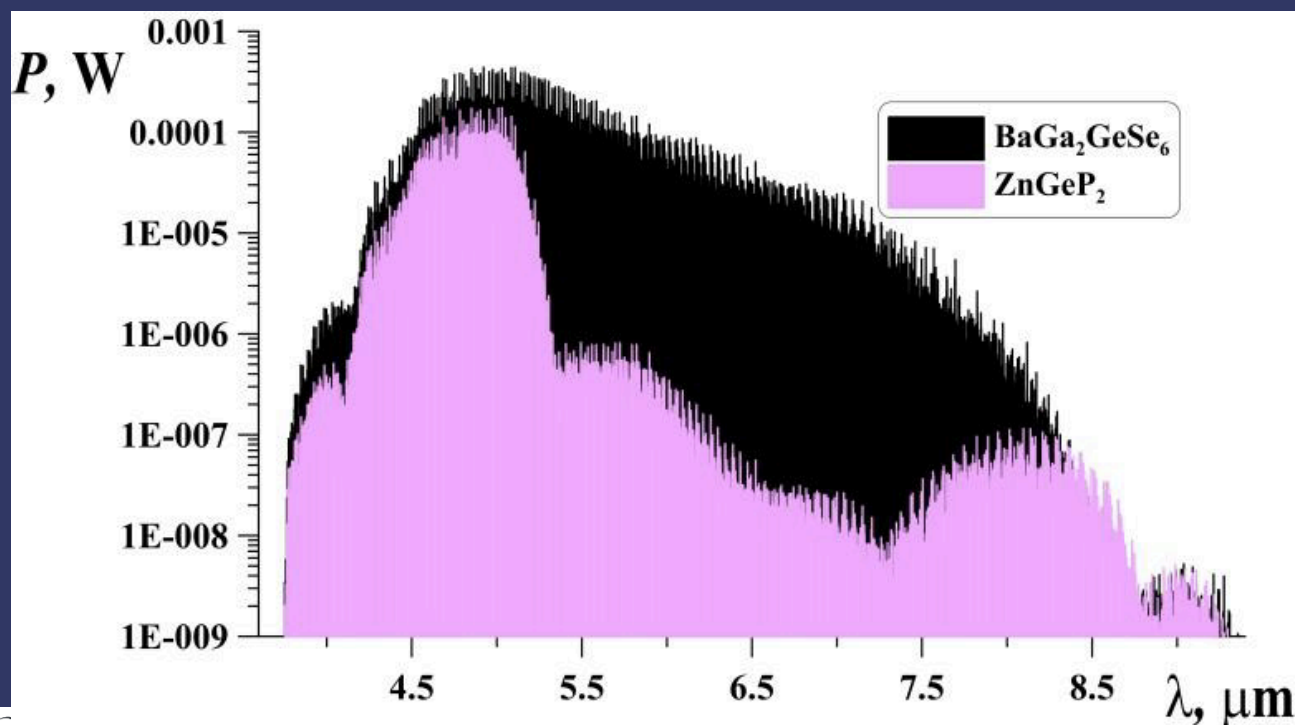
# The scheme of the laser system as proposed for the SMART project – PRIN call



M1 – mirror HR 1064 nm, M2 – mirror HR 1262 nm, M3 – mirrors HR 1064&1262&6785 nm,  
 M4 – mirror HR 6785 nm, T1 and T2 - telescopes, BS1 – beamsplitters/beamsampler 1064  
 nm, BS2 – beamsplitters/beamsampler 1262 nm, BS3 – beamsampler 6785 nm,  
 DC1 - dichroic mirror (reflecting 1064 nm, transmitting 1262 nm), DC2 - dichroic mirror  
 (reflecting 1064 nm and 1262 nm, transmitting 6785 nm), NL – nonlinear crystals, MUs –  
 measuring units:  $\lambda$  meters, energy meters, PM – polarization mixer

Result	Time period
CrFo oscillator linewidth and wavelength shot-to-shot	End of May
Efficiency of cooling the Cr:forsterite amplifiers' stage	MidJune
Damage threshold of the available nonlinear crystals (AgGaS <sub>2</sub> , LiInS <sub>2</sub> , LiInSe <sub>2</sub> , LGS, ?BaGa <sub>4</sub> S <sub>7</sub> )	Mid June
Cr:forsterite oscillator-amplifier stable single mode operation and linewidth	End of June
Nonlinear efficiencies of the available nonlinear crystals (AgGaS <sub>2</sub> , LiInS <sub>2</sub> , LiInSe <sub>2</sub> , LGS, ?BaGa <sub>4</sub> S <sub>7</sub> , BGGS ),	End of June
Test of wavelength meter shot to shot mode with real DFG pulses	Mid July
Preliminary parameters of the 6785 nm emission (~ 500 μJ): energy, pulse widths, linewidth, divergence, M2	End of July
Maximum (achievable) energies at 6785 nm	Mid August

The DFG spectra for both crystals cover the spectral interval from 4 to 9  $\mu\text{m}$  quite densely. The peak power of DFG radiation in BGGSe crystal (up to 0.35 mW) is higher than that of ZnGeP<sub>2</sub> crystal (up to 0.16 mW) in the whole considered spectral range. Hence, we can come to the conclusion that under considered conditions the conversion efficiency of second stage in BGGSe crystal is at least 2 times higher than that of ZnGeP<sub>2</sub> crystal. Thus, for the broadband two-stage frequency conversion of multi-line CO laser radiation, the new BGGSe crystal is twice more efficient than "standard" nonlinear ZnGeP<sub>2</sub> crystal.



# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - high energy MIR fine-tunable laser
  - pulsed high intensity muon beam
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

# FAMU Key elements

## Muon Beam at RIKEN-RAL

### Beam properties

surface  $\mu^+$  (20~30MeV/c) and  
decay  $\mu^+ / \mu^-$  (20~120MeV/c)

typical beam size 10cm<sup>2</sup>

$\bar{x}\Delta p/p$  FWHM 10%(decay), 5%(surface)

Double pulse structure

(Choice of single pulse

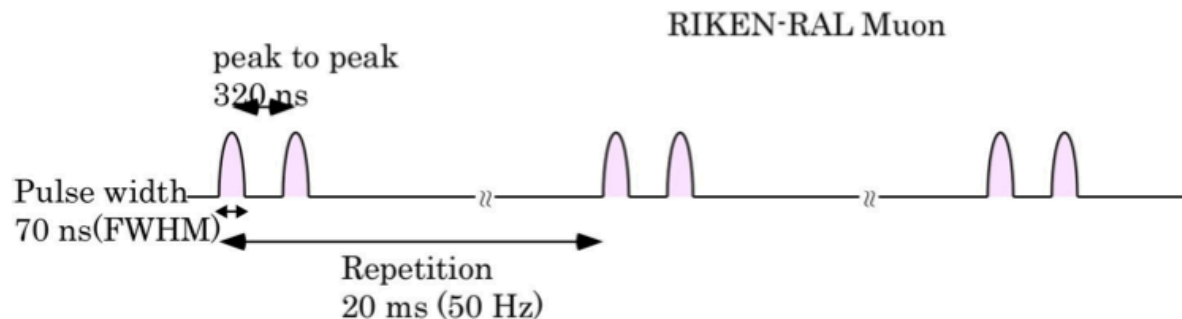
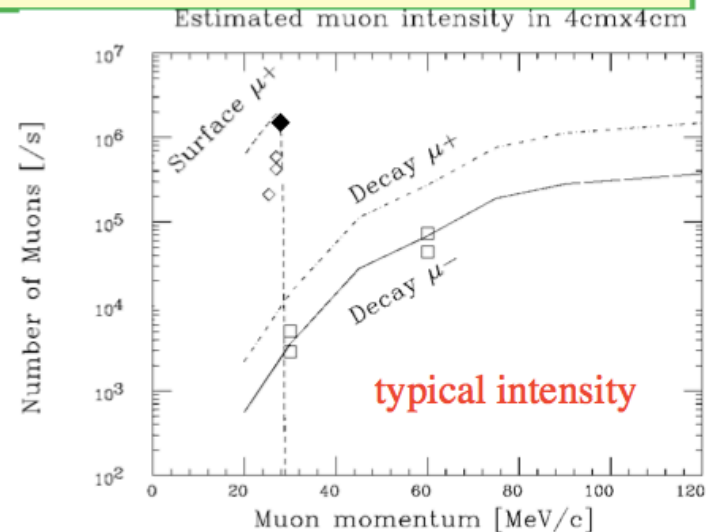
with magnetic kicker (<30 MeV/c))

### Operation

160 days/y of ISIS beam time

~40 days for UK

~120 days for RIKEN



# beam density enhancement

- Muon beam density enhancement was observed in a number of experiments carried out both at RIKEN-RAL (UK) and at TRIUMF (Canada) laboratories .
- They used several tapered tubes working with muon grazing angle: glass tubes, copper, gold plated copper.



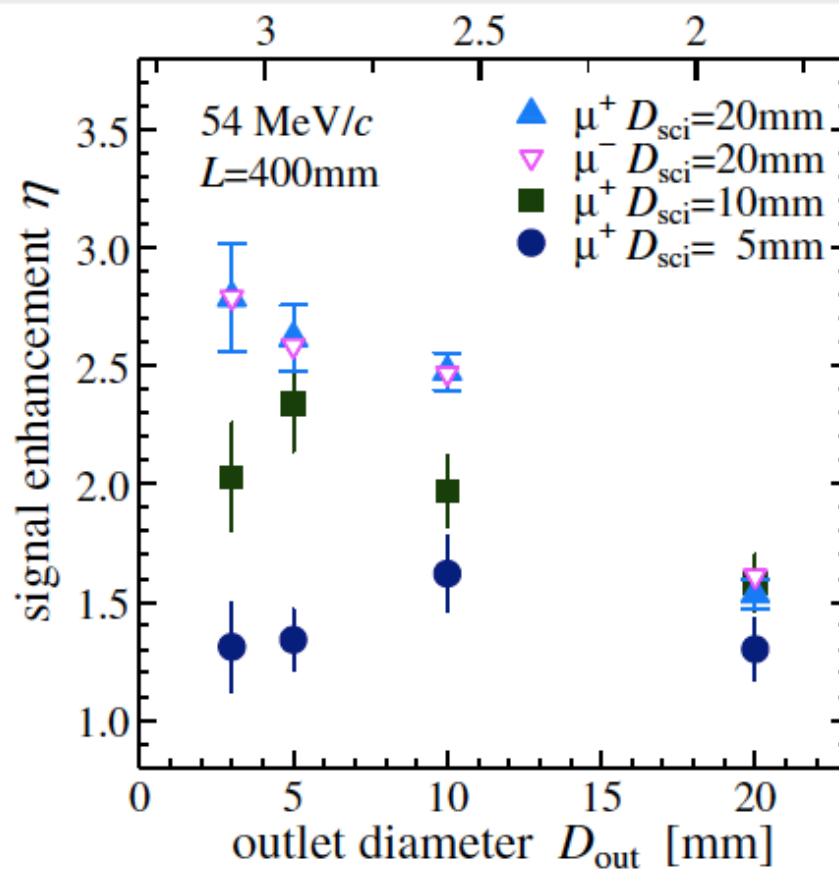
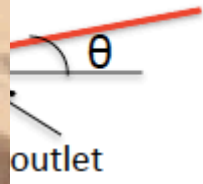


Fig. 2. (Color online) The signal enhancement factor  $\eta$  as a function of outlet diameter  $D_{out}$  for 54 MeV/c muons with  $L = 400$  mm tubes for  $D_{sci} = 5, 10,$  and  $20$  mm. The error bar includes statistical and systematic errors (see text). Data points with “T” shape error bar are the mean of multiple measurements and those with “|” shape error bar are measured only once. Error bars of  $\mu^-$  data are omitted for clarity.

