

Exact solution for an accelerated relativistic fluid

Motivation:

Find the exact form of the stress-energy tensor at thermodynamic equilibrium

ECT*, FARH workshop
May 8 2018

THERMODYNAMIC EQUILIBRIUM IN SPECIAL RELATIVITY

General covariant form
of the density operator

$$\hat{\rho} = \frac{e^{-\int d\Sigma_\mu \hat{T}^{\mu\nu} \beta_\nu - \zeta \hat{J}^\mu}}{Z}$$

Equilibrium : $\nabla_\mu (\hat{T}^{\mu\nu} \beta_\nu - \zeta \hat{J}^\mu) = 0 \Rightarrow$

$$\begin{aligned} \nabla_\mu \beta_\nu + \nabla_\nu \beta_\mu &= 0 \\ \partial_\mu \zeta &= 0 \end{aligned}$$

Flat spacetime : $\beta_\mu = \underbrace{b_\mu}_{\text{constant t.l. vector}} + \underbrace{\omega_{\mu\nu} x^\nu}_{\text{constant antisymm. tensor}}$

constant t.l. vector

constant antisymm.
tensor

thermal vorticity

$$\Rightarrow \hat{\rho} = e^{-b_\mu \hat{P}^\mu + \frac{1}{2} \omega_{\mu\nu} \hat{J}^{\mu\nu}}$$

10 constants
10 generators
10 independent
Killing vectors

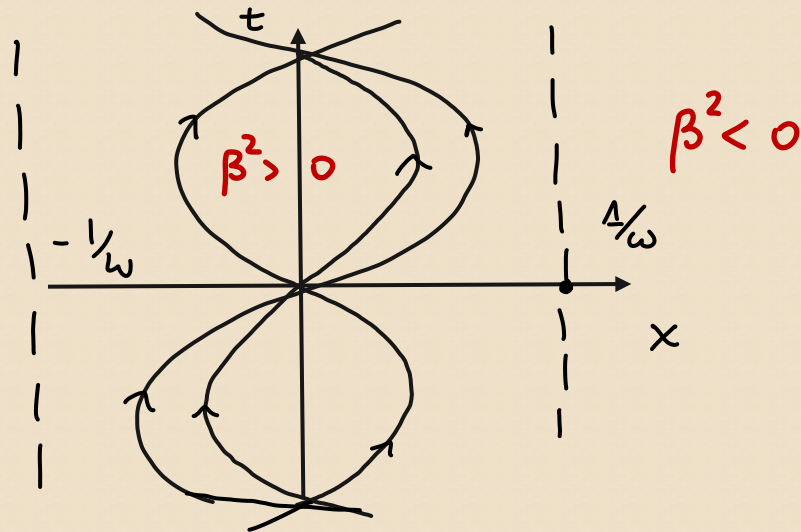
EXAMPLE : ROTATION

$$b = \frac{1}{T_0} (1, \underline{0}) \quad \omega = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & \omega_{\frac{T_0}{2}} & 0 \\ 0 & -\omega_{\frac{T_0}{2}} & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

$$\beta = \frac{1}{T_0} (1, \vec{\omega} \times \vec{r})$$

$$\hat{\rho} = \frac{1}{Z} e^{-\hat{H}_{\frac{T_0}{2}} + \omega_{\frac{T_0}{2}} \hat{J}_z}$$

$$\frac{dx^\mu}{d\tau} = \beta^\mu \quad \text{field lines}$$



NOTE : $\nabla_\mu \beta_\nu + \nabla_\nu \beta_\mu = 0$

implies $\nabla_\mu^\perp u_\nu + \nabla_\nu^\perp u_\mu = 0$

Shear = 0

2nd EXAMPLE: ACCELERATION

$$b = \frac{1}{T_0} (1, \underline{0}) \quad \omega = \begin{pmatrix} 0 & 0 & 0 & a_{\hat{K}_2} \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ -\frac{a}{T_0} & 0 & 0 & 0 \end{pmatrix}$$

