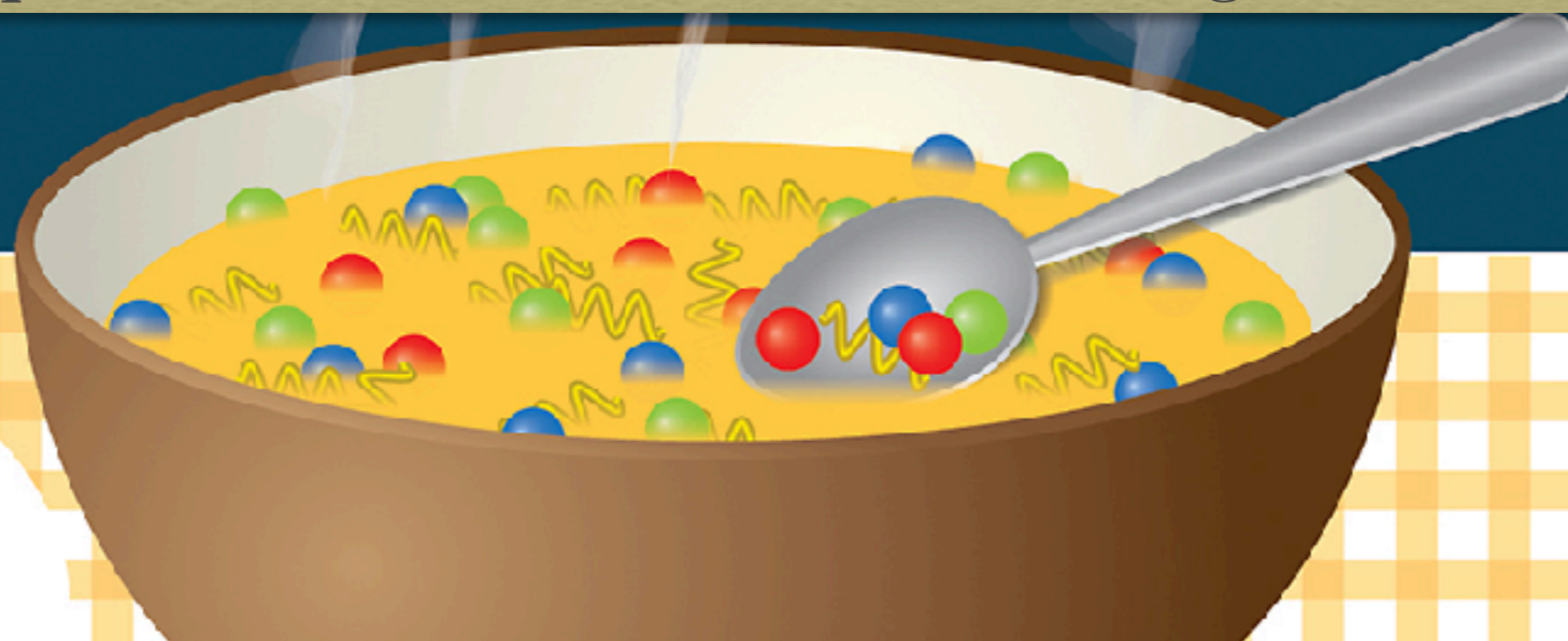




Illinois Center for Advanced Studies of the Universe



Studying $q\bar{q}$ pairs in the initial state as a background for the CME



Jacquelyn Noronha-Hostler
University of Illinois Urbana-Champaign

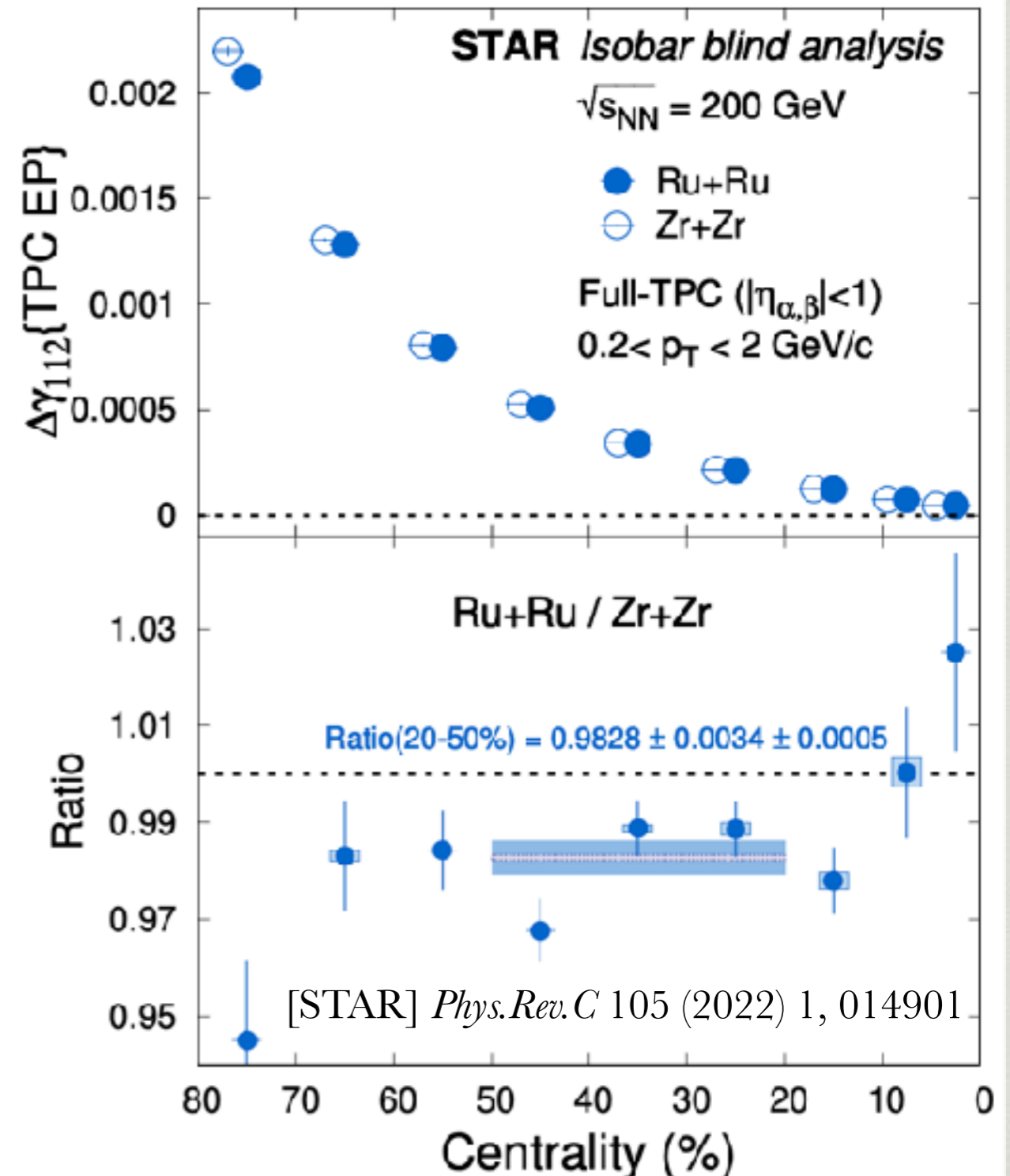
ECT* HOLOGRAPHIC PERSPECTIVES ON
CHIRAL TRANSPORT

Charge effects in the isobar run

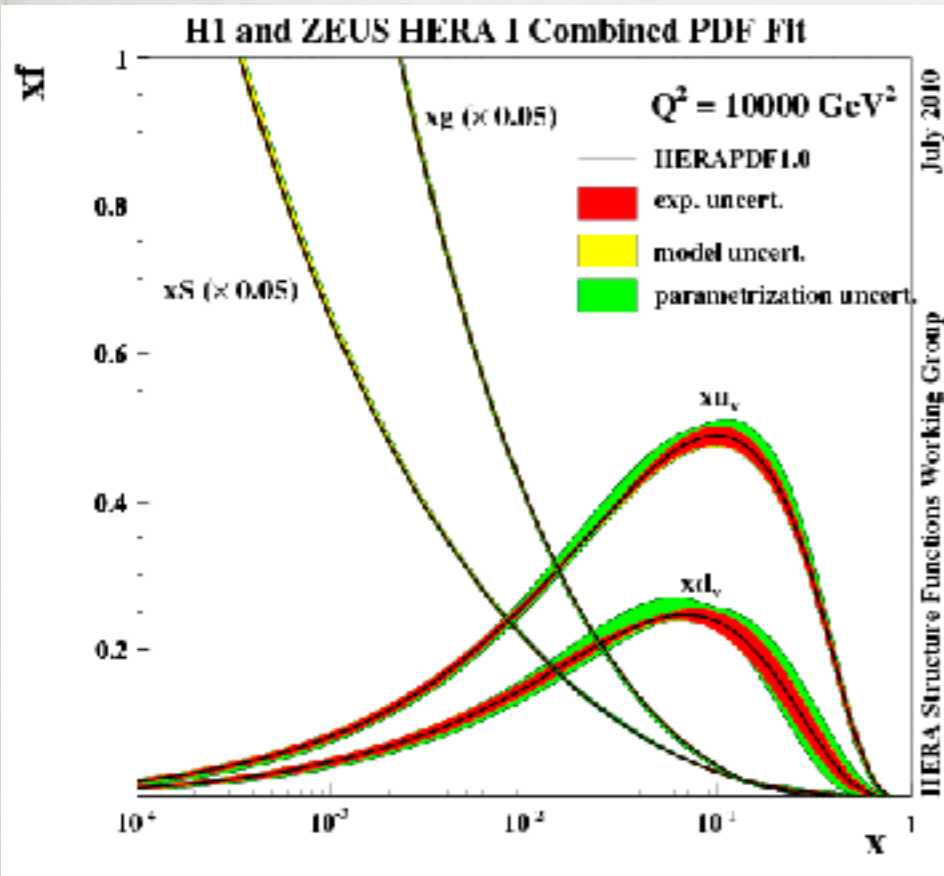
Charge correlates observables where α, β are particles of interest e.g. same-sign/opposite sign charges

$$\gamma_{ijk} = \langle \cos \left(i\phi_{\alpha} + j\phi_{\beta} - k\Psi_k \right) \rangle$$

What role do charge fluctuations from $g \rightarrow q\bar{q}$ play?

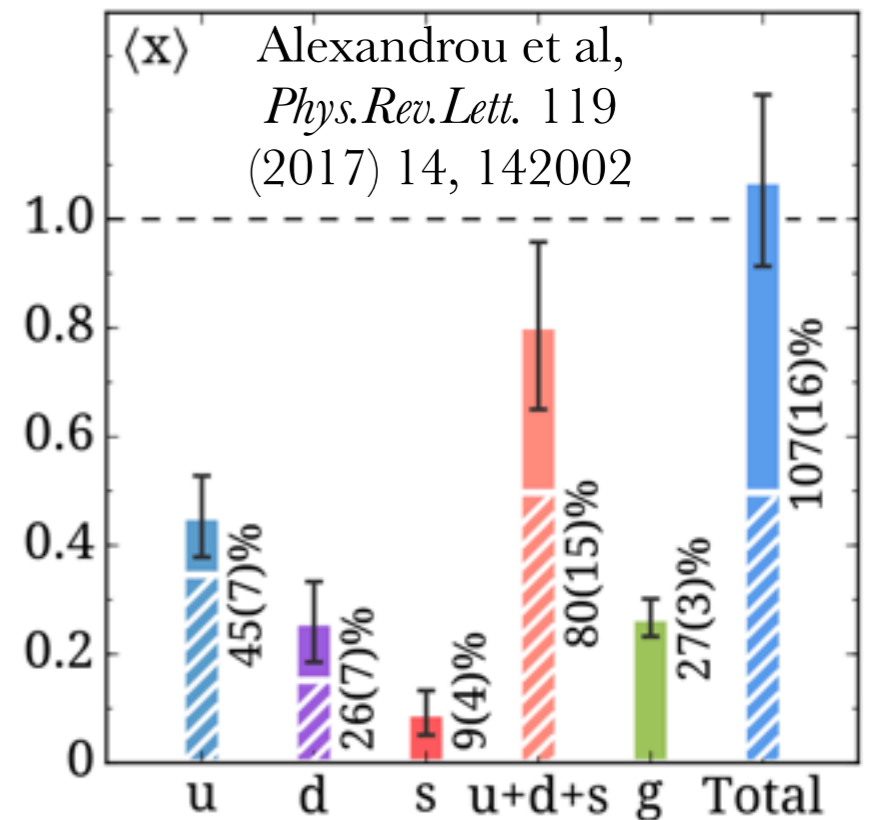
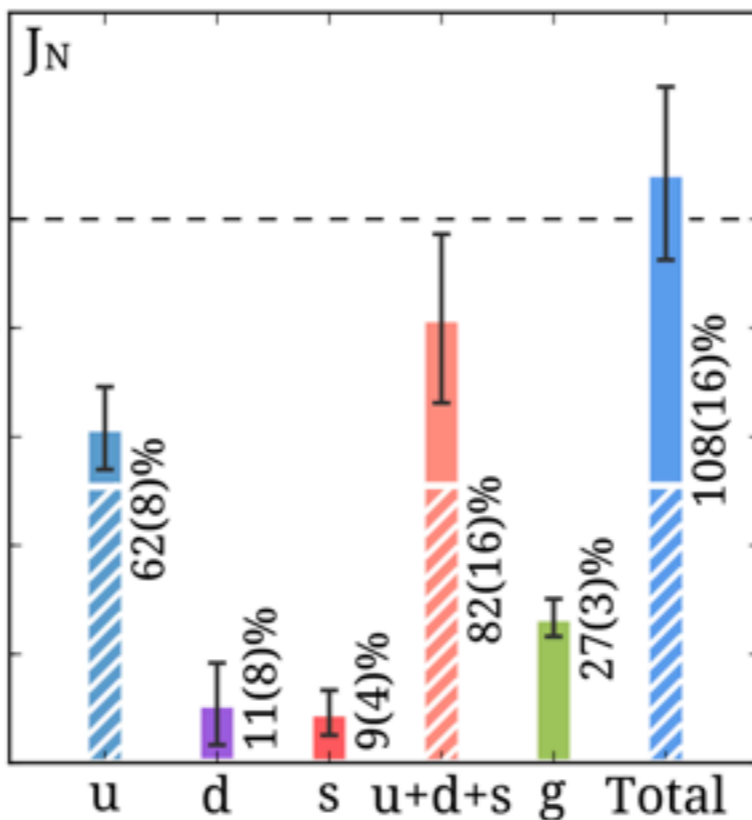


