Recent advancements in radiation detectors for precision experiments

(and what do we use them for)

Suitable detector(s) for each radiation



What do we want (could) measure...



... and how



NOT INTERESTING FOR OUR RANGES

Photon absorption is a binary process

No energy loss like for charged particles



What do we want (could) measure...



Suitable detector(s) for each radiation





Semiconductor detectors



Semiconductor detectors



Physicians must be smart and clever....



Semiconductor detectors



Semiconductor detectors



Perfect linearity and easy calibration

Large area and geometrical efficiency

Fast readout for triggering

Suitable for 4-20 keV (450 µm thickness)

Resolution limited to ~120 eV





Silicon Drift Detectors for kaonic atoms



$$E_{1s} \simeq m_{red} c^2 \frac{\alpha^2 Z^2}{2}$$

$$n \simeq \sqrt{\frac{m_{red}}{m_e}} n_e$$





SIDDHARTA Collaboration / Physics Letters B 704 (2011) 113–117



Thicker SDDs could be used to extend the working range

This is not trivial and there are technological limitations to overcome Measurement of 1s level shift and width of KHe is one of the most wanted measurement in our community but the transition is expected at ~30 keV



"Nuclear E2 resonance effects in kaonic molybdenum isotopes" – Symposium, A. Scordo, Frascati (online), 08/04/2022



First XRF tests with known targets show very promising results

Efficiency @ 60 keV is increased of 100%

1-2 mm SDDs already financed by INFN CSN3

800µm and 1mm SDDs prototypes already produced by FBK for ARDESIA (INFN)

CdZnTe detectors



CZT: first tests @ DAΦNE

Goal: background and resolution assessment in machine environment (first time)

SIDDHARTA-2 Luminosity Monitor





22/06/2022:

First prototype installed in DAΦNE

Promising results obtained ON BEAM

First technical paper submitted

New opportunities for kaonic atoms measurements from CdZnTe detectors

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CZT: proposal for new measurements at $DA\Phi NE$

E. Friedman et al. / Nuclear Physics A579 (1994) 518-538

Table 1			
Compilation	of K ⁻	atomic data	a

Nucleus	Transition	e (keV)	Γ (keV)	Y	Γ_{μ} (eV)
He $3 \rightarrow 2$	3→2	-0.04 ± 0.03	~	<u> </u>	_
		-0.035 ± 0.012	0.03 ± 0.03	-	-
Li	3→2	0.002 ± 0.026	0.055 ± 0.029	0.95 ± 0.30	-
Be	3→2	-0.079 ± 0.021	0.172 ± 0.58	0.25 ± 0.09	0.04 ± 0.02
¹⁰ B	$3 \rightarrow 2$	-0.208 ± 0.035	0.810 ± 0.100	-	_
¹¹ B	$3 \rightarrow 2$	-0.167 ± 0.035	0.700 ± 0.080	-	-
С	3→2	~0.590 ± 0.080	1.730 ± 0.150	0.07 ± 0.013	0.99 ± 0.20
0	4 → 3	-0.025 ± 0.018	0.017 ± 0.014	-	-
Mg	4 → 3	-0.027 ± 0.015	0.214 ± 0.015	0.78±0.06	0.08 ± 0.03
Al	4 → 3	-0.130 ± 0.050	0.490 ± 0.160	-	-
		-0.076 ± 0.014	0.442 ± 0.022	0.55 ± 0.03	0.30 ± 0.04
Si $4 \rightarrow 3$	4 → 3	-0.240 ± 0.050	0.810 ± 0.120	-	-
		-0.130 ± 0.015	0.800 ± 0.033	0.49 ± 0.03	0.53 ± 0.06
P	4 → 3	-0.330 ± 0.08	1.440 ± 0.120	0.26 ± 0.03	1.89 ± 0.30
S	4 → 3	-0.550 ± 0.06	2.330 ± 0.200	0.22 ± 0.02	3.10 ± 0.36
		-0.43 ± 0.12	2.310 ± 0.170	_	_
		-0.462 ± 0.054	1.96 ± 0.17	0.23 ± 0.03	2.9 ± 0.5



With CdZnTe detectors the present database on kaonic atoms can be updated and renewed, and new important measurements can be done as well

HPGe detectors



High resolutions in a very wide energy range

Cryogenic cooling is needed

Subject to radiation damage

The maximum depletion depth for the planar detectors is limited to 1-2 cm.

