

Non-perturbative characteristics of (QCD) spectral functions at finite temperature

Peter Lowdon

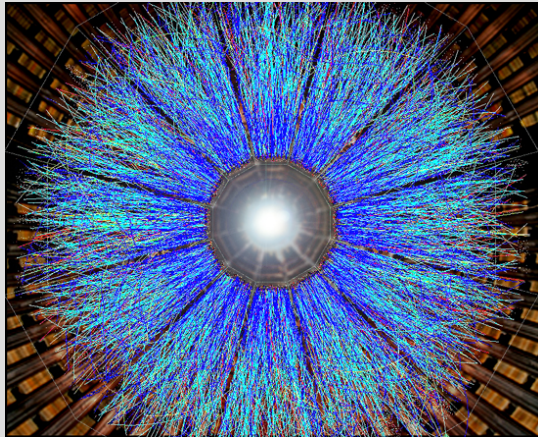
(Goethe University Frankfurt)

Talk outline

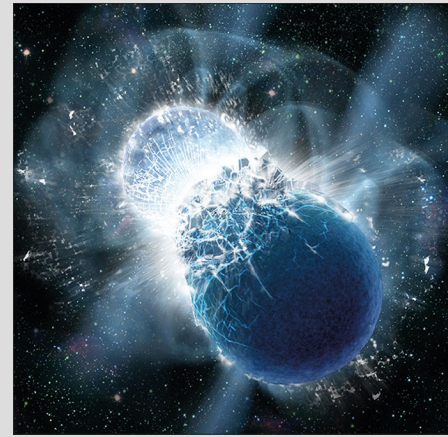
1. QFT beyond the vacuum
2. The importance of spectral functions
3. Causality constraints
4. Spectral properties from Euclidean data
5. Revisiting $T > 0$ perturbation theory

1. QFT beyond the vacuum

- To describe physical phenomena in “extreme environments” one must understand of how QFT applies to systems that are hot, dense, or both



[Brookhaven National Lab]



[Skyworks Digital Inc.]

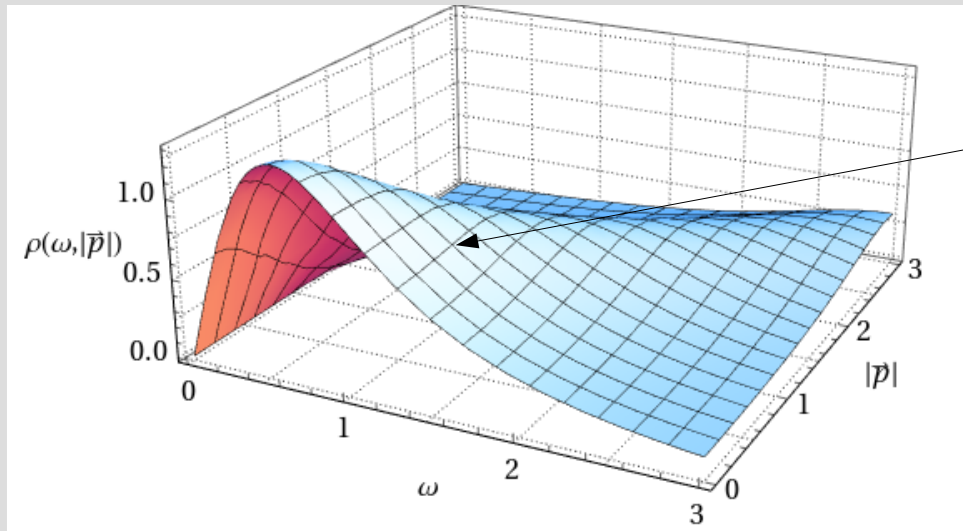
- Correlation functions are the building blocks of *any* QFT \rightarrow they encode the dynamical properties that arise due to changes in temperature $T=1/\beta$ or density. In this talk we will restrict to vanishing density.

$$\langle \Omega_\beta | \phi(x_1) \cdots \phi(x_n) | \Omega_\beta \rangle$$

2. The importance of spectral functions

- At finite T spectral functions $\rho(\omega, \vec{p})$ play a particularly important role

$$\rho(\omega, \vec{p}) = \int d^4x e^{i(\omega x_0 - \vec{p} \cdot \vec{x})} \langle \Omega_\beta | [\phi(x), \phi(0)] | \Omega_\beta \rangle$$



Peak locations and their dispersion are related to the dynamics of the medium and the underlying degrees of freedom of the theory

- Spectral functions also enter into the calculation of numerous important observables (transport coefficients, particle production rates, etc.)