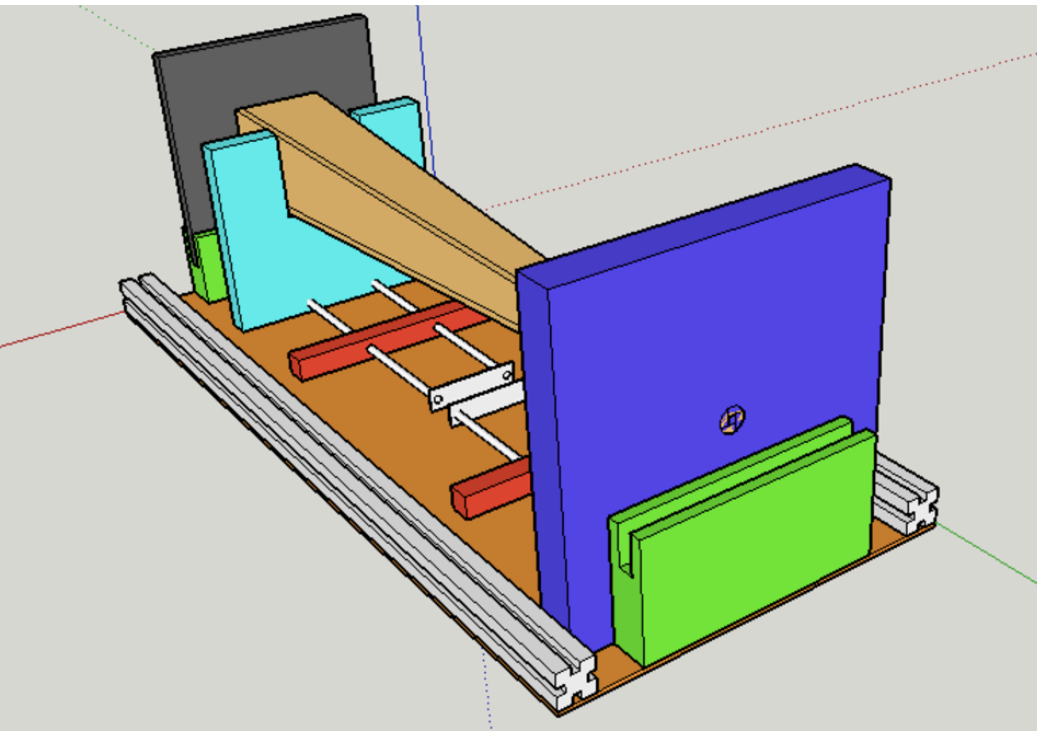
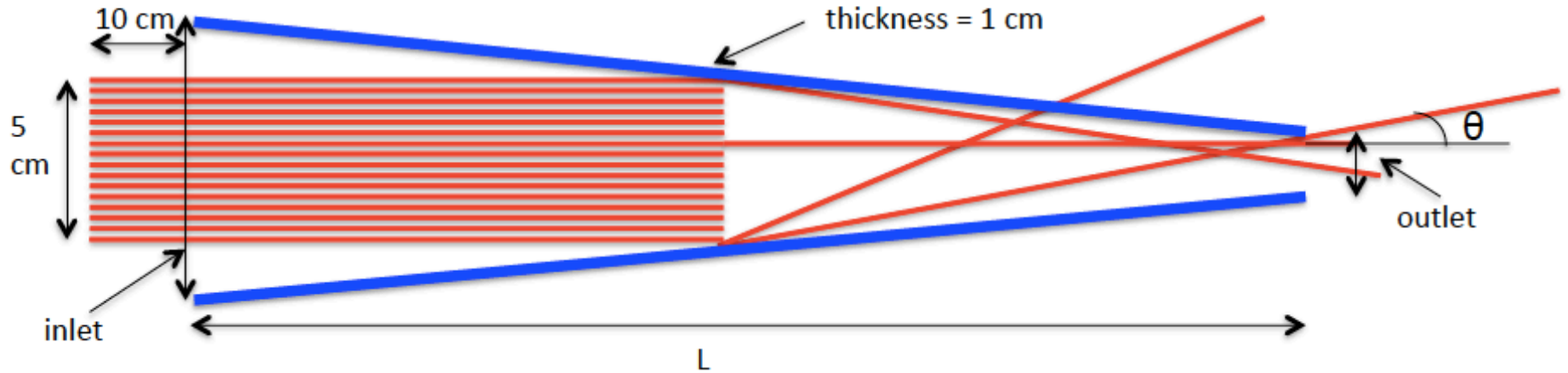
$$\eta = \begin{cases} V_{\text{with}}/V_{\text{without}} & \text{for } D_{\text{sci}} \leq D_{\text{out}}, \\ V_{\text{with}}/V_{\text{without}, d \leq D_{\text{out}}} & \text{for } D_{\text{sci}} > D_{\text{out}}, \end{cases} \quad (1a)$$

(1b)

- We decided to investigate the possibility to have also in FAMU such a density enhancement, thus recovering inside the gas target the part of the beam otherwise lost .
- We profited of a full day of run with RIKEN-RAL beam in Port 1 in March 2018 to collect data to be compared with GEANT4 simulation.
- Several experimental configurations were realized, profiting of the presence of three hodoscopes to monitor the beam intensity before and after the focusing optics: one made of polished copper, the other one of gold plated glass.



## Simulation set-up (optics material = Copper, Trapezoid tube)

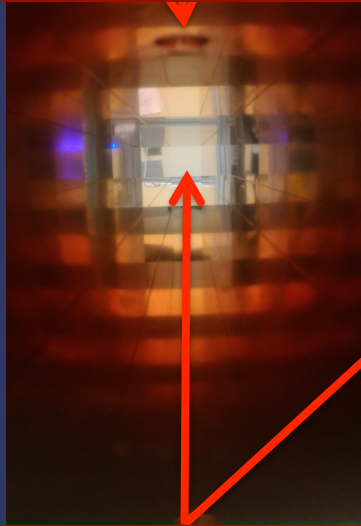


- Copper focusing optics realized in Pavia workshop and then assembled in Port1: mylar and kapton foils used for vacuum tightness. Copper optics dimensions (inlet: 3 cm, outlet: 1 cm, length: 1 m) were chosen after preliminary GEANT4 simulation. Dry scroll pump allowed to operate it with a stable pressure of 23 mbar.
- Another optics with the same dimensions has been realized with gold-plated glass by INAF-Brera group, having expertise with optics for X-ray focusing onto satellites.

