



Femtoscopy of Light Nuclei in Au+Au Collisions at RHIC-STAR

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EXOTICO: EXOTIC atoms meet nuclei Collisions for a new frontier precision era in low-energy strangeness nuclear physics

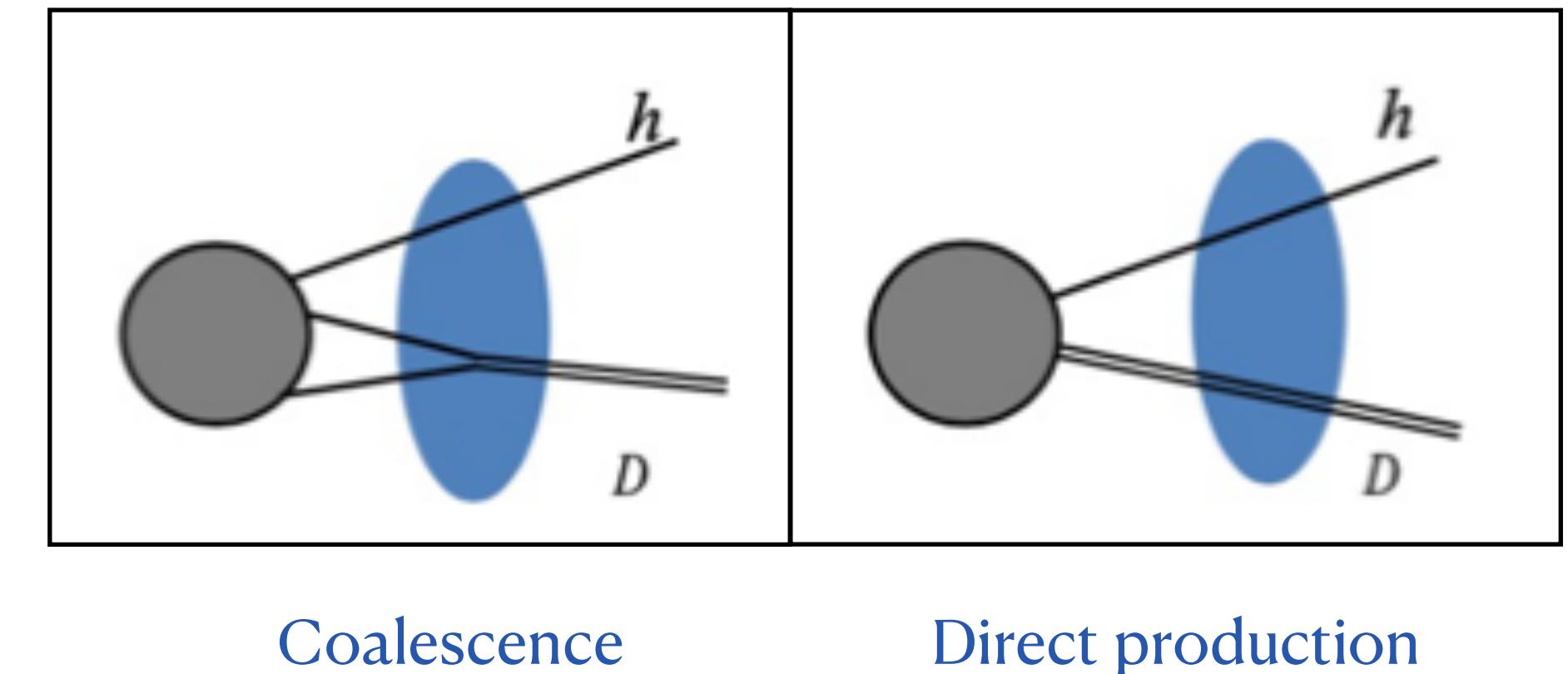
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4. Particle Identification
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Introduction



Two-particle correlations at small relative momenta contain information about the space-time characteristics of the emitting source and final-state interactions effects

- ✓ Formation mechanism of light nuclei are under debate
 - ⇒ Coalescence : final-state interaction
 - ⇒ Thermal : produced directly from fireball
 - ⇒ A systematic measurements of light nuclei femtoscopy may help to investigate

