Trapless (⁸Li) Beta Decay Study Or Hen – MIT



Laboratory for Nuclear Science @

Jefferson-Lab National Accelerator Facility

- Virginia, USA
- Electron beam
 [12 GeV; ~80 uA;
 high polarization]
- 4 experimental halls
- Approved program for coming ~decade
- Leading to EIC



Short-Range Correlations (SRC)













NN Interaction

MMMMMMM Nucleon Sub-Structure



Duer, PRL (2019); Duer, Nature (2018); Hen, Science (2014); Korover, PRL (2014); Subedi, Science (2008); Shneor, PRL (2007); Piasetzky, PRL (2006); Tang, PRL (2003); <u>Review:</u> Hen RMP (2017);

Mapping the nuclear spectral function



A. Schmidt

Probing the NN Interaction



Probing the NN Interaction

A. Schmidt

LABORATORY for NUCLEAR SCIENCE

2018/19 Publications:

- Nature 566, 354 (2019)
- Nature, 560, 617 (2018)
- PRL, In-Print (2019)
- PRL 121, 092501 (2018)
- Physics Letters B 791, 242 (2019)
- Physics Letters B 785, 304 (2018)
- Physics Letters B 780, 211 (2018)
- CPC 42, 064105 (2018)

arXiv: 1811.01823; 1812.08051; 1902.06358; 1805.01981.

Nuclear Bias in Neutrino Oscillations

Nuclear Bias in Neutrino Oscillations

CLAS @ JLab

 $\diamond 4\pi$ acceptance (almost).

 \diamond Charged particles (8-143°):

- P_p > 300 MeV/c
- $P_{\pi} > 150 \text{ MeV/c}$

 \diamond Neutral particles:

- EM calorimeter (8-75°)
- TOF (8-143°)

<u>Goal:</u> Use CLAS data to study E_{beam} reconstruction and vector-current cross-sections for different energies / nuclei.

- Select clean (e,e'p) events (no pions, 2nd protons, ...),
- Reweight by *e-N / v-N* cross-section ratio.
- Analyze as 'neutrino data' (assume unknown E_{beam}),
- Study beam energy reconstruction methods,
- Compare to GENIE predictions,
- Identify regions in phase-space where energy reconstruction and GENIE predictions agree well.

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Energy Reconstruction

1. E_{QE} has worse peak resolution than $E_{Cal.}$ 2.Same tail for E_{QE} & $E_{Cal.}$

Large <u>A</u> Dependence

⁵⁶Fe is predominantly tail.
 ⁵⁶Fe is much worse than ⁴He.

Large <u>E</u> Dependence

Better reconstruction at lower energies.

Large <u>E</u> Dependence

Better reconstruction at lower energies.

