

Holographic Signatures of Resolved Cosmological Singularities: Numerical Investigations

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based on

[arXiv:1804.01387](https://arxiv.org/abs/1804.01387)

with N. Bodendorfer & J. Münch

Quantum Gravity meets Lattice QFT
Trento, 3-7 September 2018



Introduction: QG and Holography

Holographic aspects of QG actively investigated.

Main research directions (in LQG)

1. Dual statistical models from 3+0 QG partition function

[Costantino '11; Dittrich, Hnybida '13; Bonzom, Costantino, Livine '15; Dittrich, Goeller, Livine, Riello '17]

2. Spin networks, tensor networks and holographic entanglement entropy

[Han, Hung '16; Han, Huang '17; Chirco, Oriti, Zhang '17]

3. Holography in symmetry-reduced models

[Ashtekar, Wilson-Ewing '08; Bodendorfer, Schäfer, Schliemann '16; Bodendorfer, Mele, Münch '18]

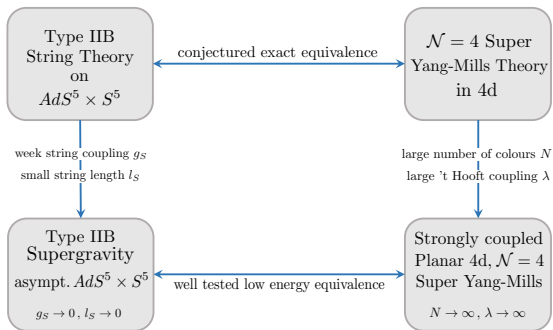
Related work: [Freidel '08; Bodendorfer, Thiemann, Thurn '11; Bodendorfer '15; Freidel, Perez, Pranzetti '16; Smolin '16; Livine '17, ...]

In this talk we will focus on the third direction.

Gauge/Gravity Duality in a Nutshell

[Maldacena '97; Gubser, Klebanov, Polyakov '98; Witten '98]

Maldacena's
original proposal:



Holographic Principle ['t Hooft '93; Susskind '95; Bousso '02; ...]

Duality between gauge and (quantum) gravity theories

- ▶ Tools for QFT (e.g, condensed matter, QCD)
- ▶ Quantum gravity models

- Gauge/Gravity mainly understood in the classical (super)gravity regime
- Singularities in classical gravity
 - ▶ limitations in application of perturbative string theory
 - ▶ no consensus on the fate of singularities in dual QFT
[Hertog, Horowitz '04, '05; Das, Michelson, Narayan, Trivedi '07; Barbón, Rabinovici '11; Craps, Hertog, Turok '12; Smolkin, Turok '12; Engelhardt, Horowitz '15, '16; . . .]
- Progress in singularity resolution from non-perturbative QG

Question

QG resolution of bulk singularity → signatures/improvements in dual QFT?

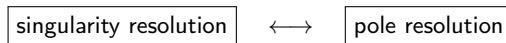
Prototype calculation: effective bulk quantum geometry in a cosmological setting

Plan of the Talk

1. Setup and strategy
2. Holographic signatures of cosmological singularities



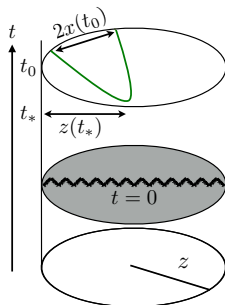
3. Resolved singularities in quantum corrected effective geometries
 - ▶ Simplifications needed for analytic treatment
 - ▶ Otherwise numerics necessary



4. Conclusions and outlook

Setup

[Engelhardt, Horowitz '14; Engelhardt, Horowitz, Hertog '15]



Bulk Kasner-AdS: $ds_5^2 = \frac{1}{z^2} (dz^2 + ds_4^2(t))$

Boundary: $ds_4^2(t) = -dt^2 + \sum_{i=1}^3 t^{2p_i} dx_i^2$, $p_i \in \mathbb{R}$

