## Searching for Light Bosons with King and Second-King Plots Optimized for $\mathrm{Li}^{+}$

Gordon W.F. Drake<br>University of Windsor<br>Windsor, Ontario, Canada<br>\section*{Collaborators}<br>Victor Marton (undergraduate student)*<br>Harvir Dhindsa (undergraduate student)*<br>PeiPei Zhang (Wuhan WIPM)<br>Zong-Chao Yan (UNB, WIPM)<br>Ken Baldwin (Australian National University)<br>Don Morton (Herzberg Institute for Astrophysics)<br>Aaron Bondy (Ph.D. Student)

Financial Support: NSERC and SHARCnet
ECT* Workshop
March 12, 2022

## OUTLINE

- Introduction and motivation.
- What is a King plot?
- Search for new physics due to an electron-neutron interaction.
- Definition of a Second-King plot for light ions such as $\mathrm{Li}^{+}$.
- Comprehensive search for the optimum King plot in $\mathrm{Li}^{+}$.
- Results and conclusions.


## Motivation

High accuracy atomic-physics measurements at low energy have the potential to complement or even extend what can be learned from high-energy particle experiments.

## Springer Handbook of atomic, Molecular, and Optical Physics: Chapters on Searches for New Physics

- Tests of Fundamental Physics (QED)

Authors: Peter Mohr and Eite Tiesinga
$-g-2$

- proton size
- Atomic Clocks and Constraints on the Fudamental Constants Authors: Savely Karshenboim, Victor Flambaum and Ekkehard Peik
- Searches for New Particles Including Dark Matter

Authors: Victor Flambaum and Yevgeny Stadnik

- Searches for New Physics

Author: Marianna Safronova
-Parity nonconservation in atoms
-Electron electric dipole moment (edm)
-Tests of CPT symmetry (particles vs. antiparticles)
-Lorentz invariance
-Tests of General Relativity

## King Plots

- Fundamental searches are made possible by very high precision measurements of isotope shifts in sequences of isotopes.
- Optical clock transitions achieved at the few Hz level [Manovitz et al. PRL 123, 203001 (2019)], and the possibility of measurements at the mHz level using entangled states and coherent optical spectroscopy [Micke et al., Nature 578, 60 (2020)].
- Recent King-plot measurements in Yb may indicate evidence for an electronneutron interaction beyond the Standard Model [Counts et al. PRL 125, 123002 (2020)]. Other measurements in Ca and Sr .
- Analytical methods extensively developed by the NSW group [Berengut et al. Phys. Rev. Res. 2, 043444 (2020)]. See also Müller et al. PRA 104, L020802 (2021).
- The goal of this work is to extend King-plot methods to light atoms where secondorder mass polarization effects are important. See Drake, Dhindsa and Marton, Phys. Rev. A 104, L060801 (2021).
- All five isotopes ${ }^{6} \mathrm{Li},{ }^{7} \mathrm{Li},{ }^{8} \mathrm{Li},{ }^{9} \mathrm{Li}$ and ${ }^{11} \mathrm{Li}$ have been extensively studied and nuclear charge radii determined from the isotope shifts [see Nörtershäuser et al. PRA 83, 012516 (2011) and Lu et al. RMP 85, 1383 (2013)].
- For $\mathrm{Li}^{+}$high precision QED theory of fine and hyperfine structure is available. See Qi et al. Phys. Rev. Lett. 125, 183002 (2020). (Recall Wei Sun's talk from yesterday.)


# King and second-King plots with optimized sensitivity for lithium ions 

G. W. F. Drake © , Harvir S. Dhindsa, and Victor J. Marton © ${ }^{*}$<br>Department of Physics, University of Windsor, Windsor, Ontario, Canada N9B 3P4

(0) (Received 16 September 2021; accepted 18 November 2021; published 6 December 2021)

King plots constructed from combinations of mass-weighted atomic isotope shifts provide a sensitive technique to search for electron-neutron interactions beyond the standard model mediated by a light boson. Using high-precision variational wave functions in Hylleraas coordinates, we present a comprehensive survey of all possible King plots arising from states of $\mathrm{Li}^{+}$up to principal quantum number $n=10$ and angular momentum $L=7$ in order to identify the ones most sensitive to new physics. A major limitation in previous work due to second-order mass polarization is eliminated by the introduction of a second-King plot defined in terms of second differences. The residual theoretical uncertainty is then of the order of $\alpha^{2}(\mu / M)^{3} \sim 0.4 \mathrm{~Hz}$, where $\alpha$ is the fine-structure constant and $\mu$ is the reduced electron mass for a nucleus of mass $M$. Test results are presented for the ${ }^{A} \mathrm{Li}^{+}$isotope sequence with $A=6,7,8,9,11$ and are compared with other methods, including the $\mathrm{Yb}^{+}$ case recently studied both experimentally and theoretically. It is shown that the second-King plots for $\mathrm{Li}^{+}$have about the same sensitivity to new physics as the $\mathrm{Yb}^{+}$case for boson masses up to about 10 keV , and nuclear size uncertainties (including nuclear polarization) are suppressed. This greatly extends the sensitivity to new physics for light two-electron systems.

DOI: 10.1103/PhysRevA.104.L060801

## What is a King Plot?

- First introduced by W. H. King J. Opt. Soc. Am. 53, 638 (1963) also W. H. King, Isotope Shifts in Atomic Spectra (Springer Science \& Business Media, 2013).
- Provides a systematic method to look for systematic trends in atomic isotope shifts.
- Basic idea: partition the total isotope shift into a mass shift (atomic structure) and a field shift (nuclear size and structure).
- Let $i$ and $j$ denote two atomic states, and $a$ a particular isotope. The transition frequency between states $i$ an $j$ can then be expanded in the form

$$
\begin{equation*}
\nu_{a}^{i j}=U^{i j}+\underbrace{V^{i j}\left(\frac{\mu}{M}\right)_{a}+W^{i j}\left(\frac{\mu}{M}\right)_{a}^{2}+\cdots}_{\text {mass shift }}+\underbrace{C^{i j} \overline{r_{a}^{2}}}_{\text {field shift }} \tag{1}
\end{equation*}
$$

where
$U^{i j}=\left(E^{i}-E^{j}\right) / \hbar$ for an infinitely heavy point nucleus (incl. relativistic and QED)
$V^{i j}=$ normal and specific mass shifts (incl. relativistic and QED)
$W^{i j}=$ second-order mass shifts (incl. relativistic and QED)
$C^{i j}=\frac{2 \pi Z e^{2}}{3} \sum_{k}\left[\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{i}-\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{j}\right]$
$\overline{r_{a}^{2}}=$ mean square nuclear charge radius