Important question: *can general spectral function characteristics be disentangled from model-dependent effects?*

3. Causality constraints

- Finite temperature QFT is *significantly* less well-understood than in vacuum, but important progress has been made for spectral functions
- For scalar fields, the fields being **local**, i.e. $[\phi(x), \phi(y)] = 0$ for $(x-y)^2 < 0$, implies the following general representation*

$$\rho(\omega, \vec{p}) = \int_0^\infty ds \int \frac{d^3 \vec{u}}{(2\pi)^2} \epsilon(\omega) \delta(\omega^2 - (\vec{p} - \vec{u})^2 - s) \tilde{D}_\beta(\vec{u}, s)$$

“Thermal spectral density”

- This is the $T > 0$ generalisation of the well-known *Källén-Lehmann* spectral representation

$$\rho(\omega, \vec{p}) \xrightarrow{\beta \rightarrow \infty} 2\pi \epsilon(\omega) \int_0^\infty ds \delta(p^2 - s) \rho(s)$$

e.g. $\rho(s) = \delta(s - m^2)$ for a massive free theory

- Determining the properties of $\tilde{D}_\beta(\mathbf{u}, s)$ is clearly key to understanding how in-medium effects manifest themselves in $\rho(\omega, \mathbf{p})$

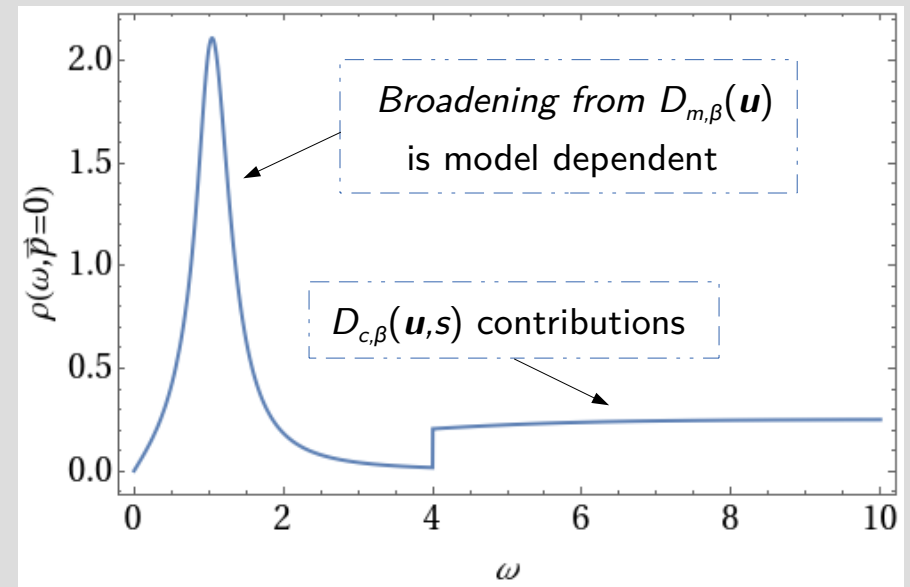
* See: J. Bros and D Buchholz, Z. Phys. C 55 (1992), Ann. Inst. H.Poincaré Phys.Theor. 64 (1996)

3. Causality constraints

- Proposition: the medium contains “Thermoparticles”: particle-like constituents which differ from collective quasi-particle excitations, and show up as **discrete** contributions [Bros, Buchholz, *NPB* 627 (2002)]

$$\tilde{D}_\beta(\vec{u}, s) = \tilde{D}_{m,\beta}(\vec{u}) \delta(s - m^2) + \tilde{D}_{c,\beta}(\vec{u}, s)$$

- Thermoparticle components reduce to those of a vacuum particle state with mass m in the limit $T \rightarrow 0$
- Non-trivial “Damping factor” $\tilde{D}_\beta(\mathbf{u})$ results in thermally-broadened peaks in the spectral function, i.e. parametrises the effects of collisional broadening
- Component $\tilde{D}_{c,\beta}(\mathbf{u}, s)$ contains all other types of excitations, including those that are *continuous* in s



4. Spectral properties from Euclidean data

- In many instances *Euclidean* data is used to calculate $T > 0$ observables, e.g. spectral functions $\rho_\Gamma(\omega, \mathbf{p})$ from $C_\Gamma(\tau, \vec{x}) = \langle O_\Gamma(\tau, \vec{x}) O_\Gamma(0, \vec{0}) \rangle_T$ where O_Γ is some particle-creating operator

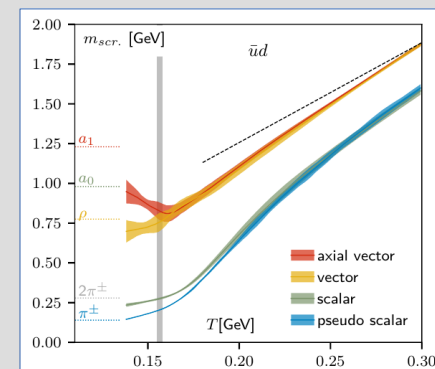
$$\tilde{C}_\Gamma(\tau, \vec{p}) = \int_0^\infty \frac{d\omega}{2\pi} \frac{\cosh \left[\left(\frac{\beta}{2} - |\tau| \right) \omega \right]}{\sinh \left(\frac{\beta}{2} \omega \right)} \rho_\Gamma(\omega, \vec{p})$$

