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# Phenomenology of the nucleon internal pressure



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Revealing emergent mass | Hervé MOUTARDE

### Sep. 16, 2022

UNIVERSITE PARIS-SACLAY

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### Perturbative and nonperturbative QCD. Study hadron structure to shed new light on nonperturbative QCD.





## **Energy-momentum tensor**

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### Energy-momentum tensor. Quark and gluon contributions.



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Isolating  $d_1$ 

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#### Abbreviations

- EMT defined from the invariance under space and time translations.
- Quark and gluon contributions

$$\begin{aligned} T_{q}^{\mu\nu} &= \bar{q}\gamma^{\mu}\frac{i}{2}\overset{\leftrightarrow}{\mathrm{D}}q\\ T_{g}^{\mu\nu} &= -F^{\mu\lambda}F^{\nu}{}_{\lambda} + \frac{1}{4}\eta^{\mu\nu}F^{2} \end{aligned}$$

with  $\stackrel{\leftrightarrow}{D}$  the symmetric covariant derivative and  $F^{\mu\nu}$  the field strength tensor.

$$T^{\mu\nu} = \sum_{a} T^{\mu\nu}_{a} \ (a = q, g).$$

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### Parameterization: massive spin-1/2 target. Introduction of 5 GFFs.



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Local, gauge-invariant, asymmetric EMT:

$$\left\langle p', s' \right| T^{\mu\nu}_{a}(0) \left| p, s \right\rangle = \bar{u}(p', s') \begin{cases} \frac{P^{\mu}P^{\nu}}{M} A_{a}(t) + M\eta^{\mu\nu} \bar{C}_{a}(t) \\ + \frac{\Delta^{\mu}\Delta^{\nu} - \eta^{\mu\nu}\Delta^{2}}{M} C_{a}(t) \\ + \frac{P^{\{\mu}i\sigma^{\nu\}\Delta}}{4M} \left[ A_{a}(t) + B_{a}(t) \right] \\ + \frac{P^{[\mu}i\sigma^{\nu]\Delta}}{4M} D_{a}(t) \end{cases}$$



### 3D profile of GFFs. Localization in the Wigner sense.



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Define distribution of a physical quantity inside a system, by first localizing the system in both position and momentum space.

Breit frame where 
$${\it P}^{\mu}=({\it P}^{0},ec{0})$$
 and  $\Delta^{\mu}=(0,ec{\Delta})$ 

$$\langle T^{\mu\nu}_{a} \rangle_{\mathrm{BF}(\vec{r})} = \int \frac{\mathrm{d}^{3}\Delta}{(2\pi)^{3}} e^{-i\vec{\Delta}\vec{r}} \left[ \frac{\langle p', s | T^{\mu\nu}_{a}(0) | p, s \rangle}{2P^{0}} \right]_{\vec{P}=\vec{0}}$$

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### Gravitational form factors. Definition of pressure.



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• Matrix element in the Breit frame (a = q, g):

$$\left\langle \frac{\Delta}{2} \left| \mathcal{T}_{a}^{\mu\nu}(0) \right| - \frac{\Delta}{2} \right\rangle = \mathcal{M} \left\{ \eta^{\mu 0} \eta^{\nu 0} \left[ \mathcal{A}_{a}(t) + \frac{t}{4M^{2}} \mathcal{B}_{a}(t) \right] \right. \\ \left. + \eta^{\mu\nu} \left[ \bar{\mathcal{C}}_{a}(t) - \frac{t}{M^{2}} \mathcal{C}_{a}(t) \right] + \frac{\Delta^{\mu} \Delta^{\nu}}{M^{2}} \mathcal{C}_{a}(t) \right\}$$

# Anisotropic fluid in **relativistic hydrodynamics**: $\Theta^{\mu\nu}(\vec{r}) = [\varepsilon(r) + p_t(r)] u^{\mu}u^{\nu} - p_t(r)\eta^{\mu\nu} + [p_r(r) - p_t(r)] \chi^{\mu}\chi^{\nu}$ where $u^{\mu}$ and $\chi^{\mu} = x^{\mu}/r$ .

Define isotropic pressure and pressure anisotropy:

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 $p(r) = \frac{p_r(r) + 2 p_t(r)}{3}$  $s(r) = p_r(r) - p_t(r)$ 

🖾 Lorcé et al. (2019)



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### Mechanical properties of hadrons. Pressure from gravitational form factors.



