# Efficient Bosonic and Fermionic Sinkhorn Algorithms for Non-Interacting Ensembles in One-body Reduced Density Matrix Functional Theory in the Canonical Ensemble 

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Manuscript arXiv:2205.15058 [chem-phys] (under review)
Code: https://www.github.com/DerkKooi/bfsinkhorn
See also Sarina Sutter her talk and arXiv:2209.11663 by S.M. Sutter and K.J.H. Giesbertz

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3. K.J.H. Giesbertz and M. Ruggenthaler. Physics Reports 806, 1-47 (2019), DOI: 10.1016/j.physrep.2019.01.010

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- When using a (non-interacting) reference system: canonical reference may be "closer" to interacting system than grand canonical reference

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- Use (canonical) entropy to model correlation energy ${ }^{[4]}$

VUK $=$

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5

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W[\gamma] & \left.=W_{0} \text { GCl } \gamma\right]+W_{c}[\gamma] \\
& \text { Will come back later! }
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6
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Canonical: $\quad A^{\beta}[h]=\min _{\hat{\Gamma} \in \mathscr{H}_{N} \otimes \mathscr{H}_{N}, \operatorname{Tr}(\hat{\Gamma})=1}\left(\operatorname{Tr}(\hat{\Gamma} \hat{H})+\frac{1}{\beta} \operatorname{Tr}(\hat{\Gamma} \log (\hat{\Gamma}))\right)$

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- Furthermore, for any interaction (also no interaction):[3,5] $h \leftrightarrow \gamma$

3. K.J.H. Giesbertz and M. Ruggenthaler. Physics Reports 806, 1-47 (2019), DOI: 10.1016/i.physrep.2019.01.010
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- Constrained search:

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A_{\mathrm{int}}^{\beta}[\gamma]=W^{\beta}[\gamma]-\frac{1}{\beta} S^{\beta}[\gamma]=\min _{\hat{\Gamma} \rightarrow \gamma}\left(\operatorname{Tr}(\hat{\Gamma} \hat{W})+\frac{1}{\beta} \operatorname{Tr}(\hat{\Gamma} \log (\hat{\Gamma}))\right)
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- Non-interacting Hamiltonian[8], but pathological

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8. K.J.H. Giesbertz and E.J. Baerends.
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& \left.\mu\right|_{\epsilon_{p}} ^{-\frac{-}{-}-}-0<n_{p}<N \quad 0<n_{p}<1
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& \min _{\hat{\Gamma}_{0}^{\beta} \rightarrow \gamma} \frac{1}{\beta} \operatorname{Tr}\left(\hat{\Gamma}_{0}^{\beta} \log \left(\hat{\Gamma}_{0}^{\beta}\right)\right)=\frac{1}{\beta} \min _{\hat{\Gamma}_{0} \rightarrow \gamma} \operatorname{Tr}\left(\hat{\Gamma}_{0} \log \left(\hat{\Gamma}_{0}\right)\right)=-\frac{1}{\beta} S_{0}\left[\left\{n_{p}\right\}\right] \\
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\hat{H}_{0}^{\beta}=\sum_{p} \epsilon_{p}^{\beta}\left[\left\{n_{q}\right\}\right] \hat{n}_{p} \quad \epsilon_{p}^{\beta}=\frac{\beta^{\prime}}{\beta} \epsilon_{p}^{\beta^{\prime}}
\end{gathered}
$$

- Define a non-interacting free energy approximation $A_{0}^{\beta}[\gamma]$ :

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A_{0}^{\beta}[\gamma]=W_{0}[\gamma]-\frac{1}{\beta} S_{0}\left[\left\{n_{p}\right\}\right]
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- Note that the functional for the (grand) canonical ensembles are not the same: constrained search over different Hilbert space
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& n_{p}=\frac{1}{e^{\beta\left(\epsilon_{p}-\mu\right)} \mp 1} \longleftrightarrow \epsilon_{p}\left[n_{p}\right]=-\frac{1}{\beta} \log \left(\frac{n_{p}}{1 \pm n_{p}}\right)
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VU $/{ }^{m=}$

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17


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- We also would like to have $W_{0}[\gamma]$ for the Canonical ensemble

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\operatorname{Tr}\left(\hat{\Gamma}_{0} \hat{S}^{2}\right) \neq \operatorname{Tr}\left(\hat{\Gamma} \hat{S}^{2}\right)
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- We can alleviate this by restricting our Hilbert (sub)space

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Less variational
freedom



11. E.H. Lieb. Phys. Rev. Lett. 46, 457 (1981), DOI:
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- As of yet: only evaluation on exact 1-RDMs at zero temperature
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- In preparation: a reference that takes into account part of the interaction



27
VUK=

- Finite temperature 1-RDMFT in the canonical ensemble was studied numerically for the first time
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- Bosonic and Fermionic Sinkhorn algorithms were derived and implemented in the bfsinkhorn package
- The algorithms were shown to be efficient and perform well for both "simulated" and ground-state 1-RDMs
- A study of the corresponding canonical approximation to the interaction $W_{0}[\gamma]$ revealed interesting behaviour w.r.t. grand canonical

- Klaas J.H. Giesbertz, Paola Gori Giorgi, Evert Jan Baerends, Sarina M. Sutter and Mauricio Rodríguez Mayorga for insightful discussions


