

Testing collapse models and gravity with levitated optomechanics

ECT* workshop

Quantum sensing and fundamental physics with levitated mechanical systems

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UNIVERSITÀ
DEGLI STUDI
DI TRIESTE



Outline

Collapse models

Gravity models

Motivation

Theory

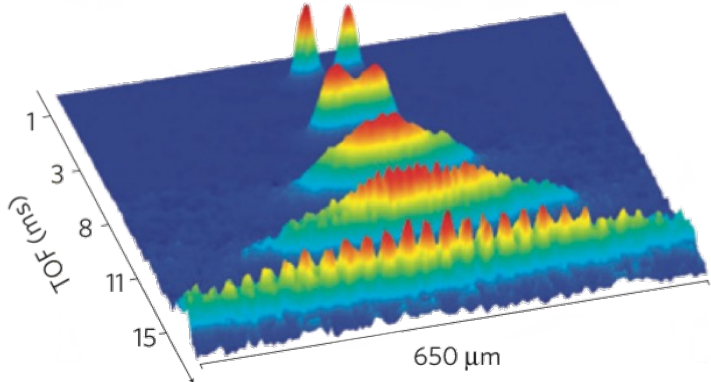
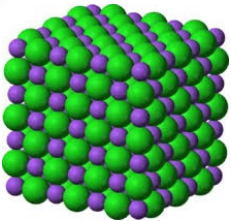
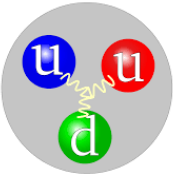
Experiments

Collapse models

Limits of the Quantum Superposition Principle

Quantum World Classical World

Micro

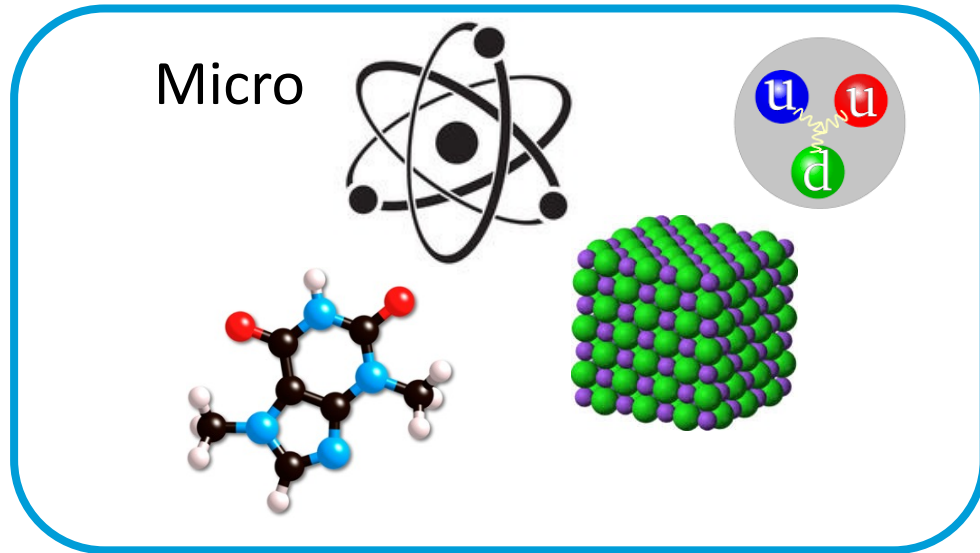


Macro



Standard Quantum Mechanics

Quantum World



Classical World



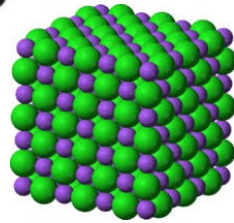
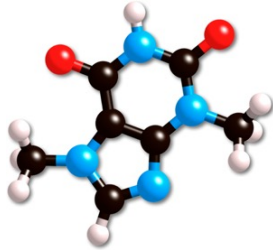
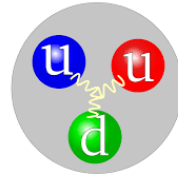
“The Copenhagen interpretation assumes a mysterious division between the microscopic world governed by quantum mechanics and a macroscopic world of apparatus and observers that obeys classical physics.”

Quantum Mechanics wanna be

Quantum World

Classical World

Micro



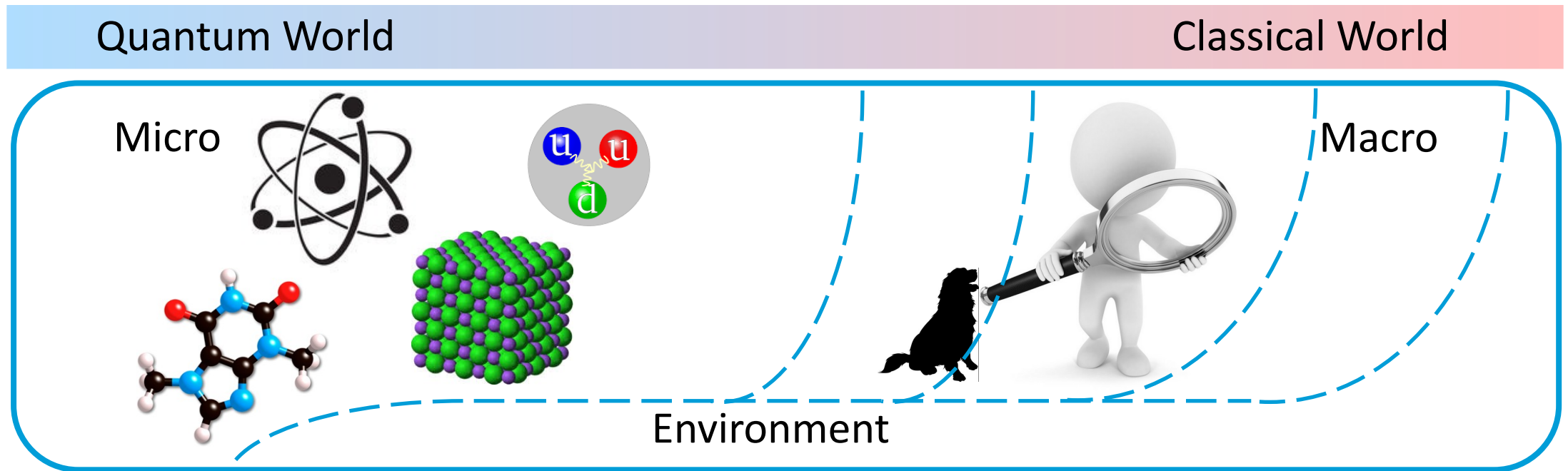
Macro



“What exactly qualifies some physical systems to play the role of 'measurer'?”

John Bell, Against 'measurement', *Physics World*, *Phys. World* **3** (8) 33 (1990)

Quantum Mechanics + Decoherence



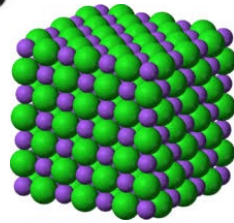
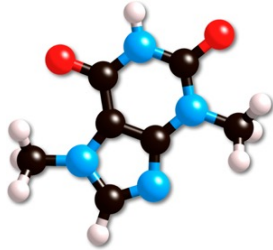
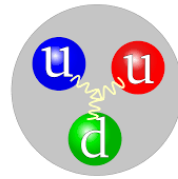
The division system-environment is arbitrary, and similarly to the division quantum-classical in the Copenhagen interpretation.

Possible solutions

Quantum World

Classical World

Micro



Macro



Bohmian Mechanics

Many Worlds

Collapse Models

Collapse models – a modified quantum theory

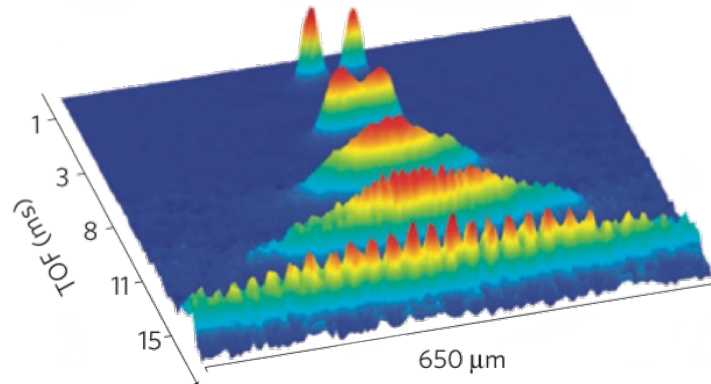
Stochastic noise

Non-linear wavefunction dynamics

Negligible microscopic action
No effective collapse
Quantum systems

Strong macroscopic action
Rapid collapse
Systems behave classically

Amplification mechanism



General structure of the dynamical equation

$$d|\psi_t\rangle = \left[-\frac{i}{\hbar}\hat{H}dt + \int d^3\mathbf{x} (\hat{M}(\mathbf{x}) - \langle\hat{M}(\mathbf{x})\rangle_t) dW_t(\mathbf{x}) - \frac{1}{2} \int d^3\mathbf{x} d^3\mathbf{y} \mathcal{D}(\mathbf{x} - \mathbf{y}) \prod_{\mathbf{q}=\mathbf{x},\mathbf{y}} (\hat{M}(\mathbf{q}) - \langle\hat{M}(\mathbf{q})\rangle_t) dt \right] |\psi_t\rangle$$

Q. Hamiltonian

Measured operator
(mass density, position, ...)

Non-linear term

White noise

Noise's spatial correlation

Two main models

Continuous Spontaneous Localization (CSL) model

Fully phenomenological model

$$\mathcal{D}_{\text{CSL}}(\mathbf{x} - \mathbf{y}) = \frac{\lambda}{m_0^2} \exp(-|\mathbf{x} - \mathbf{y}|^2/4r_C^2)$$

 λ collapse rate r_C correlation length

Diósi-Penrose model

Gravity-related model

$$\mathcal{D}_{\text{DP}}(\mathbf{x} - \mathbf{y}) = \frac{G}{\hbar} \frac{1}{|\mathbf{x} - \mathbf{y}|}$$

 R_0 spatial cutoff
gravity regularization at small distances

Experiments

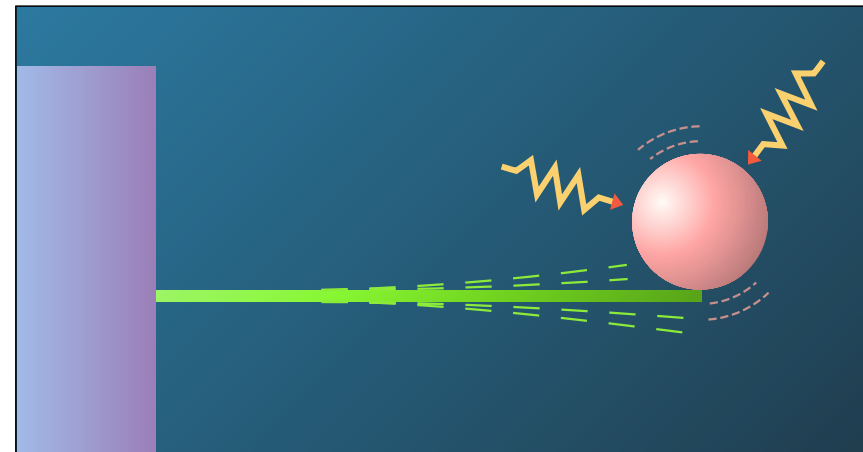
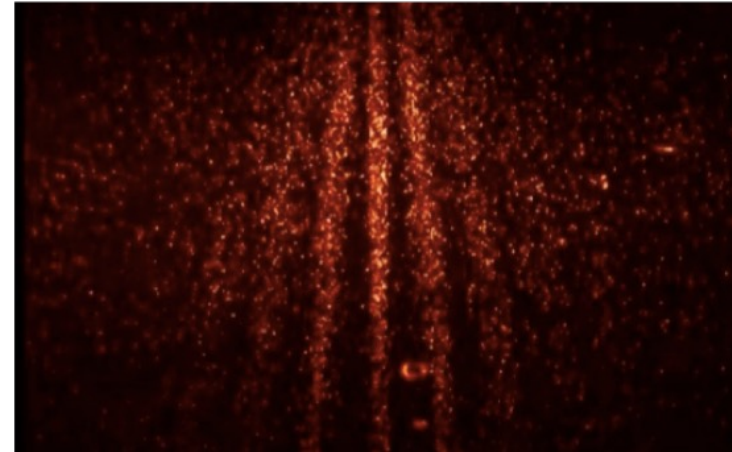
Destruction of quantum superposition

Interferometric Experiments

$$\Delta V = \frac{V_{\max} - V_{\min}}{V_{\max} + V_{\min}}$$

Extra jigglng due to collapse noise

$$S_{xx}(\omega) = \frac{1}{4\pi} \int d\Omega \langle \{ \tilde{x}(\omega), \tilde{x}(\Omega) \} \rangle$$



