

The butterfly effect in a holographic chiral system

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Lanzhou University



Holographic perspectives on chiral transport

ECT, March 13

Motivation

- Quantum chaos is associated with energy dynamics (holographic system)
- In chiral systems, energy transport through the CME

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- In chiral systems, energy transport through the CME

→ Any connection between “Quantum chaos” and “CME” ?

Quantum chaos

- One important aspect of thermal systems:

Atypicality → evolution → typicality

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initially similar (but orthogonal) states → evolve → to be quite different:

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Scrambling time is conjectured as $t \sim \beta \log S$

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[Sekino, Susskind 0808.2096]

[Maldacena, Shenker, Stanford 1503.01406]

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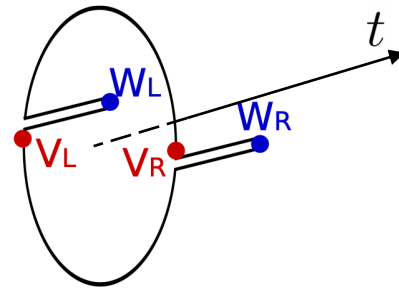
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Quantifying the butterfly effect

Exponential decrease of “Out-of-time-ordered correlators” (OTOC)

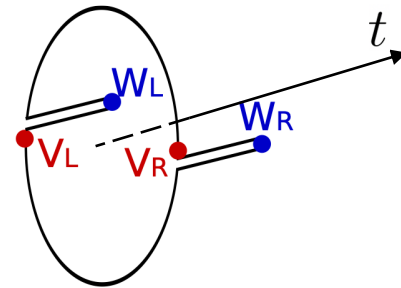
$$\text{OTOC} = F(t) = \langle \Psi | V_L W_R(t) V_R W_L(t) | \Psi \rangle$$



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- For large N holographic CFTs

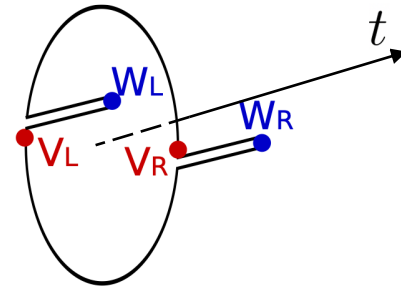
$$F(t) = f_0 - \frac{f_1}{N^2} \exp \frac{2\pi}{\beta} t + \mathcal{O}(N^{-4})$$

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Lyapunov exponent

Butterfly velocity

- The squared commutator

$$\begin{aligned} C(t_w, |x - y|) &= \text{tr} \{ \rho(\beta) [W_x(t_w), W_y]^\dagger [W_x(t_w), W_y] \} \\ &= 2 - 2 \text{Re} \langle TFD | W_y W_x(t_w) W_y W_x(t_w) | TFD \rangle \end{aligned}$$

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Butterfly speed

[Roberts, Shenker, Stanford, 1409.8180]

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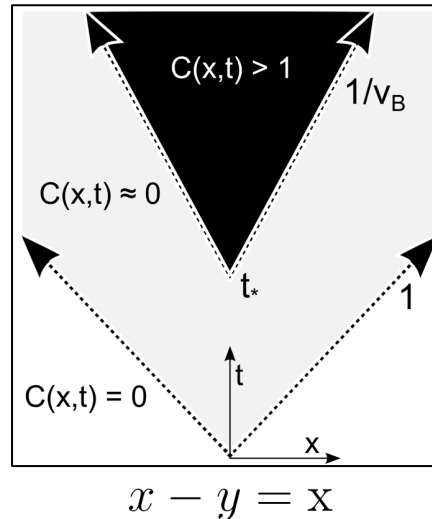
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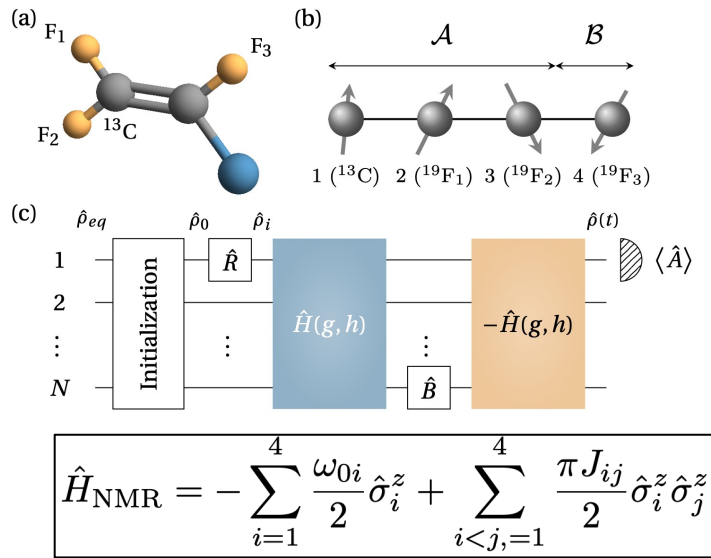
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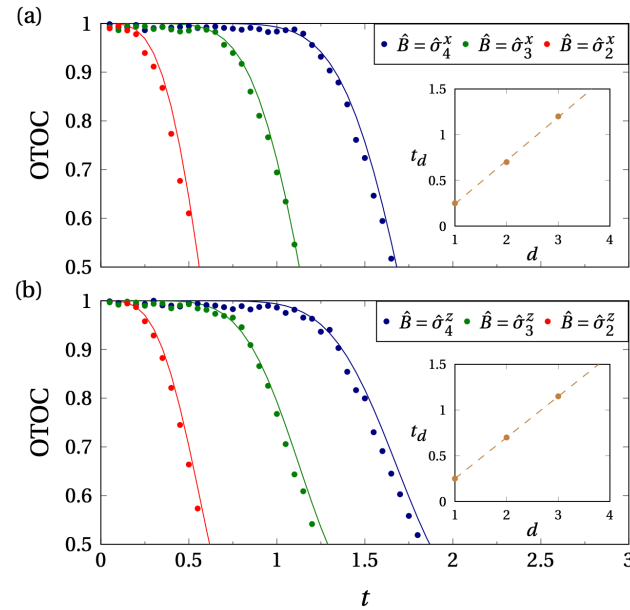


OTOC and v_B from experiment

Ising spin chain on a nuclear magnetic resonance (NMR) quantum simulator



[Li, Fan Wang, Ye, Zeng, Zhai, Peng, Du 1609.01246]



Several other experiments:

[Garttner, Bohnet, Safavi, Wall, Bollinger, Rey 1608.08938]

[Cao, Zhu, Del Campo 2111.12475]

[Swingle, Bentsen, Schleier-Smith, Hayden 1602.06271]

[...]

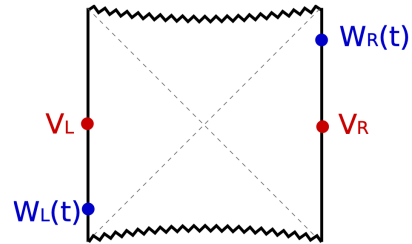
OTOC from eternal black hole

[Maldacena 0106112]

[Van Raamsdonk 1005.3035]

