Dynamical phase transitions in collective neutrino oscillations

Alessandro Roggero



(OBC)

# Neutrino's roles in supernovae

• efficient energy transport away from the shock region (burst)



#### regulation of electron fraction in ν-driven wind (nucleosynthesis)



figures from Janka et al. (2007)energy deposition to revive the stalled shock (explosion)



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# Neutrino oscillations in astrophysical environments

We know that neutrinos can display flavor oscillations in vacuum, does it matter in a core-collapse supernova?

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## Neutrino oscillations in astrophysical environments

We know that neutrinos can display flavor oscillations in vacuum, does it matter in a core-collapse supernova?

- energy deposition behind shock and in the wind proceeds through charge-current reactions (large differences in  $\nu_e \nu_{\mu/\tau}$ )
- neutrino oscillation rates can get enhanced through elastic forward scattering with high density external matter (MSW effect)



### Neutrino-neutrino forward scattering

Fuller, Qian, Pantaleone, Sigl, Raffelt, Sawyer, Carlson, Duan, ...



- diagonal contribution (A) does not impact flavor mixing
- off-diagonal term (B) equivalent to flavor/momentum exchange between two neutrinos
  - total flavor is conserved

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Important effect if initial distributions are strongly flavor dependent



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**Coherent Neutrinos** 

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# Two-flavor approximation and the iso-spin Hamiltonian

Consider two active flavors  $(\nu_e, \nu_x)$  and encode flavor amplitudes for a neutrino with momentum  $p_i$  into an SU(2) iso-spin:

 $|\Phi_i\rangle = \cos(\eta_i)|\nu_e\rangle + \sin(\eta_i)|\nu_x\rangle \equiv \cos(\eta_i)|\uparrow\rangle + \sin(\eta_i)|\downarrow\rangle$ 

A system of  ${\cal N}$  interacting neutrinos is then described by the Hamiltonian

$$H = \sum_{i} \frac{\Delta m^2}{4E_i} \vec{B} \cdot \vec{\sigma}_i + \lambda \sum_{i} \sigma_i^z + \frac{\mu}{2N} \sum_{i < j} \left( 1 - \cos(\phi_{ij}) \right) \vec{\sigma}_i \cdot \vec{\sigma}_j$$

• vacuum oscillations:  $\vec{B} = (\sin(2\theta_{mix}), 0, -\cos(2\theta_{mix}))$ • interaction with matter: • neutrino-neutrino interaction: • dependence on momentum direction:  $\vec{B} = (\sin(2\theta_{mix}), 0, -\cos(2\theta_{mix}))$   $\lambda = \sqrt{2}G_F \rho_e$ • neutrino-neutrino interaction: • dependence on momentum direction:  $\mu = \sqrt{2}G_F \rho_{\nu}$ • dependence on momentum direction: • dependence on momentum direct

for a full derivation, see e.g. Pehlivan et al. PRD(2011)

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# Beyond mean field effects

# Beyond mean field effects

 increasing effort in tackling the problem using a variety of methods: diagonalization, tensor networks and semiclassical approaches
 Cervia et al. (2021), Patwardhan et al. (2021), AR (2021)<sup>2</sup>, Xiong (2022), Martin, AR, et al. (2022), AR, Rrapaj, Xiong (2022), Lacroix et al. (2022), ...

• Great potential for many-body simulations on quantum devices

Hall, AR, et al. (2021), Yeter-Aydeniz et al. (2022), Illa & Savage (2022), Amitrano, AR, et al. (2022)

### Dynamical phase transitions

Heyl et al. PRL (2013), Heyl PRL (2015), Heyl RPP (2018)

#### Quantum quench protocols

**(**) the system starts as the ground-state of an initial Hamiltonian  $H_0$ 

**2** at time t = 0 we switch to a different Hamiltonian H and evolve

Dynamical critical behavior encoded in Loschmidt echo

$$\mathcal{L}(t) = \left| \langle \Psi_0 | e^{-iHt} | \Psi_0 \rangle \right|^2 \xrightarrow{N \gg 1} e^{-N\lambda(t)}$$

Loschmidt rate  $\lambda(t)$  plays a similar role as the free energy in equilibrium.

$$H(h) = -\sum_{\langle ij\rangle} Z_i Z_j + h \sum_i X_i$$

- start in ground-state for  $h \to \infty$
- quench across critical point at h = 1



Heyl PRL (2015)

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Schmitt & Heyl SciPost Phys (2018)

# DPT for systems with degenerate ground spaces

Heyl PRL (2014)

$$H_{XXZ} = J \sum_{i} \left[ X_i X_{i+1} + Y_i Y_{i+1} + \Delta Z_i Z_{i+1} \right]$$

- $\bullet$  disordered gapless phase for  $\Delta < 1$
- $\bullet$  anti-ferromagnetic phase for  $\Delta>1$
- critical point at  $\Delta=1$

$$|\Psi_0\rangle = |\uparrow\downarrow\uparrow\downarrow\cdots\rangle \qquad |\Psi_0'\rangle = |\downarrow\uparrow\downarrow\uparrow\cdots\rangle$$