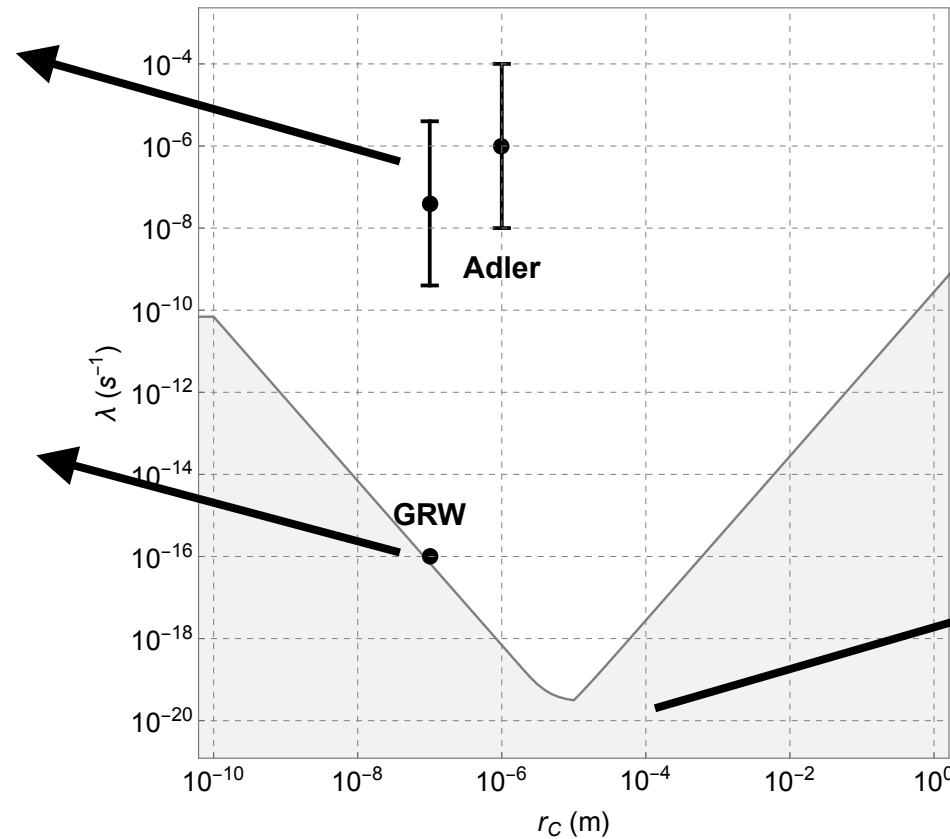
Parameters' space

Adler's proposed values

Adler, J. Phys. A **40**, 2935 (2007)

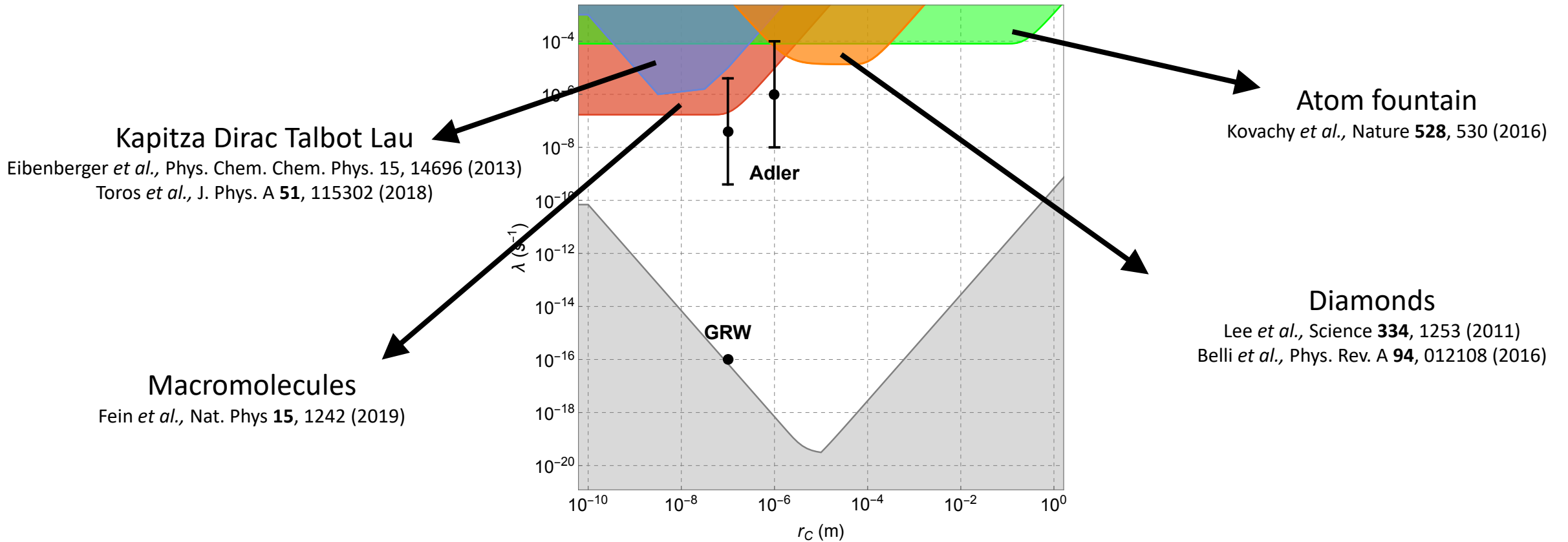
GRW's proposed values

Ghirardi *et al.*, Phys. Rev. D **34**, 470 (1986)



Theoretical lower bound
Toros *et al.*, Phys. Lett. A **381**, 3921 (2017)

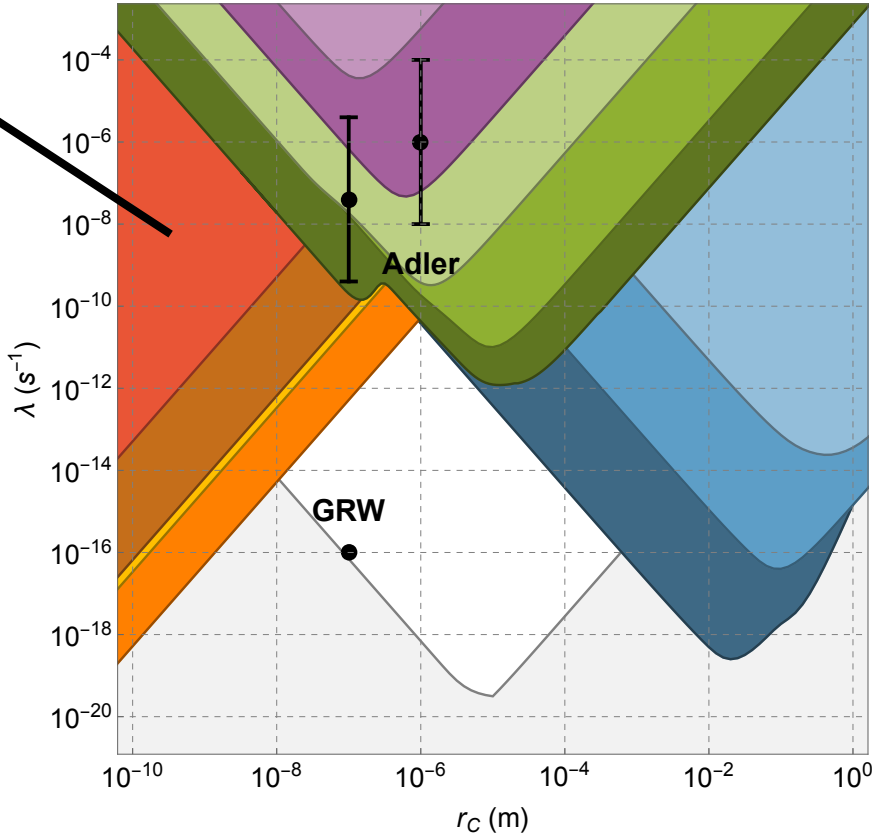
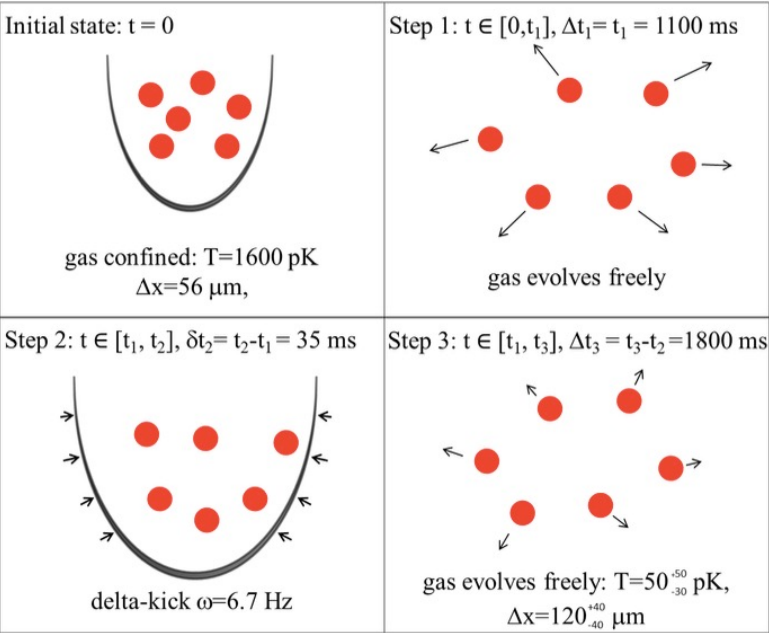
Interferometric Tests



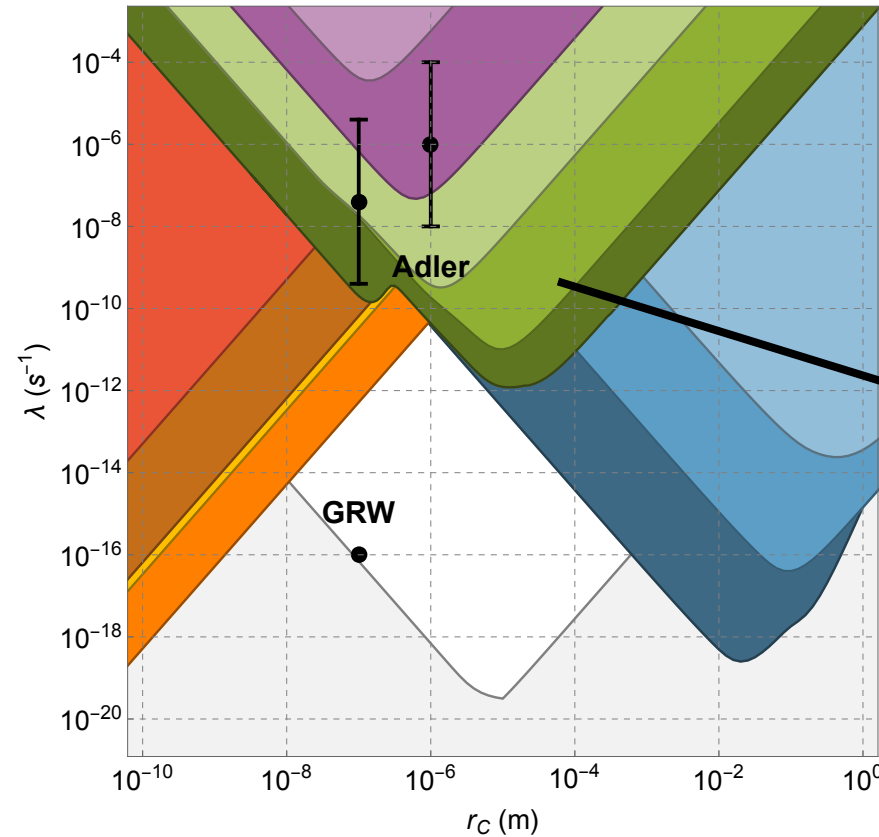
Non-interferometric Tests

Cold Atoms

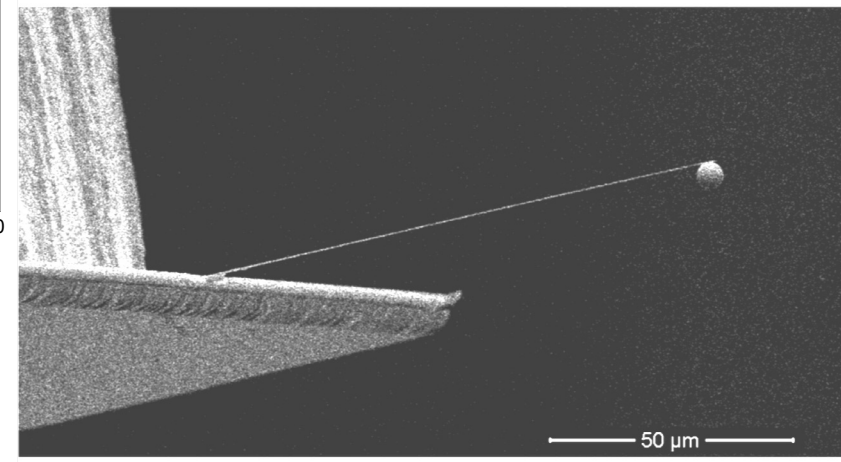
Kovachy *et al.*, Phys. Rev. Lett. **114**, 143004 (2015)
Bilardello *et al.*, Physica A, **462**:764-782 (2016)