[Shenker, Stanford 1412.6987]

$$\langle V_L W_R(t) V_R W_L(t) \rangle$$



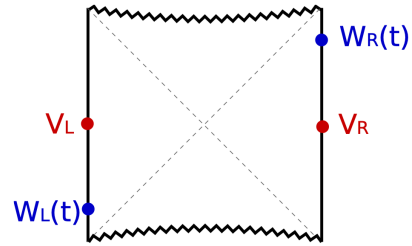
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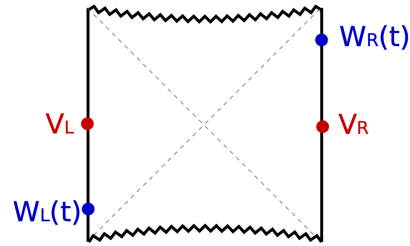
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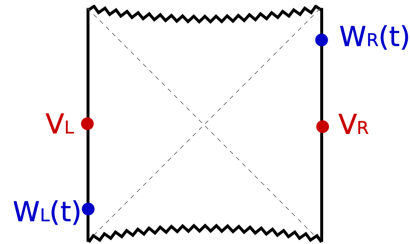
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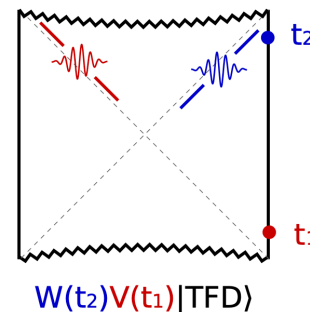
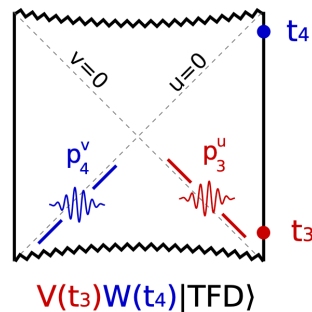
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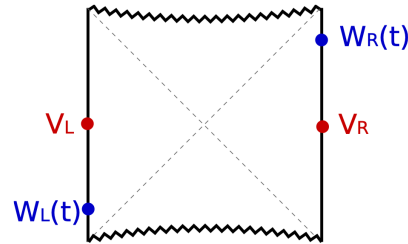
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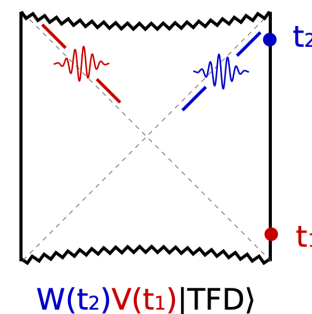
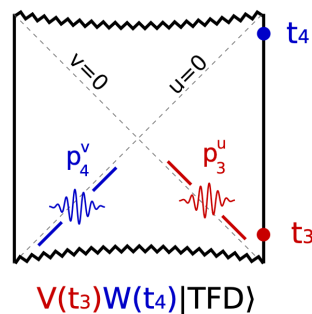
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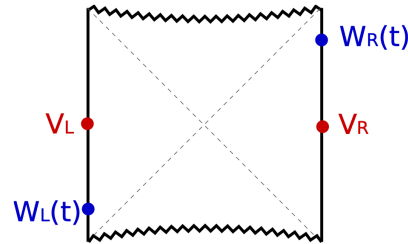
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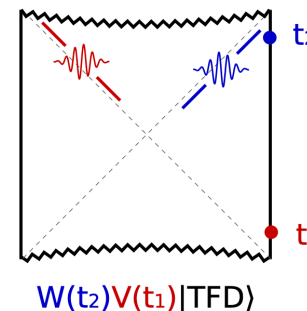
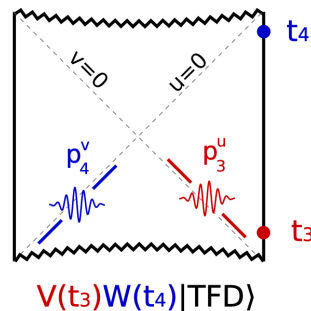
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$$|p_1^u, x_1; p_2^v, x_2\rangle_{out} \approx e^{i\delta(s,b)} |p_1^u, x_1; p_2^v, x_2\rangle_{in} + |\chi\rangle$$

Shock wave geometry in AdS

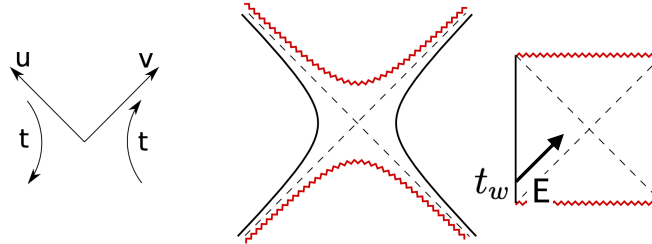
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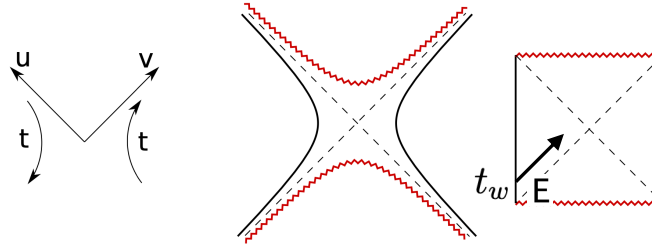
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$$T_{uu} = \frac{E}{\ell_{AdS}^{d+1}} e^{2\pi t_w / \beta} \delta(u) a_0(x)$$