Nucleon  
internal  
pressure
Write dictionary between quantum and fluid pictures:
$$\frac{\varepsilon_{a}(r)}{M} = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ A_{a}(t) + \bar{C}_{a}(t) + \frac{t}{4M^{2}} \left[ B_{a}(t) - 4C_{a}(t) \right] \right\}$$
Energy-  
momentum  
tensor
Gravitational form
$$\frac{P_{r,a}(r)}{M} = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\bar{C}_{a}(t) - \frac{4}{r^{2}} \frac{t^{-1/2}}{M^{2}} \frac{d}{dt} \left( t^{3/2} C_{a}(t) \right) \right\}$$
Nucleon EOS
Experiments
Phenomenology  $t, a(r)$ 
Strategy
$$M = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\bar{C}_{a}(t) + \frac{4}{r^{2}} \frac{t^{-1/2}}{M^{2}} \frac{d}{dt} \left[ t \frac{d}{dt} \left( t^{3/2} C_{a}(t) \right) \right]$$
Strategy
$$M = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\bar{C}_{a}(t) + \frac{2}{r^{2}} \frac{t^{-1/2}}{M^{2}} \frac{d}{dt} \left[ t \frac{d}{dt} \left( t^{3/2} C_{a}(t) \right) \right]$$
Areas for
improvement
CFF fits
GFF transfor
$$M = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\bar{C}_{a}(t) + \frac{2}{3} \frac{t}{M^{2}} C_{a}(t) \right\}$$
Conclusion
Abbreviations
$$M = \int \frac{d^{3}\vec{\Delta}}{(2\pi)^{3}} e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\frac{4}{r^{2}} \frac{t^{-1/2}}{M^{2}} \frac{d^{2}}{dt^{2}} \left( t^{5/2} C_{a}(t) \right) \right\}$$

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### Mechanical properties of hadrons. Pressure from gravitational form factors.





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### Equation of state.

Elaborating on the relation between energy and pressure.



#### Nucleon internal pressure

■ Simple multiple models: dipole for GFFs *A* and *C*, tripole for GFFs *B* and *C*.





Neutron stars





### Equation of state.

Elaborating on the relation between energy and pressure.



#### Nucleon internal pressure

Abbreviations

# Simple multiple models: dipole for GFFs A and C, tripole for GFFs B and C.



Lorcé et al.

(2019)





### Equation of state.

Elaborating on the relation between energy and pressure.



#### Nucleon internal pressure

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### From the nucleon to compact stars?









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Link between GPDs and	GFF	s	
$\int \mathrm{d}x  x \mathbf{H}^q(x,\xi,t)$	=	$\mathbf{A}^{\mathbf{q}}(t) + 4\xi^2 \mathbf{C}^{\mathbf{q}}(t)$	
$\int \mathrm{d}x  \mathbf{x} \mathbf{E}^q(\mathbf{x}, \xi, t)$	=	$B^q(t) - 4\xi^2 C^q(t)$	
	⊿ J:	i (1997), ⁄ Goeke	(2001)





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Link	between GPDs and	GFF	s	
	$\int \mathrm{d}x  x \boldsymbol{H}^{\boldsymbol{q}}(x,\xi,t)$	=	$\boldsymbol{A}^{\boldsymbol{q}}(t) + 4\xi^2 \boldsymbol{C}^{\boldsymbol{q}}(t)$	
	$\int \mathrm{d}x  x \boldsymbol{E}^q(x,\xi,t)$	=	$\boldsymbol{B}^{\boldsymbol{q}}(t) - 4\xi^2 \boldsymbol{C}^{\boldsymbol{q}}(t)$	
		⊿∎ J	i (1997), ⁄ Goeke	(2001)











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Abbreviations

Link between GPDs and GFFs  $\int dx x H^q(x,\xi,t) = A^q(t) + 4\xi^2 C^q(t)$   $\int dx x E^q(x,\xi,t) = B^q(t) - 4\xi^2 C^q(t)$   $\not \equiv \text{ Ji (1997), } \not \equiv \text{ Goeke (2001)}$ 

### Deeply Virtual Compton Scattering (DVCS)





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Link between GPDs and GFFs  $\int dx x \mathcal{H}^{q}(x,\xi,t) = \mathcal{A}^{q}(t) + 4\xi^{2} C^{q}(t)$   $\int dx x \mathcal{E}^{q}(x,\xi,t) = \mathcal{B}^{q}(t) - 4\xi^{2} C^{q}(t)$ 

🖾 Ji (1997), 🖾 Goeke (2001)

### Deeply Virtual Compton Scattering (DVCS)



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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Experiments
Nucleon EOS
3D distribution
factors
Gravitational form

Nonperturbative

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- Strategy
- CFF global fit
- Pressure forces
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- uncertainties

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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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factors				
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- A - E - N



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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### Exclusive processes of current interest. Factorization, universality and need for high luminosity.