Kasner conditions: $\sum_i p_i = 1 = \sum_i p_i^2$

$p_i < 0$ geodesics bent towards the singularity

Equal-time 2-point correlator

Geodesic approximation [Balasubramanian, Ross '00]

$$\langle \mathcal{O}(x) \mathcal{O}(-x) \rangle = \exp(-\Delta L_{ren})$$

$\Delta \equiv$ conformal weight of massive scalar operator \mathcal{O}

$L_{ren} \equiv$ renormalized length of bulk geodesic connecting x and $-x$

Renormalized Geodesic Length

fixed z metric

$$\frac{1}{z^2} \left(-dt^2 + a^2(t) dx^2 + \dots \right)$$

----- \rightarrow

boundary metric

$$dt^2 + a^2(t) dx^2 + \dots$$

- Two-point correlator:

$$\langle \mathcal{O} \mathcal{O} \rangle_{CFT} = z^{-2\Delta} \langle \phi \phi \rangle_{bulk}$$

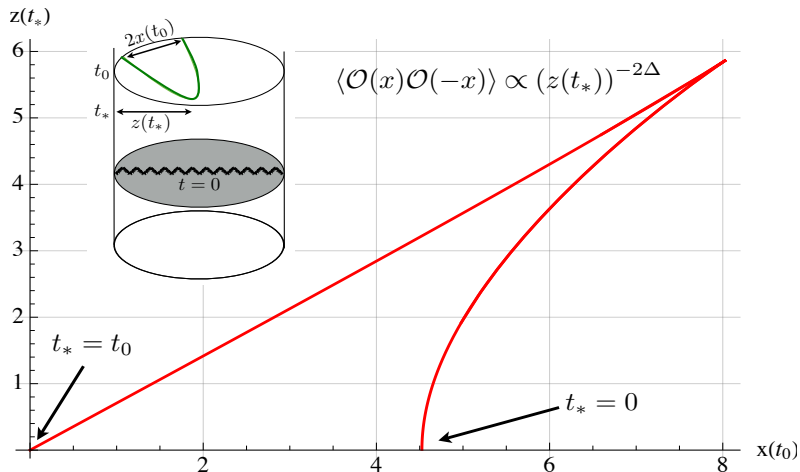
- Geodesic approximation:

$$\langle \phi \phi \rangle_{bulk} = \exp(-\Delta L) \quad , \quad L \stackrel{\epsilon \rightarrow 0}{\approx} 2 \log(2z(t_*)) - 2 \log \epsilon$$

$$\langle \mathcal{O}(-x) \mathcal{O}(x) \rangle_{CFT} = \exp(-\Delta L_{ren}) = (2z(t_*))^{-2\Delta}$$

Holographic Signature of Cosmological Singularity

[Engelhardt, Horowitz '14; Engelhardt, Horowitz, Hertog '15]



Finite distance pole in two-point correlator

Classical metric

$$ds_5^2 = \frac{1}{z^2} (dz^2 + ds_4^2(t)) \quad , \quad ds_4^2(t) = -dt^2 + \sum_{i=1}^3 t^{2p_i} dx_i^2$$

- no large curvatures associated with z -direction
- curvature singularity in t -direction



Quantum corrected metric

$$ds_5^2 = \frac{1}{z^2} (dz^2 + ds_4^2(t, z)) \quad , \quad ds_4^2(t, z) = -dt^2 + a(t, z)^2 dx^2 + \dots$$

- z -direction classical
- effective QG for 4d part s.t. singularity resolved

In principle effective metric from full 5d quantum Einstein eqs. \rightarrow here ansatz

- singularity resolution in $ds_4^2 \rightarrow$ singularity-free ds_5^2
(possible z-dependence of ds_4^2 must be small!)
- solution of 5d Einstein eqs. up to quantum corrections
- bulk QG effects relevant at 5d-Planck scale
- classical Kasner boundary metric
- quantum bounce with Kasner transitions
[Gupt, Singh '12; Ashtekar, Wilson-Ewing '08; Wilson-Ewing '17]

Quantum Corrected Metric: A First Example

Quantum corrected metric [Bodendorfer, Schäfer, Schliemann '16]

$$ds_5^2 = \frac{1}{z^2} (dz^2 + ds_4^2(t)) \quad , \quad a(t) = \frac{a_{\text{ext}}}{\lambda^p} (t^2 + \lambda^2)^{p/2}$$

