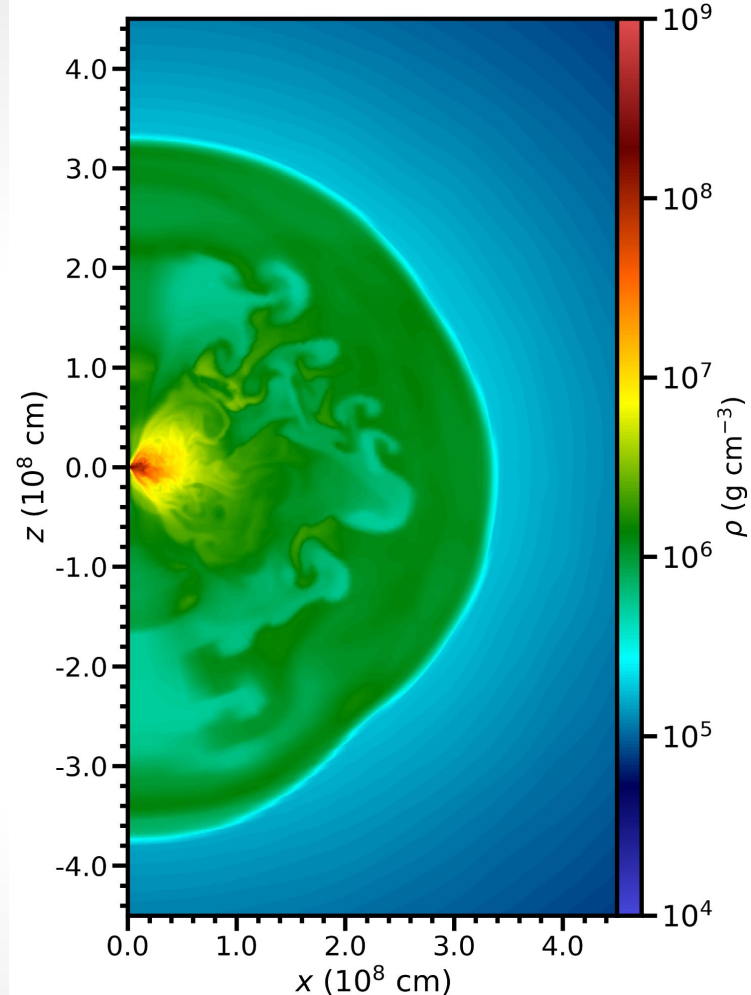


Dynamics of Collapsar Disk Outflows

Coleman Dean and
Rodrigo Fernández



Collapsars: Rotating Core Collapse Supernovae

Wolf-Rayet progenitor (low metallicity)

Core collapse

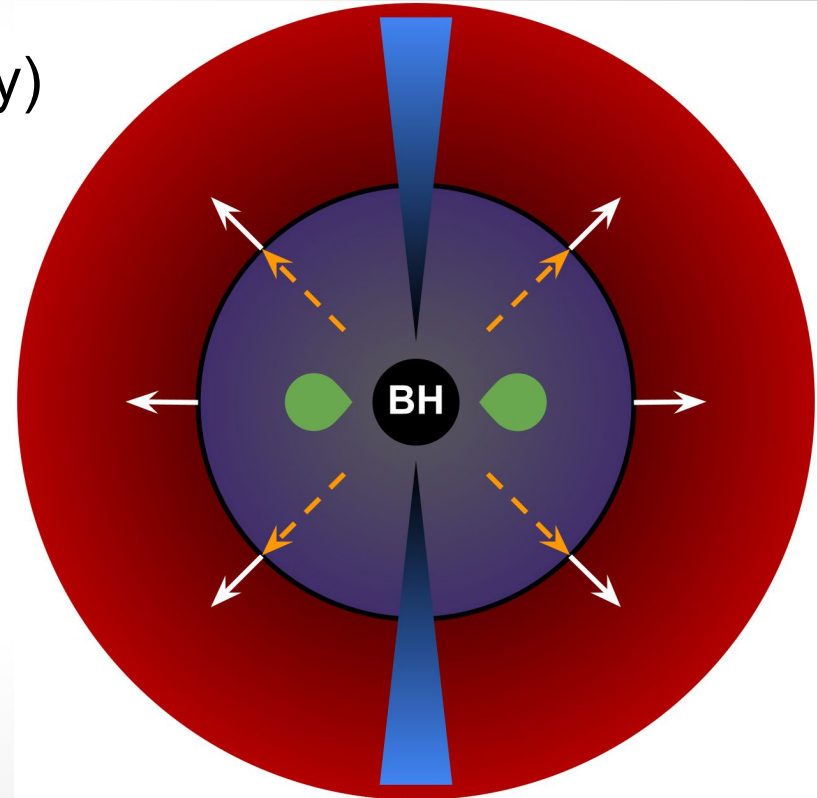
black hole - accretion disk system

Short timescale heavy element

production (Mathews and Cowen 1990)

BNS: $t_{\text{delay}} \sim 100 \text{ Myr}$

Collapsar: $t_{\text{delay}} \sim 1 \text{ Myr}$



Heavy Element Production

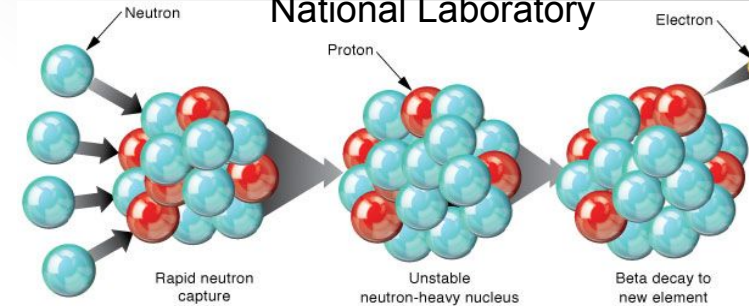
Rapid neutron capture process
(r-process)

($\sim 1/2$ of elements heavier than iron)

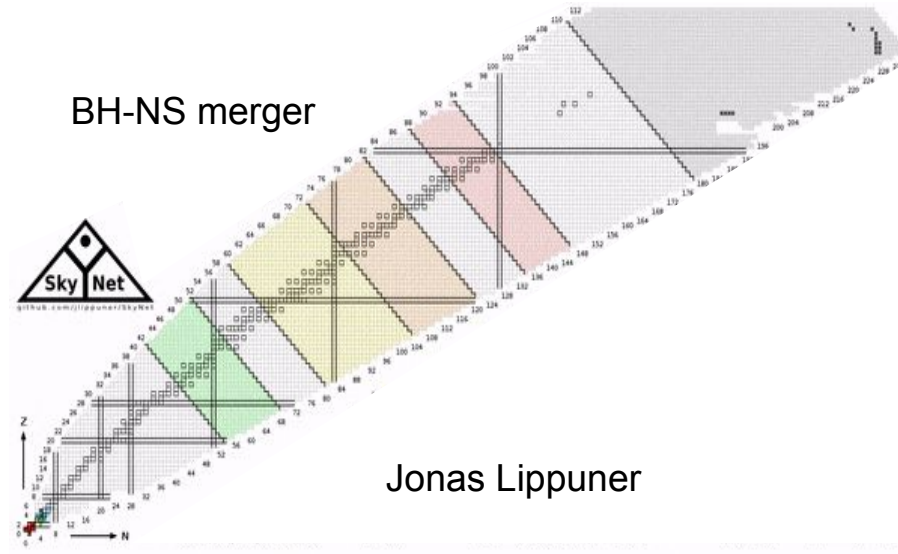
$Y_e < 0.25$ for heavy r-process
elements

Neutronization

Lawrence Livermore
National Laboratory



BH-NS merger



Jonas Lippuner

Computational Challenge

Large range of spatial and temporal scales

Magnetic effects

(turbulence, heating of disk)

Neutrino effects

(neutronization, r-process)

General relativistic effects

(potential around BH)

$t_{\text{orb,isco}}$	~ 10 ms
$t_{\text{shock breakout}}$	~ 100 s
r_{ISCO}	$\sim 10^6$ cm
$R_{\text{progenitor}}$	$\sim 10^{11}$ cm

