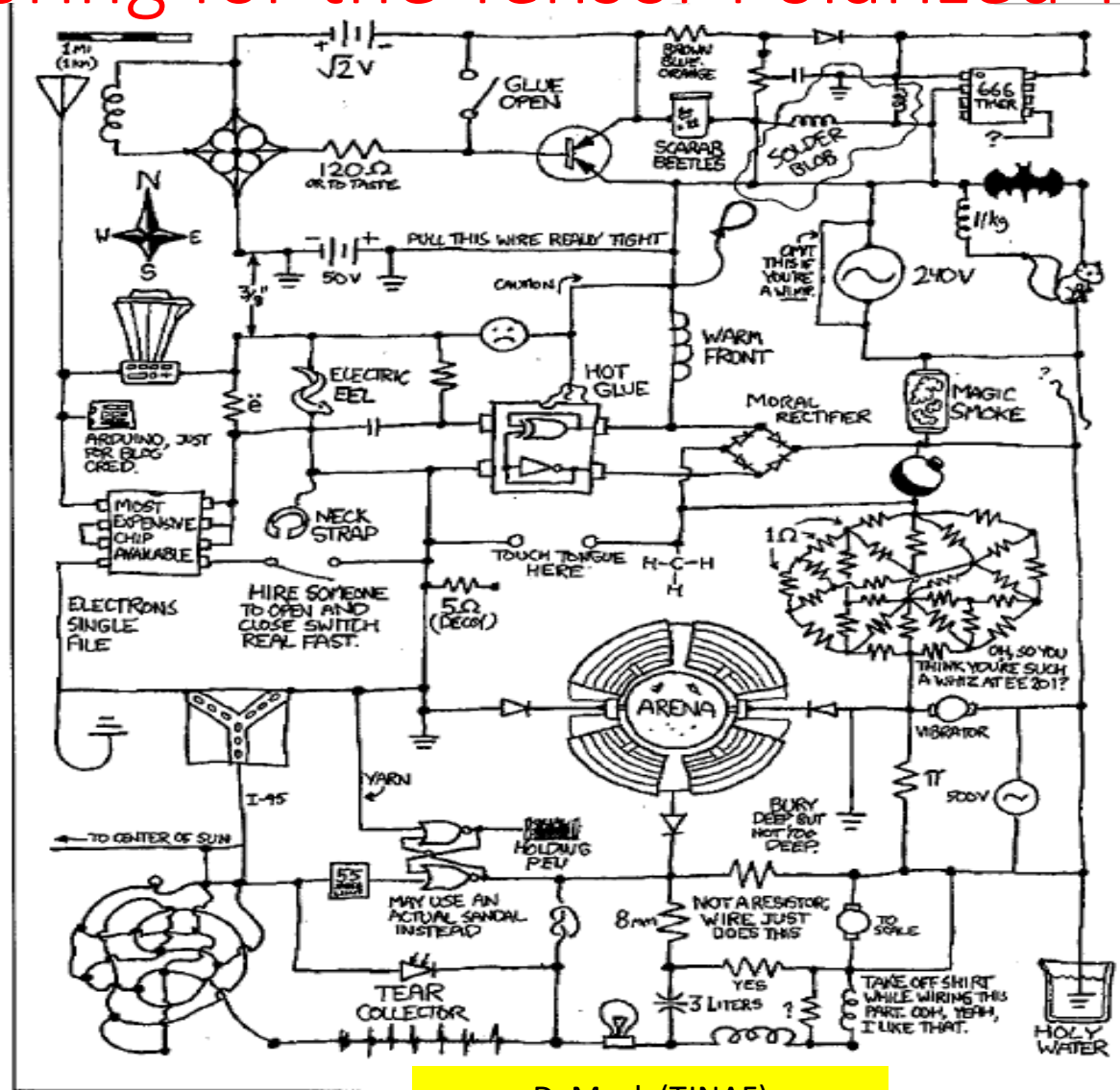


Charge Monitoring for the Tensor Polarized Target Program



D. Mack (TJNAF)
ECT* Tensor Spin Observables
Trento, Italy
July 12, 2023

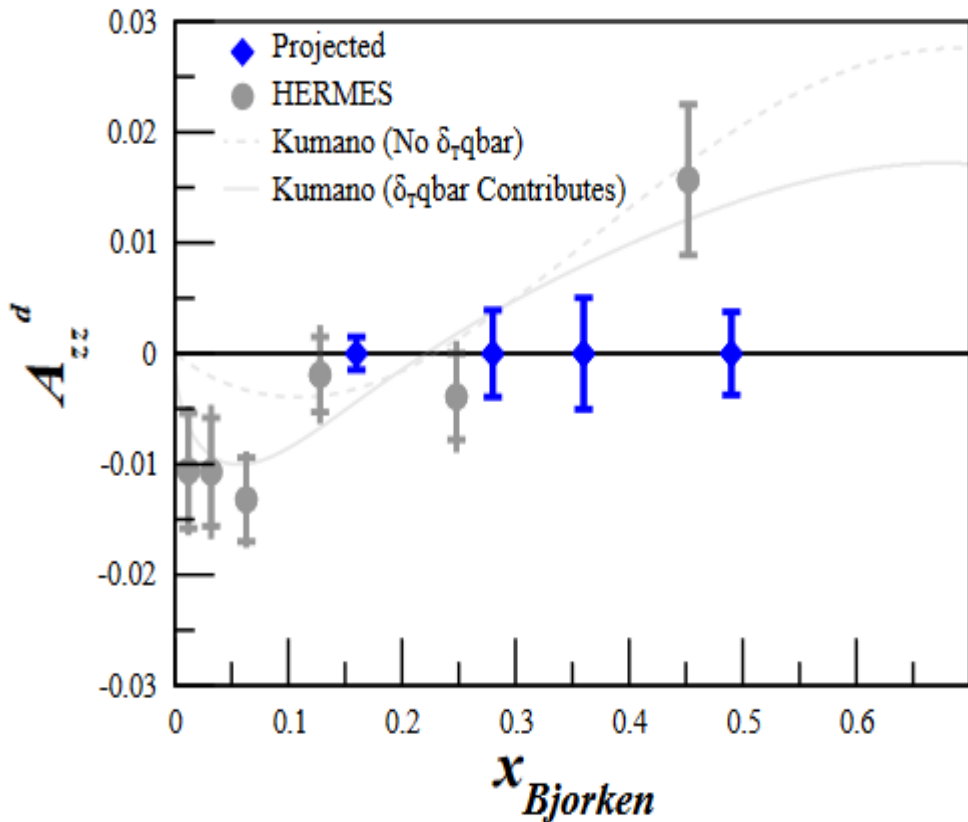
What is my error budget for charge monitoring?

Expected Size of the A_{zz}^d Asymmetry and Its Error

PR12-13-011

The Deuteron Tensor Structure Function b_1

A Proposal to Jefferson Lab PAC-40
(Update to PR12-11-110)



The model of Kumano predicts the largest value of A_{zz}^d .

This model is consistent with the HERMES data.

A significantly non-zero observation of A_{zz}^d is in principle feasible.

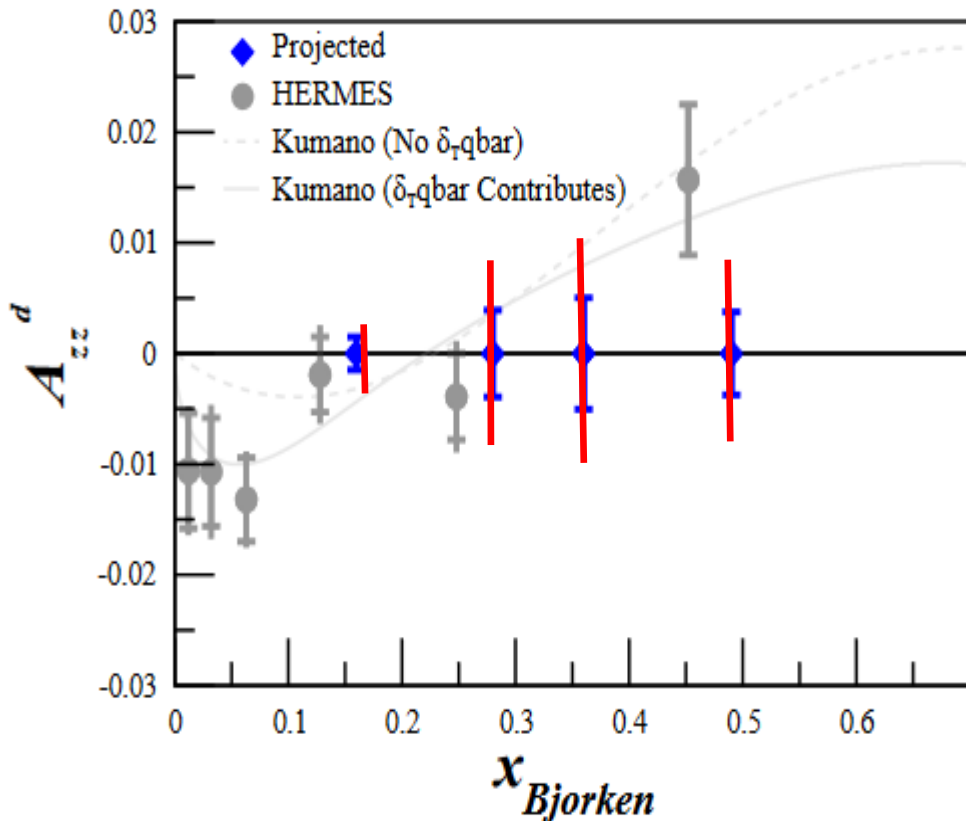
x_{bj}	Projected dA_{zz_stat} (from Table 2 of the proposal*)
0.16	1.5e-3
0.28	3.9e-3
0.36	5e-3
0.49	3.7e-3

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If we blow these errors by a factor of 2, it will be hard to distinguish the Kumano model from the null hypothesis.

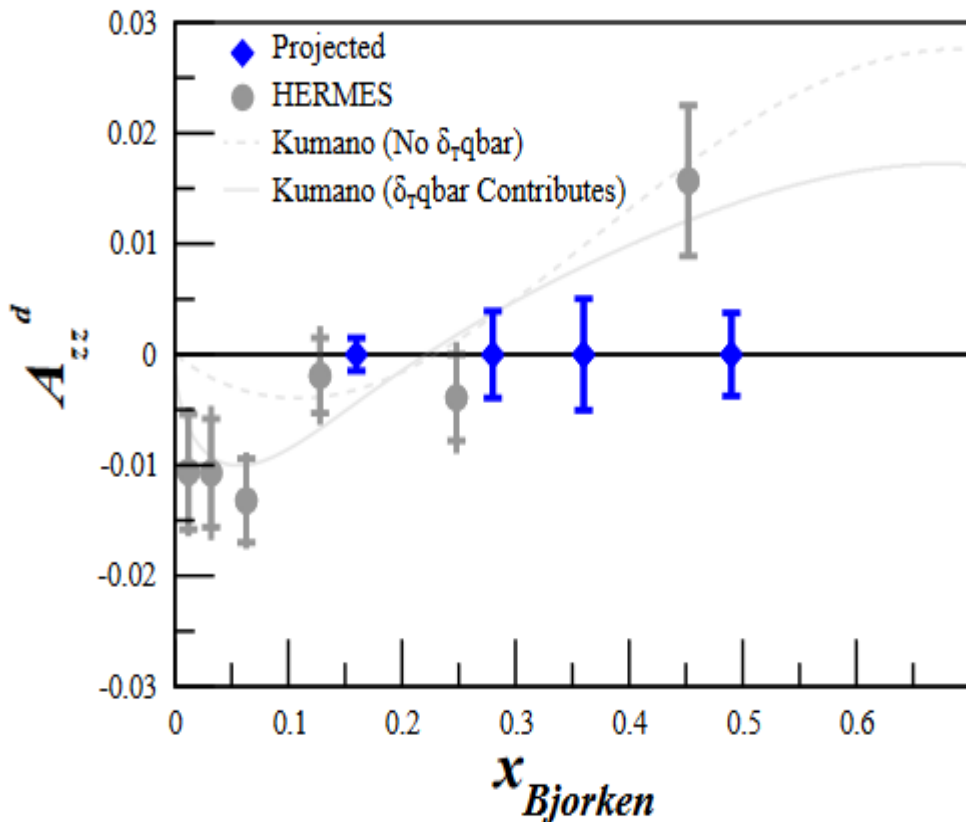
Expected Size of the Error on the Measured Asymmetry

PR12-13-011

The Deuteron Tensor Structure Function b_1

A Proposal to Jefferson Lab PAC-40
(Update to PR12-11-110)

I converted the projected dA_{zz_stat} to $dA_{measured}$ to see how large these statistical errors are before inflation by the corrections for target polarization and dilution.



x_{bj}	Projected dA_{zz_stat} (from Table 2 of the proposal)	$dA_{measured}$ $= f * P_{tgt} * dA_{zz_stat}$ $= 0.285 * 0.2 * dA_{zz_stat}$
0.16	1.5e-3	8.6e-5
0.28	3.9e-3	2.2e-4
0.36	5e-3	2.9e-4
0.49	3.7e-3	2.1e-4

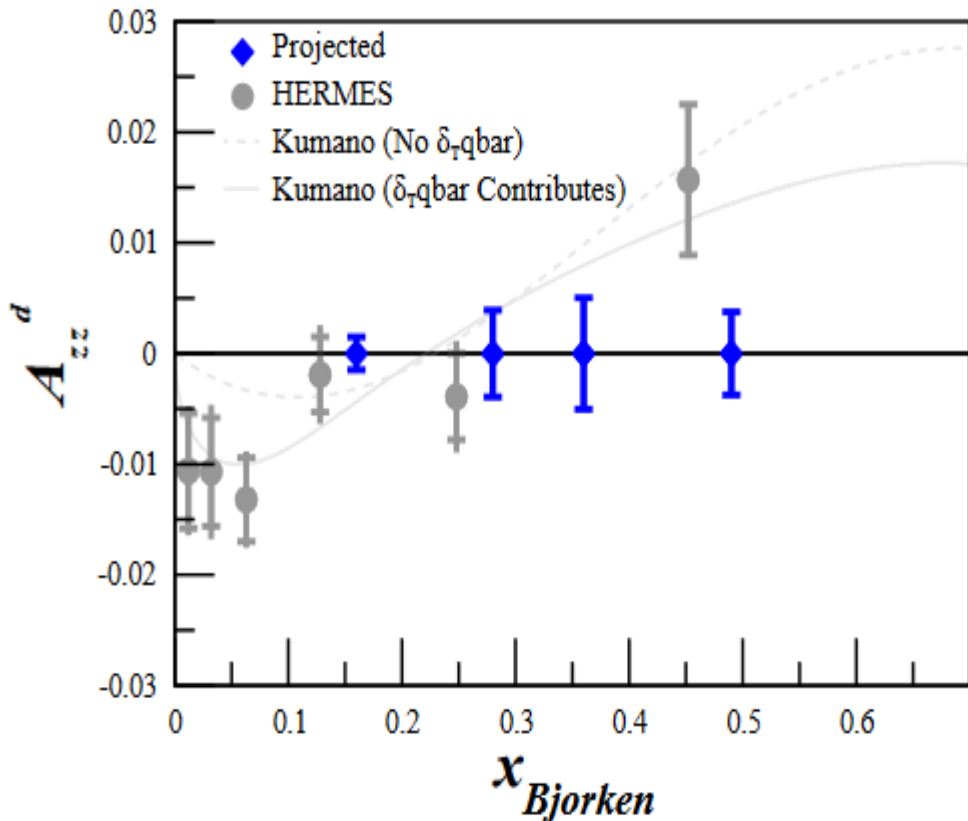
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x_{bj}	Projected dA_{zz_stat} (from Table 2 of the proposal)	$dA_{measured}$ $= f * P_{tgt} * dA_{zz_stat}$ $= 0.285 * 0.2 * dA_{zz_stat}$	Beam Days
0.16	1.5e-3	8.6e-5	6
0.28	3.9e-3	2.2e-4	9
0.36	5e-3	2.9e-4	15
0.49	3.7e-3	2.1e-4	30

The lowest x_b point is most challenging:

In only ~12 calendar days, all other random errors arguably need to average down to no larger than 2/3 of this, or ~5.7e-5.