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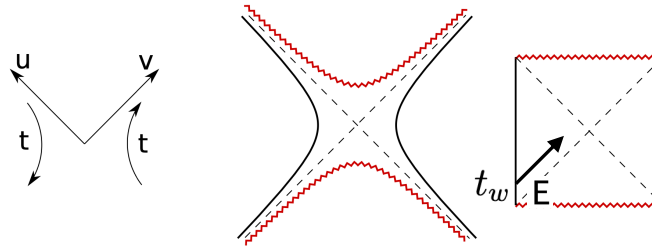
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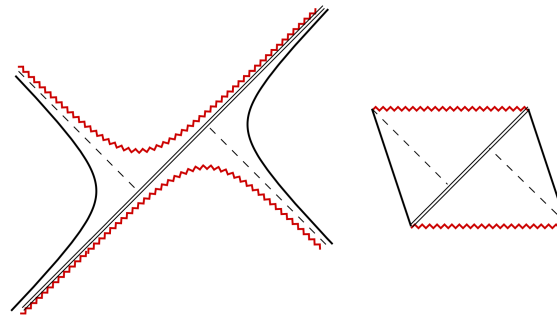
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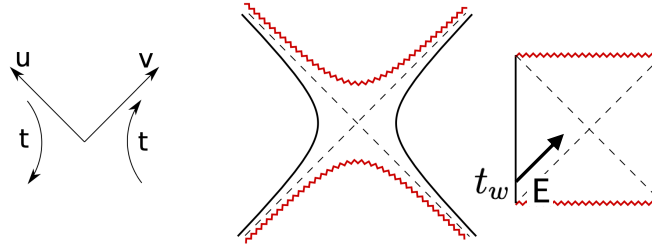
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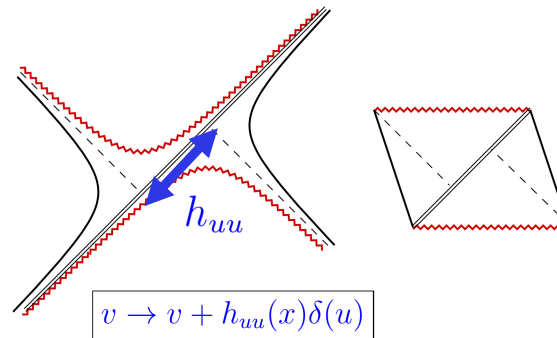
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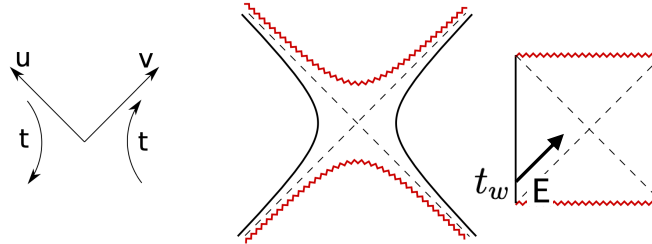
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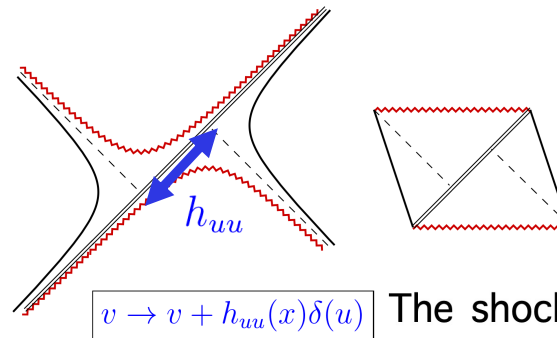
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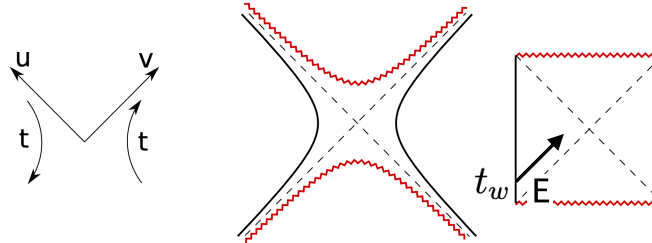
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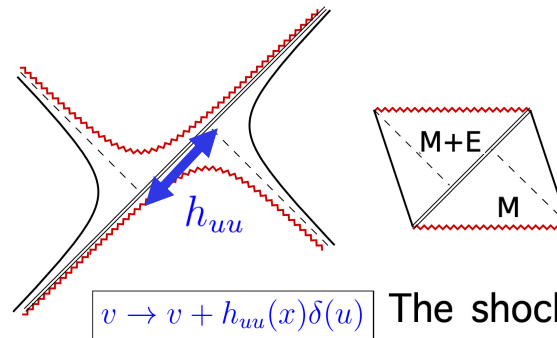
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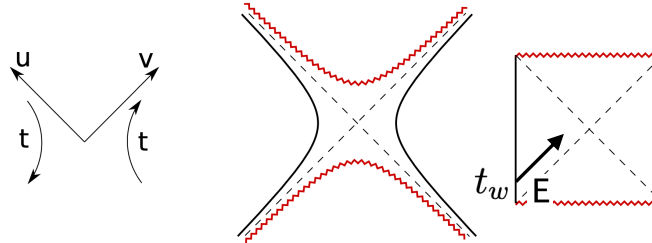
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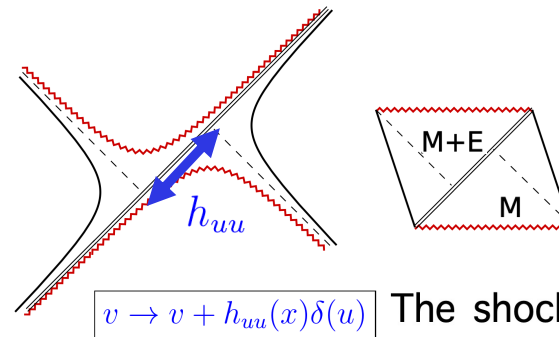
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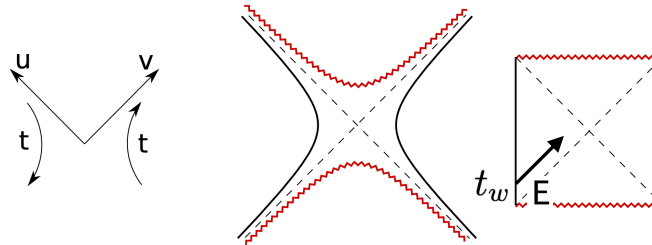
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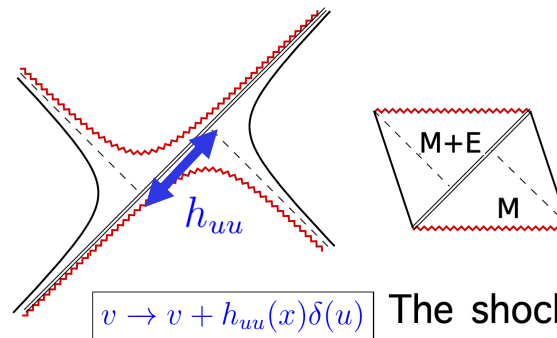
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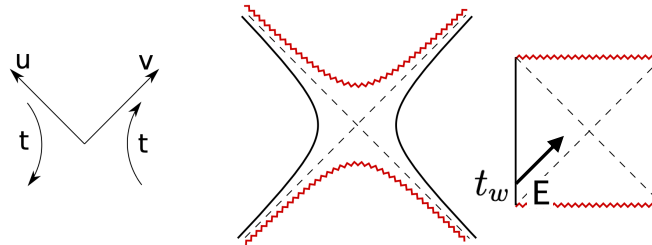
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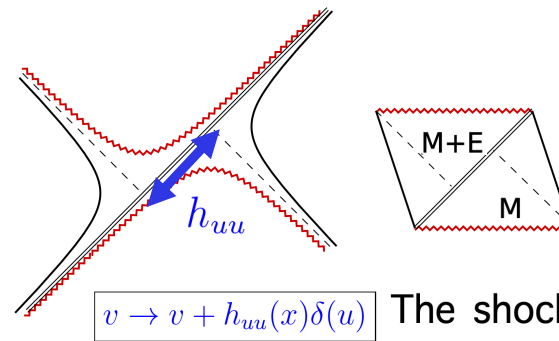
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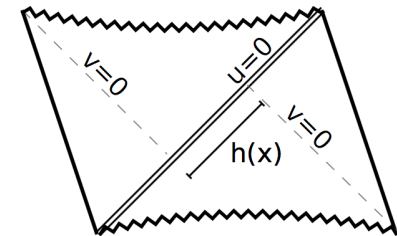
$$\delta(s, b) = S_{cl} = \frac{1}{2} \int d^{d+1}x \sqrt{-g} h_{uu} T^{uu} \sim \frac{4\pi G_N}{r_0^{d-3}} E^2 e^{\frac{2\pi}{\beta} t_w} h_{uu}(x_{12})$$

[Shenker, Stanford 1412.6987]

Butterfly velocity from BH

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- Putting $T_{uu} = \frac{E}{\ell_{AdS}^{d+1}} e^{2\pi t_w/\beta} \delta(u) a_0(x)$ on the RHS of Einstein equations
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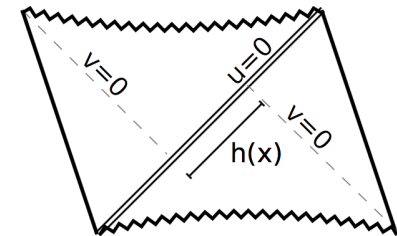


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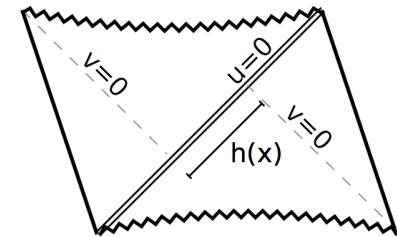


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- Which for $|x| \gg 1$ has the solution

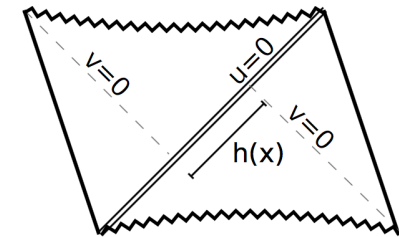
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Butterfly velocity from BH

[Shenker, Stanford 1306.0622]
 [Shenker, Stanford 1412.6987]

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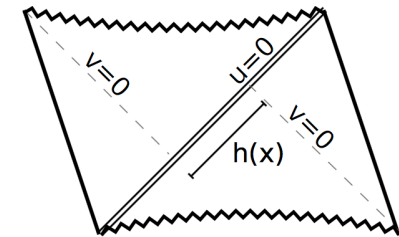
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By comparison:

➤ One finds

$$\lambda = 2\pi T$$

$$v_B = \sqrt{\frac{D-1}{2(D-2)}}$$

Our goal

We want to calculate v_B

in a holographic chiral system in the presence of a B

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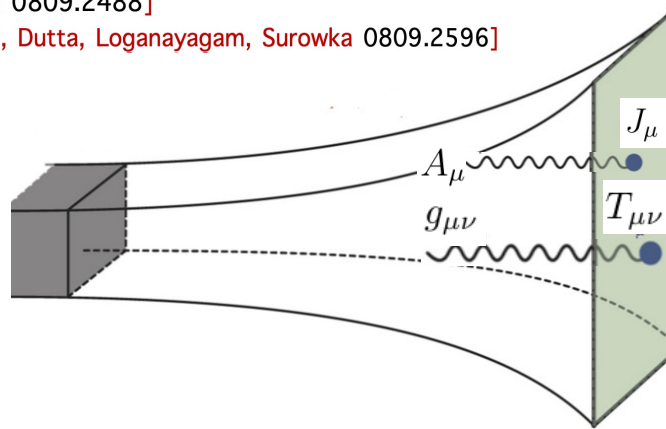
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[Erdmenger, Haack, Kaminski, Yarom 0809.2488]