Recent similar works:

Just et al. 2022

Fujibayashi et al. 2022

Fujibayashi et al. 2023

Our Model

Model	$M_{\text{ZAMS}} (M_{\odot})$	$M_{\text{cc}} (M_{\odot})$	$Z (Z_{\odot})$
16TI	16	13.95	0.01
35OC	35	28.07	0.1

Long term disk outflow simulations (no relativistic jet)

Two-dimensional Hydrodynamics (Helmholtz EOS)

Log radial grid - $\cos \theta$ grid

Newtonian multipole self-gravity + Artemova ($L = 0$)

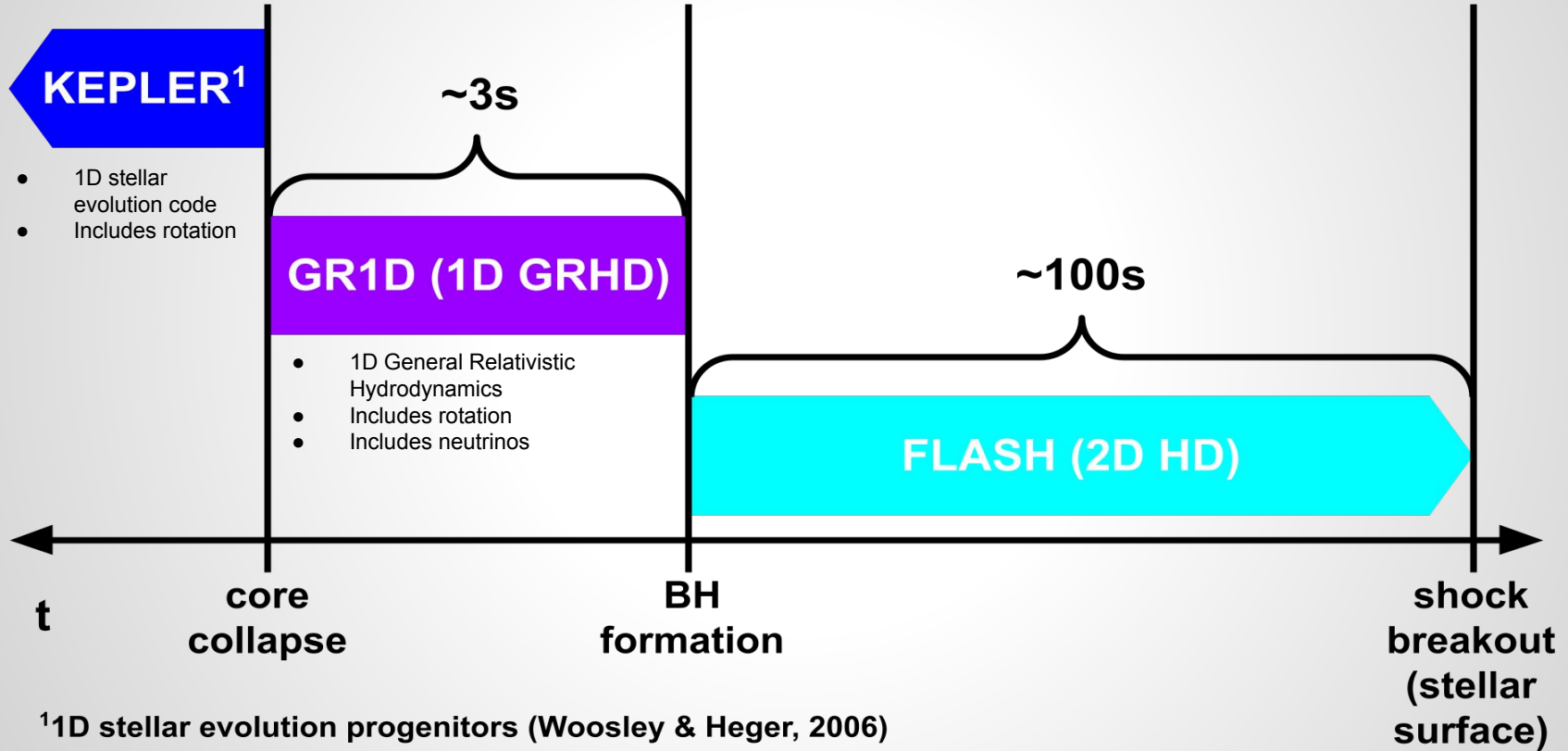
∞ -viscosity (turbulence \rightarrow heating in the disk)

(Shakura & Sunyaev 1973)

3 species neutrino leakage (annular lightbulb) + absorption

19 isotope nuclear network + Nuclear Statistical Equilibrium solver

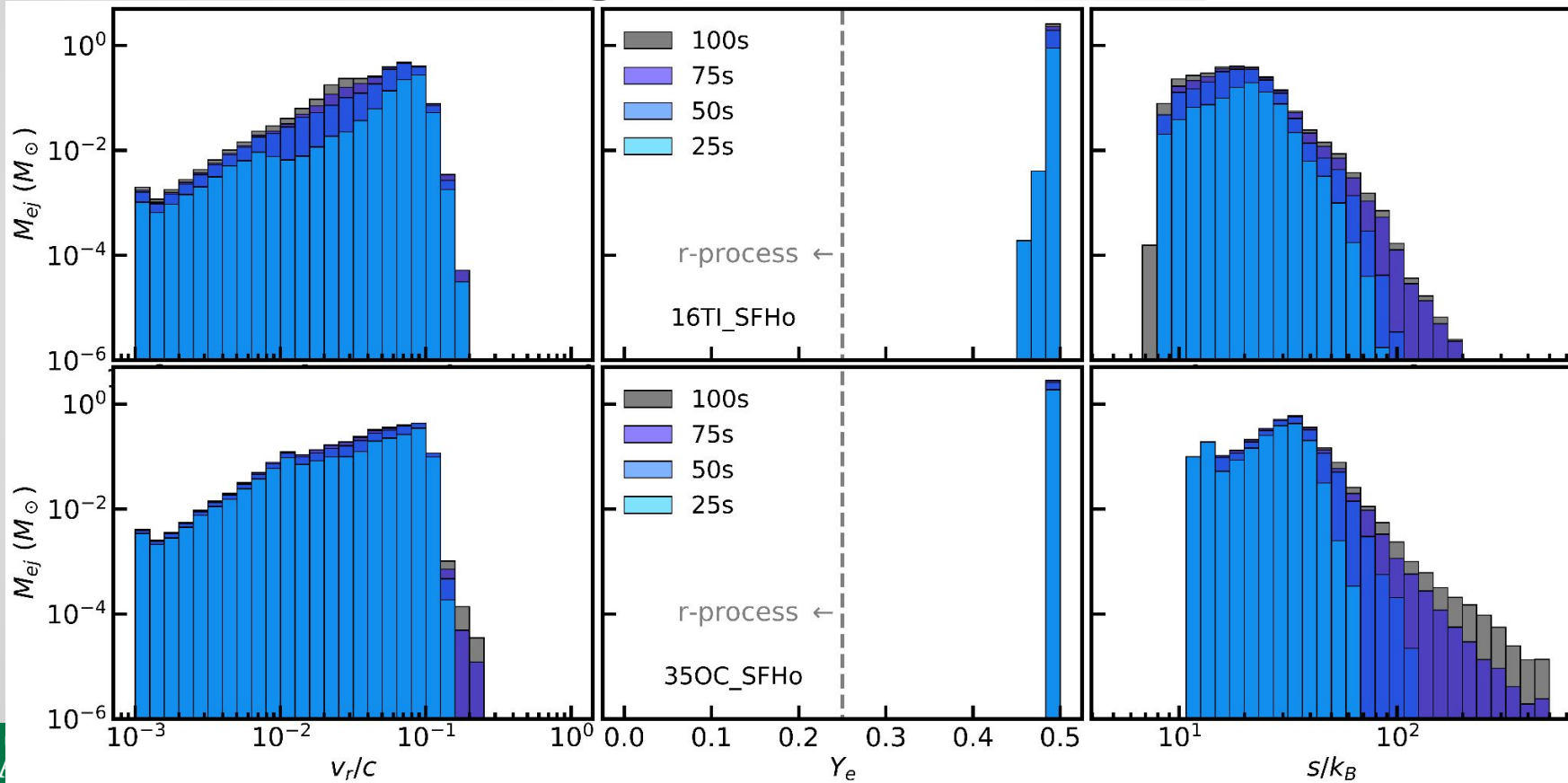
Our Model



¹1D stellar evolution progenitors (Woosley & Heger, 2006)