Slow Cycles (Once per Beam Day)

Keeping all other random error below $5.7e-5$ in 6 days is the wrong way to think about it.

Each pair of (polarized, unpolarized) is a mini-experiment.

If we're cycling once per Beam Day, then we "only" need to keep all other random errors **below $1.4e-4$** in each mini-experiment.

X_{bj}	Projected dA_{zz_stat} (from Table 2 of the proposal*)	Projected $dA_{measured}$ $= f * Ptgt * dA_{zz_stat}$ $= 0.285 * 0.2 * dA_{zz_stat}$	$dA_{measured}$ Per Beam Day =column 3 x $\sqrt{\text{Beam Days}}$
0.16	$1.5e-3$	$8.6e-5$	$2.1e-4$
0.28	$3.9e-3$	$2.2e-4$	$6.6e-4$
0.36	$5e-3$	$2.9e-4$	$1.1e-3$
0.49	$3.7e-3$	$2.1e-4$	$1.2e-3$

Not-So-Slow Cycles (Once per Beam Hour)

If we're cycling once per Beam Hour, then we "only" need to keep all other random errors below $6.7e-4$ in each mini-experiment.

X_{bj}	Projected dA_{zz_stat} (from Table 2 of the proposal*)	Projected $dA_{measured}$ $= f * Ptgt * dA_{zz_stat}$ $= 0.285 * 0.2 * dA_{zz_stat}$	$dA_{measured}$ Per Beam Day =column 3 x $\sqrt{\text{Beam Hours}}$
0.16	$1.5e-3$	$8.6e-5$	$1.0e-3$
0.28	$3.9e-3$	$2.2e-4$	$3.2e-3$
0.36	$5e-3$	$2.9e-4$	$5.4e-3$
0.49	$3.7e-3$	$2.1e-4$	$5.6e-3$

There Are Many Potential Random Errors at the $O(1)E-4$ Level So Charge Monitoring Cannot Consume the Entire Budget!

Most of the random errors in the Yield calculation are:

Target stuff:	Bead settling or slumping changing the fill factor
	Target field variations changing the acceptance
	Target temperature differences between polarized and unpolarized
Detector stuff:	Errors in corrections for drifts in the deadtime
	Errors in corrections for drifts in the tracking efficiency
	Unmitigated drifts in the Gas Cerenkov thresholds
	Unmitigated drifts in the Pb Glass thresholds
Accelerator stuff:	Unmitigated drifts in beam energy
	Unmitigated drifts in beam position on target
Charge:	Charge normalization

Perhaps half of these were mentioned in the proposal.

Let's say charge monitoring takes half the error budget for additional random errors. (This allows for 4 errors of similar magnitude, added in quadrature, to saturate the total random error budget.) Then:

- if we do daily pairs, I have to do $7e-5$ or better, every day.
- if we do hourly pairs, I have to do $3.4e-4$ or better, every hour.

The answer to the question: What is my error budget for charge monitoring?

Hand-wavingly, we need to do $O(1)e^{-4}$ in the charge monitoring for each pair.

(ditto for other non-statistical random errors)

This specification is driven by the lowest x_{bj} point.

Faster target cycling than 1 beam day will make my job easier. But 1 clock hour cycles would hurt the statistical error bar.

Perhaps 2-4 cycles per shift will be a reasonable compromise?

Overview of BCMs and Their Stability

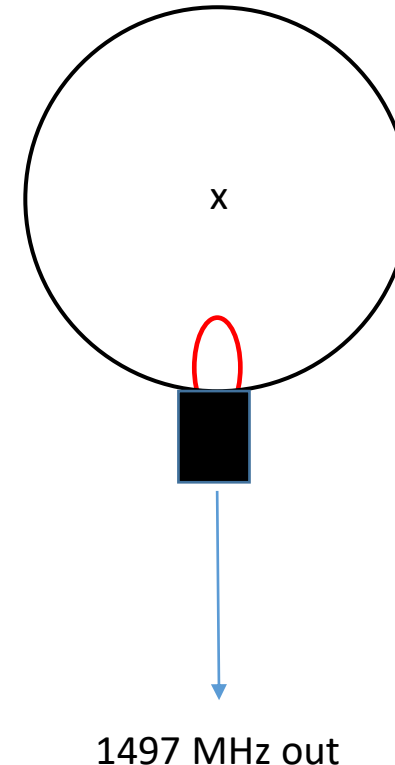
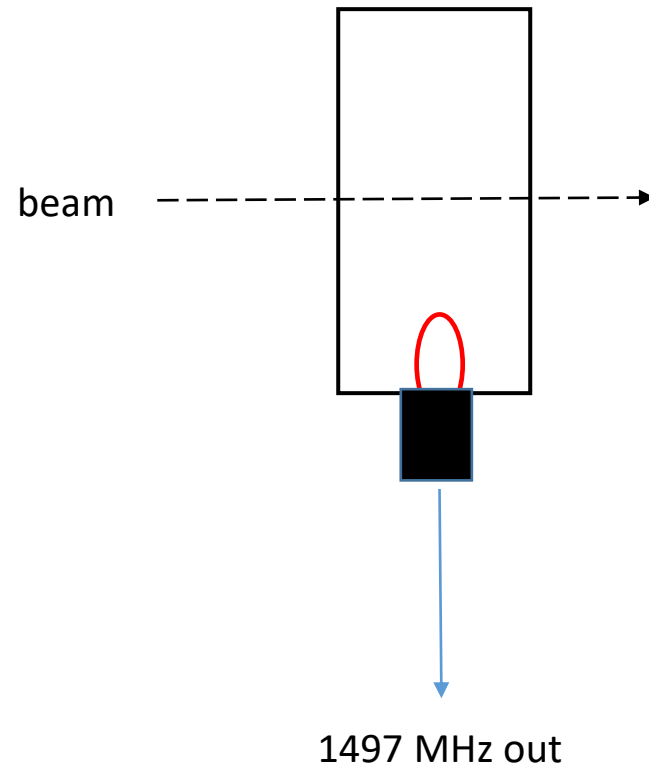
Beam Cavity Monitors (BCMs)

Jlab BCMs are resonant pillbox cavities operating in TM_{010} mode at 1.497 GHz. (the 3rd harmonic of the 499 MHz bunch frequency)

They are completely passive: when beam passes through them, it excites a mode with just the right phase of longitudinal electric field to extract power from the beam.

These BCMs act like $O(100)$ kOhm resistors, so linear processing of the RF to a voltage will be linearly proportional to the beam current.

An RF engineer would spec this as 1 μ A \rightarrow -40 dBm (or 0.1 microWatts). Even at 100 μ A the extracted power is only $I^2 \cdot R = 1$ mWatt, so there is no significant nonlinearity from heating.



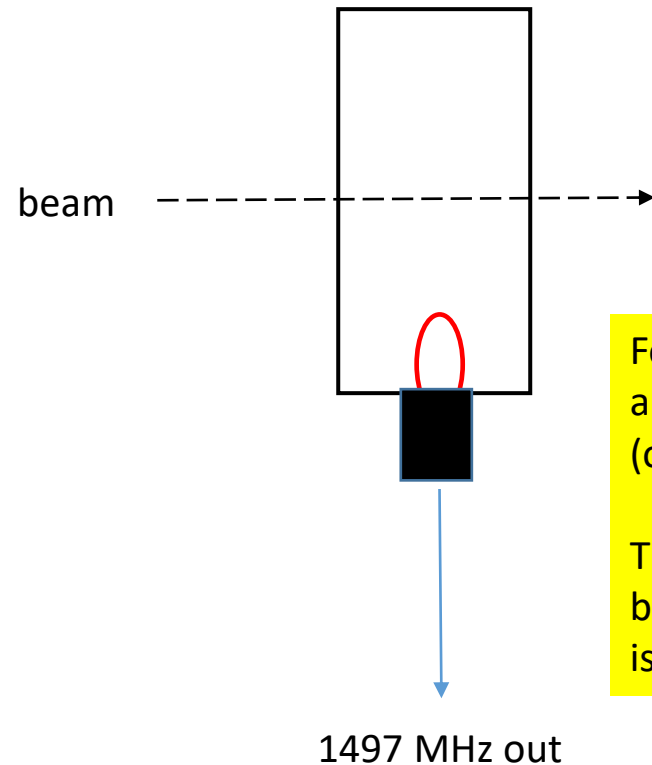
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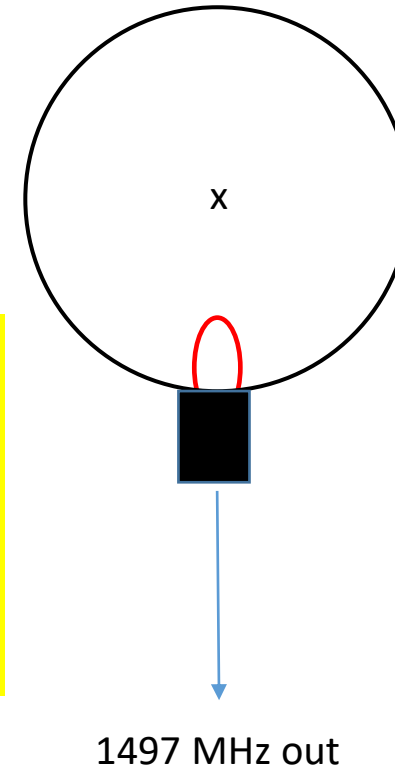
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For polarized ND3 target operation, a beam current of 100 nA is -60dBm (or 1 nanoWatt).

This is not as hard as radio-astronomy, but the BCM signal level is getting small-ish and more amplification is needed.



Making Temperature-stable BCMs

The TM_{010} mode is in principle independent of the cavity length, but does depend on the radius.

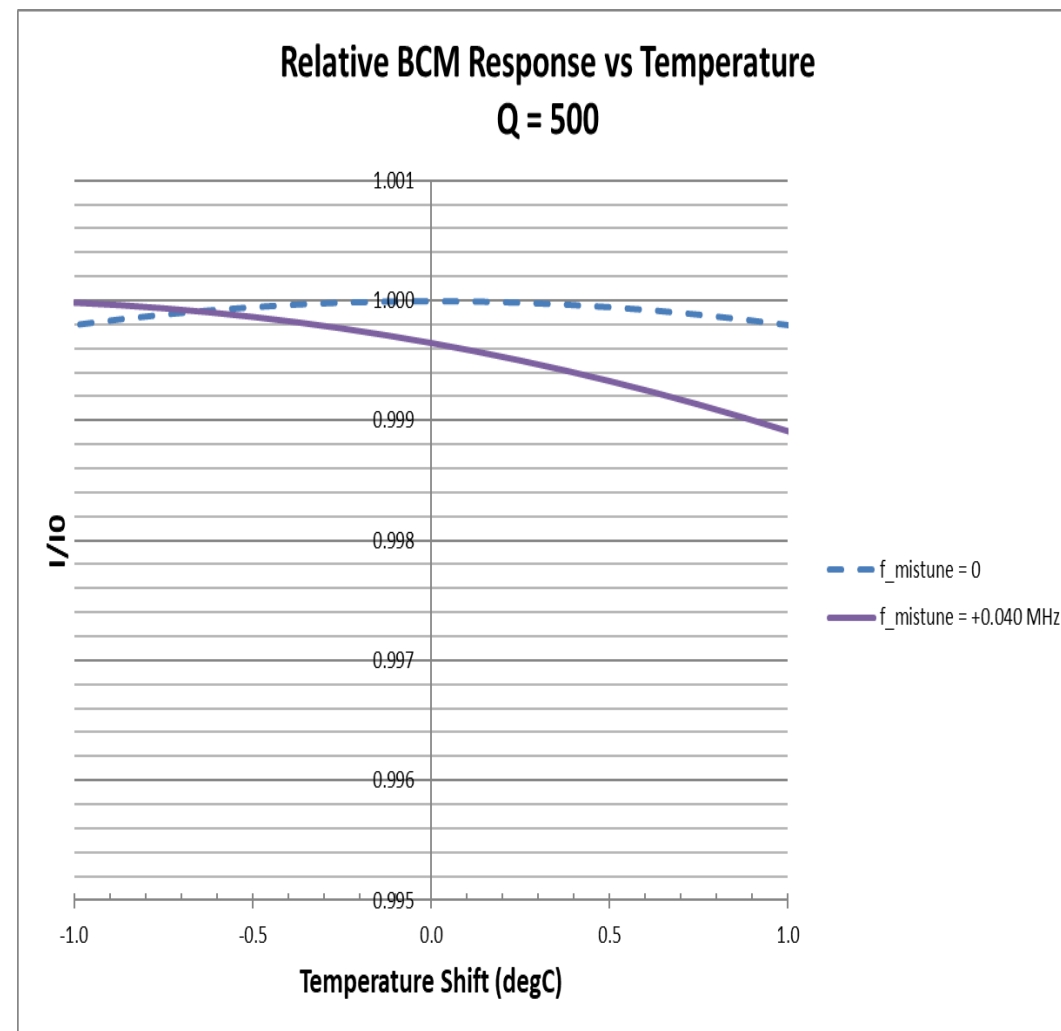
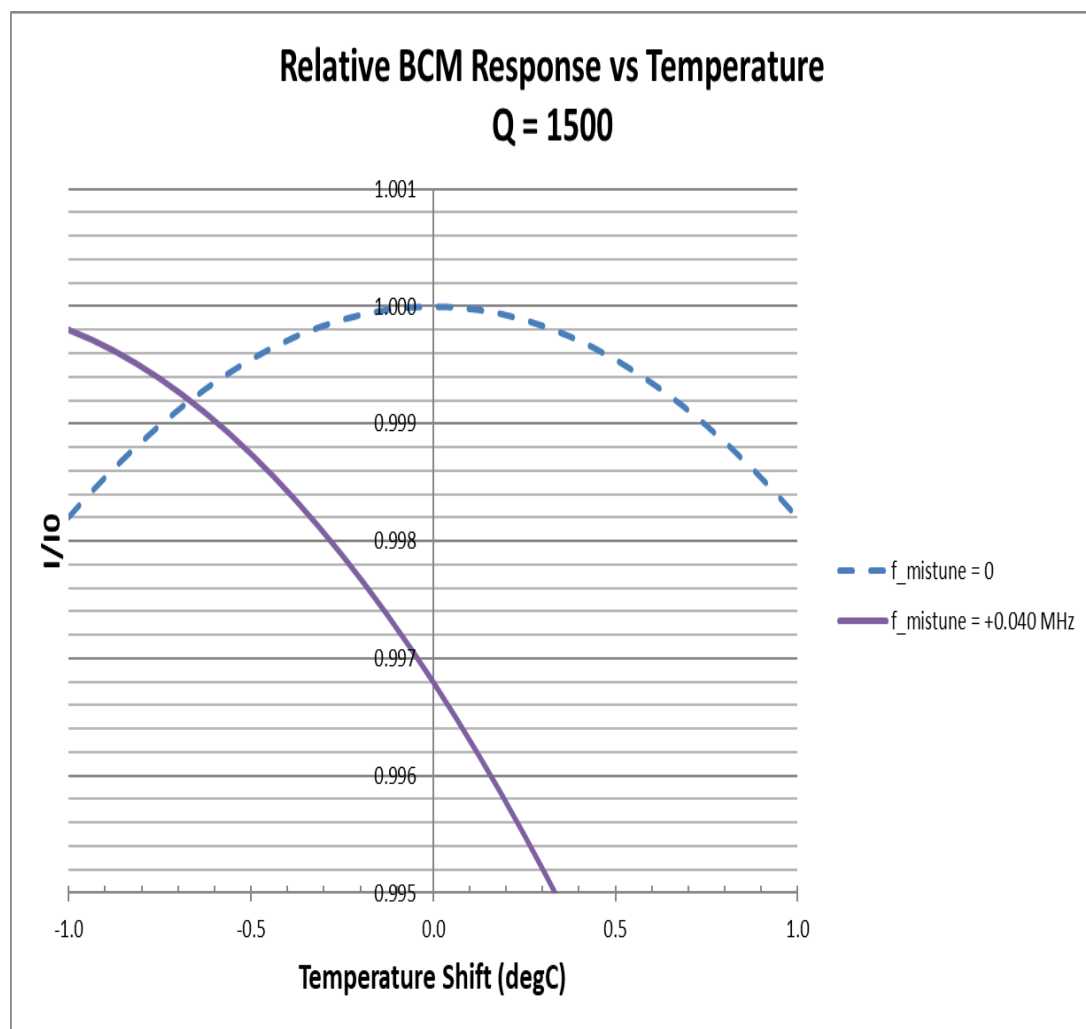
For charge measurements with excellent long term stability, one has to ensure that thermal expansion and contraction cycles don't change the radius and cause the cavity to wander on and off resonance. The problem is worse for higher Q cavities, since their resonance peak is narrower ($\Delta f = f_0/Q$). The problem is also worse if the cavity is tuned slightly off the 1.497 GHz of the beam.