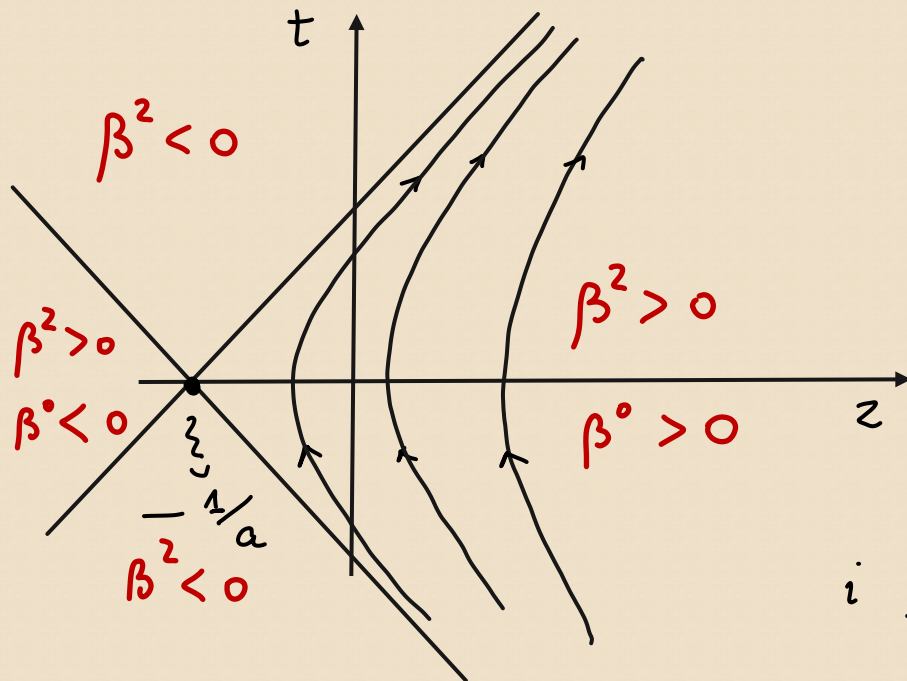
$$\hat{\rho} = \frac{1}{2} e^{-\hat{H}_{\hat{K}_2} + a_{\hat{K}_2} \hat{K}_z}$$

$$\beta^\mu = \frac{1}{T_0} (1 + az, 0, 0, at)$$

$$T = \frac{1}{\sqrt{\beta^2}} = \frac{T_0}{ka} \quad k = \sqrt{(1+az)^2 - t^2 a^2}$$

$$A^\mu = \frac{1}{k^2} (at, 0, 0, 1 + az)$$

$$A^2_{\hat{K}_2} = -\frac{a^2}{T_0^2} = \text{const}$$



NOTE :