Nederlandse Organisatie voor Wetenschappelijk Onderzoek

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- Financial support:
-Netherlands Organisation for Scientific Research under Vici grant 724.017.001 (Paola Gori Giorgi)


Nederlandse Organisatie voor Wetenschappelijk Onderzoek

- Manuscript: arXiv:2205.15058 (under review)


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- bfsinkhorn: https://www.github.com/DerkKooi/bfsinkhorn

1. T. Baldsiefen, A. Cangi and E.K.U. Gross. Phys. Rev. A. 92, 052514 (2015), DOI: $10.1103 /$ physreva. 92.052514
2. T. Baldsiefen and E.K.U. Gross. Comp. and Theo. Chem. 1003, 114 (2013), DOI: 10.1016/j.comptc. 2012.09 .001
3. K.J.H. Giesbertz and M. Ruggenthaler. Physics Reports 806, 1-47 (2019), DOI: 10.1016/j.physrep.2019.01.010
4. J. Wang and E.J. Baerends. Phys. Rev. Lett. 128, 013001 (2022), DOI: 10.1103/PhysRevLett. 128.013001
5. S.M. Sutter and K.J.H. Giesbertz. arXiv:2209.11663 [math-ph]
6. A.J. Coleman, Rev. Mod. Phys. 35, 668 (1963). DOI: 10.1103/RevModPhys.35.668
7. S.M. Valone. Phys. Rev. B 44, 1509 (1991), DOI: 10.1103/PhysRevB.44.1509
8. K.J.H. Giesbertz and E.J. Baerends. J. Chem. Phys. 132, 194108 (2010), DOI: 10.1063/1.3426319
9. H. Barghati, J. Yu and A.D. Maestro. Phys. Rev. Res. 2, 043206 (2020), DOI: 10.1103/physrevresearch. 2.043206
10. P. Borrmann and G. Franke. J. Chem. Phys. 98, 2484 (1993), DOI: 10.1063/1.464180
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$$
\begin{gathered}
h_{p p}-\frac{1}{\beta} \frac{\partial S_{0}\left[\left\{n_{q}\right\}\right]}{\partial n_{p}}+\frac{\partial W_{0}[\gamma]}{\partial n_{p}}+\frac{A_{c}^{\beta}[\gamma]}{\partial n_{p}}=0 \\
\epsilon_{p}^{0, \beta(i+1)}=h_{p p}^{(i)}+\frac{\partial W_{0}\left[\gamma^{(i)}\right]}{\partial n_{p}}+\frac{\partial A_{c}^{\beta}\left[\gamma^{(i)}\right]}{\partial n_{p}^{(i)}} \longrightarrow n_{p}^{(i+1)}\left[\left\{\epsilon_{q}^{0, \beta(i+1)}\right\}\right]
\end{gathered}
$$

- We need to optimise w.r.t. NOONs $\left\{n_{p}\right\}$ and NOs $\left\{\phi_{p}(\mathrm{x})\right\}$
- Taking the derivative of $A^{\beta}[\gamma]$ w.r.t. $n_{p}$ :

$$
\begin{gathered}
h_{p p}-\frac{1}{\beta} \frac{\partial S_{0}\left[\left\{n_{q}\right\}\right]}{\partial n_{p}}+\frac{\partial W_{0}[\gamma]}{\partial n_{p}}+\frac{A_{c}^{\beta}[\gamma]}{\partial n_{p}}=0 \\
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\end{gathered}
$$

- Optimisation over NOs yields effective one-particle Schrödinger equation
- Analogous for Grand Canonical


31
VU/ $\mathrm{F}^{\mathrm{m}}$




- Shift by a constant and either match the strongly or weakly occupied



35
vu $=$

- At finite temperature we have instead:
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Canonical:

$$
\hat{\Gamma}^{\beta}=\frac{e^{-\beta \hat{H}}}{Z^{\beta}}, \quad Z^{\beta}=\operatorname{Tr}\left(e^{-\beta \hat{H}}\right)
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Grand Canonical: $\quad \hat{\Gamma}^{\beta, \mu}=\frac{e^{-\beta(\hat{H}-\mu \hat{N})}}{\mathscr{Z}^{\beta, \mu}}, \quad \mathscr{Z}^{\beta, \mu}=\operatorname{Tr}\left(e^{-\beta(\hat{H}-\mu \hat{N})}\right)$

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- Note the different Hilbert spaces (single particle space $\mathfrak{h}$ )
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Grand Canonical: $\quad \mathscr{F}=\bigoplus_{N=0}^{\infty} \mathscr{H}_{N} \quad \quad \hat{\Gamma}^{\beta, \mu} \in \mathscr{F} \otimes \mathscr{F}$

$$
\left.\begin{array}{l}
\mathcal{S}_{z}=\left\{\begin{array}{ll}
\left|\Phi_{P}\right\rangle \in \mathscr{H}_{N} \mid & \hat{S}_{z}\left|\Phi_{P}\right\rangle=S_{z}\left|\Phi_{P}\right\rangle
\end{array}\right\} \\
\mathcal{S}_{z}=\left\{\left|\Xi_{P}\right\rangle \in \mathscr{H}_{N}\left|\quad \hat{S}^{2}\right| \Xi_{P}\right\rangle=S^{2}\left|\Xi_{P}\right\rangle
\end{array}\right\}
$$