Mass Tracking

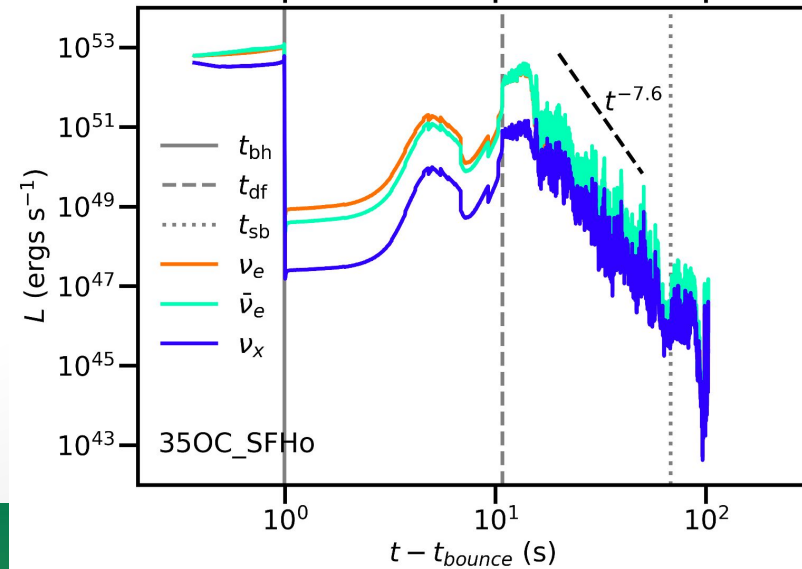
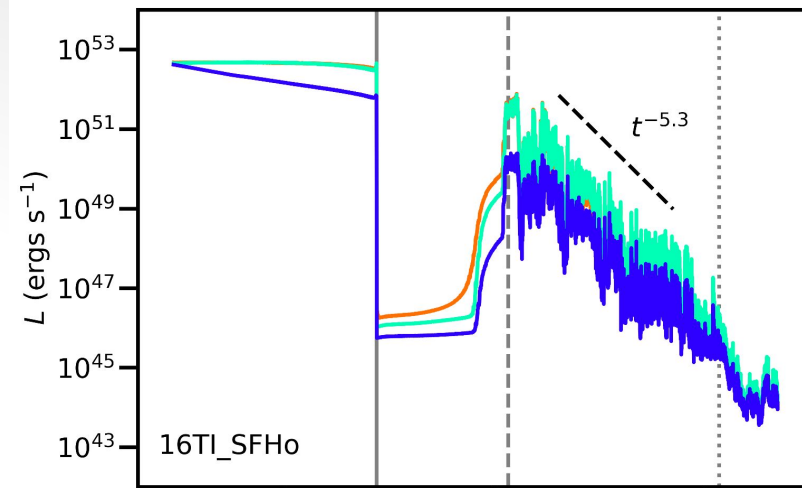
$$r_{ej} = 10^9 \text{ cm}$$