- The isotope shift between two different isotopes $a$ and $b$ is then

$$
\nu_{a}^{i j}-\nu_{b}^{i j}=\underbrace{V^{i j}\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}\right]+W^{i j}\left[\left(\frac{\mu}{M}\right)_{a}^{2}-\left(\frac{\mu}{M}\right)_{b}^{2}\right]+\cdots}_{\text {mass shift }}+\underbrace{C^{i j}\left(\overline{r_{a}^{2}}-\overline{r_{b}^{2}}\right)}_{\text {field shift }}
$$

Divide by $\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}$ to obtain the King coordinate

$$
\begin{aligned}
n_{a b}^{i j} & =\frac{\nu_{a}^{i j}-\nu_{b}^{i j}}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}} \\
& =V^{i j}+W^{i j}\left[\left(\frac{\mu}{M}\right)_{a}+\left(\frac{\mu}{M}\right)_{b}\right]+\cdots+\frac{C^{i j}\left(\overline{r_{a}^{2}}-\overline{r_{b}^{2}}\right)}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}}
\end{aligned}
$$

See example King plot.

- A third isotope $c$ defines a King slope

$$
\begin{aligned}
S_{b c} & =\frac{n_{a c}^{i^{\prime} j^{\prime}}-n_{a b}^{i i^{\prime} j^{\prime}}}{n_{a c}^{i j}-n_{a b}^{i j}} \\
& \rightarrow \frac{C^{i^{\prime} j^{\prime}}}{C^{i j}} \text { if } W^{i^{\prime} j^{\prime}}=W^{i j}=0 \\
& \equiv \frac{\sum_{k}\left[\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{i^{\prime}}-\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{j j}\right]}{\sum_{k}\left[\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{i}-\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{j}\right]}
\end{aligned}
$$

- If $c$ is replaced by a fourth isotope $d$, then there should be no change in slope.

King Plot


- The isotope shift between two different isotopes $a$ and $b$ is then

$$
\nu_{a}^{i j}-\nu_{b}^{i j}=\underbrace{V^{i j}\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}\right]+W^{i j}\left[\left(\frac{\mu}{M}\right)_{a}^{2}-\left(\frac{\mu}{M}\right)_{b}^{2}\right]+\cdots}_{\text {mass shift }}+\underbrace{C^{i j}\left(\overline{r_{a}^{2}}-\overline{r_{b}^{2}}\right)}_{\text {field shift }}
$$

Divide by $\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}$ to obtain the King coordinate

$$
\begin{aligned}
n_{a b}^{i j} & =\frac{\nu_{a}^{i j}-\nu_{b}^{i j}}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}} \\
& =V^{i j}+W^{i j}\left[\left(\frac{\mu}{M}\right)_{a}+\left(\frac{\mu}{M}\right)_{b}\right]+\cdots+\frac{C^{i j}\left(\overline{r_{a}^{2}}-\overline{r_{b}^{2}}\right)}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}}
\end{aligned}
$$

See example King plot.

- A third isotope $c$ defines a King slope

$$
\begin{aligned}
S_{b c} & =\frac{n_{a c}^{i^{\prime} j^{\prime}}-n_{a b}^{i i^{\prime} j^{\prime}}}{n_{a c}^{i j}-n_{a b}^{i j}} \\
& \rightarrow \frac{C^{i^{\prime} j^{\prime}}}{C^{i j}} \text { if } W^{i^{\prime} j^{\prime}}=W^{i j}=0 \\
& \equiv \frac{\sum_{k}\left[\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{i^{\prime}}-\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{j j}\right]}{\sum_{k}\left[\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{i}-\left\langle\delta\left(\mathbf{r}_{k}\right)\right\rangle_{j}\right]}
\end{aligned}
$$

- If $c$ is replaced by a fourth isotope $d$, then there should be no change in slope.


## Searches for New Physics

There are (at least) three sources of nonlinearity in the King plot:

1. For light atoms such as He and $\mathrm{Li}^{+}$, the term $W^{i j}\left[\left(\frac{\mu}{M}\right)_{a}+\left(\frac{\mu}{M}\right)_{b}\right]$ is not small.
2. There are higher-order corrections to the field shift term.
3. There is an additional electron-neutron interaction due to exchange of light bosons beyond the Standard Model.

Item \#1 can be handled by taking second differences to eliminate the $W^{i j}$ term to form a second-King plot, or by means of generalized King plots, as discussed by Berengut et al. Phys. Rev. Research 2, 043444 (2020), and similarly for item \#2.
Nuclear size uncertainties were recently discussed by Müller et al. Phys. Rev. A 104, L020802 (2021).
Item \#3 is the principal topic of the present work.

## New Light Bosons

- As discussed by Delaunay et al. [PRD 96,093001 (2017)] a possible candidate for dark matter is a light boson that propagates an electron-neutron interaction. It depends on the number of neutrons in the nucleus, and so varies from one isotope to the next, just as the electric field varies with the number of protons. The corresponding hypothetical Yukawa potential is

$$
V_{m_{\phi}}=\frac{N_{I} y_{e} y_{n}}{4 \pi r} e^{-\gamma r}
$$

where $\gamma=1 / \alpha a_{m_{\phi}}$ and $\alpha a_{m_{\phi}}$ is the Compton wavelength for the boson of mass $m_{\phi}, a_{m_{\phi}}=\hbar^{2} /\left(m_{\phi} e^{2}\right)$ is the corresponding Bohr radius, $y_{e} y_{n} /(4 \pi)$ is a coupling constant and $N_{I}$ is the number of neutrons for isotope $I$, analogous to $Z e^{2} /\left(4 \pi \epsilon_{0}\right)$ for then electron-nucleus interaction.