→ Determine $\rho_\Gamma(\omega, \mathbf{p})$ given $\tilde{C}_\Gamma(\tau, \mathbf{p})$: *problem is ill-conditioned, need more information!*

- Another quantity of interest in lattice studies is the *spatial* correlator

$$C_\Gamma(x_3) = \int_{-\infty}^{\infty} dx_1 \int_{-\infty}^{\infty} dx_2 \int_{-\frac{\beta}{2}}^{\frac{\beta}{2}} d\tau C_\Gamma(\tau, \vec{x}) = \int_{-\infty}^{\infty} \frac{dp_3}{2\pi} e^{ip_3 x_3} \int_0^\infty \frac{d\omega}{\pi\omega} \rho_\Gamma(\omega, p_1 = p_2 = 0, p_3)$$

- Large- x_3 behaviour $C_\Gamma(x_3) \sim \exp(-m_{scr}|x_3|)$ used to extract “screening masses” $m_{scr}(T)$



[HotQCD collaboration, PRD 100 (2019)]

4. Spectral properties from Euclidean data

Goal: Use the additional constraints imposed by causality to better understand how spectral features manifest themselves in Euclidean data

- Causality implies a general connection between the spatial correlator and thermal spectral density [P.L., *PRD* 106 (2022); P.L., O. Philipsen, *JHEP* 10, 161 (2022)]

$$C(x_3) = \frac{1}{2} \int_0^\infty ds \int_{|x_3|}^\infty dR e^{-R\sqrt{s}} D_\beta(R, s)$$

Thermal spectral density
in position space

- Thermoparticle states give rise to $C(x_3)$ contributions that are particularly significant in the large- x_3 region

$$C(x_3) \approx \frac{1}{2} \sum_{i=1}^n \int_{|x_3|}^\infty dR e^{-m_i R} D_{m_i, \beta}(R)$$

- Once the damping factors of these states are known one can use the $T > 0$ spectral representation to compute their analytic contribution to $\rho(\omega, \mathbf{p})$

4. Spectral properties from Euclidean data

- Can now apply these relations to QCD lattice data [P.L., O. Philipsen, 2022]
 - Use data [Rohrhofer et al. *PRD* 100 (2019)] for spatial correlator $C_{PS}(x_3)$ of light-quark pseudo-scalar meson operator $\mathcal{O}_{PS}^a = \bar{\psi}\gamma_5\frac{\tau^a}{2}\psi$

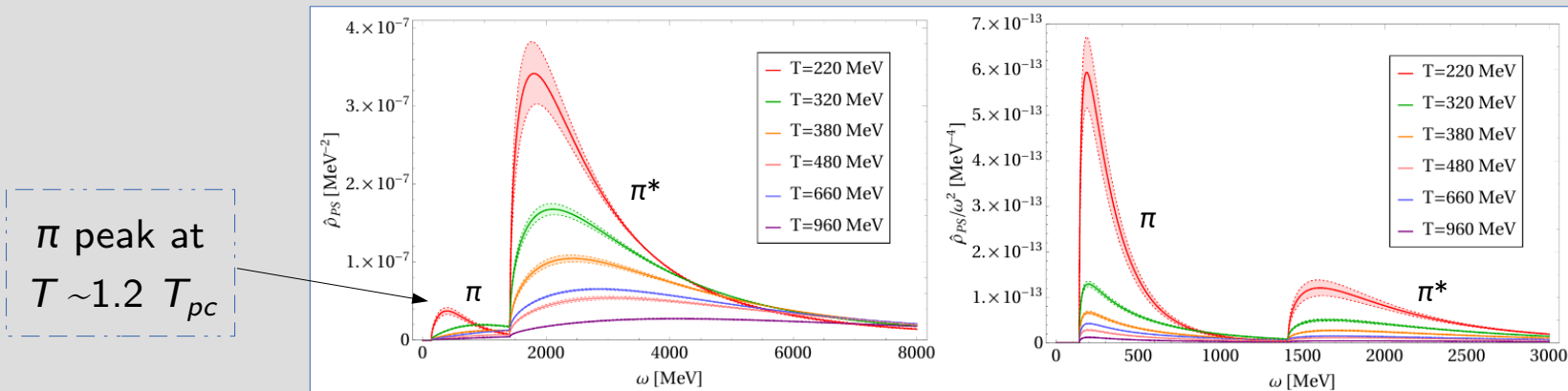
Step 1: Perform fits to $C_{PS}(x_3)$ data to obtain the functional dependence at different temperatures ($T = 220-960$ MeV) → $c_1 \exp(-m_\pi x_3) + c_2 \exp(-m_{\pi^*} x_3)$ describes data well