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### Compton Form Factors. DVCS amplitude in the Bjorken regime.



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### Bjorken regime : large ${\it Q}^2$ and fixed ${\it xB}\simeq 2\xi/(1+\xi)$

- Partonic interpretation relies on factorization theorems.All-order proofs for DVCS.
- GPDs depend on a (arbitrary) factorization scale  $\mu_{F}$ .
  - **Consistency** requires the study of **different channels**.

GPDs enter DVCS through **Compton Form Factors** :

$$\mathcal{F}(\xi, t, Q^2) = \int_{-1}^{1} \mathrm{d}x \, T\left(x, \xi, \alpha_{\mathcal{S}}(\mu_F), \frac{Q}{\mu_F}\right) F(x, \xi, t, \mu_F)$$

for a given GPD F.

Kernels T derived at NLO and (partially) NNLO.
Relitsky and Willo

🛆 Belitsky and Müller (1998)

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\land Braun et al. (2022)

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• CFF  $\mathcal{F}$  is a **complex function**.

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## Phenomenology

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### What is the proton internal pressure? Identifying the concepts.





# What is the proton internal pressure? Identifying the concepts.





# What is the proton internal pressure? Identifying the concepts.





# What is the proton internal pressure? Identifying the concepts.





### Almost all existing DVCS data sets. 2600+ measurements of 30 observables published during 2001-17.



Nucleon	No.	Collab.	Year	Ref.	Observa	ble	Kinematic dependence	No. of points used / all
internal	1	HERMES	2001	40	$A_{LU}^+$		$\phi$	10 / 10
pressure	2		2006	41	$A_C^{\cos i\phi}$	i = 1	t	4/4
	3		2008	42	$A_C^{\cos i\phi}$	i = 0, 1	$x_{\rm Bj}$	18 / 24
					$A_{UT, DVCS}^{\sin(\phi - \phi_S) \cos i\phi}$	i = 0		
					$A_{UT I}^{\sin(\phi - \phi_S) \cos i\phi}$	i = 0, 1		
Energy-					$A_{UT}^{\cos(\phi-\phi_S)\sin i\phi}$	i = 1		
momentum	4		2009	43	$A_{LU,I}^{\sin i\phi}$	i = 1, 2	$x_{\rm Bj}$	35 / 42
tensor				-	$A_{LU,DVCS}^{\sin i\phi}$	i = 1		
Gravitational form				_	$A_C^{\cos i\phi}$	i=0,1,2,3		
factors	5		2010	44	$A_{UL}^{+,\sin i\phi}$	i=1,2,3	$x_{\rm Bj}$	18 / 24
2D distribution				_	$A_{LL}^{+,\cos i\phi}$	i=0,1,2		
SD distribution	6		2011	45	$A_{LT,DVCS}^{\cos(\phi-\phi_S)\cos i\phi}$	i = 0, 1	$x_{\rm Bj}$	24 / 32
Nucleon EUS					$A_{LT,DVCS}^{\sin(\phi-\phi_S)\sin i\phi}$	i = 1		
Experiments					$A_{LT,I}^{\cos(\phi-\phi_S)\cos i\phi}$	i = 0, 1, 2		
Dhamanalana					$A_{LT,I}^{\sin(\phi-\phi_S)\sin i\phi}$	i = 1, 2		
Phenomenology	7		2012	46	$A_{LU,I}^{\sin i\phi}$	i = 1, 2	$x_{\rm Bj}$	35 / 42
Strategy					$A_{LU,DVCS}^{\sin i\phi}$	i = 1		
CFF global fit				_	$A_C^{\cos i\phi}$	i=0,1,2,3		
Pressure forces	8	CLAS	2001	47	$A_{LU}^{-,\sin i\phi}$	i = 1, 2	_	0 / 2
Models: systematic	9		2006	48	$A_{UL}^{-, \text{on } i\psi}$	i = 1, 2		2 / 2
uncertainties	10		2008	49	ALU A		φ	283 / 737
	19		2009	51	$A^ A^ A^-$		φ	22 / 33
Areas for	13		2015	52	$d^4\sigma_{TTT}$		$\phi$	1333 / 1933
improvement	14	Hall A	2015	34	$\Delta d^4 \sigma_{LH}^-$		ф ф	228 / 228
CEE fitz	15		2017	35	$\Delta d^4 \sigma_{LU}^{LU}$		$\phi$	276 / 358
	16	COMPASS	2018	36	$d^3 \sigma_{UU}^{\pm}$		t	2/4
GFF t-profile	17	ZEUS	2009	37	$d^3 \sigma^+_{UU}$		t	4 / 4
Isolating d <sub>1</sub>	18	H1	2005	38	$d^3\sigma^+_{UU}$		t	7/8
Constanting	19		2009	39	$d^{a}\sigma_{UU}$		t	12 / 12
Conclusion							SUM:	2624 / 3996
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### Almost all existing DVCS data sets. 2600+ measurements of 30 observables published during 2001-17.