Data – Generator Comparisons

Fe	e⁻ Data	ν GENIE
2.2 GeV	26%	62%
4.4 GeV	14%	62%

Fraction of Fe(e, e'p) and Fe(ν, μ^-p) events with E_{Cal} within 5% of E_{beam}

Projected Implications to DUNE

•Reconstruct with vA Neut or eA data

Today: The Standard Model

 $-\frac{1}{2}\partial_{\nu}g^a_{\mu}\partial_{\nu}g^a_{\mu} - g_s f^{abc}\partial_{\mu}g^a_{\nu}g^b_{\mu}g^c_{\nu} - \frac{1}{4}g^2_s f^{abc}f^{adc}g^b_{\mu}g^c_{\nu}g^d_{\mu}g^e_{\nu} +$ $\frac{1}{2}ig_s^2(\bar{q}_i^{\sigma}\gamma^{\mu}\bar{q}_i^{\sigma})g_{\mu}^a + \bar{G}^a\partial^2G^a + g_sf^{abc}\partial_{\mu}\bar{G}^aG^bg_{\mu}^c - \partial_{\nu}W_{\mu}^+\partial_{\nu}W_{\mu}^- M^2 W^+_{\mu} W^-_{\mu} - \frac{1}{2} \partial_{\nu} Z^0_{\mu} \partial_{\nu} Z^0_{\mu} - \frac{1}{2e^2} M^2 Z^0_{\mu} Z^0_{\mu} - \frac{1}{2} \partial_{\mu} A_{\nu} \partial_{\mu} A_{\nu} - \frac{1}{2} \partial_{\mu} H \partial_{\mu} H \frac{1}{2}m_{h}^{2}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} - \frac{1}{2c^{2}}M\phi^{0}\phi^{0} - \beta_{h}[\frac{2M^{2}}{q^{2}} + \frac{1}{2}m_{h}^{2}M\phi^{0}\phi^{0} - \frac{1}{2}m_{h}^{2}M\phi^{0}\phi$ $\frac{2M}{g}H + \frac{1}{2}(H^2 + \phi^0\phi^0 + 2\phi^+\phi^-)] + \frac{2M^4}{g^2}\alpha_h - igc_w[\partial_\nu Z^0_\mu(W^+_\mu W^-_\nu W^{+}_{\mu}W^{-}_{\mu}) - Z^{0}_{\nu}(W^{+}_{\mu}\partial_{\nu}W^{-}_{\mu} - W^{-}_{\mu}\partial_{\nu}W^{+}_{\mu}) + Z^{0}_{\mu}(W^{+}_{\nu}\partial_{\nu}W^{-}_{\mu} W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{w}[\partial_{\nu}A_{\mu}(W_{\mu}^{+}W_{\nu}^{-} - W_{\nu}^{+}W_{\mu}^{-}) - A_{\nu}(W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{+}W_{\mu}^{-})]$ $W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+}) + A_{\mu}(W_{\nu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}^{-}W_{\nu}^{+}W_{\nu}^{-} +$ $\frac{1}{2}g^2W^+_{\mu}W^-_{\nu}W^-_{\mu}W^-_{\nu} + g^2c^2_w(Z^0_{\mu}W^+_{\mu}Z^0_{\nu}W^-_{\nu} - Z^0_{\mu}Z^0_{\mu}W^+_{\nu}W^-_{\nu}) +$ $g^{2}s_{w}^{2}(A_{\mu}W_{\mu}^{+}A_{\nu}W_{\nu}^{-} - A_{\mu}A_{\mu}W_{\nu}^{+}W_{\nu}^{-}) + g^{2}s_{w}c_{w}[A_{\mu}Z_{\nu}^{0}(W_{\mu}^{+}W_{\nu}^{-} - A_{\mu}A_{\mu}W_{\nu}^{+}W_{\nu}^{-})]$ $W^{+}_{\nu}W^{-}_{\mu}) - 2A_{\mu}Z^{0}_{\mu}W^{+}_{\nu}W^{-}_{\nu}] - g\alpha[H^{3} + H\phi^{0}\phi^{0} + 2H\phi^{+}\phi^{-}] \frac{1}{8}g^2\alpha_h[H^4 + (\phi^0)^4 + 4(\phi^+\phi^-)^2 + 4(\phi^0)^2\phi^+\phi^- + 4H^2\phi^+\phi^- + 2(\phi^0)^2H^2]$ $gMW^+_{\mu}W^-_{\mu}H - \frac{1}{2}g\frac{M}{c^2}Z^0_{\mu}Z^0_{\mu}H - \frac{1}{2}ig[W^+_{\mu}(\phi^0\partial_{\mu}\phi^- - \phi^-\partial_{\mu}\phi^0) - \psi^0_{\mu}]$ $W_{\mu} \left(\phi^{0} \partial_{\mu} \phi^{+} - \phi^{+} \partial_{\mu} \phi^{0} \right) + \frac{1}{2} g \left[W_{\mu}^{+} (H \partial_{\mu} \phi - \phi - \partial_{\mu} H) - W_{\mu} (H \partial_{\mu} \phi^{+} + \partial_{\mu} \phi) \right]$ $\phi^{+}\partial_{\mu}H)$] + $\frac{1}{2}g\frac{1}{c_{\nu}}(Z^{0}_{\mu}(H\partial_{\mu}\phi^{0}-\phi^{0}\partial_{\mu}H)-ig\frac{s^{2}_{\mu}}{c_{\nu}}MZ^{0}_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+})+$