Motivated by sea quarks and lattice QCD

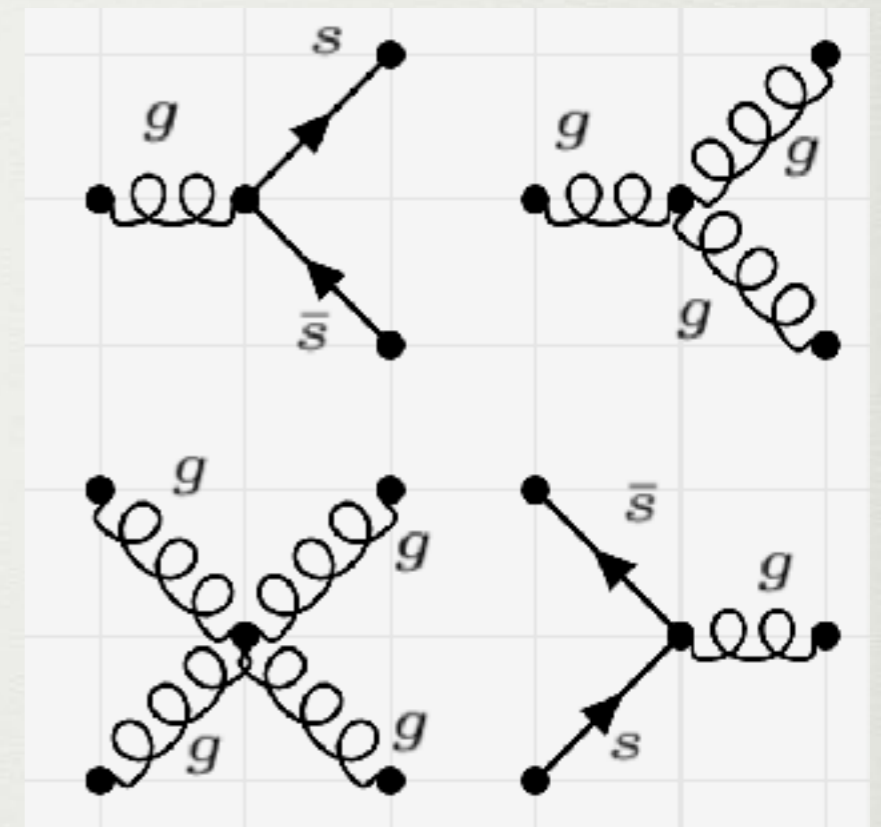
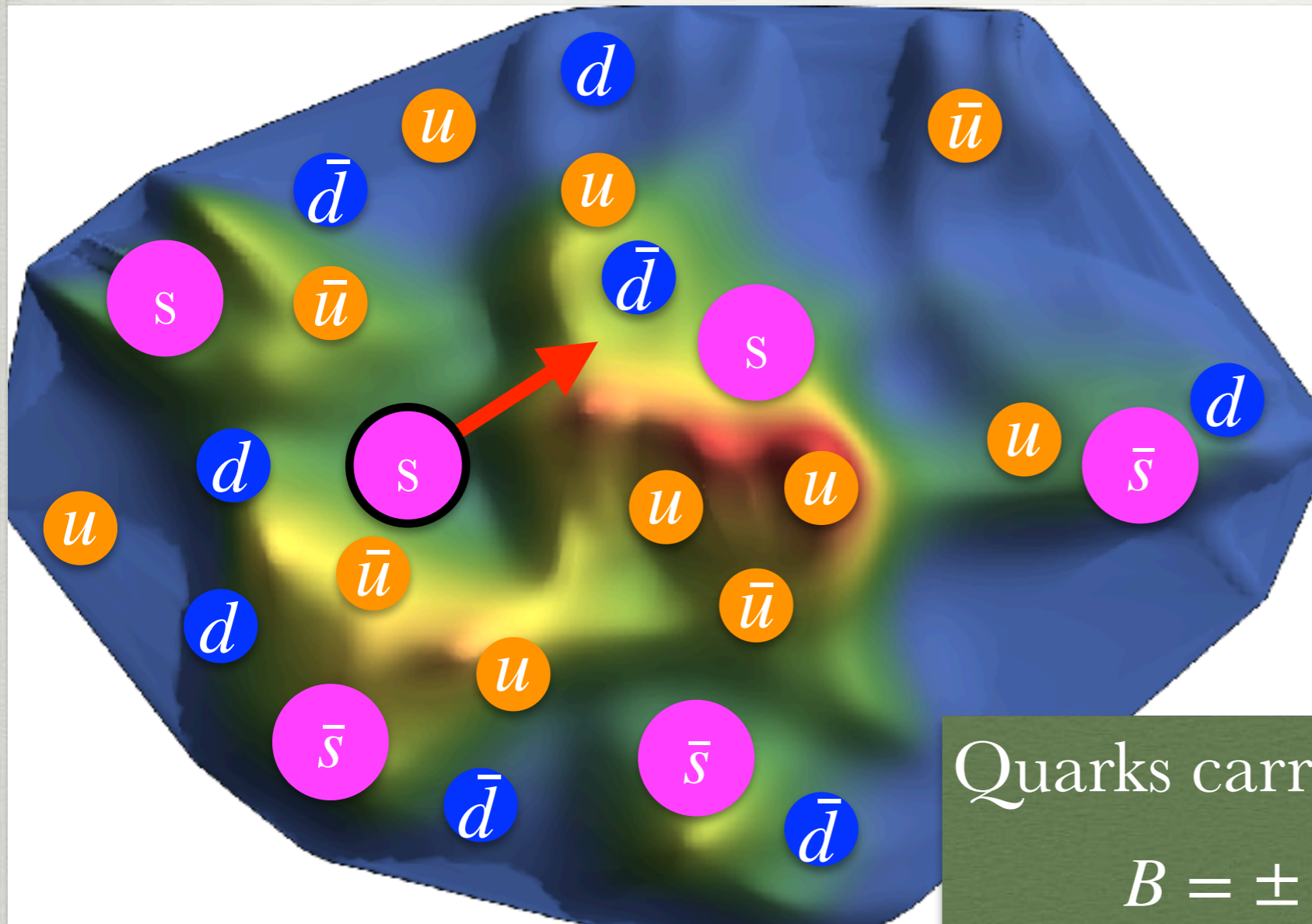


Strange quarks contribute nearly 10% to the proton spin and momentum

Sea quarks contributions grow at low-x



Beyond simple (gluon) initial conditions: BSQ charges



Quarks carry ~ 3 conserved charges:

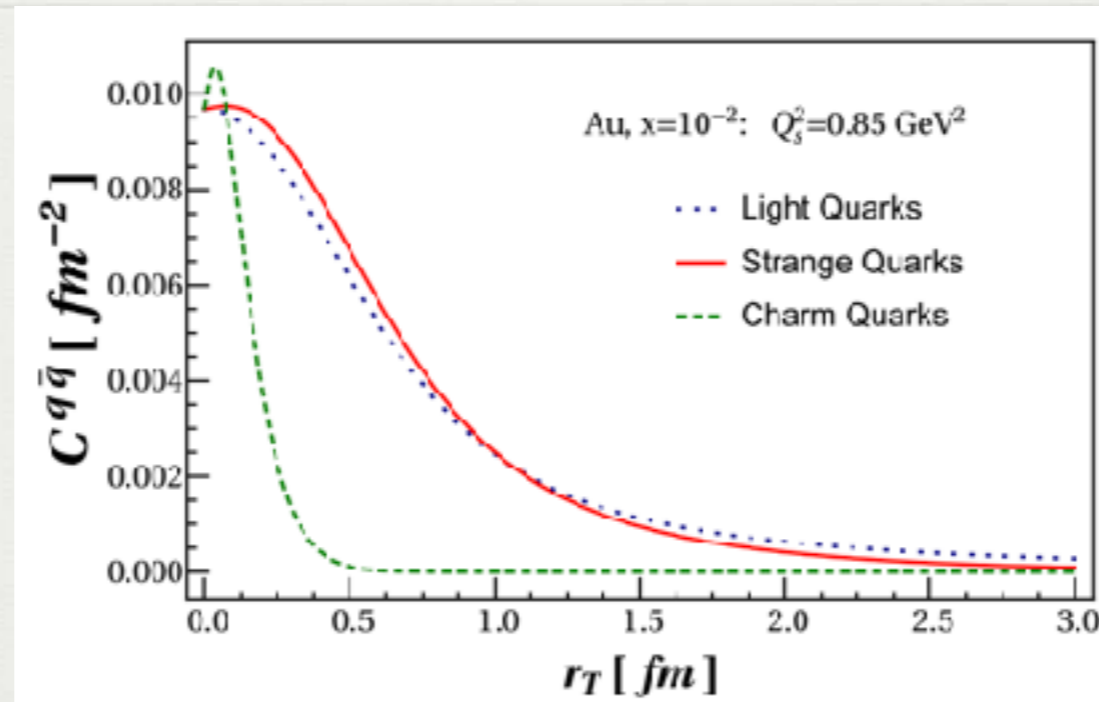
$$B = \pm \frac{1}{3} \text{ Baryon number}$$

$$S = \pm 1 \text{ Strangeness}$$

$$Q = \pm \frac{1}{3} \text{ (d,s) or } Q = \pm \frac{2}{3} \text{ (u)}$$

Splittings into $q\bar{q}$ pairs

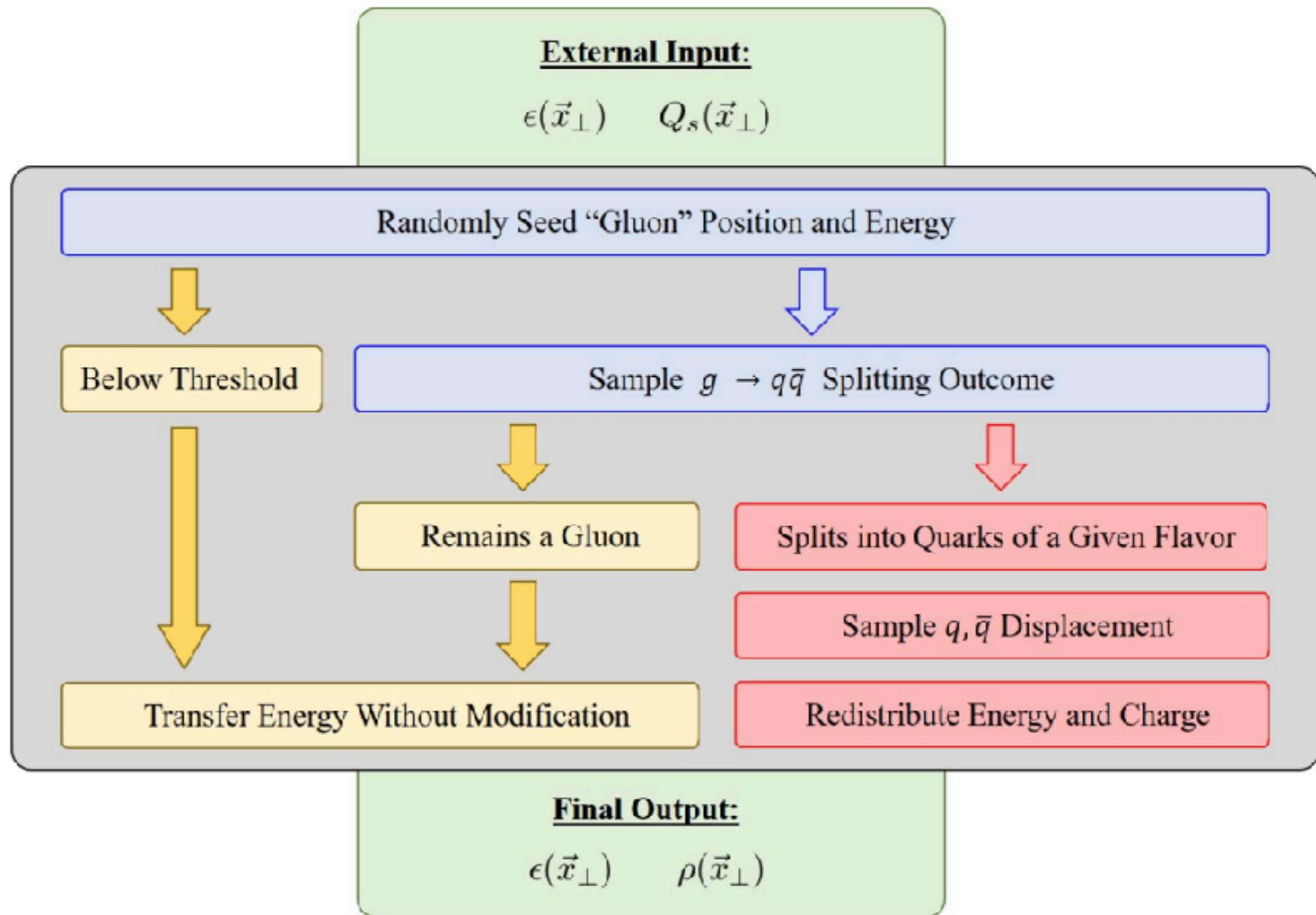
Martinez, Sievert, Wertepny JHEP 02 (2019) 024; JHEP 1807 (2018) 003



