# Radiative Corrections in Super-Rosenbluth Experiments

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ECT Workshop on "Radiative corrections from medium to high energy experiment"

#### July 19, 2022



**Electron-Proton Elastic Scattering** 





"Super Rosenbluth" Experiment



#### PhD on OLYMPUS

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Phys.Rev.Lett. 118, 092501 (2017)

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CLAS12 Backward Angle Neutron Detector

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- JLab Positron Working Group
  - See recent white paper

#### EPJ A Topical Issue: An Experimental Program with Positron Beams at Jefferson Lab

Published on 14 April 2022



The perspective of high energy and high duty cycle polarized positron beams in complement to the existing CEBAF (Continuous Electron Beam Accelerator Facility) electron beams is attracting a lot of interest. Following the proof-of-principle PEPPo (Polarized Electrons for Polarized Positrons)

Eur. Phys. J. A (2022)



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 e.g., Stanford ... Mainz A1



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- Super-Rosenbluth
  - e.g., JLab Hall A
     Qattan et al., (2005)



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#### Coincidence

- e.g., OLYMPUS
- Radiative generator + coincidence cuts
- Super-Rosenbluth
  - e.g., JLab Hall A Qattan et al., (2005)
  - RC Formulae?
  - Simulate + integrate over all e<sup>-</sup> phase space?

# A. Afanasev sent me his paper on Super-Rosenbluth RCs yesterday.

"QED radiative corrections to asymmetries of elastic *ep* scattering in hadronic variables" A. V. Afanasev et al., Phys.Lett.B 514, pp. 269-278 (2001)

- Appears to answer many of the questions I raise today.
- I have not coded up the analytic expressions yet but I will!



M. E. Christy et al. (JLab Hall C), PRC 70, 015206 (2004)



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I. A. Qattan et al., PRL 94, 142301 (2005)

- Experiment E01-001
- Beam energies: 1.91, 2.62, 2.84, 3.77, 4.70 GeV
- 4 cm liquid hydrogen target
- Q<sup>2</sup> = 2.64, 3.20, 4.10 GeV<sup>2</sup>
- Hall A High-Resolution Spectrometer (HRS)
  - $\mathbf{a}$   $\approx$  6 msr acceptance
  - $\delta p/p \approx 10^{-4}$





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## NE-18 Radiative Monte Carlo Generator

R. Ent et al., Phys. Rev. C 64, 054610 (2001)

- Adapted into SIMC, widely used at JLab
- Multi-photon (exponentiated), pure peaking approximation
- Loop corrections to  $\alpha^3$  (*non-exponentiated*)
- Randomly sample radiation from each particle, update kinematics



## My simulated pseudo experiment

Rosenbluth separations at 2.64, 3.20, and 4.10 GeV (same as Hall A)
 In-plane acceptance of ±1mrad (much smaller than Hall A)
 Assume std. dipole form factors



#### Tail-shape in simulation



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# While magnitude depends on kinematics, proton RCs tend to be flatter in $\Delta p$ .











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#### True

How about with the OLYMPUS generator?

## The OLYMPUS generator used two approaches.



- Distinguish between near-elastic and tail.
- near elastic:  $\frac{d\sigma}{d\Omega}_{\text{meas.}} = \frac{d\sigma}{d\Omega}_{\text{Born}} \times [1 + \delta(\Delta E)]$
- tail: tree-level bremsstrahlung cross section



## The OLYMPUS generator used two approaches.



- 2 Exponentiated approach
  - Based on prev. work by J. M. Friedrich, J. C. Bernauer at Mainz A1

## Exponentiated Approach

Assumptions:

- Multi-photon kinematics can be well-approximated by single-photon bremsstrahlung kinematics
- Differential cross section takes an exponentiated form:

$$d^5\sigma = rac{d\sigma}{d\Omega}_{\mathsf{Born}} e^{\delta} \left( \partial_{ec{
ho}_\gamma} \delta 
ight)$$

• The differential part of  $\delta$  is well-approximated

$$\partial_{\vec{p}_{\gamma}}\delta \longrightarrow \frac{d^{5}\sigma}{d\Omega_{e}d\Omega_{\gamma}E_{\gamma}}_{\text{Brems.}}/\frac{d\sigma}{d\Omega}_{\text{Born}}$$

•  $\delta$  given by standard prescription (e.g. Mo-Tsai)  $d^{5}\sigma = \frac{d^{5}\sigma}{d\Omega_{e}d\Omega_{\gamma}E_{\gamma}} e^{\delta(E_{\gamma})}$ 

#### Proton spectrum within the OLYMPUS generator



#### Proton spectrum within the OLYMPUS generator



#### Where are the electrons going?



The problem seems to come from "barely virtual" Compton scattering

Q<sup>2</sup>≈0



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Recap

- Super-Rosenbluth RCs have the challenge of integrating radiation "all the way down to zero."
- Within peaking framework, Super-Rosenbluth RCs are smaller/flatter.



- Super-Rosenbluth RCs have the challenge of integrating radiation "all the way down to zero."
- Within peaking framework, Super-Rosenbluth RCs are smaller/flatter.
- Numerical pitfalls for full bremss. cross section.



#### Conclusions

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- Until I get a second model working, no clue about the model dependence of SIMC approach in Qattan et al.
- Credit to A. Afanasev for already solving this back in 2001 (and thank you for sending me the paper!).