- ✓ The cavities were therefore made of stainless steel (SS) which has a low thermal coefficient of expansion.
- ✓ The loaded Q was a fairly modest ~ 1500 .
(Even a good conductor struggles at 1.5 GHz due to the shallow skin depth, and Stainless Steel is not a good conductor.)
- ✓ The cavities are tuned within ± 20 kHz. (roughly 1 part in 10^5 of the 1.497 GHz beam frequency)
- ✓ The cavities are temperature controlled to ± 0.2 F, going thru a full cycle every 5 minutes.

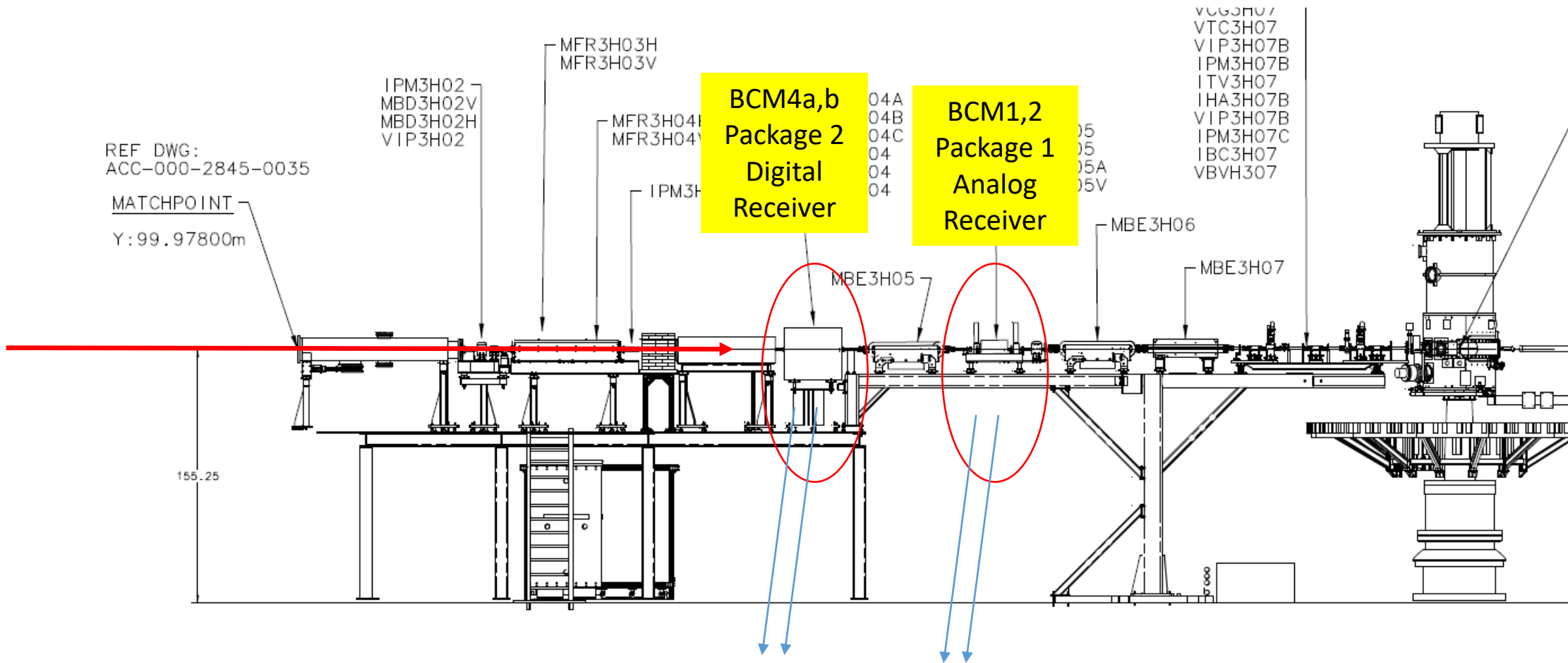
Model Stability for Q = 500 vs 1500

The resulting cavities were still too temperature sensitive for precision cross-section work! So decades ago, I asked them to be de-Q'd to ~500.

(The accelerator Machine Protection System soon followed suit. This is how most BCMs at Jlab are operated.)



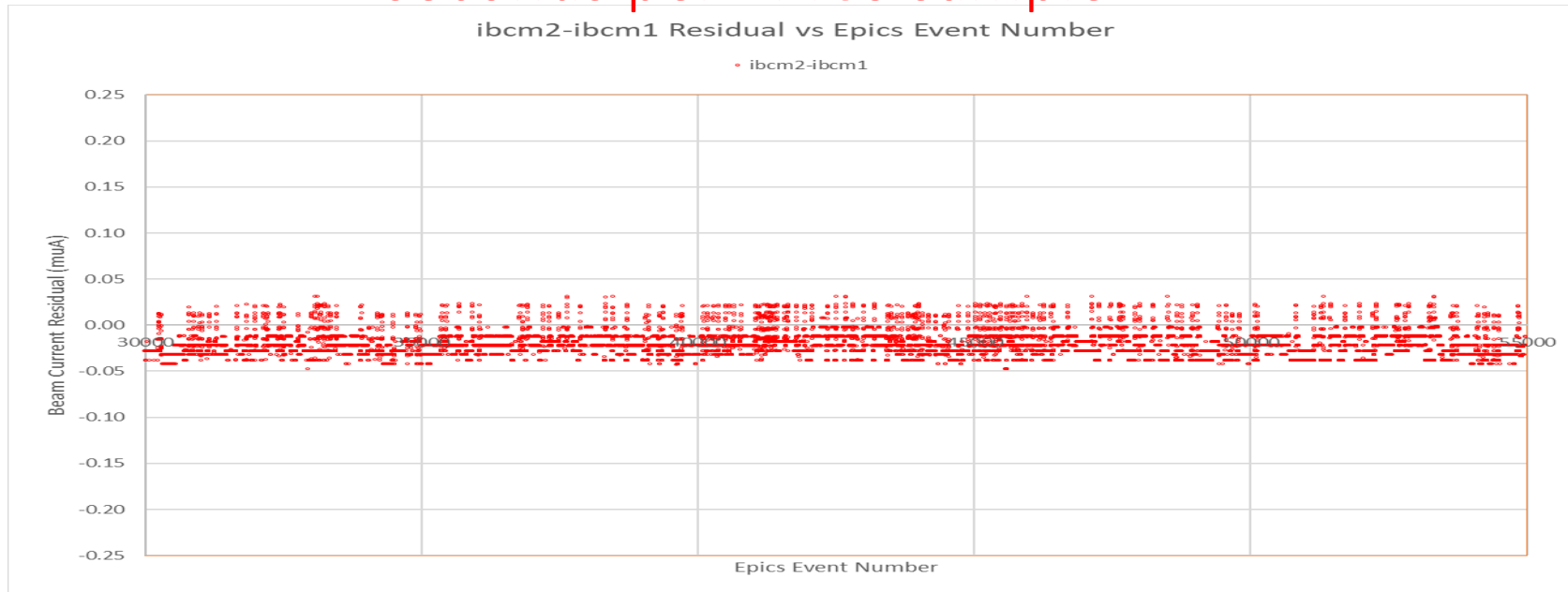
Two Thermally Separate BCM Enclosures on the Hall C Beamline



All the 1.497 GHz RF cables run upstairs to the counting house.
Package 1 uses better quality RF cable.

BCM Differences During 17 Hours at 58 muA

2 seconds per EPICS sample



Package 1
(better RF cable)
BCM1 - BCM2
(less than $\pm 2 \times 10^{-4}$ relative)

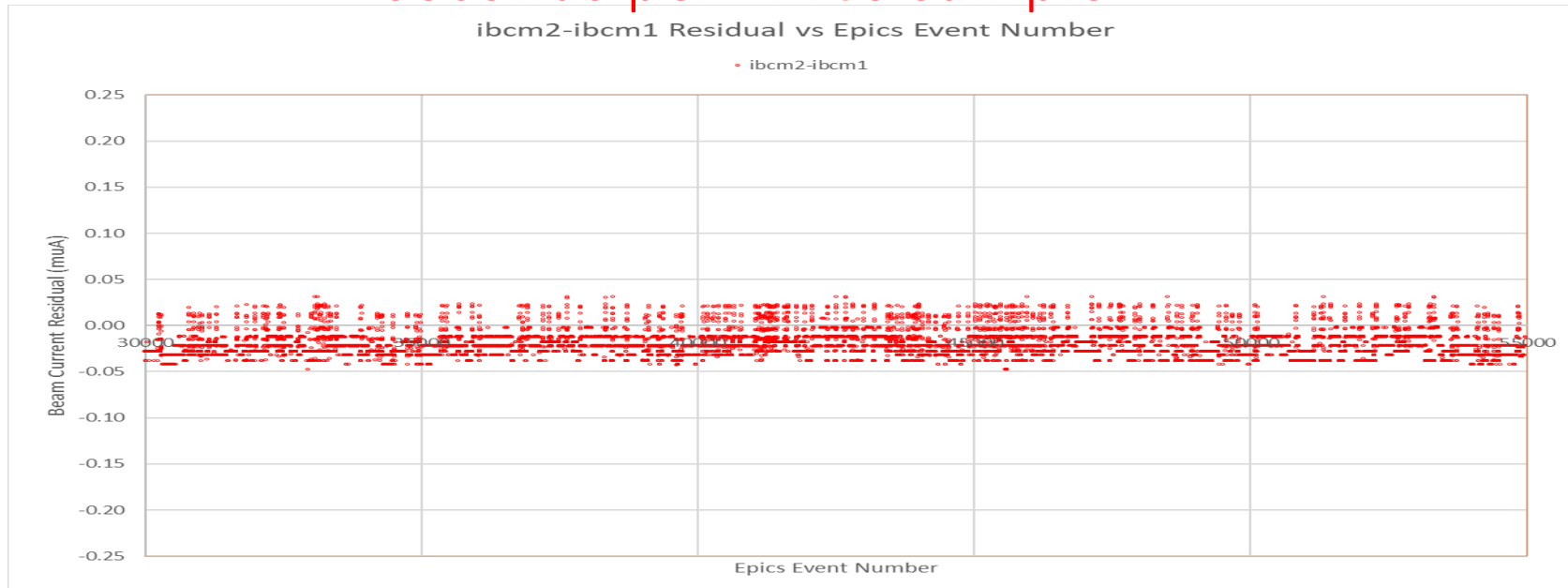
limited by V-to-F
counting noise

BCM Differences During 17 Hours at 58 muA

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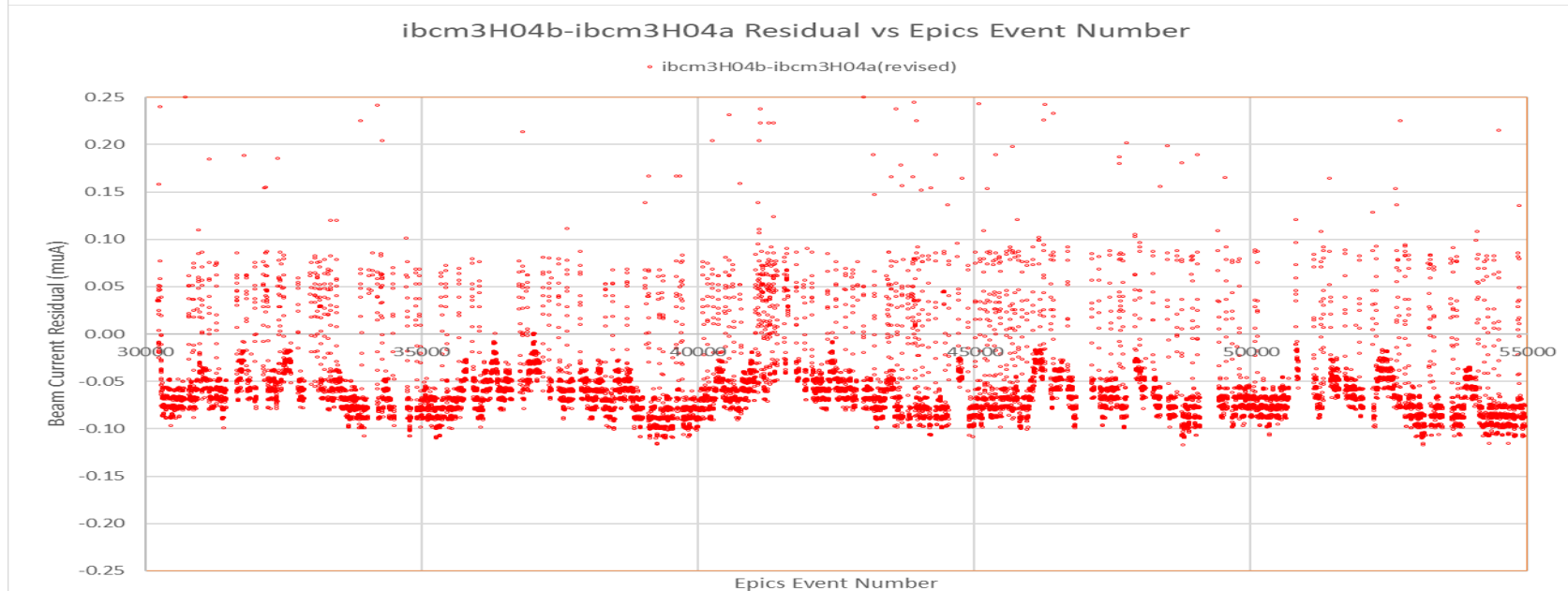
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Package 2
BCM4B – BCM4A
($\pm 4 \times 10^{-4}$ relative)
2.8 hour period

This is almost certainly
due to temperature
variations, but where is
not clear.