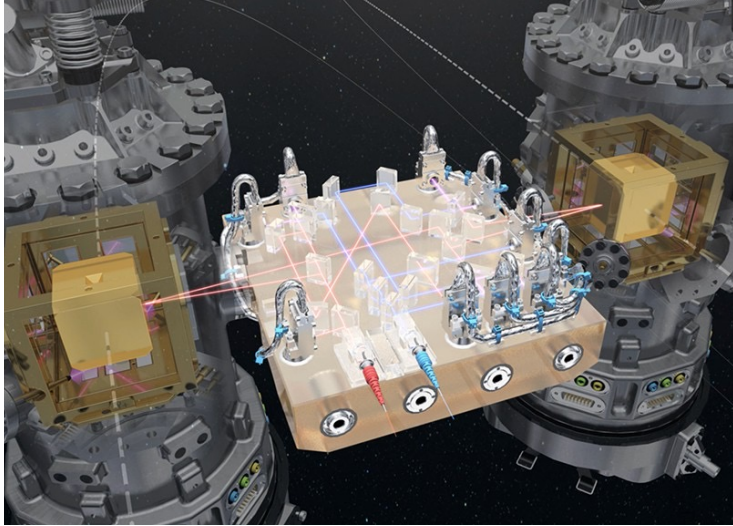
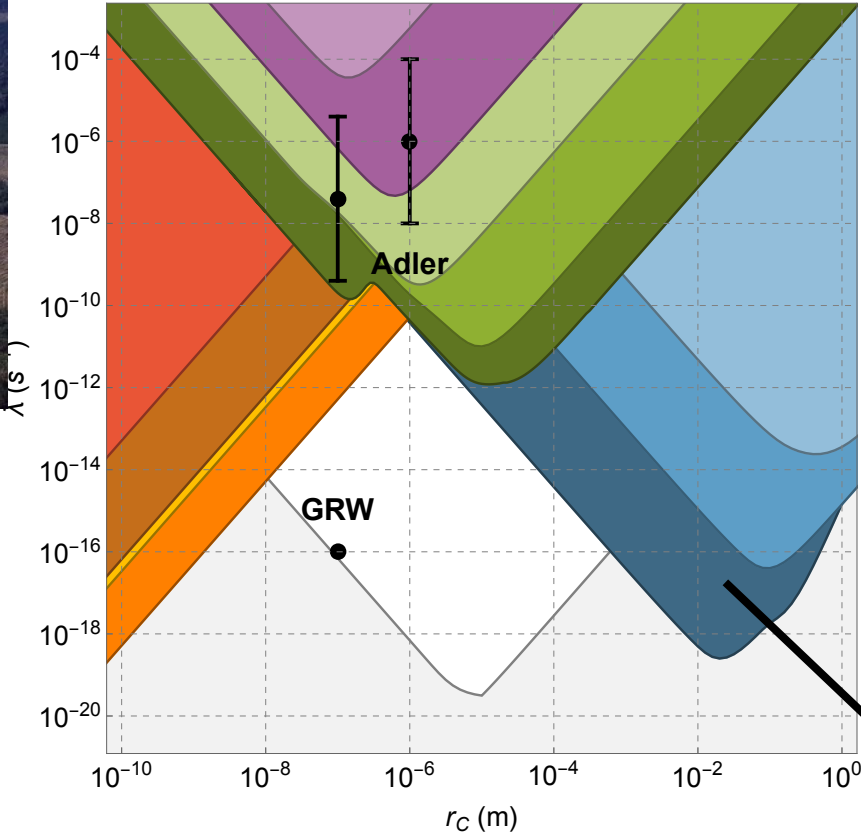
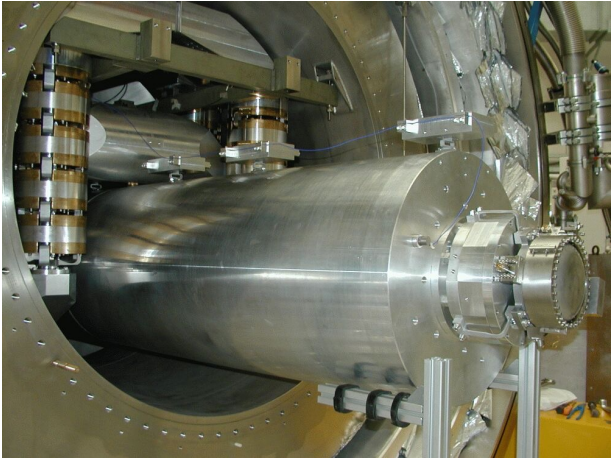
Non-interferometric Tests



Nanomechanical cantilevers
Vinante *et al.*, Phys. Rev. Lett. **116**, 090402 (2016)
Vinante *et al.*, Phys. Rev. Lett. **119**, 110401 (2017)

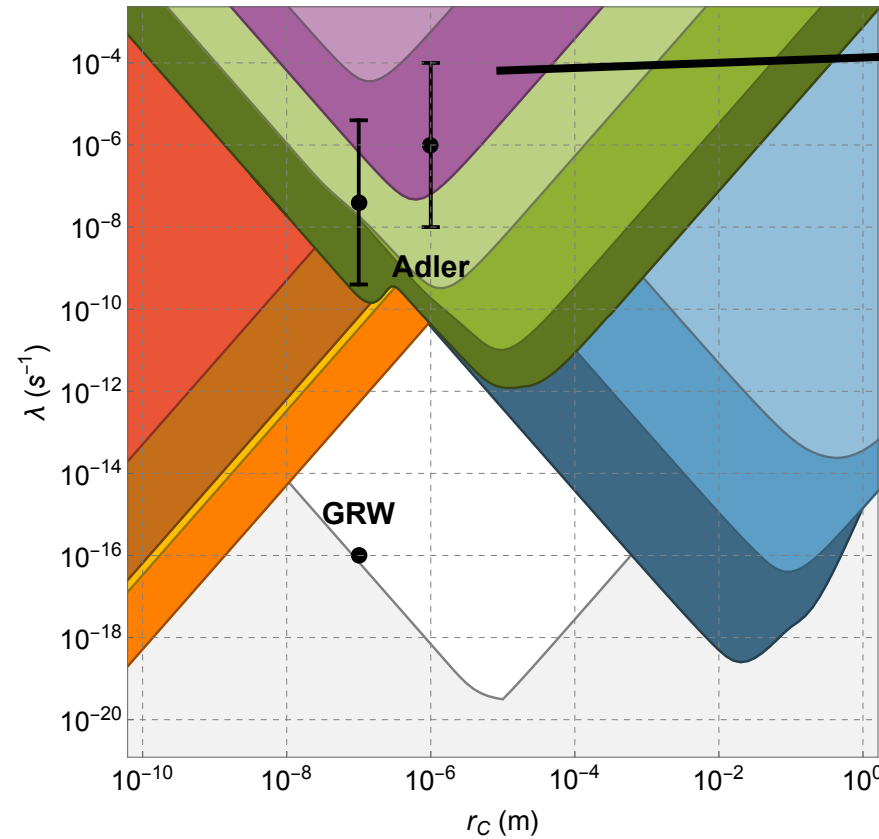
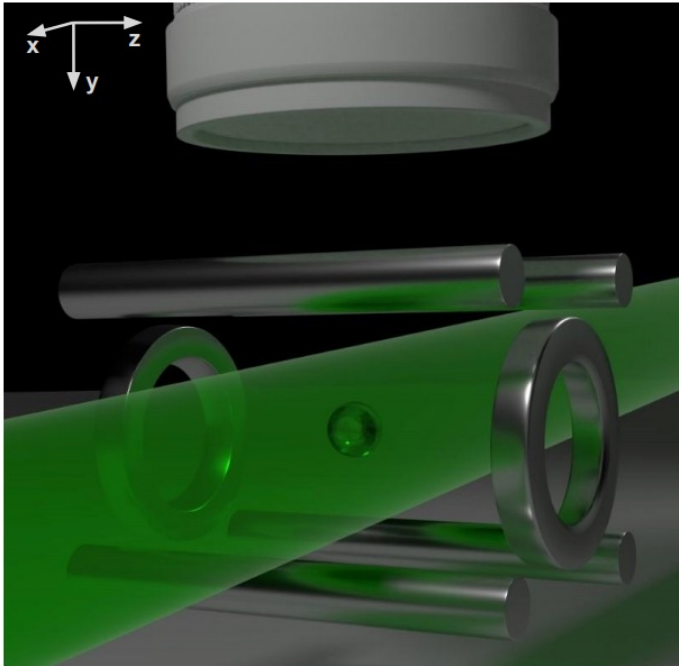


Non-interferometric Tests



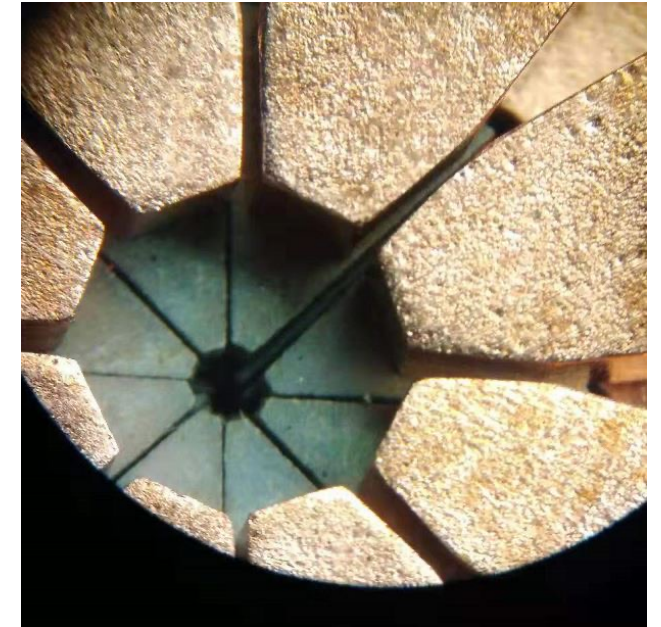
Gravitational wave detectors
Auriga, LIGO, LISA Pathfinder
Carlesso *et al*, Phys. Rev. D **94**, 124036 (2016)

Non-interferometric Tests

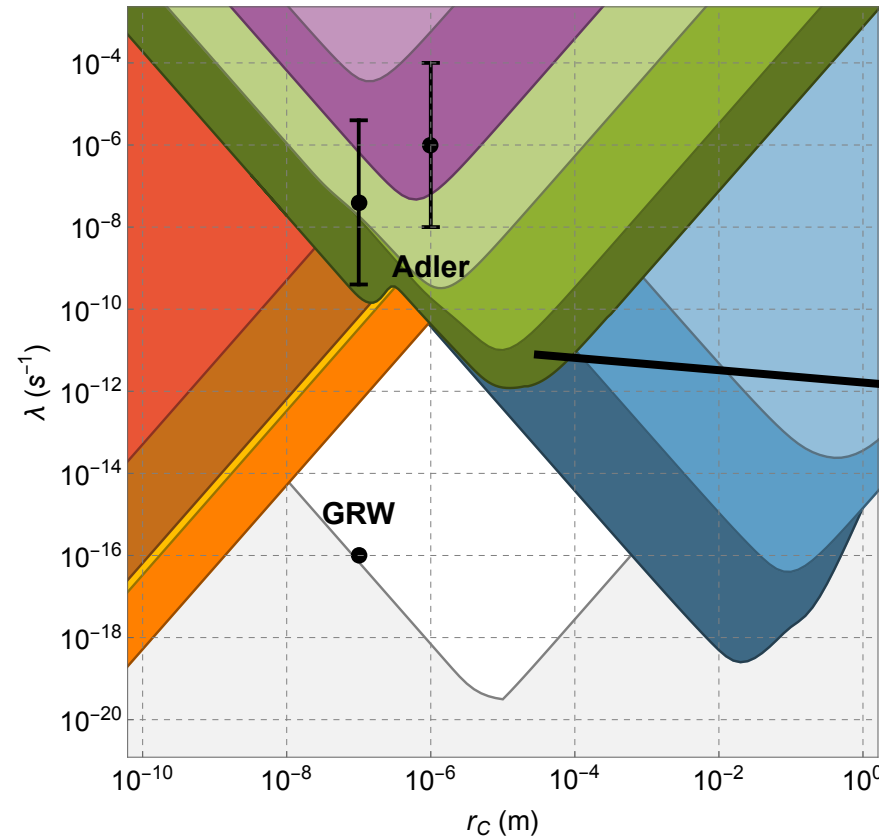
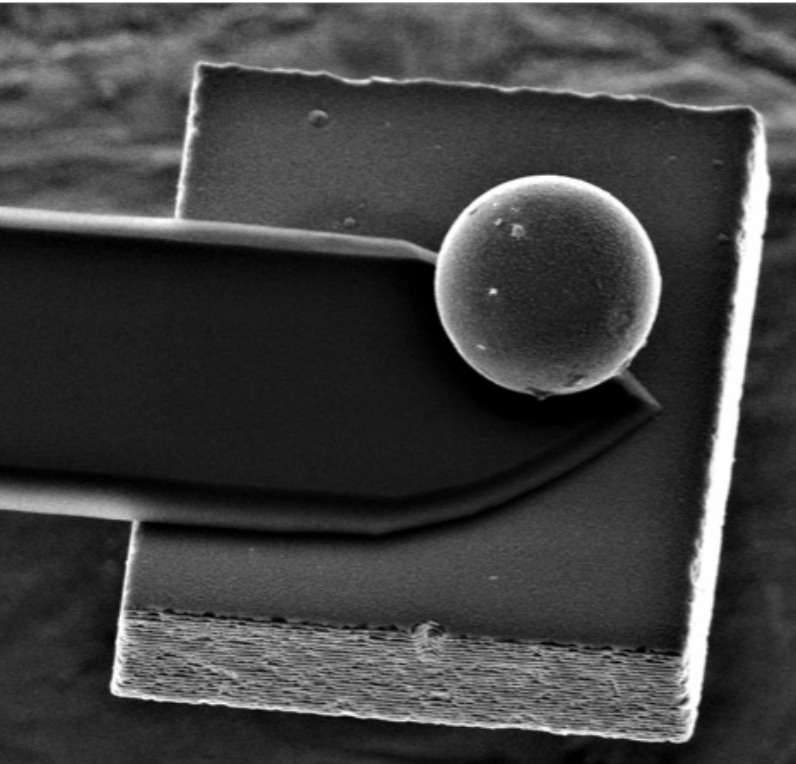