Neutrino Luminosity

Delay time measures
angular momentum
distribution

Height of second peak
depends on M/R ,
accretion rate



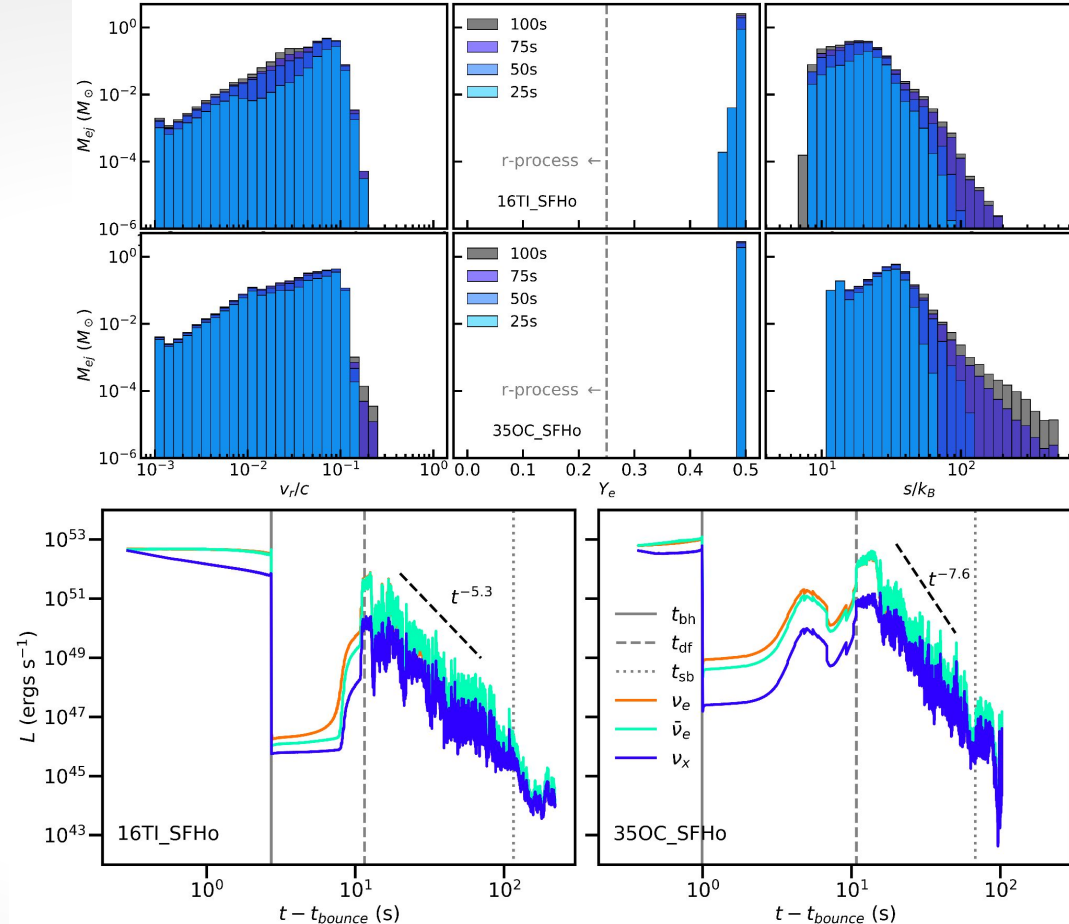
Conclusion

Insufficient neutronization in the disk wind

Neutrino Luminosity delay time diagnostic of angular momentum profile

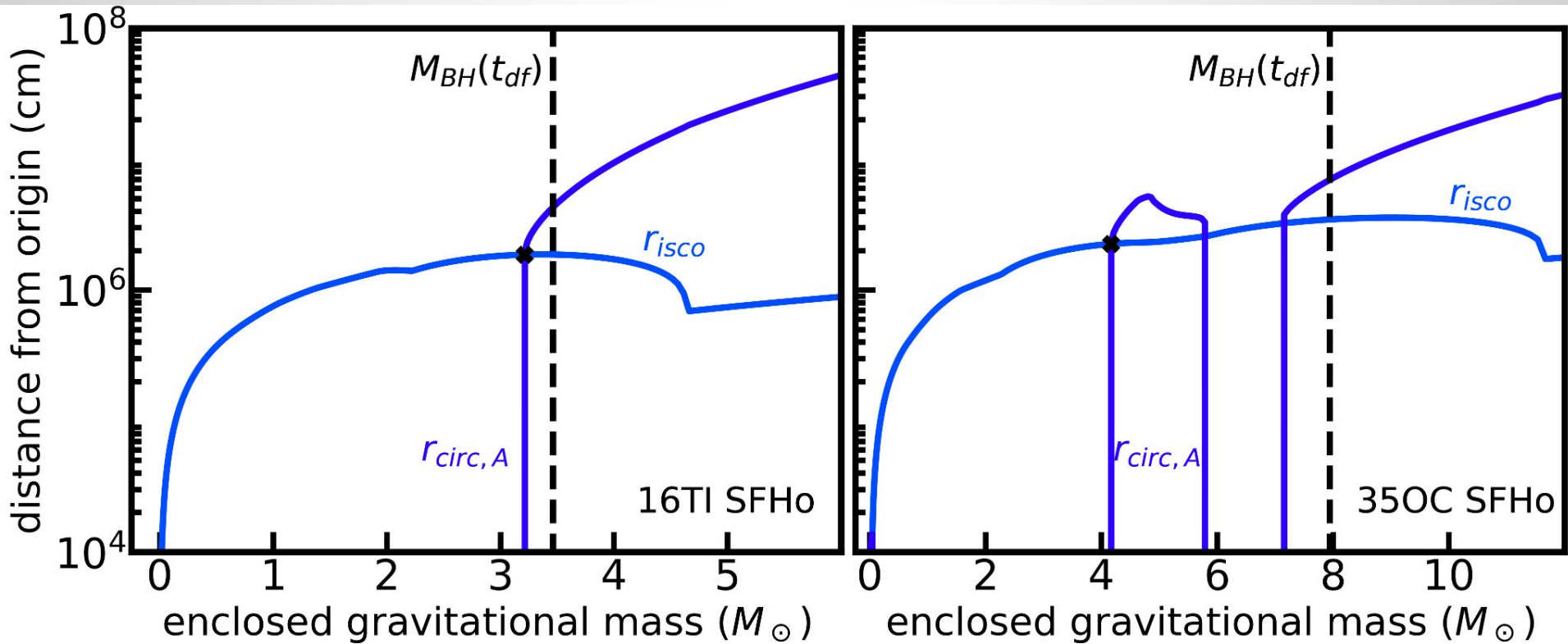
Disk wind leads to shock breakout from the star

Sufficient Ni production for type Ibc explosion

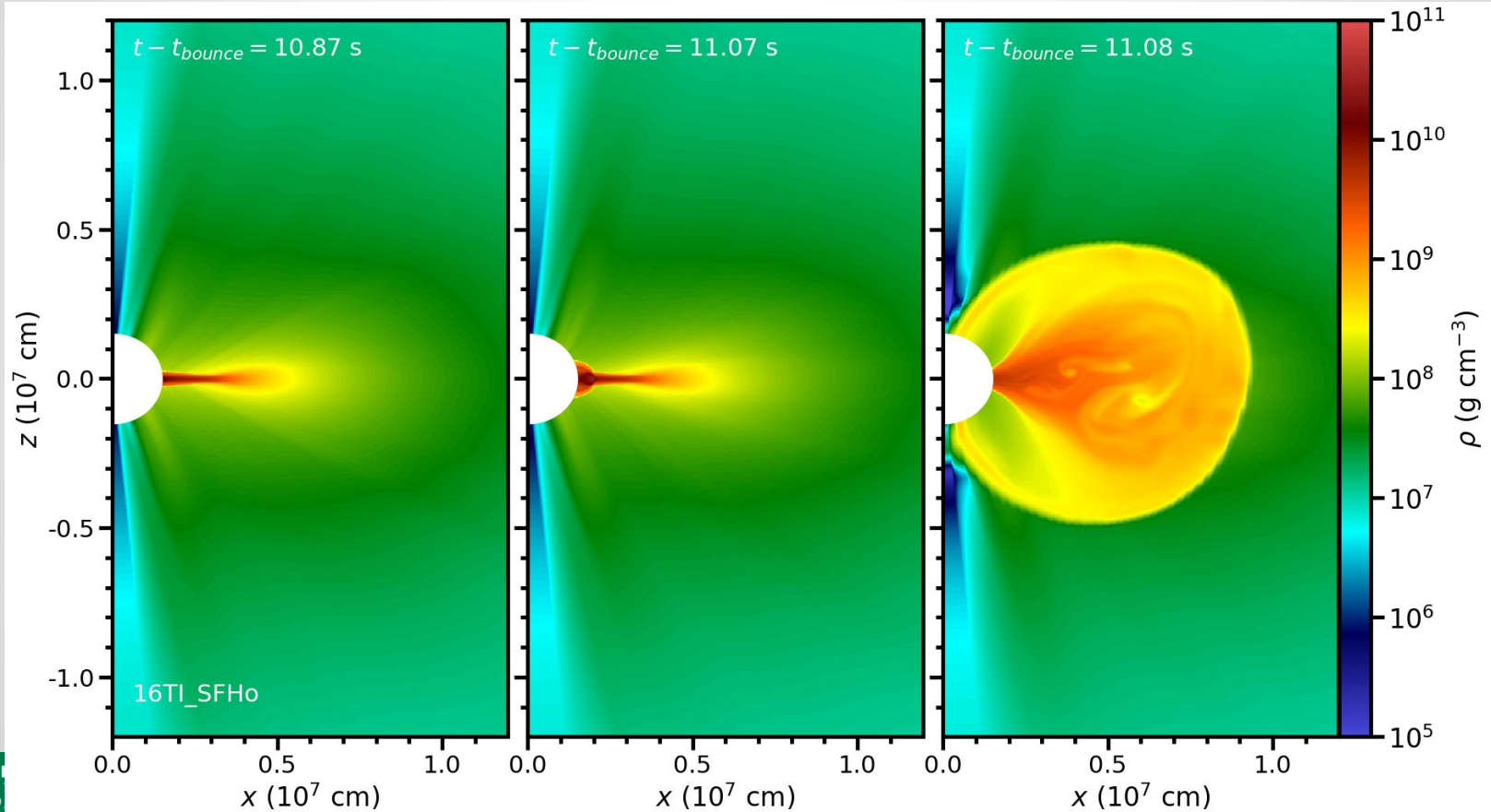


Extra Slides

Predicting accretion disk formation



“Dwarf Disk” Formation



Density colour-plot

16Ti progenitor star
(Woosley & Heger 2006)

Shock breaks out at

$$t_{\text{sb}} \sim 116 \text{ s}$$

Electron Fraction

Orange represents

$$Y_e < 0.25$$

Fed by $Y_e = 0.5$ material

Density contours

CCSNe as r-process source

Early GCE simulations of Eu suggest CCSNe necessary to explain low metallicity halo stars Eu abundance (i.e. Mathews and Cowen 1990)

Significant scatter in Eu abundances (i.e. McWilliam et al. 1995)

Hierarchical galaxy merger / low star formation efficiency might explain Eu abundances on BNS timescales (i.e. Ishimaru et al. 2015)

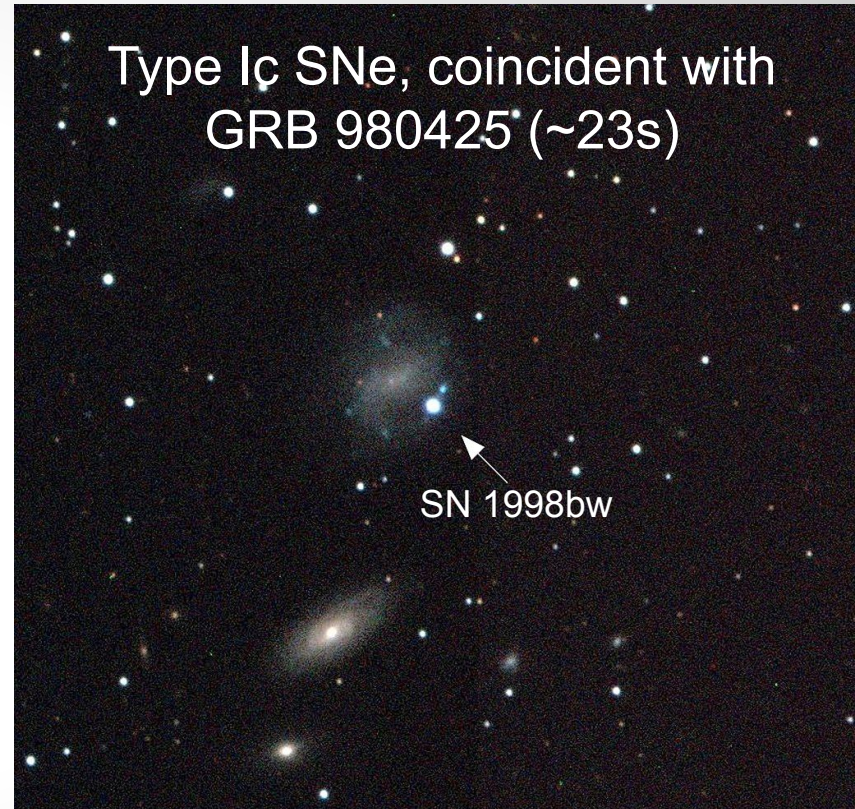
Type Ic SNe

No hydrogen lines

Weak/no helium lines

Core collapse of a stripped star
such as a wolf-rayet star

Some coincident with long GRBs



Type Ic SNe, coincident with
GRB 980425 (~23s)