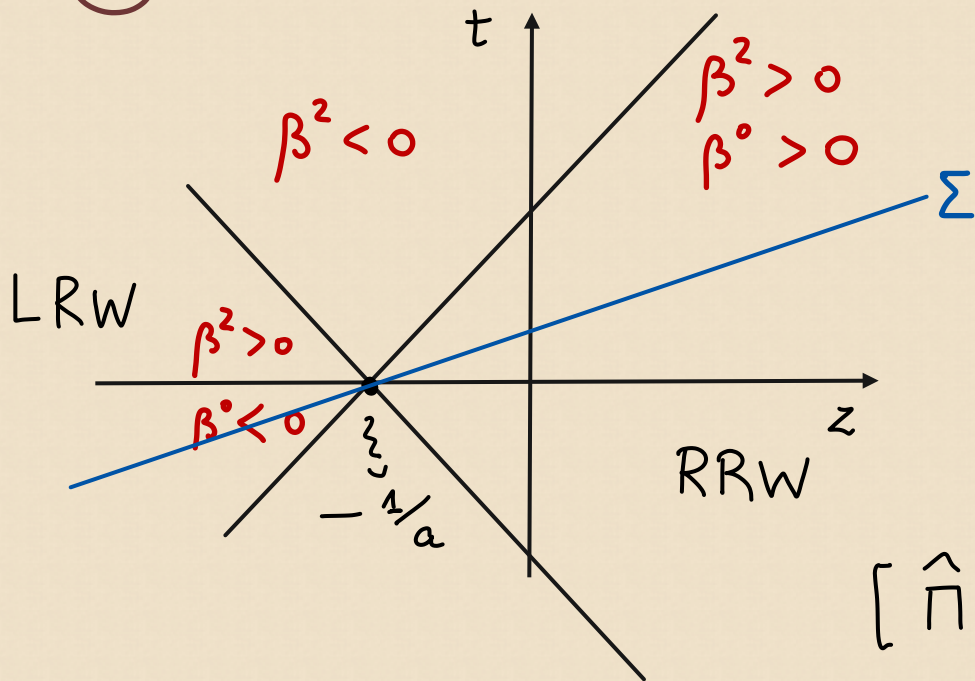
$$i \frac{d\hat{K}_t}{dt} = [\hat{K}_t, \hat{H}] + i \frac{\partial \hat{K}_t}{\partial t} = -i \hat{P}_z + i \hat{P}_z = 0$$

SOLUTION OF ACCELERATED TH. EQ.

F.B., Phys. Rev. D 97, 085013 (2018) 1712.08031

$\text{tr}(\hat{\rho} \text{ Anything})$ can be calculated exactly for a free field theory in the Right Rindler Wedge

①



$$\frac{\hat{H}}{T_0} - \frac{a}{T_0} \hat{K}_z = \int d\Sigma_r \hat{T}^{\mu\nu} \beta_\nu =$$

$$= \int_{z > -1/a} d\Sigma_r \hat{T}^{\mu\nu} \beta_\nu + \int_{z < -1/a} d\Sigma_r \hat{T}^{\mu\nu} \beta_\nu$$

$$\equiv \frac{\hat{\Pi}_R}{T_0}$$

$$- \frac{\hat{\Pi}_L}{T_0}$$

$$[\hat{\Pi}_R, \hat{\Pi}_L] = 0$$

because $\beta = 0$
in $\begin{cases} z = -1/a \\ t = 0 \end{cases}!$

②

$$\hat{\rho} = \frac{e^{-\hat{H}/T_0 + a/T_0 \hat{K}_z}}{Z} = \hat{\rho}_R \otimes \hat{\rho}_L = \frac{e^{-\hat{\Pi}_R/T_0}}{Z_R} \otimes \frac{e^{-\hat{\Pi}_L/T_0}}{Z_L}$$

Therefore: $\langle \hat{O}(x) \rangle_{RRW} = \text{tr}(\hat{\rho}_R \hat{O}(x)) = \frac{\text{tr}(e^{-\hat{\Pi}_R/T_0} \hat{O}(x))}{Z_R}$

③ SOLVE KLEIN-GORDON EQUATION IN RINDLER COORDINATES

L. Crispino, A. Higuchi, G. Matsas, Rev. Mod. Phys. 80 787 (2008)

$$\hat{\psi}(RRW) = \int_0^\infty d\omega d^2 k_T \left[u_{\omega \vec{k}_T}(x) a_R(\omega \vec{k}_T) + \text{h.c.} \right]$$