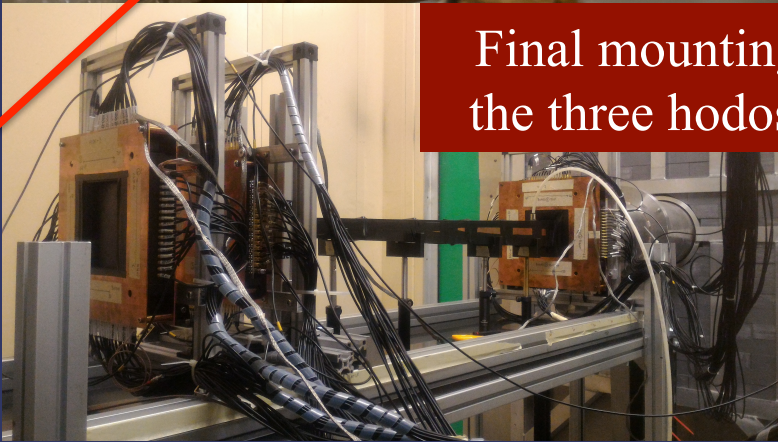


Copper optics before assembly

Inlet for tube evacuation

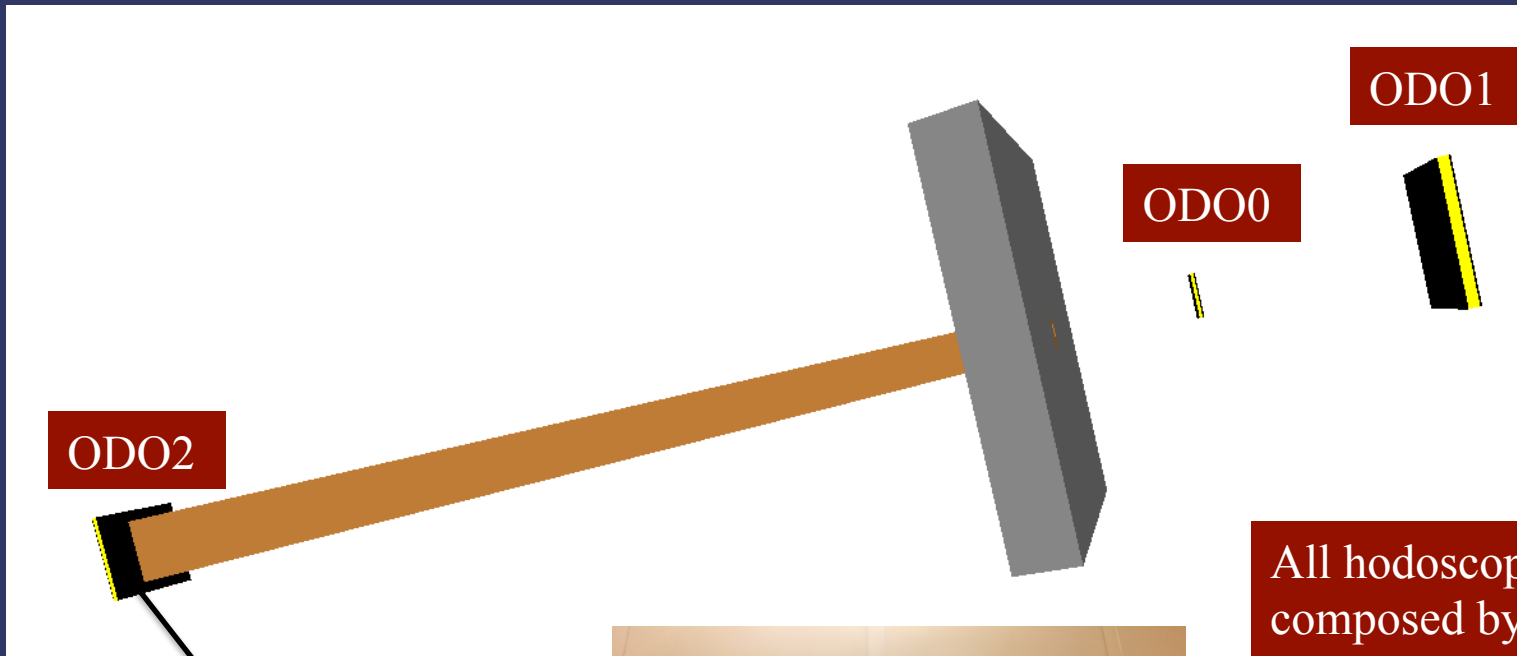


1 cm outlet



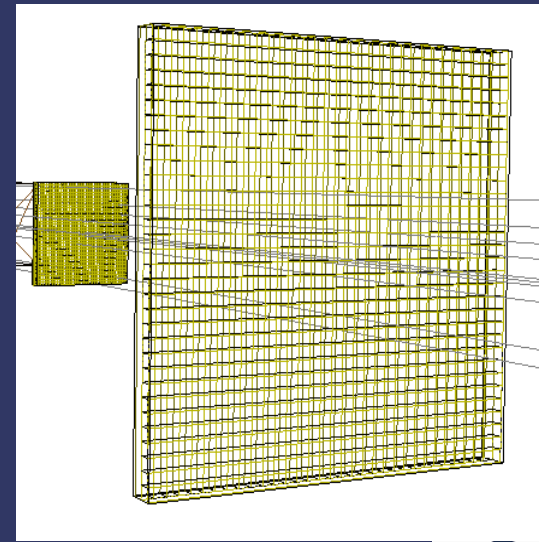
Final mounting with the three hodoscopes

- 1 full day of data taken with several setups:
  - ✓ Muon beam momentum scan: from 30 MeV/c to 80 MeV/c.
  - ✓ Use of a 1.4 cm polyethylene thickness before the optics to reduce the beam momentum.
  - ✓ Use of several foils of known thickness for calorimetric beam momentum measurements.
  - ✓ Copper optics in air and vacuum.
  - ✓ Glass optics in air.
- **A total of about 100 runs, 5000 events each**, has been collected for further analysis.
- The goal is the measurement of **beam muon flux** starting from the charge deposited by the muons inside the hodoscopes.
- To get the muon flux, two ingredients are needed:
  1. The hodoscope ADC calibration → **cosmic ray data from test run available**.
  2. Correction for low energy (non m.i.p.) muons of RIKEN-RAL → **GEANT4 simulation** of all setups is ongoing to evaluate the muon momentum in the three hodoscopes.