5 cm is required for efficient detection of MeV photons.



HPGe detectors



Suitable detector(s) for each radiation

Bragg Spectrometers



Bragg spectrometers



Photons of different energies are reflected in different positions

With a crystal and a position detector, energy spectra with ultrahigh resolution can be obtained

For monochromatic sources, also directionality could be tested

Von Hamos geometry and mosaic crystals can improve collection efficiency





FWHM of few eV with NO COOLING

Energy range between 1-20 keV (n=1, depending on the crystal)

Extremely low efficiencies (solid angle)

Bragg spectrometers: VOXES

Spectrometer developed under CSN5 Young Researcher Grant (2016-2018)



Crystal spectrometers: VOXES

High precision measurements with VOXES in LNF Lab

≥2000 $S_0' = 0.54 \text{ mm}$ E(Cu Ka,) : 8047,90 +/- 0,19 eV **%1800** $\Delta \theta' = 0,27^{\circ}$ E(Cu Ka): 8027,48 +/- 0,36 eV Counts / 3. σ(Cu Kα,): 5,15 +/- 0,13 eV $\rho_c(mm)$ value (eV) $S_0'/\Delta\theta' (mm,^{\circ})$ Parameter 1600 $S_0 = 0 \text{ mm}$ E(Ni Kβ) : 8264,64 +/- 0,21 eV $\sigma(K\alpha_{1,2})$ $4,17\pm 0,16$ 0,3/0,24 $\Delta \theta = 0.2^{\circ}$ σ(Ni Kβ) : 6,02 +/- 0,24 eV 1400 $\delta(K\alpha_1)$ 0, 110,6/0,44E(Zn Ka,): 8639,07 +/- 0,52 eV $\delta(K\alpha_2)$ 0,18 0,6/0,44 $\rho = 103,4 \text{ mm}$ E(Zn Kα) : 8615,49 +/- 0,66 eV 1200 $4,05\pm 0,13$ 0,3/0,18 $\sigma(K\alpha_{1,2})$ σ(Zn Kα,): 6,20 +/- 0,34 eV $\delta(K\alpha_1)$ 0,09 0,7/0,341000F γ^2 /ndf : 1.34 $\delta(K\alpha_2)$ 0,7/0,34 0,13 $4,02 \pm 0,08$ $\sigma(K\alpha_{1,2})$ 1,1/0,60800 $\delta(K\alpha_1)$ 0.1 1,2/0,70 600 $\delta(K\alpha_2)$ 0,15 1,2/0,70 $6,8\pm 0,07$ 0,3/0,16 $\sigma(K\alpha_{1,2})$ 400 $\delta(K\alpha_1)$ 0,6/0,32 0.070,6/0,32 $\delta(K\alpha_2)$ 0,1200 $4,77 \pm 0,05$ 0,3/0,16 $\sigma(K\alpha_{1,2})$ 0년 $\delta(K\alpha_1)$ 0,04 0,7/0,328000 8200 8400 8600 $\delta(K\alpha_2)$ 0,7/0,320,07 X-ray energy (eV) $3,60 \pm 0,05$ 0,8/0,60 $\sigma(K\alpha_{1,2})$ Counts / 12.07 eV E(Mo Ka,): 17481,08 +/- 0,61 eV $S_0' = 1,6 \text{ mm}$ $\delta(K\alpha_1)$ 0,04 1,1/0,70 $\delta(K\alpha_2)$ 1,1/0,70 0.07 $\Delta \theta' = 0.8^{\circ}$ E(Mo Ka): 17366,47 +/- 2,02 eV 500 σ(Mo Kα,): 21,08 +/- 0,79 eV $5,15\pm 0,13$ 0,5/0,27 $\sigma(K\alpha_{1,2})$ $S_0 = 1,08 \text{ mm}$ $\delta(K\alpha_1)$ 0,6/0,22E(Nb Kβ) : 18622,42 +/- 1,42 eV 0,10 σ(Nb Kβ) : 36,95 +/- 1,32 eV $\Delta \theta = 0.65^{\circ}$ 0.210,6/0,22 $\delta(K\alpha_2)$ 400 χ^2 /ndf : 1,27 $\sigma(K\beta)$ $6,02 \pm 0,24$ 0,5/0,27 $\delta(K\beta)$ 0,6/0,22 $\rho = 77,4 \text{ mm}$ 0, 13 $6,20 \pm 0,34$ 0,5/0,27 $\sigma(K\alpha_{1,2})$ 300 $\delta(K\alpha_1)$ 0,26 0,6/0,220,6/0,22 $\delta(K\alpha_2)$ 0,42200 $21, 1 \pm 0, 8$ 1,6/0,80 $\sigma(K\alpha_{1,2})$ $\delta(K\alpha_1)$ 0.6 1,6/0,80 $\delta(K\alpha_2)$ 2,01,6/0,80 100 $36,9 \pm 1,3$ $\sigma(K\beta)$ 1,6/0,80 $\delta(K\beta)$ 1,3 1,6/0,800 18500 17000 17500 18000 19000 X-ray energy (eV)