[Banerjee, Bhattacharya, Bhattacharya, Dutta, Loganayagam, Surowka 0809.2596]



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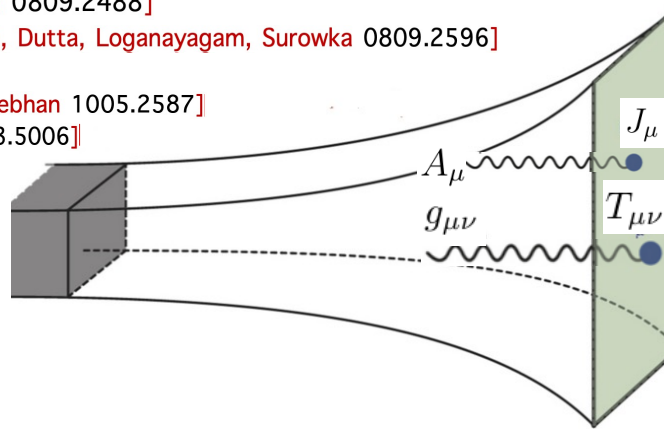
[Banerjee, Bhattacharya, Bhattacharya, Dutta, Loganayagam, Surowka 0809.2596]

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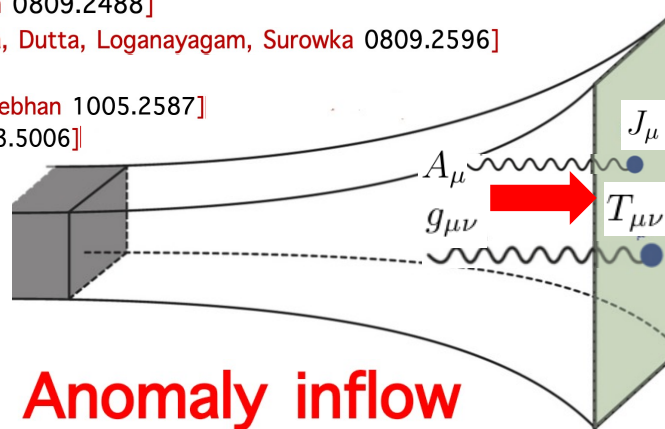
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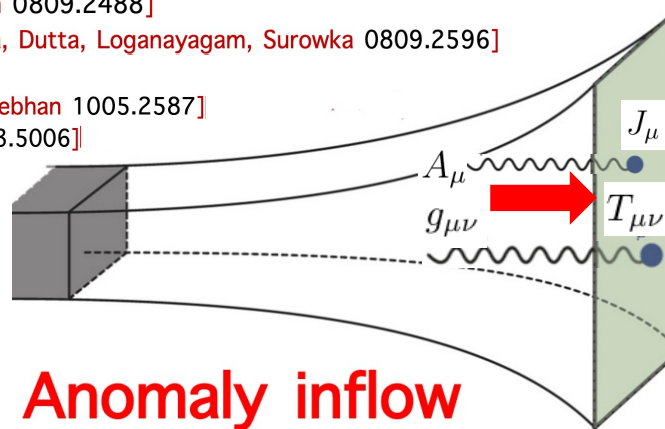
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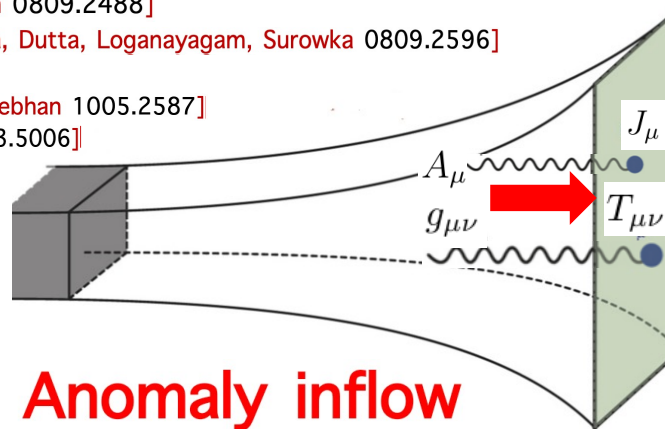
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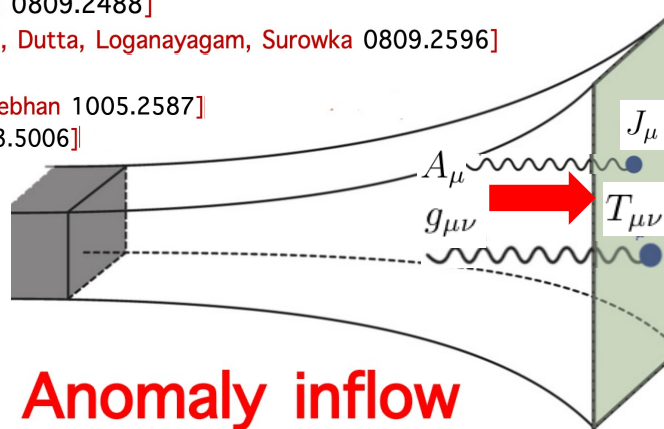
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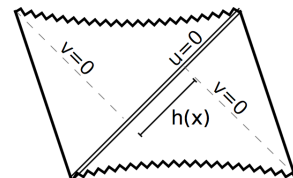
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- What we need to find: the function $h(x)$ on this background



BH solution

EOM:

$$\nabla_{\nu} F^{\nu\mu} + \frac{\kappa}{4} \epsilon^{\mu\nu\rho\alpha\beta} F_{\nu\rho} F_{\alpha\beta} = 0$$
$$R_{\mu\nu} + 4g_{\mu\nu} + \frac{1}{3} F^{\alpha\beta} F_{\alpha\beta} g_{\mu\nu} + 2F_{\mu\rho} F_{\nu}^{\rho} = 0$$

Parametrizing the Solution:

$$ds^2 = \frac{dr^2}{f(r)} - f(r)dt^2 + e^{2W_T(r)}(dx_1^2 + dx_2^2) + e^{2W_L(r)}(dx_3 + C(r)dt)^2$$

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- For example, near the horizon, $f(r)$ is given by

[NA, Tabatabaei 1910.13696]

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We perform calculations fully analytically

in a double expansion over ν and b .

Shockwave geometry

Applying $V \rightarrow V + h(x)\delta(U)$

$$ds_{\text{future}}^2 = ds_{\text{past}}^2 - A(UV)h(x)\delta(U) dU^2 - \frac{D(UV)}{V}h(x)\delta(U) dU dx_3$$
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$$\left(\partial_{\parallel}^2 + \mathbf{q}^2 \partial_{\perp}^2 + 2 \mathbf{p} \vec{b} \cdot \vec{\partial} - m_0^2 \right) h(x) \sim \frac{2B_L(0)}{A(0)} E e^{\frac{1}{2} \tilde{f}'(r_h) t_w} \delta^3(\vec{x})$$

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- Compare it with $b=nu=0$

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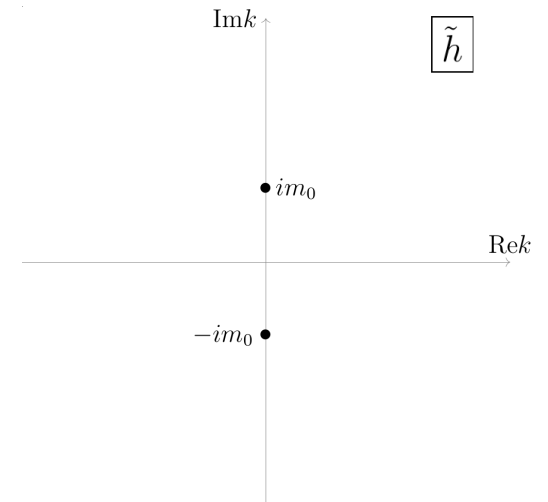
$$m_0^2 = (\pi T)^2 \left[6 + 36\nu^2 - \left(\frac{\pi^2}{6} - 1 \right) b^2 - \left(\pi^2 + \frac{92}{9} + 56\kappa^2(\log(2) - 1) \right) \nu^2 b^2 \right]$$

[NA, Tabatabaei 1910.13696]

- Compare it with $b=nu=0$

$$(-\partial_i \partial_i + m_0^2) h(x) = \frac{16\pi G_N}{A(0) \ell_{AdS}^{d-1}} E e^{\frac{2\pi}{\beta} t_w} a_0(x)$$