- The boson must be "light" because otherwise the range is too short and the potential looks like a delta-function that is indistinguishable from the regular nuclear potential.
- The useful mass range is therefore from 0 up to about 10 keV .
- If present, it would produce a kink in the King plot.


## Modified King Plot

Ignore for now the $W^{i j}$ quadratic mass-polarization term, and extend the King coordinate to read

$$
n_{a b}^{i j}=V^{i j}+C^{i j} G_{a b}+\Upsilon^{i j} N_{a b}
$$

where

$$
N_{a b}=\frac{N_{a}-N_{b}}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}}, \quad G_{a b}=\frac{\bar{r}_{a}^{2}-\bar{r}_{b}^{2}}{\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}}
$$

and $\Upsilon^{i j}$ denotes the difference in matrix elements of the Yukawa potential

$$
\Upsilon^{i j}=\langle i| V_{m_{\phi}}|i\rangle-\langle j| V_{m_{\phi}}|j\rangle
$$

For the King plot with a third isotope $c$ it is convenient to define

$$
\Delta N_{b c}=N_{a b}-N_{a c}, \quad \Delta G_{b c}=G_{a b}-G_{a c}
$$

Then after some algebra, the change in slope if $c$ is replaced by a fourth isotope $d$ is

$$
\Delta S_{c d}=\frac{C^{i^{\prime} j^{\prime}}}{C^{i j}}\left(\frac{\Upsilon^{i^{\prime} j^{\prime}}}{C^{i^{\prime} j^{\prime}}}-\frac{\Upsilon^{i j}}{C^{i j}}\right)\left(\frac{\Delta N_{b d}}{\Delta G_{b d}}-\frac{\Delta N_{b c}}{\Delta G_{b c}}\right)
$$

- Doesn't work for $\mathrm{Li}^{+}$because the $W^{i j}$ mass-polarization term is too large.
- Cure: take second differences to eliminate the $W^{i j}$ term.


## Second-King Plots

For light ions, the $W^{i j}$ term can be eliminated by forming the second-King coordinate

$$
\begin{equation*}
\kappa_{a b c}^{i j}=\frac{n_{a b}^{i j}-n_{a c}^{i j}}{\left(\frac{\mu}{M}\right)_{b}-\left(\frac{\mu}{M}\right)_{c}} \tag{2}
\end{equation*}
$$

Then, with the definition

$$
\begin{align*}
Q_{a b c}= & \frac{\overline{r_{a}^{2}}}{\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}\right]\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{c}\right]} \\
& +\frac{r_{b}^{2}}{\left[\left(\frac{\mu}{M}\right)_{b}-\left(\frac{\mu}{M}\right)_{a}\right]\left[\left(\frac{\mu}{M}\right)_{b}-\left(\frac{\mu}{M}\right)_{c}\right]} \\
& +\frac{r_{c}^{2}}{\left[\left(\frac{\mu}{M}\right)_{c}-\left(\frac{\mu}{M}\right)_{a}\right]\left[\left(\frac{\mu}{M}\right)_{c}-\left(\frac{\mu}{M}\right)_{b}\right]} \tag{3}
\end{align*}
$$

Eq. (2) becomes

$$
\begin{equation*}
\kappa_{a b c}^{i j}=W^{i j}+C^{i j} Q_{a b c} \tag{4}
\end{equation*}
$$

Factors of $Q_{a b c}-Q_{a b d}$ then cancel from numerator and denominator of the second-King slope defined by

$$
\begin{equation*}
S^{(2)}=\frac{\kappa_{a b c}^{i^{\prime} j^{\prime}}-\kappa_{a b d}^{i^{\prime} j^{\prime}}}{\kappa_{a b c}^{i j}-\kappa_{a b d}^{i j}} \tag{5}
\end{equation*}
$$

leaving $S^{(2)}=\frac{C^{C^{\prime} j^{\prime}}}{C^{i j}}$, which is the same as for the standard King slope. The disadvantage of course is that one needs a sequence five isotopes instead of four to look for a change in slope due to new physics BSM.

## Second-King plot to search for new physics BSM

With an additional boson interaction, the modified second-King coordinate becomes

$$
\begin{equation*}
\kappa_{a b c}^{i j}=W^{i j}+C^{i j} Q_{a b c}+\Upsilon^{i j} N_{a b c} \tag{6}
\end{equation*}
$$

where, in parallel with $Q_{a b c}$

$$
\begin{align*}
N_{a b c}= & \frac{N_{a}}{\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{b}\right]\left[\left(\frac{\mu}{M}\right)_{a}-\left(\frac{\mu}{M}\right)_{c}\right]} \\
& +\frac{N_{b}}{\left[\left(\frac{\mu}{M}\right)_{b}-\left(\frac{\mu}{M}\right)_{a}\right]\left[\left(\frac{\mu}{M}\right)_{b}-\left(\frac{\mu}{M}\right)_{c}\right]}  \tag{7}\\
& +\frac{N_{c}}{\left[\left(\frac{\mu}{M}\right)_{c}-\left(\frac{\mu}{M}\right)_{a}\right]\left[\left(\frac{\mu}{M}\right)_{c}-\left(\frac{\mu}{M}\right)_{b}\right]}
\end{align*}
$$

and $N_{a}, N_{b}$ and $N_{c}$ are the neutron numbers for isotopes $a, b, c$ respectively. Note that $N_{a b c}=0$ if $\left\{N_{a}, N_{b}, N_{c}\right\}$ are all equal, and so it depends only on the change in neutron number.
For brevity, define $\Delta Q_{c d}=Q_{a b c}-Q_{a b d}$ and $\Delta N_{c d}=N_{a b c}-N_{a b d}$. Once again the $W^{i j}$ term cancels from the slope of the second-King plot, resulting in

$$
\begin{align*}
S^{(2)} & =\frac{C^{i^{\prime} j^{\prime}} \Delta Q_{c d}+\Upsilon^{i^{\prime} j^{\prime}} \Delta N_{c d}}{C^{i j} \Delta Q_{c d}+\Upsilon^{i j} \Delta N_{c d}} \\
& \simeq \frac{C^{i^{\prime} j^{\prime}}}{C^{i j}}\left[1+\left(\frac{\Upsilon^{i^{\prime} j^{\prime}}}{C^{i^{\prime} j^{\prime}}}-\frac{\Upsilon^{i j}}{C^{i j}}\right) \frac{\Delta N_{c d}}{\Delta Q_{c d}}\right] \tag{8}
\end{align*}
$$

assuming in the second line that the $\Upsilon^{i j}$ term is small so that the denominator can be expanded into the numerator.