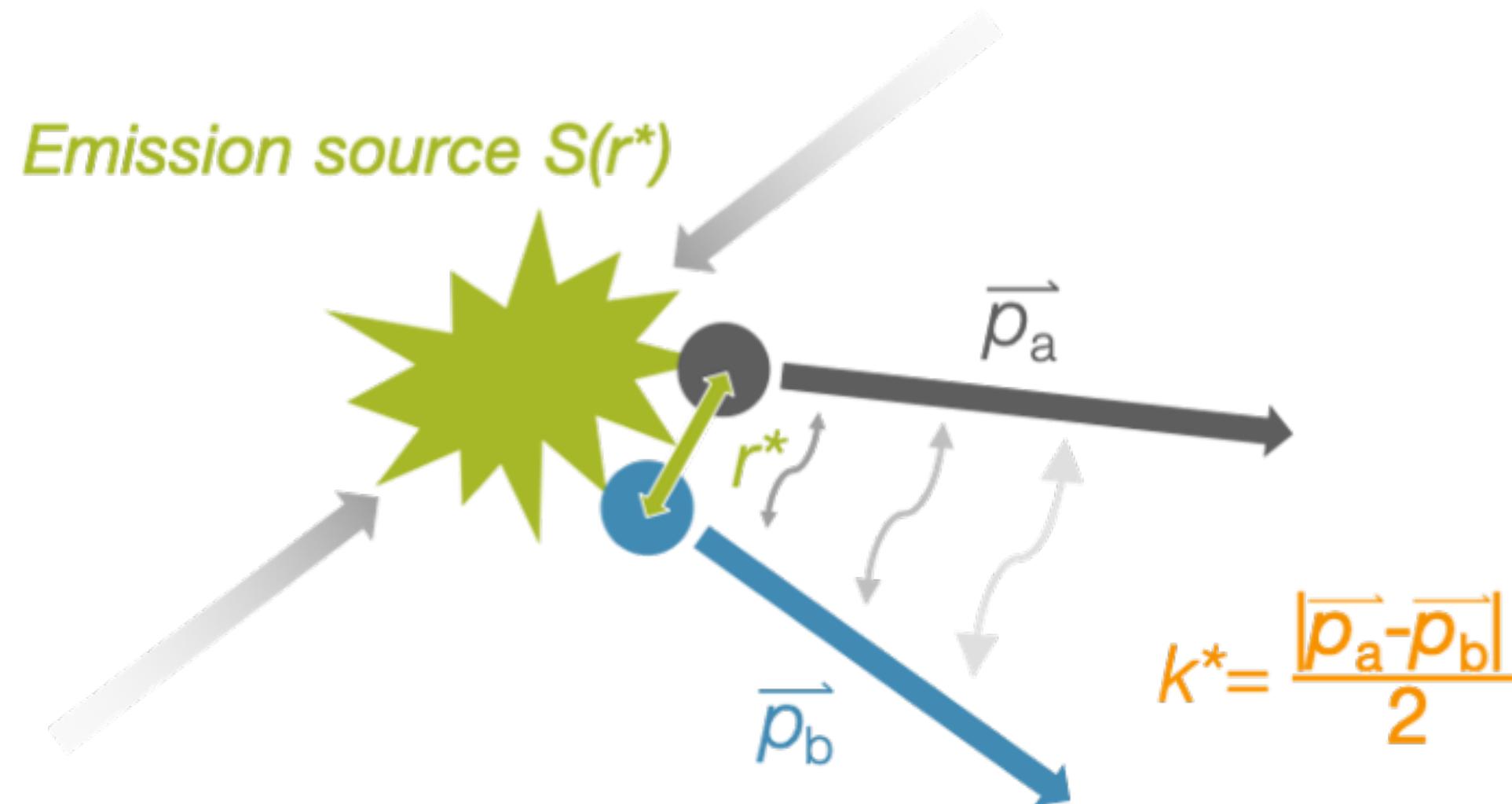


- ✓ A large amount of light nuclei produced at $\sqrt{s_{NN}} = 3 \text{ GeV}$
⇒ Allowing precision measurements of various light nuclei correlations(p-d, d-d, p- 3He , p- α , etc)

J. Cleymans et al, Phys. Rev. C 74, 034903 (2006)

K. Blum et al, Phys. Rev. C 99, 04491(2019)

Analysis Method: Femtoscopy



- ⇒ Femtoscopy (HIC) is inspired by Hanbury Brown and Twiss interferometry method (Astronomy)¹
- ⇒ Study the spatial and temporal extent of emission source
 - Quantum Statistics (Fermi-Dirac, Bose-Einstein)
 - Final-state Interactions (Coulomb, Strong interaction)
 - Collision Dynamics

✓ Two-particle correlation function:

<u>Statistical</u>	<u>Model</u>	<u>Experimental</u>
$C(\vec{k}^*) = \frac{\mathcal{P}(\vec{p}_a, \vec{p}_b)}{\mathcal{P}(\vec{p}_a)\mathcal{P}(\vec{p}_b)}$	$= \int S(\vec{r}) \Psi(\vec{k}^*, \vec{r}) ^2 d^3\vec{r}$	$= \frac{N_{\text{same}}(\vec{k}^*)}{N_{\text{mixed}}(\vec{k}^*)}$

\vec{p}_a, \vec{p}_b : Single-particle momentum

$S(\vec{r})$: Source function
 $\Psi(\vec{k}^*, \vec{r})$: Pair wave function

$k^* = \frac{1}{2} |\vec{p}_a - \vec{p}_b|$, relative momentum
 \vec{r} : relative distance