$$\begin{split} igs_w MA_\mu (\overline{W}^+_\mu \phi^- - \overline{W}^-_\mu \phi^+) - ig \frac{1-2c_e^2}{2c_e^2} Z^0_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\ igs_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4} g^2 W^+_\mu W^-_\mu [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \end{split}$$
 $\frac{1}{4}g^2 \frac{1}{c^2} Z^0_\mu Z^0_\mu [H^2 + (\phi^0)^2 + 2(2s^2_w - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s^2_w}{c} Z^0_\mu \phi^0 (W^+_\mu \phi^- + \omega^+)^2 \phi^+ \phi^-]$ $W_{\mu}^{-}\phi^{+}) - \frac{1}{2}ig^{2}\frac{s_{w}^{2}}{c}Z_{\mu}^{0}H(W_{\mu}^{+}\phi^{-} - W_{\mu}^{-}\phi^{+}) + \frac{1}{2}g^{2}s_{w}A_{\mu}\phi^{0}(W_{\mu}^{+}\phi^{-} +$
$$\begin{split} & W_{\mu}\stackrel{\phi}{\phi}^+) + \frac{1}{2} i g^2 s_w - \frac{\lambda}{\mu} H'(W_{\mu}^+ \phi^- - W_{\mu}^- \phi^+) - g^2 \frac{s_w}{2} (2c_w^2 - 1) Z_{\mu}^0 A_{\mu} \phi^+ \phi^- - g^1 s_w^2 A_{\mu} A_{\mu} \phi^+ \phi^- - \bar{e}^{\lambda} (\gamma \partial + m_{\lambda}^{\lambda}) e^{\lambda} - \bar{\nu}^{\lambda} \gamma \partial \nu^{\lambda} - \bar{u}_{\lambda}^{\lambda} (\gamma \partial + m_{\lambda}^{\lambda}) u_{\lambda}^{\lambda} - \bar{d}_{\lambda}^{\lambda} (\eta \partial + m_{\lambda}^{\lambda}) u_{\lambda}^{\lambda} - \bar{d}_{\lambda}^{\lambda} (\eta$$
 m_d^{λ} $d_i^{\lambda} + igs_w A_{\mu} [-(\bar{e}^{\lambda}\gamma e^{\lambda}) + \frac{2}{3}(\bar{u}_i^{\lambda}\gamma u_i^{\lambda}) - \frac{1}{3}(\bar{d}_i^{\lambda}\gamma d_i^{\lambda})] + \frac{ig}{4\pi} Z_{\mu}^0 [(\bar{\nu}^{\lambda}\gamma^{\mu}(1 + \bar{\nu}^{\lambda}\gamma e^{\lambda}) + \frac{2}{3}(\bar{u}_i^{\lambda}\gamma u_i^{\lambda}) - \frac{1}{3}(\bar{d}_i^{\lambda}\gamma d_i^{\lambda})] + \frac{ig}{4\pi} Z_{\mu}^0 [(\bar{\nu}^{\lambda}\gamma^{\mu}(1 + \bar{\nu}^{\lambda}\gamma e^{\lambda}) + \frac{2}{3}(\bar{u}_i^{\lambda}\gamma u_i^{\lambda}) - \frac{1}{3}(\bar{d}_i^{\lambda}\gamma d_i^{\lambda})] + \frac{ig}{4\pi} Z_{\mu}^0 [(\bar{\nu}^{\lambda}\gamma^{\mu}(1 + \bar{\nu}^{\lambda}\gamma e^{\lambda}) + \frac{2}{3}(\bar{u}_i^{\lambda}\gamma u_i^{\lambda}) - \frac{1}{3}(\bar{d}_i^{\lambda}\gamma d_i^{\lambda})] + \frac{ig}{4\pi} Z_{\mu}^0 [(\bar{\nu}^{\lambda}\gamma^{\mu}(1 + \bar{\nu}^{\lambda}\gamma e^{\lambda}) + \frac{2}{3}(\bar{\nu}^{\lambda}\gamma e^{\lambda}) +$ $(\gamma^{5})\nu^{\lambda}) + (\bar{e}^{\lambda}\gamma^{\mu}(4s_{w}^{2} - 1 - \gamma^{5})e^{\lambda}) + (\bar{u}_{j}^{\lambda}\gamma^{\mu}(\frac{4}{3}s_{w}^{2} - 