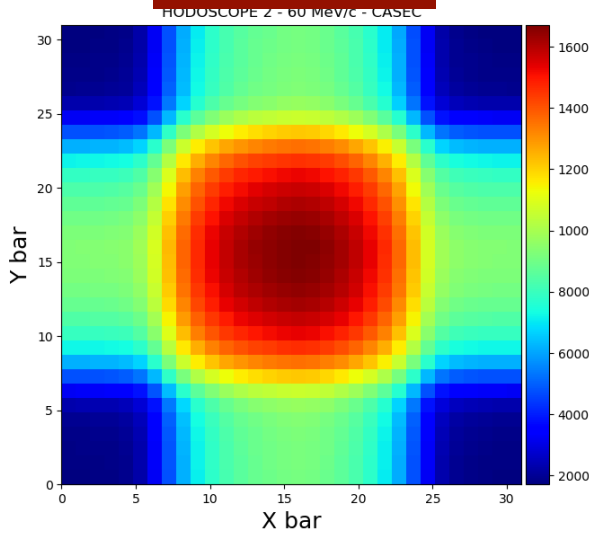


All hodoscopes are composed by 32x32 bars

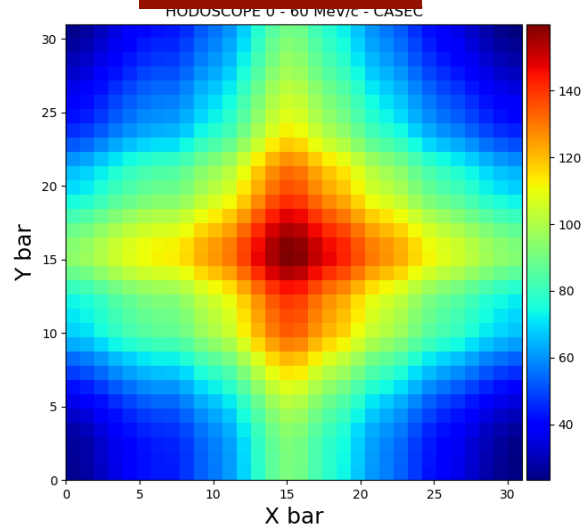
ABS in black



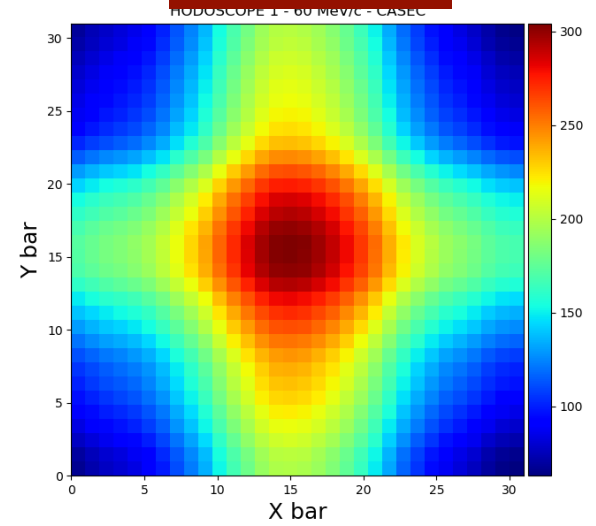
Simulatio ODO2



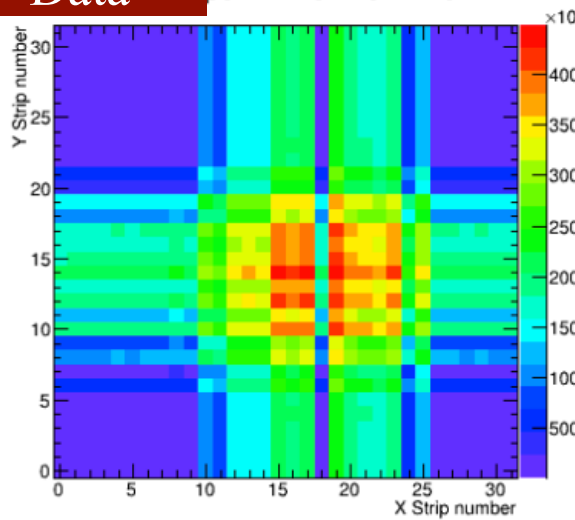
ODO0



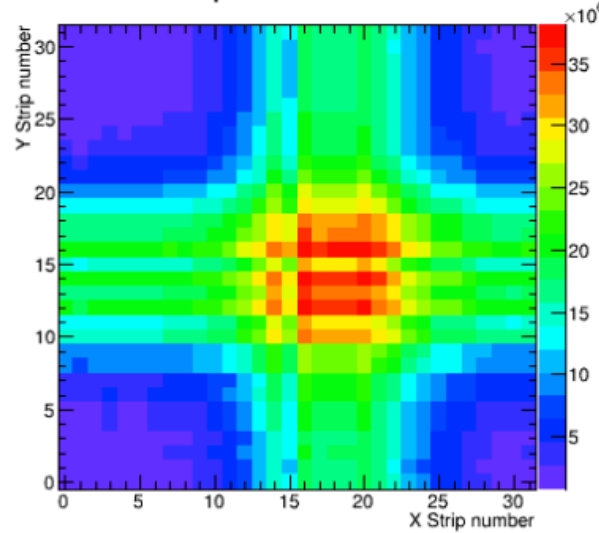
ODO1



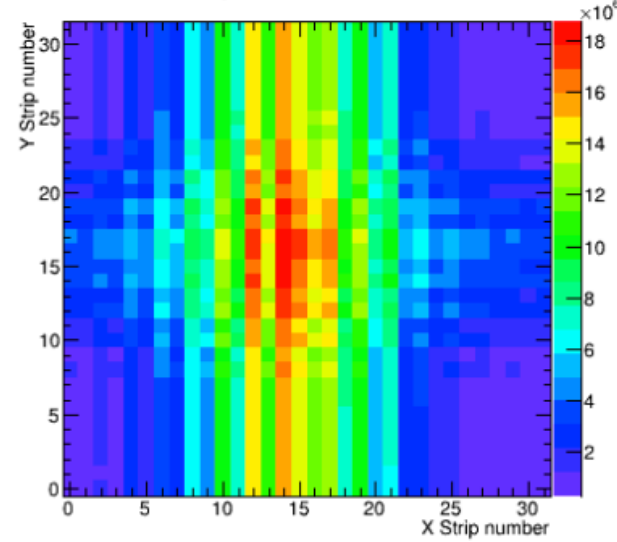
Data Hodoscope 2 X view vs. Y view



Hodoscope 0 X view vs. Y view

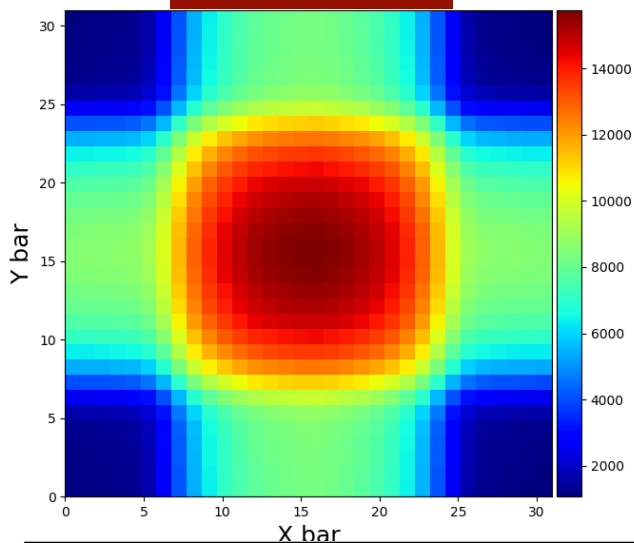


Hodoscope 1 X view vs. Y view

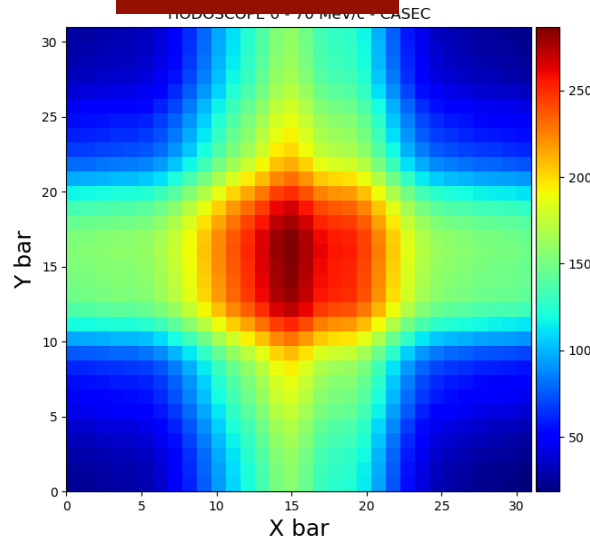




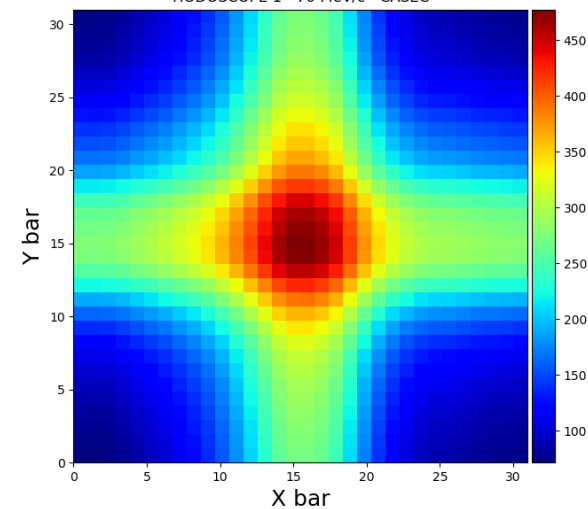
## Simulation ODO2



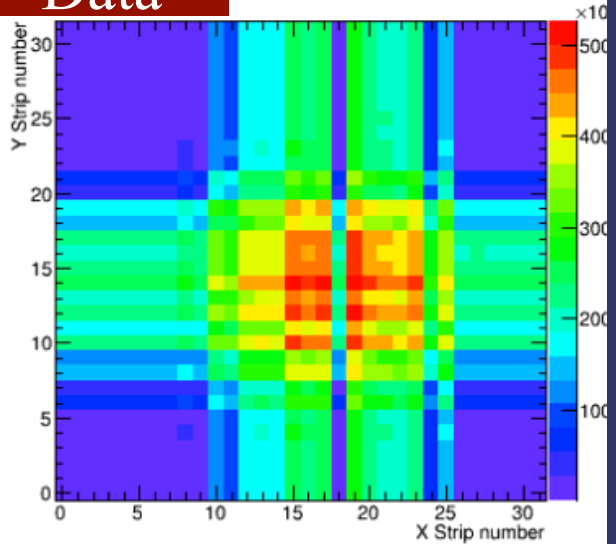
## Simulation ODO0



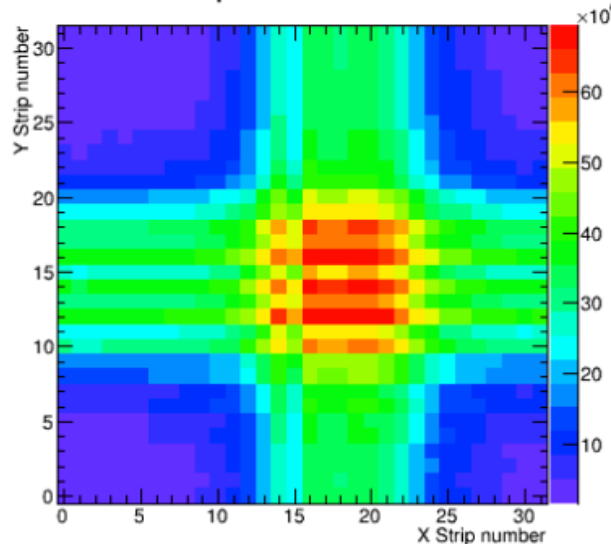
## Simulation ODO1



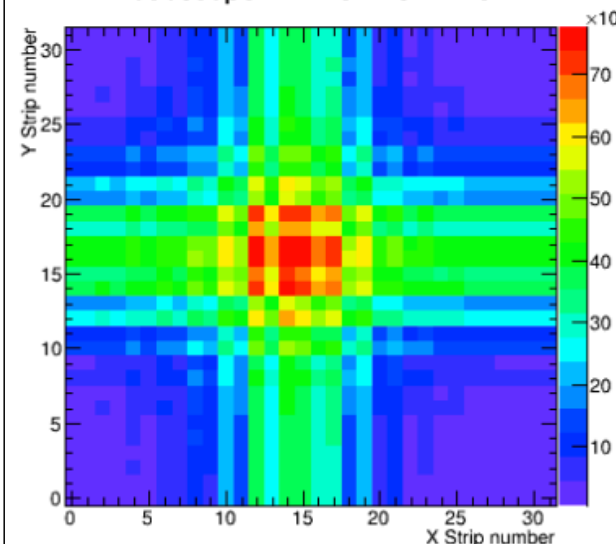
## Data Comparison: X view vs. Y view



## Data Comparison: Hodoscope 0 X view vs. Y view



## Data Comparison: Hodoscope 1 X view vs. Y view

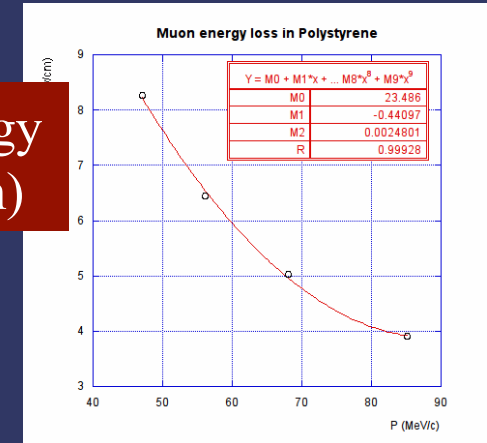
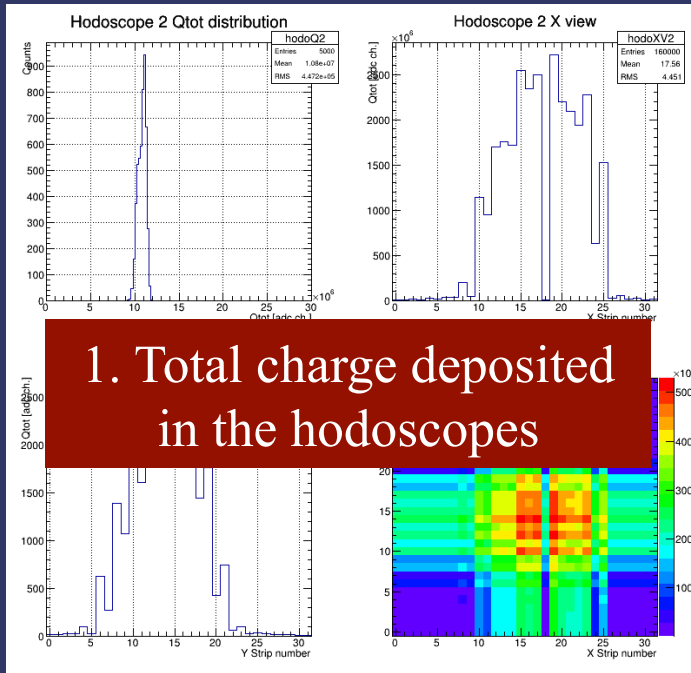


# Next steps

## 2. Calibration with cosmic ray data

## 3. Correction for low energy beam muon (G4 simulation)

## 1. Total charge deposited in the hodoscopes



- First evaluation of beam muon flux, to check the beneficial effects of the focusing optics and to choose the best beam muon momentum to operate FAMU experiment.
- Definition of the new cryogenic target as a function of the results from the focusing measurements.
- Possible new test in Port1, with the realization of a more refined focusing optics.

# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - high energy MIR fine-tunable laser
  - pulsed high intensity muon beam
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
  - Simulations
- Conclusions

# 2016 Target: a necessary trade-off

Main requirements:

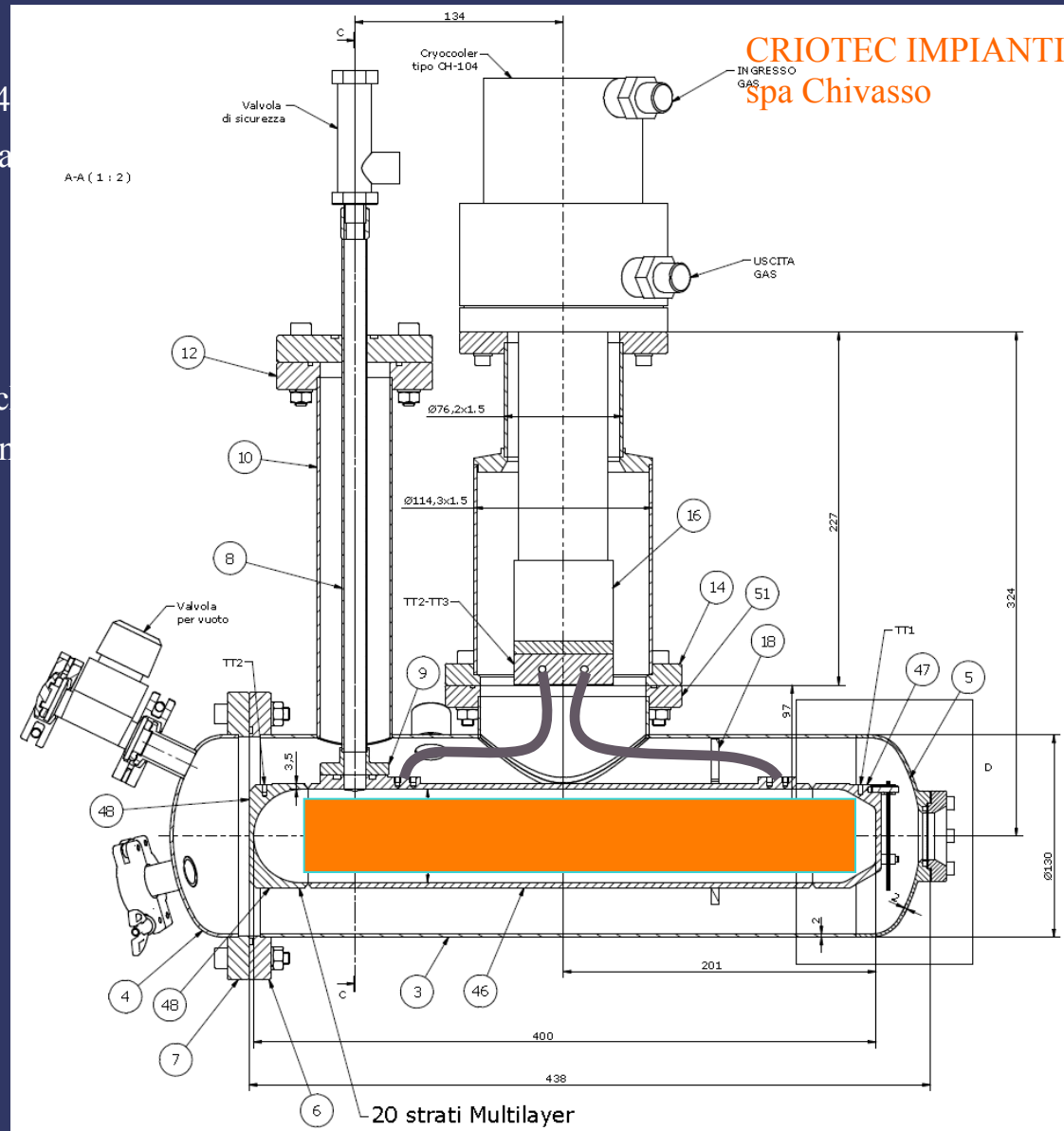
- Operating temperature range:  $40\text{ K} \leq T \leq 325\text{ K}$
- Temperature control for measurement runs at fixed  $T$  steps from 300 K to 50K
- Gas @ constant density,  $\text{H}_2$  charge pressure at room  $T$  is  $\sim 40\text{ atm}$
- International **safety** certification (Directive 97/23/CE PED)
- Minimize **walls and windows thickness**
- Target shape and dimensions to :
  - **maximize muon stop in gas**
  - **to minimize distance gas – detectors**
  - **to be compliant to allowable volume at Riken Port**
- $\text{H}_2$  compatible