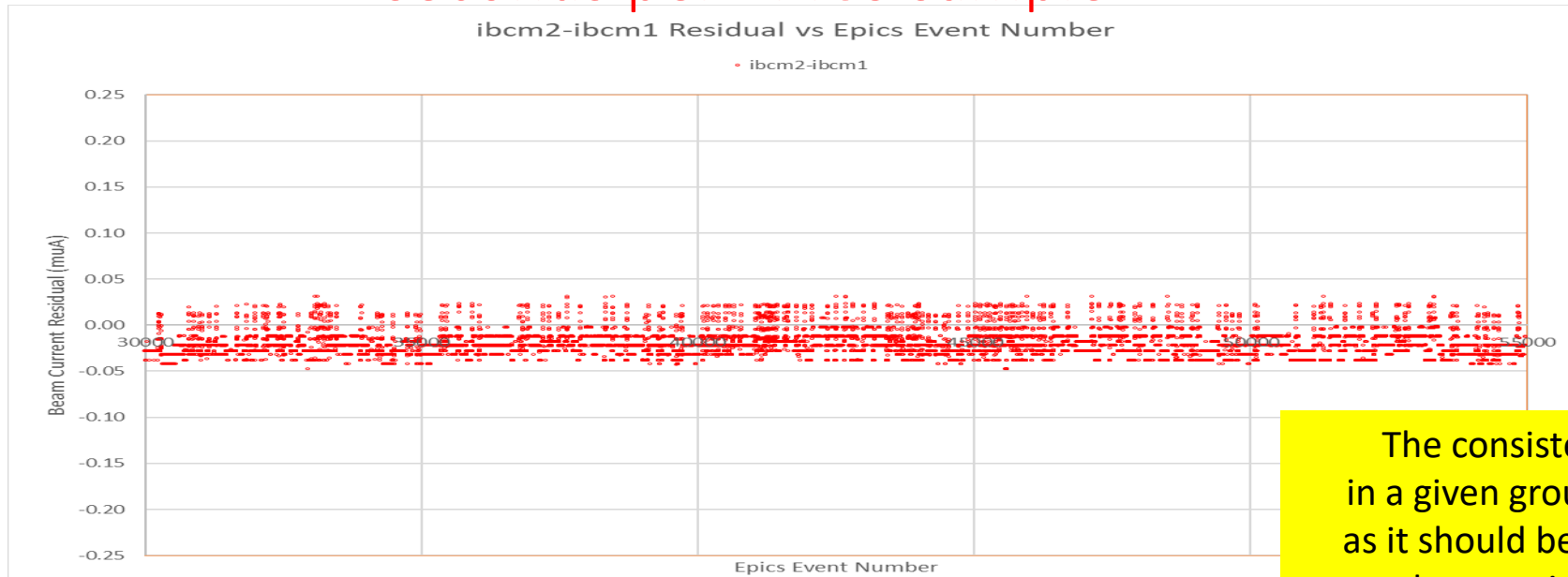


BCM Differences During 17 Hours at 58 muA

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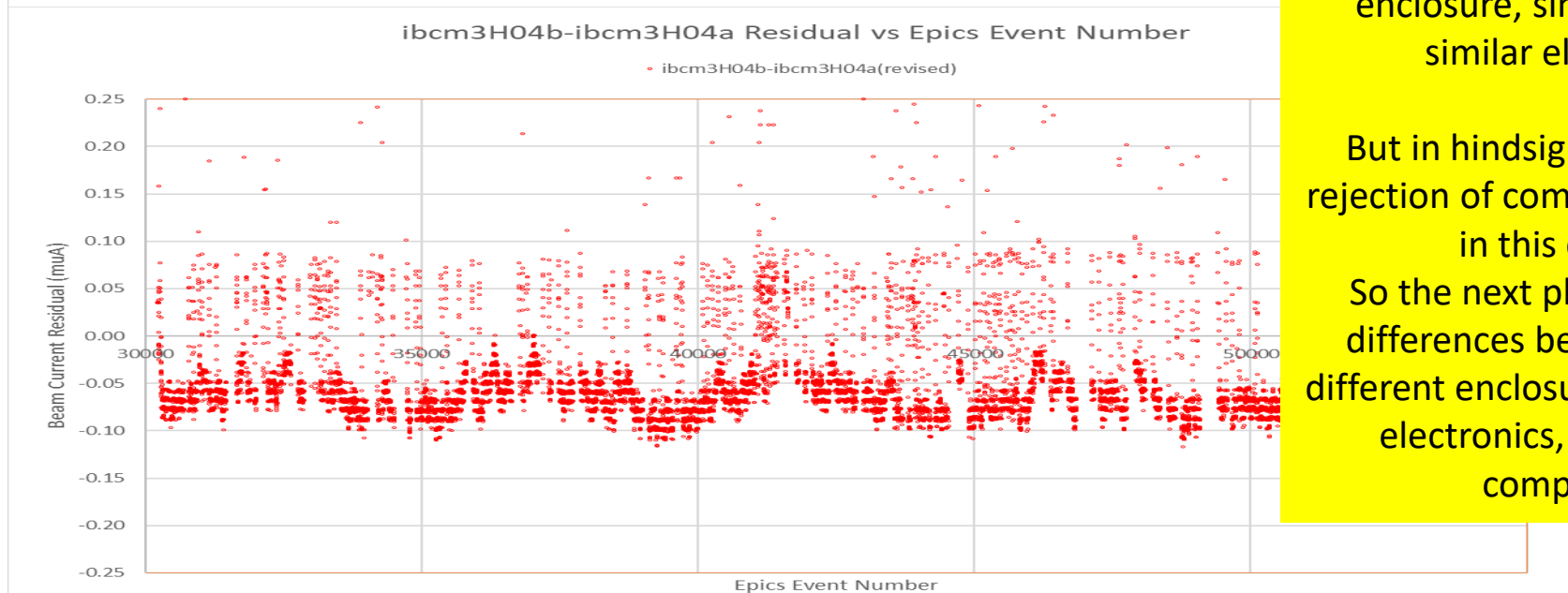
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This is almost certainly
due to temperature
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not clear.



The consistency of BCMs
in a given group is promising,
as it should be (same thermal
enclosure, similar cable run,
similar electronics).

But in hindsight, there's huge
rejection of common-mode errors
in this exercise.

So the next plot, which takes
differences between BCMs in
different enclosures with different
electronics, is a more fair
comparison.

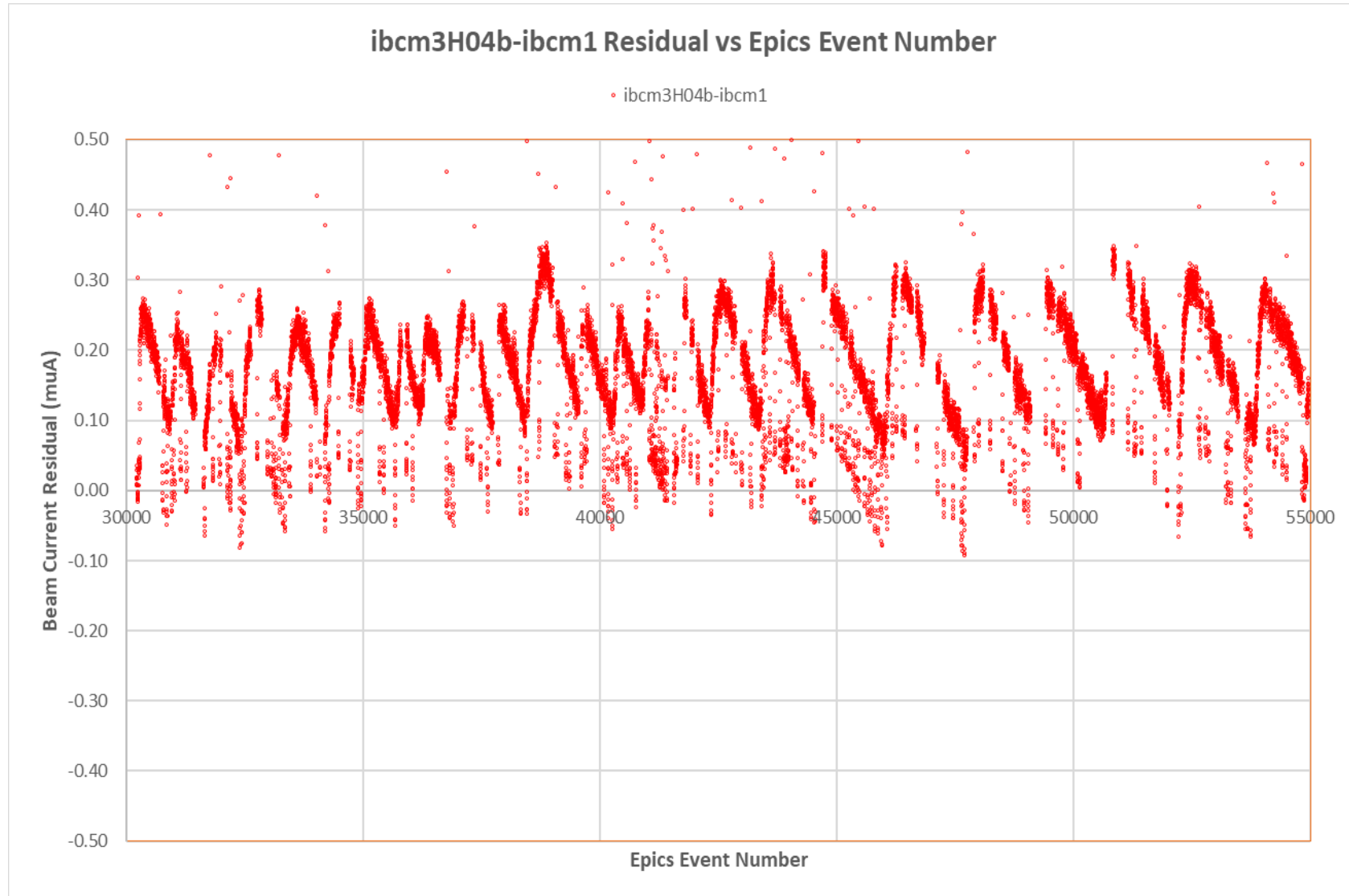
BCM Differences During 17 Hours at 58 muA

2 seconds per EPICS sample

Package 1 rel. Package 2

BCM4B - BCMa
(+0.2% relative)
25 minute period

This is 5x larger, and with
a completely different
period, than that seen on
the previous slides.



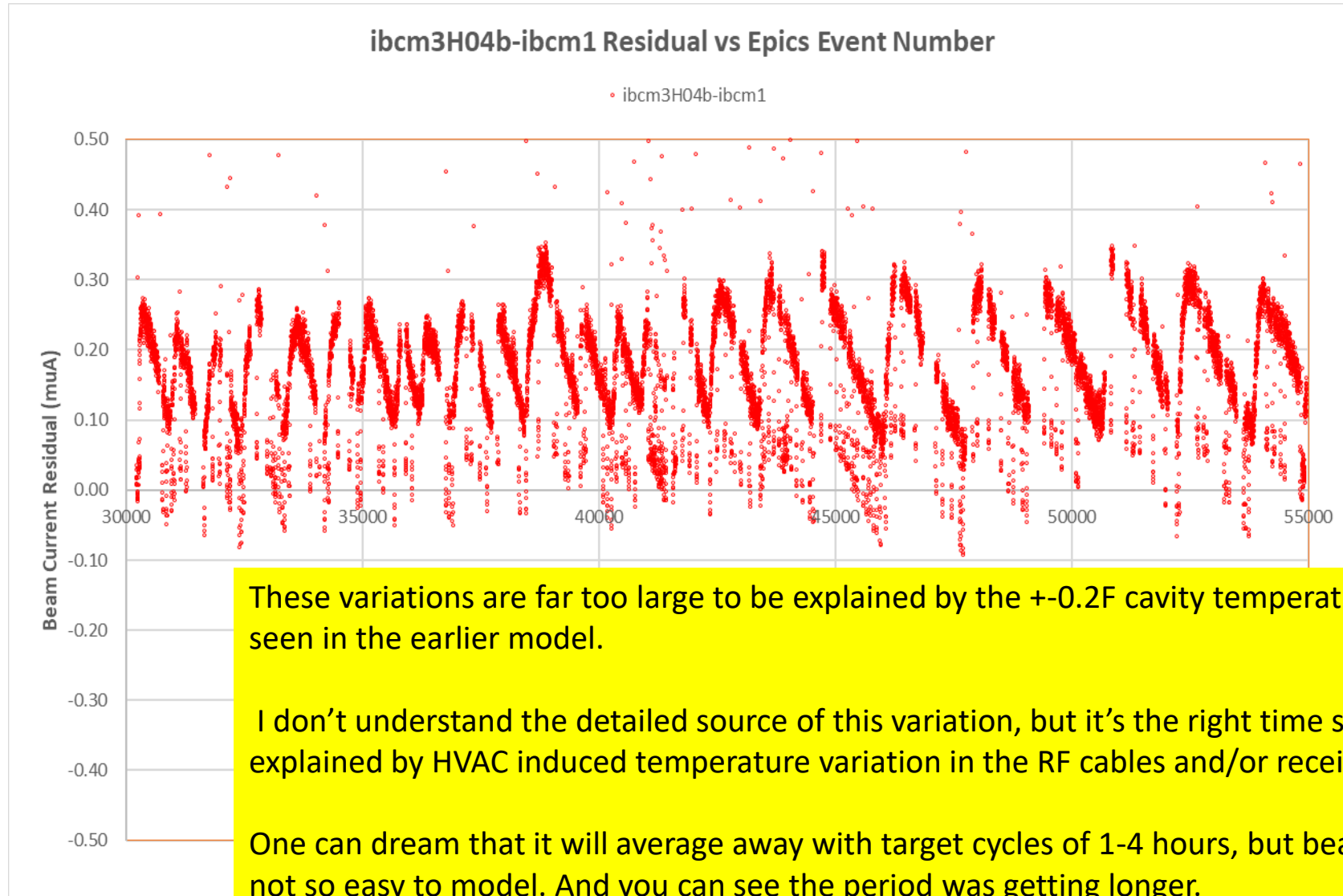
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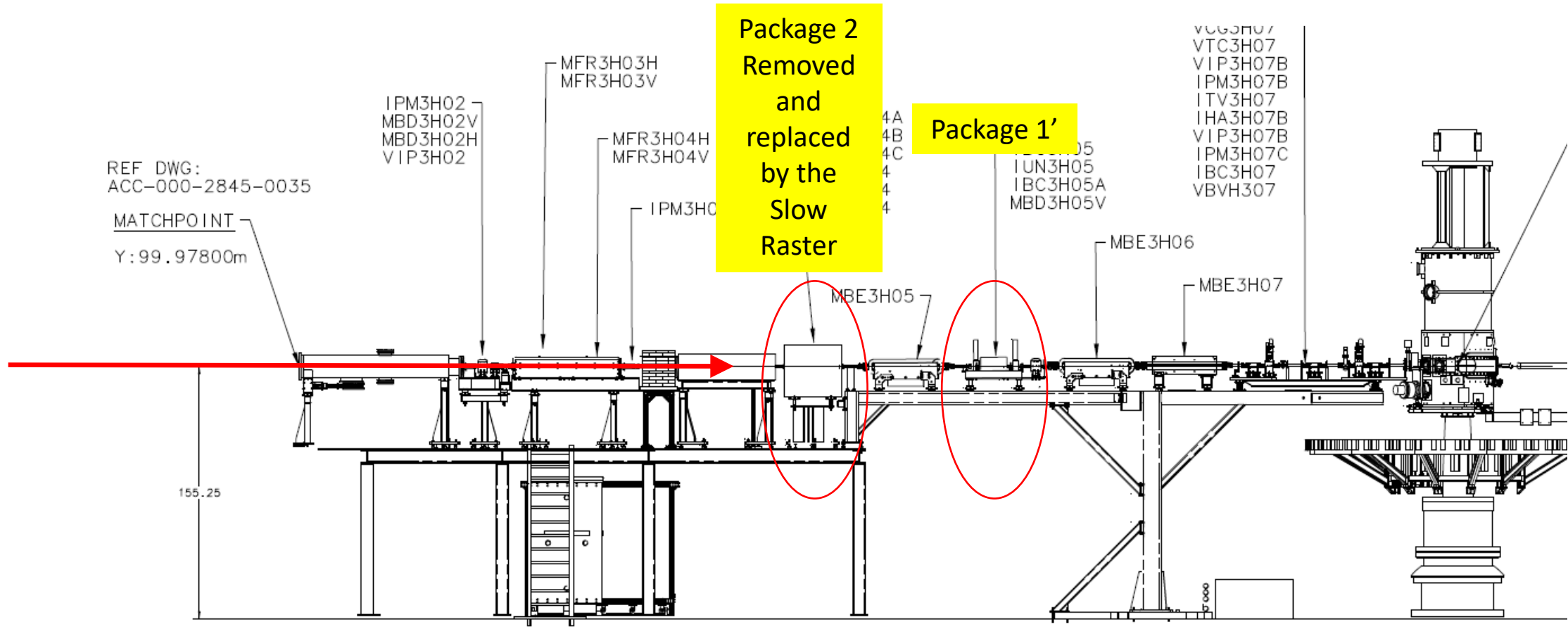
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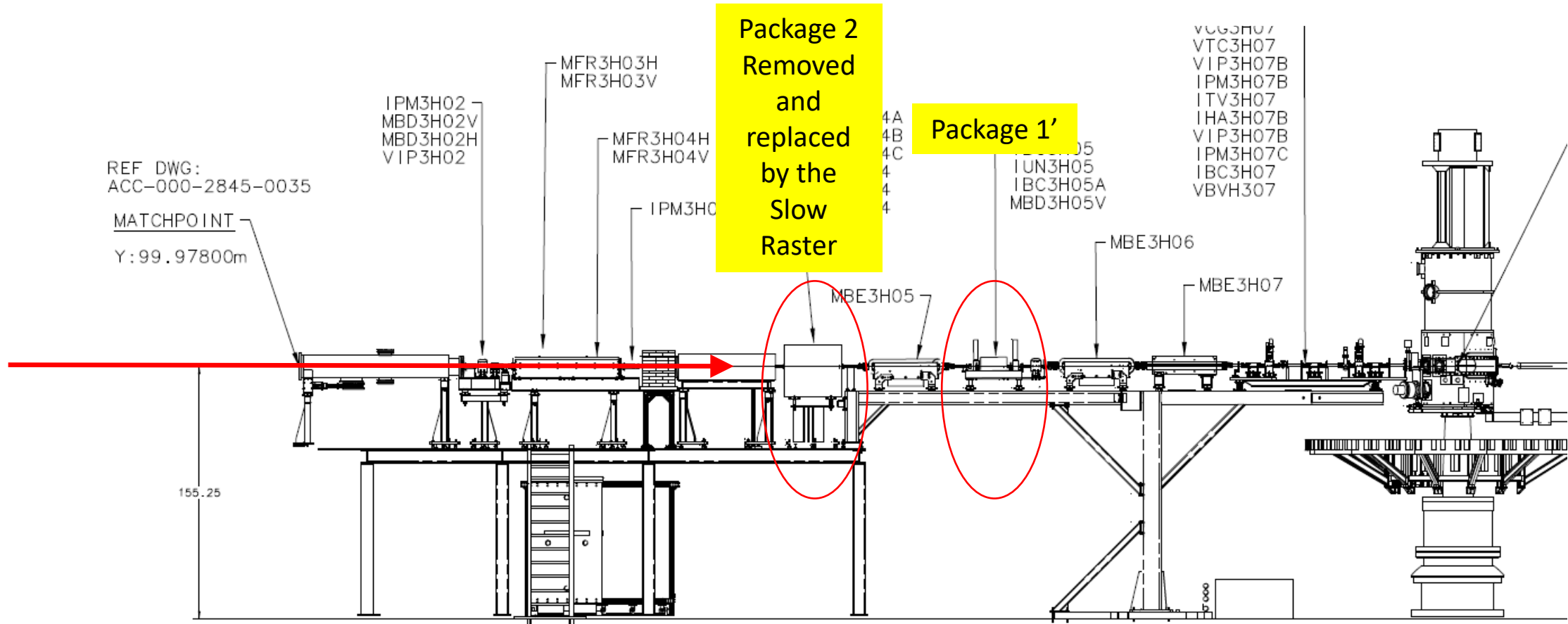
What's the BCM Situation for Polarized Target Running?