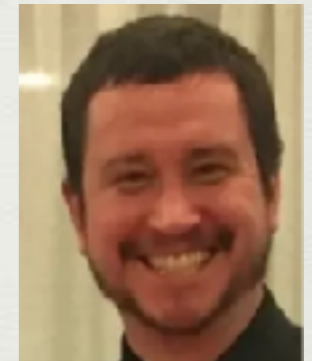
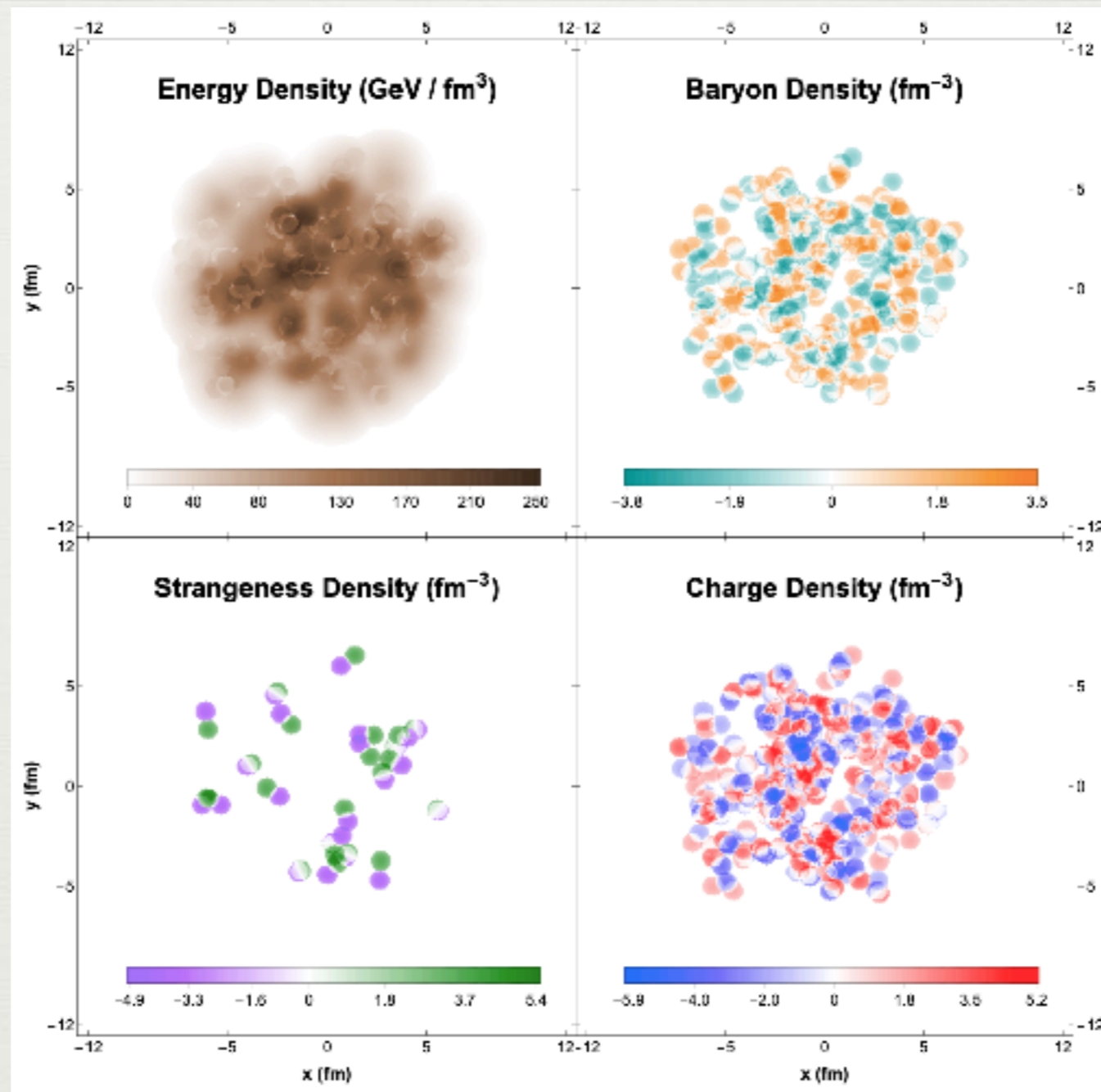
- $g \rightarrow q\bar{q}$ production from CGC (other methods possible, just need splitting probabilities)
- Spatial correlations calculated in dense-light limit
- Currently using single-pair production only (double pair production is possible, just not done yet in ICCING)

ICCING algorithm

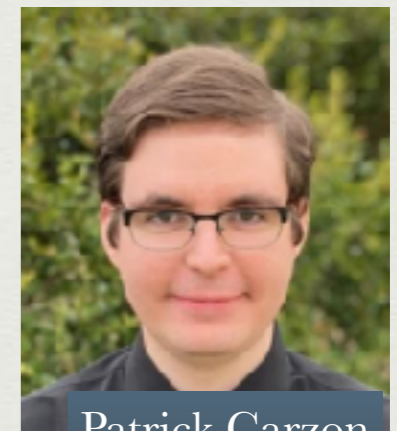
Initial Conserved Charges in Nuclear Geometry



Example ICCING initial condition Trento for PbPb $\sqrt{s_{NN}} = 5.02$ TeV



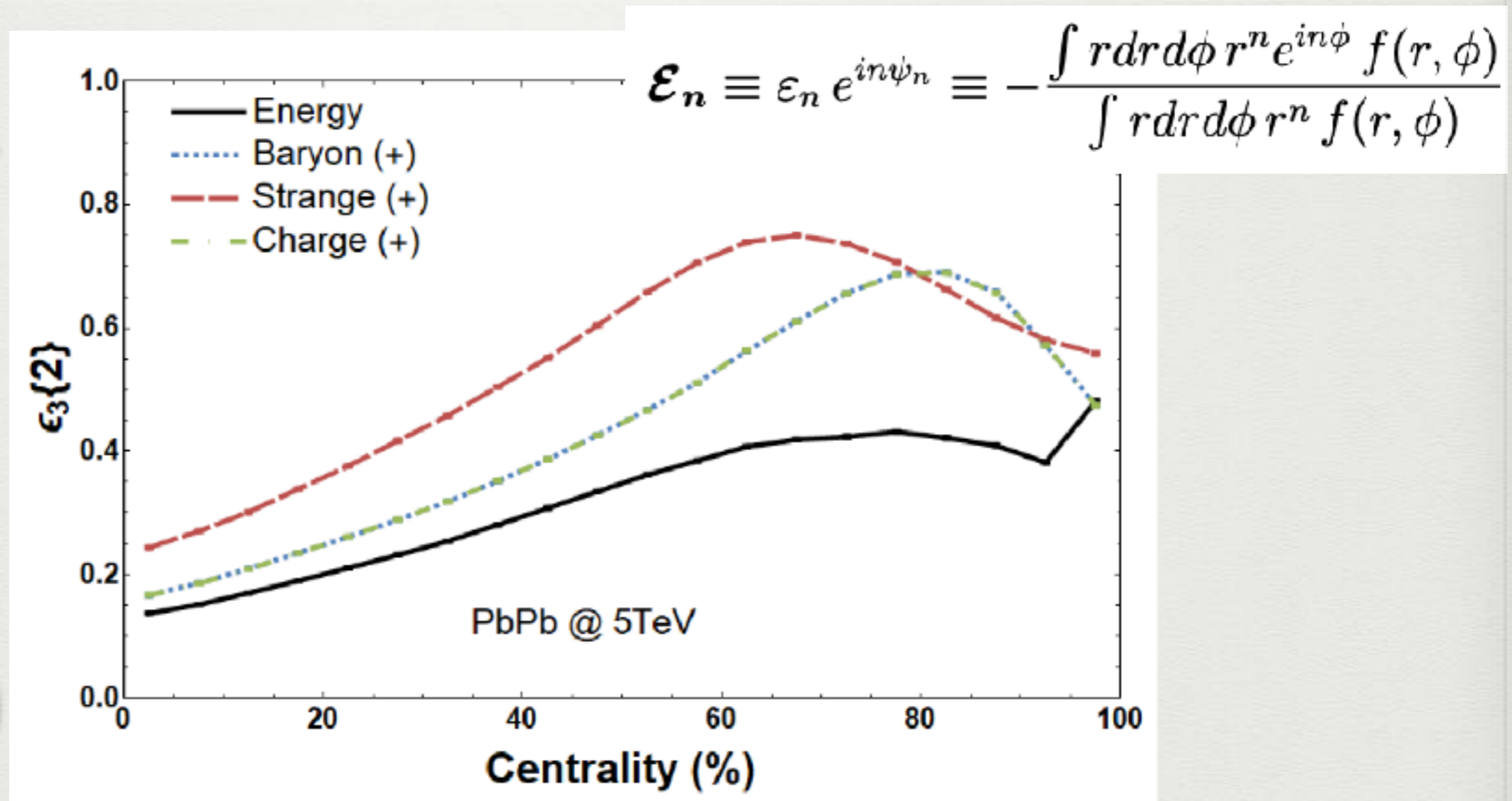
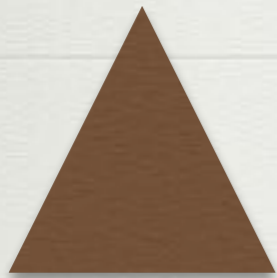
Matt Sievert
Faculty NMSU
Former Postdoc



Patrick Carzon
PhD student

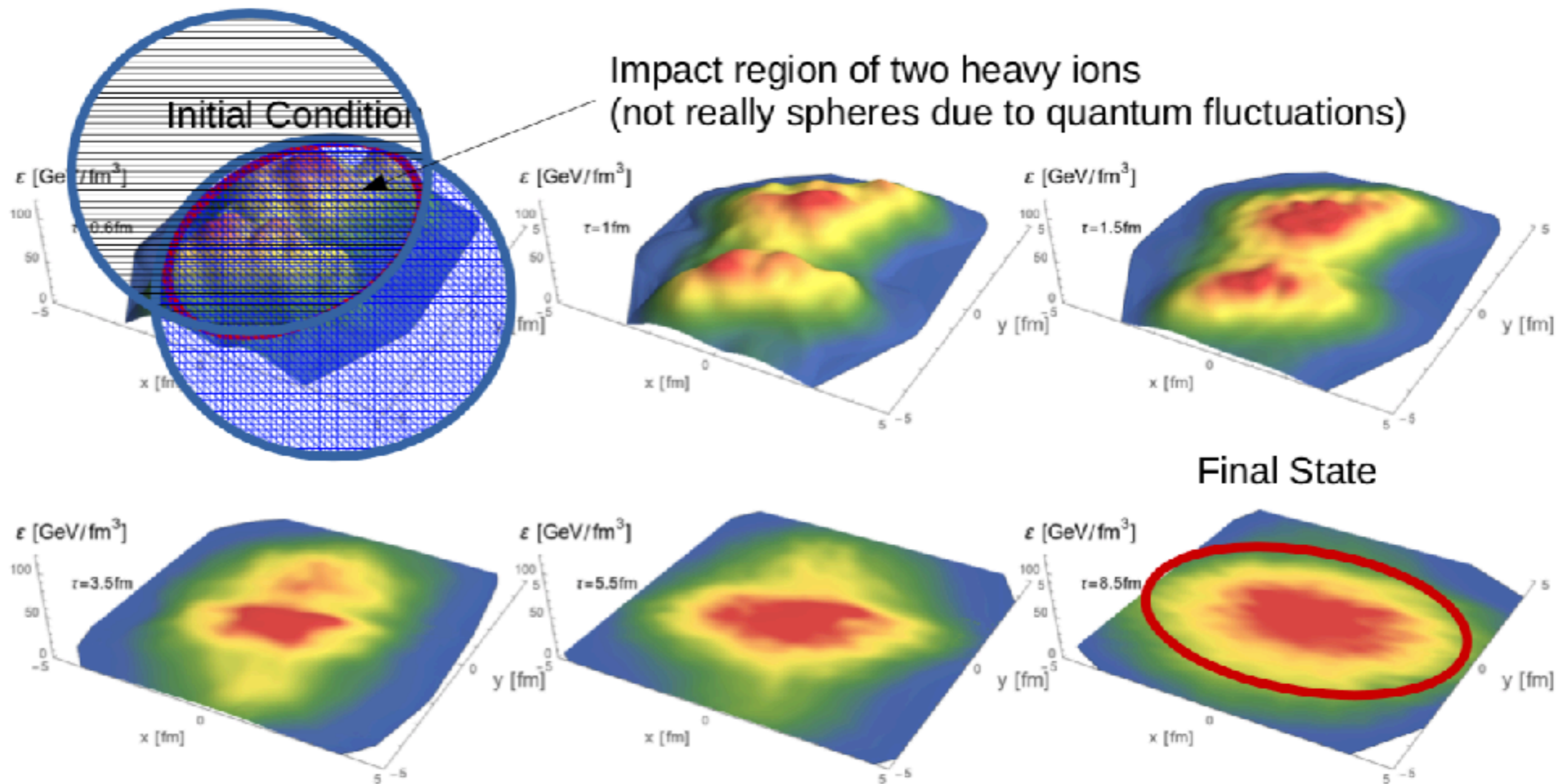
Carzon, (JNH) et al, *Phys.Rev.C* 105 (2022) 3, 034908; [1911.10272](https://arxiv.org/abs/1911.10272) [nucl-th]