By ESO - <http://www.eso.org/public/images/eso9847a/>, CC BY 4.0,
<https://commons.wikimedia.org/w/index.php?curid=15158163>

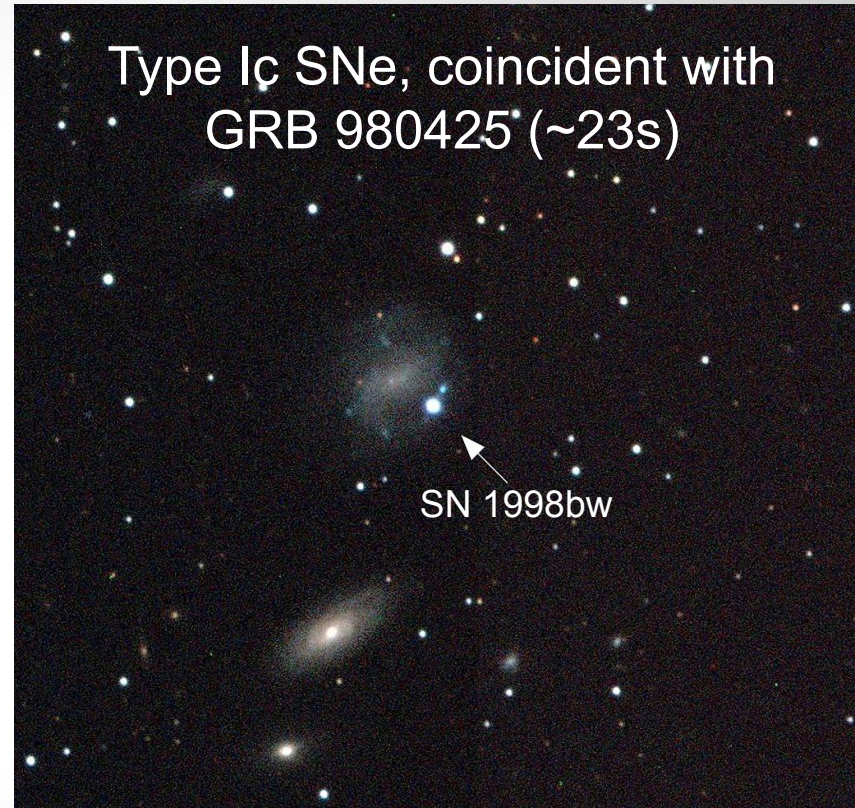
GRB Progenitors

IGRB's ($\gtrsim 2\text{s}$)

- Near Star Formation
- Death of massive stars
- Long duration GRB “engine”

Collapsars as a progenitor?

- 11 GRB-SNe (Modjaz et al. 2016)



By ESO - <http://www.eso.org/public/images/eso9847a/>, CC BY 4.0, <https://commons.wikimedia.org/w/index.php?curid=15158163>

