

# A New Neutrino Transport Module Available in FLASH-X

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# Number-Conservative Spectral $\mathcal{O}(v/c)$ Two-Moment Model<sup>1</sup>

► **Number equation**

$$\partial_t(\mathcal{D} + v^i \mathcal{I}_i) + \partial_i(\mathcal{I}^i + v^i \mathcal{D}) - \frac{1}{\varepsilon^2} \partial_\varepsilon (\varepsilon^3 \mathcal{K}^i_k \partial_i v^k) = \frac{1}{4\pi} \int_{\mathbb{S}^2} \mathcal{C}[f] d\omega$$

► **Number flux equation**

$$\begin{aligned} \partial_t(\mathcal{I}_j + v^i \mathcal{K}_{ij}) + \partial_i(\mathcal{K}^i_j + v^i \mathcal{I}_j) - \frac{1}{\varepsilon^2} \partial_\varepsilon (\varepsilon^3 \mathcal{L}^i_{kj} \partial_i v^k) \\ + (\mathcal{I}^i \partial_i v_j - \mathcal{L}^i_{kj} \partial_i v^k) = \frac{1}{4\pi} \int_{\mathbb{S}^2} \mathcal{C}[f] \ell_j d\omega \end{aligned}$$

► **Angular moments of kinetic distribution  $f$**

$$\{\mathcal{D}, \mathcal{I}^i, \mathcal{K}^{ij}, \mathcal{L}^{ijk}\}(\varepsilon, \mathbf{x}, t) = \frac{1}{4\pi} \int_{\mathbb{S}^2} f(\omega, \varepsilon, \mathbf{x}, t) \{1, \ell^i, \ell^i \ell^j, \ell^i \ell^j \ell^k\} d\omega$$

► Closed by specifying  $\mathcal{K}^{ij}$  and  $\mathcal{L}^{ijk}$  in terms of  $\mathcal{D}$  and  $\mathcal{I}^i$

► Components of fluid three-velocity  $v^i$

► Comoving-frame spherical-polar momentum coordinates  $(\omega, \varepsilon)$

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<sup>1</sup>Liau, E, Harris, Zelledge, Mezzacappa arXiv:2309.04429

$$\mathcal{C}[f, \bar{f}](\mathbf{p}) = (1 - f(\mathbf{p})) \eta(\mathbf{p}) - \chi(\mathbf{p}) f(\mathbf{p}) \quad (\text{Emission/absorption})$$

$$+ (1 - f(\mathbf{p})) \int_{V_p} R^{\text{IN}}(\mathbf{p}, \mathbf{p}') f(\mathbf{p}') dV_{p'} \quad (\text{Scattering})$$

$$- f(\mathbf{p}) \int_{V_p} R^{\text{OUT}}(\mathbf{p}, \mathbf{p}') (1 - f(\mathbf{p}')) dV_{p'}$$

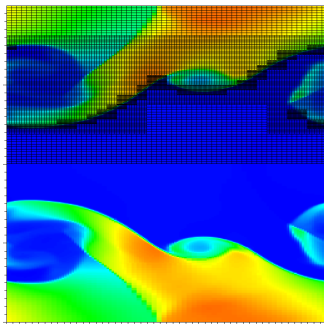
$$+ (1 - f(\mathbf{p})) \int_{V_p} R^{\text{PRO}}(\mathbf{p}, \mathbf{p}') (1 - \bar{f}(\mathbf{p}')) dV_{p'} \quad (\text{Pair processes})$$

$$- f(\mathbf{p}) \int_{V_p} R^{\text{ANN}}(\mathbf{p}, \mathbf{p}') \bar{f}(\mathbf{p}') d\mathbf{p}'$$

- ▶ Opacities depend nonlinearly on matter state (e.g.,  $\rho$ ,  $T$ , and  $Y_e$ )
- ▶ Pauli blocking factors:  $(1 - f)$
- ▶ Scattering and pair processes couple in momentum space:  $\mathcal{O}(N_p^2)$
- ▶ Pair processes couple neutrinos and antineutrinos

# Toolkit for High-Order Neutrino Rad-Hydro (THORNADO<sup>4</sup>)

- ▶ Discontinuous Galerkin (DG) methods
  - ▶ Hydrodynamics<sup>2</sup>
  - ▶ **Spectral, two-moment neutrino transport<sup>3</sup>**
- ▶ Tabulated microphysics (WEAKLIB)
  - ▶ Equations of State
  - ▶ Neutrino opacities
- ▶ GPU offloading with OpenMP or OpenACC
- ▶ Distributed parallelism and AMR through AMREX or FLASH-X