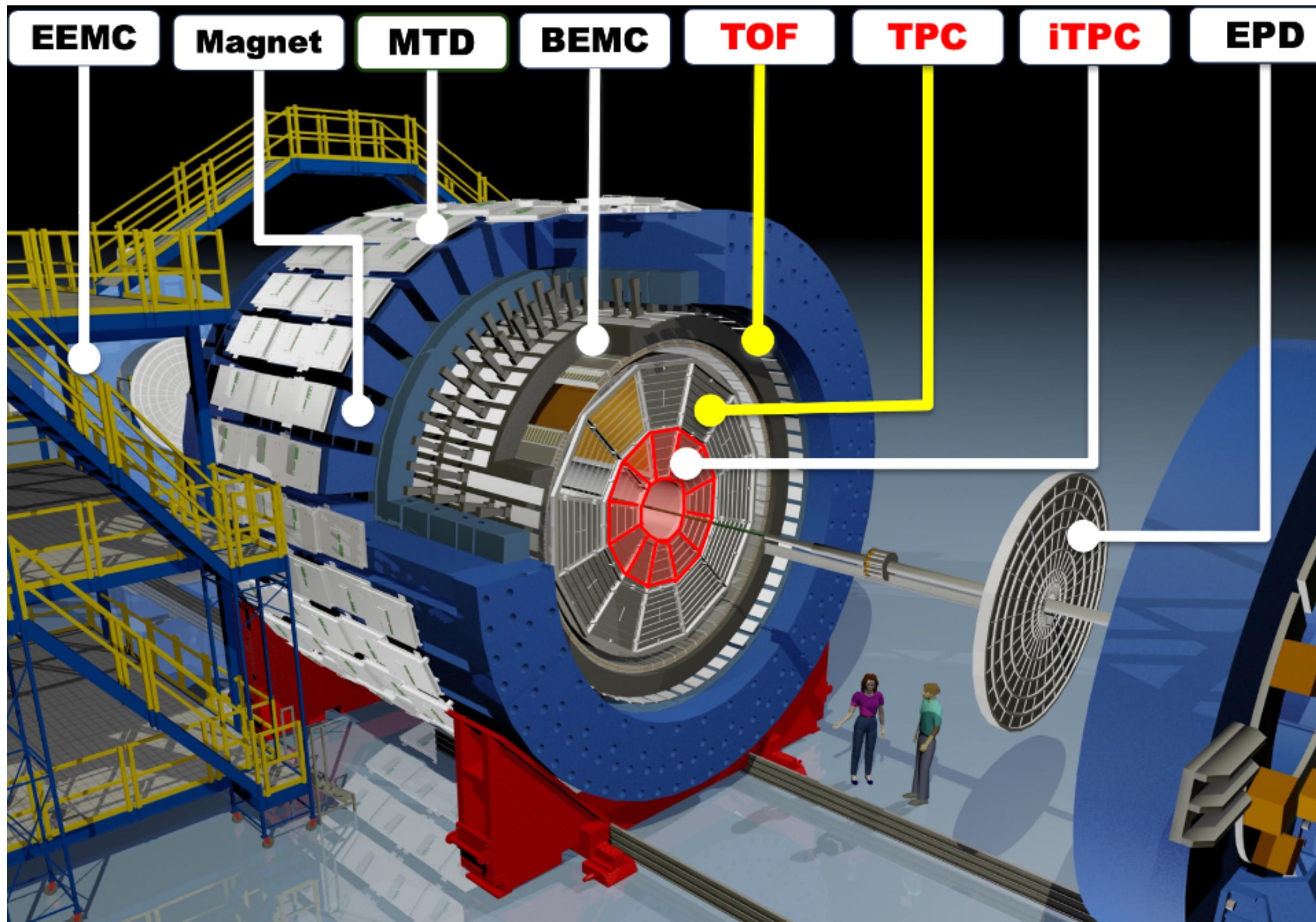
$N_{\text{same}}(\vec{k}^*)$: same event
 $N_{\text{mixed}}(\vec{k}^*)$: mixed event

>1 : Attraction
 $=1$: No Correlation
 <1 : Repulsion

¹Nature 178 1046-1048(1956)

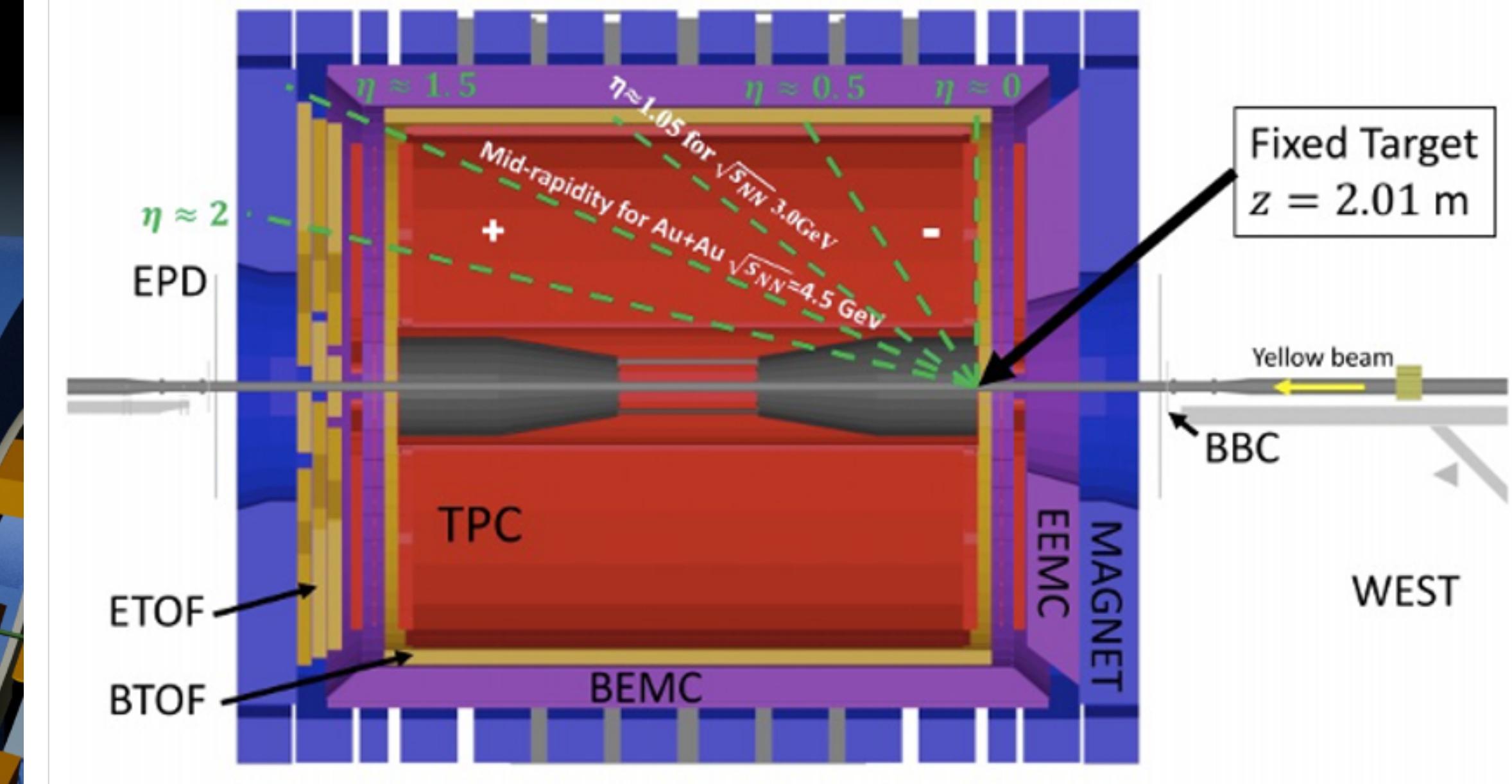
²ALICE Coll. Nature 588, 232-238 (2020)

