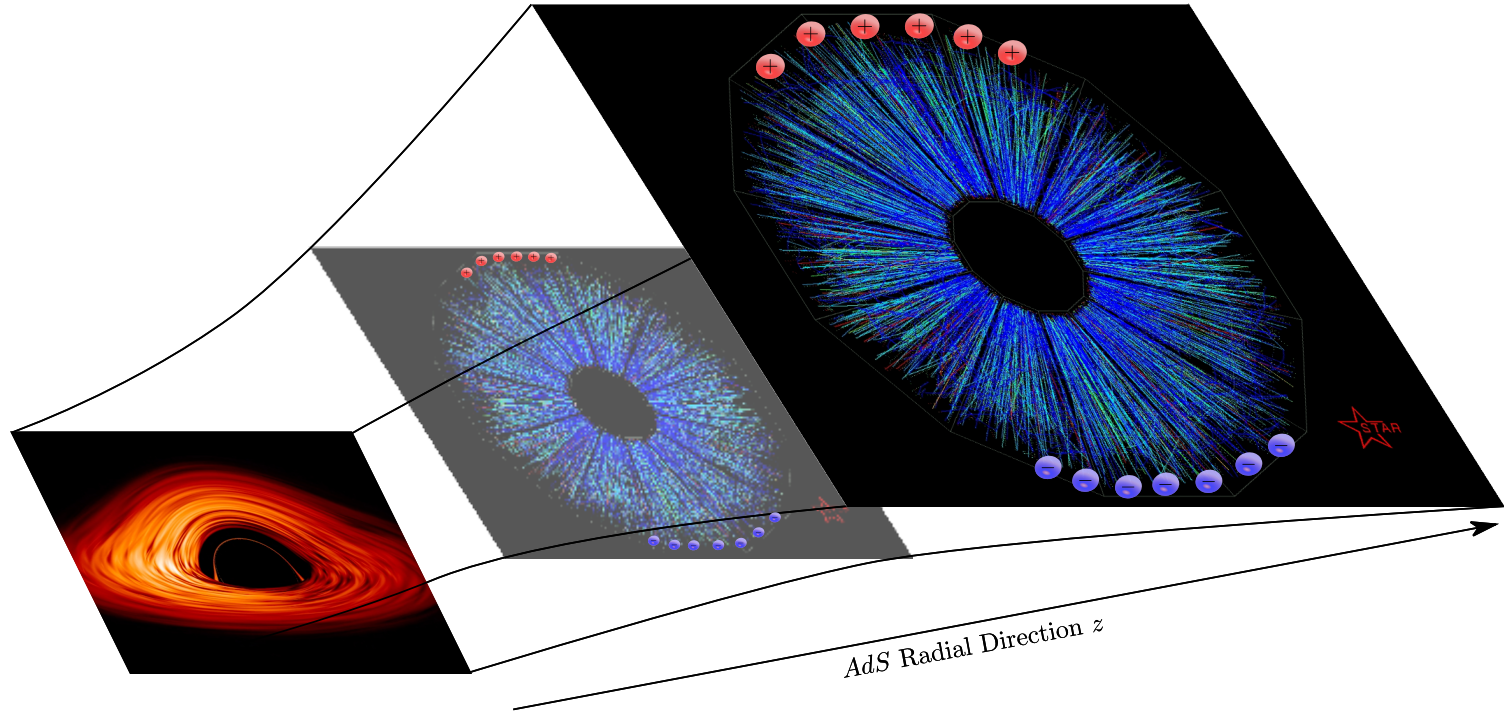


An example of convergence of hydrodynamics in strong magnetic fields



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4+1 D asymptotically locally AdS Kerr (Myers Perry) black hole in Boyer-Lindquist coordinates

$$ds^2 = \frac{(l^2 r^2 + 1)}{\rho^2 r^2} \left(-\frac{bd\phi (a^2 + r^2) \sin^2(\theta)}{\Xi_a} - \frac{ad\psi (b^2 + r^2) \cos^2(\theta)}{\Xi_b} + abdt \right)^2$$