... and, of course, all the above within time and cost constraints!

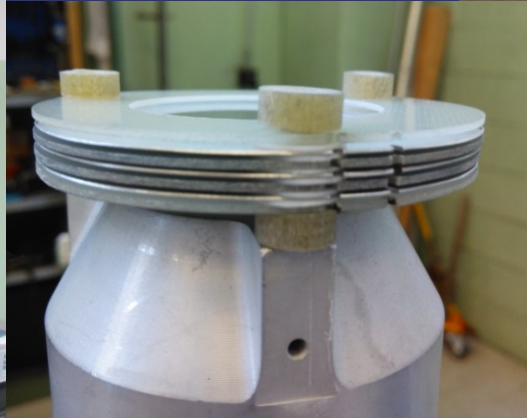
# 2016 Best solution

- Target= Inner vessel with high P gas (44
- Al alloy 6082 T6 cylinder D = 60 mm a 400 mm, inner volume of 1.08 l
- Internally Ni/Au plated (L = 280 mm)
- Cylinder side wall thickness = 3.5 mm
- Wrapped in 20 layers of MLI
- Front window D= 30 mm 2.85 mm thick
- Three discs of 0.075 mm Al foil for wire radiative shield
- 304L SS gas charging tube
- 304L SS cooler cold-end support
- G10 mechanical strut
- Two Cu straps for cooling

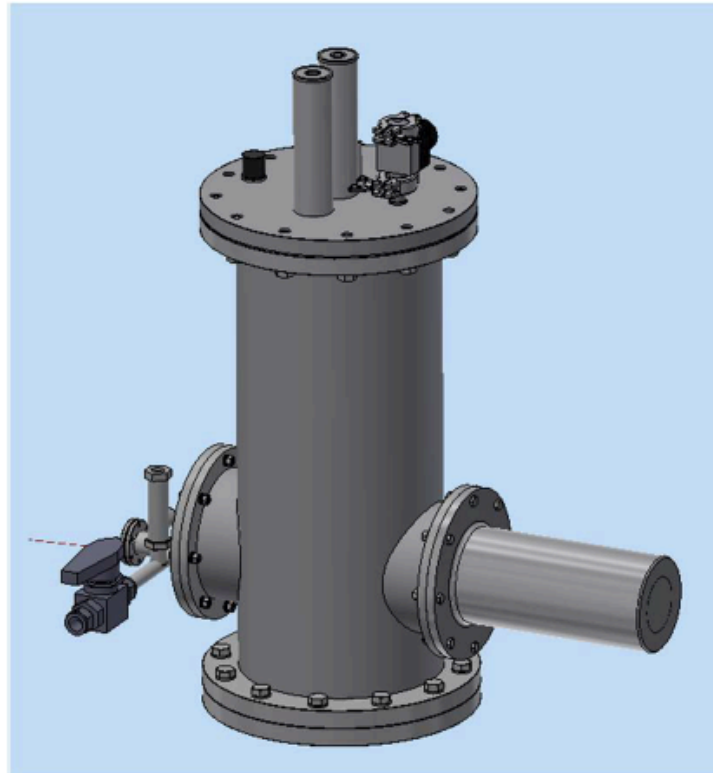
- Vacuum vessel = outer cylinder (P atm)
- Al6060 D=130 mm, 2 mm thick walls
- ≈30mm between inner/outer walls
- Flanged Al window 0.8 mm thick
- Pumping valve & harness feed-tru's



# Target in lab



# 2018 target solution under study

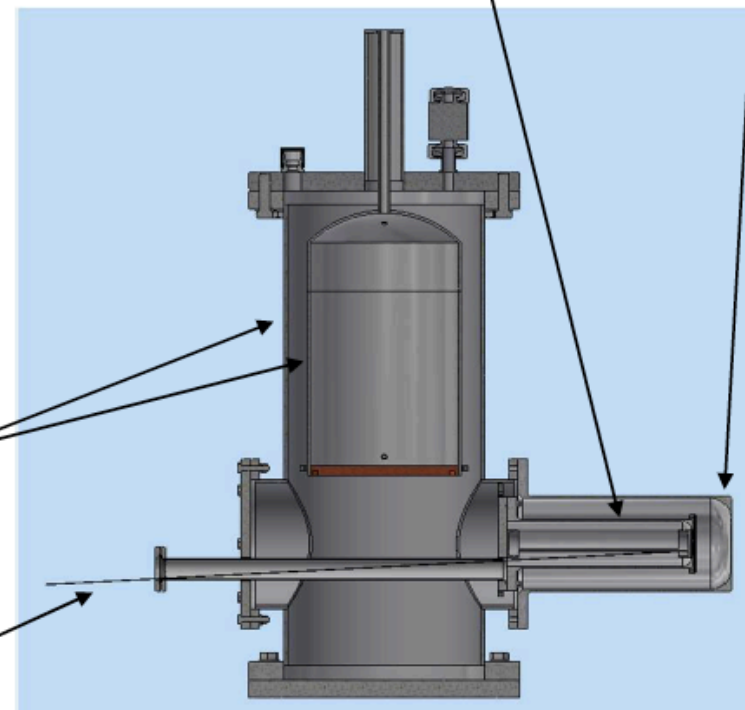


Nell'ottica di ridurre la lunghezza di percorso del laser, i diametri del criostato e del serbatoio sono stati ridotti a scapito della loro altezza

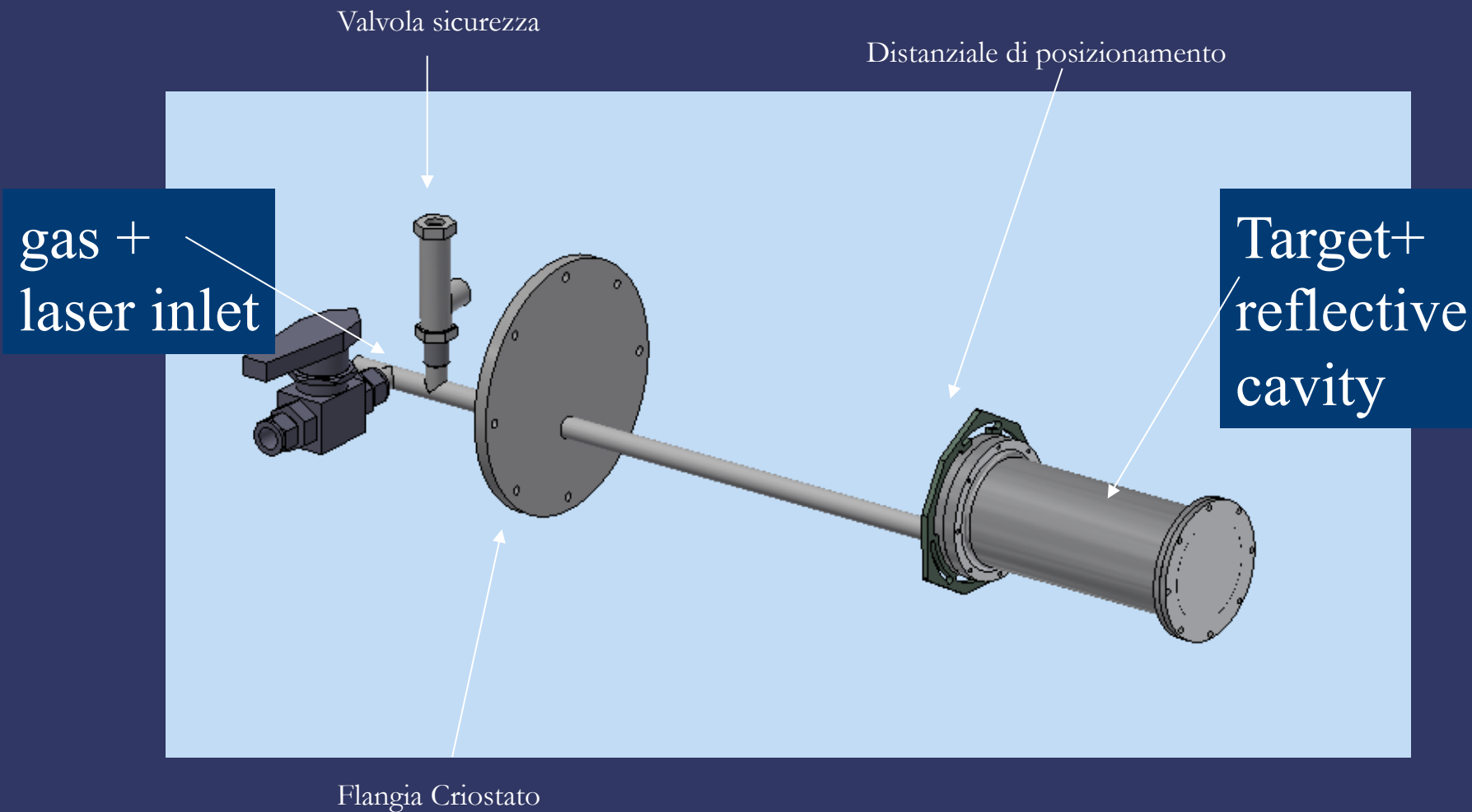
Analizzato percorso laser con inclinazione  $0.050 \text{ rad}$