1 - \gamma^{5})u_{j}^{\lambda}) +$ $(\bar{d}_{j}^{\lambda}\gamma^{\mu}(1-\frac{8}{3}s_{w}^{2}-\gamma^{5})d_{j}^{\lambda})] + \frac{ig}{2\sqrt{2}}W_{\mu}^{+}[(\bar{\nu}^{\lambda}\gamma^{\mu}(1+\gamma^{5})e^{\lambda}) + (\bar{u}_{j}^{\lambda}\gamma^{\mu}(1+\gamma^{5})e^{\lambda})]$ $\gamma^{5}C_{\lambda\kappa}d_{j}^{\kappa}$] + $\frac{ig}{2\sqrt{2}}W_{\mu}^{-}[(\bar{e}^{\lambda}\gamma^{\mu}(1+\gamma^{5})\nu^{\lambda}) + (\bar{d}_{j}^{\kappa}C_{\lambda\kappa}^{\dagger}\gamma^{\mu}(1+\gamma^{5})u_{j}^{\lambda})] +$ $\frac{ig}{2\sqrt{2}}\frac{m_e^{\lambda}}{M}\left[-\phi^+(\bar{\nu}^{\lambda}(1-\gamma^5)e^{\lambda})+\phi^-(\bar{e}^{\lambda}(1+\gamma^5)\nu^{\lambda})\right]-\frac{g}{2}\frac{m_e^{\lambda}}{M}\left[H(\bar{e}^{\lambda}e^{\lambda})+\right]$ $i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})] + \frac{ig}{2M\sqrt{2}}\phi^{+}[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\gamma^{5})d_{j}^{\kappa}) + m_{u}^{\lambda}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1+\gamma^{5})d_{j}^{\kappa})]$ $\gamma^{5}d_{j}^{\kappa}$] + $\frac{ig}{2M\sqrt{2}}\phi^{-}[m_{d}^{\lambda}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1+\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}] - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}] - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}] - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\star}(1-\gamma^{5})u_{j}^{\kappa}) - m_{u}^{$ $\frac{g}{2}\frac{m_{\lambda}^{\lambda}}{M}H(\bar{u}_{i}^{\lambda}u_{i}^{\lambda}) - \frac{g}{2}\frac{m_{d}^{\lambda}}{M}H(\bar{d}_{i}^{\lambda}d_{i}^{\lambda}) + \frac{ig}{2}\frac{m_{\lambda}^{\lambda}}{M}\phi^{0}(\bar{u}_{i}^{\lambda}\gamma^{5}u_{i}^{\lambda}) - \frac{ig}{2}\frac{m_{d}^{\lambda}}{M}\phi^{0}(\bar{d}_{i}^{\lambda}\gamma^{5}d_{i}^{\lambda}) +$ $\bar{X}^{+}(\partial^{2} - M^{2})X^{+} + \bar{X}^{-}(\partial^{2} - M^{2})X^{-} + \bar{X}^{0}(\partial^{2} - \frac{M^{2}}{2})X^{0} + \bar{Y}\partial^{2}Y +$ $igc_wW^+_\mu(\partial_\mu \bar{X}^0X - \partial_\mu \bar{X}^+X^0) + igs_wW^+_\mu(\partial_\mu \bar{Y}X^w - \partial_\mu \bar{X}^+Y) +$ $igc_wW^-_{\mu}(\partial_{\mu}X^-X^0 - \partial_{\mu}X^0X^+) + igs_wW^-_{\mu}(\partial_{\mu}X^-Y - \partial_{\mu}YX^+) +$ $igc_w Z^0_\mu (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + igs_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) \frac{1}{2}gM[\bar{X}^+X^+H + \bar{X}^-X^-H + \frac{1}{c_*^2}\bar{X}^0X^0H] + \frac{1}{2c_w^2}igM[\bar{X}^+X^0\phi^+ \bar{X}^{-}X^{0}\phi^{-}$] + $\frac{1}{2\pi}igM[\bar{X}^{0}X^{-}\phi^{+} - \bar{X}^{0}X^{+}\phi^{-}] + igMs_{w}[\bar{X}^{0}X^{-}\phi^{+} - \bar{X}^{0}X^{+}\phi^{-}]$ $\bar{X}^{0}X^{+}\phi^{-}] + \frac{1}{2}igM[\bar{X}^{+}X^{+}\phi^{0} - \bar{X}^{-}X^{-}\phi^{0}]$