Real scalar field

$$[a_R(\omega, \vec{k}_T), a_R^\dagger(\omega', \vec{k}'_T)] = \delta(\omega - \omega') \delta^2(\vec{k}_T - \vec{k}'_T)$$

$$u_{\omega \vec{k}_T}(x) = \sqrt{\frac{\sinh(\pi\omega/a)}{4\pi^4 a}} K_{i\frac{\omega}{a}}\left(\frac{m_T e^{a\xi}}{a}\right) e^{-i\omega\tau} e^{i\vec{k}_T \cdot \vec{x}_T}$$

$$t = \frac{e^{a\xi}}{a} \sinh(a\tau)$$

$$z + 1/a = \frac{e^{a\xi}}{a} \cosh(a\tau)$$

NOTE: a_R, a_L can be written as linear combinations of $a(p), a^\dagger(p)$; the so-called Bogoliubov relations

(4) Show that $\hat{\Pi}_L = \frac{1}{2} \int_0^\infty d\omega \int d^2 k_T \omega (a_{R_L}^\dagger a_{R_L} + a_{R_L} a_{R_L}^\dagger)$

\Downarrow

$$\text{tr} (\hat{\rho} a_R^\dagger(\omega, \vec{k}_T) a_R(\omega', \vec{k}'_T)) = \langle a_R^\dagger(\omega, \vec{k}_T) a_R(\omega', \vec{k}'_T) \rangle = \frac{\delta(\omega - \omega') \delta^2(\vec{k}_T - \vec{k}'_T)}{e^{\omega/T_0} - 1}$$

$$\langle a_L^\dagger a_L \rangle = +\infty \quad \text{We don't care}$$

Any quadratic operator in the RRW

$$\langle A \hat{\psi}(x) B \hat{\psi}(x) \rangle = \int_0^\infty d\omega \int d^2 \vec{k}_T f_{\omega, \vec{k}_T}(x) \frac{1}{e^{\omega/T_0} - 1} + f_{\omega, \vec{k}_T}^*(x) \left(\frac{1}{e^{\omega/T_0} - 1} + 1 \right)$$

⑤ RENORMALIZATION (FOR THE FREE FIELD)

Usual prescription for the stress-energy tensor $\langle : \hat{T}^{\mu\nu}(x) : \rangle$

But $: :_{\text{Rindler}} \neq : :_{\text{Minkowski}} \Leftrightarrow |0_R\rangle \neq |0_M\rangle$

Correct prescription $\langle \hat{T}^{\mu\nu} \rangle_{\text{Ren}} = \langle \hat{T}^{\mu\nu} \rangle - \langle 0_n | \hat{T}^{\mu\nu} | 0_n \rangle$

Unruh effect: (Bisognano Wichmann 1975)

$$\langle 0_n | A \hat{\psi} B \hat{\psi} | 0_n \rangle = \frac{\text{tr} \left(e^{\frac{-2\pi}{a}(\hat{H}_R - \hat{H}_L)} A \hat{\psi} B \hat{\psi} \right)}{Z}$$

$$\langle A \hat{\psi} B \hat{\psi} \rangle_{\text{Ren}} = \int_0^\infty d\omega \int d^2 k_r \left[f_{\omega \vec{k}_r}(x) + f_{\omega \vec{k}_r}(x)^* \right] \left(\frac{1}{e^{\omega/T_0} - 1} - \frac{1}{e^{\omega/2\pi/a} - 1} \right)$$