Finally, if the isotope $d$ is replaced by a fifth isotope $e$, then the change in slope is

$$
\begin{equation*}
\Delta S_{d e}^{(2)}=\frac{C^{i^{\prime} j^{\prime}}}{C^{i j}}\left(\frac{\Upsilon^{i^{\prime} j^{\prime}}}{C^{i^{\prime} j^{\prime}}}-\frac{\Upsilon^{i j}}{C^{i j}}\right)\left(\frac{\Delta N_{c e}}{\Delta Q_{c e}}-\frac{\Delta N_{c d}}{\Delta Q_{c d}}\right) \tag{9}
\end{equation*}
$$

This shows a clear factorization into a part determined by the nuclear properties represented by the last factor, and a part determined by the electronic wave function represented by the first two factors. It is a nonzero value for the middle factor that would represent a signal for new physics BSM.

## Sensitivity to New Physics

Assume a 1 Hz uncertainty in each of the four isotope shift measurements required to determine a King slope. The sensitivity is then defined as the value of the boson coupling constant $y_{e} y_{n}$ for which

$$
\Delta S_{d e}^{(2)} / S^{(2)}=\Delta S_{1 \mathrm{~Hz}}^{(2)} / S^{(2)}
$$

where $\Delta S_{1 \mathrm{~Hz}}^{(2)}$ is the uncertainty in $S^{(2)}$ induced by the 1 Hz measurement uncertainties added in quadrature.

Nonrelativistic Eigenvalues


Hylleraas coordinates
(Hylleraas, 1929)

The Hamiltonian in atomic units is

$$
H=-\frac{1}{2} \nabla_{1}^{2}-\frac{1}{2} \nabla_{2}^{2}-\frac{Z}{r_{1}}-\frac{Z}{r_{2}}+\frac{1}{r_{12}}
$$

Expand

$$
\Psi\left(\mathbf{r}_{1}, \mathbf{r}_{2}\right)=\sum_{i, j, k} a_{i j k} r_{1}^{i} r_{2}^{j} r_{12}^{k} e^{-\alpha r_{1}-\beta r_{2}} \mathcal{Y}_{l_{1} l_{2} L}^{M}\left(\hat{\mathbf{r}}_{1}, \hat{\mathbf{r}}_{2}\right) \pm \text { exchange }
$$

where $i+j+k \leq \Omega$ (Pekeris shell).
Diagonalize $H$ in the

$$
\phi_{i j k}=r_{1}^{i} r_{2}^{j} r_{12}^{k} e^{-\alpha r_{1}-\beta r_{2}} \mathcal{Y}_{l_{1} l_{2} L}^{M}\left(\hat{\mathbf{r}}_{1}, \hat{\mathbf{r}}_{2}\right) \pm \text { exchange }
$$

basis set.
to satisfy the variational condition

$$
\delta \int \Psi(H-E) \Psi d \tau=0
$$

## New Variational Techniques

I. Double the basis set

$$
\begin{aligned}
& \text { If } \phi_{i, j, k}(\alpha, \beta)=r_{1}^{i} r_{2}^{j} r_{12}^{k} e^{-\alpha r_{1}-\beta r_{2}} \\
& \text { then } \tilde{\phi}_{i, j, k}=a_{1} \phi_{i, j, k}\left(\alpha_{1}, \beta_{1}\right)+a_{2} \phi_{i, j, k}\left(\alpha_{2}, \beta_{2}\right) \\
& \text { asymptotic inner correlation }
\end{aligned}
$$

II. Include the screened hydrogenic function

$$
\phi_{\mathrm{SH}}=\psi_{1 s}(Z) \psi_{n L}(Z-1)
$$

explicitly in the basis set.
III. Optimize the nonlinear parameters

$$
\begin{aligned}
& \frac{\partial E}{\partial \alpha_{t}}=-2\left\langle\Psi_{\mathrm{tr}}\right| H-E\left|r_{1} \Psi\left(\mathbf{r}_{1}, \mathbf{r}_{2} ; \alpha_{t}\right) \pm r_{2} \Psi\left(\mathbf{r}_{2}, \mathbf{r}_{1} ; \alpha_{t}\right)\right\rangle \\
& \frac{\partial E}{\partial \beta_{t}}=-2\left\langle\Psi_{\mathrm{tr}}\right| H-E\left|r_{2} \Psi\left(\mathbf{r}_{1}, \mathbf{r}_{2} ; \alpha_{t}\right) \pm r_{1} \Psi\left(\mathbf{r}_{2}, \mathbf{r}_{1} ; \alpha_{t}\right)\right\rangle
\end{aligned}
$$

for $t=1,2$, with $\left\langle\Psi_{\mathrm{tr}} \mid \Psi_{\mathrm{tr}}\right\rangle=1$.
$\Psi\left(\mathbf{r}_{1}, \mathbf{r}_{2} ; \alpha_{t}\right)=$ terms in $\Psi_{\text {tr }}$ which depend explicitly on $\alpha_{t}$.
For all states up to $n=10$ and $L=7$, see Drake and Yan, PRA 46, 2378 (1992) and http://drake.sharcnet.ca for downloadable resources.