STAR Detector



- Excellent Particle Identification
- Large, Uniform Acceptance at Mid-rapidity

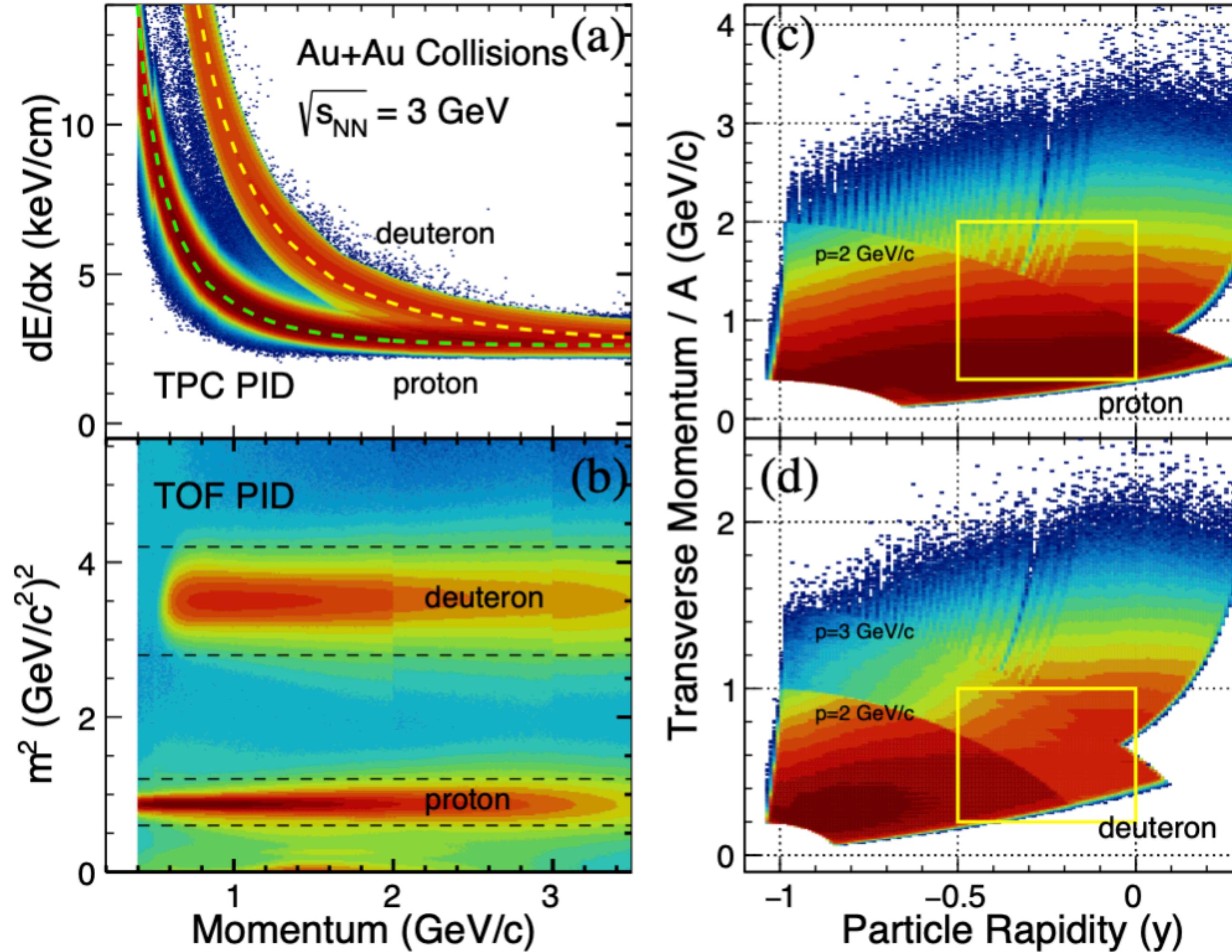
STAR Fixed-target Experiment Setup



→ Dataset

$\sqrt{s_{NN}} = 3 \text{ GeV}$, Au+Au Collisions, Fixed-target mode
~260 millions minimum bias triggered event