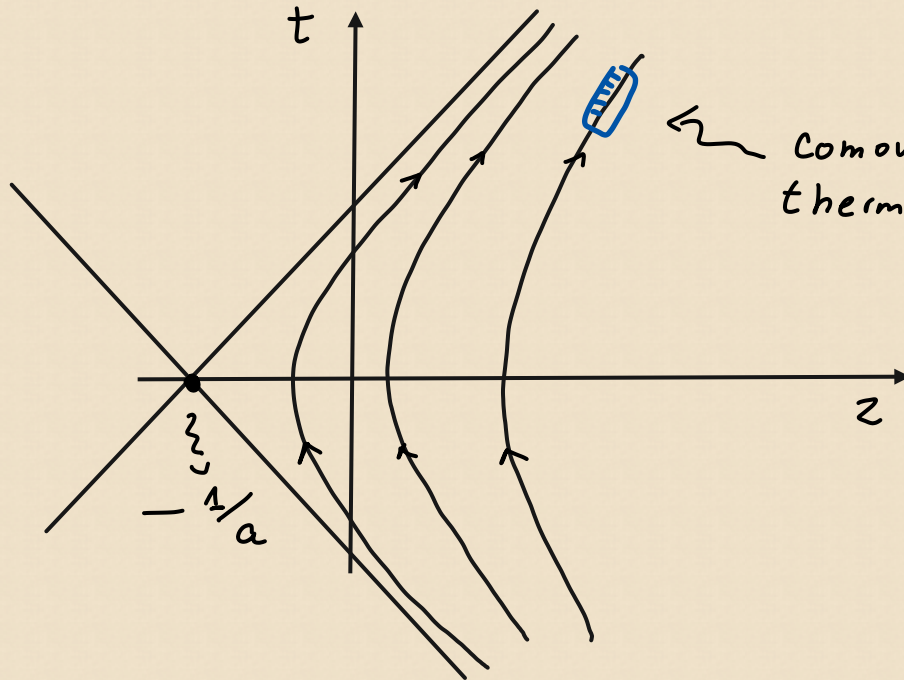
$\Rightarrow !$

Consequence:

$$T_0 = \frac{a}{2\pi} \Rightarrow T_{\text{Ren}}^{\mu\nu} = j_{\text{Ren}}^{\mu} = \dots = O(x)_{\text{Ren}} = 0$$

$$T_0 > \frac{a}{2\pi} = T_v$$

$$\text{Since } \frac{|A|^2}{T^2} = \frac{a^2}{T_0^2} \Rightarrow T_{\text{local}} > \frac{|A|}{2\pi}$$



$$T = \frac{|A|}{2\pi} \text{ in vacuum}$$

From now on

$$\alpha^\mu = \frac{A^\mu}{T}$$

$$\alpha^2 = \frac{|A|^2}{T^2} = |\alpha|^2$$

THERMAL EXPECTATION VALUES

F.B., M. Buzzioli, in preparation

$$\begin{aligned}\hat{T}_{Can}^{\mu\nu} &= \partial^\mu \hat{\psi}^+ \partial^\nu \hat{\psi} + \partial^\nu \hat{\psi}^+ \partial^\mu \hat{\psi} - g_{\mu\nu} (\partial \hat{\psi}^+ \cdot \partial \hat{\psi} - m^2 \hat{\psi}^+ \hat{\psi}) \\ &= \partial^\mu \hat{\psi}^+ \partial^\nu \hat{\psi} + \partial^\nu \hat{\psi}^+ \partial^\mu \hat{\psi} - g_{\mu\nu} \frac{1}{2} \square (\hat{\psi}^+ \hat{\psi})\end{aligned}$$

on-shell

$$u_\mu u_\nu \hat{T}^{\mu\nu} = 2 (u \cdot \partial) \hat{\psi}^+ (u \cdot \partial) \hat{\psi} - \frac{1}{2} \square (\hat{\psi}^+ \hat{\psi})$$

$$\Rightarrow \langle u_\mu u_\nu \hat{T}^{\mu\nu} \rangle_{Ren} \equiv \varepsilon = 2 \left\langle \left(\frac{d\hat{\psi}}{d\tau} \right)^2 \right\rangle_{Ren} - \frac{1}{2} \square \langle \hat{\psi}^+ \hat{\psi} \rangle_{Ren}$$

Can be calculated from the known field expansion
exactly for $m=0$

RESULT:

$$\mathcal{E} = \frac{\pi^2}{15} T^4 \left(1 + \frac{5}{2\pi^2} \alpha^2 \right) - \frac{\pi^2}{15} T_U^4 \left(1 + \frac{5}{2\pi^2} \alpha_U^2 \right)$$

$$\alpha^2 = \frac{|A|^2}{T^2}$$

NOTE : $\alpha_U = 2\pi$ $T_U = \frac{|A|}{2\pi} = \frac{T}{2\pi} \alpha$, the energy density can be written as :

$$\mathcal{E} = \frac{\pi^2}{15} T^4 \left(1 + \frac{5\alpha^2}{2\pi^2} - \frac{11}{16} \frac{\alpha^4}{\pi^4} \right)$$