Progenitor Length Scale

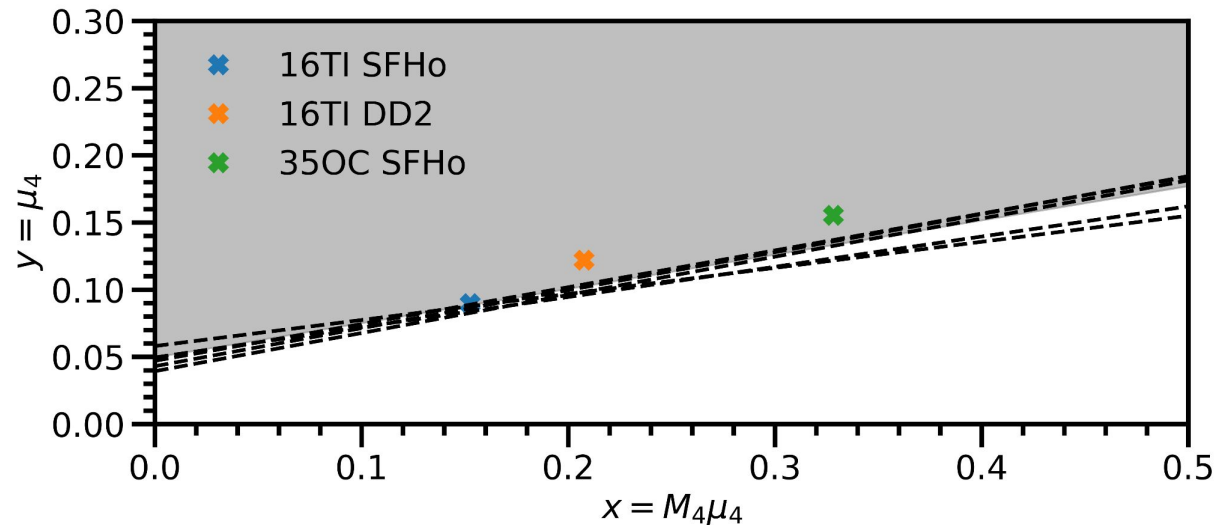


WR:
R~1-20 R_{\odot}
T_{esc}~2-40s

RSG:
R~1500 R_{\odot}
T_{esc}~3000s

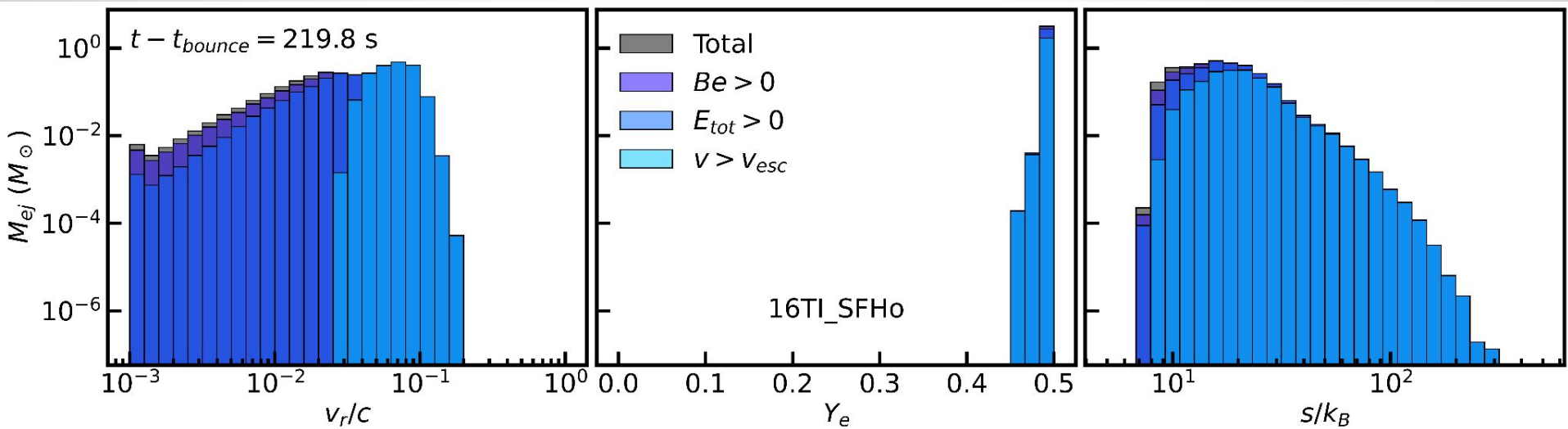
Two parameter explosion criteria

Ertl et al. 2016



350C_SFHo

Density colour-plot



Motivation

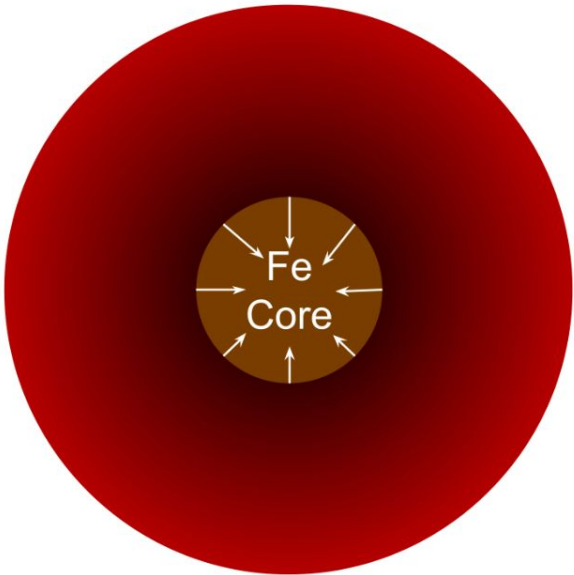
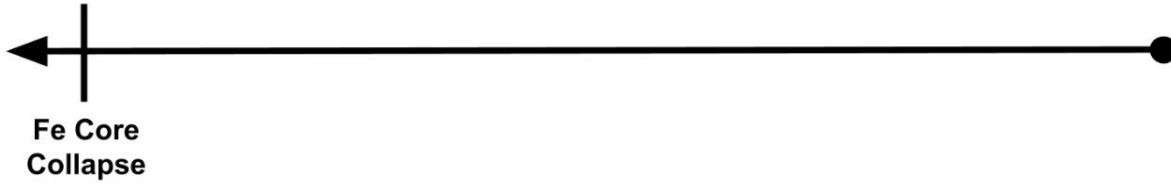
Short timescale heavy element production (Mathews and Cowen 1990)

Binary neutron star merger: $t_{\text{delay}} \sim 100 \text{ Myr}$

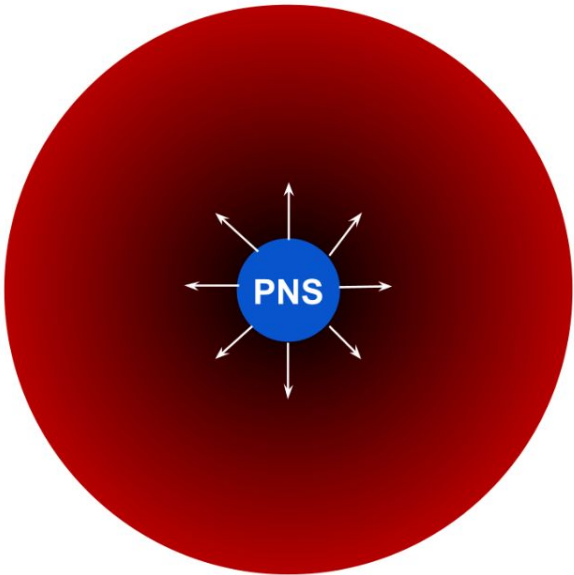
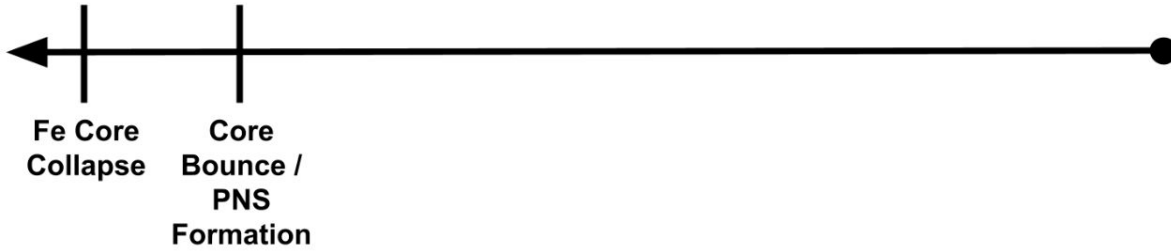
Collapsar: $t_{\text{delay}} \sim 1 \text{ Myr}$



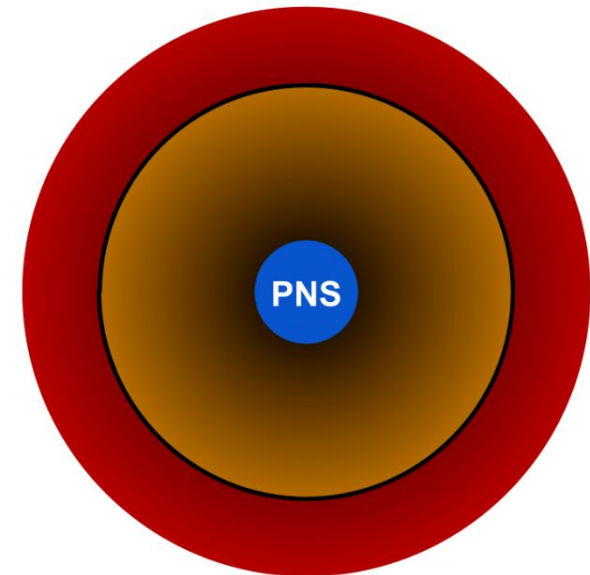
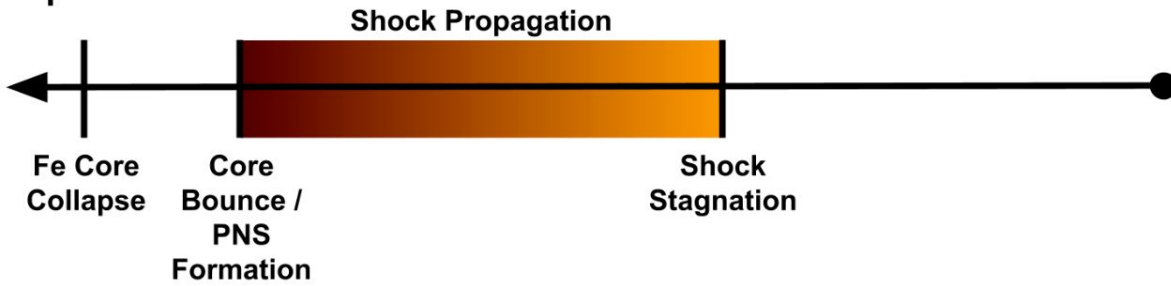
Failed Core
Collapse
Supernova



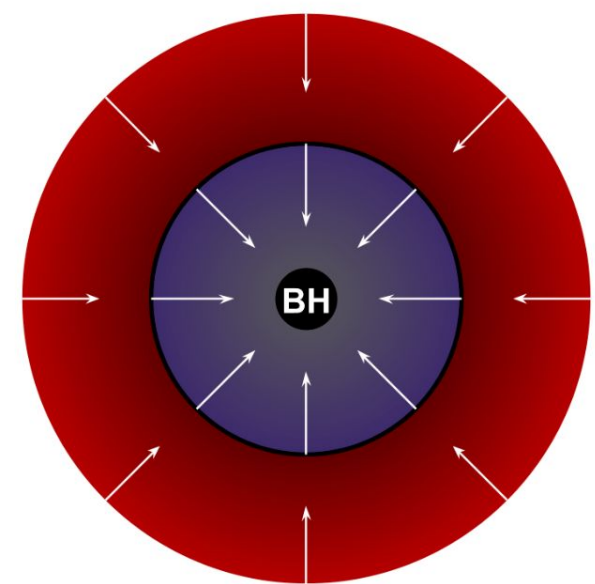
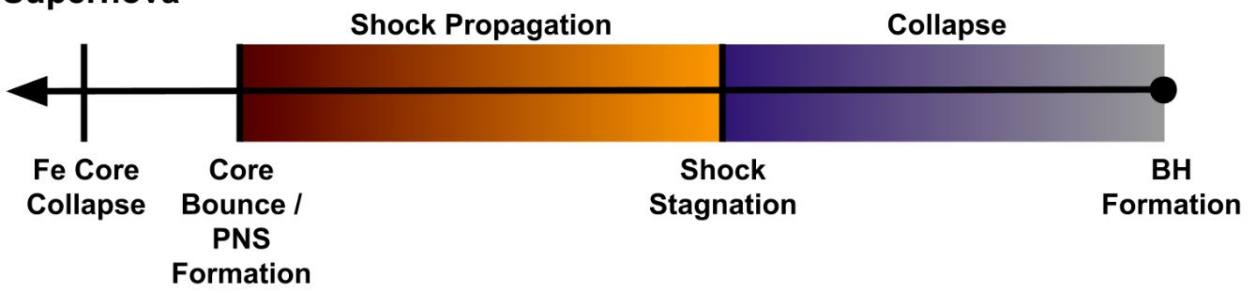
Failed Core Collapse Supernova