$$+ \frac{\Delta_\theta \sin^2(\theta)}{\rho^2} \left(adt - \frac{d\phi (a^2 + r^2)}{\Xi_a} \right)^2 - \frac{\Delta_r}{\rho^2} \left(-\frac{d\phi (a \sin^2(\theta))}{\Xi_a} - \frac{bd\psi \cos^2(\theta)}{\Xi_b} + dt \right)^2$$

$$+ \frac{\Delta_\theta \cos^2(\theta)}{\rho^2} \left(bdt - \frac{d\psi (b^2 + r^2)}{\Xi_b} \right)^2 + \frac{d\theta^2 \rho^2}{\Delta_\theta} + \frac{dr^2 \rho^2}{\Delta_r}$$

$$\Delta_r = \frac{1}{r^2} (r^2 + a^2)(r^2 + b^2)(r^2 l^2 + 1) - 2M$$

$$\Delta_\theta = 1 - a^2 l^2 \cos^2(\theta) - b^2 l^2 \sin^2(\theta)$$

$$\rho^2 = r^2 + a^2 \cos^2(\theta) + b^2 \sin^2(\theta)$$

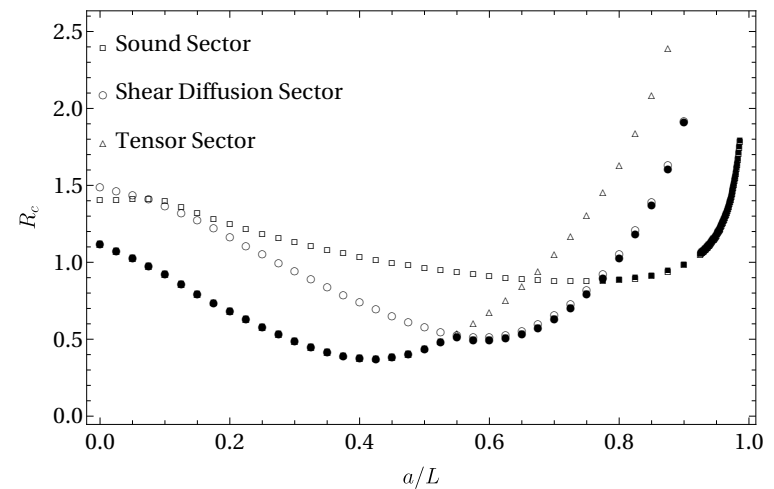
$$\Xi_a = 1 - a^2 l^2$$

$$\Xi_b = 1 - b^2 l^2$$

Magic ↓

$$ds^2 = \frac{2\mu \left(\frac{a\sigma_3}{2} + dt \right)^2}{r^2} + \frac{dr^2}{G(r)} - dt^2 \left(\frac{r^2}{l^2} + 1 \right) + \frac{1}{4} r^2 (\sigma_1^2 + \sigma_2^2 + \sigma_3^2),$$

Magic* →



[Cartwright, Garbiso-Amano, Noronha, Speranza, PRL in review]

Chaos and pole-skipping in a simply spinning plasma

[A. G. Amano, Blake, Cartwright, Kaminski, Thompson, 2023]

OTOC

$$D_{\pm}(t, \Psi) \sim 1 - cG_N \frac{\exp\left(2\pi T_{\pm} \left(t - \frac{L\Psi}{2v_{\pm}^*}\right)\right)}{4\pi|\Psi - \Omega t|}, \quad \pm\Psi \mp \Omega t > 0,$$

$$2\pi T_{\pm} = \frac{2r_0}{L^2 \sqrt{1 - a^2/L^2}} \left(1 \mp \sqrt{\frac{3}{2}} \frac{a}{L}\right),$$

$$\frac{2\pi T_{\pm}}{v_{\pm}^*} = \frac{2r_0}{L^2 \sqrt{1 - a^2/L^2}} \left(\sqrt{\frac{3}{2}} \mp \frac{a}{L}\right).$$

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Hydrodynamics – A universal, effective field theory, of the near equilibrium collective behavior of many body (quantum) systems described by the evolution of (non)-conserved quantities

What if the underlying microscopic theory is weakly coupled to U(1) gauge fields?