Contribution of 2 lowest-lying states, π and π^*

Step 2: If π and π^* are thermoparticle-type states for $T > 0$, then:

→ Fit ansatz implies $D_{m_i,\beta}(\vec{x}) = \alpha_i e^{-\gamma_i|\vec{x}|}$ with screening masses $m_i(T) = m_i(T=0) + \gamma_i(T)$, $i = \pi, \pi^*$

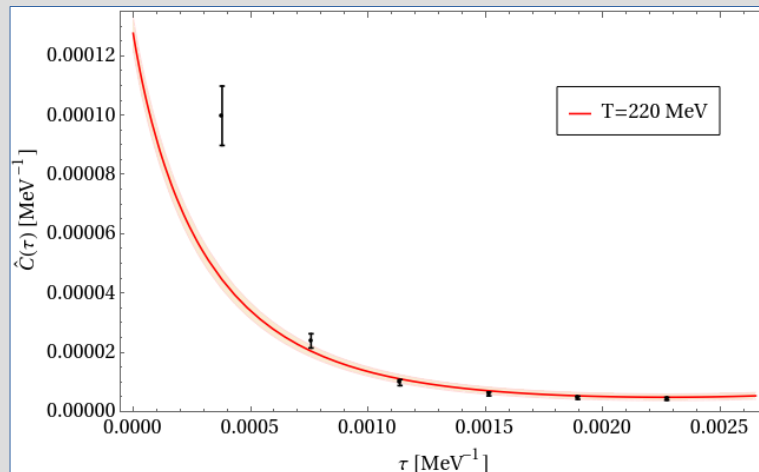
Step 3: Using $D_{m,\beta}(\mathbf{x})$ and spectral representation one can compute $\rho_{PS}(\omega, \mathbf{p})$ contributions:



4. Spectral properties from Euclidean data

Non-trivial test: Since procedure gives the full analytic structure of $\rho_{\text{PS}}(\omega, \mathbf{p})$ due to thermoparticle contributions, one can use this to **predict** the form of the corresponding temporal correlator $\tilde{C}_{\text{PS}}(\tau, \mathbf{p})$

- The spatial and temporal correlators have very different $\rho_{\text{PS}}(\omega, \mathbf{p})$ dependencies \rightarrow a *highly non-trivial check!*
- Using the $T = 220 \text{ MeV}$ $\mathbf{p} = 0$ temporal data from [Rohrhofer et al. *PLB* 802 (2020)] one obtains:

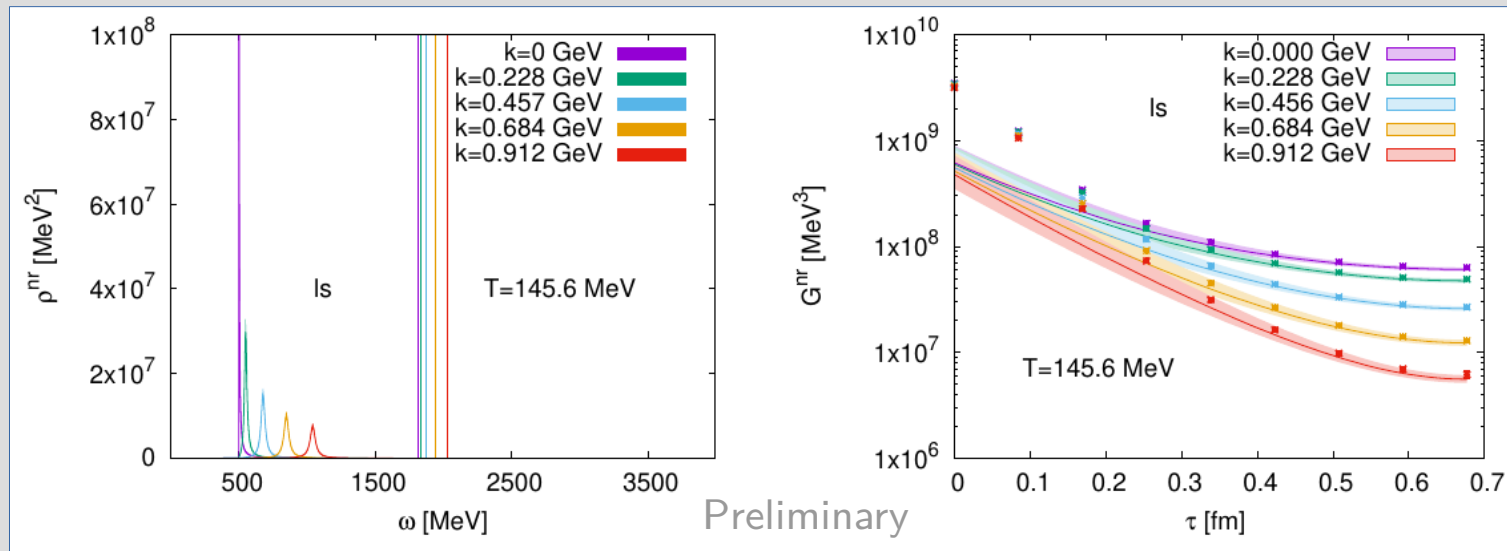


Prediction matches the data well for large τ , and then begins to undershoot \rightarrow Missing contributions from higher excited states