The Standard Model

 $\frac{2M}{g}H + \frac{1}{2}(H^2 + \phi^0\phi^0 + 2\phi^+\phi^-)] + \frac{2M^4}{c^2}\alpha_b$ $W^+_+W^-_-) - Z^0_+(W^+\partial_-W)$

$$\begin{split} & -\frac{1}{2}\partial_{\nu}g^{a}_{\mu}\partial_{\nu}g^{a}_{\mu} - g_{s}f^{abc}\partial_{\mu}g^{b}_{\nu}g^{b}_{\mu}g^{c}_{\nu} - \frac{1}{4}g^{2}_{s}f^{abc}f^{adc}g^{b}_{\mu}g^{c}_{\nu}g^{d}_{\mu}g^{c}_{\nu} + \\ & \frac{1}{2}ig^{2}_{s}(q^{c}_{\tau}\gamma^{\mu}q^{c}_{\tau})g^{a}_{\mu} + \bar{G}^{a}\partial^{2}G^{a} + g_{s}f^{abc}\partial_{\mu}\bar{G}^{a}G^{b}g^{c}_{\mu} - \partial_{\nu}W^{+}_{\mu}\partial_{\nu}W^{-}_{\mu} - \\ & M^{2}W^{+}_{\mu}W^{-}_{\mu} - \frac{1}{2}\partial_{\nu}Z^{0}_{\mu}\partial_{\nu}Z^{0}_{\mu} - \frac{1}{2c^{2}_{\nu}}M^{2}Z^{0}_{\mu}Z^{2}_{\mu} - \frac{1}{2}\partial_{\mu}A_{\nu}\partial_{\mu}A_{\nu} - \frac{1}{2}\partial_{\mu}H\partial_{\mu}H - \\ & \frac{1}{2}m^{2}_{h}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} - \frac{1}{2c^{2}_{\nu}}M\phi^{0}\phi - \beta_{h}[\frac{2M^{2}}{2d}] \end{split}$$

$$\begin{split} \frac{1}{i \sqrt{2}} \frac{M}{M} &[-\phi^+ (\tilde{\ell}^\lambda (1-\gamma^5) e^\lambda) + \phi^- (\tilde{e}^\lambda (1+\gamma^5) \nu^\lambda)] - \frac{g}{2M} \frac{m^\lambda}{M} [H(\tilde{e}^\lambda e^\lambda) + i \phi^0 (\tilde{e}^\lambda e^\lambda) + \frac{m^\lambda}{M} (\tilde{e}^\lambda e^\lambda) (1+\gamma^5) d^k_j + \frac{m^\lambda}{M} (\tilde{e}^\lambda e^\lambda) d^k_j d^k$$

But..... We still don't know Matter Anti-Matter Asymmetry, Dark Matter, Dark Energy, Black Holes, Gravity,