Two main simplifications:

1. λ related to 4d bulk Planck scale
2. No Kasner transitions

Renormalized length and two-point correlator

Geodesic eqs. analytically solvable:

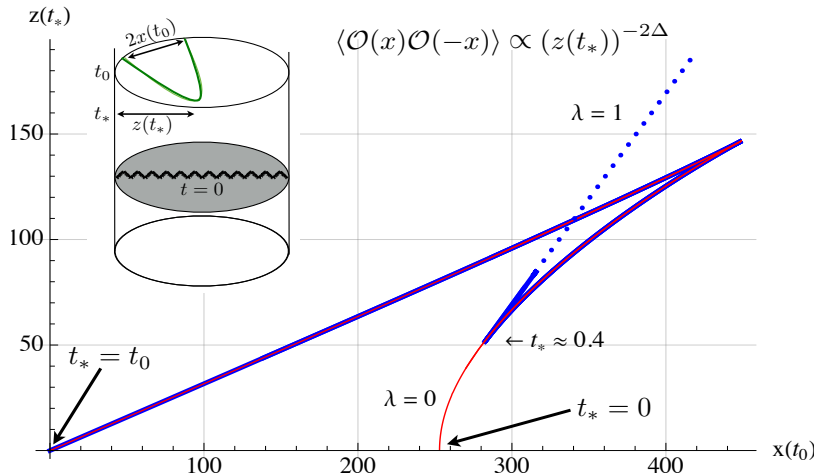
$$\pm s(z = \epsilon) \stackrel{\epsilon \rightarrow 0}{\equiv} \log(2z(t_*)) - \log(\epsilon) \quad \epsilon = \text{boundary regulator}$$

and

$$L_{\text{ren}} = 2 \log(2z(t_*)) \quad \Rightarrow \quad \langle \mathcal{O}(x) \mathcal{O}(-x) \rangle = (2z(t_*))^{-2\Delta}$$

Holographic Signature of Resolved Cosmological Singularity

[Bodendorfer, Schäfer, Schliemann '16]



No finite distance pole in two-point correlator!

Drop Simplification 1: 5d Planck Scale

Quantum corrected metric [Bodendorfer, Mele, Münch '18]

$$ds_5^2 = \frac{1}{z^2} \left(dz^2 - dt^2 + \sum_i a_i^2(t, z) dx_i^2 \right) , \quad a_i^2(t, z) = \frac{a_{\text{ext}}^2}{\lambda^{2p_i}} (t^2 + z^2 \lambda^2)^{p_i}$$

- ds_5^2 not singular at $t = 0$

$$R_{\text{Kretschmann}}^{(5)} = z^4 R_{\text{Kretschmann}}^{(4)} + \dots , \quad R_{\text{Kretschmann}}^{(4)} \sim \lambda^{-4}$$

$$\Rightarrow \text{effective 5d scale: } \lambda \rightarrow z\lambda \quad \text{s.t.} \quad R_{\text{Kretschmann}}^{(5)} = \mathcal{O}(\lambda_{5d}^{-4})$$

- z -dependence of ds_4^2 small (order of λ) and $\partial_z a \rightarrow 0$ in zeroth order in λ
- classical Kasner-AdS boundary metric recovered ($\frac{a_{\text{ext}}}{\lambda^p} \rightarrow 1$, $\lambda, z \rightarrow 0$)
- symmetries:
 - ▶ translation (e.g., for us in x -direction)
 - ▶ scaling: $t \mapsto \Lambda t$, $z \mapsto \Lambda z$, $x \mapsto \Lambda^{1-p_x} x$ ($\Lambda \in \mathbb{R}$)
- no Kasner transitions

Boundary Pole Resolution

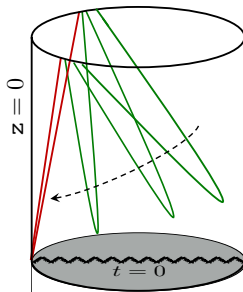
Boundary metric is singular

$$a_i^2(t, z) = \frac{a_{ext}^2}{\lambda^{2p_i}} (t^2 + \lambda^2 z^2)^{p_i} \xrightarrow{z \rightarrow 0} \frac{a_{ext}^2}{\lambda^{2p_i}} t^{2p_i}$$