- No matter the procedure, comparing temporal and spatial correlator predictions is an important test for *any* extracted spectral function

4. Spectral properties from Euclidean data

- Work is ongoing [D. Bala, O. Kaczmarek, P. Lowdon, O. Philipsen, and T. Ueding] to apply this approach to pseudo-scalar mesons involving heavier quarks (light-strange and strange-strange)



- The temporal correlator predictions are now *also* compared for $p > 0$
→ Consistent predictions are obtained in both light-strange and strange-strange channels!
- This approach is straight-forwardly generalisable to higher spin states (work in progress...)

5. Revisiting $T > 0$ perturbation theory

- It has long been understood that finite-temperature perturbation theory has complications: non-analytic contributions, IR divergences, ...
- In fact, more specifically, Weldon [*PRD* 65 (2002)] showed that the perturbative procedure in Φ^4 theory *fails* at 2-loop order because the self-energy $\Pi(k)$ has a branch point on the perturbative mass shell $k_0 = E(k)$

→ This is a generic feature of perturbative computations that use free thermal propagators, or in fact any propagators that have a real dispersion relation $p_0 = E(p)$

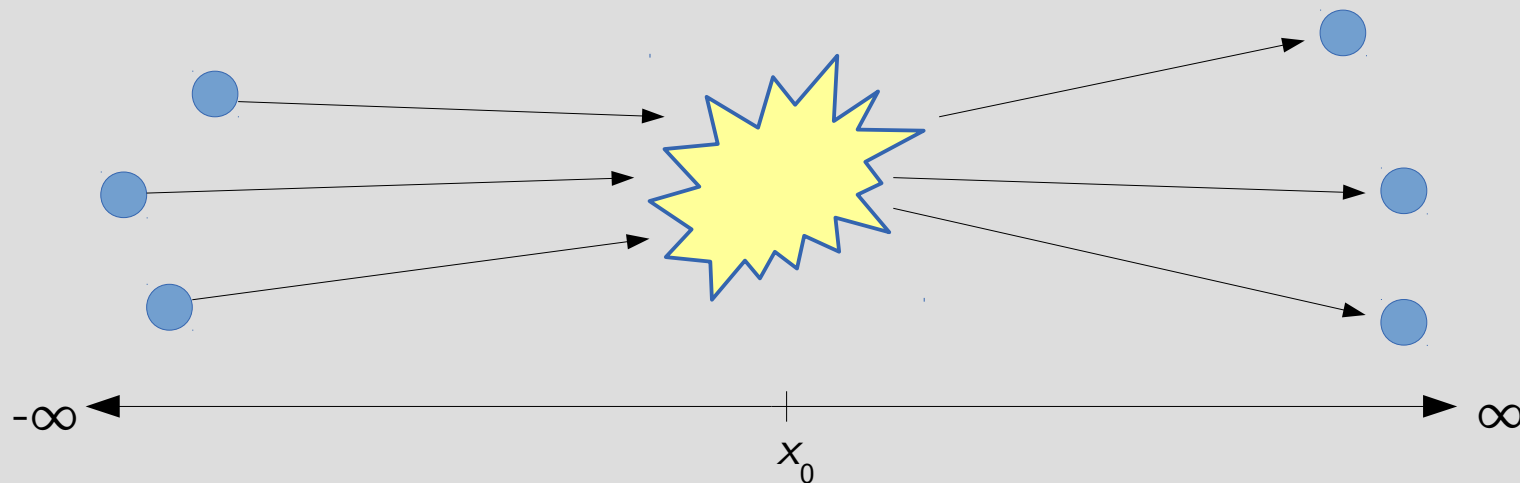
$$D_R(P) = \frac{1}{(p_0 + i\epsilon)^2 - E^2(p)}$$

- Physically, this arises due to the incompatibility of the KMS condition with on-shell states and non-zero interactions [Landsman, *Ann. Phys.* 186, 141 (1988)] (*Narnhofer-Requardt-Thirring Theorem* [Commun. Math. Phys. 92, 247 (1983)])

5. Revisiting $T > 0$ perturbation theory

Idea: Start with propagators that are off shell [Weldon, 2002]

- The logic is that interactions with the thermal medium persist, even for large times $x_0 \rightarrow$ need to take into account in the definition of scattering states



\rightarrow But how does one decide what form these propagators should take?