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<sup>2</sup>Pochik et al. (2021), ApJS, 253:21; Dunham et al. (arXiv:2307.10904)

<sup>3</sup>Chu et al. (2019), JCP, 389, 62; Laiu et al. (2021), ApJS, 253:52; Laiu et al. (arXiv:2309.04429)

<sup>4</sup>[github.com/endeve/thornado](https://github.com/endeve/thornado)

- ▶ Library for tabulated microphysics (EoS and weak interactions)
  - ▶ Tabulation in terms of matter states (e.g.,  $\rho$ ,  $T$ , and  $Y_e$ ) and neutrino energy ( $\epsilon$ )
- ▶ Basic functionality for hydrodynamics and neutrino transport algorithms
  - ▶ Interpolation on shared grids (EoS and weak interactions)
  - ▶ EoS inversions (e.g.,  $\epsilon \rightarrow T$  and  $s \rightarrow T$ )
- ▶ GPU offloading with OpenMP or OpenACC

	Process	original WeakLib	updated WeakLib
I	$e^- + p \rightleftharpoons n + \nu_e$	Bruenn (1985)	Reddy et al. (1998); Horowitz (2002)
II	$e^+ + n \rightleftharpoons p + \bar{\nu}_e$	Bruenn (1985)	Reddy et al. (1998); Horowitz (2002)
III	$e^- + A(Z, N) \rightleftharpoons A(Z-1, N+1) + \nu_e$	Fuller et al. (1982)	Langanke et al. (2003); Hix et al. (2003)
IV	$\nu + \{n, p, A\} \rightleftharpoons \nu + \{n, p, A\}$	Bruenn (1985)	Bruenn & Mezzacappa (1997); Horowitz (1997, 2002)
V	$\nu + e^\pm \rightleftharpoons \nu' + e^{\pm'}$	Bruenn (1985)	Bruenn (1985)
VI	$e^- + e^+ \rightleftharpoons \nu + \bar{\nu}$	Bruenn (1985)	Bruenn (1985)
VII	$N + N \rightleftharpoons N' + N' + \nu + \bar{\nu}$	X	Hannestad & Raffelt (1998)

<sup>5</sup>[github.com/starkiller-astro/weaklib](https://github.com/starkiller-astro/weaklib)

# Self-Gravitating Neutrino Radiation Hydrodynamics

- ▶ **Hydrodynamics** (Euler equations with nuclear EoS)

$$d_t \mathbf{u} = \mathbf{T}_u(\mathbf{u}, \Phi) + \mathbf{C}_u(\mathcal{U}, \mathbf{u})$$

Hyperbolic system with sources —  $\mathbf{u} \in \mathbb{R}^6$  per spacetime point

- ▶ **Neutrino transport** (spectral two-moment model)

$$d_t \mathcal{U} = \mathbf{T}_{\mathcal{U}}(\mathcal{U}, \mathbf{u}) + \mathbf{C}_{\mathcal{U}}(\mathcal{U}, \mathbf{u})$$

Hyperbolic system with sources —  $\mathcal{U} \in \mathbb{R}^{6 \times 4 \times 32 = 768}$  per spacetime point

- ▶ **Gravity** (Poisson equation)

$$F(\Phi, \mathbf{u}) = 0$$

Elliptic equation for scalar potential  $\Phi$

## Coupling THORNADO with FLASH-X

- ▶ First-order Lie–Trotter splitting
- ▶ FLASH-X: Euler–Poisson system with finite-volume and RK methods

$$d_t \mathbf{u} = \mathcal{T}_u(\mathbf{u}, \Phi)$$

$$F(\Phi, \mathbf{u}) = 0$$

- ▶ THORNADO: Two-moment model with DG and IMEX-RK methods
  - ▶ Phase-space advection (explicit)

$$d_t \mathcal{U} = \mathcal{T}_{\mathcal{U}}(\mathcal{U}, \mathbf{u})$$

- ▶ Collisions (implicit)

$$d_t \mathbf{u} = \mathcal{C}_u(\mathcal{U}, \mathbf{u})$$

$$d_t \mathcal{U} = \mathcal{C}_{\mathcal{U}}(\mathcal{U}, \mathbf{u})$$

- ▶ Fluid fields  $\mathbf{u}$  require finite-volume *and* DG representations

## Neutrino-Matter Solver: Moment Update

- Implicit update on primitive moments  $\mathcal{M} = (\mathcal{D}, \mathcal{I}_j)^\top$

$$(\mathcal{D} + v^i \mathcal{I}_i) = \mathcal{N}^n + \Delta t (\eta - \chi \mathcal{D})$$

$$(\mathcal{I}_j + v^i \mathcal{K}_{ij}) = \mathcal{G}_j^n - \Delta t \kappa \mathcal{I}_j$$