Coupling between thermal and electromagnetic d.o.f's is negligible – Maxwell's equations in matter

$$\begin{aligned}\nabla \cdot D &= \rho_f & \nabla \cdot B &= 0 \\ \nabla \times H &= J_f + \partial_t D & \nabla \times E &= \partial_t B\end{aligned}$$

Coupling between thermal and electromagnetic d.o.f's is non-negligible – Hydrodynamics coupled to Maxwell in matter

If one treats, the fields as *external*, then relevant hydrodynamical variables are u^μ, T, μ

$$\begin{aligned}\nabla_\mu T^{\mu\nu} &= F^{\nu\lambda} J_\lambda \\ \nabla_\mu J^\mu &= 0\end{aligned}$$

If one treats, the fields as *dynamical*, then relevant hydrodynamical variables are $u^\mu, T, \mu, E^\mu, B^\mu$

$$\begin{aligned}\nabla_\mu T^{\mu\nu} &= F^{\nu\lambda} J_\lambda^{ext} \\ \nabla_\nu (F^{\mu\nu} - M^{\mu\nu}) &= J_{ext}^\mu + J_{free}^\mu \\ \epsilon^{\mu\nu\alpha\beta} \nabla_\nu F_{\alpha\beta} &= 0\end{aligned}$$

Magnetohydrodynamics (MHD) – The fields are dynamical, the coupling between thermal and electromagnetic fields is non-negligible and the matter is electrically conducting while the electric field is screened

$$E \sim O(\partial) \quad B \sim O(1)$$

$$B/T^2 \ll 1$$

Convergence of series and plane curves – Simple Series

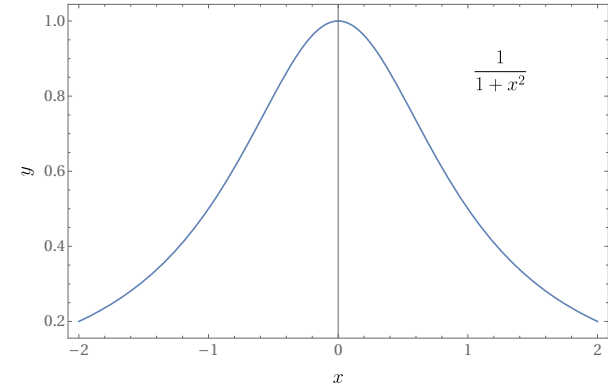
Consider the simple function $y(x) = \frac{1}{1+x^2}$

For the time being take $x \in \mathbb{R}$

It has a Taylor series expression around the origin $\sum_{n=0}^{\infty} (-1)^n x^{2n}$

Convergence? Ratio test

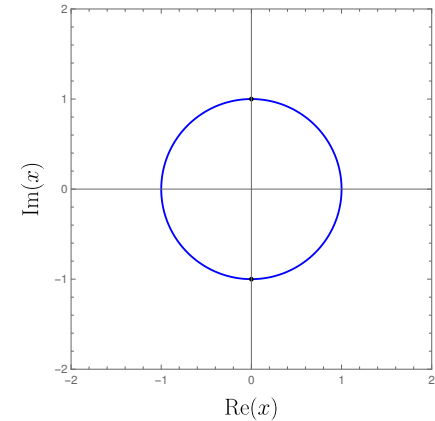
$$\lim_{n \rightarrow \infty} \left| \frac{(-1)^{n+1} x^{(2n+2)}}{(-1)^n x^{2n}} \right| < 1 \quad \Rightarrow \quad |x| < 1$$



So what is the problem?