Levitated nano-oscillators

Pontin *et al.*, Phys. Rev. Res. **2**, 023349 (2020)
Zheng *et al.*, Phys. Rev. Res. **2**, 013057 (2020)

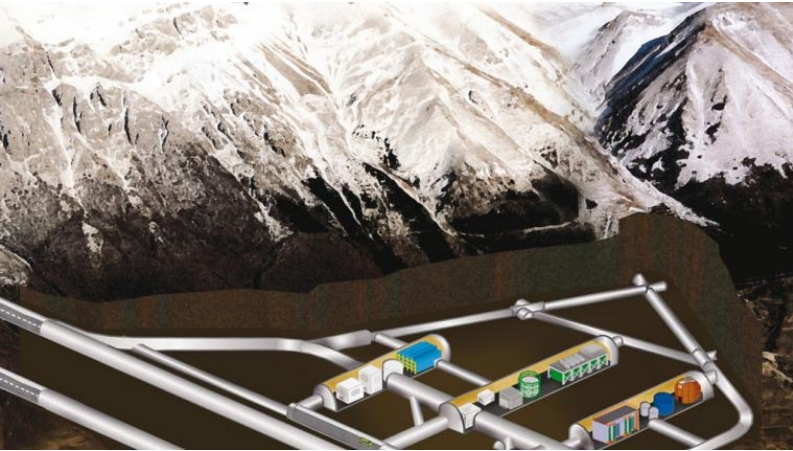


Non-interferometric Tests



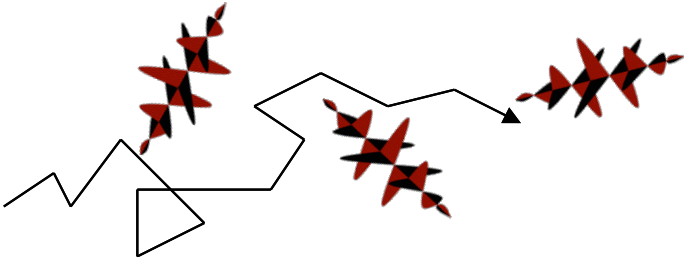
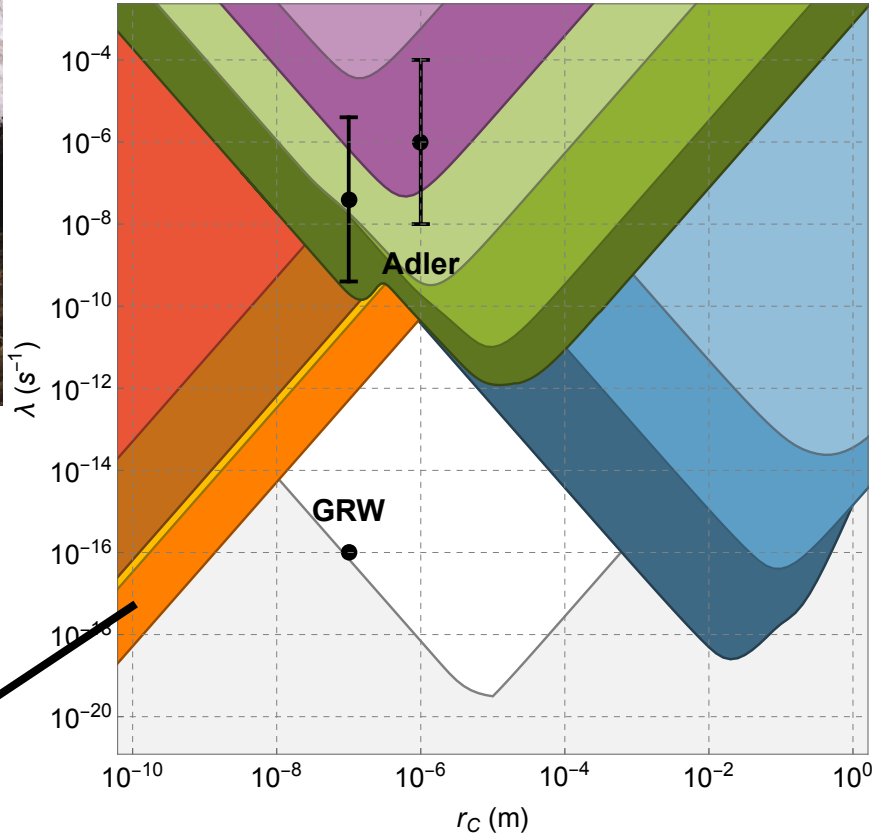
Nanomechanical cantilevers
Vinante *et al.*, Phys. Rev. Lett. **125**, 100404 (2020)

Non-interferometric Tests

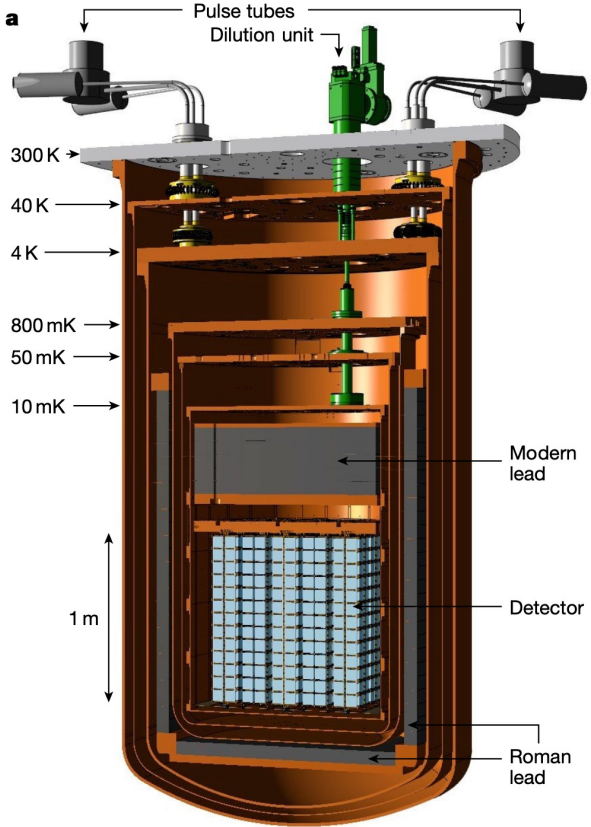
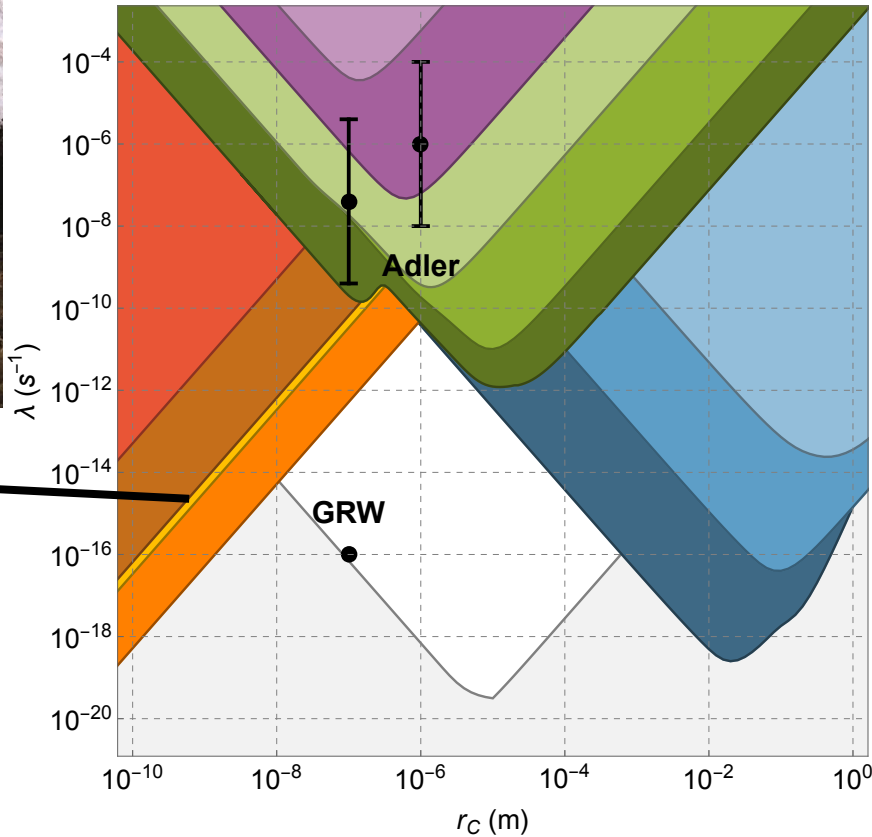
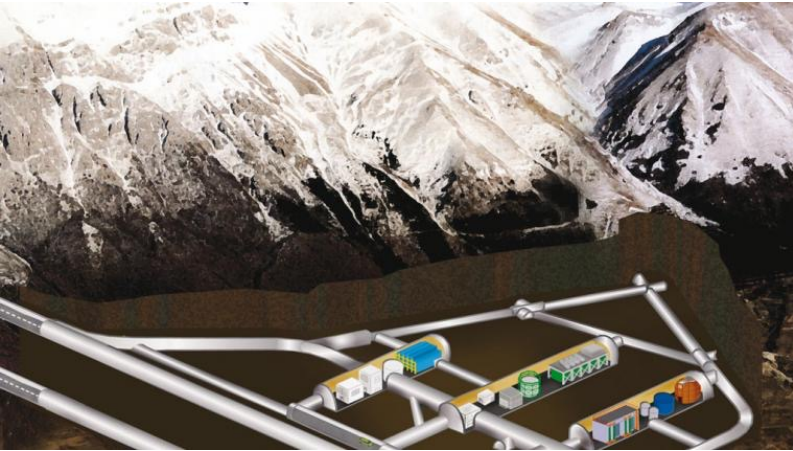


X-rays emission

Donadi *et al*, Eur. Phys. J. C **81**, 1 (2021)