Failed Core Collapse Supernova

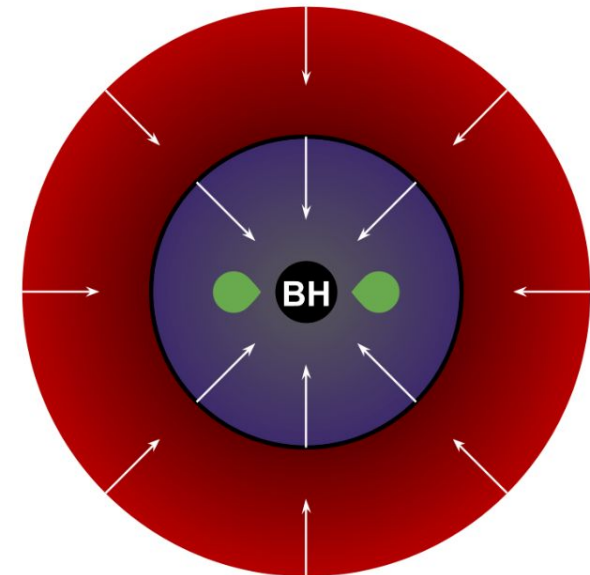
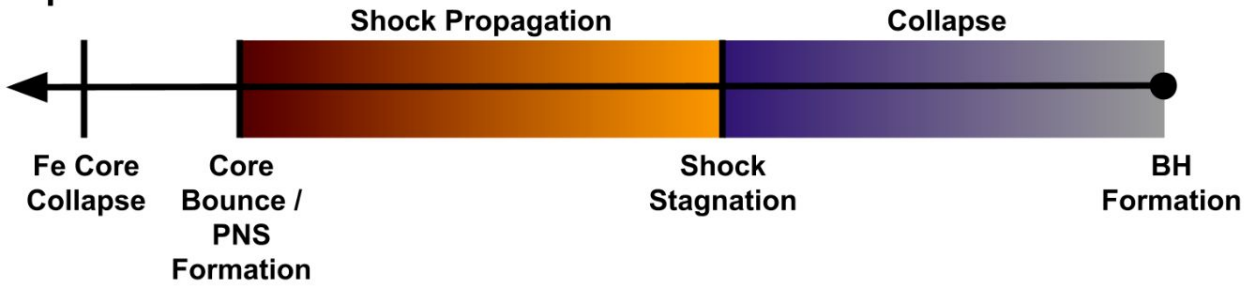


Failed Core Collapse Supernova



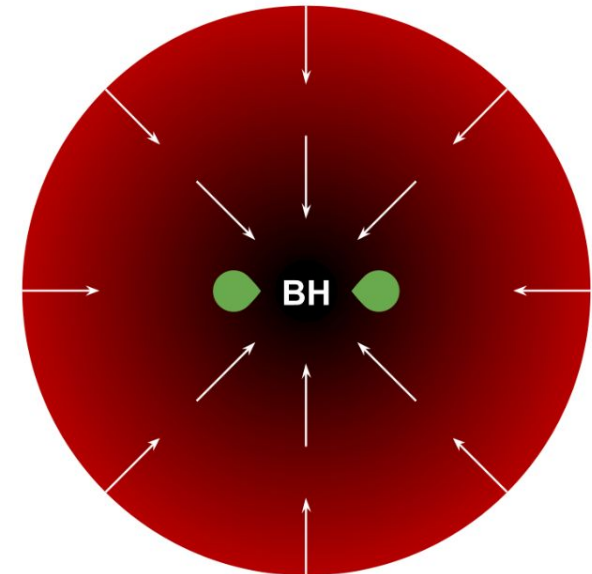
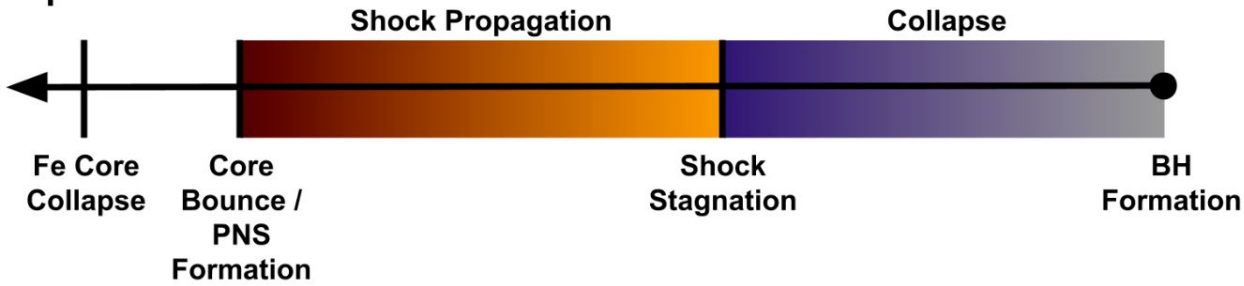
Failed Core Collapse

Supernova



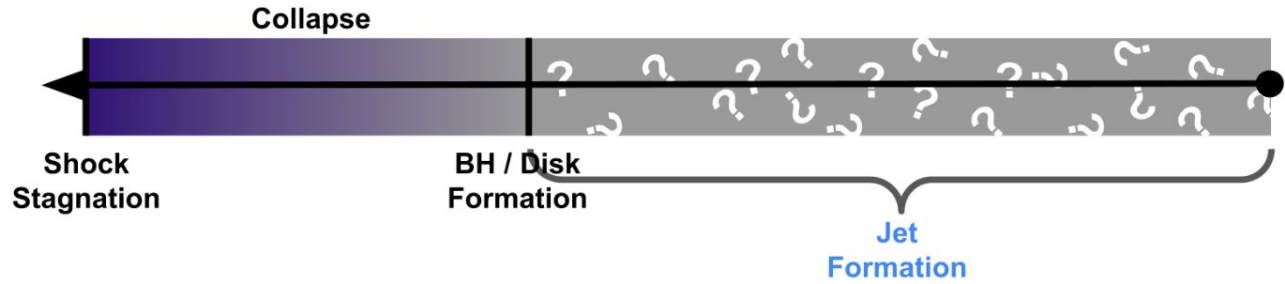
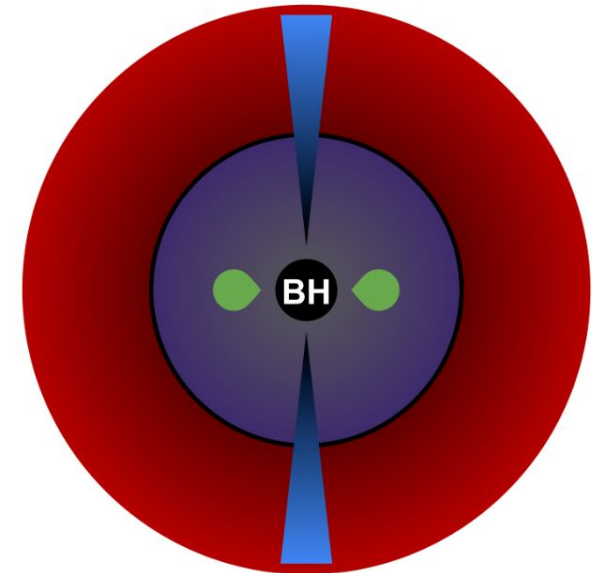
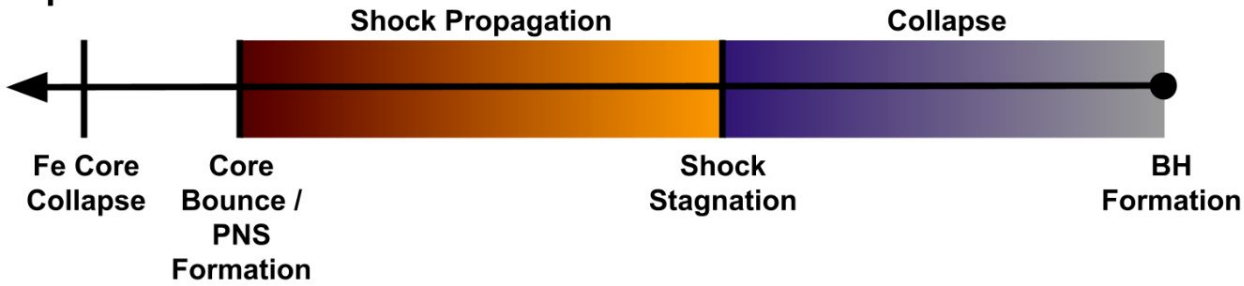
Failed Core Collapse