$\eta$ ,  $\chi$ , and  $\kappa$  depend on  $\mathcal{M}$ ,  $\bar{\mathcal{M}}$ , and  $\mathbf{u}$

- Modified Richardson iteration with step size  $\lambda = 1/(1 + |\mathbf{v}|)$

$$\mathcal{D}^{[k+1]} = \mathcal{D}^{[k]} + \lambda \frac{[\mathcal{N}^n - (\mathcal{D}^{[k]} + v^i \mathcal{I}_{s,i}^{[k]}) + \Delta t (\eta^{[k]} - \chi^{[k]} \mathcal{D}^{[k]})]}{(1 + \Delta t \chi^{[k]})}$$

$$\mathcal{I}_j^{[k+1]} = \mathcal{I}_j^{[k]} + \lambda \frac{[\mathcal{G}_j^n - (\mathcal{I}_j^{[k]} + v^i \mathcal{K}_{s,ij}^{[k]}) - \Delta t \kappa^{[k]} \mathcal{I}_j^{[k]}]}{(1 + \Delta t \kappa^{[k]})}$$

Realizability-preserving with guaranteed convergence\*

- Write as fixed-point map

$$\mathcal{M}^{[k+1]} = \mathcal{G}(\mathcal{M}^{[k]}, \mathbf{u})$$



## Neutrino-Matter Solver: Fluid Update

- Fluid system for  $\mathbf{u} = (\rho, \rho v_j, \rho \epsilon_f, \rho Y_e)^T$ : Enforce conservation laws

$$\begin{aligned}\rho &= \rho^n \\ \rho v_j &= \rho v_j^n - (S_j - S_j^n) \\ \rho \epsilon_f &= \rho \epsilon_f^n - (E - E^n) \\ \rho Y_e &= \rho Y_e^n - m_b (N - N^n)\end{aligned}$$

$S_j$ ,  $E$ , and  $N$ : **neutrino momentum, energy, and number densities**

- Write as fixed-point map

$$\mathbf{u} = \mathbf{g}(\mathcal{M}, \mathbf{u})$$

# Neutrino-Matter Solver: Nested Algorithm<sup>6</sup>

- ▶ Coupled nonlinear system

$$\mathbf{u} = \mathbf{g}(\mathcal{M}, \mathbf{u}) \quad \text{and} \quad \mathcal{M} = \mathcal{G}(\mathcal{M}, \mathbf{u})$$

- ▶ Solved in nested manner

$$\mathbf{u}^{[k+1]} = \mathbf{g}(\widehat{\mathcal{M}}^{[k]}, \mathbf{u}^{[k]}) \quad (k = 1, \dots, k_{\max}),$$

where  $\widehat{\mathcal{M}}^{[k]}$  is limit point of inner iteration sequence

$$\mathcal{M}^{[k, \ell+1]} = \mathcal{G}(\mathcal{M}^{[k, \ell]}, \mathbf{u}^{[k]}) \quad (\ell = 1, \dots, \ell_{\max}).$$

- ▶ Opacities only evaluated in outer loop
- ▶ Fixed-point iteration avoids Jacobian and solution of dense linear system
- ▶ Easy to implement and extend for additional opacities
- ▶ Anderson acceleration can be applied separately to outer and inner loops

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<sup>6</sup>Laiu, E, Chu, Harris, Messer, ApJS, 253:52

## Anderson Accelerated Fixed-Point Method<sup>7</sup>: $\mathbf{u} = \mathbf{g}(\mathbf{u})$

$$\mathbf{u}^{[1]} = \mathbf{g}(\mathbf{u}^{[0]}); \mathbf{F}^{[0]} = \mathbf{u}^{[0]} - \mathbf{g}(\mathbf{u}^{[0]});$$

**for**  $k = 1, \dots, k_{\max}$  **do**

$$m_k = \min(k, M);$$

$$\mathbf{F}^{[k]} = \mathbf{u}^{[k]} - \mathbf{g}(\mathbf{u}^{[k]});$$

**Solve**

$$\min_{\alpha_j} \left\| \sum_{j=0}^{m_k} \alpha_j \mathbf{F}^{[k-m_k+j]} \right\| \quad \text{subject to} \quad \sum_{j=0}^{m_k} \alpha_j = 1$$

**Update**

$$\mathbf{u}^{[k+1]} = \sum_{j=0}^{m_k} \alpha_j \mathbf{g}(\mathbf{u}^{[k-m_k+j]})$$