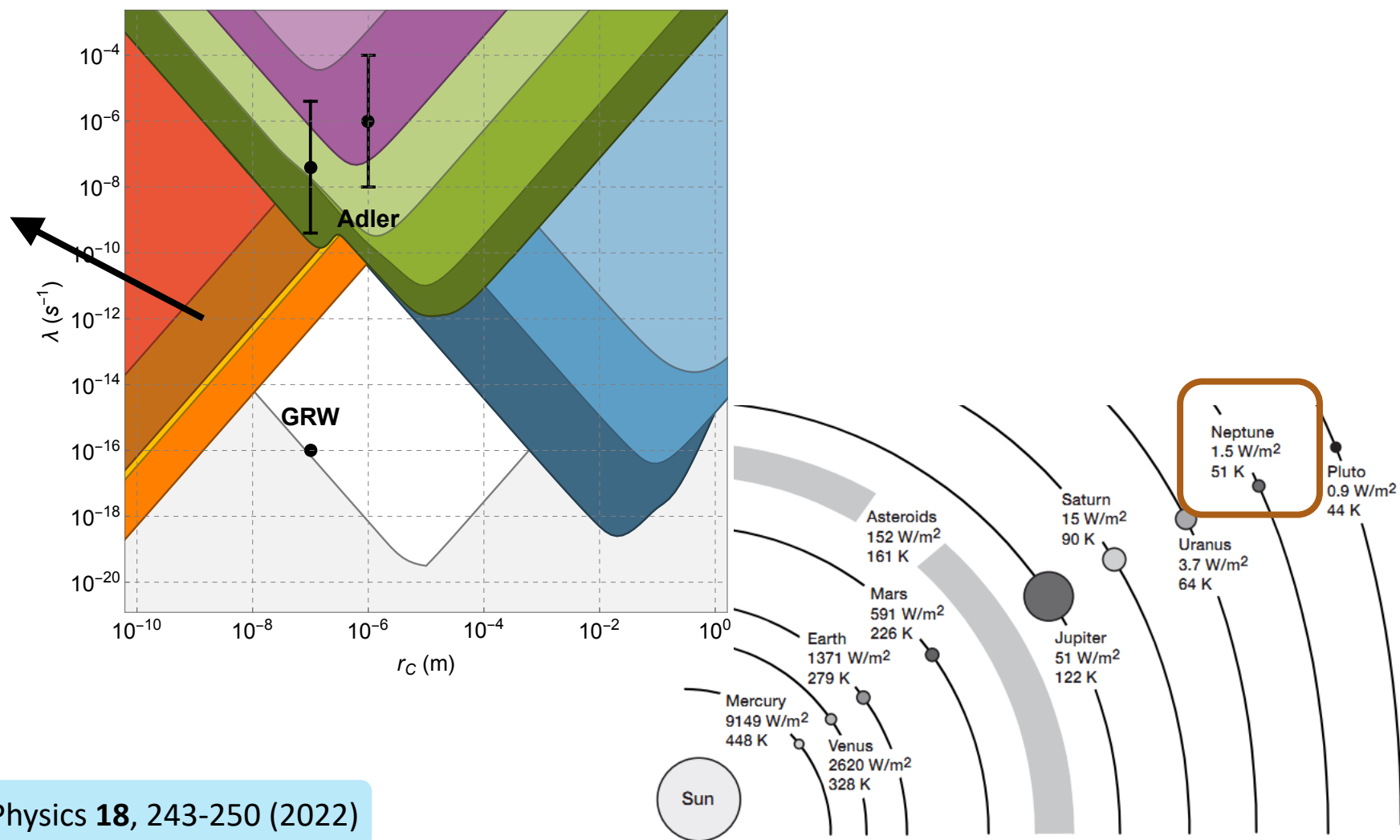
Non-interferometric Tests



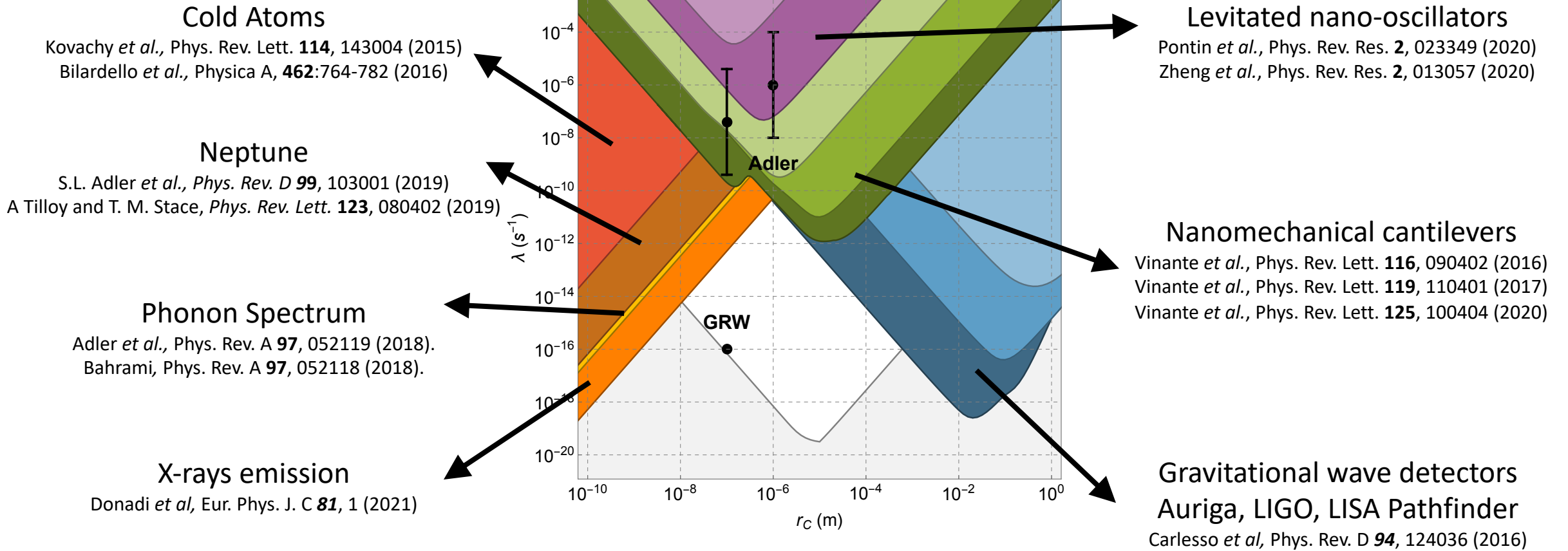
Phonon Spectrum
Adler *et al.*, Phys. Rev. A **97**, 052119 (2018).
Bahrami, Phys. Rev. A **97**, 052118 (2018).

Non-interferometric Tests

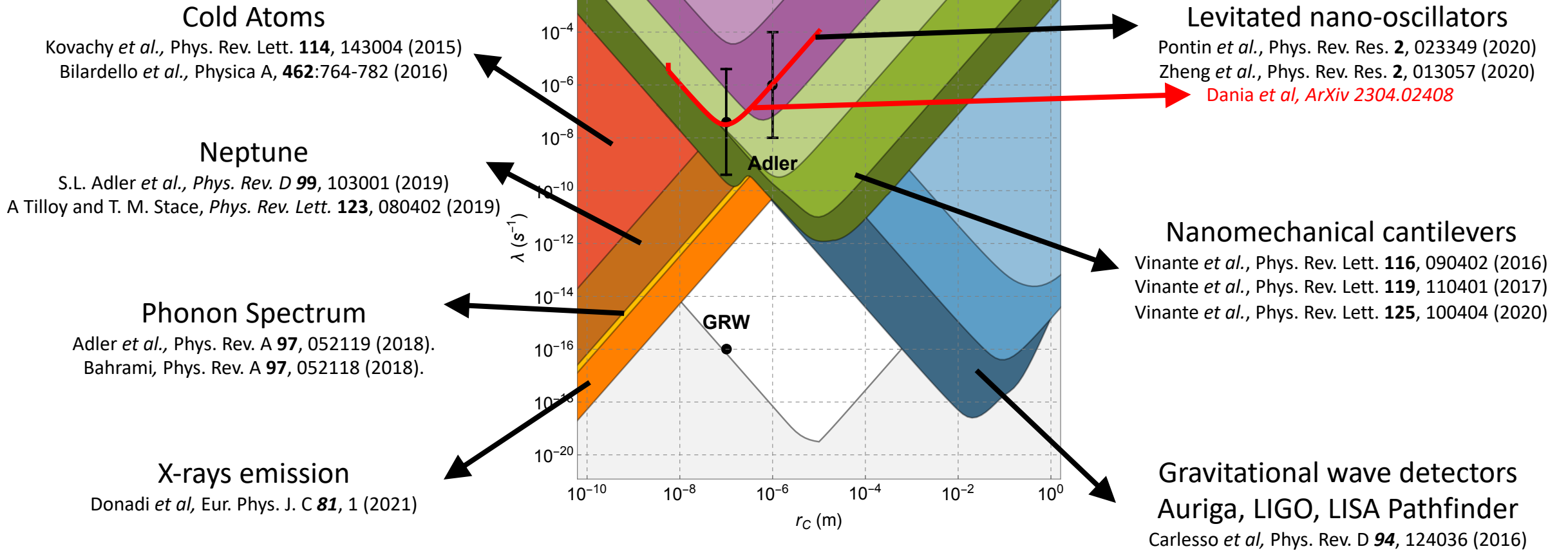
Neptune
S.L. Adler *et al.*, *Phys. Rev. D* **99**, 103001 (2019)



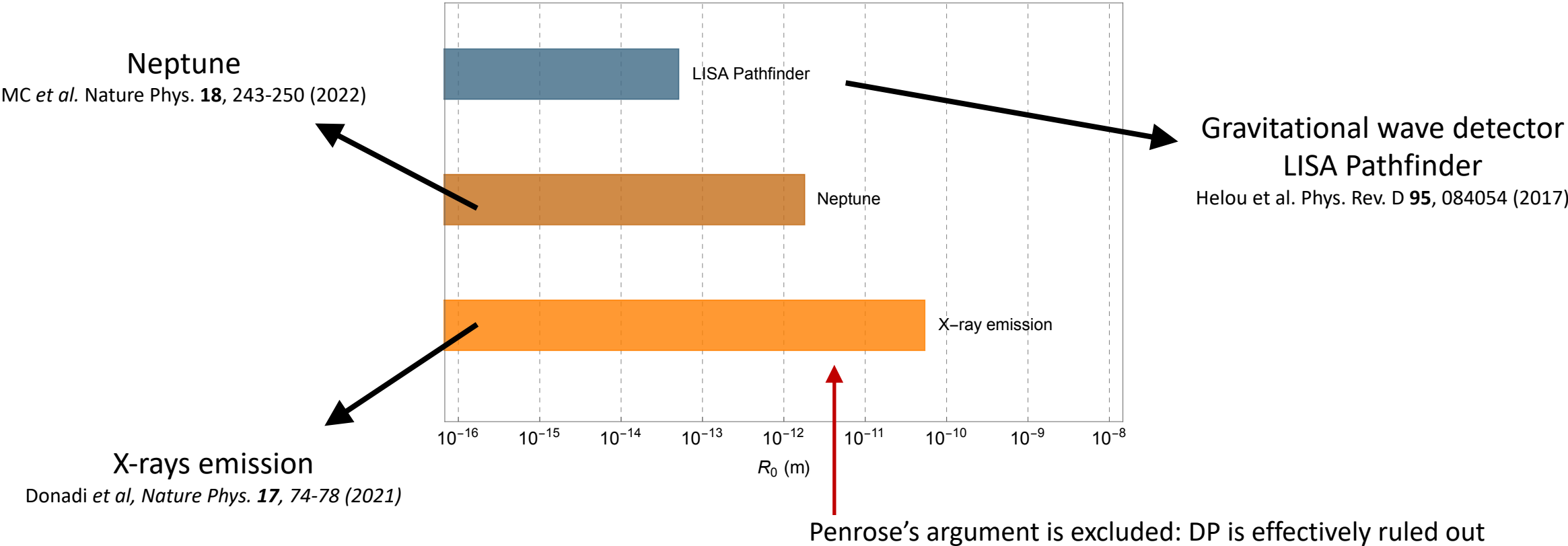
Non-interferometric Tests



Non-interferometric Tests



Non-interferometric Tests - DP



Future? Space experiments!

Option on ground: drop towers

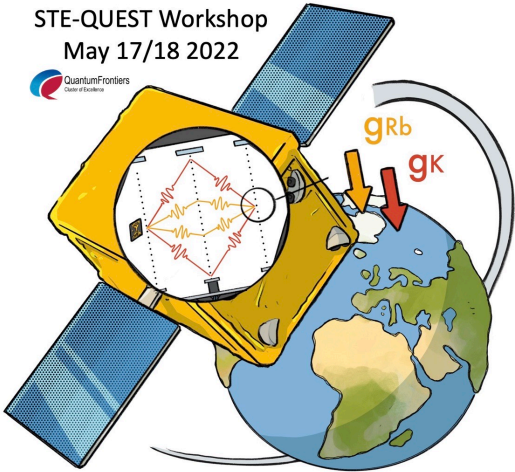


Bremen drop tower:
Up to 4.6s of free-fall time, 9.2s with the catapult
3 runs/day

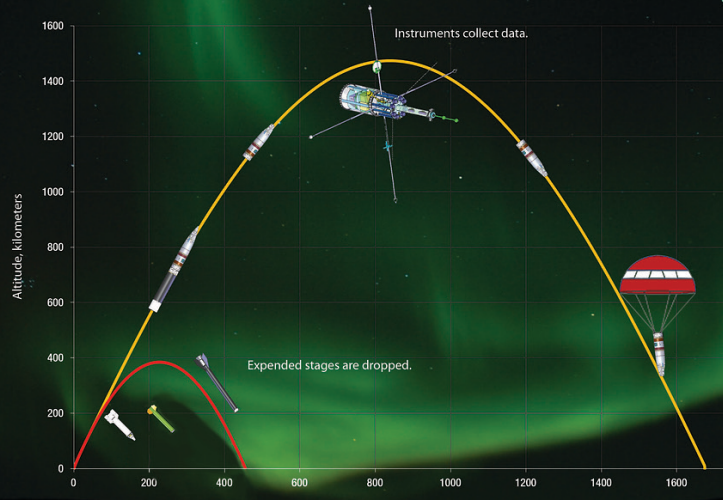
Einstein elevator:
Around 4s of free-fall time with the catapult
300 runs/day

Belenchia *et al.*, *Phys. Rep.* **951**, 1 (2022)

STE-QUEST Workshop
May 17/18 2022



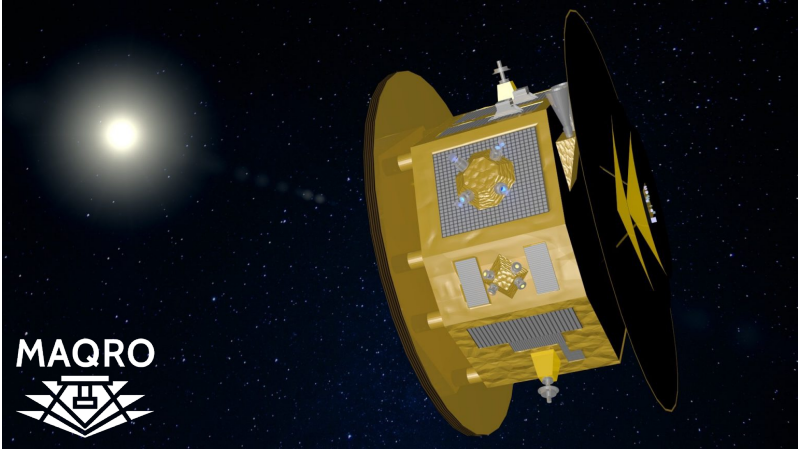
Options in space



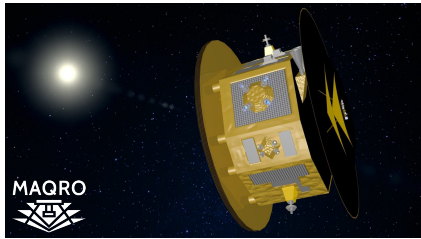
Sounding rockets:
Around 5-10 minutes of free-fall time



International Space Station (ISS)