## Mass Scaling

$$
H=-\frac{\hbar^{2}}{2 M} \nabla_{X}^{2}-\frac{\hbar^{2}}{2 m} \nabla_{x_{1}}^{2}-\frac{\hbar^{2}}{2 m} \nabla_{x_{2}}^{2}-\frac{Z e^{2}}{\left|\mathbf{X}-\mathbf{x}_{1}\right|}-\frac{Z e^{2}}{\left|\mathbf{X}-\mathbf{x}_{2}\right|}+\frac{e^{2}}{\left|\mathbf{x}_{1}-\mathbf{x}_{2}\right|}
$$

Transform to centre-of-mass plus relative coordinates $\mathbf{R}, \mathbf{r}_{1}, \mathbf{r}_{2}$

$$
\begin{aligned}
\mathbf{R} & =\frac{M \mathbf{X}+m \mathbf{x}_{1}+m \mathbf{x}_{2}}{M+2 m} \\
\mathbf{r}_{1} & =\mathbf{X}-\mathbf{x}_{1} \\
\mathbf{r}_{2} & =\mathbf{X}-\mathbf{x}_{2}
\end{aligned}
$$

and ignore centre-of-mass motion. Then

$$
H=-\frac{\hbar^{2}}{2 \mu} \nabla_{r_{1}}^{2}-\frac{\hbar^{2}}{2 \mu} \nabla_{r_{2}}^{2}-\frac{\hbar^{2}}{M} \nabla_{r_{1}} \cdot \nabla_{r_{2}}-\frac{Z e^{2}}{r_{1}}-\frac{Z e^{2}}{r_{2}}+\frac{e^{2}}{\left|\mathbf{r}_{1}-\mathbf{r}_{2}\right|}
$$

where $\mu=\frac{m M}{m+M}$ is the electron reduced mass.

Expand

$$
\begin{aligned}
\Psi & =\Psi_{0}+\frac{\mu}{M} \Psi_{1}+\left(\frac{\mu}{M}\right)^{2} \Psi_{2}+\cdots \\
\mathcal{E} & =\mathcal{E}_{0}+\frac{\mu}{M} \mathcal{E}_{1}+\left(\frac{\mu}{M}\right)^{2} \mathcal{E}_{2}+\cdots
\end{aligned}
$$

The zero-order problem is the Schrödinger equation for infinite nuclear mass

$$
\left\{-\frac{1}{2} \nabla_{\rho_{1}}^{2}-\frac{1}{2} \nabla_{\rho_{2}}^{2}-\frac{Z}{\rho_{1}}-\frac{Z}{\rho_{2}}+\frac{1}{\left|\boldsymbol{\rho}_{1}-\boldsymbol{\rho}_{2}\right|}\right\} \Psi_{0}=\mathcal{E}_{0} \Psi_{0}
$$

The "normal" isotope shift is

$$
\Delta E_{\text {normal }}=-\frac{\mu}{M}\left(\frac{\mu}{m}\right) \mathcal{E}_{0} \quad 2 R_{\infty}
$$

The first-order "specific" isotope shift is

$$
\Delta E_{\text {specific }}^{(1)}=-\frac{\mu}{M}\left(\frac{\mu}{m}\right)\left\langle\Psi_{0}\right| \nabla_{\rho_{1}} \cdot \nabla_{\rho_{2}}\left|\Psi_{0}\right\rangle \quad 2 R_{\infty}
$$

The second-order "specific" isotope shift is

$$
\Delta E_{\text {specific }}^{(2)}=\left(-\frac{\mu}{M}\right)^{2}\left(\frac{\mu}{m}\right)\left\langle\Psi_{0}\right| \nabla_{\rho_{1}} \cdot \nabla_{\rho_{2}}\left|\Psi_{1}\right\rangle \quad 2 R_{\infty}
$$

## ASYMPTOTIC EXPANSION FOR $\left\langle\boldsymbol{\delta}\left(\boldsymbol{r}_{1}\right)+\boldsymbol{\delta}\left(\boldsymbol{r}_{2}\right)\right\rangle$

Core Polarization Model (Dalgarno, Drachman)

- neglect exchange.
- Rydberg electron moves in the field generated by the polarizable core.

$$
V(x)=-\frac{Z-1}{x}+\Delta V(x)
$$



Illustration of the physical basis for the asymptotic expansion method in which the Rydberg electron moves in the field generated by the polarized core.

$$
\Delta V(x)=-\frac{c_{4}}{x^{4}}-\frac{c_{6}}{x^{6}}-\frac{c_{7}}{x^{7}}-\frac{c_{8}}{x^{8}}-\frac{c_{9}}{x^{9}}-\frac{c_{10}}{x^{10}}+\cdots
$$

For example, $c_{4}=\frac{1}{2} \alpha_{1}$

$$
c_{6}=\frac{1}{2}\left(\alpha_{2}-6 \beta_{1}\right)
$$

where
$\alpha_{1}=\frac{9}{2 Z^{4}}$ is the dipole polarizability,
$\alpha_{2}=\frac{15}{Z^{6}}$ is the quadrupole polarizability,
$\beta_{1}=\frac{43}{8 Z^{6}}$ is a nonadiabatic correction.

For the $\delta$-function [Drake, PRA 45, 70 (1992)]

$$
\left\langle\delta\left(\mathbf{r}_{1}\right)\right\rangle=\frac{Z^{3}}{\pi a_{\mu}^{3}}\left[\frac{1}{2}-\frac{62}{Z^{6}}\left\langle x^{-4}\right\rangle_{n L}+\frac{1447}{4 Z^{8}}\left\langle x^{-6}\right\rangle_{n L}+\cdots\right]
$$

and the finite mass correction due to mass polarization is

$$
\Delta\left\langle\delta\left(\mathbf{r}_{1}\right)\right\rangle=\frac{Z^{3}}{\pi a_{\mu}^{3}}\left(\frac{\mu}{M}\right)\left[-\frac{124}{Z^{6}}\left\langle x^{-4}\right\rangle_{n L}+\frac{4789}{2 Z^{8}}\left\langle x^{-6}\right\rangle_{n L}+\cdots\right]
$$

All calculations can be done analytically, using methods of Dalgarno and Stewart (195660 ) and Cohen and Dalgarno (1961-66), especially the "Dalgarno Interchange Theorem."