Particle Identification

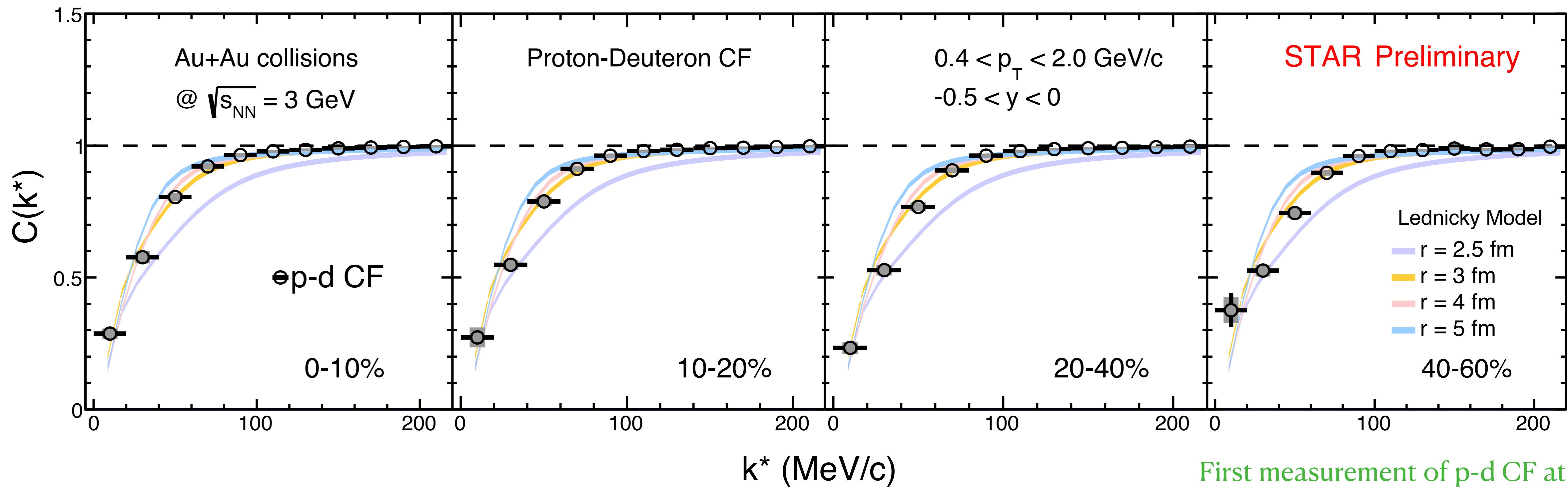


- ⇒ PID using TPC (low momentum) and TOF (high momentum)
- ⇒ Excellent acceptance within middle rapidity window
- ⇒ High purity particle sample
 - Proton purity > 99%
 - Deuteron purity > 97%
 - Pair purity: > 96% (pd), > 95% (dd)

Proton-Deuteron Femtoscopy



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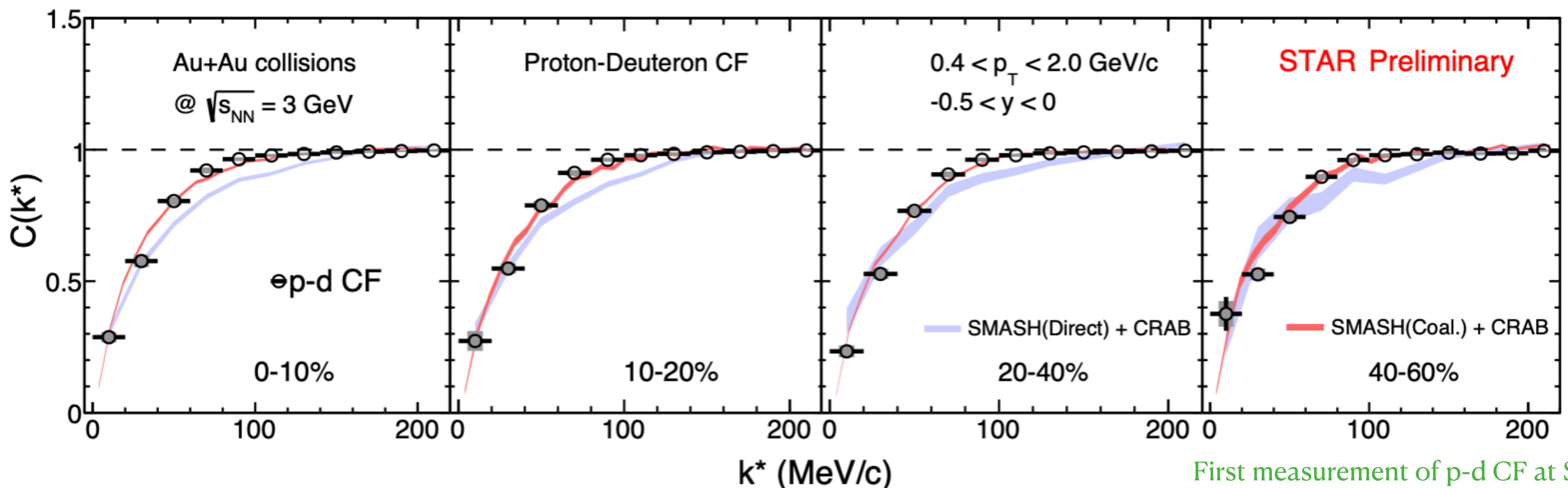
First measurement of p-d CF at STAR

- ⇒ Clear depletion at small k^* range seen in data
- ⇒ Data compared with Lednicky&Lyuboshitz model
- A spherical source size with $r_g = 3 - 4 \text{ fm}$ is consistent with data