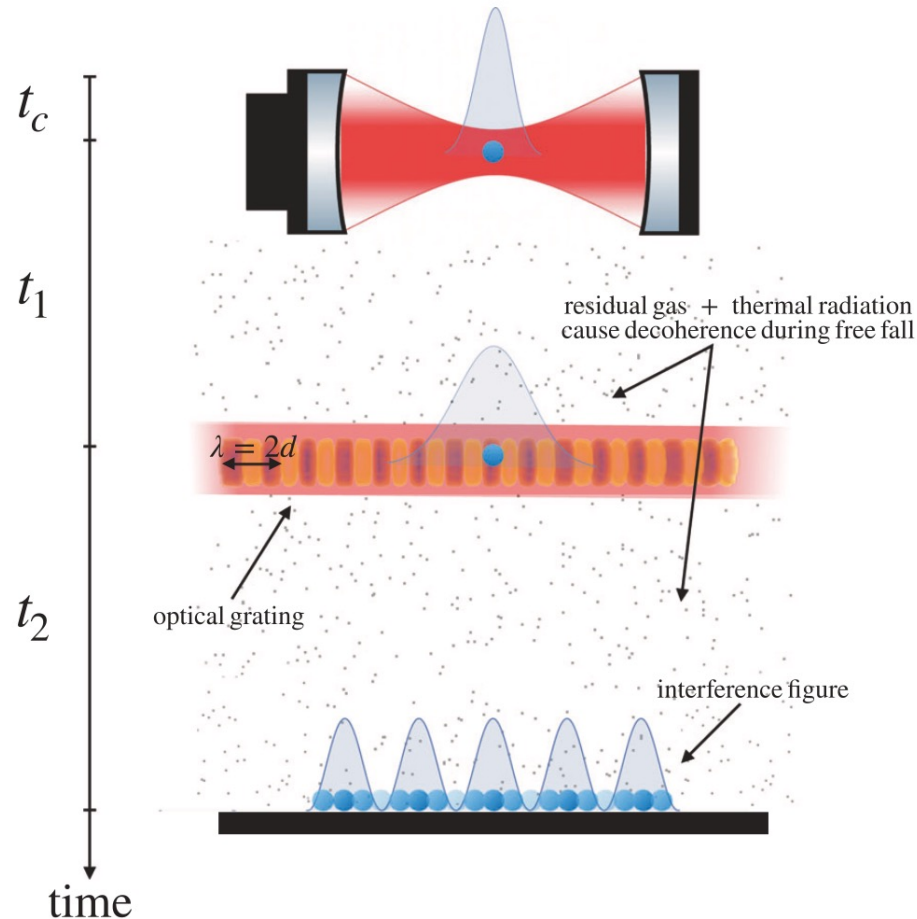


Dedicated experiments: STE-QUEST; MAQRO & QFFP



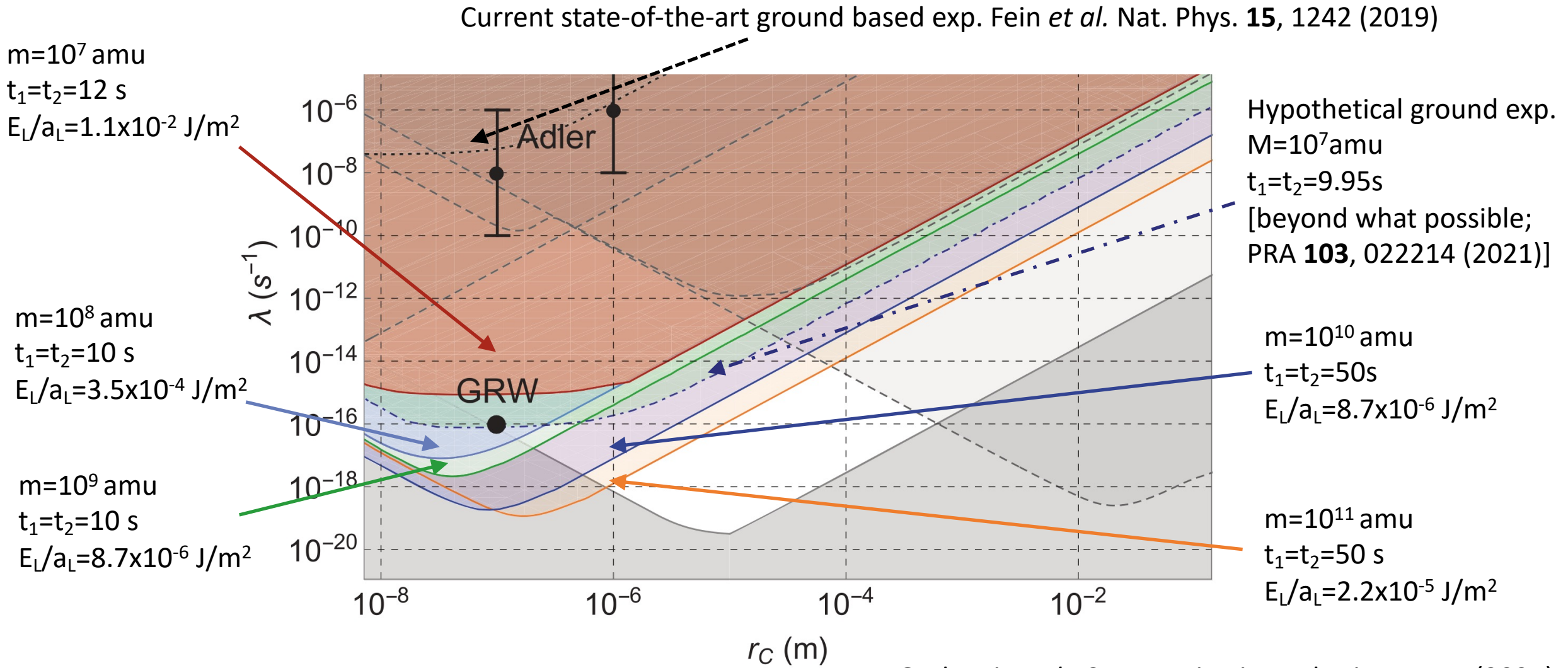
Interferometric experiments in space with nanoparticles

MAQRO and QPPF investigate the possibility of performing near-field interferometric scheme with “large” particles on dedicated scientific space missions



- (a) Nanoparticle is trapped and cooled in an optical cavity
- (b) It is released and let evolve freely for a time t_1 . Its coherent length needs to cover 2 adjacent “slits” of the optical grating
- (c) A retro-reflected laser provides a pure-phase grating
- (d) Free evolution of time t_2 needed to form the interference pattern
- (e) Measurement of the particle via optical detection

Kaltenbaek *et al.*,
 EPJ Q. Tech. **3**, 5 (2016);
 Gasbarri *et al.*,
 Comm. Phys. **4**, 155 (2021);
 Belenchia *et al.*,
 Phys. Rev. A **100**, 033813 (2019)



Gravity models

Problems at hand:

how does the gravitational field of a superposition look like?

how two quantum system interact gravitationally?

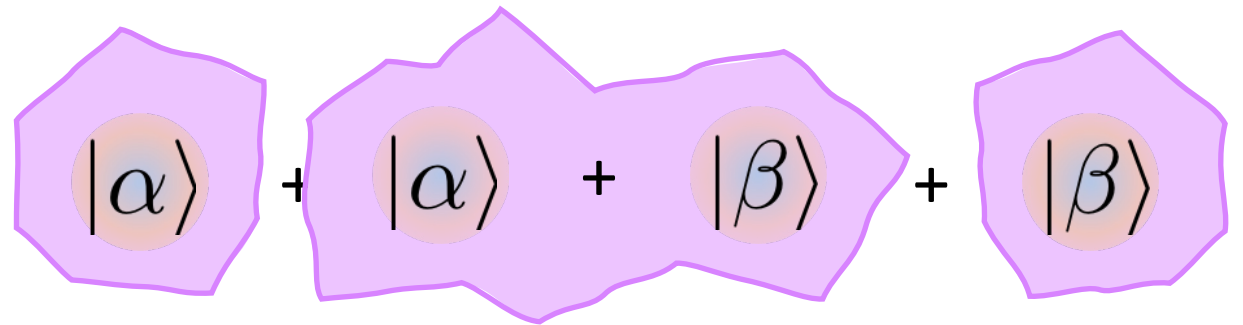
Quantum scenario

Semi-classical scenario

Something else?

Prepare a “massive superposition”

$$|\alpha\rangle + |\beta\rangle$$



Gravity models

Problems at hand: how does the gravitational field of a superposition look like?
 how two quantum system interact gravitationally?

Quantum scenario

Semi-classical scenario

Something else?

How to discriminate these two alternatives with low-energy experiments?

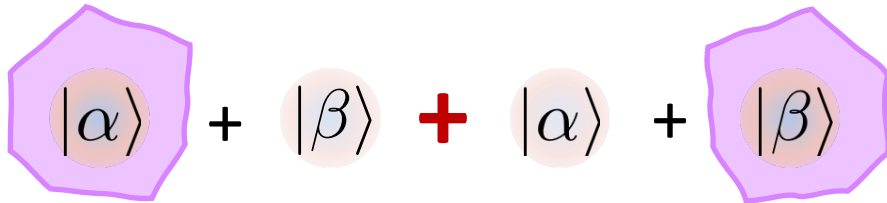
Different approaches based on the question we pose

- Entanglement generation (quantum scenario)
- Superposition of the gravitational field (quantum and classical scenario)
- Presence of extra noises (classical scenario)

How does the gravitational field of a superposition looks like?

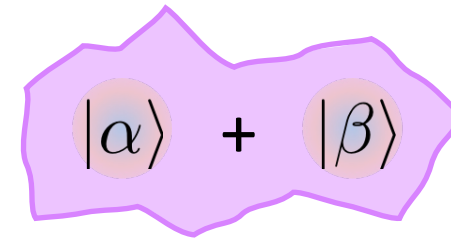
$$\psi(\mathbf{r}_1) = \frac{1}{\sqrt{2}}(\alpha(\mathbf{r}_1) + \beta(\mathbf{r}_1))$$

Quantum scenario



$$V_\gamma = -Gm_1 m_2 \int d^3 \mathbf{r}_1 \frac{|\gamma(\mathbf{r}_1, t)|^2}{|\mathbf{r}_1 - \mathbf{r}_2|}$$

Semi-Classical scenario




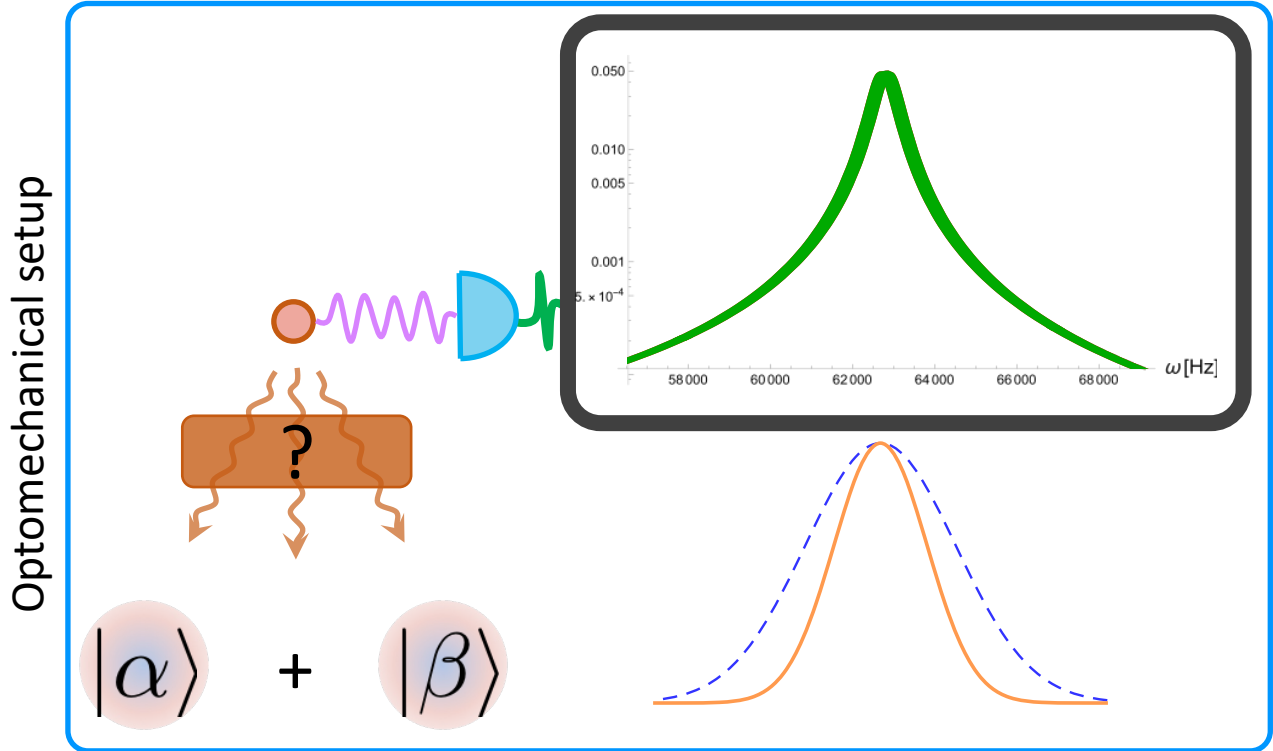
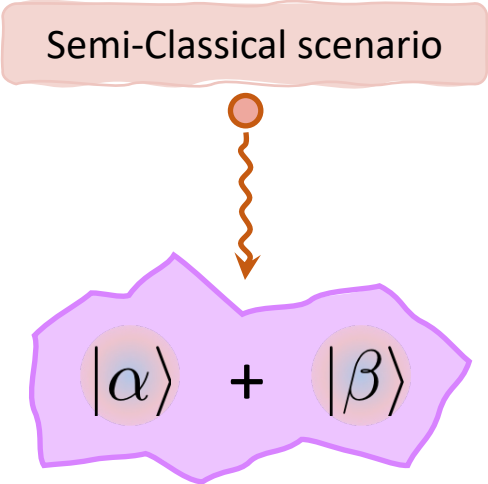
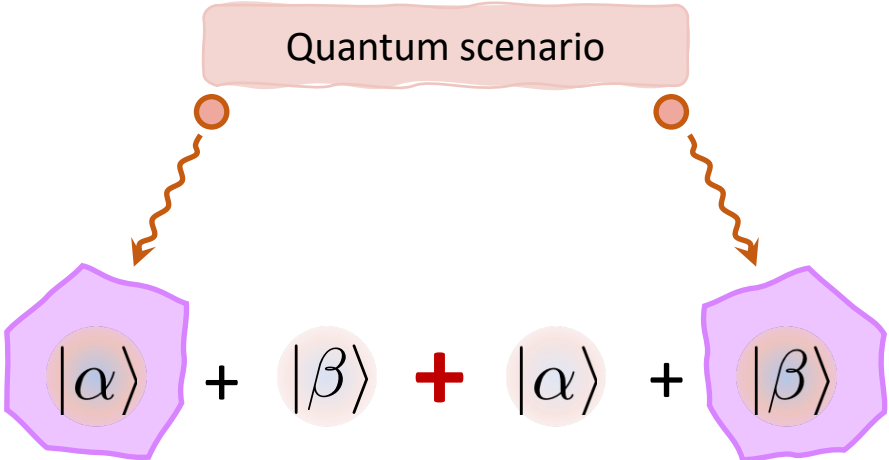
$$V_{cl} = -Gm_1 m_2 \int d^3 \mathbf{r}_1 \frac{|\psi(\mathbf{r}_1, t)|^2}{|\mathbf{r}_1 - \mathbf{r}_2|}$$