Supernova



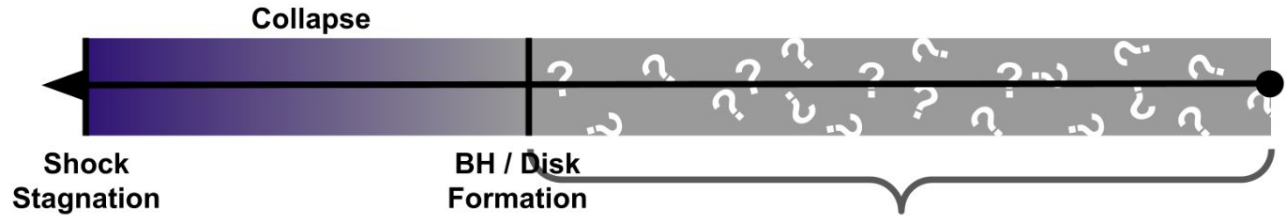
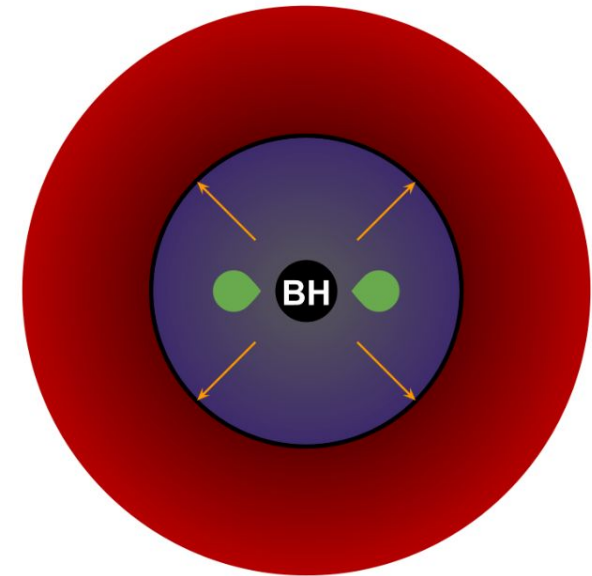
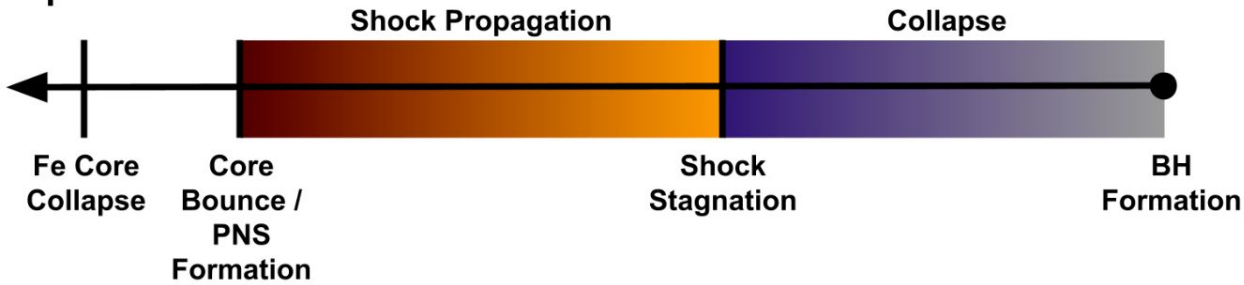
Failed Core Collapsar

Collapse Supernova



Failed Core Collapsar

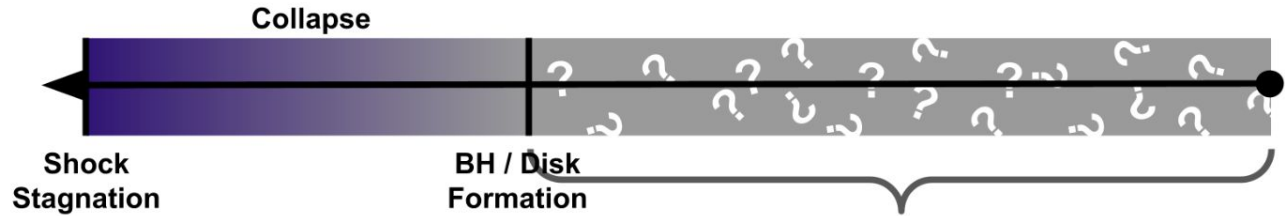
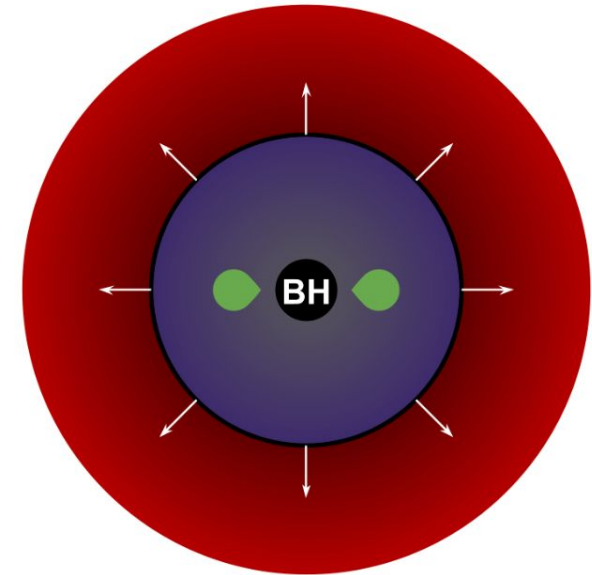
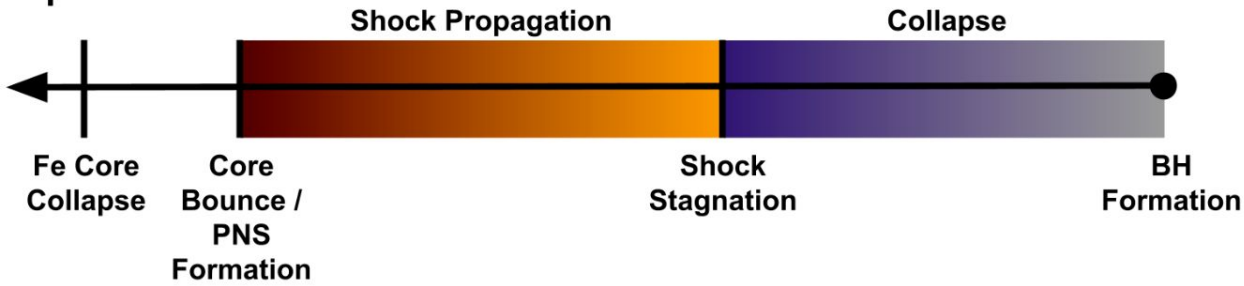
Collapse Supernova



Disk Wind
Launched

Failed Core Collapsar

Collapse Supernova



SN
Revived

Failed Core Collapse

Supernova