The Standard Model

 $\begin{array}{l} -\frac{1}{2}\partial_{\nu}g^{a}_{\mu}\partial_{\nu}g^{a}_{\mu} - g_{s}f^{abc}\partial_{\mu}g^{a}_{\nu}g^{b}_{\nu}g^{c}_{\nu} - \frac{1}{4}g^{a}_{s}f^{abc}f^{adc}g^{b}_{\mu}g^{c}_{\nu}g^{d}_{\mu}g^{c}_{\nu} + \\ \frac{1}{2}ig^{2}_{s}(q^{\sigma}_{i}\gamma^{\mu}q^{\sigma}_{j})g^{a}_{\mu} + \bar{G}^{a}\partial^{2}G^{a} + g_{s}f^{abc}\partial_{\mu}\bar{G}^{a}G^{b}g^{c}_{\mu} - \partial_{\nu}W^{a}_{\mu}\partial_{\nu}W^{-}_{\mu} - \end{array}$ $M^2 W^+_{\mu} W^-_{\mu} - \frac{1}{2} \partial_{\nu} Z^0_{\mu} \partial_{\nu} Z^0_{\mu} - \frac{1}{2c^2} M^2 Z^0_{\mu} Z^0_{\mu} - \frac{1}{2} \partial_{\mu} A_{\nu} \partial_{\mu} A_{\nu} - \frac{1}{2} \partial_{\mu} H \partial_{\mu} H \frac{1}{2}m_{h}^{2}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} - \frac{1}{2c^{2}}M\phi^{0}\phi^{0} - \beta_{h}\left[\frac{2M^{2}}{c^{2}}\right]$

The standard model is incomplete; Quarks II C New physics <u>MUST</u> be out there...

 $\frac{2M}{a}H + \frac{1}{2}(H^2 + \phi^0\phi^0 + 2\phi^+\phi^-)] + \frac{2M^4}{a}\alpha_\mu$

 $W^{+}_{*}W^{-}_{*}) - Z^{0}_{*}(W^{+}\partial$

 $\frac{1}{2\sqrt{2}} \frac{M}{M} \left[-\phi^+ (\bar{\nu}^{\lambda}(1-\gamma^5)e^{\lambda}) + \phi^- (\bar{e}^{\lambda}(1+\gamma^5)\nu^{\lambda}) \right] - \frac{g}{2} \frac{m_{\mu}^2}{M} \left[H(\bar{e}^{\lambda}e^{\lambda}) + \phi^- (\bar{e}^{\lambda}(1+\gamma^5)\nu^{\lambda}) \right]$ $i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})] + \frac{ig}{2M\sqrt{2}}\phi^{+}[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\gamma^{5})d_{j}^{\kappa}) + m_{u}^{\lambda}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1+\gamma^{5})d_{j}^{\kappa})]$ $\gamma^{5}d_{j}^{\kappa}] + \frac{ig}{2M\sqrt{2}}\phi^{-}[m_{d}^{\lambda}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1+\gamma^{5})u_{i}^{\kappa}) - m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{i}^{\kappa}]$ $\frac{g}{2}\frac{m_{\star}^{\lambda}}{M}H(\bar{u}_{i}^{\lambda}u_{i}^{\lambda}) - \frac{g}{2}\frac{m_{d}^{\lambda}}{M}H(\bar{d}_{i}^{\lambda}d_{i}^{\lambda}) + \frac{ig}{2}\frac{m_{\star}^{\lambda}}{M}\phi^{0}(\bar{u}_{i}^{\lambda}\gamma^{5}u_{i}^{\lambda}) - \frac{ig}{2}\frac{m_{d}^{\lambda}}{M}\phi^{0}(\bar{d}_{i}^{\lambda}\gamma^{5}d_{i}^{\lambda}) +$ $\bar{X}^{+}(\partial^{2} - M^{2})X^{+} + \bar{X}^{-}(\partial^{2} - M^{2})X^{-} + \bar{X}^{0}(\partial^{2} - \frac{M^{2}}{c^{2}})X^{0} + \bar{Y}\partial^{2}Y$ $igc_wW^+_\mu(\partial_\mu \bar{X}^0X - \partial_\mu \bar{X}^+X^0) + igs_wW^+_\mu(\partial_\mu \bar{Y}X^- - \partial_\mu \bar{X}^+Y) +$ $igc_w W^{\mu}_{\mu}(\partial_{\mu}X^-X^0 - \partial_{\mu}X^0X^+) + igs_w W^{\mu}_{\mu}(\partial_{\mu}X^-Y - \partial_{\mu}YX^+) +$ $igc_w Z^0_\mu (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + igs_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) \frac{1}{2}gM[\bar{X}^+X^+H + \bar{X}^-X^-H + \frac{1}{c_*^2}\bar{X}^0X^0H] + \frac{1}{2c_w}\frac{2c_w}{i}gM[\bar{X}^+X^0\phi^+ - \frac{1}{2c_w}\bar{X}^0\bar{X}^0H]$ $\bar{X}^{-}X^{0}\phi^{-}$] + $\frac{1}{2\pi}igM[\bar{X}^{0}X^{-}\phi^{+} - \bar{X}^{0}X^{+}\phi^{-}] + igMs_{w}[\bar{X}^{0}X^{-}\phi^{+} - \bar{X}^{0}X^{+}\phi^{-}]$ $\bar{X}^{0}X^{+}\phi^{-}] + \frac{1}{2}igM[\bar{X}^{+}X^{+}\phi^{0} - \bar{X}^{-}X^{-}\phi^{0}]$