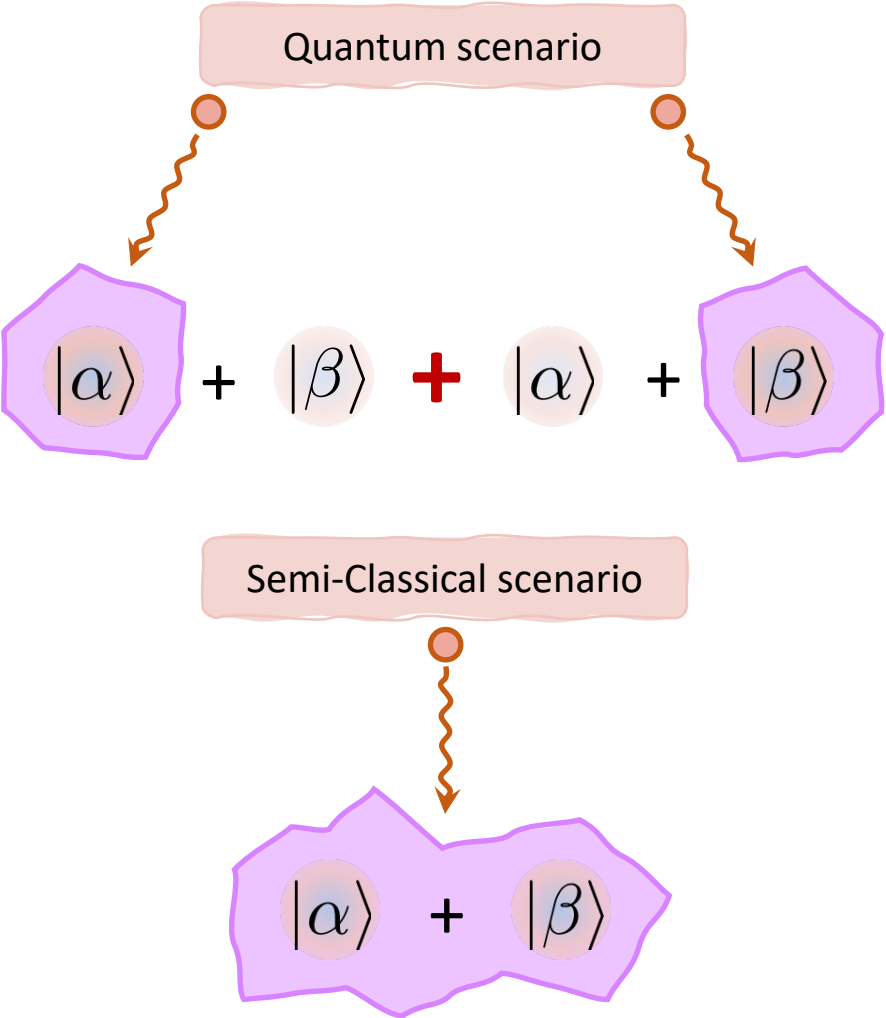
How to test the two hypothesis?

Is Gravity Quantum?

M. Bahrami,^{1,2} A. Bassi,^{1,2} S. McMillen,³ M. Paternostro,³ and H. Ulbricht⁴
ArXiv 1507.05733 (2015)

$|\alpha\rangle + |\beta\rangle$ and add a quantum probe: 





New Journal of Physics

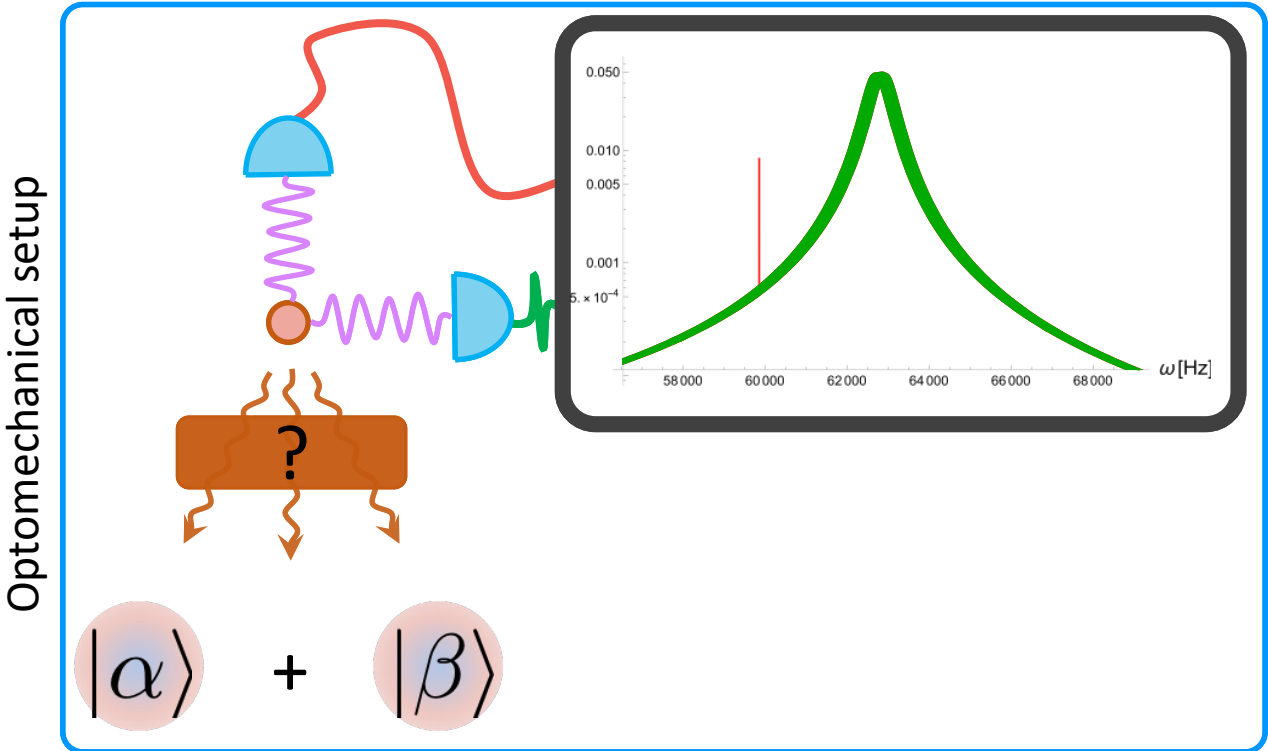
The open access journal at the forefront of physics

Testing the gravitational field generated by a quantum superposition

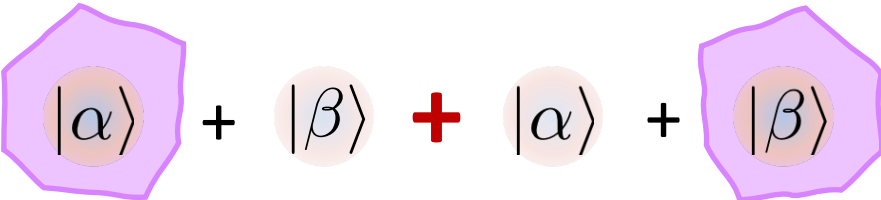
New Journal of Physics **21**, 093052 (2019)

M Carlesso^{1,2} , A Bassi^{1,2} , M Paternostro³  and H Ulbricht⁴ 

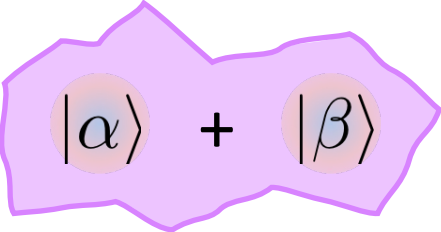
Test the motion along two directions



Quantum scenario



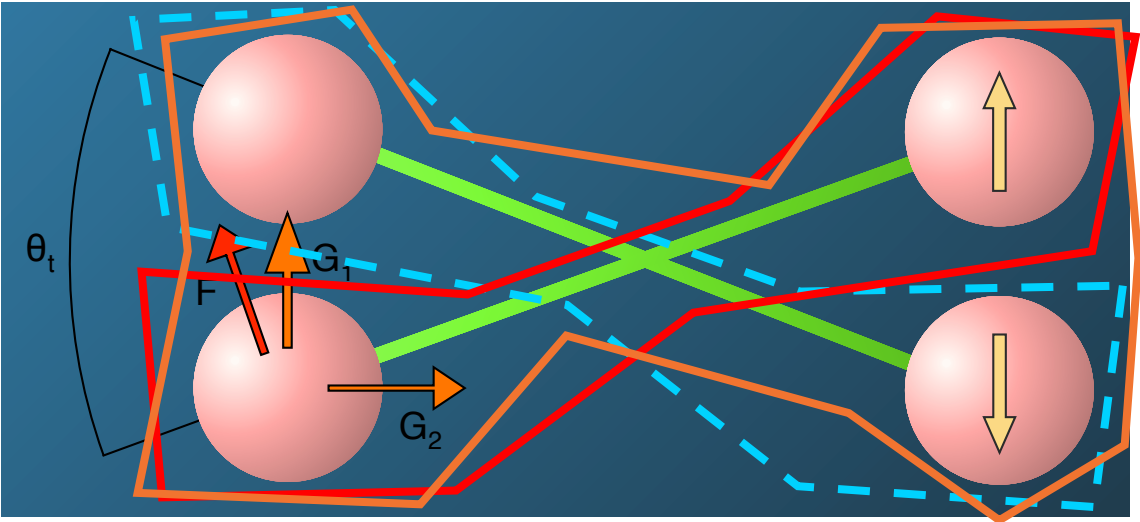
Semi-Classical scenario



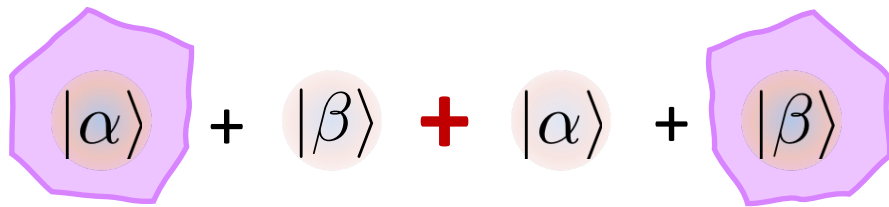
When Cavendish meets Feynman: A quantum torsion balance for testing the quantumness of gravity

Matteo Carlesso,^{1,2,*} Mauro Paternostro,^{3,4} Hendrik Ulbricht,⁵ and Angelo Bassi^{1,2}
ArXiv 1710.08695 (2017)

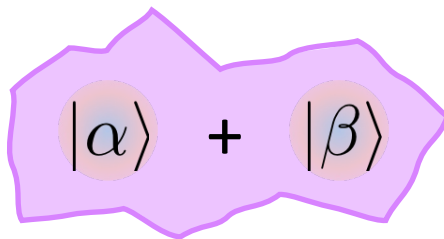
Using a single self-testing quantum system