Classical Kasner-AdS: [Engelhardt, Horowitz, Hertog '15]

$$z_* \rightarrow 0 \text{ as } t_* \rightarrow 0$$

family of space-like geodesics approaches a light-like boundary geodesic probing high-curvature regime



Quantum corrected bulk: **boundary geodesics isolated** (more later).

Solving Geodesic Equations

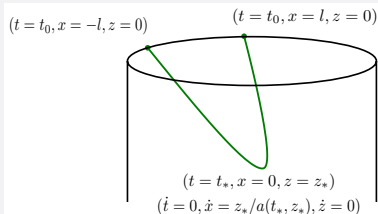
geodesic Eqs. not decoupled \longrightarrow numerics necessary!

Strategy

boundary-value problem



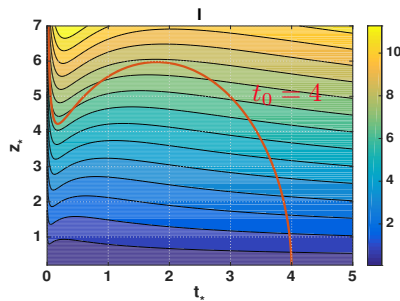
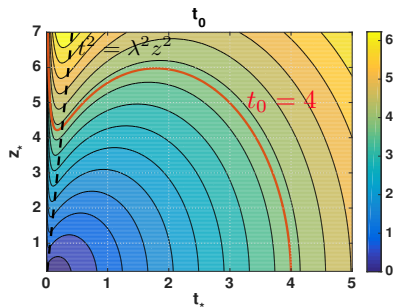
initial-value problem



- initial values at the turning point only in terms of t_* and z_*
- solve IVP numerically and find $t_0(t_*, z_*)/l(t_*, z_*)$
- we are interested in fixed t_0 -values and vary $l \rightarrow$ find relation $z_*(l)$

Solving Geodesic Equations

Key point: relate turning point data (t_*, z_*) with boundary data (t_0, l)



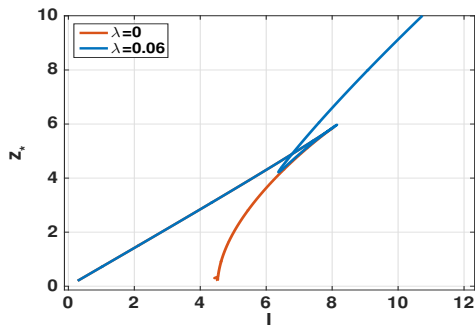
$$p = -1/4, \quad \lambda = 0.06, \quad a_{ext} = \lambda^p$$

Note

- Classical case: bulk geodesics approaching a null boundary geodesic ($z_* \rightarrow 0$ for $t_* \rightarrow 0$)
- Quantum case: z_* grows as $t_* \rightarrow 0 \Rightarrow$ null boundary geodesic not limit of bulk geodesics

Solving Geodesic Equations

Reading off the values of l along t_0 -level line, we get the relation $z_*(l)$:



$$p = -1/4 \quad , \quad a_{ext}/\lambda^p = 1$$

$$\text{---} \quad \lambda = 0$$

$$\text{---} \quad \lambda = 0.06$$

last step: check $L_{ren} \sim \log(z(t_*)) \quad \Rightarrow \quad \langle \mathcal{O}(-x)\mathcal{O}(x) \rangle \propto (z(t_*))^{-2\Delta}$

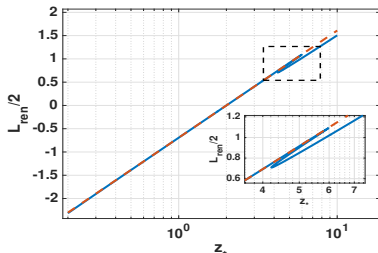
Renormalized Geodesic Length & Two-point Correlator