Analizzata la possibilità di evitare la flangia di estremità del Target

Studio copertura Target

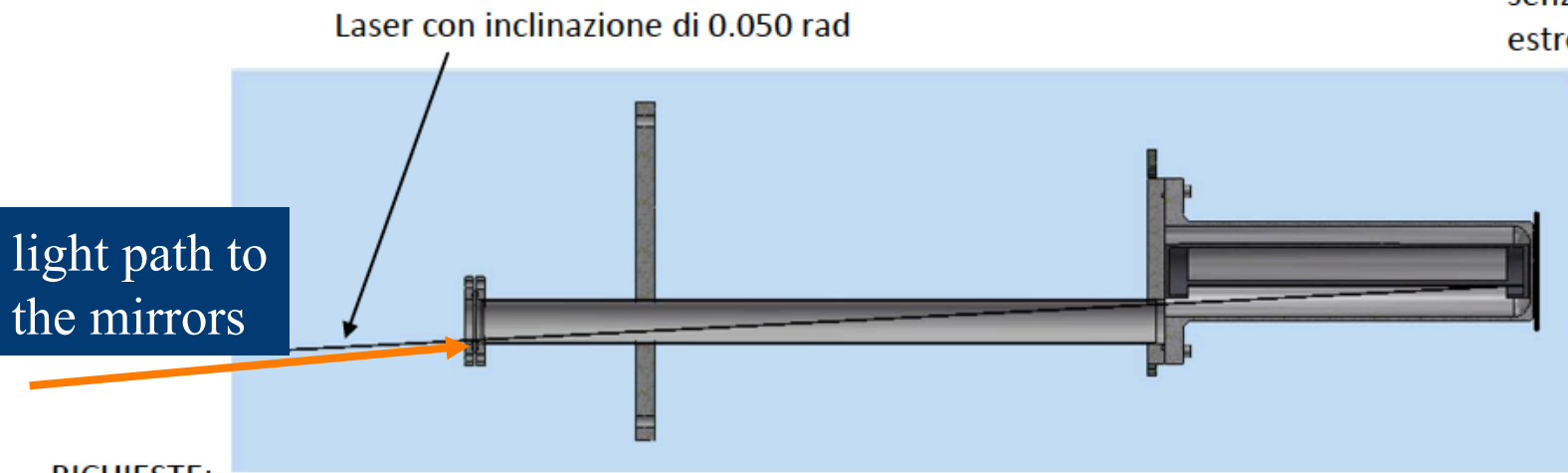
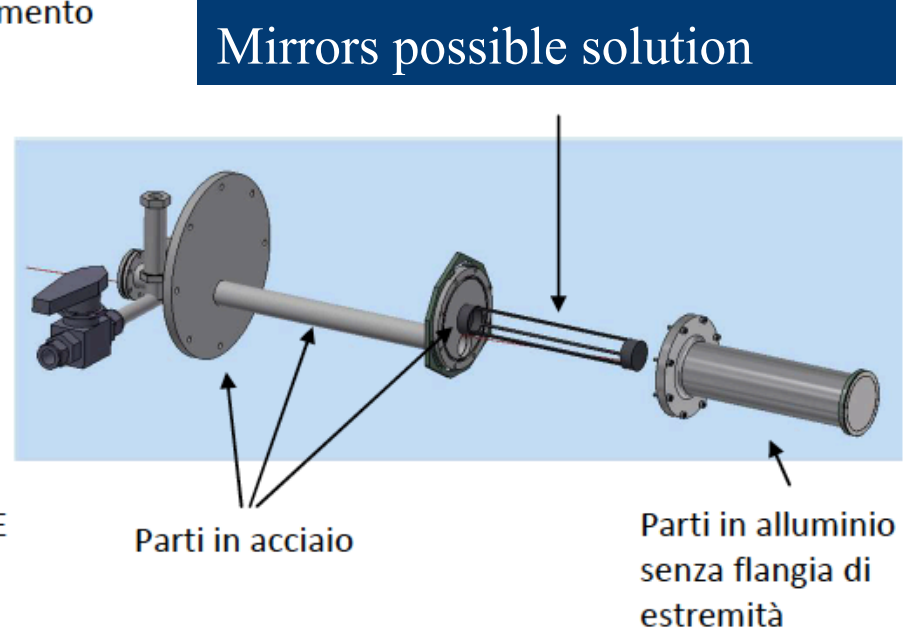
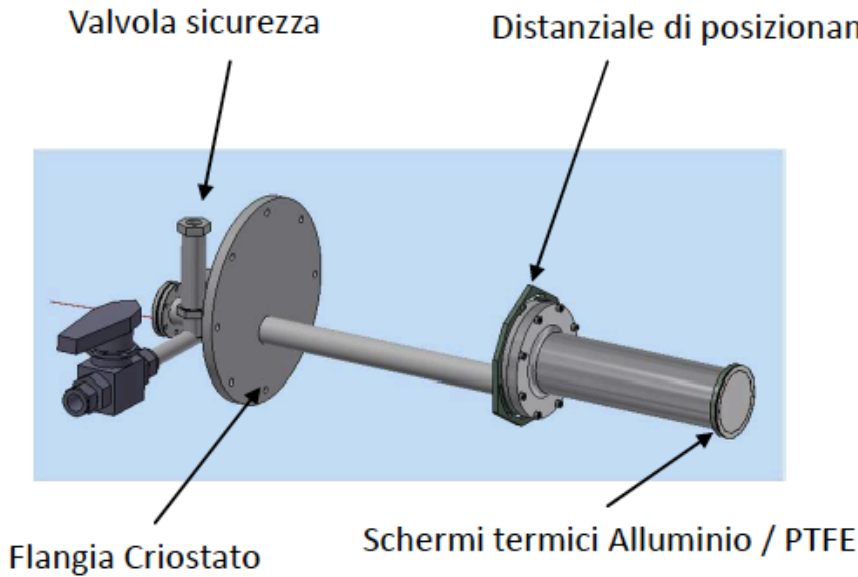


# low temperature gas target very preliminary approach for simulation studies





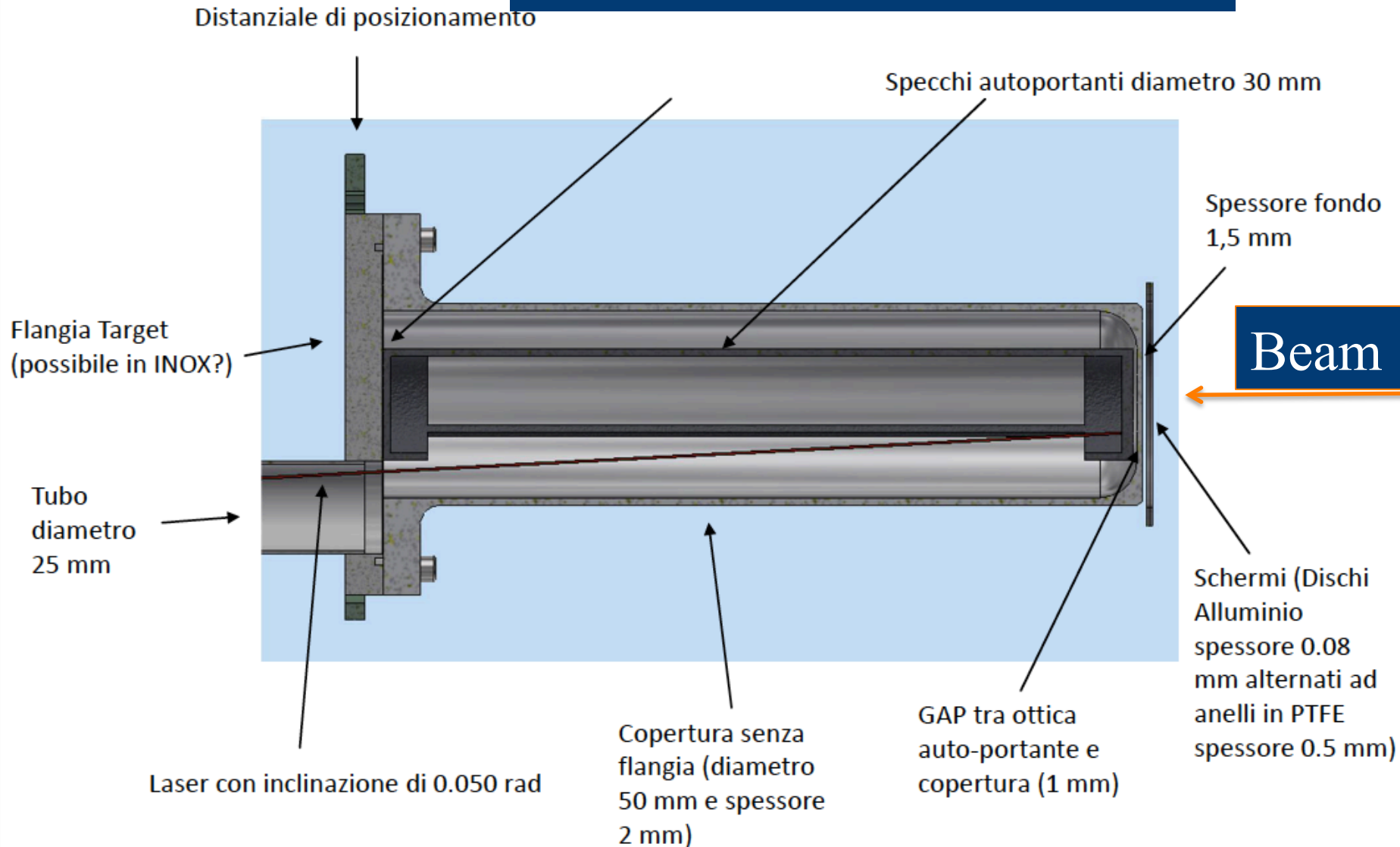
# Very preliminary approach under simulation studies



**RICHIESTE:**

- possibilità di ridurre angolo massimo del laser (magari a 0.030 rad) in modo ridurre i diametri del tubo e della flangia del Target