Quantum scenario



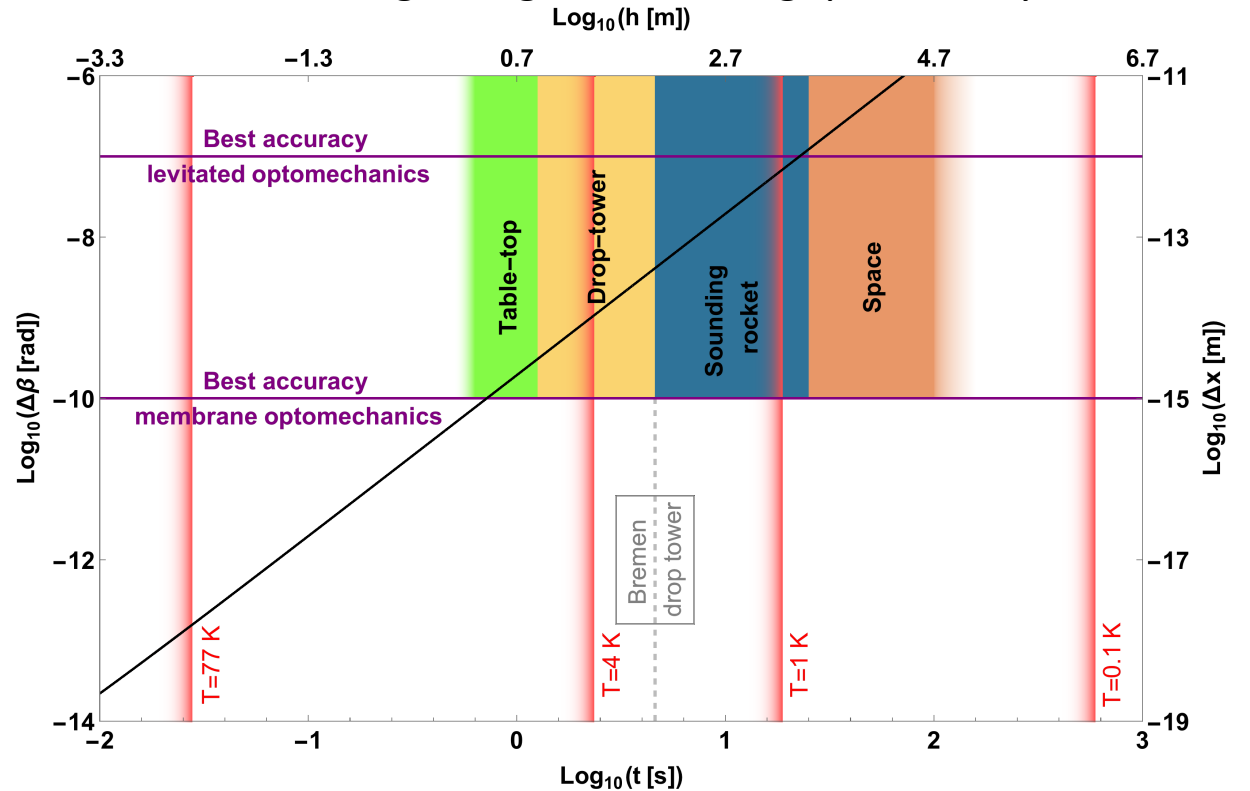
Semi-Classical scenario



When Cavendish meets Feynman: A quantum torsion balance for testing the quantumness of gravity

Matteo Carlesso,^{1,2,*} Mauro Paternostro,^{3,4} Hendrik Ulbricht,⁵ and Angelo Bassi^{1,2}
 ArXiv 1710.08695 (2017)

Using a single self-testing quantum system



How two quantum system interact gravitationally?

I.e., How to reconcile the action of a classical mediator (gravity) with a quantum object?

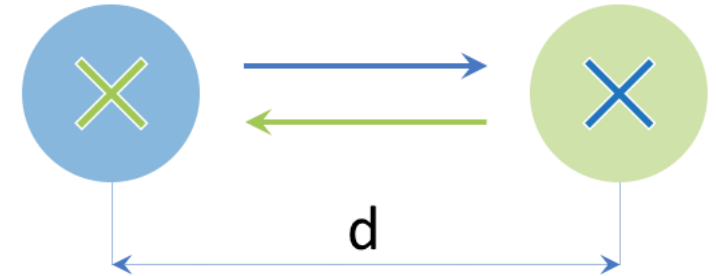
Eventually, in quantum mechanics, one can do only two things:

- Apply a unitary evolution
- Measure the expectation values

A classical channel model for gravitational decoherence

D Kafri¹, J M Taylor¹ and G J Milburn^{2,3}

New Journal of Physics **16**, 065020 (2014)



KTM model: Two harmonic oscillators, at distance d , which interact gravitationally

The “quantum” linearised interaction will be

$$\hat{H} = \hat{H}_0 + \hat{H}_{\text{grav}} \quad \hat{H}_0 = \sum_{\alpha=1}^2 \frac{\hat{p}_{\alpha}^2}{2m_{\alpha}} + \frac{1}{2}m_{\alpha}\omega_{\alpha}^2\hat{x}_{\alpha}^2 \quad \hat{H}_{\text{grav}} = K\hat{x}_1\hat{x}_2 \quad K = \frac{Gm_1m_2}{d^3}$$

1) Weak measurement of the position

$$r_{\alpha} = \langle \hat{x}_{\alpha} \rangle + \frac{\hbar}{\sqrt{\gamma_{\alpha}}} \frac{dW_{\alpha,t}}{dt}$$

How classical gravity acts:

2) Feedback dynamics

$$\hat{H}_{\text{grav}} \longrightarrow \hat{H}_{\text{fb}} = \chi_{12}r_1\hat{x}_2 + \chi_{21}r_2\hat{x}_1$$

KTM model: full dynamics 2 particles

$$\frac{d\hat{\rho}}{dt} = -\frac{i}{\hbar} [\hat{H}_0 + K\hat{x}_1\hat{x}_2, \hat{\rho}] - \sum_{\alpha=1}^2 \left(\frac{\gamma_{\alpha}}{8\hbar^2} + \sum_{\substack{\beta=1 \\ \beta \neq \alpha}}^2 \frac{K^2}{2\gamma_{\beta}} \right) [\hat{x}_{\alpha}, [\hat{x}_{\alpha}, \hat{\rho}]]$$

We mimicked the gravitational interaction!

An diffusive noise is the price to pay

Mimimising wrt $\gamma = \gamma_1 = \gamma_2$ one obtains $-\frac{K}{2\hbar} \sum_{\alpha=1}^2 [\hat{x}_{\alpha}, [\hat{x}_{\alpha}, \hat{\rho}]]$, with $K = \frac{Gm_1m_2}{d^3}$

One can test the diffusion, and thus the model

Levitated optomechanics is a perfectly suitable platform for it

Principle of least decoherence for Newtonian semiclassical gravity

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Same idea as in the KTM model, but with continuous mass distributions

$$\hat{H}_{\text{grav}} = \frac{1}{2} \int d\mathbf{x} d\mathbf{y} \mathcal{V}(\mathbf{x} - \mathbf{y}) \hat{\mu}(\mathbf{x}) \hat{\mu}(\mathbf{y}) \quad \text{with} \quad \mathcal{V}(\mathbf{x} - \mathbf{y}) = -\frac{G}{|\mathbf{x} - \mathbf{y}|}$$

is substituted to
 1) measurement
 2) feedback protocol

Full dynamics

$$\frac{d\hat{\rho}}{dt} = -\frac{i}{\hbar} [\hat{H}_0 + \hat{H}_{\text{grav}}, \hat{\rho}] - \int d\mathbf{x} d\mathbf{y} D(\mathbf{x} - \mathbf{y}) [\hat{\mu}(\mathbf{x}), [\hat{\mu}(\mathbf{y}), \hat{\rho}]] \quad D(\mathbf{x} - \mathbf{y}) = \left[\frac{\gamma}{8\hbar^2} + \frac{1}{2} (\mathcal{V} \circ \gamma^{-1} \circ \mathcal{V}) \right] (\mathbf{x} - \mathbf{y})$$

For $\gamma(\mathbf{x} - \mathbf{y}) = -2\hbar\mathcal{V}(\mathbf{x} - \mathbf{y})$ one gets the Diosi-Penrose model, which is the minimum decoherence model

Should all these models be disregarded?

The model describes an indefinite increase of energy due to gravity $\langle \hat{H} \rangle_t = \frac{\hbar G \sum_k m_k}{4\sqrt{\pi} R_0^3} t$

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How one can implement dissipation

Case of the TD model with minimal decoherence, i.e. the Diosi-Penrose model

$$\hat{A}(\mathbf{x}) = \hat{\mu}(\mathbf{x}) + i\hat{\mu}_I(\mathbf{x}) \quad \text{with} \quad \hat{\mu}_I(\mathbf{x}) = -\frac{\hbar\beta}{4} \nabla_{\mathbf{x}} \hat{\mathbf{J}}(\mathbf{x})$$

mass density current

$$\hat{A}(\mathbf{x}) = \sum_{k=1}^N \frac{m_k}{(2\pi\hbar)^3} \int d^3q e^{-\frac{i}{\hbar} \mathbf{q} \cdot (\mathbf{x} - \hat{\mathbf{x}}_k) - \frac{R^2}{2\hbar^2} [(1 + \alpha_k) \mathbf{q} + 2\alpha_k \hat{\mathbf{p}}_k]^2}$$

The model has a more fundamental flavour compared to standard dissipative version of DP

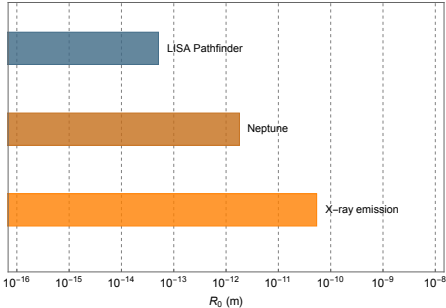
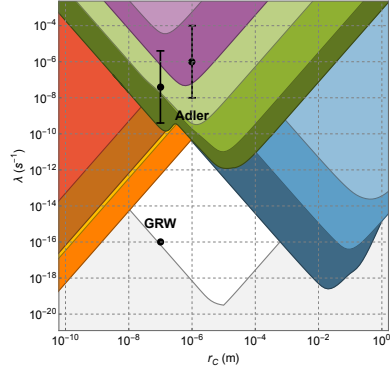
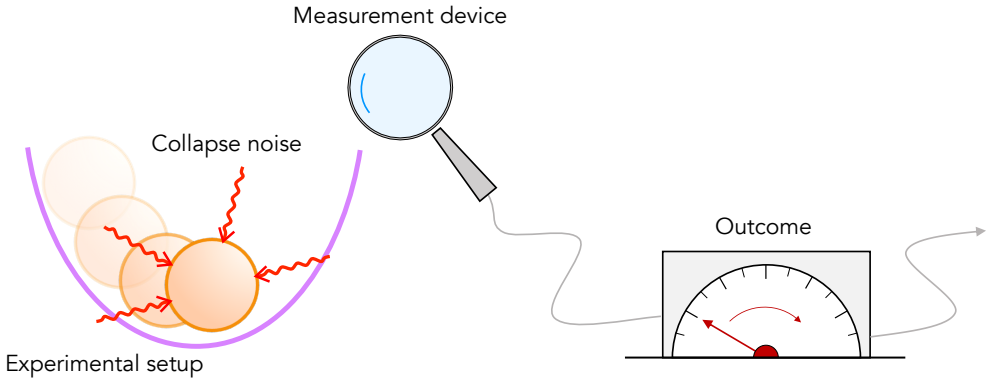
The same choice can be considered to construct the collapse operator in dissipative collapse models as the Continuous Spontaneous Localisation (CSL) model

In both cases, one obtains an evolution for the energy reading

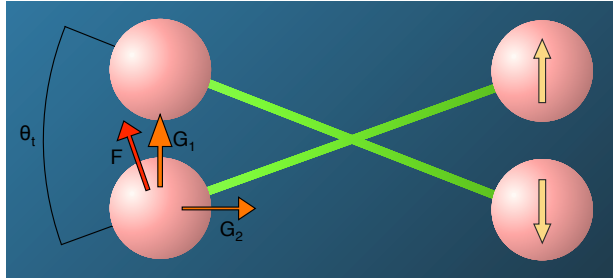
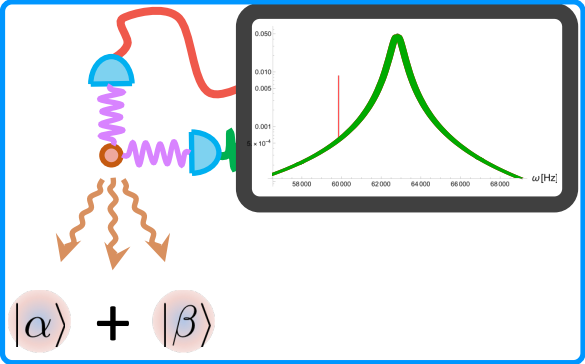
$$\frac{d}{dt} \langle \hat{H} \rangle_t = P - \Gamma \langle \hat{H} \rangle_t \quad \text{with} \quad P, \Gamma > 0 \quad \text{which gives a finite asymptotic energy}$$

Summary

Collapse models



Gravity models



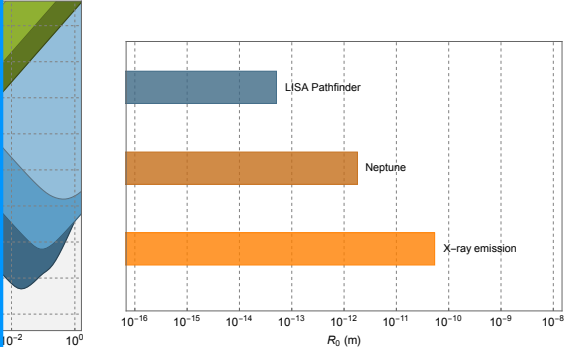
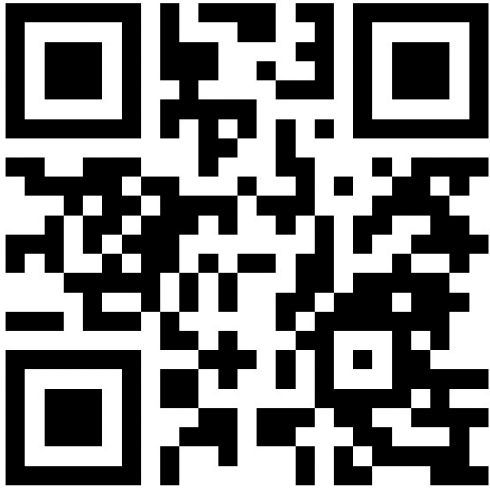
Summary

We are organising a School in Foundations

“Fundamental Problems in Quantum Physics 2023”

in Trieste, from 13 to 15 September 2023.

Info and registration at www.qmts.it/?q=fpqp2023



$$|\alpha\rangle + |\beta\rangle$$