- With decomposition $\tilde{D}_\beta(\vec{u}, s) = \tilde{D}_{m,\beta}(\vec{u}) \delta(s - m^2) + \tilde{D}_{c,\beta}(\vec{u}, s)$ one can prove that the thermoparticle component *dominates* the two-point function $\langle \Omega_\beta | \phi(x) \phi(0) | \Omega_\beta \rangle$ at large x_0 [Bros, Buchholz, *NPB* 627 (2002)]

\rightarrow *Thermoparticles are a natural asymptotic thermal state candidate!*

5. Revisiting $T > 0$ perturbation theory

Idea: thermal scattering states are defined by imposing an asymptotic field condition [Bros, Buchholz, *NPB* 627 (2002)]:

Asymptotic fields Φ_0 are assumed to satisfy dynamical equations, but only at large x_0

In Φ^4 theory

$$(\partial^2 + m^2)\phi_0(x) + \frac{\lambda}{3!}\phi_0^3(x) \xrightarrow{|x_0| \rightarrow \infty} 0$$

“Asymptotic mass”

- The thermoparticle damping factor $\tilde{D}_{m,\beta}(\mathbf{u})$ is **uniquely fixed** by the asymptotic field equation
 - This means that the non-perturbative effects experienced by thermoparticle states are controlled by the asymptotic dynamics
- Given $D_{m,\beta}(\mathbf{u})$ one can simply combine this together with the spectral representation to compute the explicit form of the thermoparticle propagator or spectral function

5. Revisiting $T > 0$ perturbation theory

- Can then start to perform perturbative calculations with *this* propagator instead of a free field propagator → suggested that this could give rise to an IR-regularised perturbative expansion for $T > 0$ [Bros, Buchholz hep-th/9511022]

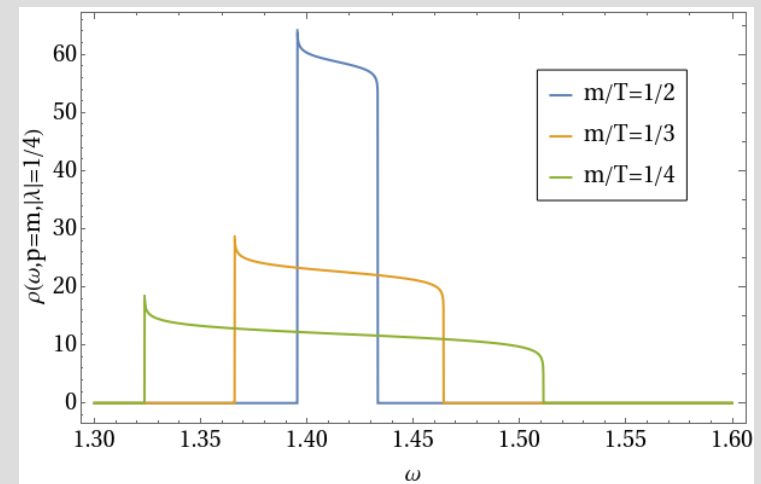
Example: Φ^4 theory [PL, O. Philipsen, *in preparation*]

→ Thermoparticle propagator:

(Width parameter $\kappa \sim \sqrt{|\lambda|T}$)

$$\tilde{G}_\beta^{(-)}(k_0, \vec{p}) = \frac{1}{4|\vec{p}|\kappa} \ln \left[\frac{-k_0^2 + m^2 + (|\vec{p}| + \kappa)^2}{-k_0^2 + m^2 + (|\vec{p}| - \kappa)^2} \right]$$

- Spectral function already has a width at 1-loop order (and is renormalisable)
- At 2-loop the thermoparticle peak plays a dominant role at low energies



Summary & outlook

- Causality imposes **non-perturbative** constraints for $T > 0$ which have significant implications
 - Spectral properties of thermal correlation functions
 - Connection between real-time observables and Euclidean correlators
 - New insights into the characteristics of perturbation theory
- So far, only real scalar fields ϕ with $T > 0$ have been considered, but this approach *can* be extended
 - Other hadronic states (baryons, exotic states, ...)
 - Higher spin fields/states (fermions, vectors, ...) *Work in progress!*
 - Non-vanishing density, $|\mu| > 0$
- Ultimately, these constraints and methods can help in gaining a better understanding of the phase structure of QCD

Backup: *Euclidean spectral relations*

- One can use the assumptions of local QFT at finite T to put constraints on the the structure of Euclidean correlation functions

→ From the KMS condition and locality:

$$\mathcal{W}_E(\tau, \vec{x}) = \frac{1}{\beta} \sum_{N=-\infty}^{\infty} w_N(\vec{x}) e^{\frac{2\pi i N}{\beta} \tau}$$