We will have less information.



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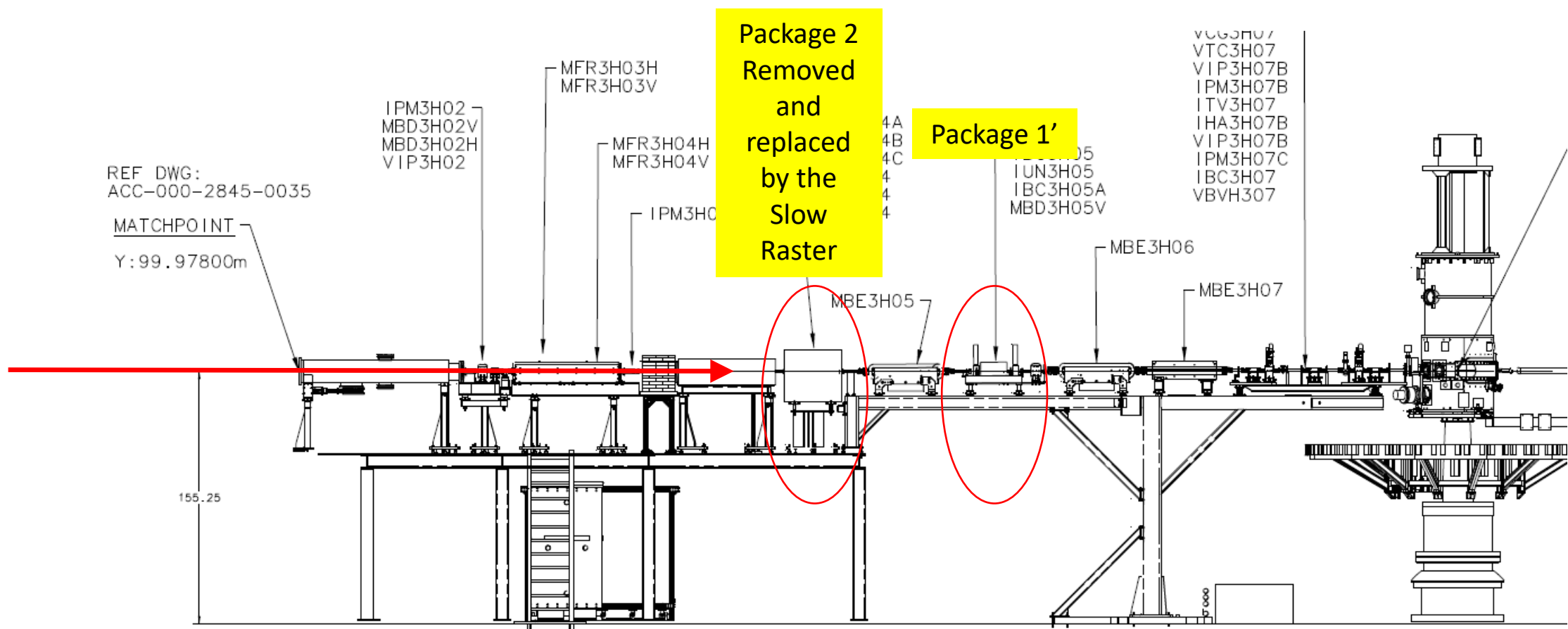
RF cables are already the highest quality except for ~10m jumpers.

Lowest risk, lowest effort concept for Package 1':

- upgrade those ~10m jumpers (moots any rad-damage questions)
- upgrade the thermal enclosure
- split the RF 50:50 upstairs to feed 2 analog and 2 digital channels

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- upgrade the thermal enclosure
- split the RF 50:50 upstairs to feed 2 analog and 2 digital channels

I can do some tempco calculations and temperature checks. But if I don't find a smoking gun to explain the +0.2% variation, the only choice in the data analysis the students will have is between analog and digital receivers. So fingers crossed on the temperature stability of the cavities, cables, and tuners of Package 1'.

Section Summary on BCMs and Their Stability

Measurement of the beam current with non-intercepting BCMs to achieve pair-level random errors at the $\pm 0.1\%$ level is not a given.

And lest we forget, our goal is $O(0.01)\%$!

BCM cavity stability is probably a solved problem (low tempco Stainless Steel, low Q, regulated temperature, accurate tuning).

But we still have to run 1.497 GHz through an inherently lossy and temperature-dependent cable to a location where the receiver electronics won't get rad-damaged. And the receivers will have their own tempcos.

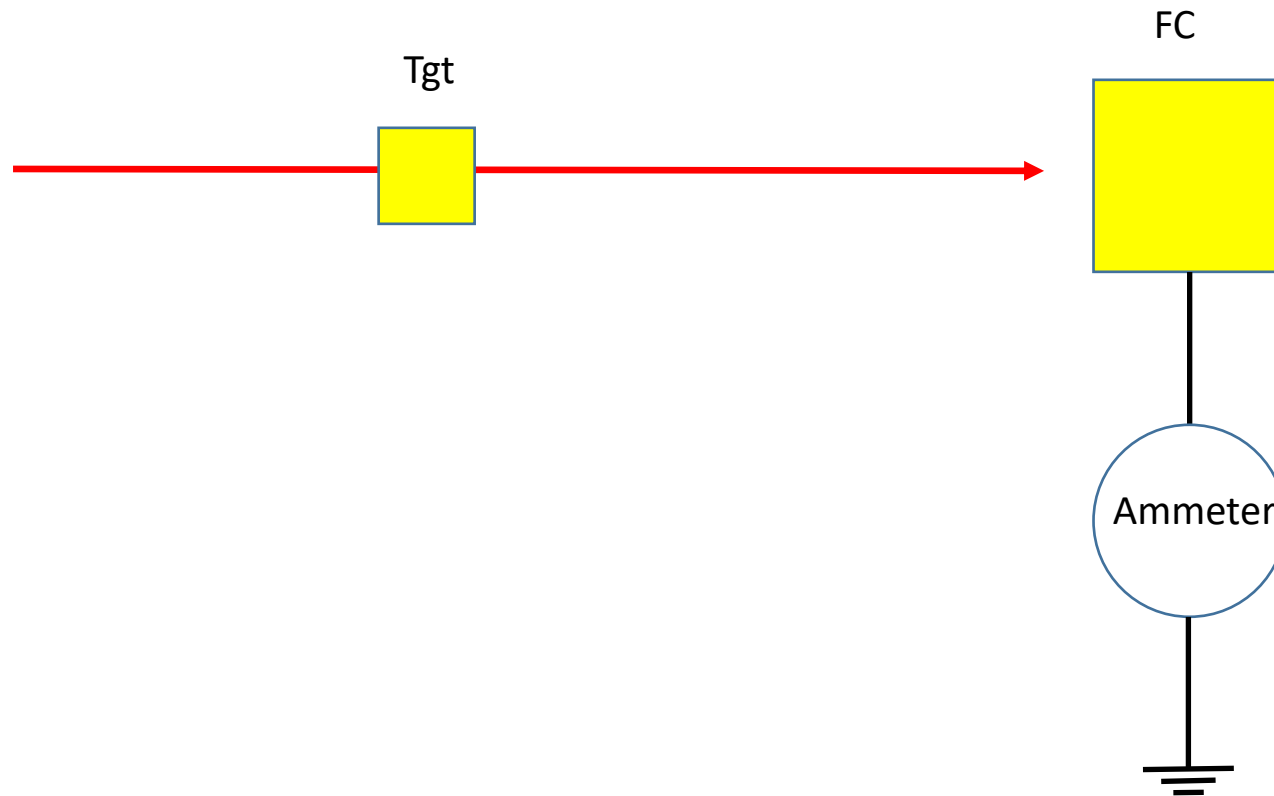
It's probably not impossible to achieve $O(0.01)\%$, but we should hedge our bets.

I'd Like to sell you on the idea of a ~~Timeshare~~ Faraday Cup



Basic Idea of a Faraday Cup (FC)

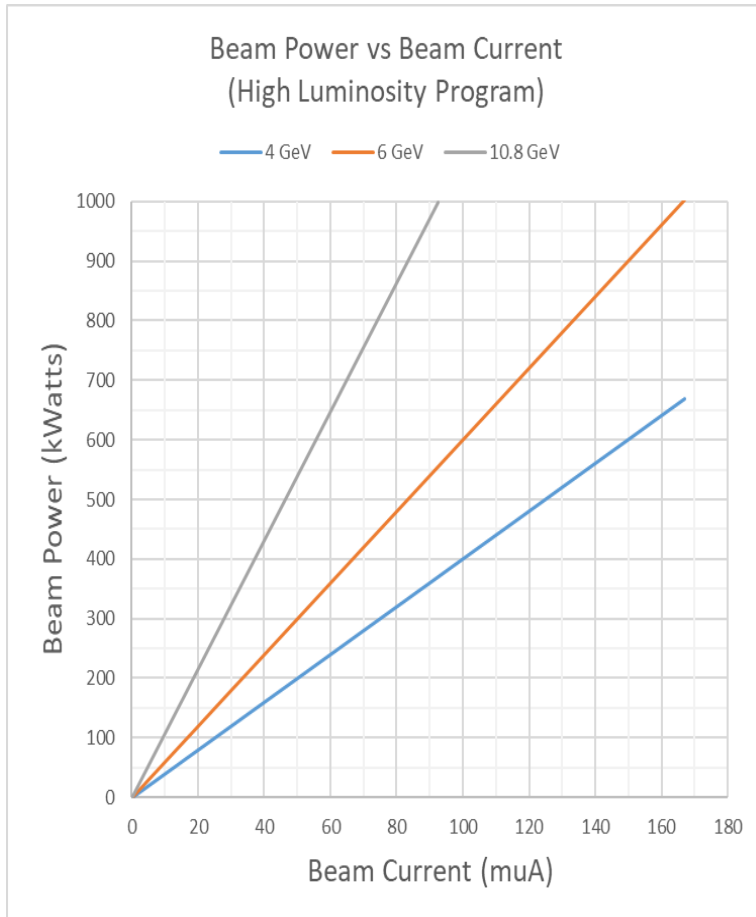
Simple in principle: one captures the vast majority of the beam charge, and measures the current.



Beam Power Levels

The FC also has to absorb most of the beam energy. But electrons shower, leading to maximum energy deposition at $\sim 5-6 X_0$.