Abbreviations

also as close as possible to the valence region. 15 / 33

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### Modeling of $\mathcal{H}$ , $\widetilde{\mathcal{H}}$ , $\mathcal{E}$ and $\widetilde{\mathcal{E}}$ . Independent descriptions of real and imaginary parts.



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- Gravitational form factors 3D distribution
- Nucleon EOS
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Pressure forces Models: systematic uncertainties

#### Areas for improvement

CFF fits GFF t-profile Isolating d<sub>1</sub>

Conclusion

- Real and imaginary parts of CFFs parameterized by neural networks.
- Propagation of uncertainties through replica method and evaluation of 68 % confidence levels.





### Pressure forces from DVCS measurements. A first-principle connection.



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Abbreviations

1 Expand D-term on Gegenbauer polynomials

$$D_{\text{term}}^{q}(z, t, \mu_{F}^{2}) = (1 - z^{2}) \sum_{\text{odd } n} d_{n}^{q}(t, \mu_{F}^{2}) C_{n}^{3/2}(z)$$

**2** Write dispersion relation for CFF (true at all pQCD orders)

$$\mathcal{C}_{\mathcal{H}}(t, Q^2) = \operatorname{Re}\mathcal{H}(\xi) - \frac{1}{\pi} \int_0^1 \mathrm{d}\xi' \operatorname{Im}\mathcal{H}(\xi') \left(\frac{1}{\xi - \xi'} - \frac{1}{\xi + \xi'}\right)$$

3 Compute subtraction constant

$$\mathcal{C}_{H}^{q,g}(t,Q^2) = \frac{2}{\pi} \int_{1}^{+\infty} \mathrm{d}\omega \operatorname{Im} T^{q,g}(\omega) \int_{-1}^{1} \mathrm{d}z \, \frac{D^{q,g}(z)}{\omega - z}$$

🛆 Diehl & Ivanov (2007)

4 Retrieve GFF

$$d_1^q(t,\mu_F^2) = 5C_q(t,\mu_F^2)$$

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### Pressure forces from DVCS measurements. A first-principle connection.



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**2** Write dispersion relation for CFF (true at all pQCD orders)

$$\mathcal{C}_{\mathcal{H}}(t, Q^2) = \operatorname{Re}\mathcal{H}(\xi) - \frac{1}{\pi} \int_0^1 \mathrm{d}\xi' \operatorname{Im}\mathcal{H}(\xi') \left(\frac{1}{\xi - \xi'} - \frac{1}{\xi + \xi'}\right)$$

**3** Compute subtraction constant at LO

$$\mathcal{C}_{H}(t, Q^{2}) = 4 \sum_{q} e_{q}^{2} \sum_{\text{odd } n} d_{n}^{q}(t, \mu_{F}^{2} \equiv Q^{2})$$

⁄ Diehl & Ivanov (2007)



 $d_1^q(t,\mu_F^2) = 5C_q(t,\mu_F^2)$ 

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### Subtraction constant from measurements. EIC prospect: determination over a wide kinematic domain.

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### Pressure forces from DVCS measurements. Working assumptions.