ICCING leads to distinct differences in initial state geometry



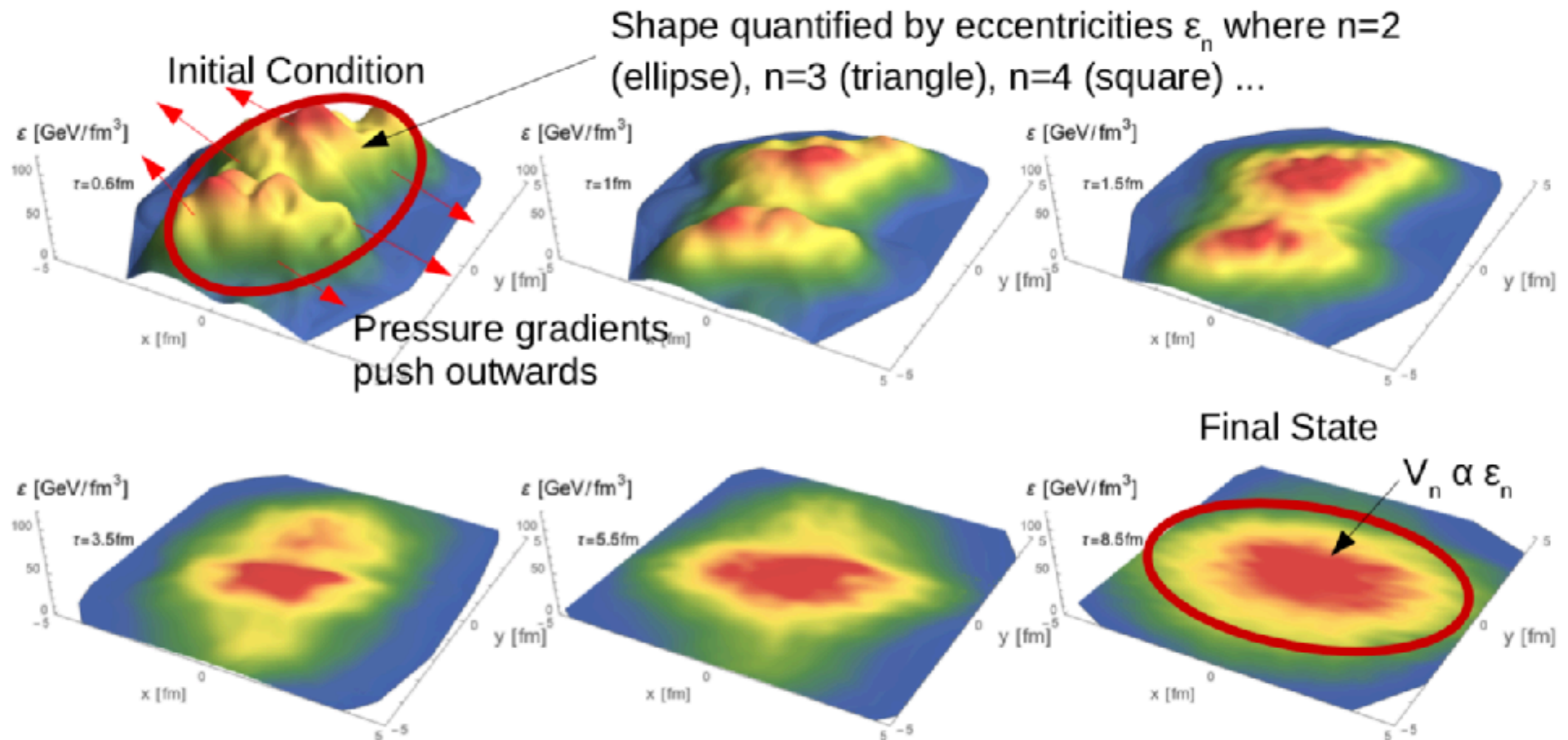
Here only positive charge is considered to calculate eccentricities.
Eccentricities of BSQ charges ill-defined (working in progress)

How does $\mathcal{E}_n \rightarrow V_n$



Eccentricities ϵ_2 's are directly related to the final measured flow observables v_n 's

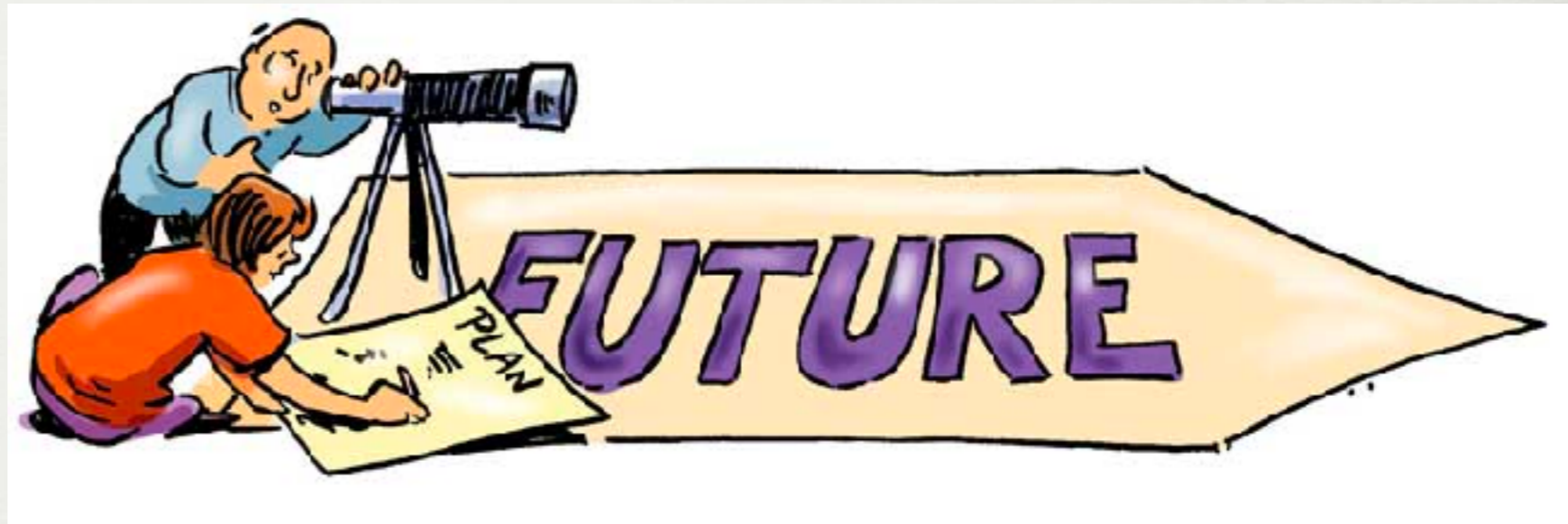
How does $\mathcal{E}_n \rightarrow V_n$



Eccentricities \mathcal{E}_2 's are directly related to the final measured flow observables v_n 's

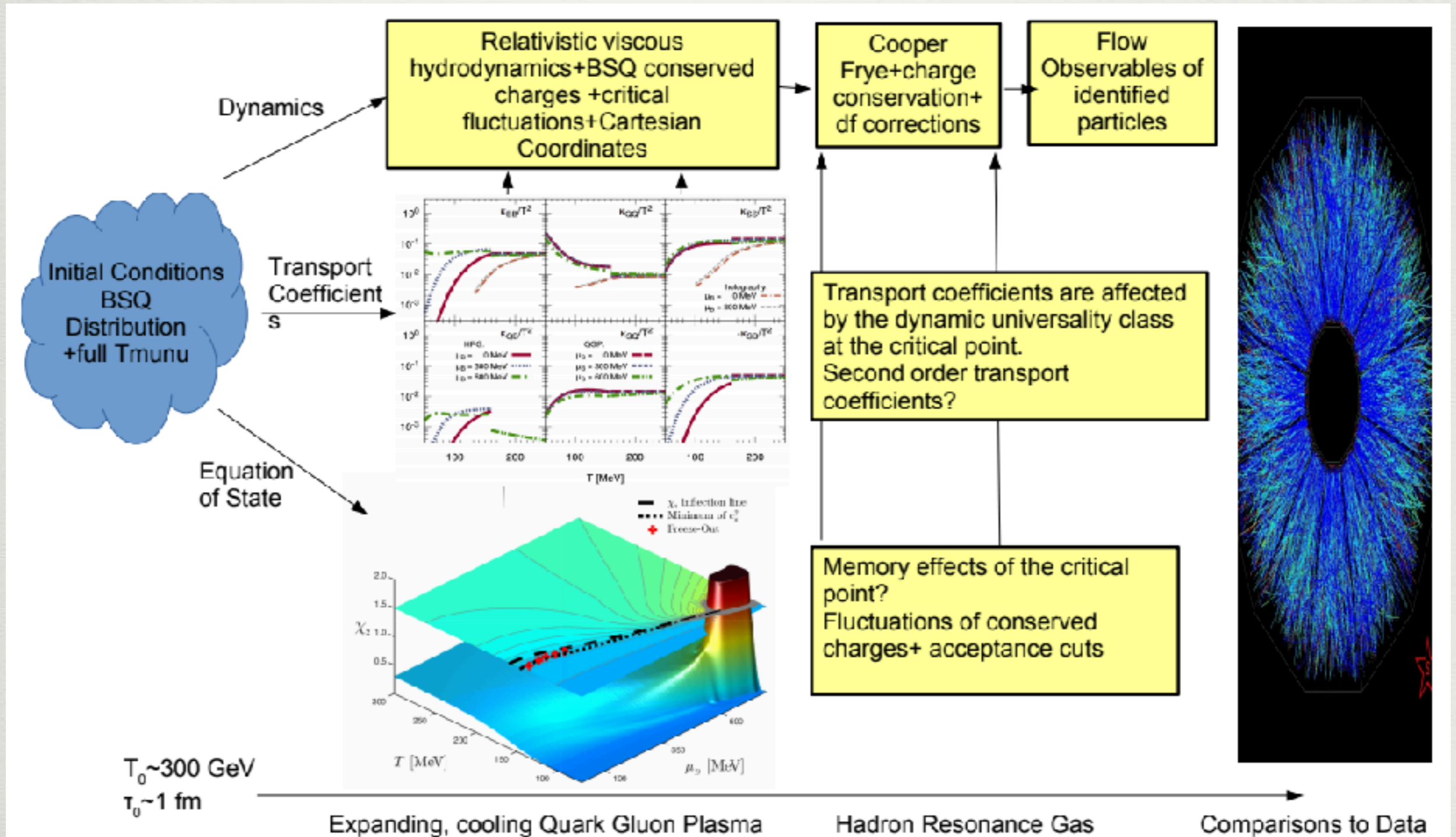
Future possibilities with ICCING

- Spin
- Full $T^{\mu\nu}$ and Q^μ
- Small systems
- Include double-pair production
- Finite net-densities?
- Inclusion of the chiral anomaly?

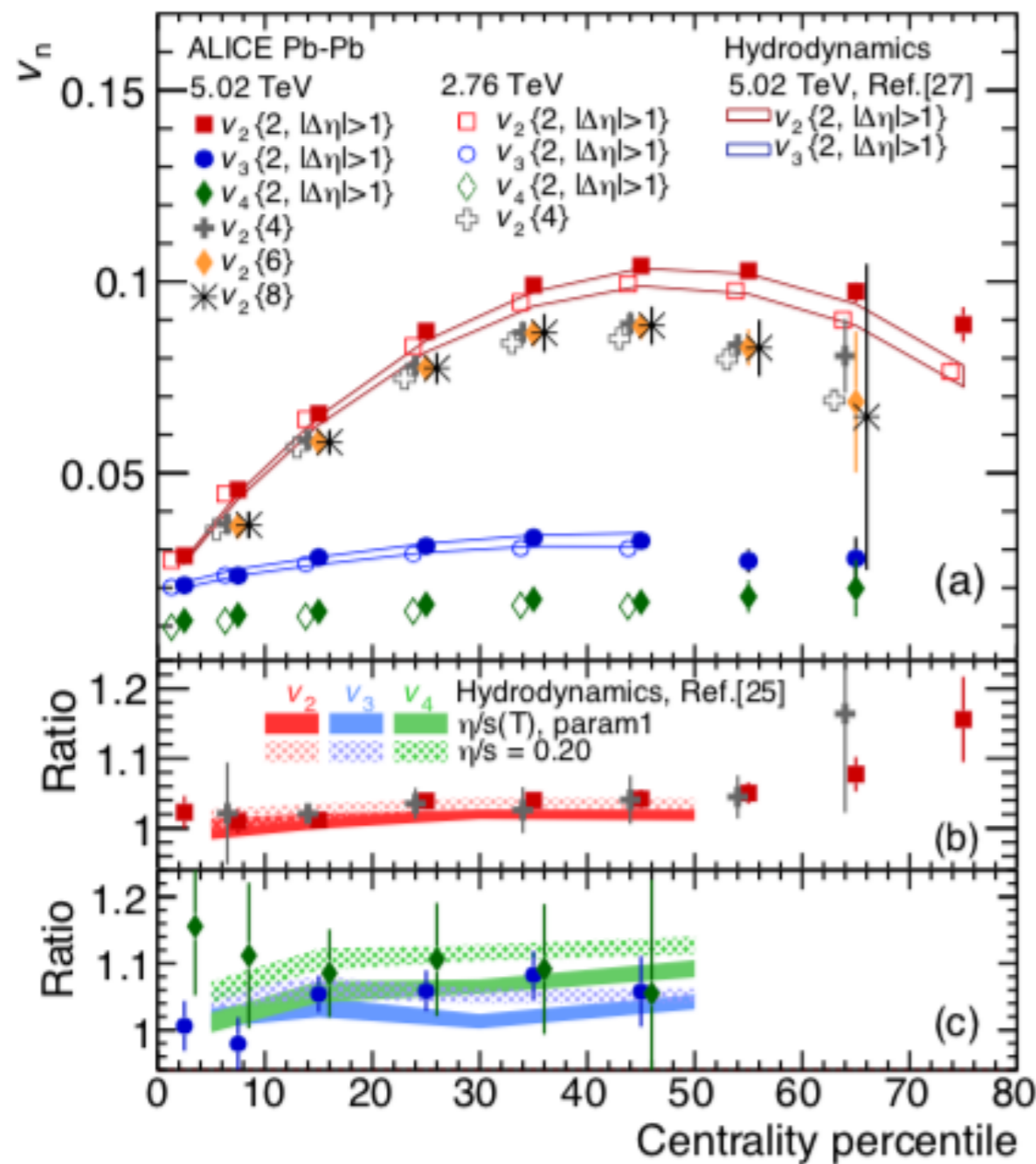


Need relativistic viscous hydrodynamics + 3 conserved charges (BSQ)

Review: Dexheimer, Noronha, JNH, Ratti, Yunes, *J.Phys.G* 48 (2021) 7, 073001



Converting v-USPhydro to include BSQ conserved charges



v-USPhydro used to make predictions at the $\sim 1\%$ level.