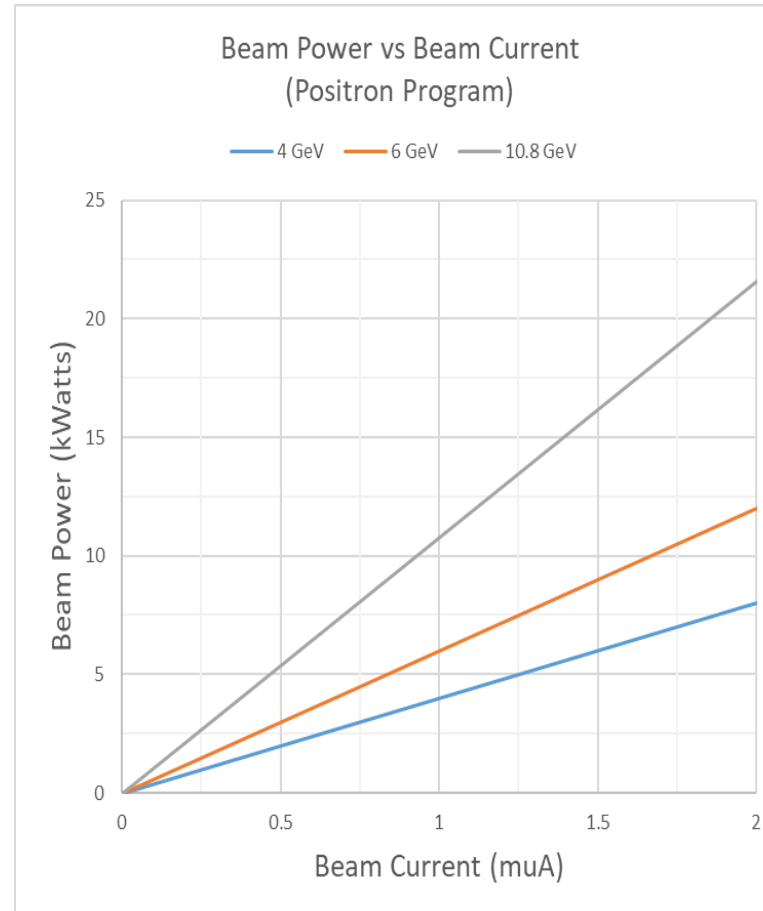
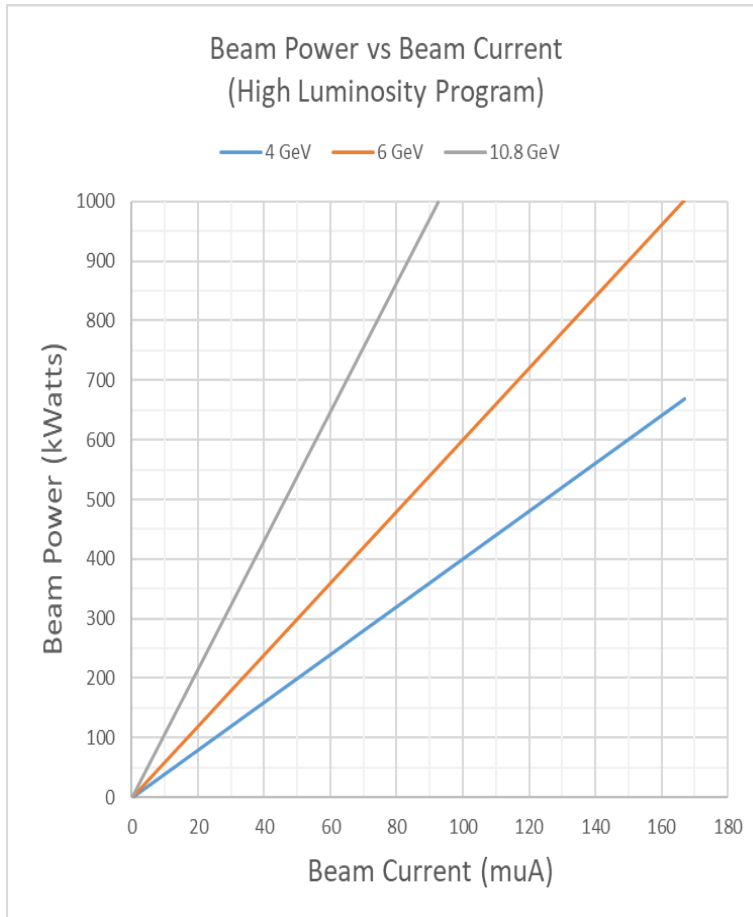
- For the Hall A/C high luminosity program reaching ~ 1 Mwatt, this arguably made a FC impractical.



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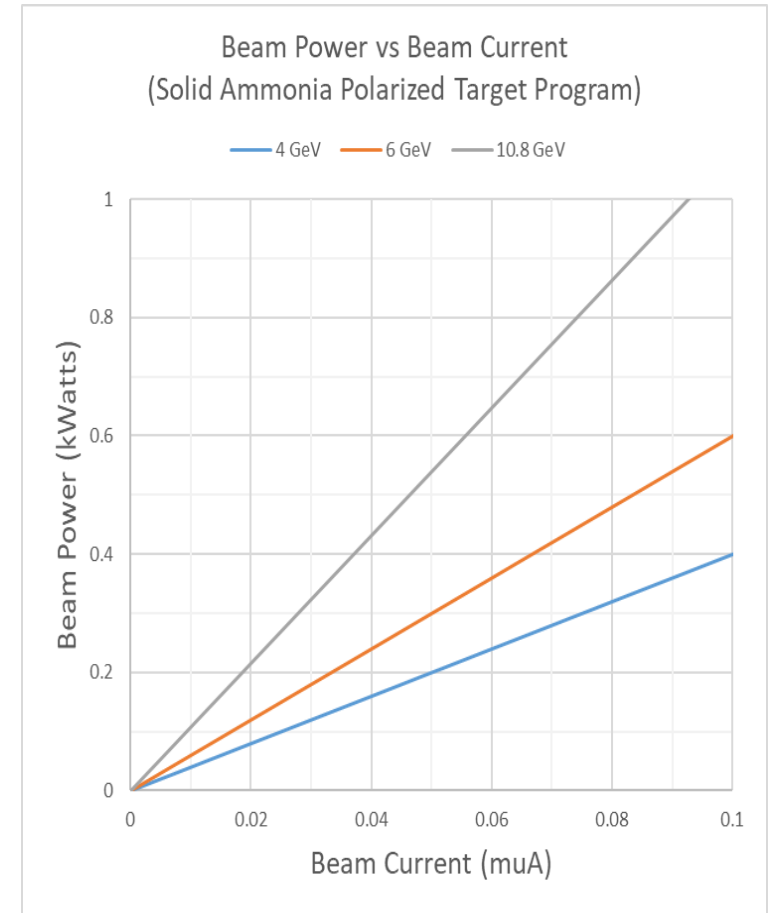
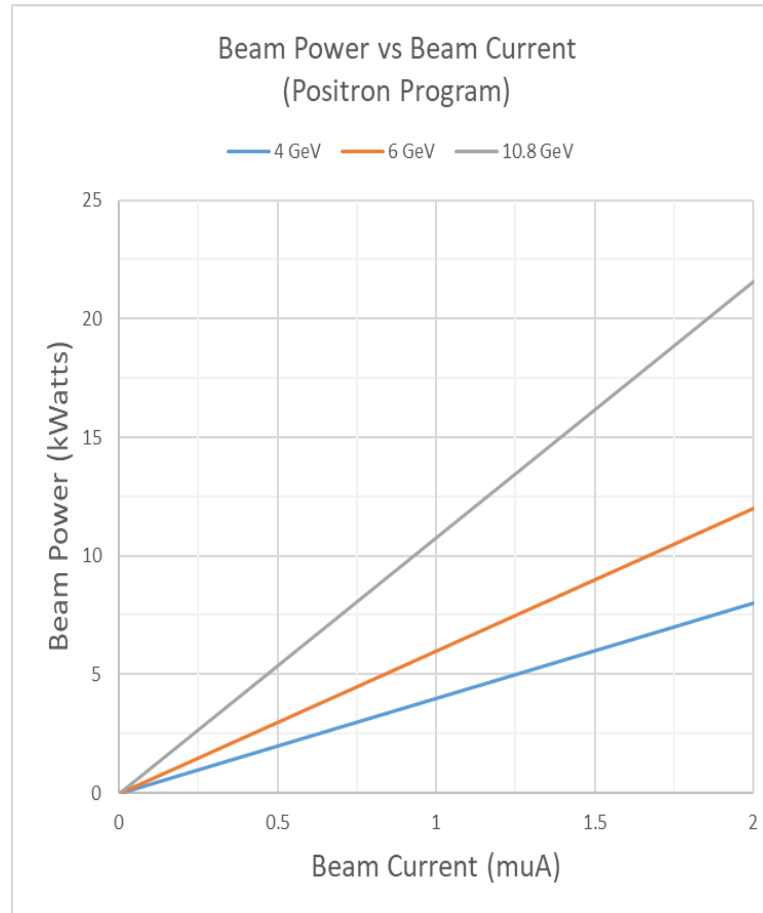
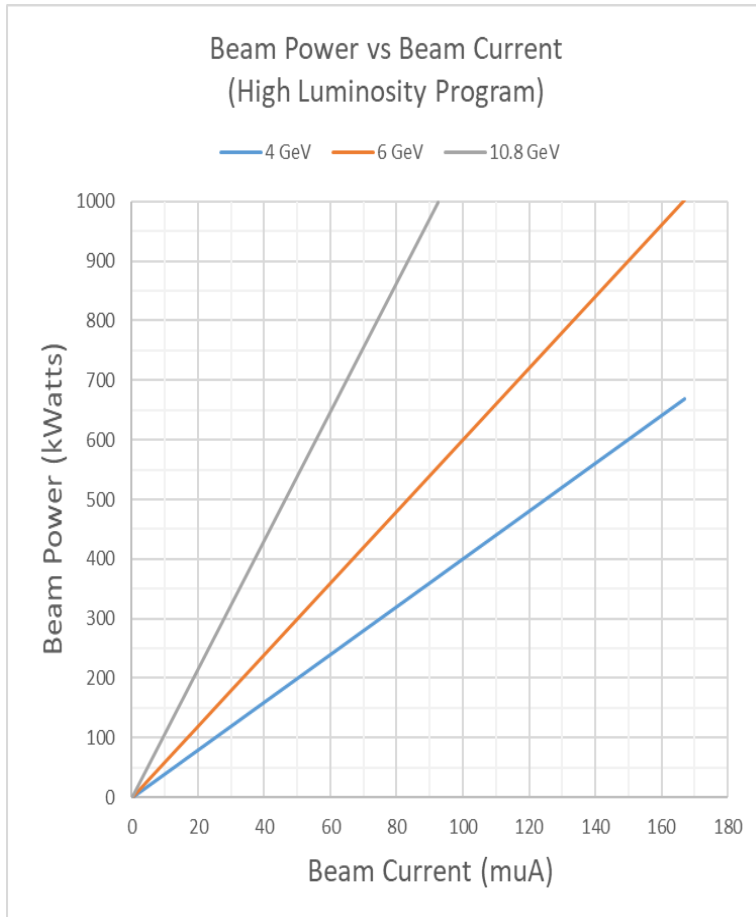
- For the Hall A/C high luminosity program reaching ~ 1 MWatt, this arguably made a FC impractical.
- For the anticipated positron program reaching ~ 22 kWatts, a FC looks challenging but likely feasible. (Solving the calibration problem!)



Beam Power Levels

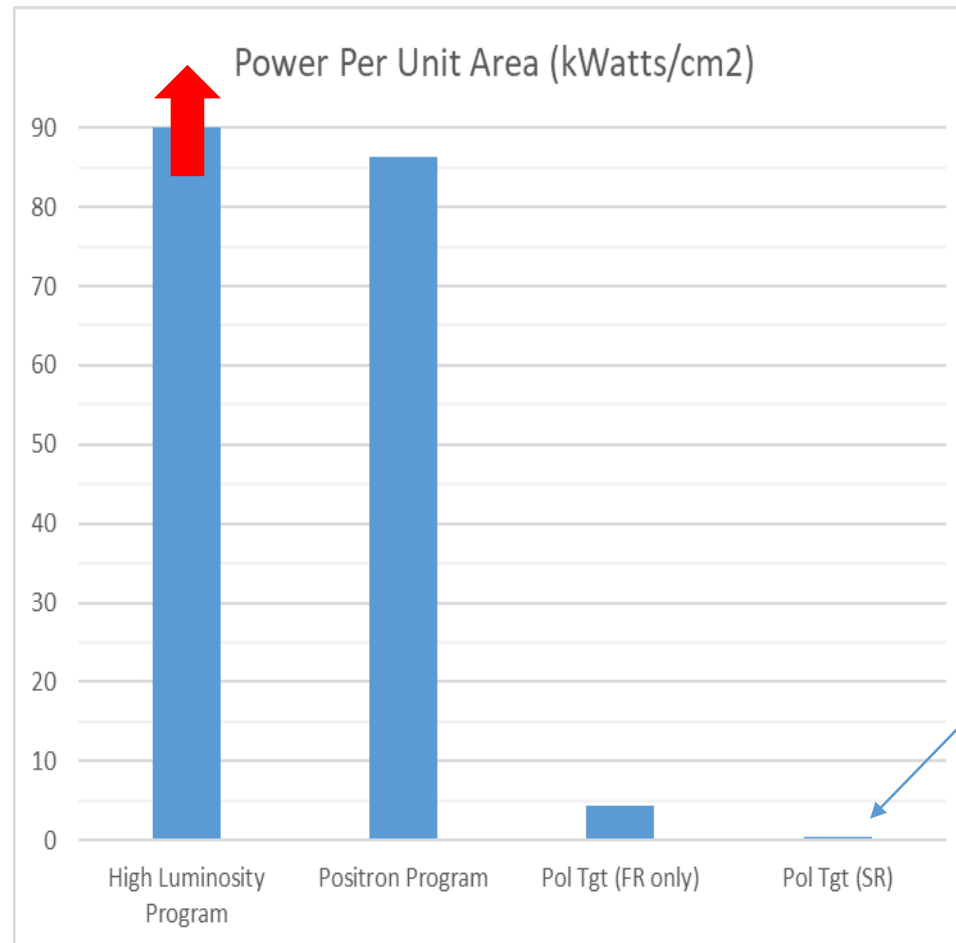
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- For the Hall A/C high luminosity program reaching ~ 1 Mwatt, this arguably made a FC impractical.
- For the anticipated positron program reaching ~ 22 kWatts, a FC looks challenging but likely feasible. (Solving the calibration problem!)
- For the solid ammonia polarized target program at less than ~ 1 kWatt, a FC looks relatively easy.



Beam Power Per Unit Area

In terms of maximum beam power per unit area at the face of the FC, with the Slow Raster enabled, the Pol Tgt Program value is at most 0.3 kWatts/cm². It seems there's no way to melt anything.



With the slow raster enabled, we are here.

Classical Errors in FCs

We stand on the shoulders of giants who have already recognized and solved most systematic effects in FC's for GeV-scale electron beams.

Problem	Mitigations	Comment
Charged Backscatter	Re-entrant geometry (to recapture backscatter); Graphite facing (low secondary emission coefficient); Transverse B field (traps low energy electrons)	Only a problem for b1 if it's time dependent. So design to keep backscatter small, and keep the FC guts under high vacuum.
Electron Shower Penetration	Need enough radiation lengths in radial and longitudinal directions.	Not a source of random error for b1.
Leakage paths to ground	Excellent vacuum to suppress neutralization by ions; Insulating stand-offs; Disconnect any ion gauges during operation	Only a problem for b1 if vacuum goes bad.
Target out-scattering	Entrance hole needs to be large enough that the measured current is insensitive to small changes in beam position.	Given the large size of the beam at the FC due to the Slow Raster and thick target, this needs careful analysis.
Asymmetric charge losses	Secondary e- emission off the rear face due to muons; Probably more mu- than mu+ (cuz more neutrons?); Proton escape matters while neutron losses do not;	Not a source of random error for b1.

The possibly unique feature here is the low current of 100 nA at 100% duty factor. (1e-4 would be 10 pA, which is a very small current.)
Measuring the current of old, pulsed beams was sometimes 4 orders of magnitude less sensitive to small, time-varying current offsets!

The good news is one gets to integrate for thousands of seconds in b1. But life gets harder below the 1 nA level.

SLAC “High” Power Faraday Cup (1 kWatt CW)

You can see the many features already mentioned: large but re-entrant aperture, high vacuum jacket, insulating stand-offs, oodles of radiation lengths, etc.

With no cooling water, and apparently sketchy FEA, they were unclear on the melting threshold so kept average power < 1 kWatt.