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#### Conclusion

Abbreviations

**1** Subtraction constant assumed equal to  $d_1$ .

- **2** Equal values for light quark contributions  $d_1^{uds}$ .
- Radiative generation of gluon d<sup>g</sup><sub>1</sub> and charm d<sup>c</sup><sub>1</sub> contributions.
- **4** Tripole Ansatz for the *t*-dependence of  $d_1$ .



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🛆 Dutrieux et al.



### Pressure forces from DVCS measurements. Working assumptions.



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- **4** Tripole Ansatz for the *t*-dependence of  $d_1$ .



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µ<sup>2</sup> [GeV<sup>2</sup>]





## ▲ Dutrieux et al. (2021)

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### Pressure forces from DVCS measurements. Working assumptions.



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- **4** Tripole Ansatz for the *t*-dependence of  $d_1$ .

### Summary of existing determinations

No.	Marker in Fig. 3	$\sum_{q} d_{1}^{q}(\mu_{F}^{2})$	$_{\rm in~GeV^2}^{\mu_{\rm F}^2}$	# of flavours	Type	Ref
1	0	$-2.30 \pm 0.16 \pm 0.37$	2.0	3	from experimental data	[13
2		$0.88 \pm 1.69$	2.2	2	from experimental data	[14
3	0	-1.59	4	2	<i>t</i> -channel saturated model	[55
		-1.92	4	2	<i>t</i> -channel saturated model	55
4		-4	0.36	3	$\chi QSM$	30
5	$\nabla$	-2.35	0.36	2	$\chi QSM$	[10
6		-4.48	0.36	2	Skyrme model	56
7	Ħ	-2.02	2	3	LFWF model	[57
8	$\otimes$	-4.85	0.36	2	$\chi QSM$	]58
9	$\oplus$	$-1.34 \pm 0.31$	4	2	lattice QCD (MS)	[59
		$-2.11 \pm 0.27$	4	2	lattice QCD (MS)	[59

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■ No justification to truncate the subtraction constant expansion to its first term and assume that it is the *d*<sub>1</sub> coefficient related to the energy-momentum tensor.





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- No justification to truncate the subtraction constant expansion to its first term and assume that it is the  $d_1$ coefficient related to the energy-momentum tensor.
- Leading contributions of d<sub>1</sub> and d<sub>3</sub>, higher order terms neglected.





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#### Conclusion

- No justification to truncate the subtraction constant expansion to its first term and assume that it is the *d*<sub>1</sub> coefficient related to the energy-momentum tensor.
- Leading contributions of d<sub>1</sub> and d<sub>3</sub>, higher order terms neglected.
- 3 active quark flavors (*uds*), and radiative generation of *c* contribution.





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Abbreviations

No justification to truncate the subtraction constant **expansion to its first term** and assume that it is the  $d_1$ coefficient related to the energy-momentum tensor.

• Leading contributions of  $d_1$  and  $d_3$ , higher order terms neglected.

3 active quark flavors (uds), and radiative generation of c contribution.

4 parameters

 $d_1^{uds}(\mu_0) = d_3^{uds}(\mu_0) = d_1^g(\mu_0) = d_3^g(\mu_0)$ 

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### Pressure forces from DVCS measurements. From leading order to next-to-leading order.



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# 4 parameters $d_1^{uds}(0.1 \text{GeV}^2)$ $d_1^g(0.1 \text{GeV}^2)$ $d_3^{uds}(0.1 \text{GeV}^2)$ $d_3^g(0.1 \text{GeV}^2)$

### Investigate 3 fitting scenarios

	Pressure to From leading ord	rces from DV er to next-to-leading	cs measuremei	nts.
Nucleon internal pressure		Investigate 3	fitting scenarios	
Energy- momentum tensor Gravitational form factors 3D distribution Nucleon EOS Experiments	4 parameters $d_1^{uds}(0.1 \text{GeV}^2)$ $d_1^g(0.1 \text{GeV}^2)$	1. Nominal fit: 1 free p. Gluon a	arameter for light qua nd charm radiatively	arks. generated.
Phenomenology Strategy	$d_2^{uds}(0.1 \mathrm{GeV}^2)$		LO	NLO
CFF global fit Pressure forces	$d^{g}(0.1 \text{GeV}^{2})$	$d_1^{uds}(0.1 {\rm GeV}^2)$	$-0.7 \pm 1.7$	$-0.8 \pm 2.0$
Models: systematic uncertainties	<i>u</i> <sub>3</sub> (0.10ev)	$d_1^{uds}(2 \text{GeV})$	$-0.5\pm1.2$	$-0.5 \pm 1.4$
Areas for		$d_1^{g}(2 \text{GeV})$	$-0.6\pm1.6$	$-0.7\pm1.9$
CFF fits	Free	$d_1^{\rm C}(2{\rm GeV})$	$-0.002 \pm 0.0005$	$-0.002 \pm 0.006$
Isolating d <sub>1</sub>	Fixed			
Conclusion			Dutrieux et al.,	in preparation
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### Pressure forces from DVCS measurements. From leading order to next-to-leading order.