ALICE Phys.Rev.Lett. 116 (2016) no.13, 132302

v-USPhydro predictions: JNH et al, Phys.Rev. C93 (2016) no.3, 034912

EKRT predictions: Niemi et al, Phys. Rev. C 93, 014912 (2016)



Travis Dore
Postdoc Bielefeld



Dekra Almaalol
Postdoc

Equation of motion: BSQ diffusion in relativistic hydrodynamics

Almaalol, Dore, JNH [arXiv:2209.11210 [hep-th]]

Ensure only positive entropy production

$$\tau_\pi \dot{\pi}^{\mu\nu} + \pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} + \frac{\tau_\pi \pi^{\mu\nu}}{2}\theta - \frac{\tau_\pi}{2\beta_\pi} \dot{\beta}_\pi \pi^{\mu\nu} - \frac{2\eta}{\beta} \left(\gamma_1^q \nabla^{\langle\mu} n_q^{\nu\rangle} + \frac{1}{2} n_q^{\langle\mu} \nabla^{\nu\rangle} \gamma_1^q \right)$$

Israel-Stewart

DNMR

$$\tau_\Pi \dot{\Pi} + \Pi = - \left(\zeta + \frac{\tau_\Pi}{2} \Pi \right) \theta - \frac{\tau_\Pi}{2\beta_\Pi} \dot{\beta}_\Pi \Pi - \frac{\zeta}{\beta} \left(\gamma_0^q D_\mu n_q^\mu + \frac{1}{2} n_q^\mu \nabla_\mu \gamma_0^q \right)$$

$$\tau_{qq'} \dot{n}_{q'}^\mu + n_q^\mu = - \kappa_{qq'} \nabla^\mu \alpha_{q'} + \frac{\tau_{qq'} n_{q'}^\mu}{2} \theta - \frac{\tau_{qq'}}{2\beta_{qq'}} \dot{\beta}_{qq'} n_{qq'}^\mu - \frac{\kappa_{qq'}}{\beta} \left(\gamma_0^{qq'} \nabla^\mu \Pi - \frac{\Pi}{2} \nabla^\mu \gamma_0^{qq'} \right) - \frac{\kappa_{qq'}}{\beta} \left(\gamma_1^{qq'} \nabla_\nu \pi^{\mu\nu} + \frac{\pi^{\mu\nu}}{2} \nabla_\nu \gamma_1^{qq'} \right)$$

CCAKE: new hydrodynamic code with all BSQ coefficients

Linear thermodynamic stability constraints for BSQ hydro

Almaalol, Dore, JNH [arXiv:2209.11210 [hep-th]]

Following the method from [1]

- (1) derive Lyapunov functional for S^μ in Landau frame
- (2) generalize BSQ fluid
- (3) extract the conditions for positive energy functional

7 stability conditions on both EOS and transport coefficients

Diffusion / $\pi^{\mu\nu}$ coupling

$$\varepsilon + p \geq \frac{T\eta}{\tau_\pi} \lambda^2 (\delta_{n\pi}^q + 1)^2 - \rho_q^2 \beta_n^{qq'}$$

Diffusion matrix

EoS: Moving beyond Lattice QCD

Taylor series provides cross-over Equation of State $\mu_B \lesssim 450$ MeV

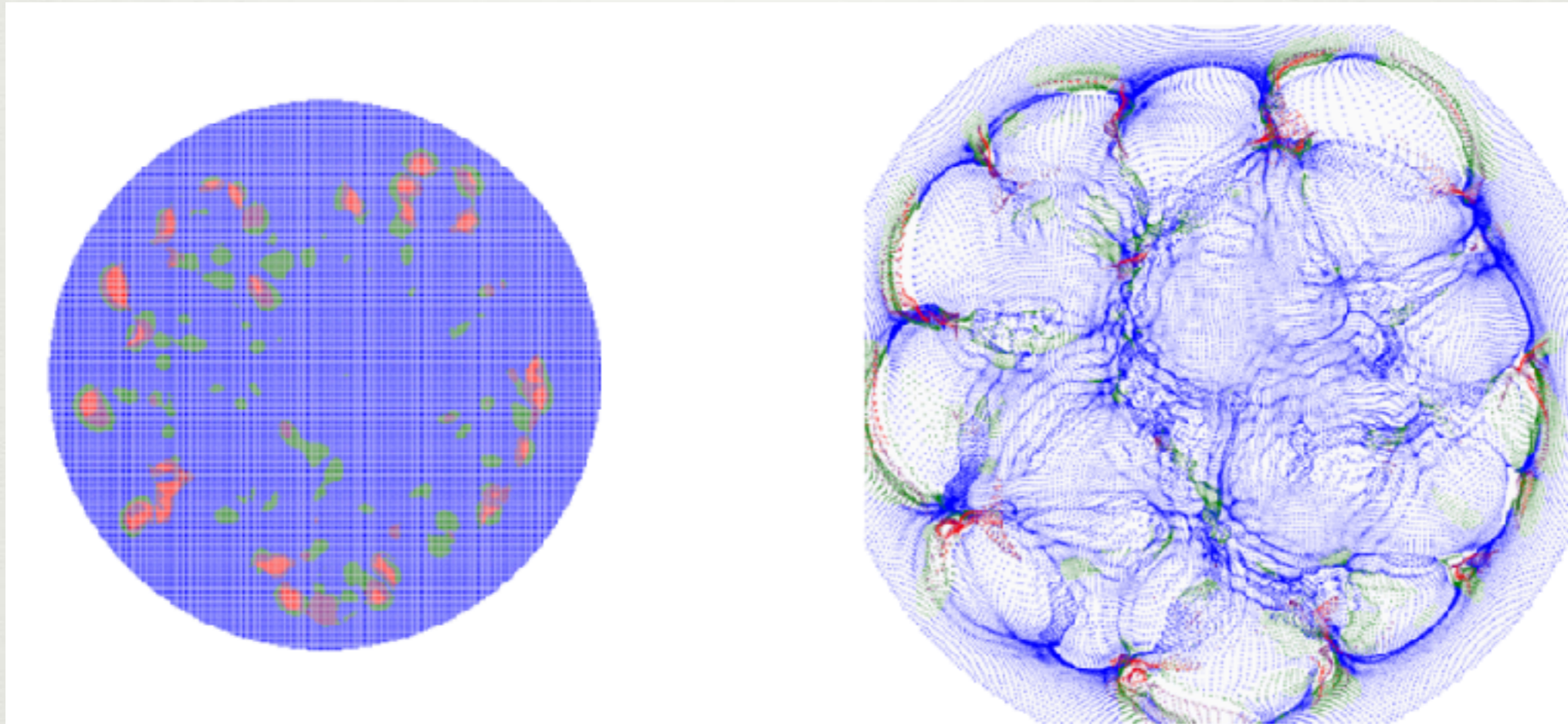
$$\frac{p(T, \mu_B, \mu_Q, \mu_S)}{T^4} = \sum_{i,j,k} \frac{1}{i!j!k!} \chi_{ijk}^{BSQ} \left(\frac{\mu_B}{T}\right)^i \left(\frac{\mu_Q}{T}\right)^j \left(\frac{\mu_S}{T}\right)^k$$

JNH, Paolo Parotto, Claudia Ratti, Jamie Stafford Phys.Rev.C 100 (2019) 6, 064910

- Taylor series not suited for critical point analyses
- Other expansion series have issues far from the critical point: Critelli et al *Phys.Rev.D* 96 (2017) 9, 096026
- Currently only have Lattice QCD results for the full BSQ expansion up to $\mathcal{O}(\mu/T)^4$

Current lattice QCD EOS not enough

Almaalol, Carzon, Dore, Espino, Mroczek, Plumberg, Salinas San Martin, Spsychalla, Sievert, JNH to appear soon



Blue lattice QCD

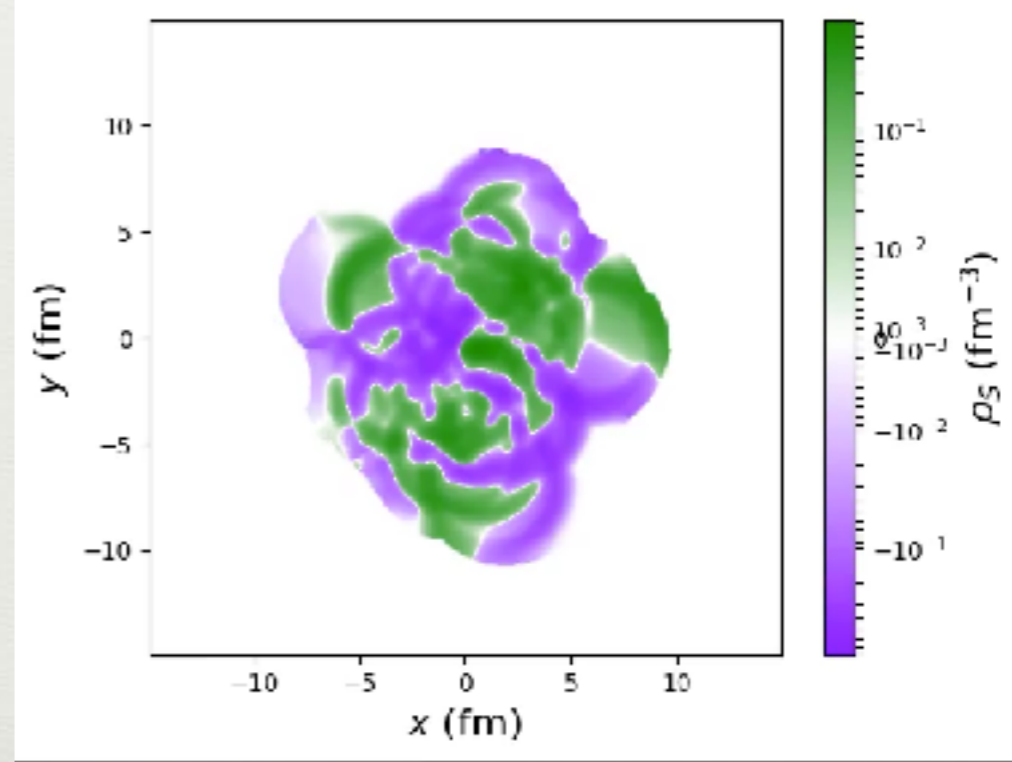
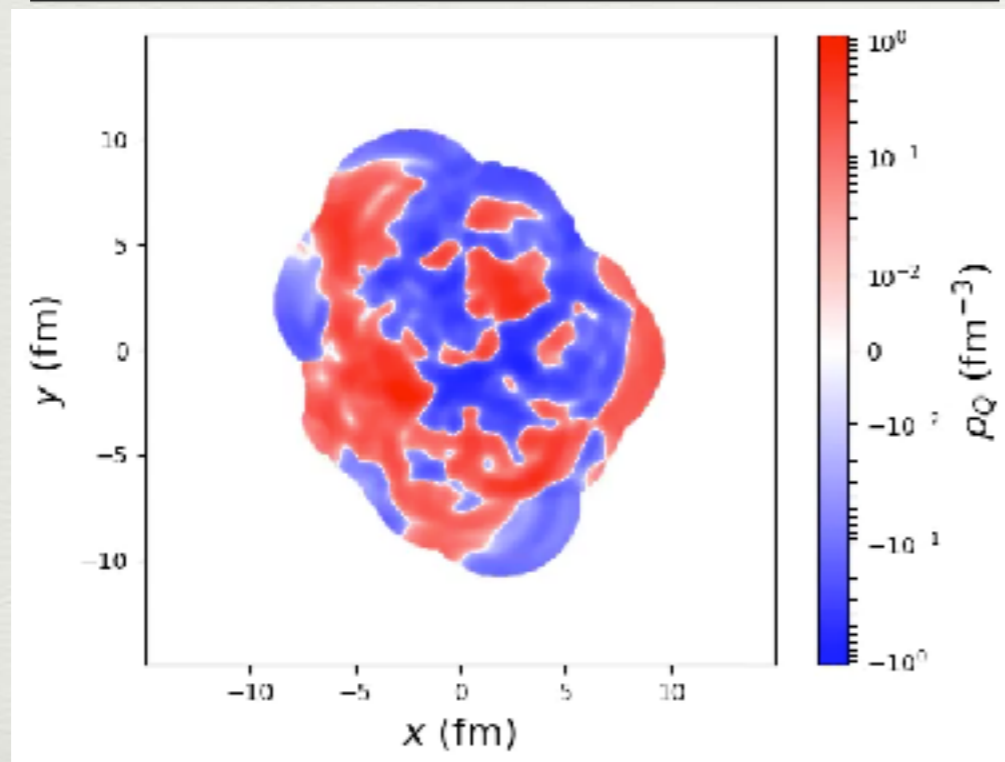
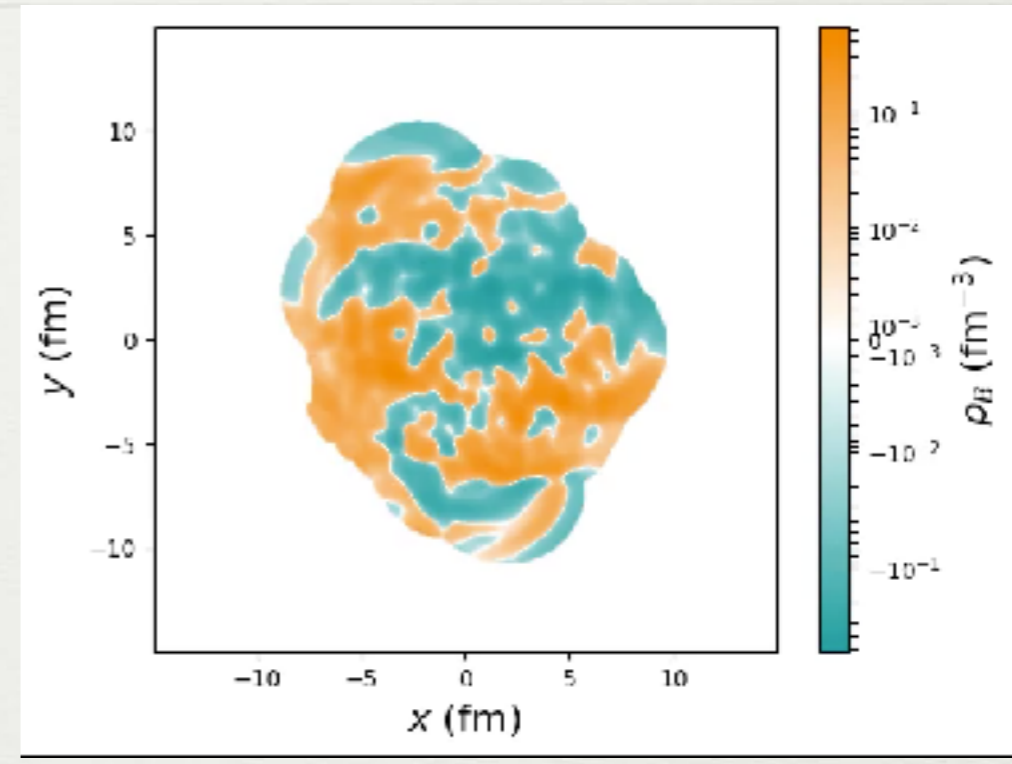
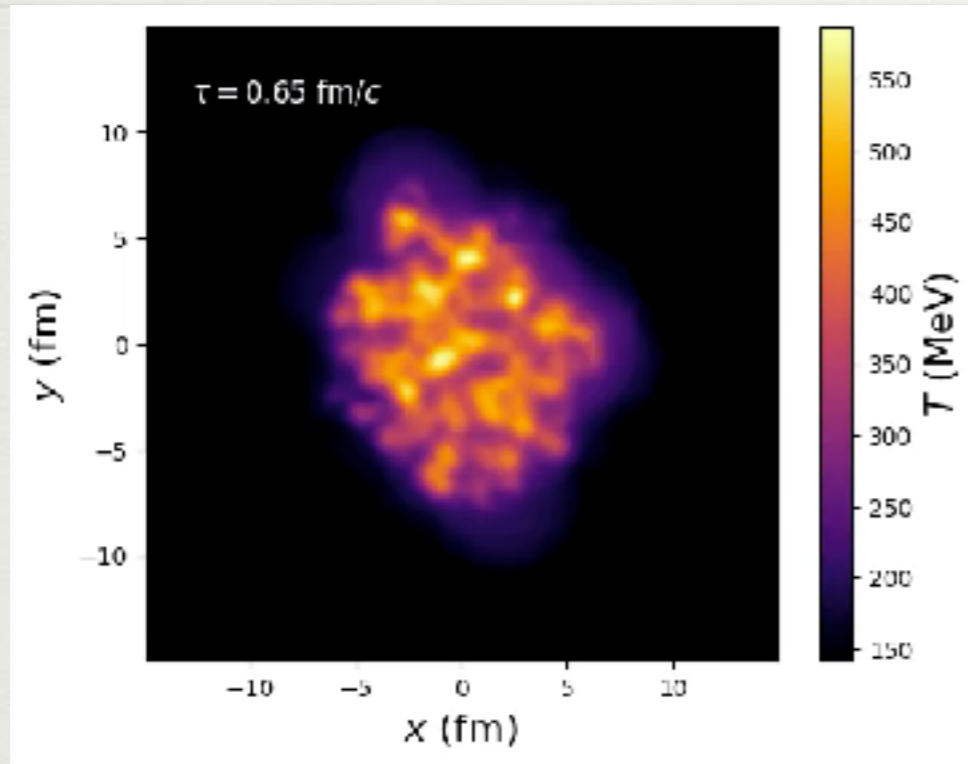
Green tanh non-conformal