#### Loschmidt Echo for degenerate ground-states

$$\mathcal{L}_0(t) = \left| \langle \Psi_0 | e^{-itH} | \Psi_0 \rangle \right|^2 \quad \mathcal{L}_1(t) = \left| \langle \Psi'_0 | e^{-itH} | \Psi_0 \rangle \right|^2 \,,$$

 $\mathsf{DPT} \Leftrightarrow \mathsf{non-analytic} \text{ behavior of the total echo } \mathcal{L}(t) = \mathcal{L}_0(t) + \mathcal{L}_1(t)$ 

### DPT for systems with degenerate ground spaces II

$$\mathcal{L}_0(t) = \left| \langle \Psi_0 | e^{-itH} | \Psi_0 \rangle \right|^2 \quad \mathcal{L}_1(t) = \left| \langle \Psi'_0 | e^{-itH} | \Psi_0 \rangle \right|^2$$

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Both scale exponentially in system size, but with different rates, there is a kink forming if the order between  $\mathcal{L}_0(t)$  and  $\mathcal{L}_1(t)$  changes at some  $t = t^*$ 



### Simple neutrino model

Friedland & Lunardini (2003), AR (2021)

$$H = \frac{1}{2N} \sum_{i < j} \vec{\sigma}_i \cdot \vec{\sigma}_j = \frac{1}{N} S^2 + const.$$

Initialize system in  $|\Psi(0)\rangle = |\downarrow\rangle^{\otimes N/2} \otimes |\uparrow\rangle^{\otimes N/2}$  and compute the flavor persistence  $p(t) = (1 - \langle \Psi(t) | \sigma_1 | \Psi(t) \rangle)/2$  for increasing system size



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Simple neutrino model II

$$H(x) = \frac{x}{2N}S^2 + (1-x)\sum_{a \in \mathcal{A}}\sum_{b \in \mathcal{B}}Z_aZ_b ,$$

start at x = 0 and evolve with x = 1. State is  $|\Psi_0\rangle = |\downarrow\rangle^{\otimes N/2} \otimes |\uparrow\rangle^{\otimes N/2}$ .



Crossing time  $t^*$  diverges as  $\sqrt{N} \Rightarrow$  no evolution for a large system!

### Many-body speedup in unphysical model

Bell, Rawlinson, Sawyer PLB (2003), AR (2021)

$$H_{BRS} = \frac{1}{2N} \sum_{i < j} \mathcal{J}_{ij} \left( X_i X_j + Y_i Y_j + \Delta Z_i Z_j \right)$$

with  $\mathcal{J}_{ij} = J_{AA}$  for (i, j) in  $\mathcal{A}$  or  $\mathcal{B}$  and  $\mathcal{J}_{ij} = J_{AB}$  otherwise. Our initial state is (degenerate) gs of  $H_{BRS}$  in the limit  $\Delta \gg 1$  and  $J_{AA} < J_{AB}$ 



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## Many-body speedup in a physical model

To engineer a "DPT" we can ensure the system crosses a critical point

$$H = -\frac{\delta_{\omega}}{2} \left( \sum_{i \in \mathcal{A}} \sigma_i^z - \sum_{i \in \mathcal{B}} \sigma_i^z \right) + \frac{\mu}{2N} \sum_{i < j} \vec{\sigma}_i \cdot \vec{\sigma}_j ,$$

AFM ( $\mu > 0$ ) transition at  $\delta_{\omega} = 0$  between gapped phases FM ( $\mu < 0$ ) transitions at  $\delta_{\omega} = \pm \mu$  between gapped and gapless phases

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# Summary and perspectives

- collective neutrino oscillations are an interesting **strongly coupled** many-body system driven by the **weak interaction**
- there seems to be a strong connection between oscillation time scales and dynamical phase transitions
  - explains seemingly conflicting results from the past
  - alternative way of understanding the appearance of bipolar modes
- can be generalized to more complicated geometries
  - "fast" modes with three beams can also be understood in terms of a DPT almost identical to the one present in the two beam case
- useful to exploit semi-classical methods to have a better understanding of the DPT in these systems
- great system to explore with fast advancing quantum technologies, Hamiltonian is two-local but all-to-all  $\rightarrow$  best suited for trapped-ions

# Many-body speedup in a physical model II

To engineer a "DPT" we can ensure the system crosses a critical point

$$H = -\frac{\delta_{\omega}}{2} \left( \sum_{i \in \mathcal{A}} \sigma_i^z - \sum_{i \in \mathcal{B}} \sigma_i^z \right) + \frac{\mu}{2N} \sum_{i < j} \vec{\sigma}_i \cdot \vec{\sigma}_j ,$$



Phase diagram for the BRS model

$$M_{AB}^{Z} = \frac{1}{N} \langle Z_{A} Z_{B} \rangle \qquad M_{AB}^{XY} = \frac{1}{N} \left( \langle X_{A} X_{B} \rangle + \langle Y_{A} Y_{B} \rangle \right)$$

$$M_{AB}^{X} = 0 \qquad M_{AB}^{X} = 0 \qquad M_{AB}^{Y} = 0 \qquad M_{$$