Proton-Deuteron Femtoscopy



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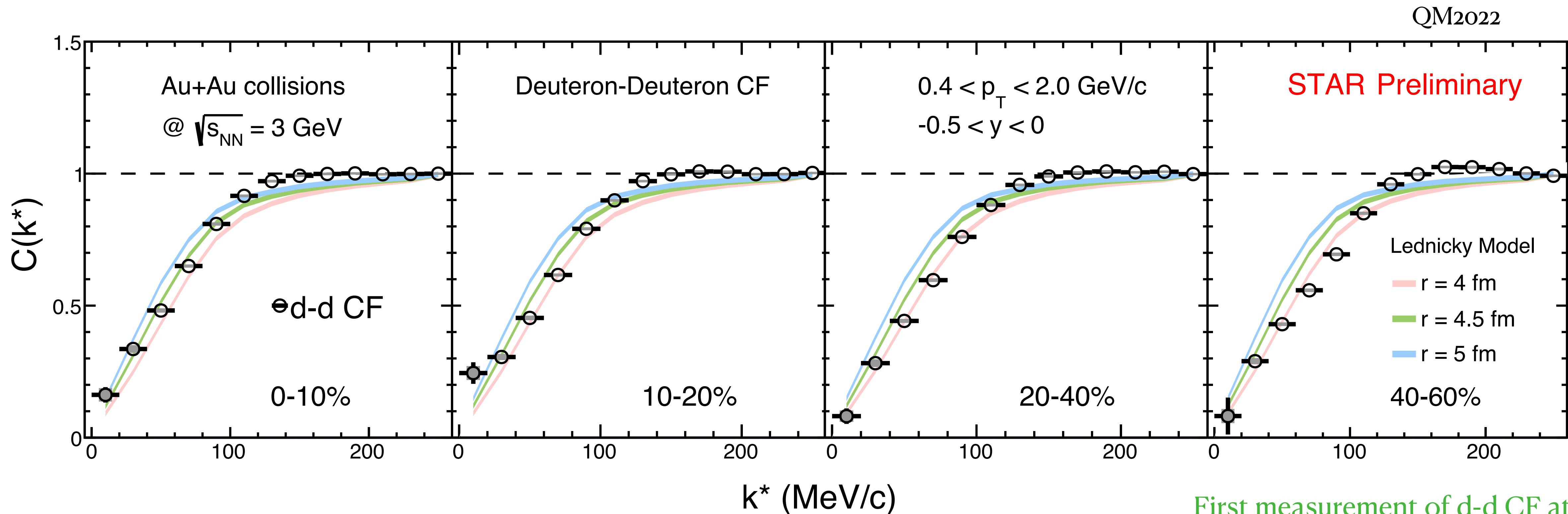


⇒ Compared with SMASH + Correlation After burner (CRAB) model
CF calculated with coalescence of deuterons is in better agreement with data
Support the deuteron formation at 3 GeV is dominated by coalescence

Deuteron-Deuteron Femtoscopy



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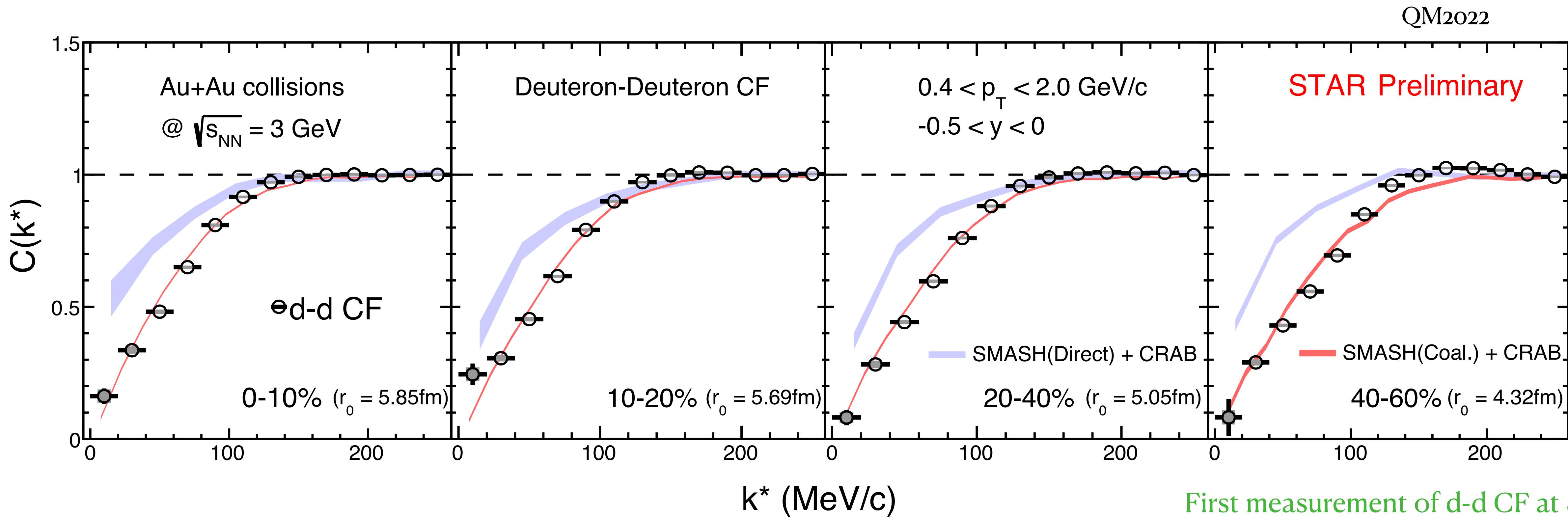


- ⇒ Clear depletion at small k^* range seen in data
- ⇒ Data compared with Lednicky & Lyuboshitz model
- A spherical source size with $r_g = 4-5 \text{ fm}$ is consistent with data

Deuteron-Deuteron Femtoscopy



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First measurement of d-d CF at STAR

⇒ Compared with SMASH + Correlation After burner (CRAB) model

CF calculated with coalescence of deuterons is in better agreement with data

Support the deuteron formation at 3 GeV is dominated by coalescence

Summary



- ✓ First measurement of p-d and d-d correlation functions with good precision from STAR
- ✓ p-d and d-d correlation functions qualitatively described by L&L model
- ✓ p-d and d-d CF described better by the model including coalescence
 - ⇒ Light nuclei are likely to be formed via coalescence at $\sqrt{s_{NN}} = 3 \text{ GeV}$

⇒ With high precision data, the properties of strong interactions among light nuclei pairs is achievable
⇒ More inputs from theory are needed
⇒ In the 2nd phase of BES, STAR has collected 10-20 times more data in Au+Au collisions at the energy range $\sqrt{s_{NN}} = 3 - 19.6 \text{ GeV}$. These data allow us to perform precision light nuclei femtoscopy analysis !