Table 3 Best achieved resolutions and precisions summary.

Element

Fe

Cu

Cu

Ni

Zn

Mo

Nb

77,5

103,4

206,7

77,5

103,4

206,7

103,4

103,4

103.4

77.5

77,5

VOXES: applications





VOXES: applications in DAΦNE

A new setup including several spectrometer arms could allow for new and very precise measurements of kaonic atoms transitions both from solid and gaseous targets





Suitable detector(s) for each radiation

Transition Edge Sensors (TES): Microcalorimetry





Transition Edge Sensors

Photon absorption is used to rise the temperature of a thin film of superconducting material above the Tc

 ΔT is proportional to the photon energy which can be derived with extremely high accuracy a thermal detector measuring the energy of an incident x-ray photon as a temperature rise (= $E/C \sim 1 \text{ mK}$)





Transition Edge Sensors

Co Ka1

Fit w/o LE tail

Co Ka3,4

 $\langle \Delta E \rangle = 6.1 \text{ eV}$

Co Ka2

two-stage pulse tube (60K, 3K)

50 mK cryostat

(model : HPD 102 DENALI)

(double-stage salt pills : GGG 1K, FAA 50mK)

Ultra-High resolution (FHWM ~0,03% @ 5900 eV)

Acceptable geometrical efficiency (small active areas)

Extremely high costs Non-trivial calibration





(a)

Fit with LE tail

 $<\Delta E > = 5.2 \text{ eV}$

104

103

Counts / 1 eV

Transition Edge Sensors



What do we want (could) measure...



Strip & Pixel detectors

With strip detectors, 1D position spectra can be obtained

2D spectra are obtained from double sided strip detectors or Pixel Detectors (like CCD)





What do we want (could) measure...





Multi Pixel Photon Counters



No Cooling No radiation damage Working within magnetic fields

Few photons can be measured Visible photon range (some attempts with direct X-rays)

Saturation effects (non-linearity)

Signal "quantization" is even visible on an oscilloscope



Electron charge can be measured (for students)

Radiation detection technology is evolving very fast, and new experiments become feasible

Triggers for new experiments on fundamental physics are very welcome