**end**

**Algorithm 1:** Anderson Accelerated Fixed-Point Iteration

- ▶ Uses information from previous iterates to improve convergence rate
- ▶ Memory  $M$  typically small. We use  $M = 2$  or  $3$  ( $M = 1$  is Picard iteration)

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<sup>7</sup>Toth & Kelley (2015), SIAM J. Numer. Anal, 53, 805

# Collisional Relaxation

- ▶ Space homogeneous,  $\nu_e + \bar{\nu}_e$ , tabulated Bruenn 85 opacities
  - ▶ Emission/Absorption, Iso-energetic scattering, NES, and Pairs
- ▶ **Goals:** (i) Relaxation to equilibrium, (ii) iteration counts, and (iii) GPU timings

## Problem Specifications

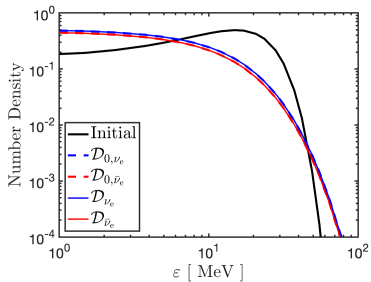
- ▶  $\Omega^\varepsilon = [\varepsilon_{\min}, \varepsilon_{\max}] = [0, 300]$  MeV
- ▶ Gaussian initial spectrum

$$\mathcal{D}_0(\varepsilon) = \frac{1}{2} \times \exp \left[ - \frac{(\varepsilon - 2k_B T)^2}{200 \text{ MeV}} \right]$$

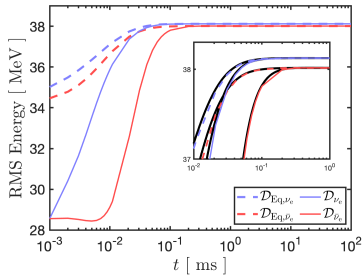
Forward-isotropic distribution with  $|\mathcal{I}_0|/\mathcal{D}_0 = 0.5$

- ▶ Initial matter states with low and high collisionality
  - ▶  $\rho_0 = 10^{12} \text{ g cm}^{-3}$ ,  $\mathbf{v}_0 = (0.1 c, 0, 0)^T$ ,  $T_0 = 7.6 \text{ MeV}$ ,  $Y_{e,0} = 0.14$
  - ▶  $\rho_0 = 10^{14} \text{ g cm}^{-3}$ ,  $\mathbf{v}_0 = (0.1 c, 0, 0)^T$ ,  $T_0 = 15 \text{ MeV}$ ,  $Y_{e,0} = 0.27$
- ▶  $N^\varepsilon = 16$  (geometric;  $\Delta\varepsilon_1 = 1.9 \text{ MeV}$ )
- ▶ Evolve to equilibrium:  $t = 100 \text{ ms}$ .  $\Delta t = 10^{-3} \text{ ms}$  (low) and  $t = 1 \text{ ms}$  (high).  $\Delta t = 10^{-3} \text{ ms}$

# Collisional Relaxation

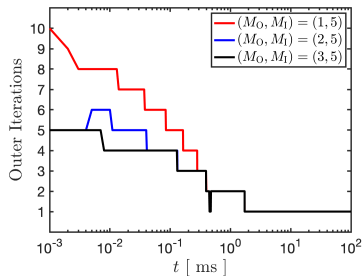


(a) Initial and Final Spectra

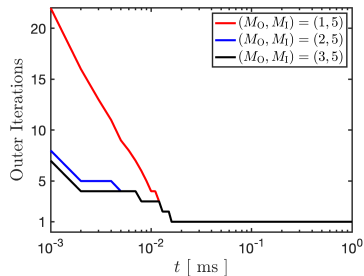


(b) Evolution of RMS Energies

# Collisional Relaxation: Iteration Counts

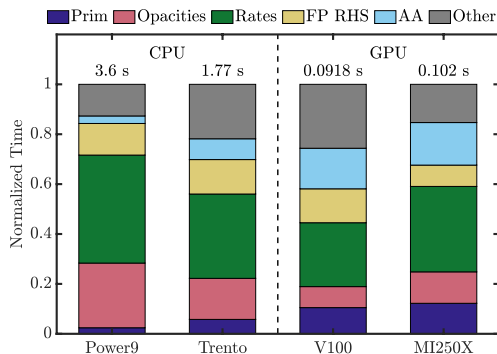


(a) Outer Iterations (Low Collisionality)



(b) Outer Iterations (High Collisionality)