Back in grade school we learned that the obstruction to convergence is a singular point

Complexify x and solve for the poles $x = \pm i$



Consider fluctuations around a homogeneous-isotropic equilibrium state,

$$T^{\mu\nu} = (T^{\mu\nu})_{eq} + \delta T^{\mu\nu}$$

Combine the EOM with constitutive relations and Fourier expand.

—————▶ Linear system $L(q^2, \omega)_{\alpha\mu} \delta T^{0\mu} = 0$

Nontrivial solutions if:

—————▶ $\det(L(q^2, \omega)_{\alpha\mu}) = P(q^2, \omega) = 0$

Define: *Hydrodynamic Spectral Curve* $P(q^2, \omega) = \det(L_{\alpha\mu})$

If $P(q^2, \omega)$ is analytic at $(q, \omega) = (q_0, \omega_0)$ then by the implicit function theorem

$\omega(q^2)$ can be in general expanded as a Puiseux series,

$$\omega(q^2) = \sum_{i=0}^{\infty} a_k (q^2 - q_0^2)^{i/m}$$

with radius of convergence R given by the distance to the nearest critical point q_c defined by $P(q_c^2, \omega_c) = 0, \partial_\omega P(q_c^2, \omega_c) = 0$ with $R = |q_0 - q_c|$

AdS/CFT – Background solutions

Asymptotically AdS₅ solutions dual to a $\mathcal{N} = 4$ SYM theory in the presence of an external magnetic field and uniform charge density.

$$S = \frac{1}{16\pi G_5} \int d^5x \sqrt{-g} \left(R - 2\Lambda - \frac{L^2}{4} F^{\mu\nu} F_{\mu\nu} \right) - \frac{\gamma}{6} \int A \wedge F \wedge F + S_{ct}$$

For $\gamma = 2/\sqrt{3}$ the theory is a consistent truncation of supergravity, and the chiral anomaly coefficient of the dual theory is then given by $C = \gamma/6$

We will put this term to zero for now

The dual energy momentum tensor and current obey the (non)-conservation laws

$$\nabla_\mu T^{\mu\nu} = F^{\nu\mu} J_\mu$$

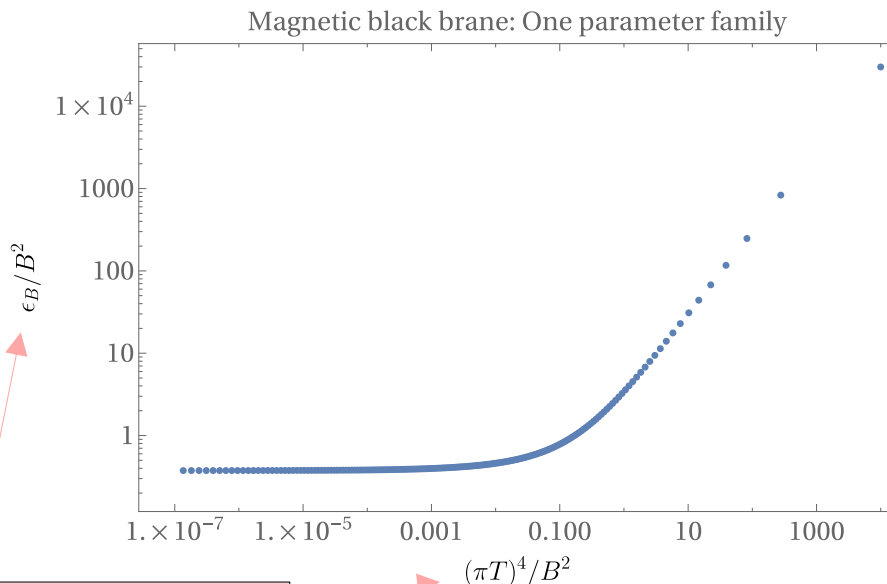
$$\nabla_\mu J^\mu = \frac{3}{4} C \epsilon^{\alpha\beta\gamma\delta} F_{\alpha\beta} F_{\gamma\delta}$$

We will take the ansatz of the metric and gauge field as the following, in Edd. Fink. coordinates

$$ds^2 = \frac{1}{z^2} \left(-U(z) d\nu^2 - 2d\nu dz + v(z)^2 (dx_1^2 + dx_2^2) + w(z)^2 dx_3^2 \right)$$

$$A = \frac{B}{2} (-x_2 dx_1 + x_1 dx_2)$$

Leads to a one parameter family of solutions parameterized by B/T^2



Dimensionless ratios are important!