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Shockwave geometry

Applying $V \rightarrow V + h(x)\delta(U)$

$$ds_{\text{future}}^2 = ds_{\text{past}}^2 - A(UV)h(x)\delta(U) dU^2 - \frac{D(UV)}{V}h(x)\delta(U) dU dx_3$$

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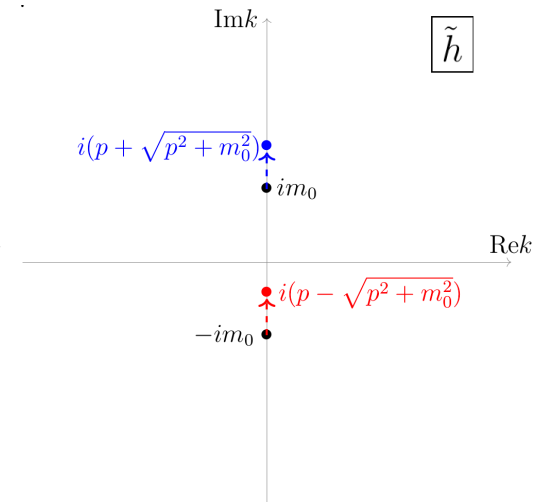
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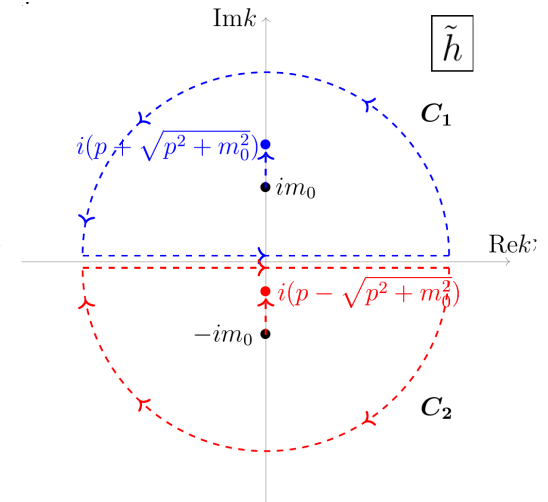
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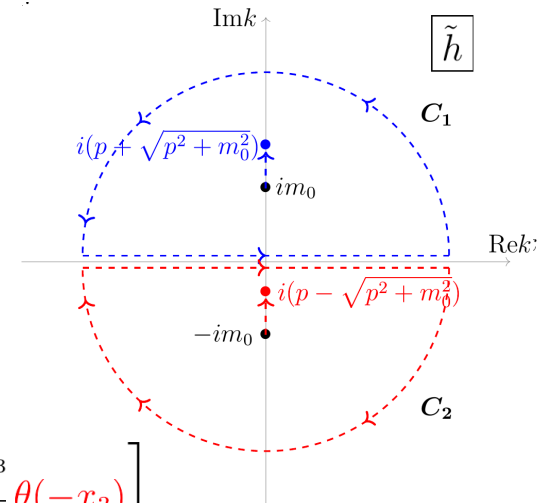
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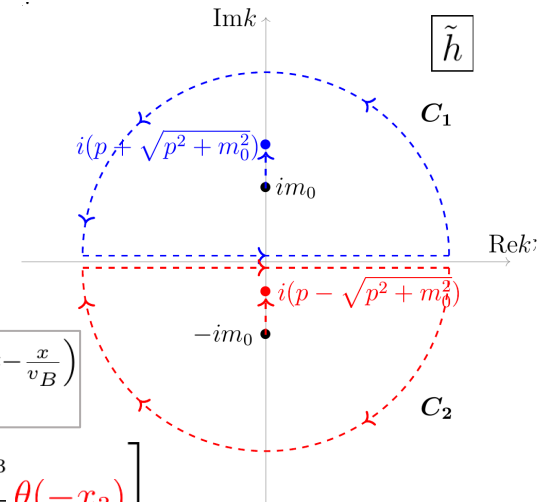
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$$C(t, x) = 2 - 2 \langle W(t, \vec{x}) W(0) W(t, \vec{x}) W(0) \rangle_{\beta} \sim \frac{1}{N^2} e^{\lambda \left(t - \frac{x}{v_B} \right)}$$

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Asymmetry

Two different velocities

$$x_3 > 0 : \quad v_B^{L_1} = \frac{2\pi T}{m_0^2} \left(\sqrt{\mathfrak{p}^2 + m_0^2} - \mathfrak{p} \right)$$

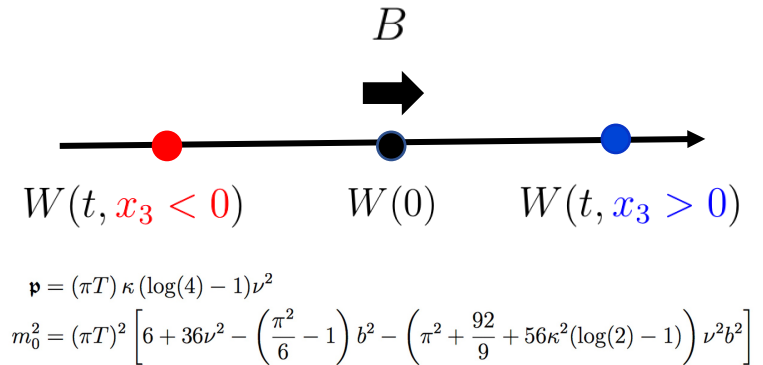
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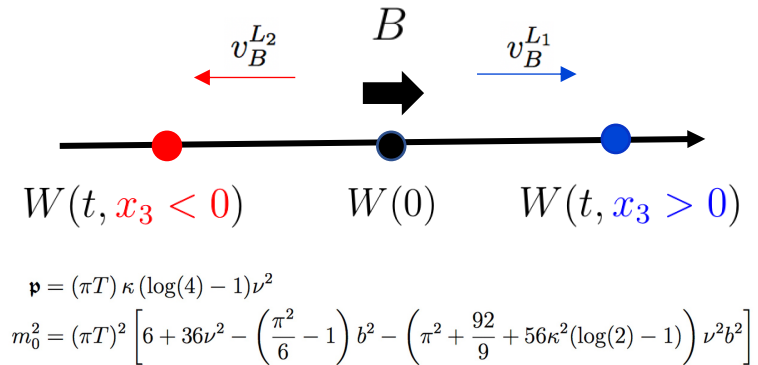


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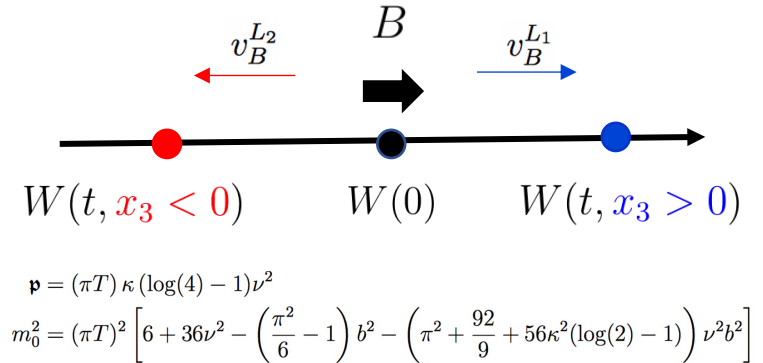


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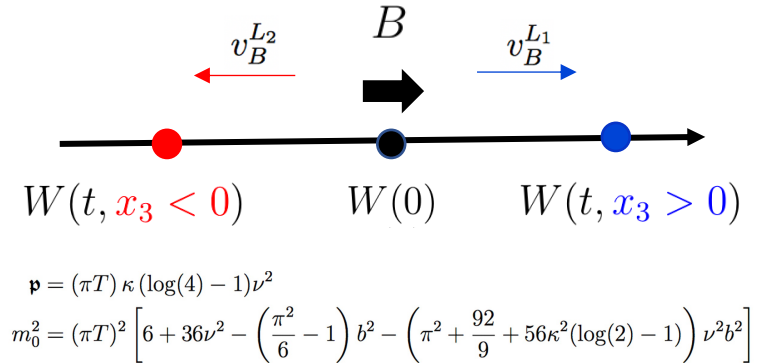
$$v_B^{L1,2} = \pm \sqrt{\frac{2}{3}} \left(1 - \frac{\mu^2}{3(\pi T)^2} \right) - \frac{2}{3} \kappa (\log(4) - 1) \frac{\mu^2 B}{(\pi T)^4} \pm \left(\frac{\pi^2 - 6}{36\sqrt{6}} - \frac{\pi^2 + 18(-4\kappa^2(\log(4) - 2))}{108\sqrt{6}} \frac{\mu^2}{(\pi T)^2} \right) \frac{B^2}{(\pi T)^4}$$