Renormalization of geodesic length

- Conformal boundary ($z = 0$) lies at infinity \rightarrow geodesic length diverges
- Renormalization procedure needed:
 1. Evaluate geodesic up to $\epsilon = z_{UV}/t \ll 1$
 2. Subtract divergent contribution ($\epsilon \rightarrow 0$)

$$L_{ren} = L - L_0 \quad , \quad L_0 = 2 \log \left(\left| \frac{z_{UV}}{t} \right| \right)$$

Two-point correlator



$$\langle \mathcal{O}(-l) \mathcal{O}(l) \rangle \propto e^{-\Delta L_{ren}} = (z_*)^{-2\Delta}$$

+

$z_* \neq 0$ for finite non-zero l

\Downarrow

pole resolved

Drop Simplification 2: Inclusion of Kasner Transitions

Inclusion of Kasner transitions highly non trivial:

- Find a metric satisfying all requirements discussed before not straightforward
- We provided two examples of quantum corrected metrics partially satisfying them



pole resolution qualitatively the same!

Summary

- Bulk singularity resolution $\overset{dual}{\longleftrightarrow}$ boundary finite-distance pole resolution
- Application of LQG in Gauge/Gravity-framework possible
- New tools for holography/improvements in dual field theory

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Future work

- 5d-Loop quantization
- Look for independent field theory calculations, e.g. via lattice QFT
- Generalize to other spacetimes, e.g. black holes (work in progress)
- Long term goal: finite N gauge theory

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Thank you for your attention!

Quantum corrected metric

$$ds_5^2 = \frac{1}{z^2} (dz^2 - dt^2 + a(t, z)^2 dx^2 + \dots)$$

with

$$a(t, z)^2 = \frac{a_{\text{ext}}^2}{\lambda^{2p}} (t^2 + \lambda^2 z^2)^p \exp \left[2\Delta p \sinh^{-1} \left(\frac{t}{z\lambda} \right) \right] , \quad \Delta p \in \mathbb{R}$$

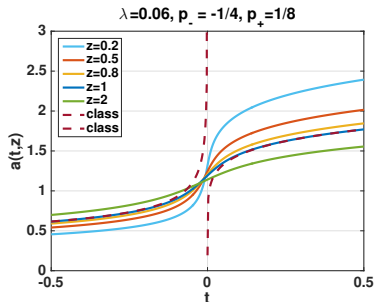
- ds_4^2 : $a(t, z) = \frac{a_{\text{ext}}}{\lambda^p} (t^2 + \lambda^2 z^2)^{p/2} \exp \left[2\Delta p \sinh^{-1} \left(\frac{t}{\lambda} \right) \right]$
 - limiting curvature mimetic gravity [Chamseddine, Mukhanov '16]
 - toy model effective homog. and isotr. LQG [Bodendorfer, Schäfer, Schliemann '17]
- $\lambda \rightarrow z\lambda$ (quantum effects at 5d Planck scale)
- implement transition $p_- = p - \Delta p \rightarrow p_+ = p + \Delta p$
- $\Delta p = 0 \rightarrow$ 5d Planck scale quantum corrected metric with no transitions

Inclusion of Kasner Transitions

For $\left| \frac{t}{z\lambda} \right| \gg 1$:

$$a(t, z)^2 \simeq \frac{a_{\text{ext}}^2}{\lambda^{2p}} \left(\frac{2}{\lambda z} \right)^{\pm 2\Delta p} t^{2p_{\pm}}, \quad p_{\pm} = p \pm \Delta p$$

\Rightarrow **transition** $p_- \rightarrow p_+$ at the bounce ($t=0$).



boundary 4d classical Kasner

✗

solution of 5d Einstein eqs. up to $\mathcal{O}(\lambda)$

✗

Kretschmann condition $R_{\text{Kretsch.}}^{(5)} = \mathcal{O}(\lambda_{5d}^{-4})$