[D'Hoker, Kraus, 2009]

[D'Hoker, Kraus, 2009]

[D'Hoker, Kraus, 2010]

[Ammon, Leiber, Macedo, 2016]

[Gauntlett, Varela, 2007]

[Buchel, Liu, 2006]

[Gauntlett, Colgain, Varela, 2006]

Metric and gauge field perturbations of the geometry correspond to deviations away from thermal equilibrium. i.e. hydrodynamics

1. Linearize around a background geometry

$$g_{\mu\nu} = g_{\mu\nu}^{(0)} + \epsilon g_{\mu\nu}^{(1)} + O(\epsilon^2), \quad A_\mu = A_\mu^{(0)} + \epsilon A_\mu^{(1)} + O(\epsilon^2)$$

2. Solve the fluctuations equation of motion

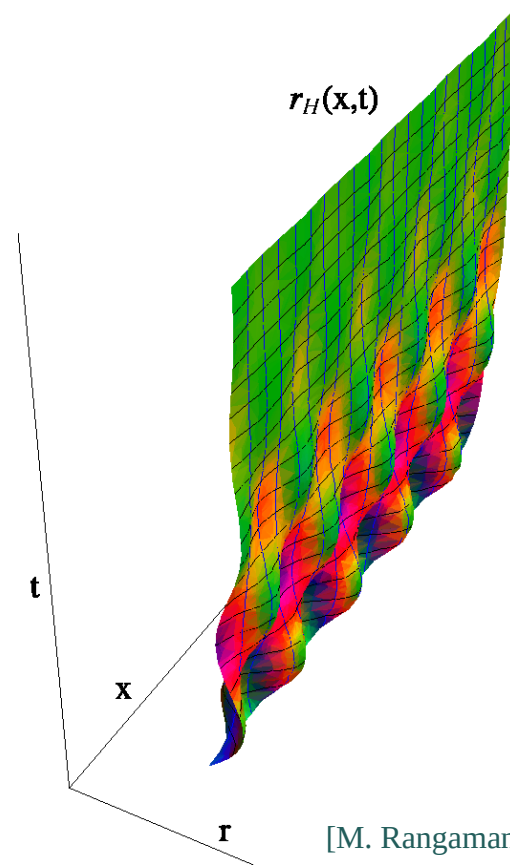
$$0 = \frac{1}{2} \nabla_\mu \nabla_\nu g^{(1)} - \frac{1}{2} \nabla^\lambda \nabla_\lambda g_{\mu\nu}^{(1)} + \nabla^\lambda \nabla_{(\mu} g_{\nu)\lambda}^{(1)} - \frac{2\Lambda}{D-2} g_{\mu\nu}^{(1)} - s(A^{(0)}, A^{(1)}),$$

$$0 = 8\partial_\mu \left[\sqrt{-g^{(0)}} \left(\frac{1}{2} (g^{(0)\alpha\beta} g_{\beta\nu}^{(1)}) F^{(0)\mu\nu} + F^{(1)\mu\nu} + (g^{(0)\mu\alpha} g^{(1)\nu\beta} + g^{(1)\mu\alpha} g^{(0)\nu\beta}) F_{\alpha\beta}^{(0)} \right) \right] \\ + \gamma \epsilon^{\nu\alpha\beta\lambda\sigma} \left(F_{\alpha\beta}^{(0)} F_{\lambda\sigma}^{(1)} + F_{\alpha\beta}^{(1)} F_{\lambda\sigma}^{(0)} \right)$$