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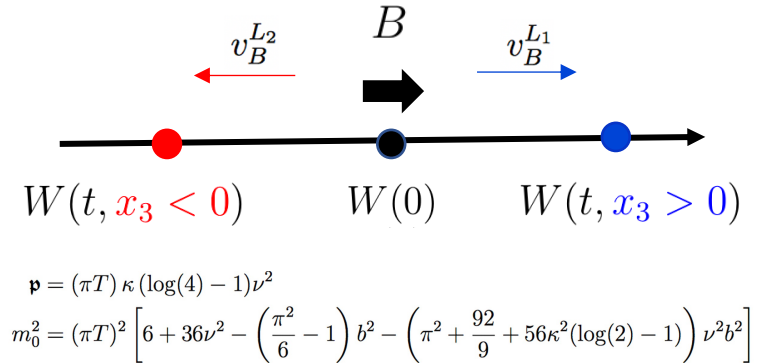
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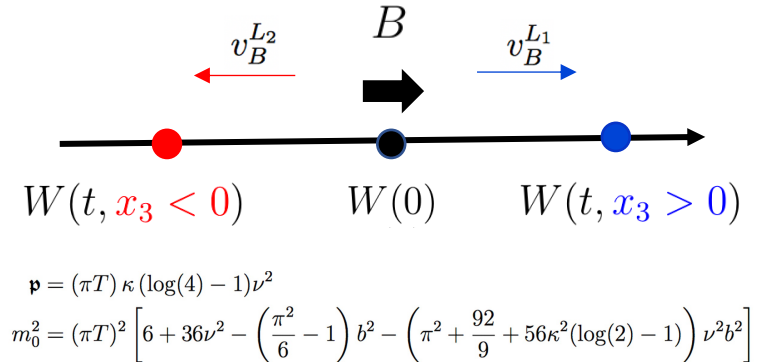
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This is actually a diagnostic of the chiral anomaly. [NA, Tabatabaei 1910.13696]

Connection to Chiral Transport

Hydrodynamic spectrum in a chiral system

Constitutive relations:

$$T^{\mu\nu} = wu^\mu u^\nu + pg^{\mu\nu} + \sigma_\epsilon^{\mathcal{B}}(u^\mu B^\nu + u^\nu B^\mu) + \sigma_\epsilon^{\mathcal{V}}(u^\mu \omega^\nu + u^\nu \omega^\mu),$$

$$J^\mu = nu^\mu + \sigma^{\mathcal{B}} B^\mu + \sigma^{\mathcal{V}} \omega^\mu,$$

[Neiman, Oz 1011.5107]

[Kharzeev, Yee 1105.6360]

[Banerjee, Bhattacharya, Bhattacharya, Jain, Minwalla, Sharma 1203.3544]

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[Ammon, Kaminski, Koirala, Leiber, Wu 1701.0565]

[NA, Naderi, Taghinavaz 1712.06175]

[Ammon, Greiner, Hernandez, Kaminski, Koirala, Leiber, Wu 2102.09183]

Hydro modes:

$$\omega = \pm \frac{nB}{w} + \frac{nB}{w^2} \sigma_\epsilon^{\mathcal{V}} k$$

$$\omega = \frac{B}{w} \frac{1}{[\beta, \alpha]} \left(w(\alpha_1 \partial_\mu + \alpha_2 \partial_T) \sigma^{\mathcal{B}} - n(\alpha_1 \partial_\mu + \alpha_2 \partial_T) \sigma_\epsilon^{\mathcal{B}} \right) k$$

$$\omega = \pm c_s k + \frac{B}{2w} \frac{[\gamma, \alpha]}{[\beta, \alpha]} \left(-1 + \frac{(n\alpha_2 - w\beta_2) \partial_T - (n\alpha_1 - w\beta_1) \partial_\mu}{n[\gamma, \alpha] - w[\gamma, \beta]} \right) \sigma^{\mathcal{B}} k$$

$$+ \frac{B}{w} \left(1 - \frac{[\gamma, \beta]}{[\alpha, \beta]} \right) \sigma_\epsilon^{\mathcal{B}} k + \frac{B}{2w} \frac{[\gamma, \alpha]}{[\beta, \alpha]} \left(\frac{(n\alpha_1 - w\beta_1) \partial_\mu - (n\alpha_2 - w\beta_2) \partial_T}{n[\gamma, \alpha] - w[\gamma, \beta]} \right) \sigma_\epsilon^{\mathcal{B}} k$$

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$$\alpha_1 = \left(\frac{\partial \epsilon}{\partial T} \right)_\mu$$

$$\alpha_2 = \left(\frac{\partial \epsilon}{\partial \mu} \right)_T$$

$$\beta_1 = \left(\frac{\partial n}{\partial T} \right)_\mu$$

$$\beta_2 = \left(\frac{\partial n}{\partial \mu} \right)_T$$

$$\gamma_1 = \left(\frac{\partial p}{\partial T} \right)_\mu$$

$$\gamma_2 = \left(\frac{\partial p}{\partial \mu} \right)_T$$

Special case: holographic chiral system

Thermodynamics: [NA, Ghazi, Taghinavaz, Tavakol, 1812.11310]

$$\begin{aligned}\epsilon &= \frac{N_c^2}{8\pi^2} (3(\pi T)^4 + 12(\pi T)^2 \mu^2 + 8\mu^4) + \frac{N_c^2 B^2}{4\pi^2} \left((1 - \log(\frac{\pi T}{\Delta})) - \frac{2}{3} \frac{\mu^2}{\pi T^2} (8 \log(2) - 3) \right) \\ p &= \frac{N_c^2}{24\pi^2} (3(\pi T)^4 + 12(\pi T)^2 \mu^2 + 8\mu^4) + \frac{N_c^2 B^2}{4\pi^2} \left(\log(\frac{\pi T}{\Delta}) + \frac{2}{3} \frac{\mu^2}{\pi T^2} (8 \log(2) - 3) \right) \\ n &= \frac{N_c^2}{3\pi^2} (3(\pi T)^2 \mu + 4\mu^3) + \frac{N_c^2 B^2}{3\pi^2} \frac{\mu}{(\pi T)^2} (8 \log(2) - 3)\end{aligned}$$

Transport coefficients: [Landsteiner, Megias, Pena-Benitez 1207.5808]

$$\begin{aligned}\sigma^B &= -\frac{\kappa}{2\pi G_5} \mu = \frac{2N_c^2}{\pi^2 \sqrt{3}} \mu \\ \sigma^V = \sigma_\epsilon^B &= -\frac{\kappa}{2\pi G_5} \mu^2 = \frac{N_c^2}{\pi^2 \sqrt{3}} \mu^2 \\ \sigma_\epsilon^V &= -\frac{\kappa}{6\pi G_5} \mu^3 = \frac{2N_c^2}{3\pi^2 \sqrt{3}} \mu^3\end{aligned}$$

Special case: holographic chiral system

Thermodynamics: [NA, Ghazi, Taghinavaz, Tavakol, 1812.11310]

$$\begin{aligned}\epsilon &= \frac{N_c^2}{8\pi^2} (3(\pi T)^4 + 12(\pi T)^2 \mu^2 + 8\mu^4) + \frac{N_c^2 B^2}{4\pi^2} \left((1 - \log(\frac{\pi T}{\Delta})) - \frac{2}{3} \frac{\mu^2}{\pi T^2} (8 \log(2) - 3) \right) \\ p &= \frac{N_c^2}{24\pi^2} (3(\pi T)^4 + 12(\pi T)^2 \mu^2 + 8\mu^4) + \frac{N_c^2 B^2}{4\pi^2} \left(\log(\frac{\pi T}{\Delta}) + \frac{2}{3} \frac{\mu^2}{\pi T^2} (8 \log(2) - 3) \right) \\ n &= \frac{N_c^2}{3\pi^2} (3(\pi T)^2 \mu + 4\mu^3) + \frac{N_c^2 B^2}{3\pi^2} \frac{\mu}{(\pi T)^2} (8 \log(2) - 3)\end{aligned}$$