But..... We still don't know Matter Anti-Matter Asymmetry, Dark Matter, Dark Energy, Black Holes, Gravity,

Tensor Currents in ⁸Li Beta Decay

Searching Under the Lamppost ...

Measuring Everything we can:

- Energy Spectra
- Angular Correlations
- Half-Lives
- Polarizations

... Constraining New Physics **Comparing to Theory and Probing:** • Non V-A Contribution (S • Right-handed Currents (V+A • Massive Neutrinos. • CKM Unitary.

New Physics in β Decay

New Physics in β Decay

Standard Model:

$$H_{\beta} = (\bar{\psi}_{n}\gamma_{\mu}\psi_{p})(C_{V}\bar{\psi}_{e}\gamma^{\mu}\psi_{\nu} + C_{V}'\bar{\psi}_{e}\gamma^{\mu}\gamma_{5}\psi_{\nu}) -(\bar{\psi}_{n}\gamma_{\mu}\gamma_{5}\psi_{p})(C_{A}\bar{\psi}_{e}\gamma^{\mu}\gamma_{5}\psi_{\nu} + C_{A}'\bar{\psi}_{e}\gamma^{\mu}\psi_{\nu}) C_{V} = C_{V}' = 1$$

$$C_A = C'_A = 1.26$$

A-V vs. Tensor Currents

36
A-V vs. Tensor Currents in ⁸Li

Kinematical distributions sensitive to the current:

- Energy distribution of recoiling ion (from two alpha measurement).
- Angle between neutrino and electron.

Resulting Constrain: C_T/C_A < 10% (95% C.L.)





Standard β Decay Experiment







Standard β Decay Experiment





Why Trap?

- Cold and Dilute:
 - No 'smearing'
 - Less interactions
- Well localized vertex
- Isotope selectivity

Standard β Decay Experiment

Produce Radioactive Atoms (Produce, Transport, Neutralize)



Downside of Trapping:

- Complicated experimental setup.
 Limited number of Isotopes.
 Low Statistics.
 Cold and Diluted and D
 - No 'smearing'

Analyze the Data and Compare to		Less interactions
SM Prediction	٠	Well localized vertex

Isotope selectivity

Avoiding the Trap

1. Use nucleus with 'high energy' decays. => ${}^{8}Li \rightarrow e^{-} + \bar{v}_{e} + {}^{8}Be(\rightarrow \alpha + \alpha)$

- Easy to produce using conventional neutron sources.
- Two ~1.5 MeV alphas in the final state.
- 2. Device a 'trap-less' measurement scheme

TPC

Cathode Mesh eeρ. Ground Mesh E = Readout **66666666** 2kV/mm Anode 700V

⁸Li Physics (Tensor vs. V-A)