Nucleon internal pressure

4 parameters

 $d_1^{uds}(0.1 \,{\rm GeV}^2)$ 

 $d_3^{uds}(0.1 \,{\rm GeV}^2)$ 

 $d_1^{g}(0.1 \, \text{GeV}^2)$ 

 $d_3^{g}(0.1 \,{\rm GeV}^2)$ 

Free

Fixed

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### Investigate 3 fitting scenarios

- 2. Alternative fit:
  - 2 free parameters: light quarks and gluons.
  - Charm radiatively generated.

	LO	NLO
$d_1^{uds}(0.1 {\rm GeV}^2)$	$-6.2 \pm 14$	$-0.4\pm2.3$
$d_1^{\mathrm{g}}(0.1\mathrm{GeV}^2)$	$68 \pm 152$	$6.3\pm22$
$d_1^{uds}(2 {\rm GeV})$	$-0.7\pm1.2$	$0.4 \pm 2.8$
$d_1^{g}(2{ m GeV})$	$51\pm111$	$5.3\pm19$
$d_1^c(2 {\rm GeV})$	$0.2 \pm 0.4$	$0.02\pm0.06$

### Dutrieux et al., in preparation

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### Pressure forces from DVCS measurements. From leading order to next-to-leading order.



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Free
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4 parameters

 $d_1^{uds}(0.1 \,{\rm GeV}^2)$ 

 $d_{3}^{uds}(0.1 \,{\rm GeV}^2)$ 

 $d_1^{g}(0.1 \, \text{GeV}^2)$ 

 $d_{3}^{g}(0.1 \,{\rm GeV}^2)$ 

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#### Abbreviations

### Investigate 3 fitting scenarios

- 2. Alternative fit:
  - 2 free parameters: light quarks and gluons.
  - Charm radiatively generated.



Decorrelation of  $d_1^g$  and  $d_1^{uds}$  at NLO.

# Dutrieux *et al.*, *in preparation*

Correlation at NLO

R = 0.23043

m = 1.73655

b = 7.15348

10.0

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### Pressure forces from DVCS measurements. From leading order to next-to-leading order.



Nucleon internal		Investigate 3	fitting scenario	os
pressure		3. Alternative	fit:	
Energy- momentum tensor Gravitational form	4 parameters	<ul><li>2 free particular de la construcción de la</li></ul>	arameters: $d_1$ and not charm radiative	$d_3$ for light ely generated.
3D distribution	$d_1^{uds}(0.1 { m GeV}^2)$		LO	NLO
Experiments	$d_1^{g}(0.1 {\rm GeV}^2)$	$d_1^{uds}(0.1 \text{GeV}^2)$	$16 \pm 37$	$15 \pm 34$
Phenomenology Strategy	$d^{uds}(0.1 \text{ CoV}^2)$	$d_3^{uds}(0.1 \text{GeV}^2)$	$-26\pm59$	$-18 \pm 39$
CFF global fit	$u_3 (0.1 \text{GeV})$	$d_1^{uds}(2 \text{GeV})$	$11 \pm 25$	$11 \pm 23$
Models: systematic uncertainties	$d_3^{g}(0.1 {\rm GeV}^2)$	$d_1^{g}(2 \text{GeV})$	$15 \pm 34$	$15 \pm 32$
Areas for		$d_1^c(2 \text{GeV})$	$-0.05\pm0.1$	$-0.05\pm0.1$
CFF fits		$d_3^{uds}(2 \text{GeV})$	$-11\pm26$	$-7.7\pm17$
GFF t-profile	Free	$d_3^{g}(2 \text{GeV})$	$-1.8 \pm 3.9$	$-1.2 \pm 2.6$
Conclusion	Fixed	$d_3^{c}(2 {\rm GeV})$	$-0.04\pm0.01$	$-0.003 \pm 0.007$
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### Pressure forces from DVCS measurements. From leading order to next-to-leading order.