Subject to the boundary conditions:

1. Ingoing at the horizon (gives infalling wave)
2. Vanishing at the conformal boundary (Dirichlet)
3. Vanishing partial derivative of the frequency (Neumann-ish)

3. Extract the critical momentum



AdS/CFT – Sectors and mode decomposition

A general fluctuation can then be represented as

$$g_{\mu\nu}^{(1)}(z, x^i) = \int d^4k e^{-ik_i x^i} g_{\mu\nu}^{(1)}(z, k^i), \quad A_\mu^{(1)}(z, x^i) = \int d^4k e^{-ik_i x^i} A_\mu^{(1)}(z, k^i)$$

Align the momentum with the magnetic field $\vec{k} \parallel \vec{B}$ leaving us with a residual $SO(2)$

Spin	Fields	$\begin{cases} B \rightarrow -B \\ \gamma \rightarrow -\gamma \end{cases}$	$\begin{cases} B \rightarrow -B \\ k \rightarrow -k \end{cases}$	Hydrodynamic modes $\lim_{q \rightarrow 0} w(q) = 0$
2^+	$g_{x_1 x_2}, g_{x_1 x_1} - g_{x_2 x_2}$	$2^+ \rightarrow 2^+$	$2^+ \rightarrow 2^+$	X
1^+	$g_{\nu x_+}^{(1)}, g_{x_3 x_+}^{(1)}, A_{x_+}^{(1)}$	$1^+ \rightarrow 1^-$	$1^+ \rightarrow 1^-$	✓
1^-	$g_{\nu -}^{(1)}, g_{x_3 -}^{(1)}, A_-^{(1)}$	$1^- \rightarrow 1^+$	$1^- \rightarrow 1^+$	✓
0^+	$g_{x_3 x_3}^{(1)}, g_{x_1 x_1}^{(1)} + g_{x_2 x_2}^{(1)}, g_{\nu x_3}^{(1)}$ $g_{\nu\nu}^{(1)}, A_\nu^{(1)}, A_{x_3}^{(1)}$	$0^+ \rightarrow 0^+$	$0^+ \rightarrow 0^+$	✓

AdS/CFT – Constructing solutions

In Edd. Fink. Coordinates the fields are regular at the horizon $g_{ab}^{(1)} = (z-1)^\alpha z^\beta g_{ab}^{(1),Reg}(z)$ for $\alpha = 0$

Decompose the fluctuation equations

$$M(q^2, \mathfrak{w})\phi = (M_0(q^2) + M_1(q^2)\mathfrak{w} + M_2(q^2)\mathfrak{w}^2 + \dots)\phi = 0$$

$$M_i = M_i^2(q^2)\frac{d^2}{dz^2} + M_i^1(q^2)\frac{d}{dz} + M_i^0(q^2)$$

ϕ -fields of a given sector

$$\mathfrak{w} = \frac{\omega}{2\pi T}, \quad q = \frac{q}{2\pi T}$$

To obtain QNM modes one solves $P(q^2, \mathfrak{w}) = \det(M(q^2, \mathfrak{w})) = 0$ See for instance [Jansen, 2017]