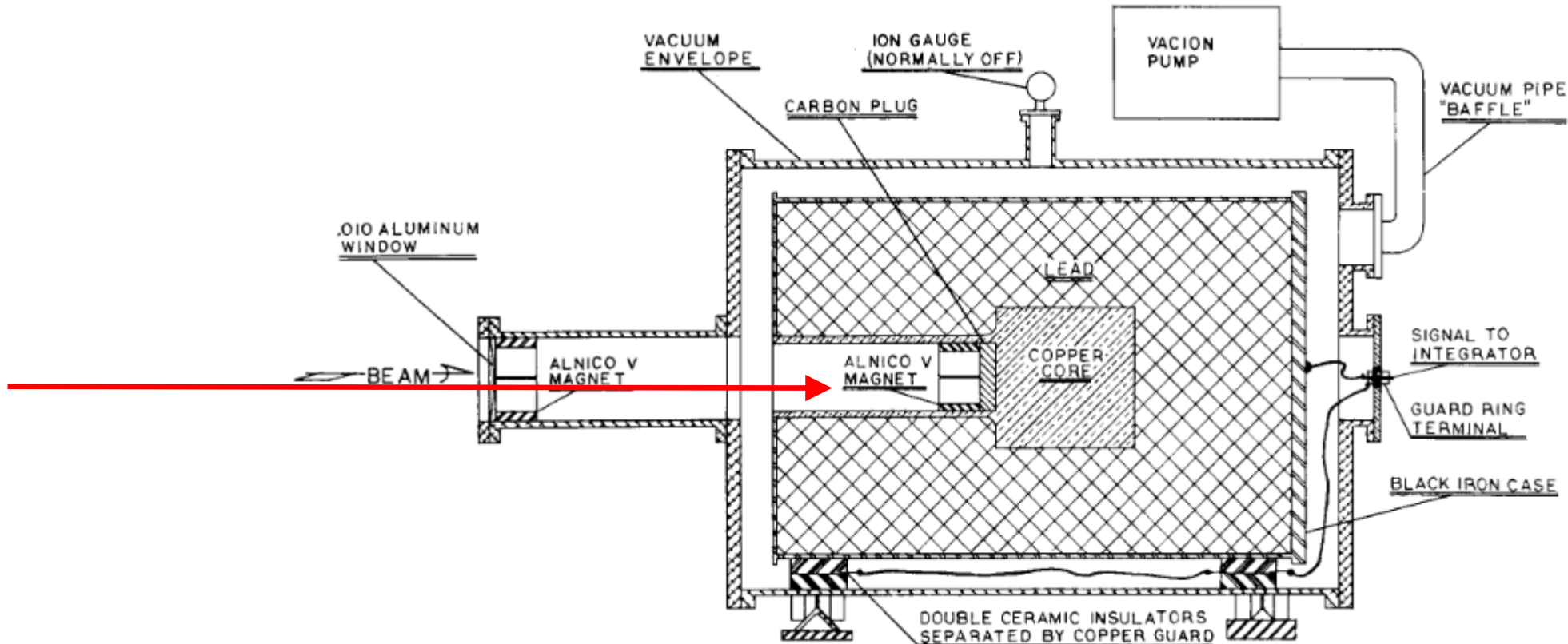


Fig. 1. Schematic diagram of the 20-GeV Faraday cup. The scale is indicated by the copper core which is 10" deep and 10" in dia.

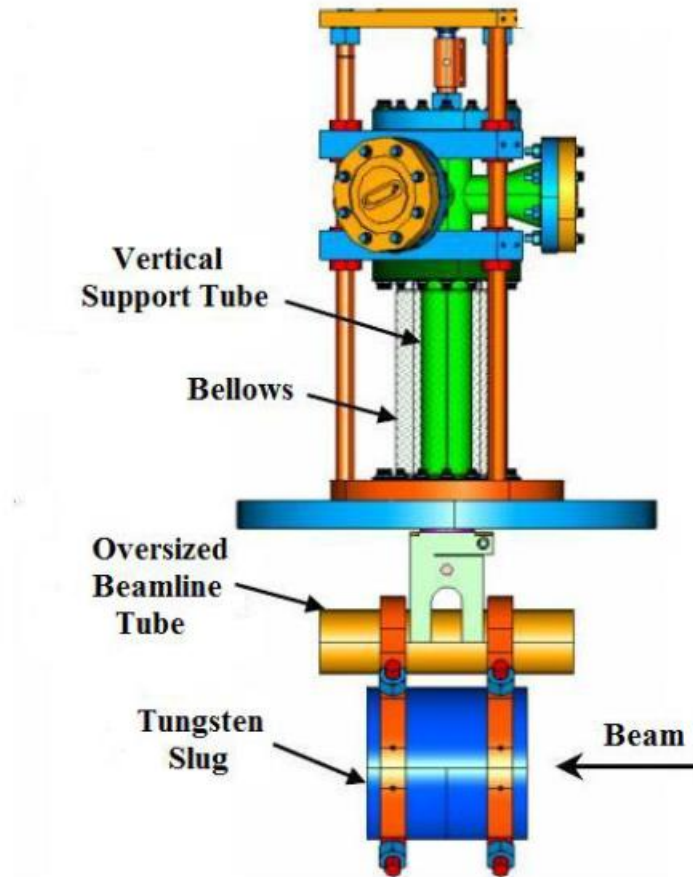
For 20 GeV,
they wanted
72 X_0 long
and
46 X_0 radially.

(0.01% shower
penetration
loss, which
seems overkill)

*D. Young, “A High Precision Faraday Cup and Quantameter for SLAC”, SLAC-PUB-264 (January 1967); and NIM 52 (1967) 1-14.

Tungsten-Copper Calorimeter

Left-over from an old Hall A project.



95% W, 5% Cu

16cm long ($45.7 X_0$)
8cm radius ($22.9 X_0$)

Very thin wrt SLAC FC.

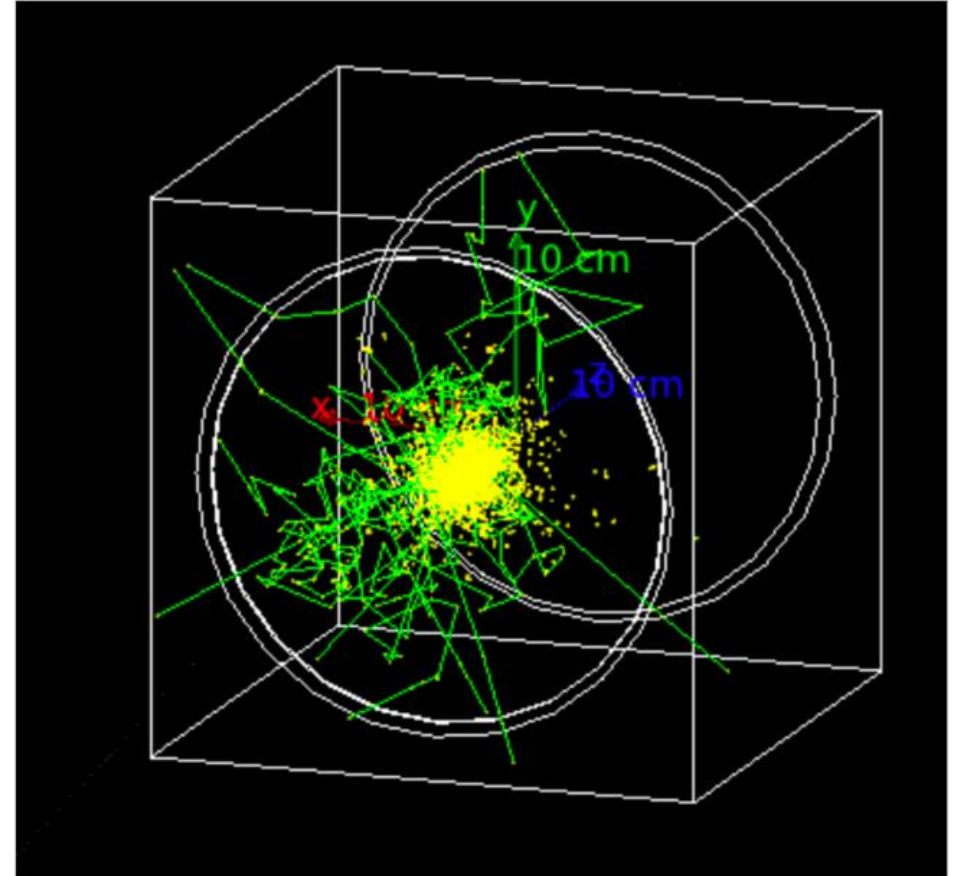
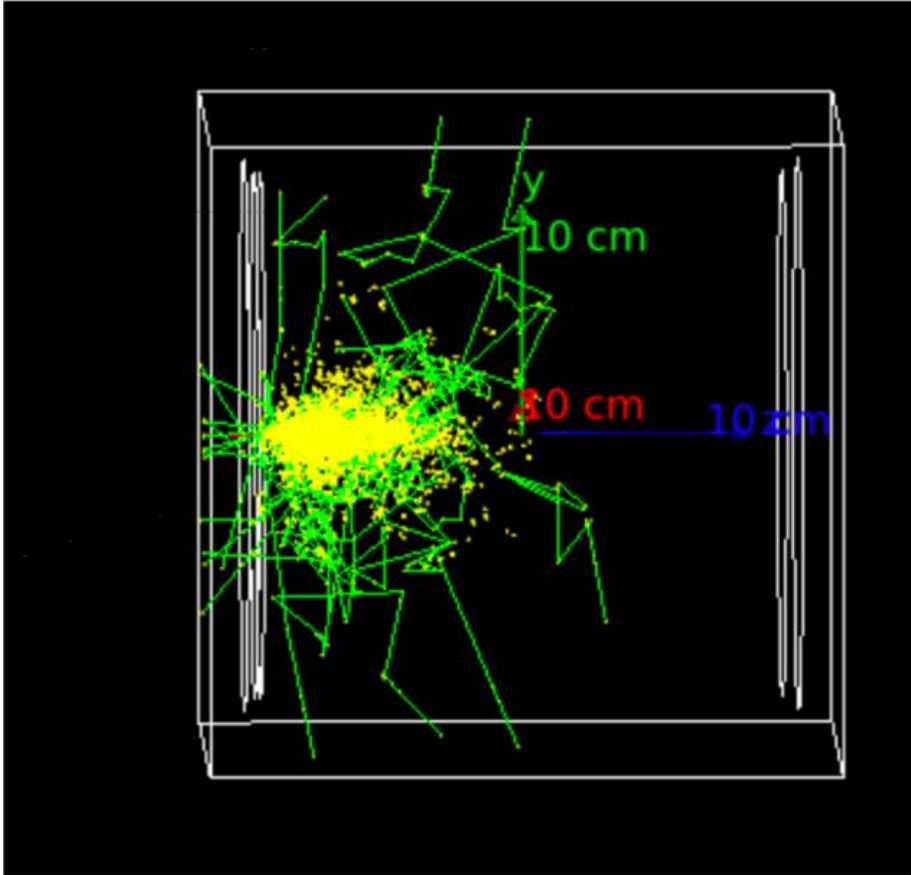
But given the large proposed relative error on A_{zz} , charge losses of 1% would be no problem as long as any drifts are small during a target cycle.

Hole is the wrong size (0.5cm radius, and 2.5cm deep).

Since it is mildly activated, drilling to the right size and aspect ratio could be a pain.

Figure 2: Slug with vertical tube.

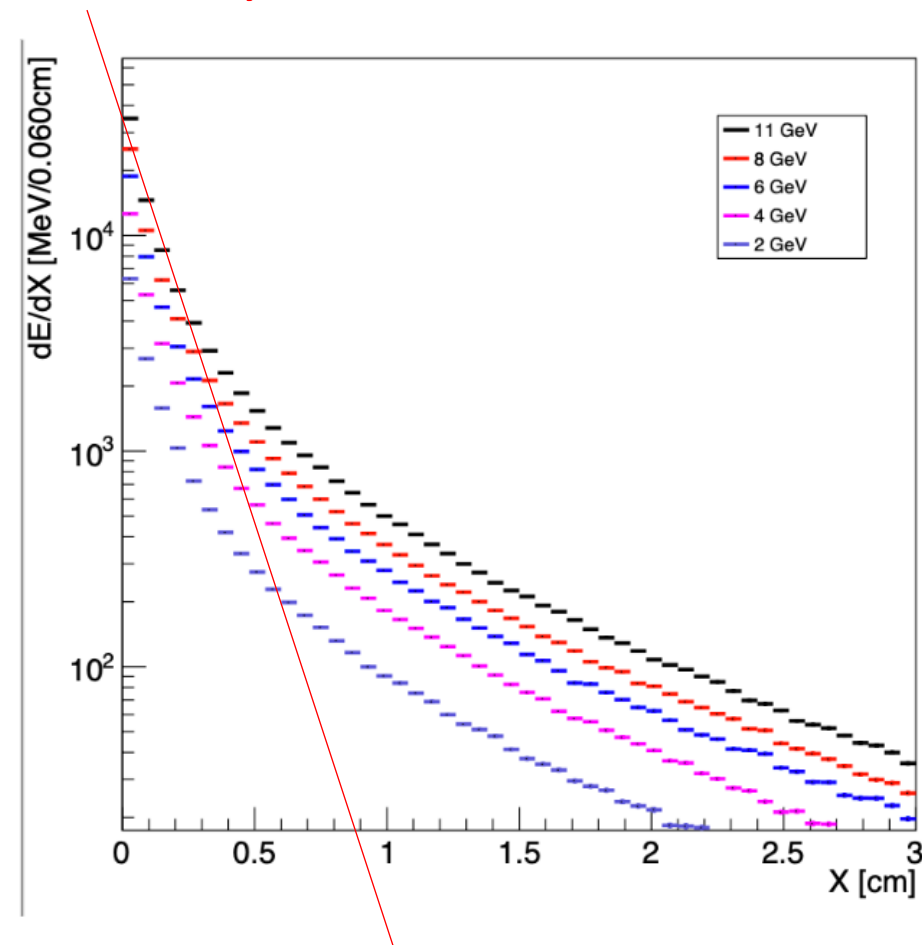
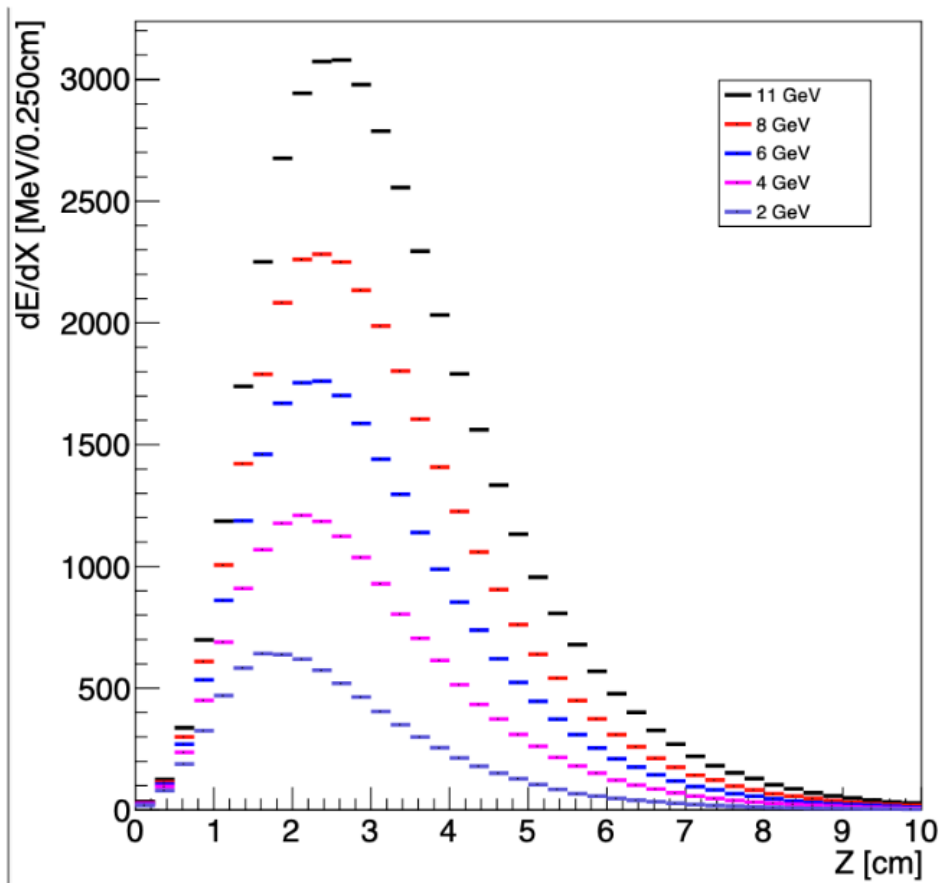
Related G4 Simulations by Nathaly Santiesteban



Qualitatively Noteworthy:

- Longitudinally, the shower peaks at shallow depth, so not much energy escapes out the back. (Cannot use W-Cu at high power though lol!)
- There's some escape from the front, which mostly appears to be gammas.
- Radially, some escape is easily observed, again mostly appearing to be gammas.
 - Gamma rays of $O(1)$ MeV have surprising penetrating power, so Compton scattering losses can matter for absolute charge monitoring.