Red conformal EOS

Lattice QCD EOS breaks down with μ_B/T so low T , high n_B hard to capture

Hydrodynamics with BSQ charges

Almaalol, Carzon, Dore, Espino, Mroczek, Plumberg, Salinas San Martin, Spsychalla, Sievert, JNH to appear soon

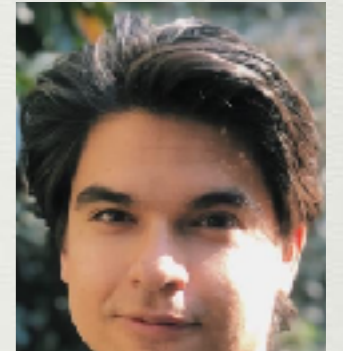
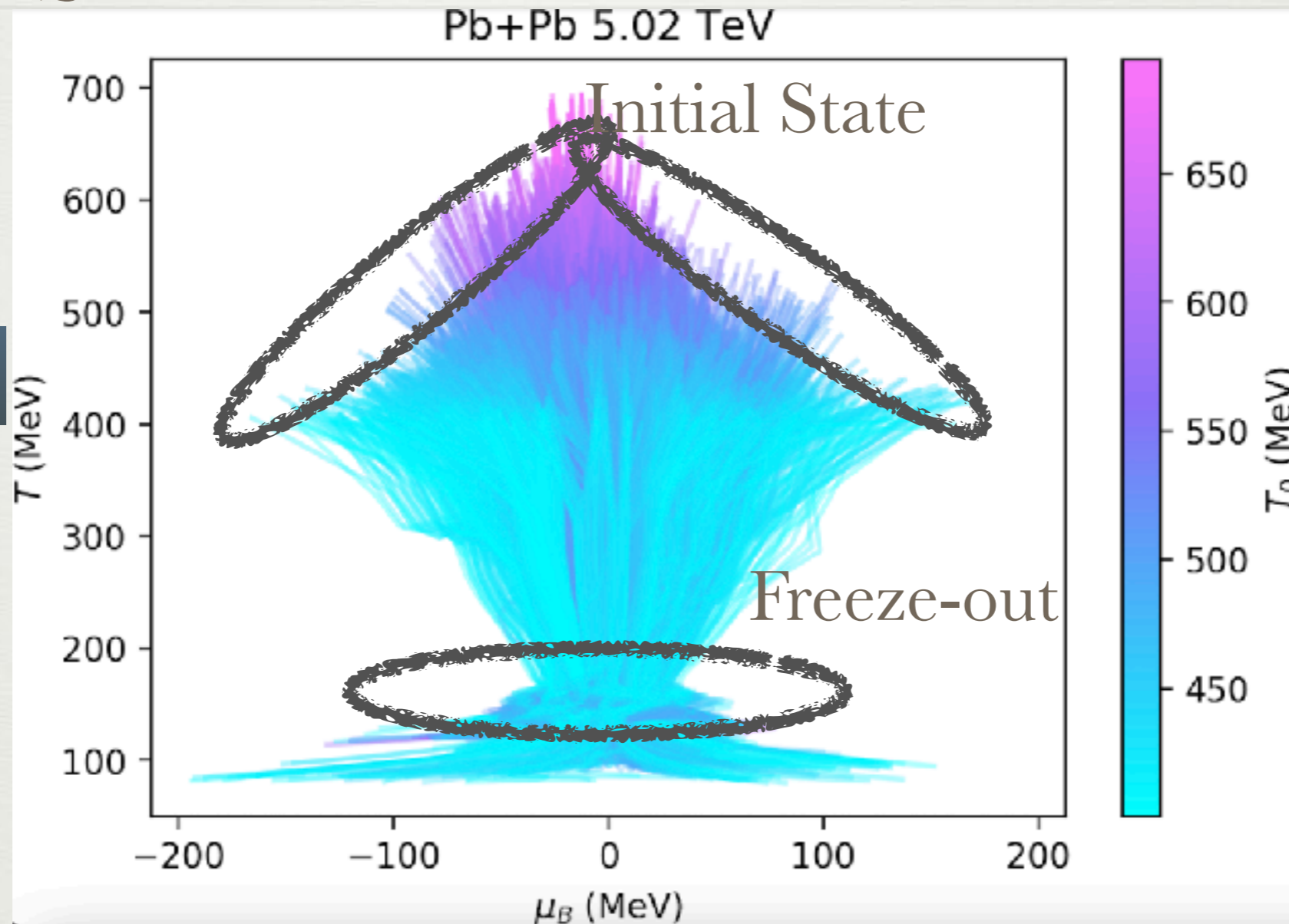


BSQ fluctuations large at the LHC

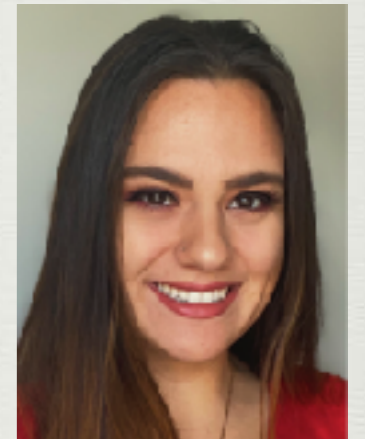
$$\varepsilon + B + S + Q$$



Christopher Plumberg
Pepperoni University



Jordi Salinas
San Martín
PhD student



Debora Mroczek
PhD Student

BSQ fluctuations just from $g \rightarrow q\bar{q}$.
net- $B=0$, local baryon number fluctuations large

Almaalol, Carzon, Dore, Espino, Mroczek, Plumberg, Salinas San Martin, Spsychalla, Sievert, JNH to appear soon

End goal



Calculate γ_{ijk} charge correlations for all BSQ possibilities

Ideas for other quantities that are sensitive to this physics as well

Holography wish list

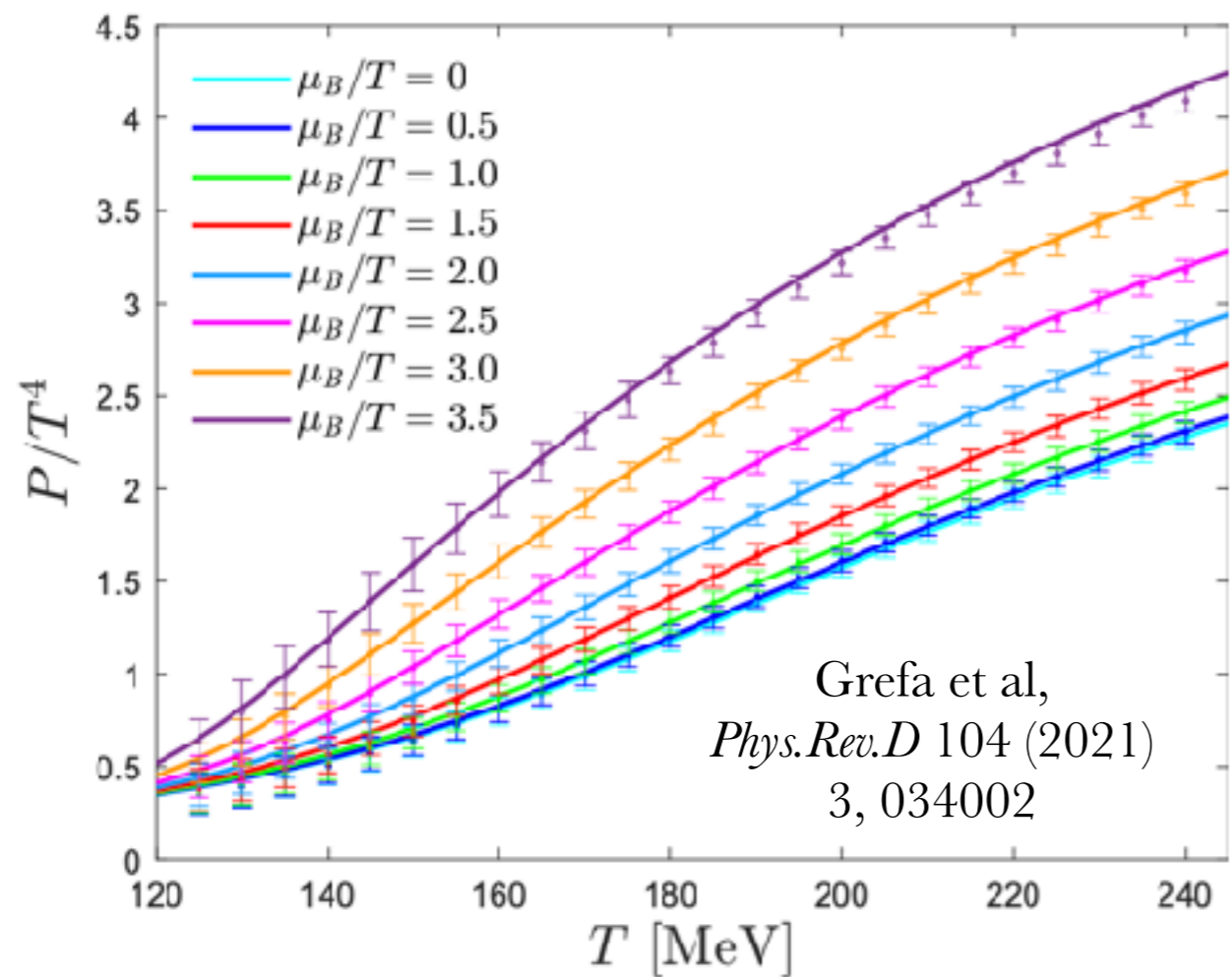
- Equation of state with BSQ conserved charges (looks like QCD)
- Transport coefficients with BSQ conserved charges: shear, bulk, diffusion matrix (6 independent terms)
- Initial state $T^{\mu\nu}$ and B^μ, S^μ, Q^μ for the 3 out-of-equilibrium currents