See G.W.F. Drake, Adv. At. Mol. Opt. Phys. 31, 1 (1993).

$$
\left\langle x^{-4}\right\rangle_{n L}=\frac{16(Z-1)^{4}\left[3 n^{2}-L(L+1)\right]}{n^{5}(2 L+3)(2 L+2)(2 L+1)(2 L)(2 L-1)}
$$

TABLE II. Comparison of values for $\pi(\delta(\mathrm{r})\rangle-4$ (units of $10^{-6}$ a.u.). [Drake, PRA 45,70 (1992)]
Variational calculation

| $L$ | $n$ | Singlet | Triplet | Asymptotic expansion | Difference ${ }^{\text {a }}$ |
| :--- | ---: | :--- | :--- | :--- | :--- |
| 3 | 4 | $-35.1830(1)^{\text {b }}$ | $-35.3973(1)$ | $-35.6(6)$ | $0.3(6)$ |
| 3 | 5 | $-20.0527(3)$ | $-20.2325(4)$ | $-20.3(4)$ | $0.2(4)$ |
| 3 | 6 | $-12.2447(1)$ | $-12.3732(2)$ | $-12.42(27)$ | $0.11(27)$ |
| 3 | 7 | $-7.9536(2)$ | $-8.0446(3)$ | $-8.07(19)$ | $0.07(19)$ |
| 3 | 8 | $-5.43373(4)$ | $-5.4992(1)$ | $-5.52(13)$ | $0.05(13)$ |
| 3 | 9 | $-3.8672(1)$ | $-3.91533(2)$ | $-3.93(10)$ | $0.04(10)$ |
| 3 | 10 | $-2.8453(4)$ | $-2.88182(2)$ | $-2.89(7)$ | $0.03(7)$ |
| 4 | 5 | $-4.88644(2)$ | $-4.88678(1)$ | $-4.8866(22)$ | $0.0000(22)$ |
| 4 | 6 | $-3.1356(1)$ | $-3.1359(1)$ | $-3.1360(25)$ | $0.0002(25)$ |
| 4 | 7 | $-2.0917(3)$ | $-2.0917(3)$ | $-2.0918(22)$ | $0.0001(22)$ |
| 4 | 8 | $-1.4520(1)$ | $-1.4523(1)$ | $-1.4522(18)$ | $0.0001(18)$ |
| 4 | 9 | $-1.042727(1)$ | $-1.04447(1)$ | $-1.0444(14)$ | $0.0000(14)$ |
| 4 | 10 | $-0.77400(1)$ | $-0.77412(2)$ | $-0.7741(11)$ | $0.0000(11)$ |
| 5 | 6 | $-1.0039(2)$ | $-1.0040(2)$ | $-1.00393(1)$ | $-0.0000(2)$ |
| 5 | 7 | $-0.69615(2)$ | $-0.69616(2)$ | $-0.69613(2)$ | $-0.00002(2)$ |
| 5 | 8 | $-0.49416(1)$ | $-0.49415(1)$ | $-0.49414(3)$ | $-0.00002(3)$ |
| 5 | 9 | $-0.3603(2)$ | $-0.3603(2)$ | $-0.36043(3)$ | $0.00010(20)$ |
| 5 | 10 | $-0.26971(7)$ | $-0.26965(6)$ | $-0.26972(2)$ | $0.00004(9)$ |
| 6 | 7 | $-0.26836(1)$ | $-0.26837(1)$ | $-0.268369(3)$ | $0.000004(14)$ |
| 6 | 8 | $-0.19656(1)$ | $-0.19656(2)$ | $-0.196564(1)$ | $0.000004(22)$ |
| 6 | 9 | $-0.14611(4)$ | $-0.14611(3)$ | $-0.146131(1)$ | $0.000021(50)$ |
| 6 | 10 | $-0.11074(1)$ | $-0.11075(1)$ | $-0.110741(1)$ | $-0.000004(15)$ |
| 7 | 8 | $-0.086575(3)$ | $-0.086575(3)$ | $-0.0865752(5)$ | $0.0000002(42)$ |
| 7 | 9 | $-0.0660478(3)$ | $-0.0660483(6)$ | $-0.0660460(4)$ | $-0.0000020(8)$ |
| 7 | 10 | $-0.050881(3)$ | $-0.050881(3)$ | $-0.0508804(3)$ | $-0.0000006(42)$ |
| 8 | 9 |  |  | $-0.0321558(1)$ |  |
| 8 | 10 |  |  |  | $-0.0253141(1)$ |

${ }^{2}$ Difference between the singlet-triplet average and the asymptotic value.

TABLE II. Comparison of values for $\pi(\delta(\mathbf{r})\rangle-4$ (units of $10^{-6}$ a.u.). [Drake, PRA 45, 70 (1992)]
Variational calculation