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Abbreviations

# $d_1^{uds}(0.1 \,{\rm GeV}^2)$ $d_1^{g}(0.1 \,{\rm GeV}^2)$ $d_3^{uds}(0.1 \,{\rm GeV}^2)$ $d_3^{g}(0.1 \,{\rm GeV}^2)$

Free

Fixed

# 4 parameters

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**Strong correlation of**  $d_1^{uds}$  and  $d_3^{uds}$ 

LO and NLO.

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both at

## Investigate 3 fitting scenarios

- 3. Alternative fit:
  - 2 free parameters: d<sub>1</sub> and d<sub>3</sub> for light quarks.
  - Gluon and charm radiatively generated.





Correlation at NLO

### Areas for improvement

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### Increase the physics input in the global fit. An example of the bias-variance trade-off.



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So far the CFF fit gathering most of the world DVCS measurements relies on an independent modeling of the CFF real and imaginary parts by neural networks.

• Convenient because of the **dimensionality** of the problem but yields **large statistical uncertainties**.

🛆 Moutarde *et al*. (2019)

 Conversaly a fit to the same data with a physically motivated parameterization still required *ad hoc* assumptions.

🛆 Moutarde *et al*. (2018)

 Many first-principle constraints expressed at the GPD level are not implemented at the CFF level.

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### Increase the physics input in the global fit. An example of the bias-variance trade-off.





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## Cez

### Increase the physics input in the global fit. An example of the bias-variance trade-off.



Nucleon internal pressure

Energy-

tensor Gravitational form

factors 3D distribution

momentum

Nucleon EOS

- Next step requires a (challenging) GPD global fit to world data.
- On the long run, need more experimental data to
  - Increase the *Q*<sup>2</sup>-lever arm.
    - Provide a better handle on the real part of  $\mathcal{H}$ .
    - Improve the accuracy of existing measurements.
    - Probe the kinematic regions insufficiently constrained.



Cea

### Relax modeling assumptions on $d_1(t)$ . Shape of pressure distribution not set by the current fit...



Nucleon internal pressure

#### Energymomentum tensor

Gravitational form factors 3D distribution

Nucleon EOS

Experiments

#### Phenomenology

Strategy

CFF global fit

Pressure forces

Models: systematic uncertainties

### Areas for improvement

CFF fits

GFF t-profile

Isolating  $d_1$ 

#### Conclusion

Abbreviations

- $d_1(t,\mu_F) = \frac{d_1(\mu_F)}{\left(1 \frac{t}{\Lambda^2}\right)^{\alpha}}$  Remind  $d_1^q(t,\mu_F^2) = 5C_q(t,\mu_F^2).$
- Plug in pressure anisotropy

Use multipole Ansatz

$$\frac{5(r)}{M} \propto \int \frac{\mathrm{d}^3 \vec{\Delta}}{(2\pi)^3} \, e^{-i\vec{\Delta} \cdot \vec{r}} \left\{ -\frac{4}{r^2} \frac{t^{-1/2}}{M^2} \frac{\mathrm{d}^2}{\mathrm{d}t^2} \left( t^{5/2} \, d_1(t) \right) \right\}$$

- Normalization d<sub>1</sub>(µ<sub>F</sub>) set by fit.
- Position of node in *r* depends on Λ.
- 🛆 Dutrieux et al. (2021)



C02

### Relax modeling assumptions on $d_1(t)$ . Shape of pressure distribution not set by the current fit...



#### Nucleon internal pressure

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Conclusion

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- Normalization set by fit.
- Position of node in r depends on Λ.

### \land Dutrieux et al. (2021)

- **Asymptotic** information on |t|-dependence from perturbative QCD. *But how large is "asymptotic"*?
- Factorization constraint: Q<sup>2</sup> ≫ |t|. Most of the experimental data used as fit input has low |t|.
- Need for more experimental data points.