Stay tuned for the RHIC BES-II !

Thank you for your attention !

Back up slides

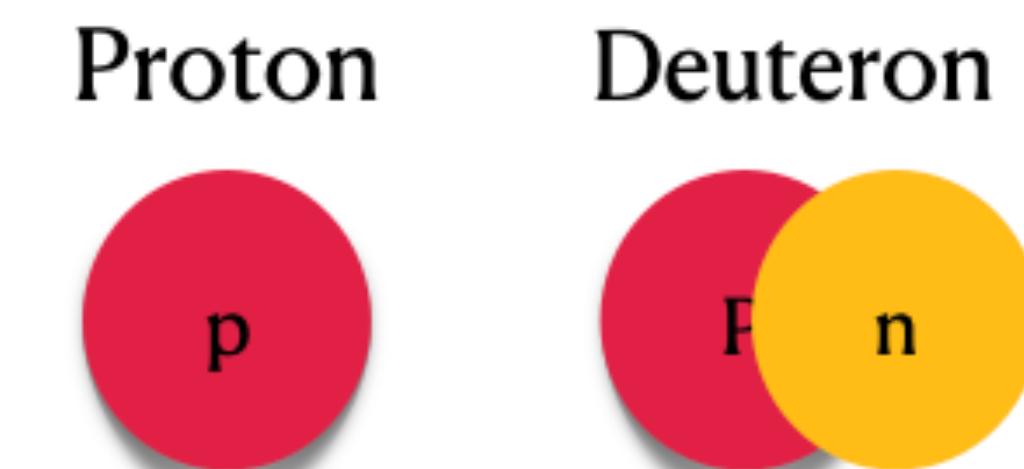
⇒ Phase Space of proton / neutron /deuteron is produced via SMASH transport model

⇒ Coalescence ($p + n \rightarrow d$):

⇒ Formation probability calculated via Wigner Function

$$f_2(\rho, \mathbf{p}_\rho) = 8g_2 \exp \left[-\frac{\rho^2}{\sigma_\rho^2} - \mathbf{p}_\rho^2 \sigma_\rho^2 \right],$$

$$\rho = \frac{1}{\sqrt{2}}(\mathbf{x}'_1 - \mathbf{x}'_2), \quad \mathbf{p}_\rho = \sqrt{2} \frac{m_2 \mathbf{p}'_1 - m_1 \mathbf{p}'_2}{m_1 + m_2},$$



⇒ Related to the relative coordinate and relative momentum of proton and neutron

⇒ Interaction

⇒ SMASH does not contain FSI interaction between two particles

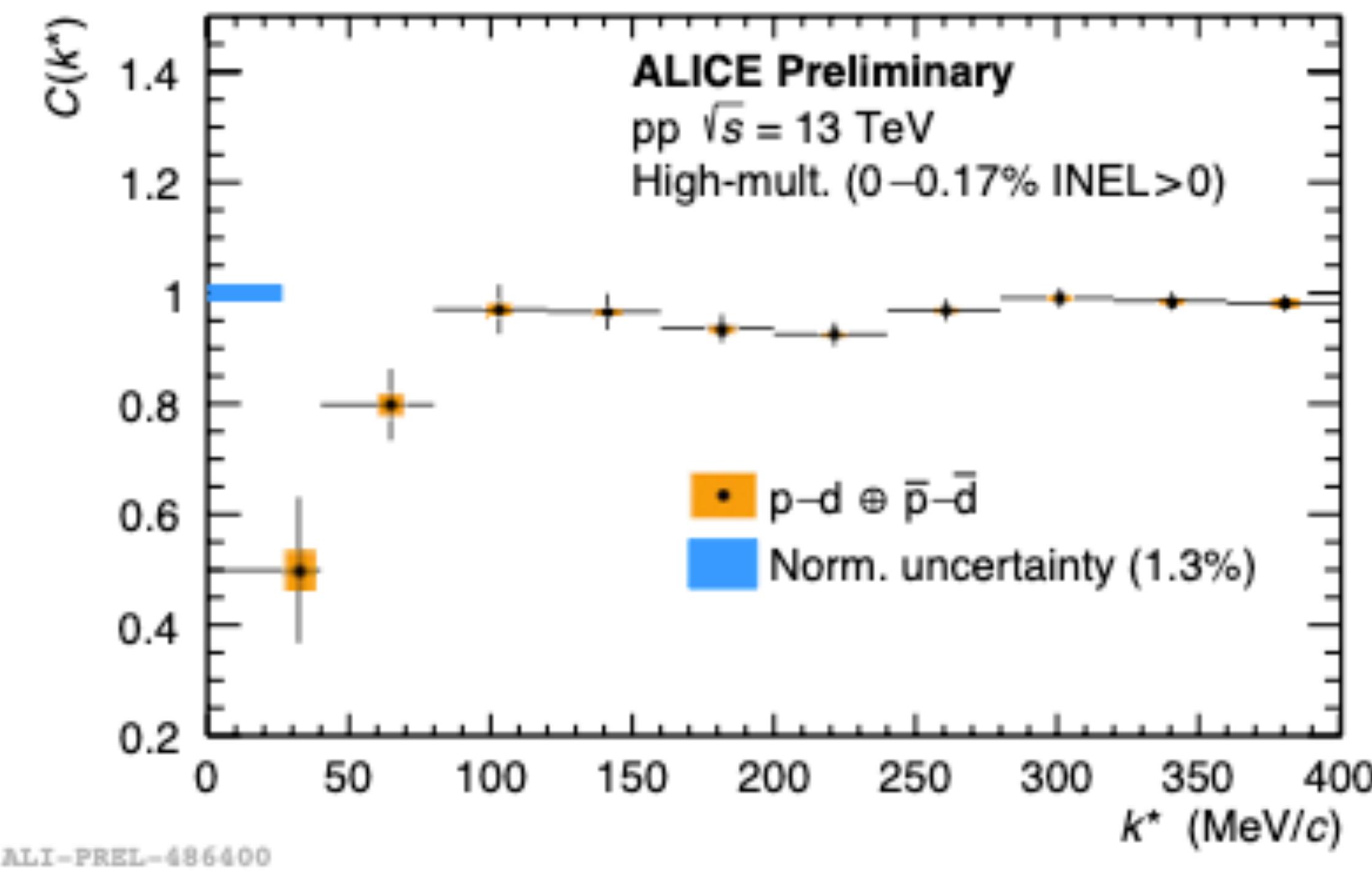
⇒ All of Coulomb, QS and SI potentials are introduced via CRAB