| $L$ | $n$ | Singlet | Triplet | Asymptotic expansion | Difference $^{\mathrm{a}}$ |
| :--- | ---: | :--- | :--- | :--- | :--- |
| 3 | 4 | $-35.1830(1)^{\mathrm{b}}$ | $-35.3973(1)$ | $-35.6(6)$ | $0.3(6)$ |
| 3 | 5 | $-20.0527(3)$ | $-20.2325(4)$ | $-20.3(4)$ | $0.2(4)$ |
| 3 | 6 | $-12.2447(1)$ | $-12.3732(2)$ | $-12.42(27)$ | $0.11(27)$ |
| 3 | 7 | $-7.9536(2)$ | $-8.0446(3)$ | $-8.07(19)$ | $0.07(19)$ |
| 3 | 8 | $-5.43373(4)$ | $-5.4992(1)$ | $-5.52(13)$ | $0.05(13)$ |
| 3 | 9 | $-3.8672(1)$ | $-3.91533(2)$ | $-3.93(10)$ | $0.04(10)$ |
| 3 | 10 | $-2.8453(4)$ | $-2.88182(2)$ | $-2.89(7)$ | $0.03(7)$ |
| 4 | 5 | $-4.88644(2)$ | $-4.88678(1)$ | $-4.8866(22)$ | $0.0000(22)$ |
| 4 | 6 | $-3.1356(1)$ | $-3.1359(1)$ | $-3.1360(25)$ | $0.0002(25)$ |
| 4 | 7 | $-2.0917(3)$ | $-2.0917(3)$ | $-2.0918(22)$ | $0.0001(22)$ |
| 4 | 8 | $-1.4520(1)$ | $-1.4523(1)$ | $-1.4522(18)$ | $0.0001(18)$ |
| 4 | 9 | $-1.04427(1)$ | $-1.04447(1)$ | $-1.0444(14)$ | $0.0000(14)$ |
| 4 | 10 | $-0.77400(1)$ | $-0.77412(2)$ | $-0.7741(11)$ | $0.0000(11)$ |
| 5 | 6 | $-1.0039(2)$ | $-1.0040(2)$ | $-1.00393(1)$ | $-0.0000(2)$ |
| 5 | 7 | $-0.69615(2)$ | $-0.69616(2)$ | $-0.69613(2)$ | $-0.00002(2)$ |
| 5 | 8 | $-0.49416(1)$ | $-0.49415(1)$ | $-0.49414(3)$ | $-0.00002(3)$ |
| 5 | 9 | $-0.3603(2)$ | $-0.3603(2)$ | $-0.36043(3)$ | $0.00010(20)$ |
| 5 | 10 | $-0.26971(7)$ | $-0.26965(6)$ | $-0.26972(2)<---$ | $0.00004(9)$ |
| 6 | 7 | $-0.26836(1)$ | $-0.26837(1)$ | $-0.268369(3)<---$ | $0.000004(14)$ |
| 6 | 8 | $-0.19656(1)$ | $-0.19656(2)$ | $-0.196564(1)$ | $0.000004(22)$ |
| 6 | 9 | $-0.14611(4)$ | $-0.14611(3)$ | $-0.146131(1)$ | $0.000021(50)$ |
| 6 | 10 | $-0.11074(1)$ | $-0.11075(1)$ | $-0.110741(1)$ | $-0.000004(15)$ |
| 7 | 8 | $-0.086575(3)$ | $-0.086575(3)$ | $-0.0865752(5)$ | $0.0000002(42)$ |
| 7 | 9 | $-0.0660478(3)$ | $-0.0660483(6)$ | $-0.0660460(4)$ | $-0.0000020(8)$ |
| 7 | 10 | $-0.050881(3)$ | $-0.050881(3)$ | $-0.0508804(3)$ | $-0.0000006(42)$ |
| 8 | 9 |  |  | $-0.0321558(1)$ |  |
| 8 | 10 |  |  | $-0.0253141(1)$ |  |
|  |  |  |  |  |  |

[^0]TABLE IV. Comparison of values for the specific-mass correction to $\pi\langle\delta(\mathbf{r})\rangle$ (units of $10^{-6} \mu / \mathrm{M}$ a.u. ). [Drake, PRA 45, 70 (1992)]

| $L$ | $n$ | Variational calculation |  | Asymptotic expansion | Difference ${ }^{\text {a }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Singlet | Triplet |  |  |
| 3 | 4 | $-67(2)$ | -64(2) | -68.8(3.6) | 3.3(4.6) |
| 3 | 5 | -39(1) | -36(1) | -38.9(2.7) | 1.4(3.0) |
| 3 | 6 | -23(1) | -21(2) | -23.7(1.8) | 1.7(2.9) |
| 3 | 7 | -16(1) | -13(1) | -15.3(1.2) | 0.8(1.9) |
| 3 | 8 | $-10(1)$ | -8(2) | -10.5(9) | 1.5(24) |
| 3 | 9 | -7.0(4) | -6.5(2) | -7.4(6) | 0.7(7) |
| 3 | 10 | -5.1(4) | -4.3(4) | -5.5(5) | 0.8(8) |
| 4 | 5 | $-9.3(6)$ | -9.2(6) | -9.61(1) | 0.4(8) |
| 4 | 6 | -6.0(7) | -6.1(2) | -6.13(1) | 0.1(7) |
| 4 | 7 | -3.9(2) | -3.6(2) | -4.08(1) | 0.3(3) |
| 4 | 8 | -3.2(4) | -3.5(7) | -2.828(8) | 0.5(8) |
| 4 | 9 | -2.7(1.0) | $-2.2(5)$ | -2.031(6) | $-0.4(1.1)$ |
| 4 | 10 | -1.6(4) | $-1.5(4)$ | -1.505(5) | $-0.5(6)$ |
| 5 | 6 | -2.0(3) | $-2.1(2)$ | -1.9949(3) | -0.05 (36) |
| 5 | 7 | -1.4(1) | -1.4(1) | -1.3805(4) | -0.02(14) |
| 5 | 8 | -0.9(4) | -1.0(4) | -0.9788 (3) | 0.03(56) |
| 5 | 9 | -0.6(4) | $-0.5(4)$ | -0.7134 (3) | 0.16(56) |
| 5 | 10 | -0.57 (5) | $-0.59(2)$ | $-0.5336(2)$ | $-0.05(6)$ |
| 6 | 7 | -0.56 (7) | -0.55(5) | -0.535 14(2) | $-0.02(8)$ |
| 6 | 8 | $-0.35(2)$ | -0.31(5) | -0.39161(2) | 0.06(5) |
| 6 | 9 | -0.27(1) | -0.27(1) | $-0.29097(2)$ | 0.02(2) |
| 6 | 10 | -0.16 (5) | $-0.17(5)$ | $-0.22042(2)$ | 0.05(8) |
| 7 | 8 | -0.17 (1) | $-0.18(2)$ | $-0.172879(2)$ | $0.003(22)$ |
| 7 | 9 | $-0.131(2)$ | -0.126(6) | -0.131 825(2) | $-0.004(6)$ |
| 7 | 10 | -0.103(3) | -0.103(2) | -0.101 525(2) | $-0.002(4)$ |
| 8 | 9 |  |  | -0.064 253 6(2) |  |
| 8 | 10 |  |  | -0.0505697 (3) |  |

${ }^{\text {a }}$ Difference between the singlet-triplet average and the asymptotic value.