✓

Qualitative LQC behavior

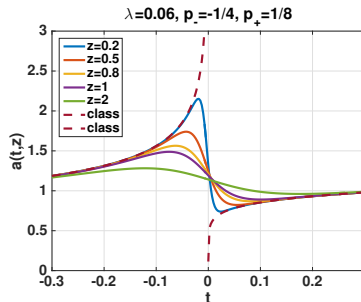
✓

symmetries (translation, scaling)

✓

Alternative Metric for Kasner Transitions

$$a(t, z)^2 = \frac{a_{\text{ext}}^2}{\lambda^{2p}} (t^2 + \lambda^2 z^2)^p \exp \left[2\Delta p \sinh^{-1} \left(\frac{t}{z\lambda} \right) \right] \cdot (\lambda z)^{2\Delta p \tanh \left(\frac{t}{z\lambda} \right)}$$



boundary 4d classical Kasner



solution of 5d Einstein eqs. up to $\mathcal{O}(\lambda)$



qualitative LQC behavior



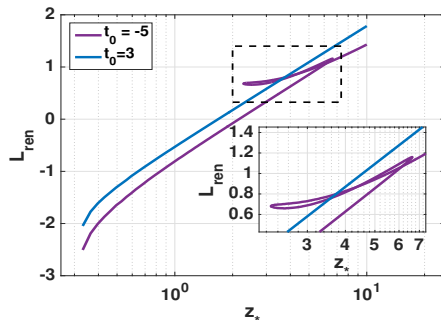
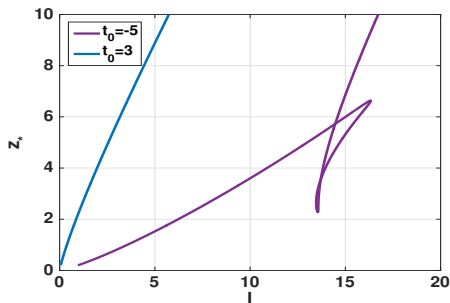
Kretschmann condition $R_{\text{Kretsch.}}^{(5)} = \mathcal{O}(\lambda_{5d}^{-4})$



Numerical Results with Kasner Transitions

Two kinds of solutions:

- starting at $t_0 < 0$, bent towards the resolved singularity and passing it
- starting at $t_0 > 0$, bent away from the resolved singularity



Finite-distance pole in two-point correlator resolved!

Numerical Checks

cut-off dependence ?

No Kasner Transitions

Z_{UV}	$\exp(-L_{ren})$	$\exp(-L_{ren})$
0.05	0.761810	0.917559
0.06	0.761841	0.917597
0.07	0.761860	0.917661
0.08	0.761896	0.917716
0.09	0.761943	0.917783
0.1	0.761972	0.917897

$$\lambda = 0.06, p = -\frac{1}{4}$$

Kasner Transitions

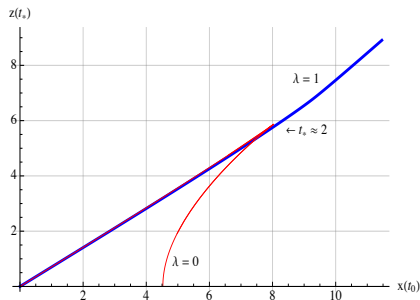
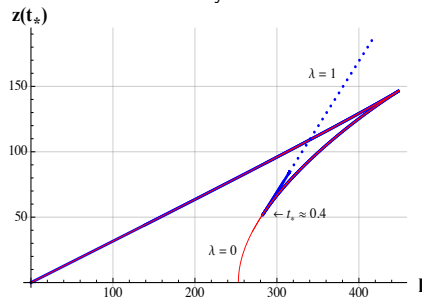
Z_{UV}	$\exp(-L_{ren})$	$\exp(-L_{ren})$
0.05	0.367113	0.308506
0.06	0.367149	0.308554
0.07	0.367185	0.308602
0.08	0.367229	0.308661
0.09	0.367267	0.308714
0.1	0.367314	0.308774

$$\lambda = 0.06, p = -\frac{1}{16}, \Delta p = \frac{3}{16}$$

Renormalization procedure independent of z_{UV} within 0.1%

Comparison with Analytical Results: No z -dependence in ds_4^2

Analytic



Numeric