SMASH: J. Weil et al. Phys.Rev.C 94 (2016) 5, 054905

Coalescence: W.Zhao et al. Phys. Rev. C.98 (2018) 5,054905

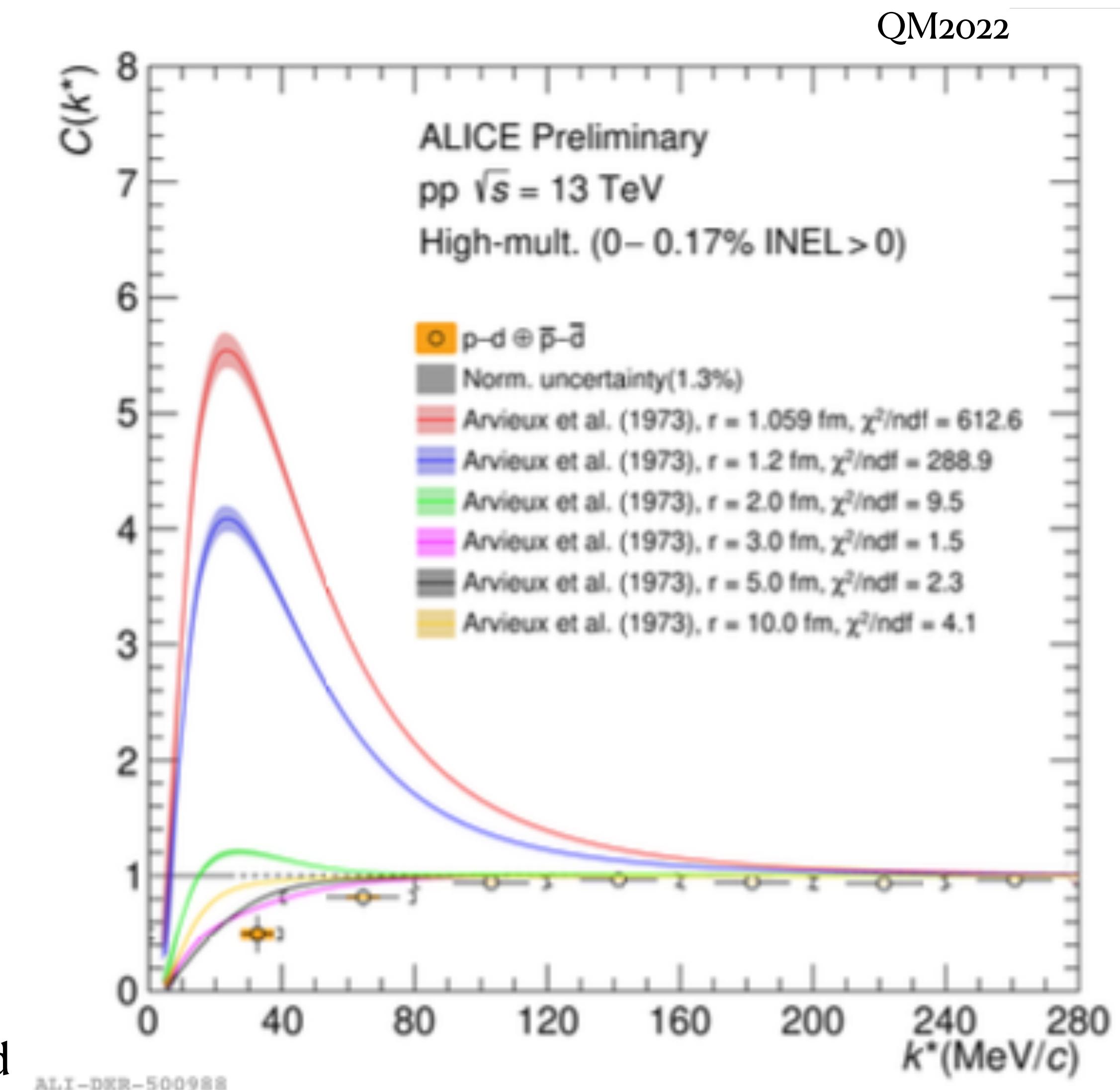
CRAB (p-d): Private conversations with Prof.Scott Pratt

Proton-Deuteron CF from ALICE



Interpretation:

- ⇒ Improved agreement with larger source sizes
- ⇒ CF becomes flat at large source size
- ⇒ The effect of attractive strong interaction in the CF is diluted



Proton-Cluster CF from HADES



M.Stefaniak, WPCF 2022

Ag+Ag @ 1.58A GeV