# Very preliminary approach under simulation studies



# OUTLINE

- FAMU background & motivations
- The method to measure the hfs
- FAMU's path 2014-2018
- 2016 muon transfer rate measurements between 100 and 300 K
- 2018 low pressure data - rates evaluation
- FAMU key ingredients optimization
  - High efficiency multi-pass optical cavity
  - high energy MIR fine-tunable laser
  - pulsed high intensity muon beam
  - cryogenic gas optical cavity target
  - best X-rays detectors (fast and accurate)
- Conclusions

# Detectors: suited for time-resolved X-ray spectroscopy

## Germanium HPGe: low energy X-rays spectroscopy

ORTEC GLP:

Energy Range: 0 – 300 keV

Crystal Diameter: 11 mm

Crystal Length: 7 mm

Beryllium Window: 0.127 mm

Resolution Warrented (FWHM):

- at 5.9 keV is 195 eV ( $T_{sh}$  6  $\mu$ s)

- at 122 keV is 495 eV ( $T_{sh}$  6  $\mu$ s)

ORTEC GMX:

Energy Range: 10 – 1000 keV

Crystal Diameter: 55 mm

Crystal Length: 50 mm

Beryllium Window: 0.5 mm

Resolution Warrented (FWHM):

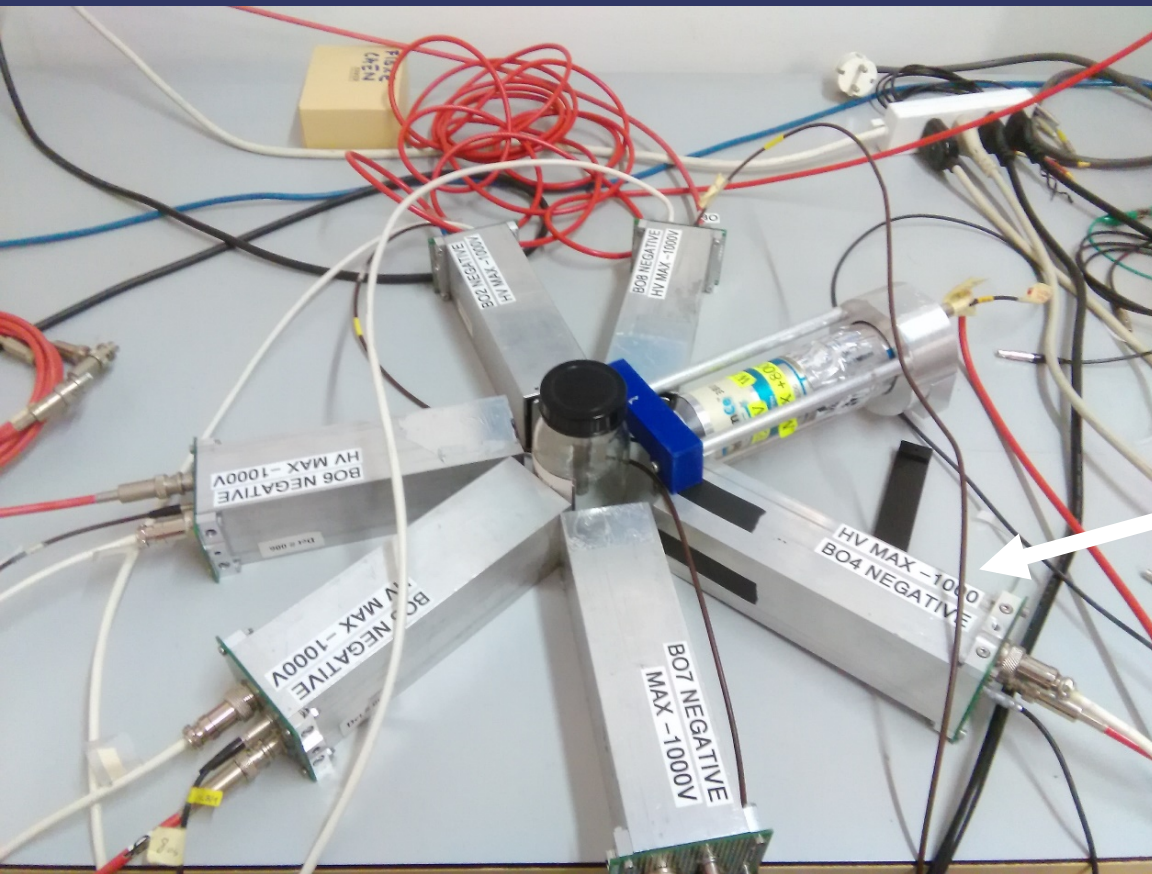
- at 5.9 keV is 600 eV ( $T_{sh}$  6  $\mu$ s)

- at 122 keV is 800 eV ( $T_{sh}$  6  $\mu$ s)



# Detectors: suited for time-resolved X-ray spectroscopy

Lanthanum bromide scintillating crystals  $[\text{LaBr}_3(\text{Ce})]$ :  
fast timing X-rays detectors

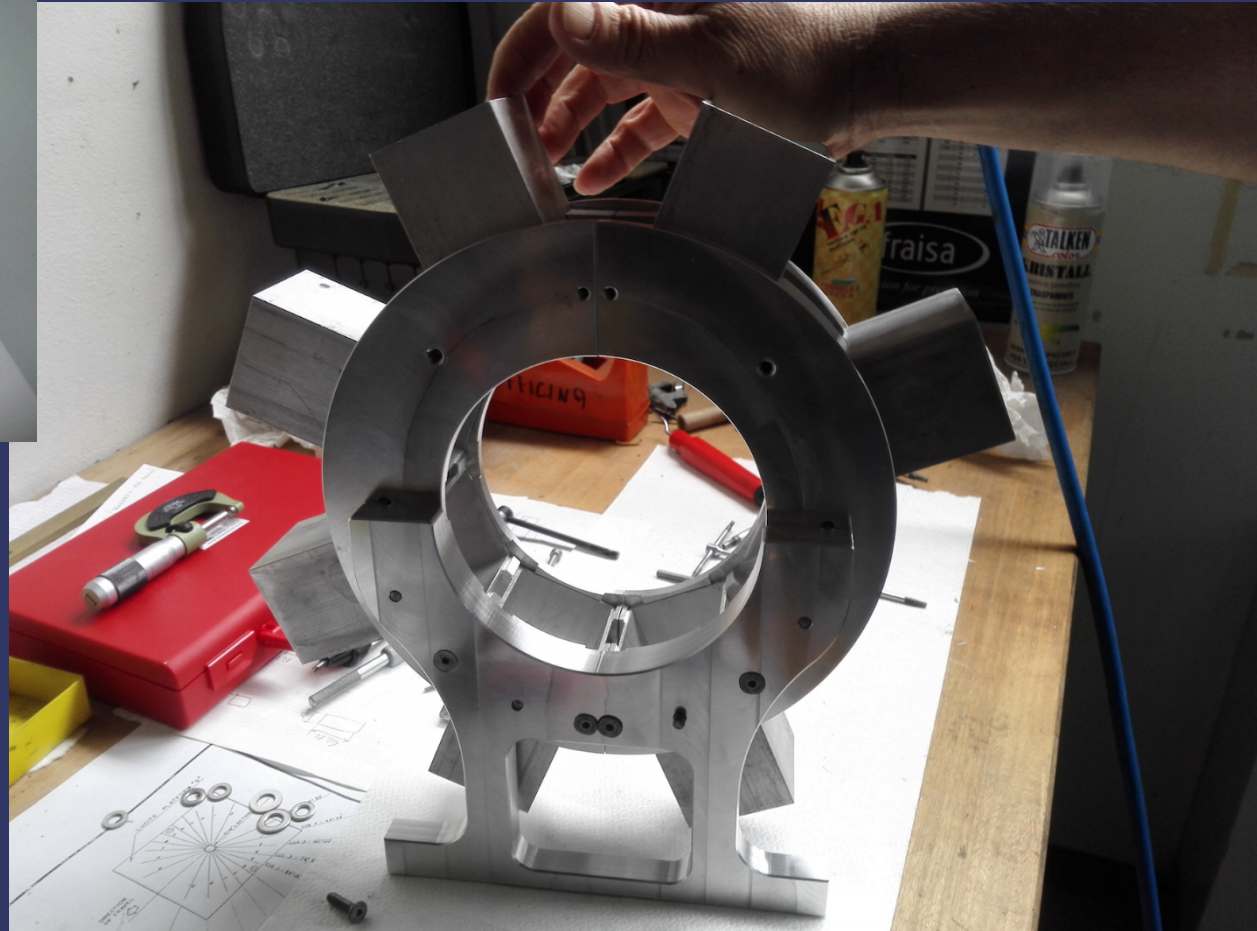
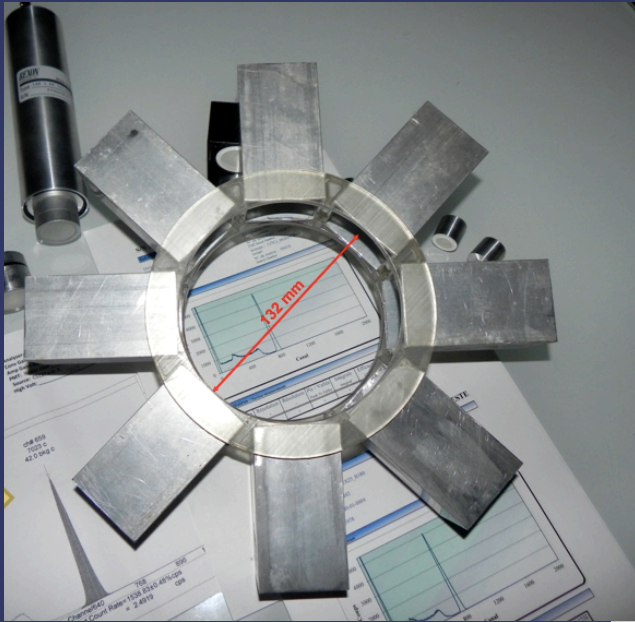


8 cylindrical 1 inch diameter  
1 inch long  $\text{LaBr}_3(5\% \text{Ce})$   
crystals  
read by PMTs.

On purpose developed fast  
electronics and fast digital  
processing signal.