Summary

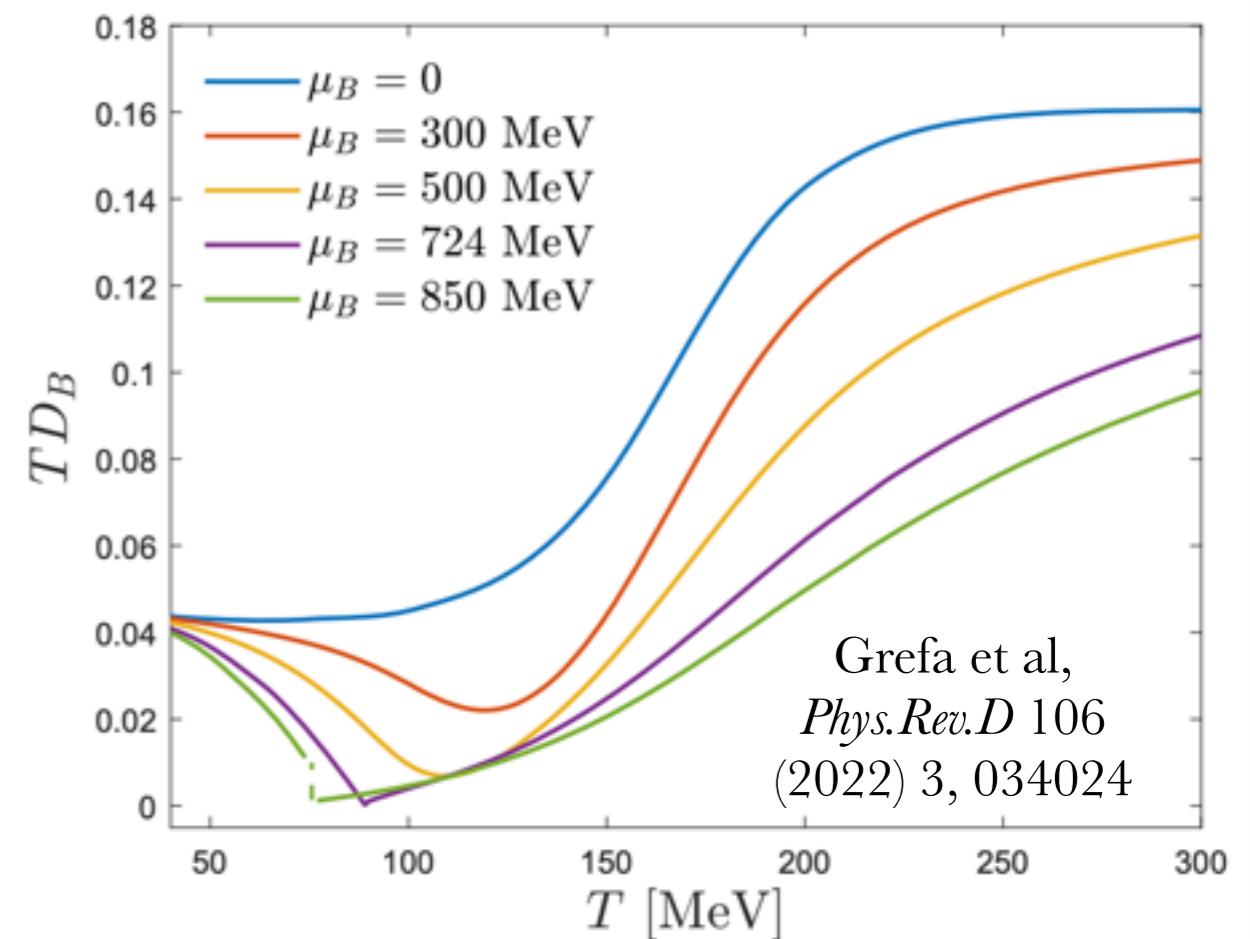
- ICCING can help to quantify background effects for the CME observables
- BSQ relativistic viscous hydrodynamics CCAKE allows for direct comparisons to data
- Preliminary results find consequences for flow
- Future: isobar study with ICCING+CCAKE

How can holography help?

Non-conformal EOS



Transport coefficients at finite μ_B

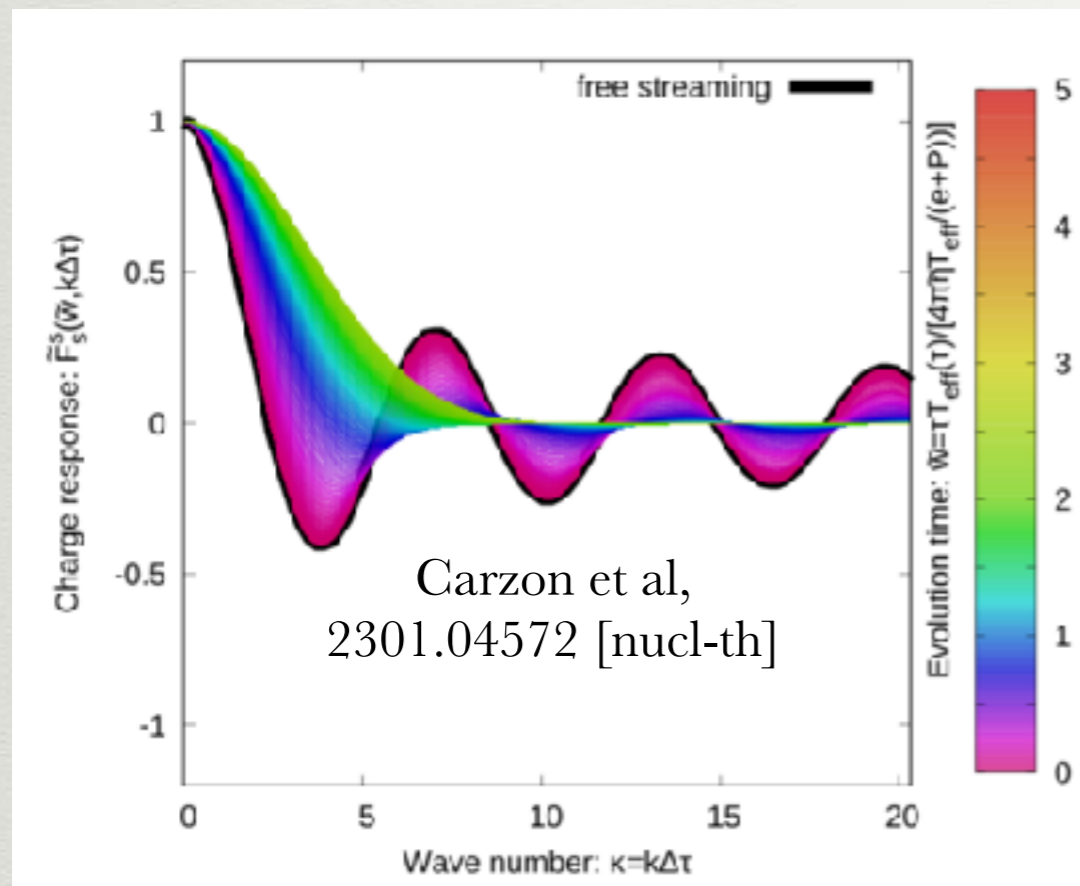


However, need 3 conserved charges. Currently only includes 1

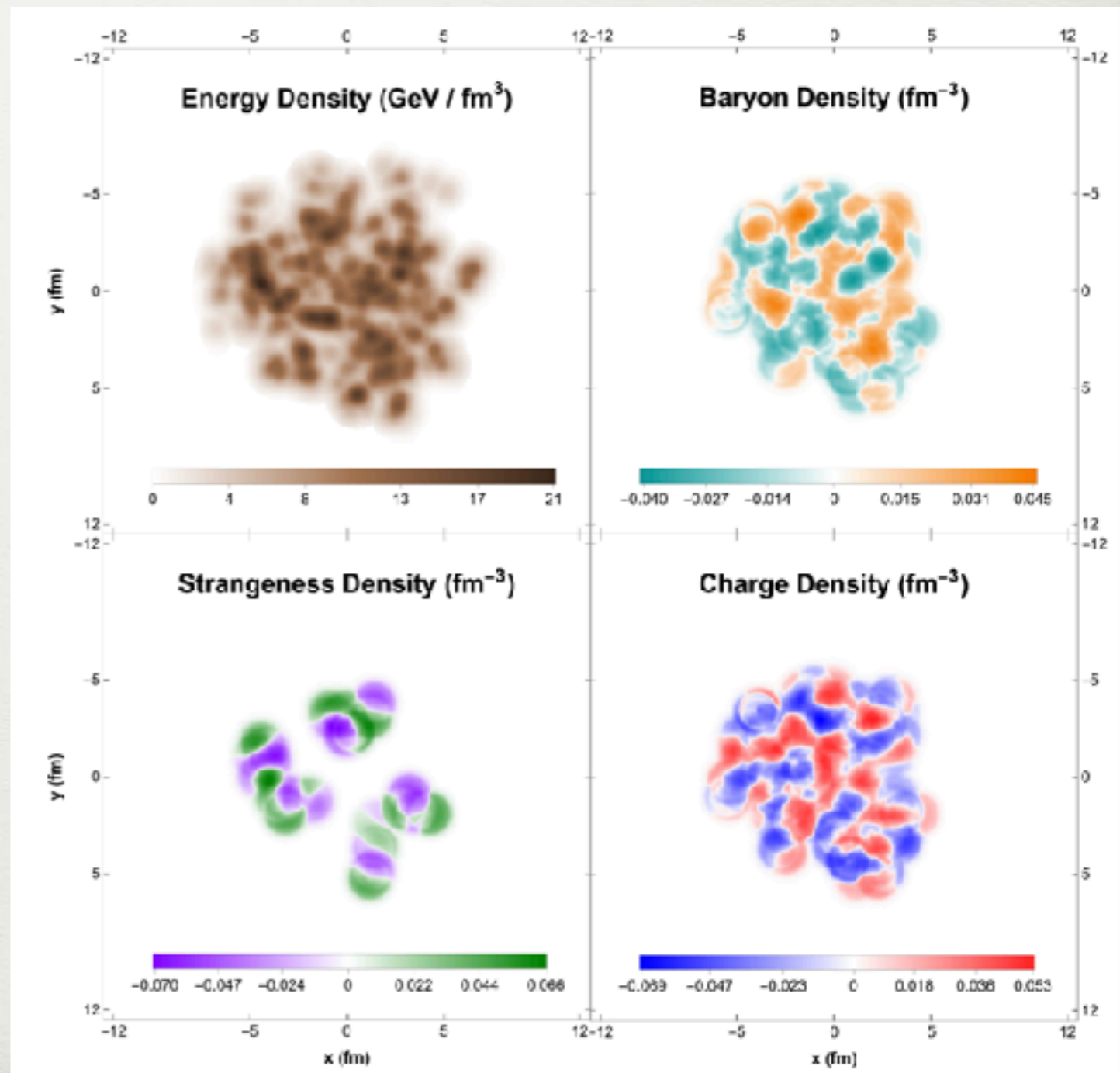
Pre-equilibrium expansion of BSQ charge

Energy and charge
Green's functions

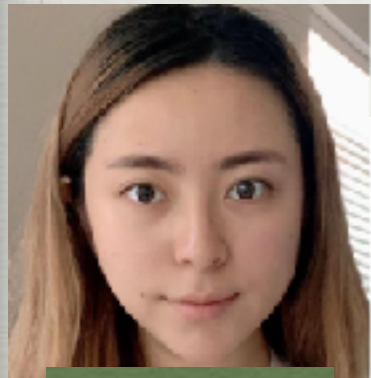
ICCING evolution after $\tau = 1 \text{ fm}$



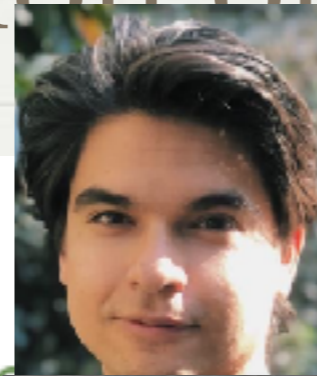
Relies on a
perturbative approach,
limits $q\bar{q}$ production



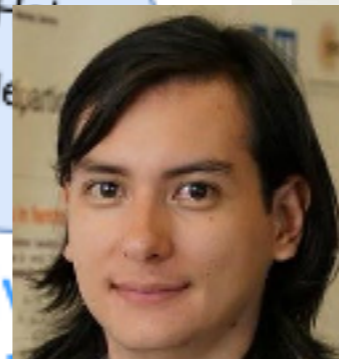
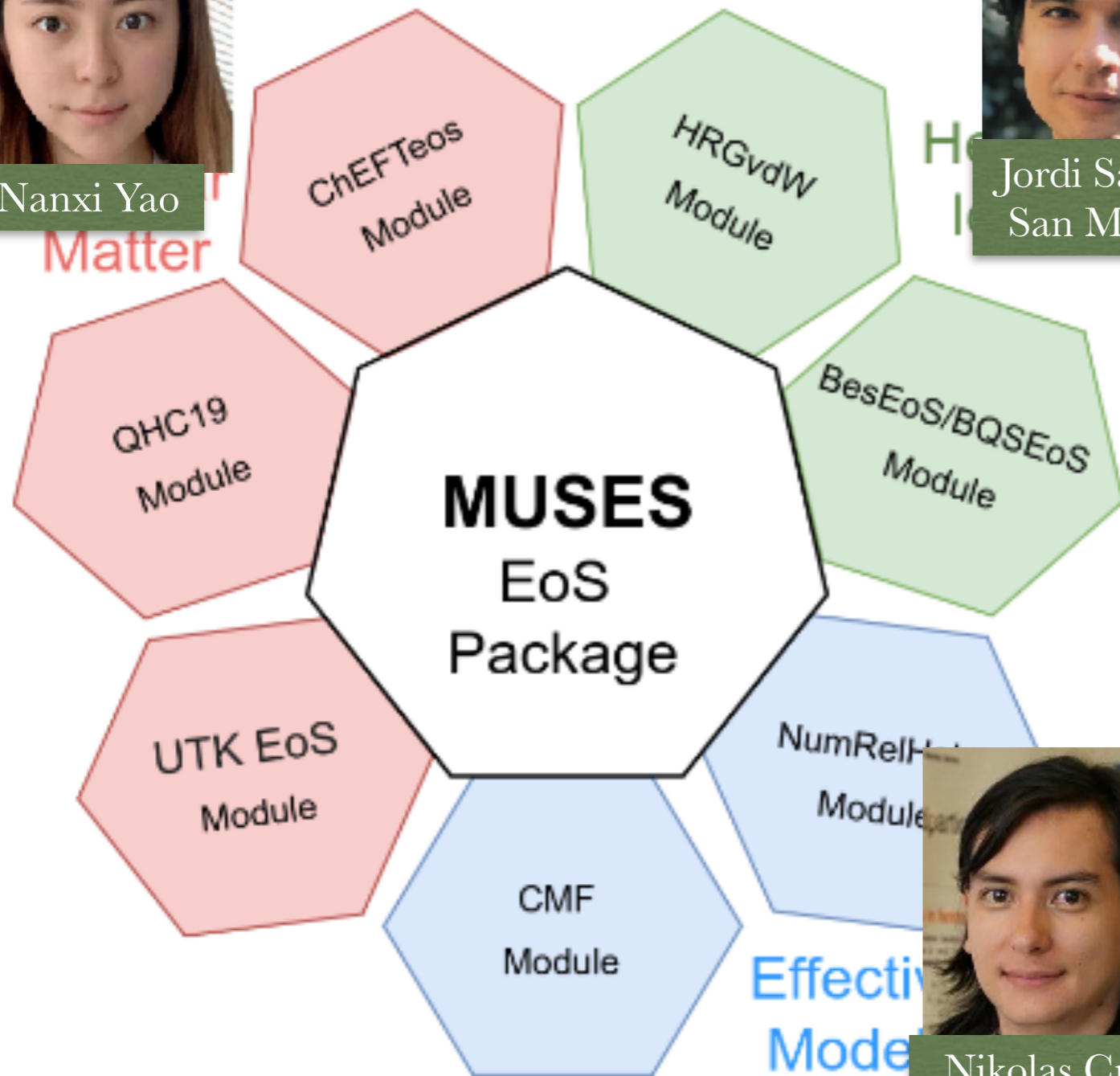
MUSES: Modular Unified Solver of the Equation of State