## Results

- Comprehensive survey of all possible King plots for states of $\mathrm{Li}^{+}$up to $n=10$ and $L=7$ (K-states).
- Counting both singlets and triplets, there are $N_{\mathrm{S}}=103$ states in this range.
- The number of possible unique King plots is

$$
\begin{aligned}
N_{\text {King }} & =\frac{1}{8} N_{\mathrm{S}}\left(N_{\mathrm{S}}-1\right)\left[N_{\mathrm{S}}\left(N_{\mathrm{S}}-1\right)-2\right] \\
& =13794378
\end{aligned}
$$

- Assume a nominal 1 Hz accuracy in the isotope shift measurements.

Contributions to the slope of the regular King plots k1, k2, and k3, and the secondKing (super-King) plots $\mathrm{k} 1^{(2)}$ and $\mathrm{k} 2^{(2)}$ for the transition pair $2{ }^{3} P_{1}-10{ }^{3} S_{1} / 2{ }^{3} S_{1}$ $10{ }^{3} S_{1}$ of $\mathrm{Li}^{+}$. The last line is the uncertainty in the slope induced by a $\pm 1 \mathrm{~Hz}$ uncertainty in each of the independent isotope shift measurements.

| Contribution | k 1 | k 2 | k3 | $\mathrm{k}^{(2)}$ | $\mathrm{k}^{(2)}$ |
| :--- | ---: | ---: | ---: | ---: | ---: |
| Infinite mass limit | -3.453226 | -3.453226 | -3.453226 | -3.453226 | -3.453226 |
| $(\mu / M)^{2}+\cdots$ | 4.101441 | 4.446124 | 2.328237 | 0.000153 | 0.000082 |
| $(\mu / M) r_{\mathrm{c}}^{2}$ | 0.000208 | 0.000154 | 0.000456 | 0.000816 | 0.000859 |
| Total | 0.648423 | 0.993052 | -1.124533 | -3.452257 | -3.452285 |
| Nucl. radius uncertainty | $\pm 0.56$ | $\pm 0.31$ | $\pm 0.06$ | $\pm 0.000037$ | $\pm 0.000012$ |
| $\pm 1 \mathrm{~Hz}$ | 0.000002 | 0.000001 | 0.000001 | 0.000009 | 0.000005 |

Transitions with highest sensitivity to electron-neutron interactions for different vales of the Yukawa parameter $\gamma$. The conversion factor from $\gamma$ to boson mass $m_{\phi}$ is $Z \alpha m_{e}=11.1868 \mathrm{keV}$ with $Z=3$.

| $\gamma\left(Z / a_{0}\right)$ | King transitions | $y_{e}{ }^{1}$ |
| :---: | :---: | :---: |
| 0.001 | $2{ }^{1} S-10{ }^{1} S / 2{ }^{3} P-10^{1} S$ | $9.6312 \times 10^{-15}$ |
| 0.002 | $2{ }^{1} S-10{ }^{1} S / 2{ }^{3} P-10{ }^{1} S$ | $9.6465 \times 10^{-15}$ |
| 0.005 | $2{ }^{1} S-10^{1} S / 2{ }^{3} P-10^{1} S$ | $9.7322 \times 10^{-15}$ |
| 0.010 | $2^{3} P-10^{3} S / 2{ }^{1} S-10^{1} S$ | 9.94 |
| 0.020 | $2{ }^{1} P-5{ }^{3} S / 2{ }^{1} P-7{ }^{1} S$ | 1.0 |
| 0.050 | $2{ }^{1} P-5{ }^{3} S / 2{ }^{1} P-7{ }^{1} S$ | 1.204 |
| 0.100 | $2{ }^{1} P-5{ }^{3} S / 2{ }^{1} P-7{ }^{1} S$ | 1.568 |
| 0.200 | $2{ }^{1} S-10^{1} S / 2{ }^{3} P-10^{3} S$ | $2.5312 \times 10^{-1}$ |
| 0.200 | $2{ }^{1} S-10^{3} S / 2{ }^{3} P-10^{3} S$ | $2.5307 \times 10^{-1}$ |
| 0.500 | $2{ }^{3} P-2{ }^{3} S / 1^{1} S-2{ }^{1} S^{\text {a }}$ | $5.4886 \times 10^{-1}$ |
| 1.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $1.2293 \times 10^{-1}$ |
| 2.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $4.4138 \times 10^{-1}$ |
| 5.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $4.8762 \times 10^{-1}$ |
| 10.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $4.6078 \times 10^{-1}$ |
| 20.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $5.5648 \times 10^{-}$ |
| 50.000 | $2^{3} S-2{ }^{3} P / 1{ }^{1} S-2{ }^{1} S$ | $1.9496 \times 10^{-}$ |
| 100.00 | $2^{3} S-2{ }^{1} P / 1{ }^{1} S-2{ }^{1} S$ | $2.7815 \times 10^{-7}$ |

${ }^{\bar{a}}$ For this and the following entries, the strongly forbidden $1{ }^{1} S-2{ }^{3} S$ transition has a slightly lower limit.

Sensitivity of second-King transitions for $\mathrm{Li}^{+}$to NP

$$
\text { for }
$$

Compare with

$$
2^{1} S-10{ }^{1} S / 2{ }^{3} P-10^{1} S
$$



## Conclusions and Discussion

- We have defined and tested a second-King plot method that eliminates secondorder mass shifts, as needed for light ions such as $\mathrm{Li}^{+}$.
- We have carried out high-precision variational calculations for all states of $\mathrm{Li}^{+}$ up to $n=10$ and $L=7$, and identified the ones most sensitive to a putative electron-neutron interaction, useful for boson masses up to about 10 keV .
- The sensitivity is about the same as for recent experiments involving $\mathrm{Yb}^{+}$for boson masses up to about 10 keV .
- The optimum King-plot combination is $2{ }^{1} S-10^{1} S / 2{ }^{3} P-10^{1} S$ at 87.53 nm and 89.87 nm respectively.
- The $7 I-10 H$ transition at 2142 nm is a special case because of strong cancellation of the field shift (independent of $Z$ ).
- The same methods can be applied to helium and other light heliumlike ions.
- Online resources are available at http://drake.sharcnet ca.



[^0]:    ${ }^{2}$ Difference between the singlet-triplet average and the asymptotic value.