- The Fourier coefficients of the Euclidean two-point function are then related to the thermal spectral density as follows [P.L., *PRD* 106 (2022)]:

$$w_N(\vec{x}) = \frac{1}{4\pi|\vec{x}|} \int_0^{\infty} ds e^{-|\vec{x}|\sqrt{s+\omega_N^2}} D_\beta(\vec{x}, s)$$

- With the Bros-Buchholz decomposition this becomes

$$w_N(\vec{x}) = \frac{1}{4\pi|\vec{x}|} \left[D_m(\vec{x}) e^{-|\vec{x}|\sqrt{m^2+\omega_N^2}} + \int_0^{\infty} ds e^{-|\vec{x}|\sqrt{s+\omega_N^2}} D_c(\vec{x}, s) \right]$$

→ *The continuous component $D_c(\mathbf{x}, s)$ is increasingly suppressed for large $|\mathbf{x}|$*

Backup: *Damping factors from asymptotic dynamics*

- Applying the asymptotic field condition for Φ^4 theory, the resulting damping factors have the form [Bros, Buchholz, 2002]:

$$\rightarrow \text{For } \lambda < 0: \quad D_{m,\beta}(\vec{x}) = \frac{\sin(\kappa|\vec{x}|)}{\kappa|\vec{x}|} \quad \rightarrow \text{For } \lambda > 0: \quad D_{m,\beta}(\vec{x}) = \frac{e^{-\kappa|\vec{x}|}}{\kappa_0|\vec{x}|}$$

where κ is defined with $r = m/T$:

$$\kappa = T\sqrt{|\lambda|}K(r), \quad K(r) = \sqrt{\int \frac{d^3\hat{q}}{(2\pi)^3 2\sqrt{|\hat{q}|^2 + r^2}} \frac{1}{e^{\sqrt{|\hat{q}|^2 + r^2}} - 1}}$$

\rightarrow The parameter κ has the interpretation of a thermal width: $\kappa \rightarrow 0$ for $T \rightarrow 0$, or equivalently κ^{-1} is mean-free path

- Now that one has the exact dependence of $D_{m,\beta}(\mathbf{x})$ on the external physical parameters, in this case T , m and λ , one can use this to calculate observables *analytically*

Backup: Analytic shear viscosity computation

- Of particular interest is the *shear viscosity* η , which measures the resistance of a medium to sheared flow

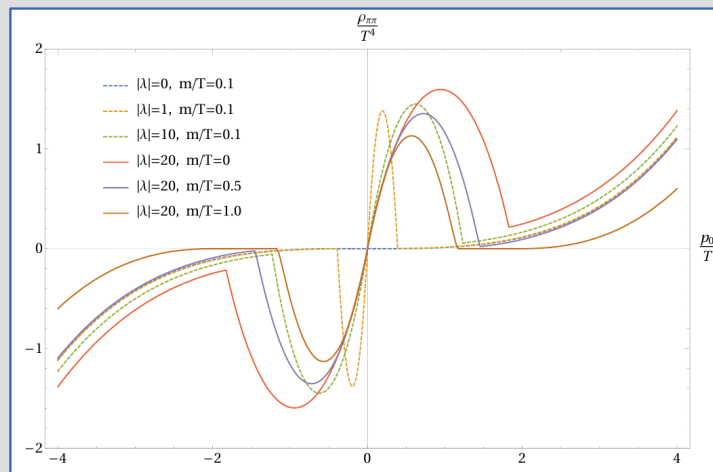
→ This quantity can be determined from the spectral function of the spatial traceless energy-momentum tensor

$$\rho_{\pi\pi}(p_0) = \lim_{\vec{p} \rightarrow 0} \mathcal{F}[\langle \Omega_\beta | [\pi^{ij}(x), \pi_{ij}(y)] | \Omega_\beta \rangle](p)$$

... and η is recovered via the *Kubo relation*

$$\eta = \frac{1}{20} \lim_{p_0 \rightarrow 0} \frac{d\rho_{\pi\pi}}{dp_0}$$

- Using $D_{m,\beta}(\mathbf{x})$ for $\lambda < 0$, the EMT spectral function $\rho_{\pi\pi}$ has the form:

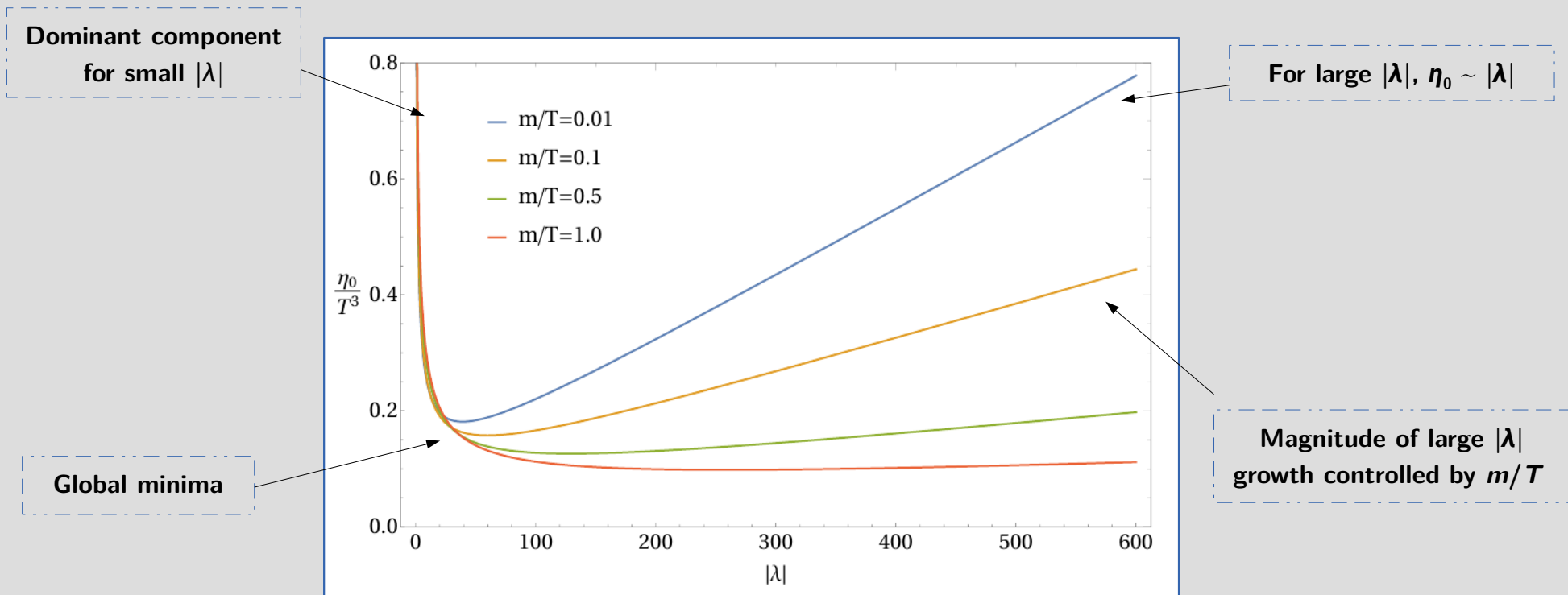


- The presence of interactions causes resonant peaks to appear → peaked when $p_0 \sim \kappa = 1/\ell$
- For $\lambda \rightarrow 0$ the free-field result is recovered, as expected
- The dimensionless ratio m/T controls the magnitude of the peaks

Backup: Analytic shear viscosity computation

- Applying Kubo's relation, the shear viscosity η_0 arising from the asymptotic states can be written [P.L., R.-A. Tripolt, J. M. Pawłowski, D. H. Rischke, *PRD* 104 (2021)]

$$\eta_0 = \frac{T^3}{15\pi} \left[\frac{\mathcal{K}_3\left(\frac{m}{T}, 0, \infty\right)}{\sqrt{|\lambda|}} + \sqrt{|\lambda|} \mathcal{K}_1\left(\frac{m}{T}, 0, \infty\right) + \frac{\mathcal{K}_4\left(\frac{m}{T}, \sqrt{|\lambda|}K\left(\frac{m}{T}\right), \sqrt{|\lambda|}K\left(\frac{m}{T}\right)\right)}{4|\lambda|} \right]$$



→ For fixed coupling, η_0/T^3 is entirely controlled by functions of m/T

Backup: *Shear viscosity from FRG data*

- Locality constraints imply that particle damping factors $D_{m,\beta}(\mathbf{x})$ can also be calculated from Euclidean momentum space data [P.L., *PRD* 106 (2022)]

$$D_{m,\beta}(\vec{x}) \sim e^{|\vec{x}|m} \int_0^\infty \frac{d|\vec{p}|}{2\pi} 4|\vec{p}| \sin(|\vec{p}||\vec{x}|) \tilde{G}_\beta(0, |\vec{p}|).$$

p-space Euclidean propagator

Holds for large separation $|\mathbf{x}|$

- In [P.L., R.-A. Tripolt, *PRD* 106 (2022)] pion propagator data from the quark-meson model (FRG calculation) was used to compute the damping factor at different values of T via the analytic relation above
- Fits to the resulting data were consistent with the form: $D_{m_\pi,\beta}(\vec{x}) = \alpha_\pi e^{-\gamma_\pi|\vec{x}|}$
- $D_{m,\beta}(\mathbf{x})$ can then be used as input for calculations, e.g. shear viscosity

Similar qualitative features to results from chiral perturbation theory