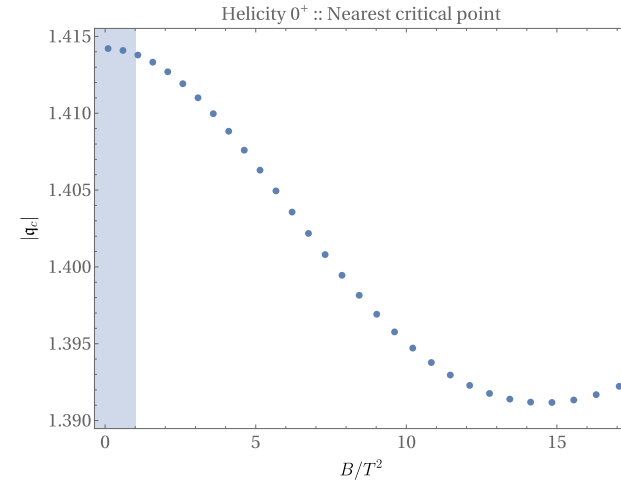
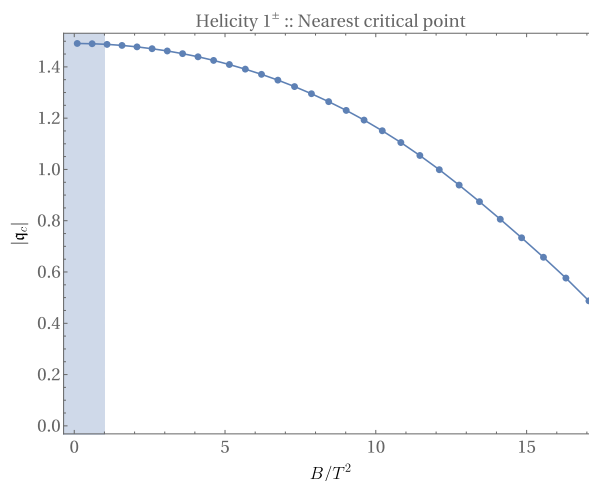
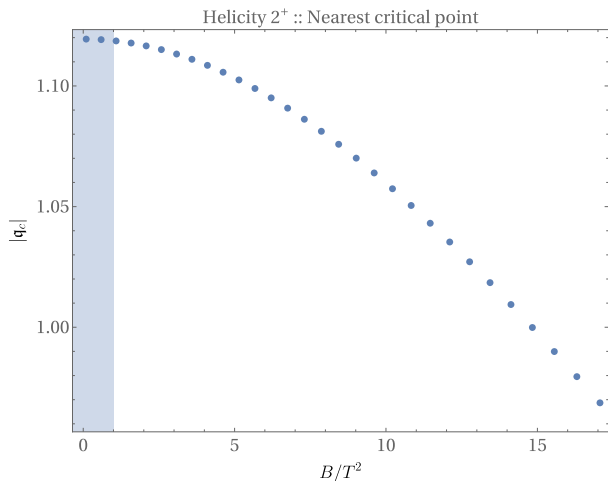
To obtain critical points one simultaneously solves $P(q^2, \mathfrak{w}) = 0, \partial_{\mathfrak{w}}P(q^2, \mathfrak{w}) = 0$

Numerical procedure:

1. Guess initial value of q^2 & \mathfrak{w} and use a pseudo-spectral Chebyshev decomposition to represent the matrix M
2. Use a Newton-Raphson root finder to find q^2 & \mathfrak{w} which satisfy the two conditions

$$(q_i^2, \mathfrak{w}_i) = (q_{i-1}^2, \mathfrak{w}_{i-1}) - J^{-1} \cdot \chi \quad \chi = (P, \partial_{\mathfrak{w}}P), \quad J_{ij} = \frac{\partial \chi_i}{\partial x^j}, \quad x = (q^2, \mathfrak{w})$$

What happens near the regime of validity of hydro in external fields? $B \ll T^2$



Helicity 2

No hydro mode

- Remains smooth
- Remains roughly the same value for regime of validity of hydro approx.

Helicity 1

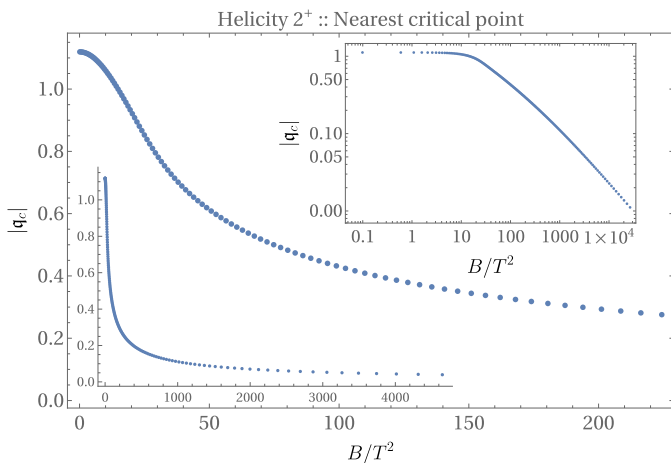
- Remains smooth
- Remains roughly the same value for regime of validity of hydro approx.
- Drops significantly for increasing B/T^2

Helicity 0

- Remains smooth
- Remains roughly the same value for regime of validity of hydro approx.