Vanishes in the Minkowski vacuum $\alpha = 2\pi$

quantum corrections $\alpha = \frac{\hbar A}{c k T}$

The coefficient of α^2 coincides with that calculated

in the quadratic expansion of $\hat{\mathcal{F}}$ { F.B., E. Grossi PRD 92 (2015) 045037
N. Bzrzegala, E. Grossi, F.B. JHEP 1710 (2017) 91

STRESS-ENERGY TENSOR

$$T^{\mu\nu} = (\varepsilon + p) u^\mu u^\nu - p g^{\mu\nu} + A \hat{\alpha}^\mu \hat{\alpha}^\nu$$

$$\alpha^\mu = \frac{A^\mu}{T}$$

$$\varepsilon = \frac{\pi^2}{15} T^4 \left(1 + \frac{5\alpha^2}{2\pi^2} - \frac{11}{16} \frac{\alpha^4}{\pi^4} \right)$$

$$p = \frac{\pi^2}{45} T^4 \left(1 - \frac{5\alpha^2}{\pi^2} + \frac{19\alpha^4}{16\pi^4} \right)$$

$$A = T^4 \left(\frac{\alpha^2}{6} - \frac{2\pi^2}{3} \right)$$

$$\lambda_4 = \frac{A}{gT^2}$$

NOTE These coefficients are specific for the Canonical stress-energy tensor

$$\hat{T}'_{\mu\nu} = \hat{T}_c^{\mu\nu} + \partial_\alpha \partial_\beta \left(\hat{Z}^{\alpha\mu, \beta\nu} + \hat{Z}^{\alpha\nu, \beta\mu} \right)$$

$$\hat{Z}^{\alpha\mu, \beta\nu} = \frac{1}{2} (g^{\beta\alpha} g^{\mu\nu} - g^{\mu\alpha} g^{\beta\nu}) \hat{\psi}^\dagger \hat{\psi}$$

CONCLUSIONS

Quantum corrections can be important

$$W(x, k) = \frac{2}{(2\pi)^4} \int d^4 y \langle : \hat{\psi}^\dagger(x + y/2) \hat{\psi}(x - y/2) : \rangle e^{-iy \cdot k}$$

$$j^\mu(x) = i \langle : \hat{\psi}^\dagger(x) \overleftrightarrow{\partial}^\mu \hat{\psi}(x) : \rangle$$

$$i \langle : \hat{\psi}^\dagger(x) \overleftrightarrow{\partial}^\mu \hat{\psi}(x) : \rangle = \int d^4 k k^\mu W(x, k)$$

$$\Rightarrow j^\mu(x) = \text{Re} \int \frac{d^3 p}{\varepsilon} p^\mu [f_c(x, p) - \bar{f}_c(x, p)]$$

$$f_c(x, p) = \frac{1}{(2\pi)^3} \int \frac{d^3 p'}{2\varepsilon'} e^{i(p-p') \cdot x} \langle \hat{a}^\dagger(p) \hat{a}(p') \rangle$$

$$\langle : \hat{T}_C^{\mu\nu} : \rangle = \text{Re} \left[\int \frac{d^3 p}{\varepsilon} \left(p^\mu p^\nu + \frac{1}{4} (ip^\mu \partial^\nu + ip^\nu \partial^\mu) + \frac{1}{4} (\partial^\mu \partial^\nu - g^{\mu\nu} \square) \right) (f_c(x, p) + \bar{f}_c(x, p)) \right]$$

\hbar

\hbar^2