Lab test

# Star-shaped support for detectors

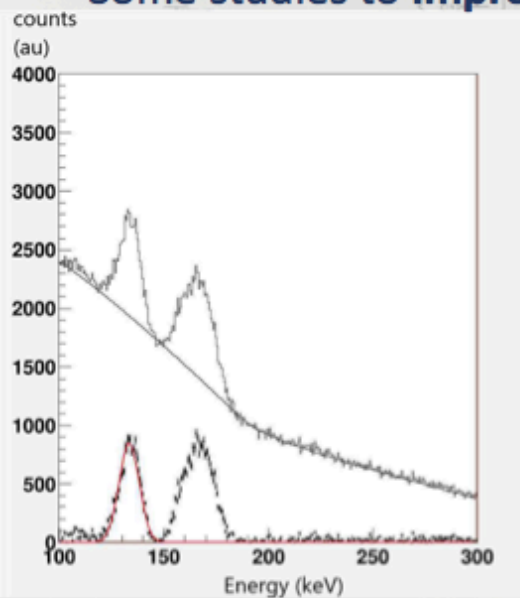


# Last steps Bologna detectors

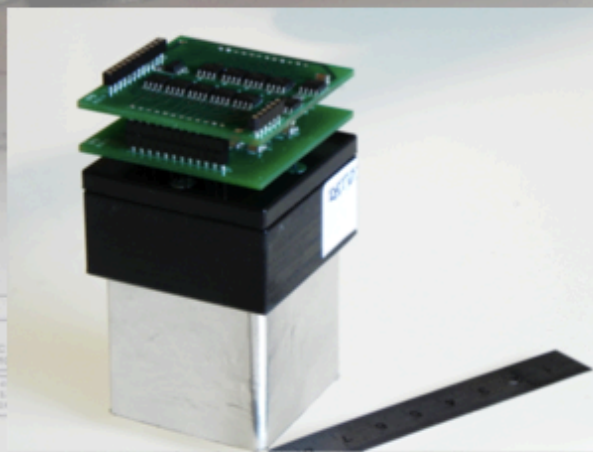
## LaBr<sub>3</sub> + (UBA) Hamamatsu photomultiplier

- **Custom active voltage divider** for high rate applications
- **8 built and on beam tested + 8 build ongoing**
- New **high coverage detector's geometry** in order to adapt to the new work in progress target.
- Some studies to **improve energy resolution**

Energy (keV)	Literature resolution	Famu detector
122 keV	7.4%	8.8%
662 keV	2.8%	3.5%



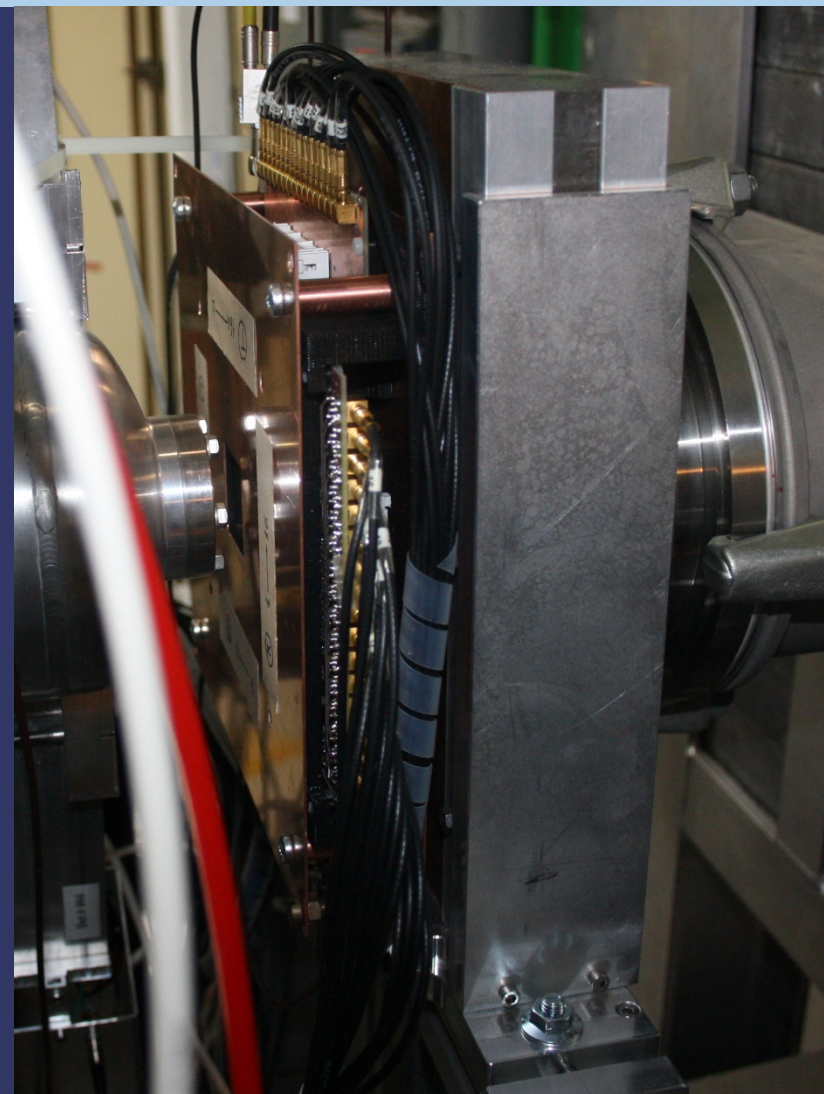
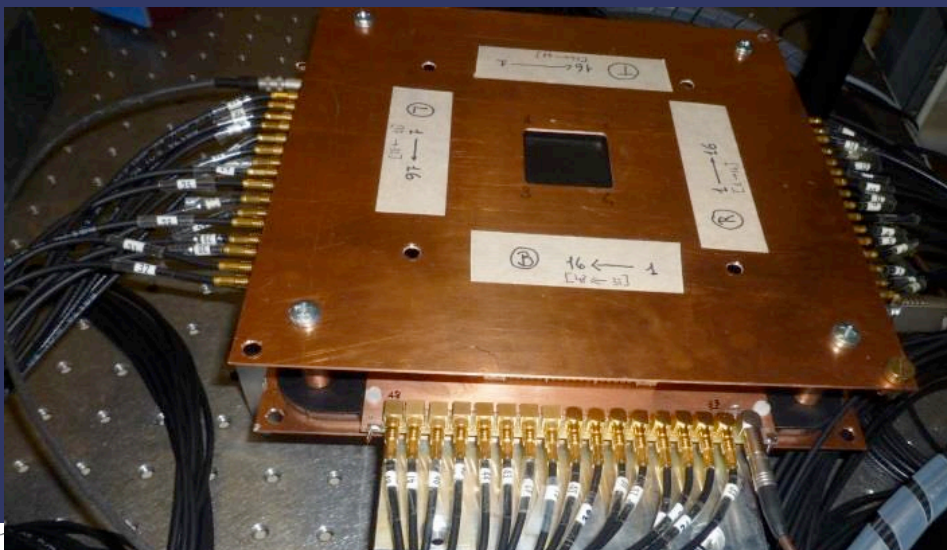
**Delayed muonic oxygen lines well resolved.**  
The 133keV line resolution is 8.5%, slightly worse than the 8.1% predicted



Article: G. Baldazzi and al., The LaBr<sub>3</sub> (Ce) based detection system for the FAMU experiment, Journal of Instrumentation 12 (2017) 03

# Hodoscope for beam shape monitoring

Final version:  
two planes (X and Y) of 32 scintillating  
fibers 1 x 1 mm<sup>2</sup> square section  
SiPM reading with fast electronics  
3D printed supports



hodoscope in the 2016 setup





RIKEN RAL 2016-7



RIKEN RAL 2014-5

the



collaboration

Jinst

PUBLISHED BY IOP PUBLISHING FOR SISSA MEDIALAB

RECEIVED: April 5, 2016

ACCEPTED: April 11, 2016

PUBLISHED: May 12, 2016

## Steps towards the hyperfine splitting measurement of the muonic hydrogen ground state: pulsed muon beam and detection system characterization



### The FAMU collaboration

A. Adamczak,<sup>a</sup> G. Baccolo,<sup>b</sup> D. Bakalov,<sup>c</sup> G. Baldazzi,<sup>d</sup> R. Bertoni,<sup>b</sup> M. Bonesini,<sup>b</sup> V. Bonvicini,<sup>e</sup> G. Campana,<sup>d</sup> R. Carbone,<sup>e</sup> T. Cervi,<sup>g,h</sup> F. Chignoli,<sup>b</sup> M. Clemenza,<sup>b</sup> L. Colace,<sup>i,j</sup> A. Curioni,<sup>b</sup> M. Danailov,<sup>e,f</sup> P. Danev,<sup>c</sup> I. D'Antone,<sup>d</sup> A. De Bari,<sup>g,h</sup> C. De Vecchi,<sup>h</sup> M. De Vincenzi,<sup>i,k</sup> M. Furini,<sup>d</sup> F. Fuschino,<sup>d</sup> K.S. Gadedjisso-Tossou,<sup>e,l</sup> D. Guffanti,<sup>e</sup> A. Iacofano,<sup>i</sup> K. Ishida,<sup>m</sup> D. Iugovaz,<sup>e</sup> C. Labanti,<sup>d</sup> V. Maggi,<sup>b</sup> A. Margotti,<sup>d</sup> M. Marisaldi,<sup>d</sup> R. Mazza,<sup>b</sup> S. Meneghini,<sup>d</sup> A. Menegolli,<sup>g,h</sup> E. Mocchiutti,<sup>e</sup> M. Moretti,<sup>b</sup> G. Morgante,<sup>d</sup> R. Nardò,<sup>h</sup> M. Nastasi,<sup>b</sup> J. Niemela,<sup>i</sup> E. Previtali,<sup>b</sup> R. Ramponi,<sup>n</sup> A. Rachevski,<sup>e</sup> L.P. Rignanese,<sup>d</sup> M. Rossella,<sup>h</sup> P.L. Rossi,<sup>d</sup> F. Somma,<sup>i,o</sup> M. Stoilov,<sup>c</sup> L. Stoychev,<sup>e,l</sup> A. Tomaselli,<sup>h,p</sup> L. Tortora,<sup>i</sup> A. Vacchi,<sup>e,q,1</sup> E. Vallazza,<sup>e</sup> G. Zampa<sup>e</sup> and M. Zuffa<sup>d</sup>

<sup>a</sup>Institute of Nuclear Physics, Polish Academy of Sciences, Radzikowskiego 152, Kraków, PL31342 Poland

<sup>b</sup>National Institute for Nuclear Physics (INFN), Sezione di Milano Bicocca, Piazza della Scienza 3, Milano, Italy

<sup>c</sup>Institute for Nuclear Research and Nuclear Energy, Bulgarian Academy of Sciences,

2016 JINST 11 P05007



Istituto Nazionale di Fisica Nucleare



The FAMU  
Yearly Collaboration Meeting Trieste

# Summary

The FAMU project is progressing towards the spectroscopy of the hyperfine splitting (hfs) in the 1S state of muonic hydrogen  $\Delta E_{\text{hfs}}(\mu^-p)_{1S}$

through a measurement of the muon transfer rate to oxygen

open problems progressively faced:

- *first measurement* of the temperature dependent muon transfer rate to Oxygen
- *innovative* and powerful laser system
- *optimized intense pulsed beam target and optical system*
- *best detectors* for energy and time observation

Looking forward to perform the initiate the spectroscopic measurements 2018-19.



Than you for your attention