Transport coefficients: [Landsteiner, Megias, Pena-Benitez 1207.5808]

$$\begin{aligned}\sigma^B &= -\frac{\kappa}{2\pi G_5} \mu = \frac{2N_c^2}{\pi^2 \sqrt{3}} \mu \\ \sigma^V = \sigma_\epsilon^B &= -\frac{\kappa}{2\pi G_5} \mu^2 = \frac{N_c^2}{\pi^2 \sqrt{3}} \mu^2 \\ \sigma_\epsilon^V &= -\frac{\kappa}{6\pi G_5} \mu^3 = \frac{2N_c^2}{3\pi^2 \sqrt{3}} \mu^3\end{aligned}$$

Hydro modes: [NA, Tabatabaei 1910.13696]

$$\begin{aligned}v_{CAW} &= 0 \\ v_{CMW} &= \frac{2}{\sqrt{3}} b (1 - 2\nu^2) \\ v_{sound} &= \frac{1}{\sqrt{3}} \left(\pm 1 + \frac{4}{3} \nu^2 b \right)\end{aligned}$$

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- Splitting of butterfly velocities might be originated just from chiral magnetic effects.

$$v_{sound} = \pm c_s + \frac{B}{2w} \frac{[\gamma, \alpha]}{[\beta, \alpha]} \left(-1 + \frac{(n\alpha_2 - w\beta_2)\partial_T - (n\alpha_1 - w\beta_1)\partial_\mu}{n[\gamma, \alpha] - w[\gamma, \beta]} \right) \sigma^B \quad (2.41)$$

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[Grozdanov, Schalm, Scopelliti 1710.00921]
 [Blake, Lee, Liu 1801.00010]
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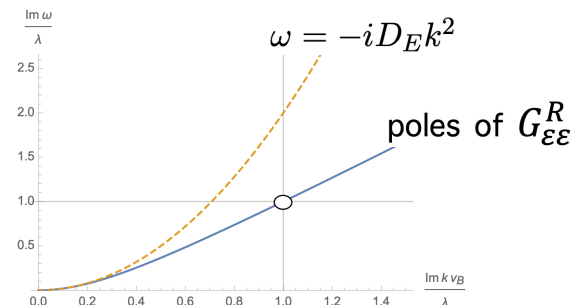
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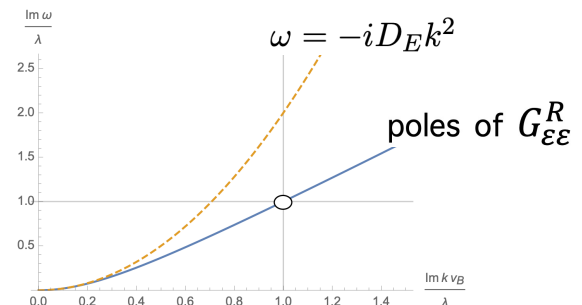
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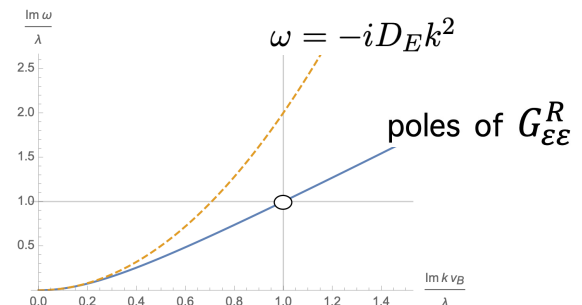
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Discussion

- Butterfly velocity is a measurable quantity in experiment.
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- Can it be regraded as a sign of CME?

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needs more exploration!

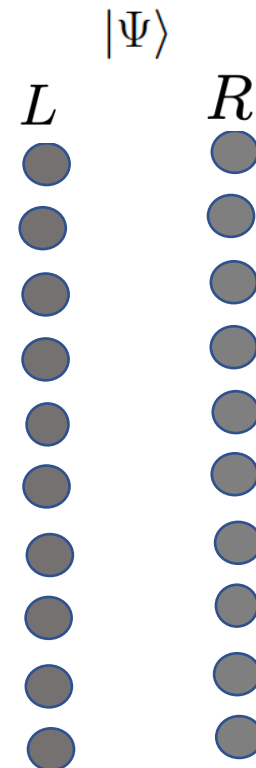
Thank you for your attention

Scrambling measures

In a qubit system with $H_L = \sum_{i=1}^{10} \left\{ \sigma_z^{(i)} \sigma_z^{(i+1)} - 1.05 \sigma_x^{(i)} + 0.5 \sigma_z^{(i)} \right\}$

Prepare the system in the thermofield state $|\Psi\rangle = \frac{1}{Z^{1/2}} \sum_n e^{-\beta E_n/2} |n\rangle_L |n\rangle_R$

[Shenker, Stanford 1306.0622]