### Increase $Q^2$ -lever arm. Evolution equations bring a slow $\log Q^2$ dependence.



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Isolating  $d_1$ 

Conclusion

Abbreviations

### Remind computation of subtraction constant at LO

$$\mathcal{C}_{H}(t, Q^{2}) = 4 \sum_{q} e_{q}^{2} \sum_{\text{odd } n} d_{n}^{q}(t, \mu_{F}^{2} \equiv Q^{2})$$

🛆 Diehl & Ivanov (2007)

 Plug LO evolution of D-term to obtain the following pattern

$$\mathcal{C}_{H}(t, Q^{2}) \propto \sum_{\text{odd } n} d_{n}(t, \mu_{F}) \left( \frac{\alpha_{s}(Q^{2})}{\alpha_{s}(\mu_{F}^{2})} \right)^{\gamma_{n}}$$

with  $\gamma_n$  computed in perturbative QCD. Since  $\alpha_s(Q^2) \propto 1/\log Q^2$ , an exact knowledge of  $\mathcal{C}_H(t,Q^2)$  on an  $Q^2$ -interval allows to exactly retrieve  $d_n$ .



# Increase Q<sup>2</sup>-lever arm.

Anomalous dimensions  $\gamma_n$  are small and take comparable values.



Nucleon internal pressure Introduce evolution operator  $\Gamma$  so that  $d_n(\mu_1) = \Gamma_n(\mu_1, \mu_2) d_n(\mu_2)$ 

#### Energymomentum tensor

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GFF t-profile

#### Conclusion

Abbreviations

Γ<sub>1</sub> and Γ<sub>3</sub> are numerically very close.





• Experimental data mostly constrain  $d_1 + d_3 + \dots$ 

▲ Dutrieux (et) al → (2021) C
H. Moutarde | Revealing emergent mass | 27 / 33



### Anomalous dimensions $\gamma_n$ and $Q^2$ -lever arm. Inverse problem and regularization.

Remind pattern of the problem



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Conclusion

Abbreviations

If  $Q^2$ -range is too small, a solution with  $d_1(t, \mu_F) + d_3(t, \mu_F) + d_5(t, \mu_F) + \ldots = 0$  can remain hidden within experimental uncertainties over the whole range  $Q^2 \in [Q^2_{\min}, Q^2_{\max}]$ .

 $\mathcal{C}_{H}(t, Q^{2}) \propto \sum_{\text{odd } p} d_{n}(t, \mu_{F}) \left( \frac{\alpha_{s}(Q^{2})}{\alpha_{s}(\mu_{F}^{2})} \right)^{n}$ 

- In practice: act as if the problem of retrieving  $d_1, d_3, \ldots$ from measurements has infinitely many solutions.
- Add extra regularization to select one solution robust with respect to statistical uncertainties.
- Today cannot reliably estimate the uncertainty associated to the neglect of  $d_3, \ldots$



### Conclusion

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### Conclusion and prospects. The phenomenological quest towards proton internal pressure.



Nucleon internal pressure

#### Energymomentum tensor

- Gravitational form factors 3D distribution Nucleon EOS
- Experiments

#### Phenomenology

- Strategy
- CFF global fit
- Pressure forces
- Models: systematic uncertainties

#### Areas for improvement

CFF fits GFF t-profile Isolating d<sub>1</sub>

Conclusion

- Concept well-defined and suitable for phenomenology.
- Strong first-principle connection between concept and experimental data.
- Need for multi-channel analysis beyond LO on a wide kinematic coverage. EIC and EIcC much needed!
- The GPD deconvolution problem is ill-posed. Huge sensitivity to numerical noise or experimental uncertainties.



gg75478317 GoGraph.com

- Development of the software ecosystem PARTONS for 3D hadron structure studies.
- Need for coordinated effort involving fits, computing chains, continuum and lattice QCD to make the best from experiments.
  - H. Moutarde | Revealing emergent mass | 30 / 33

### **Abbreviations**

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# Abbreviations used in this presentation.



Nucleon internal pressure	ANN CFF DDVCS	artificial neural network Compton form factor double deeply virtual Compton scattering
Energy- momentum tensor Gravitational form factors 3D distribution Nucleon EOS Experiments Phenoenology Strategy CFF global fit Pressure forces Models: systematic uncertainties Areas for improvement CFF fits GFF t-profile tealating d <sub>1</sub> Conclusion	DVCS DVMP DR EIC EFF GFF GFF GPD LO NLO PDF TCS	deeply virtual Compton scattering deeply virtual meson production dispersion relation electron-ion collider electron-ion collider in China elastic form factor gravitational form factor generalized parton distribution leading order next-to-leading order parton distribution function timelike Compton scattering
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