The OLIVIA Experiment



The OLIVIA Experiment



The OLIVIA Experiment (MIT/HUJI/UM) Optical Lithium V-minus-A Experiment



The OLIVIA Experiment (мпт/нил/им) Optical Lithium V-minus-A Experiment



The OLIVIA Experiment (мпт/нил/им) Optical Lithium V-minus-A Experiment



⁸Li Production

Using a DT generator: ${}^{11}B(n,\alpha)^{8}Li$



⁸Li Production

Using a DT generator: ${}^{11}B(n,\alpha)^{8}Li$



⁸Li Production

Using a DT generator: ${}^{11}B(n,\alpha)^{8}Li$









Roll over image to zoom in

All American 30-Quart Pressure Cooker Canner

by All American ★★★★ × 2,921 customer reviews | 713 answered questions

Price: \$343.00 **vprime**

FREE Delivery Tomorrow

if you order within 9 hrs 45 mins. Details Only 2 left in stock - order soon. Sold by FastSavings ~ and Fulfilled by Amazon. Gift-wrap available.

Eligible for **amazon**smile donation.

Size: 30 qt

10.5 qt	15.5 qt	21.5 qt	25 qt	30 qt
\$232.80	used from \$218.99	\$279.95 • prime	\$299.95	\$343.00 √prime

41.5 qt \$494.95 **√prime**

- The All American 30-quart pressure cooker and canner holds approximately 19 standard regular mouth pint jars or 14 standard regular mouth quart jars; Perfect Canner for all your canning needs!
- Made of durable, hand-cast aluminum with an attractive, easy to clean satin finish; Easy on-off cover; Positive action clamping wing nuts permit easy opening and closing
- Sturdy phenolic top handle; Exclusive "metal-to-metal" sealing system for a steam-tight seal; No gaskets to crack, burn, replace or clean
- Easy to read geared steam gauge; Automatic overpressure release; Settings of 5 psi, 10 psi, and 15 psi
- 19 inches high with 12-1/4-inch inside diameter; made in USA



Currently working on production & transport rate measurements

(1) Produce ⁸Li



(3) Measure the decay products



Readout

(1) Produce ⁸Li



(3) Measure the decay products



Readout











Little MITPC

- 2.8L
- 0.2 10 MeV nuclear recoil
- 4 months of data at Double Chooz far hall
- Now at MIT for OLIVIA

- 60L
- 0.3 20 MeV nuclear recoil
- 7 months of data at Double Chooz near hall
- Now taking data at FNAL

Event Readout for Alpha Track





Energy Reconstruction



Energy Reconstruction



Energy Reconstruction



SRIM & Garfield Simulations





- 1. Find farthest endpoints along x and y directions.
- 2. Find $\theta_{reconstructed}$ using plot of cos(theta) vs. width.
- 3. Find z component of track length for each $z = \frac{1}{2}$

 $\overline{\tan(\theta_{reconstructed})^*\sqrt{x^2+y^2}}$, where *x*, *y*, and *z* are the respective components of the track length (cm).

4. Find track length by taking $\sqrt{x^2 + y^2 + z^2}$



Sensitivity to ⁸Li decay Alphas

'Typical' simulated events @ 200 Torr



Sensitivity to ⁸Li decay Alphas

'Typical' simulated events @ 200 Torr


'Typical' simulated events @ 200 Torr



'Typical' simulated events @ 200 Torr



Ongoing Simulation Development:

- Simulate waveforms.
- Optimize clustering algorithm for alpha



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- Simulate waveforms.
- Optimize clustering algorithm for alpha



76

⁸Li Physics (Tensor vs. V-A)



Resolution Effects



Status and Outlook



Status and Outlook



Status and Outlook

- •MITPC is a great, proven, 3D alpha detector!
- •Based on the reported SNO rates we can measure 10⁷ decays in 2 3 months of data taking.
- •Finalizing simulations to optimize the resolutions and extract the expected C_T/C_A sensitivity.
- •Initial DT feasibility runs planed for the summer.

HAPPY TO COLLABORATE !