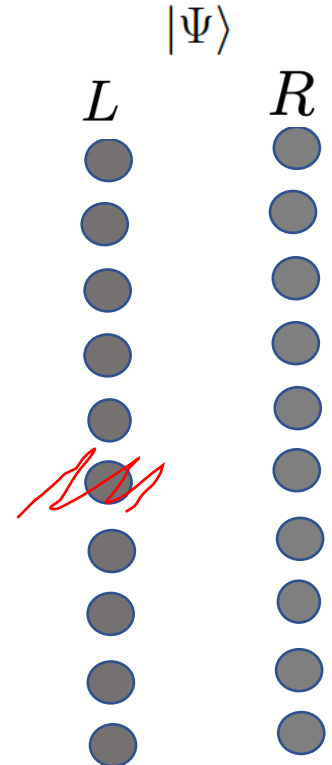
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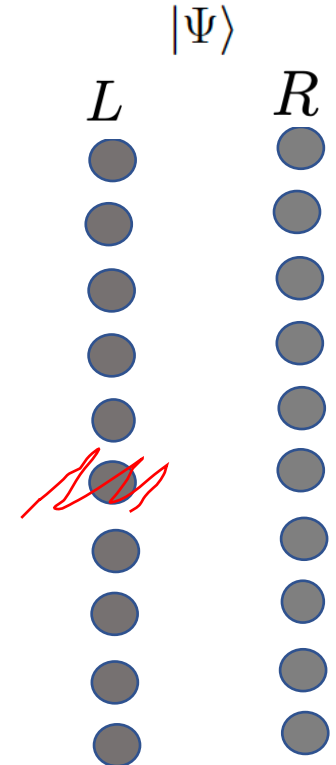
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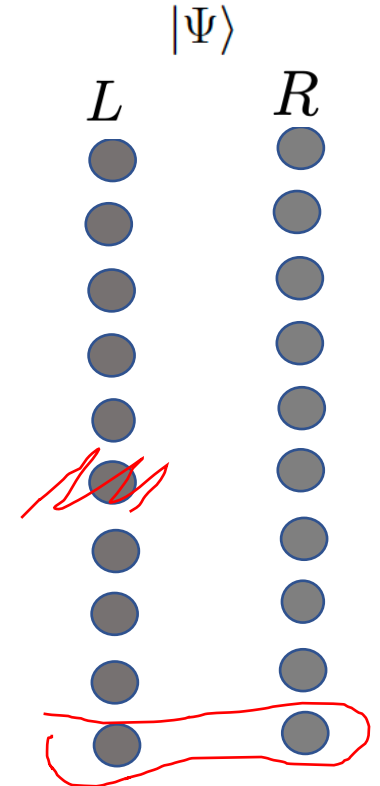
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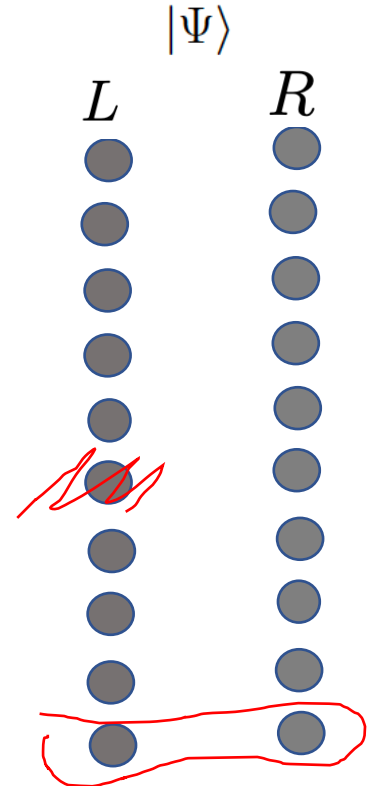
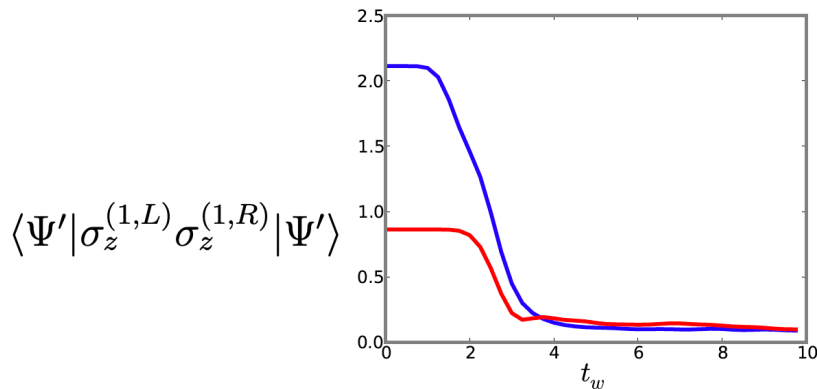
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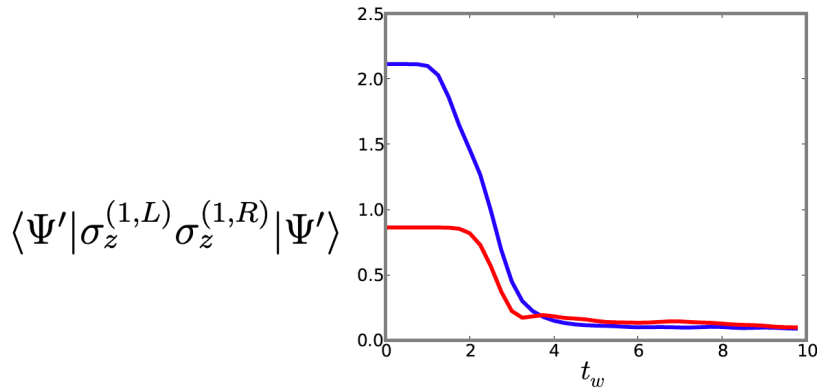
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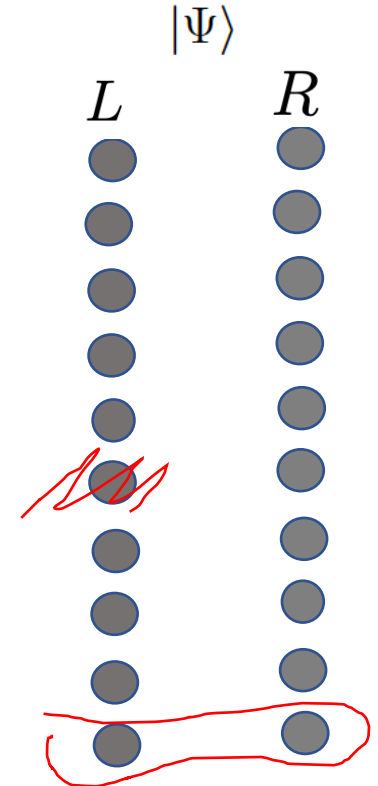
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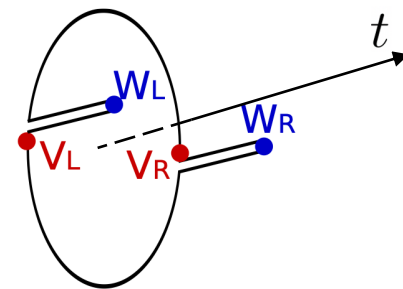
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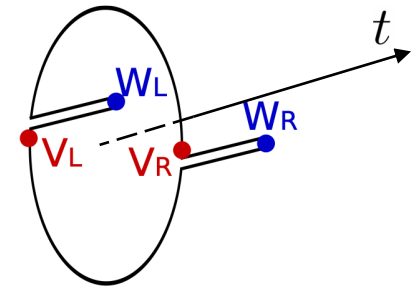
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- For large N holographic CFTs

$$F(t) = f_0 - \frac{f_1}{N^2} \exp \frac{2\pi}{\beta} t + \mathcal{O}(N^{-4})$$

[Maldacena, Shenker, Stanford 1503.01406]



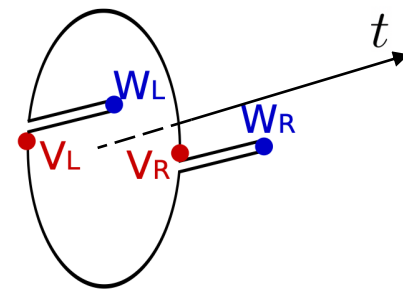
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Lyapunov exponent

Size of precursor growth

- The squared commutator

$$\begin{aligned} C(t_w, |x - y|) &= \text{tr} \{ \rho(\beta) [W_x(t_w), W_y]^\dagger [W_x(t_w), W_y] \} \\ &= 2 - 2 \text{Re} \langle TFD | W_y W_x(t_w) W_y W_x(t_w) | TFD \rangle \end{aligned}$$

- Size of precursor $s[W_x(t_w)]$ is the volume of region in y such that $C \geq 1$
= a ball centered at x of the radius

$$r[W_x(t_w)] \approx v_B(t_w - t_*)$$

$$C(t) = 2 - 2 \langle W(t, \vec{x}) V(0) W(t, \vec{x}) V(0) \rangle_\beta \sim \frac{1}{N} e^{\lambda(t - \frac{x}{v_B})}$$

Size of precursor growth

- The squared commutator

$$\begin{aligned} C(t_w, |x - y|) &= \text{tr} \{ \rho(\beta) [W_x(t_w), W_y]^\dagger [W_x(t_w), W_y] \} \\ &= 2 - 2 \text{Re} \langle TFD | W_y W_x(t_w) W_y W_x(t_w) | TFD \rangle \end{aligned}$$

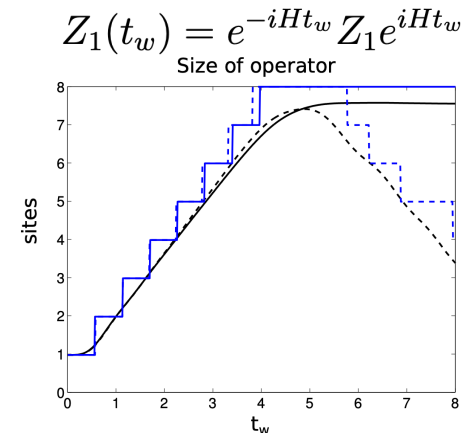
$$C(t) = 2 - 2 \langle W(t, \vec{x}) V(0) W(t, \vec{x}) V(0) \rangle_\beta \sim \frac{1}{N} e^{\lambda(t - \frac{x}{v_B})}$$

- Size of precursor $s[W_x(t_w)]$ is the volume of region in y such that $C \geq 1$
= a ball centered at x of the radius

$$r[W_x(t_w)] \approx v_B(t_w - t_*)$$

- Linear growth in time is checked numerically

in the spin chain $H = - \sum_i Z_i Z_{i+1} + g X_i + h Z_i$



Size of precursor growth

- The squared commutator

$$\begin{aligned}
 C(t_w, |x - y|) &= \text{tr} \{ \rho(\beta) [W_x(t_w), W_y]^\dagger [W_x(t_w), W_y] \} \\
 &= 2 - 2 \text{Re} \langle TFD | W_y W_x(t_w) W_y W_x(t_w) | TFD \rangle
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Butterfly speed

[Roberts, Shenker, Stanford, 1409.8180]