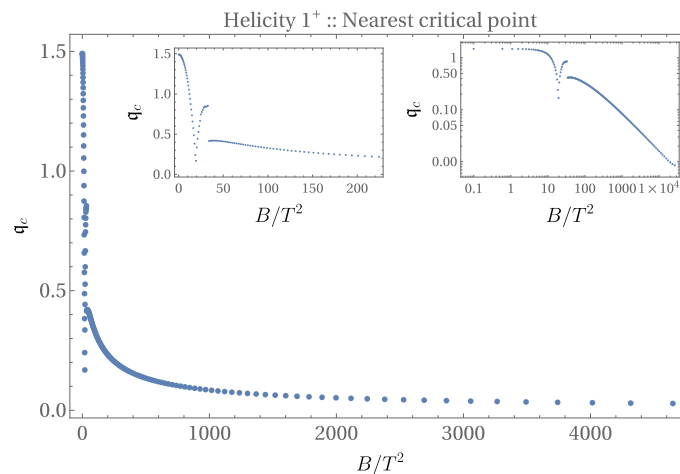
The bound on the radius of convergence remains finite and non-zero within and outside of the regime of applicability of hydro.

What happens far outside the regime of validity of hydro in external fields? $B \gg T^2$



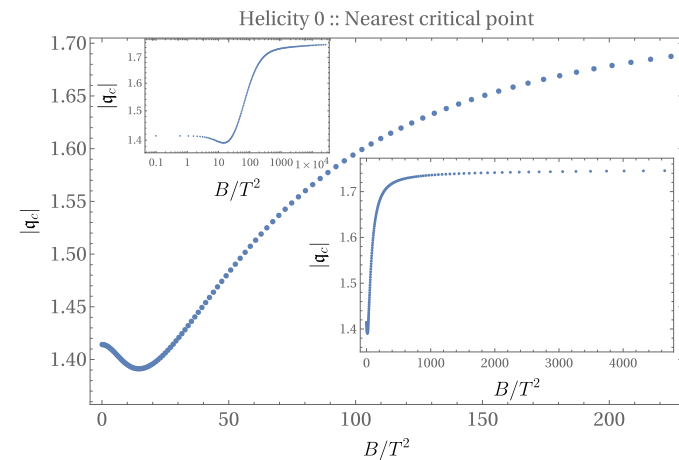
Helicity 2

- Remains smooth
- Appears to asymptote to zero



Helicity 1

- Gaps
- Appears to asymptote to zero



Helicity 0

- Remains smooth
- Appears to asymptote to finite value

The bound on the radius of convergence remains finite and non-zero for substantially longer than one might expect.

Current “observations”

- The lack of evidence of the CME in HICs necessitates deeper theoretical control to moderate the analysis and direction of experimental searches.
- The radius of convergence of the linearized hydrodynamic gradient expansion:
 - *decreases* for *shear* modes of SYM in strong external magnetic fields
 - *decreases* for *sound* modes of SYM in strong external magnetic fields
- Can this be extended to full MHD?

Stay tuned:

- Work on critical points with anomalous current is in progress.
- Full Einstein-Maxwell-Chern-Simons geometry contains a quantum critical point in its dual CFT
- Near the critical point in QCD as well as quantum critical point of EMCS in AdS, hydro should break down, these methods should provide a look at how rapidly this occurs
- *An example of convergence in hydrodynamics in strong external fields, C. Cartwright, will appear soon.*