# Collisional Relaxation: GPU Timings

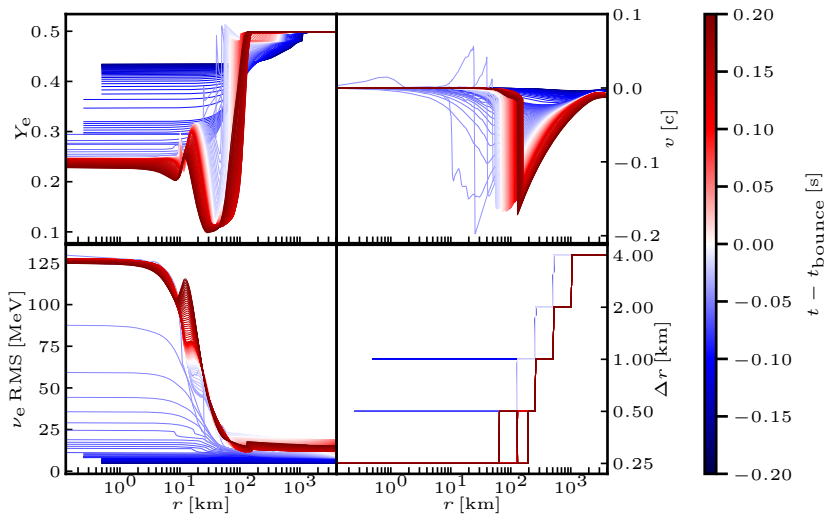


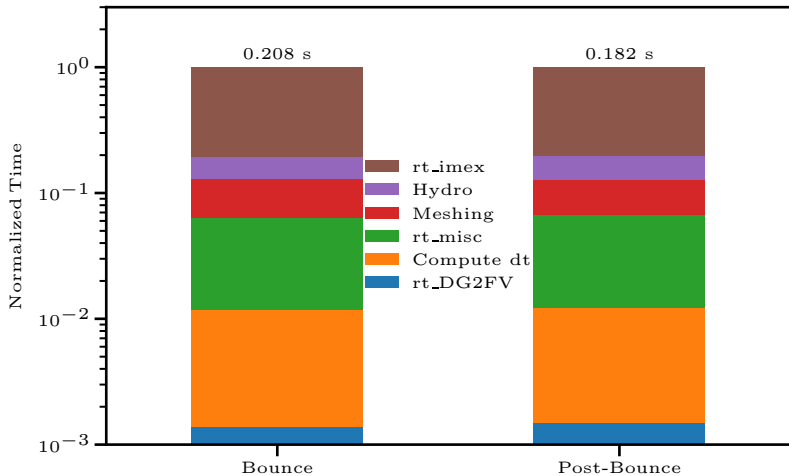
- ▶ One  $8^3 \times 16$  grid block, 2 nodes per phase-space dimension
- ▶ Bruenn 85 + Bremsstrahlung (HR98), six neutrino species
- ▶ **CPU:**
  - ▶ Summit: 7 OpenMP threads on 7 Power9 cores; NVIDIA compiler (22.5), ESSL libs
  - ▶ Frontier: 7 OpenMP threads on 7 AMD Trento; CCE compiler (15.0.1), Cray LibSci
- ▶ **GPU:**
  - ▶ Summit: NVIDIA V100 with OpenACC; NVIDIA compiler (22.5) and libs
  - ▶ Frontier: AMD MI250X with OpenMP OL; CCE compiler (15.0.1), ROCm libs (5.4.0)

# Collapse and Post-Bounce Evolution with FLASH-X+THORNADO

- ▶ 15  $M_{\odot}$  progenitor from Woosley & Heger (2007)
- ▶ Spark hydrodynamics from FLASH-X
- ▶ Spectral, two-moment neutrino transport from THORNADO
  - ▶ Six species, 16 linear elements in  $\epsilon \in [0, 300]$  MeV
  - ▶ Updated WEAKLIB opacities
- ▶ SFHo EoS tabulated with WEAKLIB
- ▶ Five AMR levels  $\Delta r = 4 - 0.25$  km







▶  $T_{\text{Trans}}/T_{\text{HD}} \sim 12$

▶ Collisions about 3 times more expensive than advection in rt-imex

# Summary

- ▶ DG-IMEX method for  $\mathcal{O}(v/c)$  two-moment model in THORNADO
  - ▶ Neutrino-matter coupling algorithm
  - ▶ Ported to use GPUs with OpenMP or OpenACC
- ▶ Interface to multi-physics simulation framework FLASH-X
  - ▶ Simulate neutrino transport in CCSN models with DG methods
- ▶ Ongoing
  - ▶ Multi-dimensional simulations
  - ▶ General relativistic model
  - ▶ Improvements to neutrino weak interaction physics (e.g., muons, inelastic scattering on nucleons)