Related G4 Simulations by Nathaly Santiesteban



Semi-quantitatively now:

- Longitudinally on linear scale, there's not much escape out the back at 10cm ($\sim 30 X_0$).
- Radially on log scale, you can see the narrow (90%) core of energy inside the exponential-ish Moliere radius, but there's a small, long tail beyond that.

Needs more study, but for relative charge monitoring, the W-Cu plug with a larger hole may be large enough.

Why Not Just Hook an Ammeter up to the Existing Beam Dump?

Many of the reasons we just mentioned:

- No backscatter reduction
- Many leakage paths to ground, including the low-conductivity cooling water and ionized air.
- Highly activated area after 30 years.

(Dimensions are in meters.)

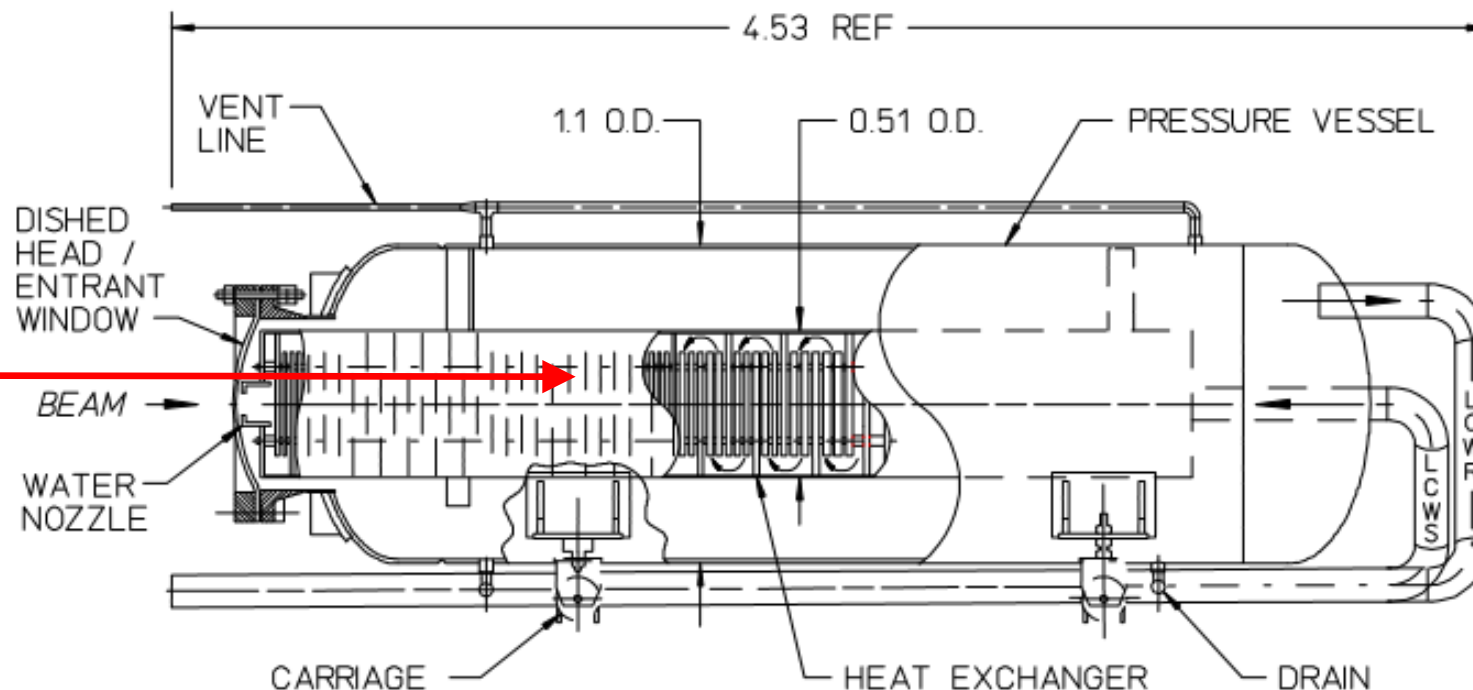


FIGURE 5. Cross sectional view of the end station beam dump.

Section summary on FC

Beam power is less than 1 kWatt. Despite high vacuum and no cooling water, FC can likely passively shed heat thru black body radiation to outer vessel (and BeO supporting insulators?).

Beam power per unit area is crazy low. Probably no issues there, even in Tungsten.

High vacuum (1E-5 Torr) is an important specification.

Can the W-Cu calorimeter be modified to be useful? It is certainly compact. Existing hole is way too small. Might not be practical to modify though since it has been in the beam and is mildly activated.

Monitoring leakage currents at the 10 pA level sounds hard. This is probably at the level of the dark current out of the injector, so we'll need to insert a beam block in the injector. (The 5 MeV FC will do just fine.)

Charge Monitoring Talk Summary

BCMs:

I need to do some tempco estimates and thermal stability tests, then decide what tweaks to make in the BCM system.

This is the devil I know, so the necessary-but-not-necessarily-sufficient improvement path is clear. But we're ultimately limited by 1.497 GHz losses in RF cables with significant tempco's. Given the likely target cycling period of a few hours (?), it would be irrationally optimistic to think this random error can get down to the O(0.01)% level.

FC:

I originally thought of the b1 measurement as a stepping stone to a FC for the positron program. We will learn a lot from our first FC in Hall C, but I can see now that the two will be QUITE different:

- b1 is low-powered at < 1 kWatt, so presumably a FC wouldn't need water-cooling.
The trick will be keep any time-dependent background currents below the O(10) pA level.
I am perhaps being irrationally optimistic, but it sounds fun!
- The positron program at < 22 kWatts will likely need water cooling for its FC.
The trick will be to design a closed cycle system with negligible (< 2nA) time varying leakage current leakage to ground.
This will be a more difficult design and R&D project.

(Comments on the order of magnitude of parity-violating asymmetry corrections to any single target cycle are in the backups. In her draft Jeopardy slides, Elena says we'll be reversing the vector part of the target polarization, so these small contributions will cancel within a pair of target cycles.)

extras

PV Asymmetry Corrections?

Since the measured asymmetries are so small, and the target has vector polarization, do we have to make corrections for PV asymmetries? (I'm assuming the worst case where the target vector polarization never gets flipped.)

It seems that PV corrections are small-ish and could probably be ignored:

X_bj	Q2	Projected dAzz_stat (from Table 2 of the proposal*)	dA_measured = f*Ptgt*dAzz_stat = 0.285*0.2*dAzz_stat	measured PV Asymmetry WAG ~100ppm*Q2*f*Pvector =1e-4*Q2*0.285*0.5 =1.43e-5*Q2	measured PV Asymmetry WAG after correcting for the correct f and the wrong polarization =1e-4*Q2*(0.5/0.2) =2.5e-4*Q2
				Compare to column 4 O(10)%	Compare to column 3 O(20)%
0.16	1.17	1.5e-3	8.6e-5	1.7e-5	2.9e-4
0.28	1.76	3.9e-3	2.2e-4	2.5e-5	4.4e-4
0.36	2.12	5e-3	2.9e-4	3.0e-5	5.3e-4
0.49	3.25	3.7e-3	2.1e-4	4.6e-5	8.1e-4

The size of the PV correction is O(10)% of the statistical error bar on the measured asymmetry. But because the vector asymmetry is larger than the tensor asymmetry, I think this gets enhanced to O(20)% of the Azz error. (This estimate could be off a factor of 2 in either direction!)

Title

What is my error budget for charge monitoring?

1. Expected size of the Azz asymmetry and its error
2. With doubled error bars
3. Expected size of the error on the measured asymmetry
4. Only have 6 days
5. Slow cycles
6. Not-so-slow cycles
7. Random error budget
8. Section summary

Overview of BCMs

1. BCM infrastructure
2. BCM infrastructure 2
3. Making temperature-stable BCMs
4. Model stability for $Q = 500$ vs 1500
5. Two thermally separate BCM enclosures
6. Bcm difference package 1
7. Bcm difference package 2
8. Bcm difference package 1 – package 2
9. Mods for polarized target running
10. Section summary

Overview of FCs

1. The basic idea of a FC
2. Beam power levels 1, 2, 3
3. Power per unit area
4. Classical errors
5. SLAC “high” power FCC
6. JLab Tungsten calorimeter
7. JLab high power beam dump
8. Section summary

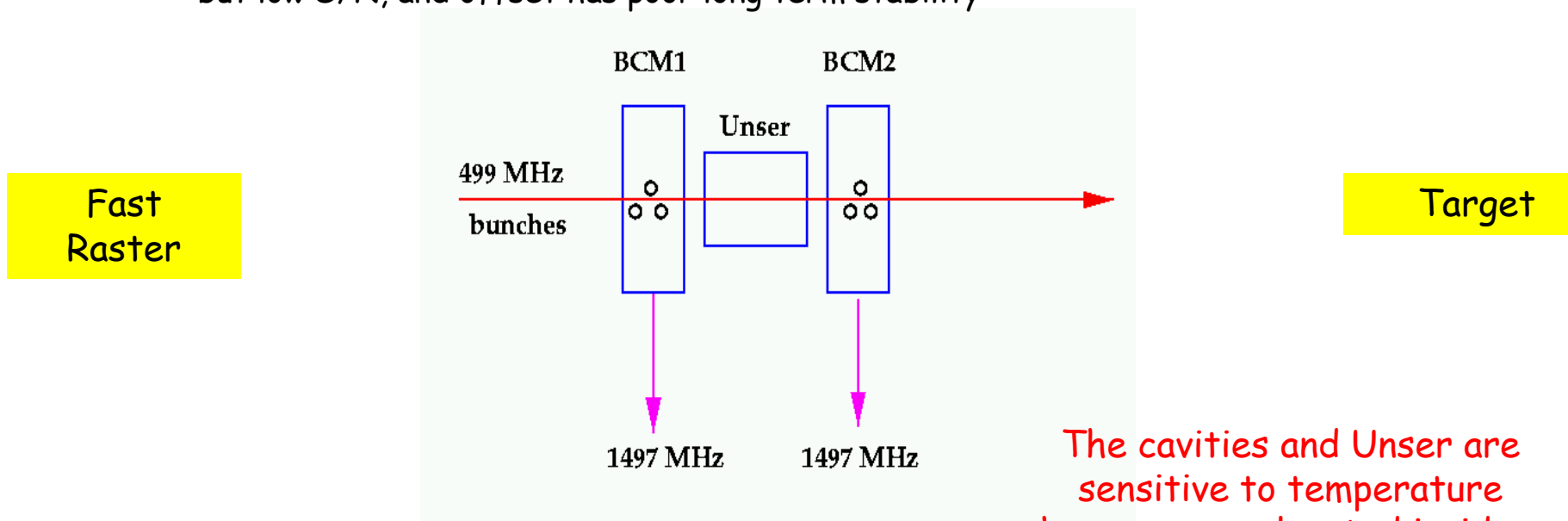
Overall summary

1. PV asymmetry corrections?

Big Picture

Two key types of non-intercepting transducers:

1. Resonant cavity monitors (pill-box, TM_{010} , Q value $\sim 1500??$) -
very high S/N,
but unfeasible to dead-reckon an accurate calibration with RF at 1497 MHz
2. Unser monitor (Parametric Current Transformer) -
extremely stable gain,
but low S/N, and offset has poor long term stability



The cavities and Unser are sensitive to temperature change, so are located inside an insulating box regulated at 110F.

With beam threading both cavities and Unser, an absolute calibration can be transferred from the Unser to the cavities.

The Unser signal is then generally ignored until the next calibration, but is available for backup.

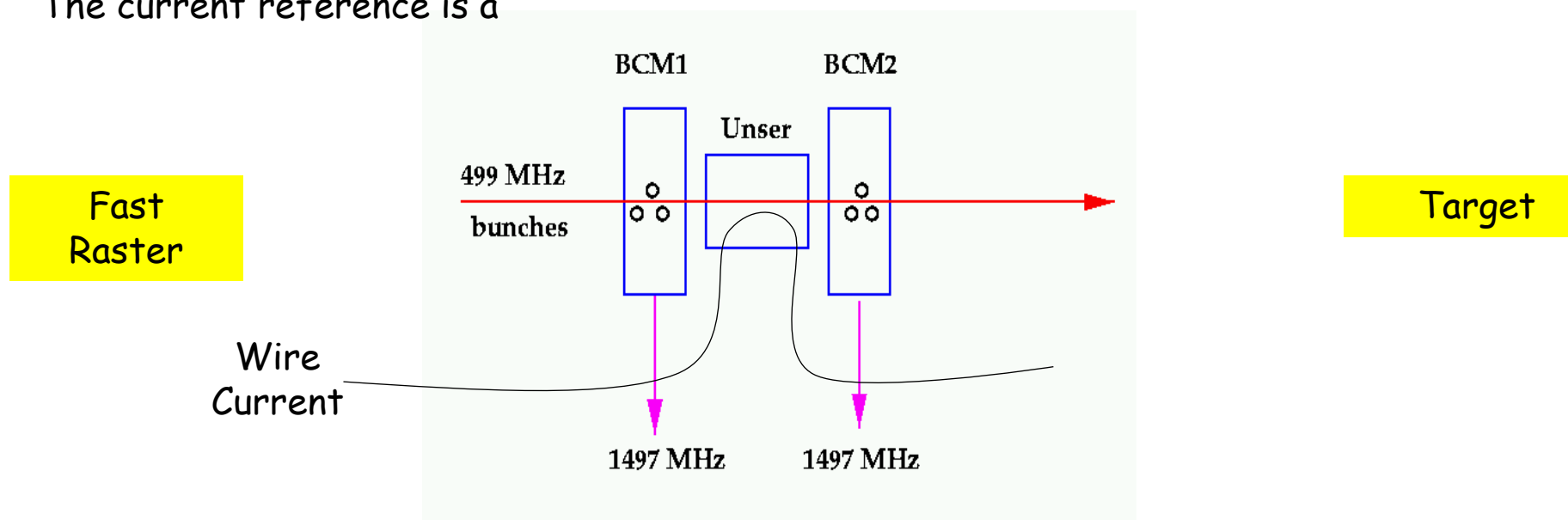
Unser

Determines absolute current for calibration of cavities

The Unser gain is stable and accurate to 0.1% out of the box (4 mV/ μ A).

We check the gain of the Unser+V/F+gated electronics chain by passing an accurate current thru a wire.

The current reference is a



Current Monitors

- The RF cavities monitors have high S/N and can be very linear, but have to be calibrated.
- The Unser monitor has a stable gain which can be verified with wire current calibrations, but it is noisy and the offset tends to drift. It is only used for calibration.
- Same beam current threads all three devices.