Nanxi Yao
Matter



Jordi Salinas
San Martín



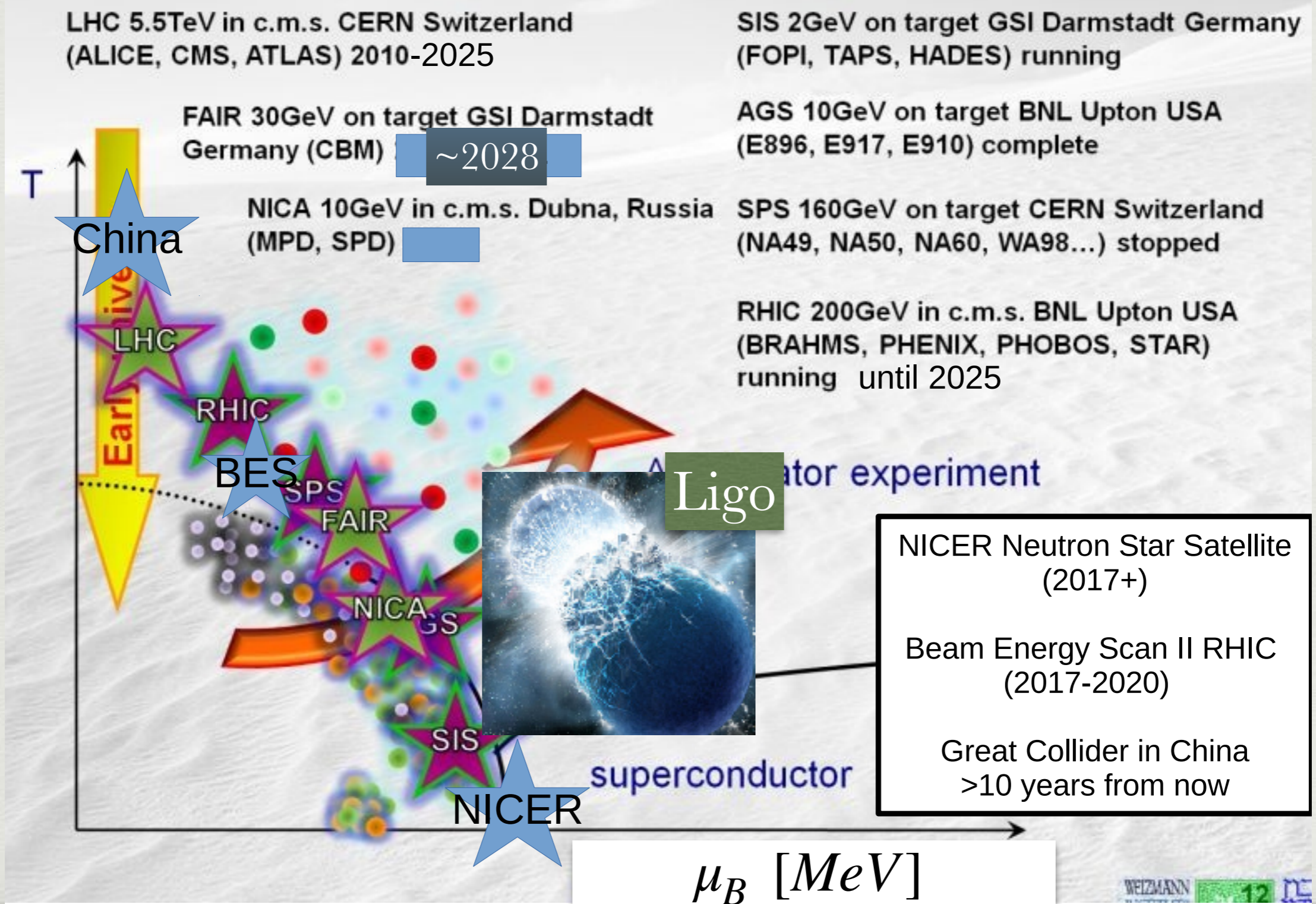
Nikolas Cruz
Camacho



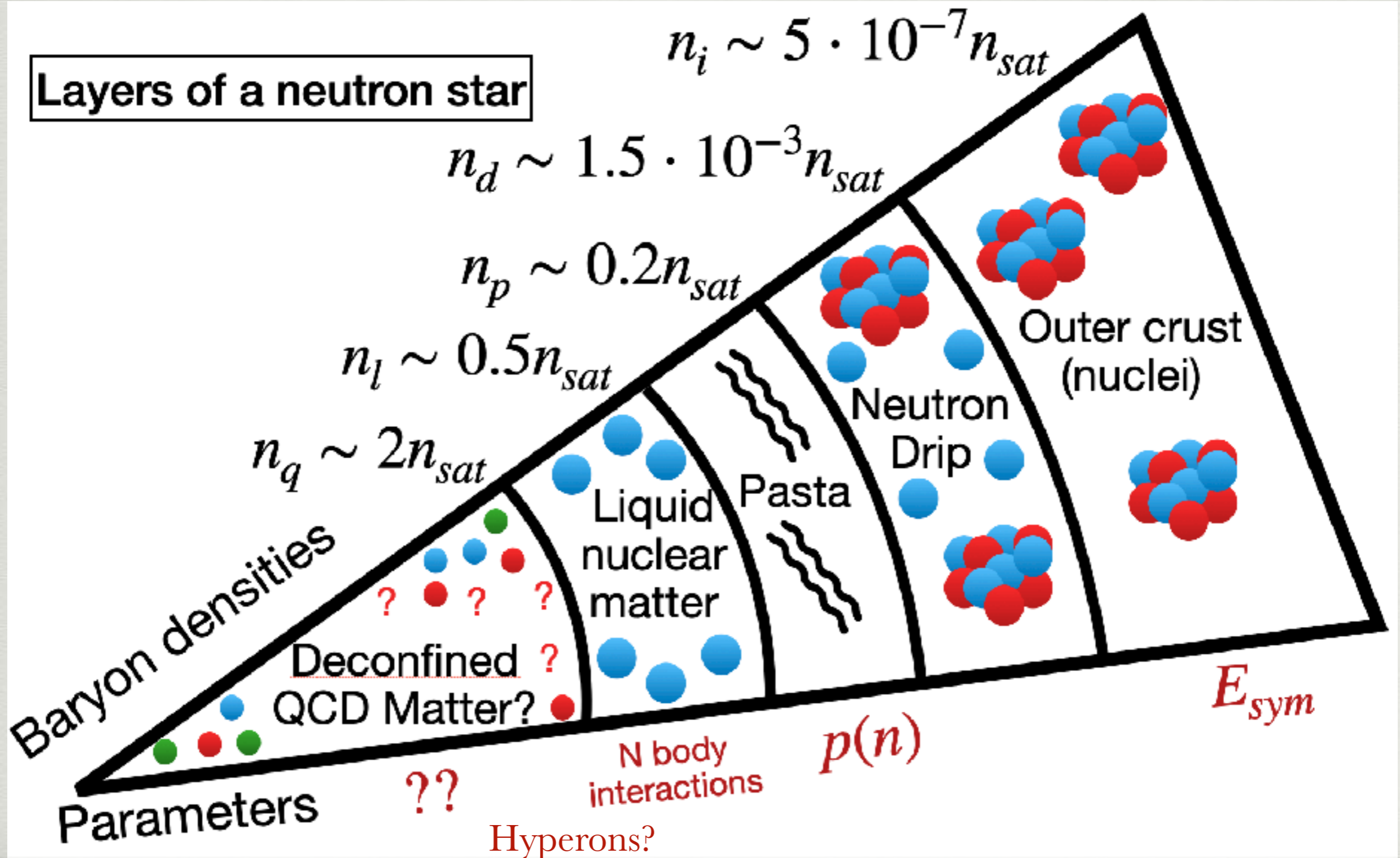
Started Fall 2021

16+ institutions, combines nuclear/computer science/gravity/astro/particle

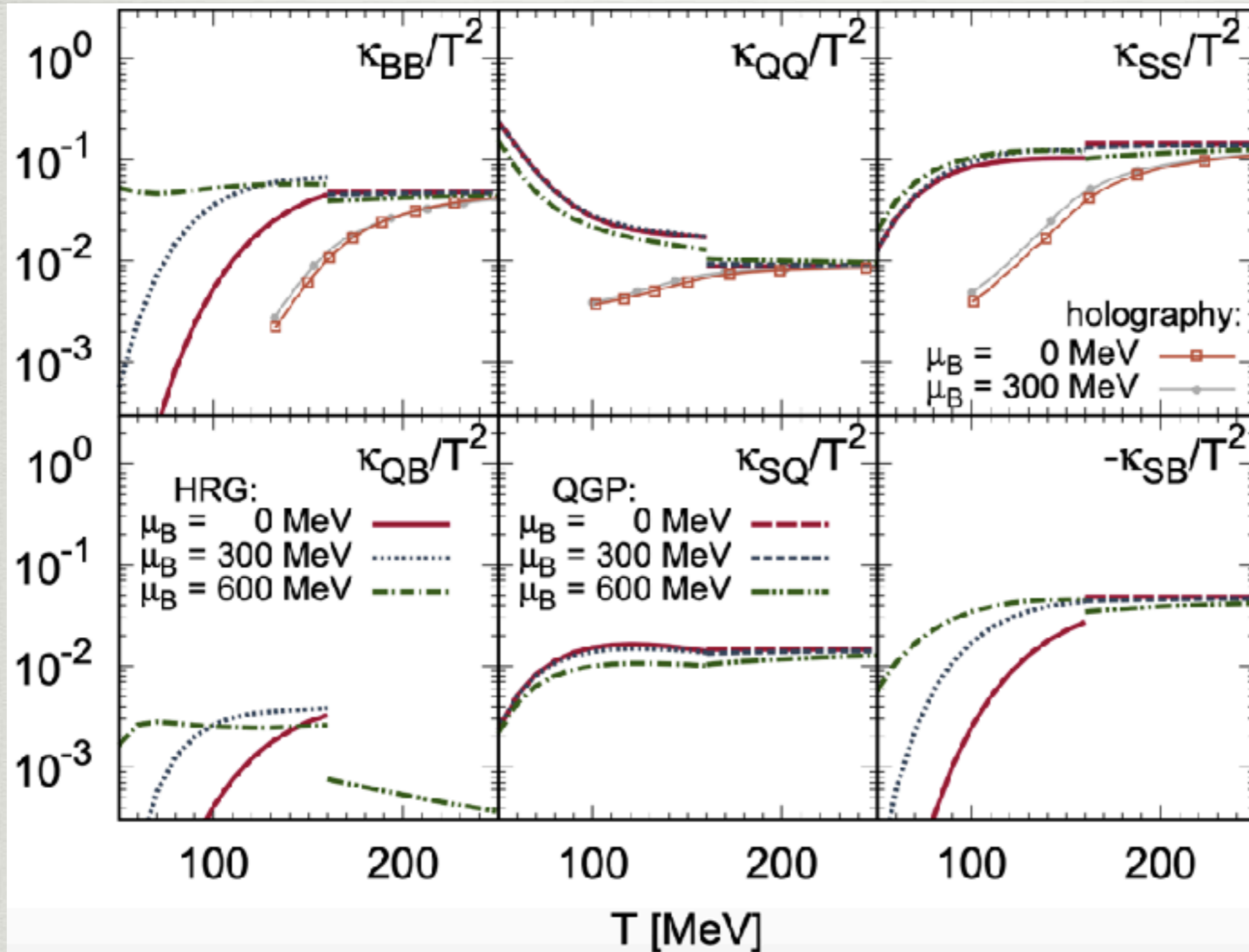
Mapping the QCD phase diagram



Degrees of freedom by density



Dynamics: BSQ Diffusion Matrix



Weakly Coupled

*Greif, Fotakis, Denicol, Greiner
Phys.Rev.Lett. 120 (2018) no.24,
242301*

Very Strongly
Coupled

*Rougemont, Critelli, JNH, Noronha,
Ratti Phys.Rev. D96 (2017) no.1,
014032*